

Light detection in neutrino and dark matter (liquid) detectors

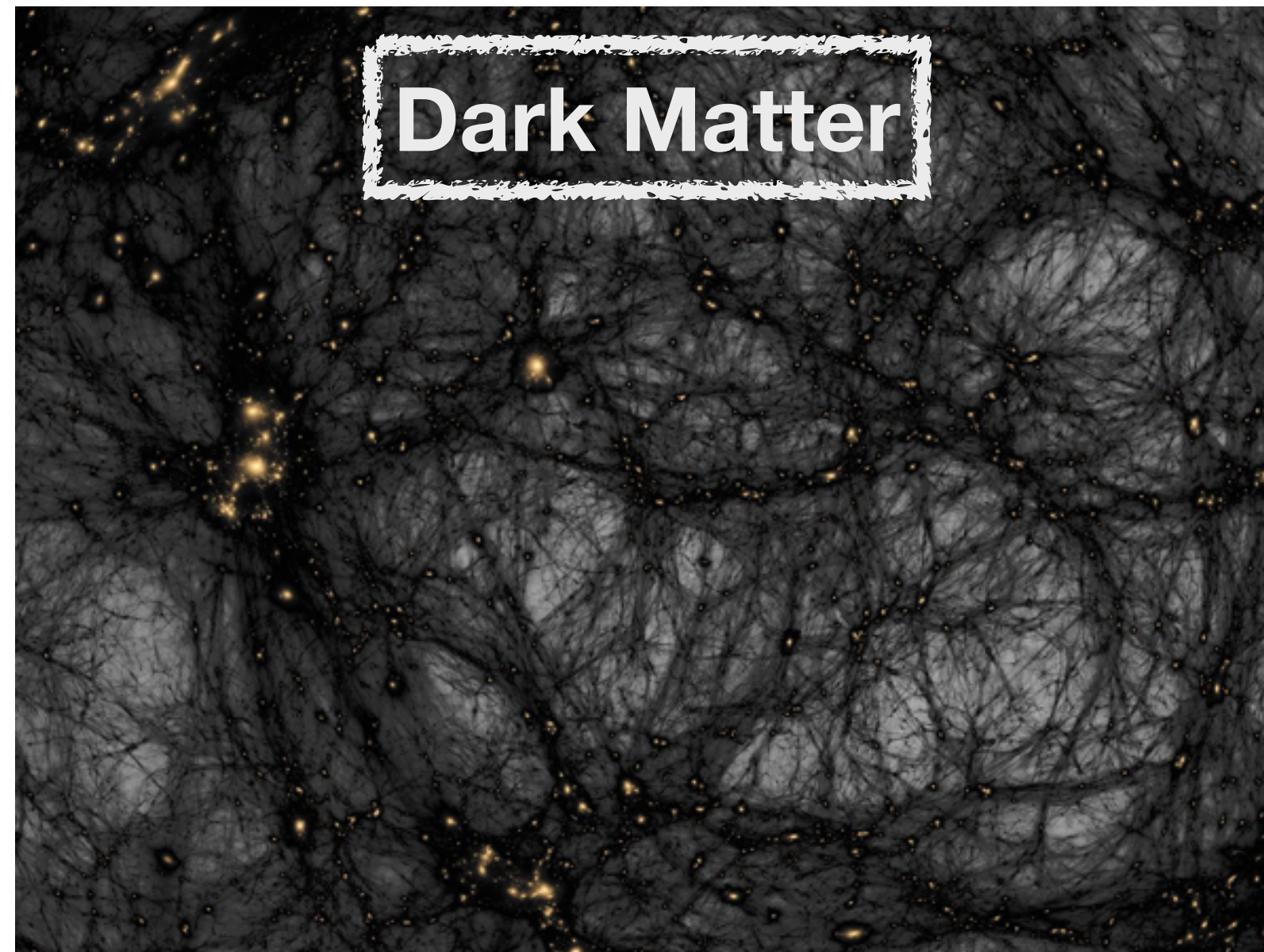
Roxanne Guénette

MANCHESTER
1824

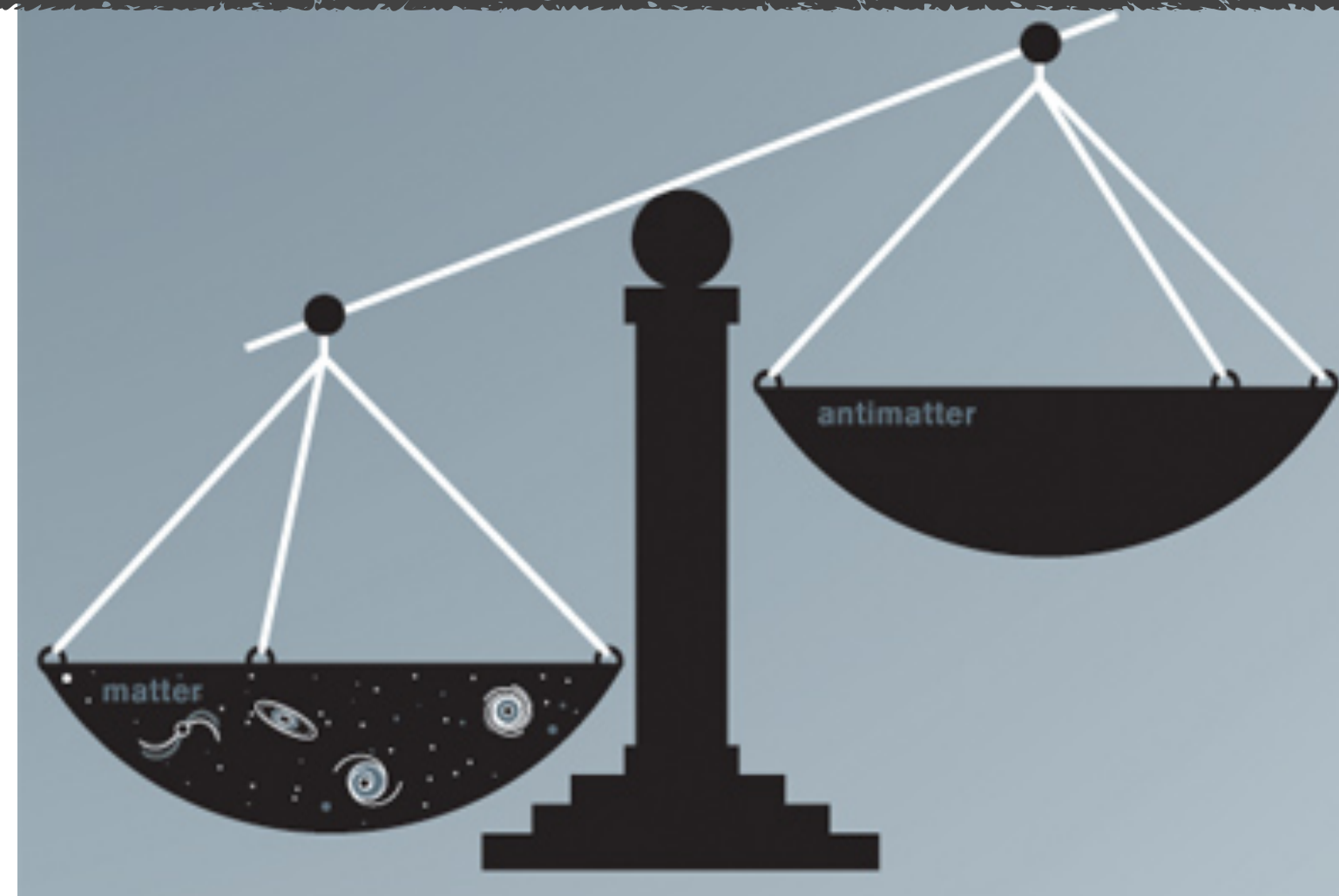
The University of Manchester

6th International Workshop on New Photon-Detectors
20 November 2024

The Big Questions of Particle Physics



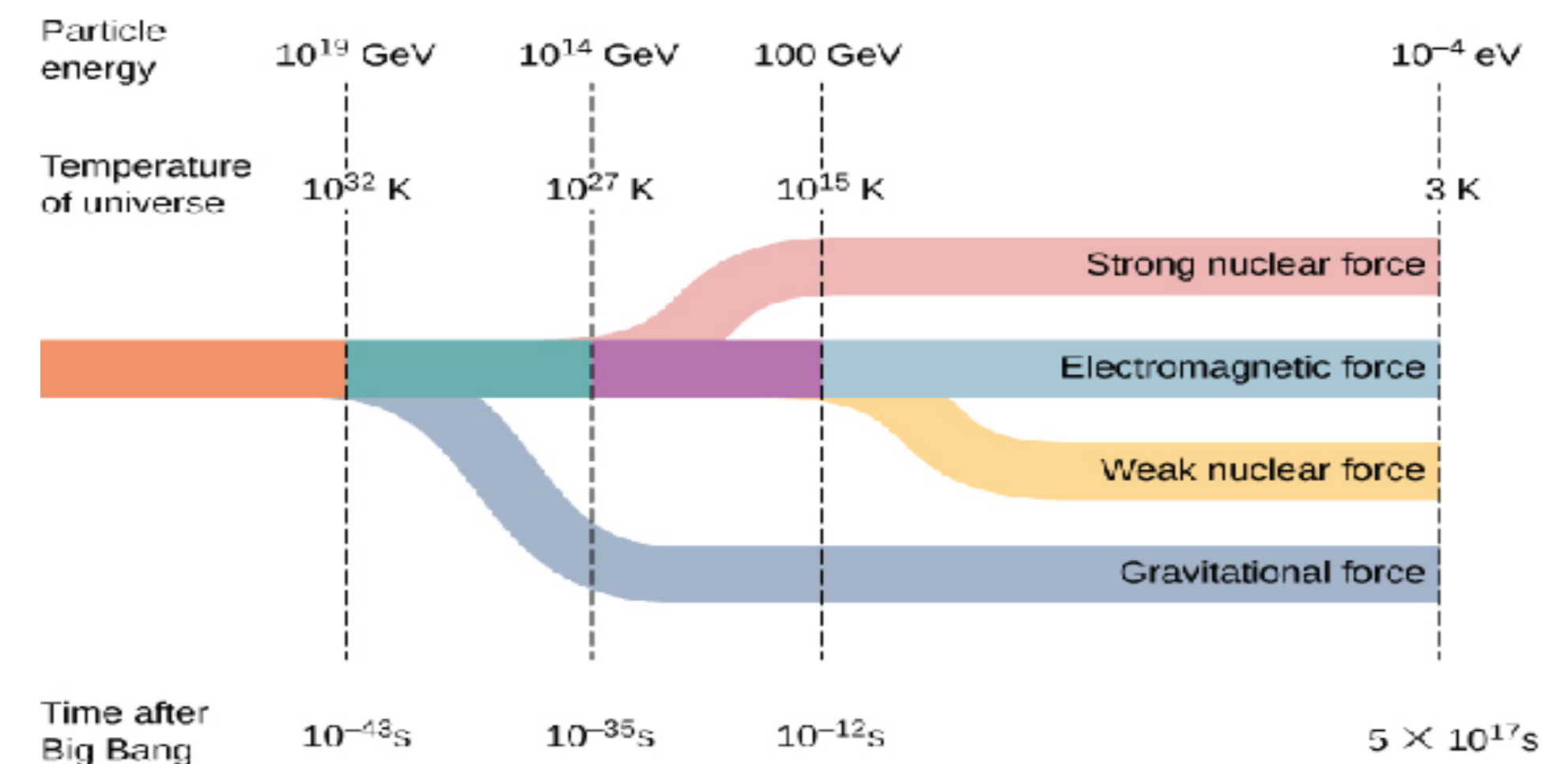
Matter/antimatter asymmetry



New Particles?

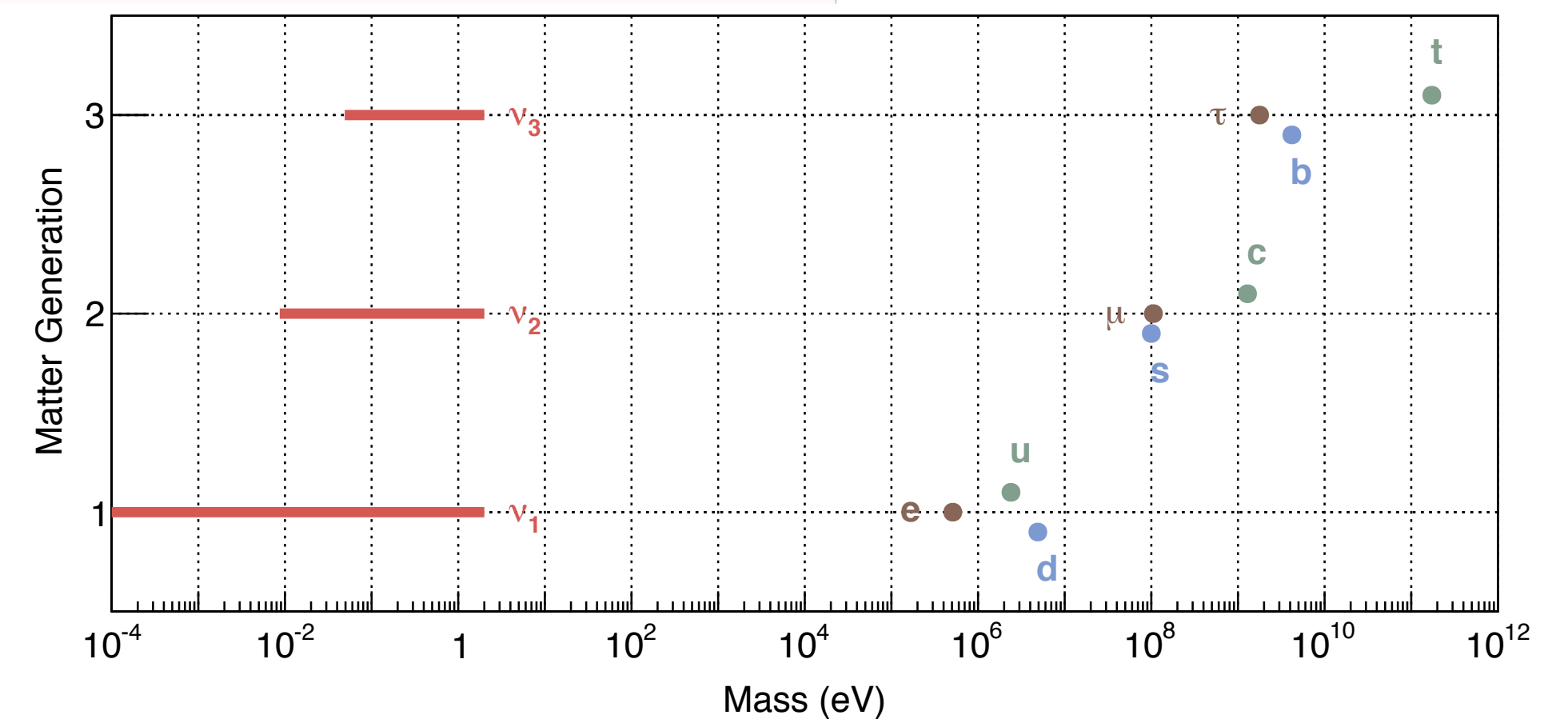
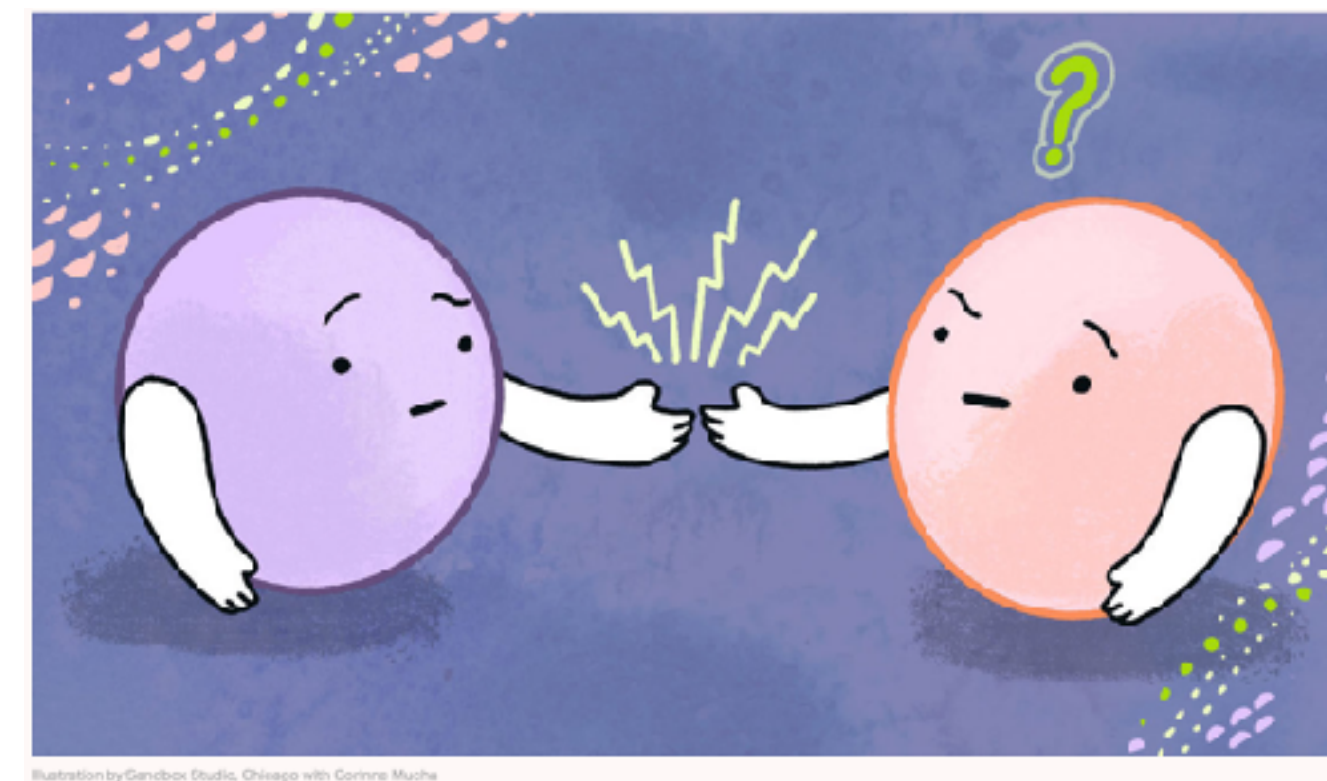
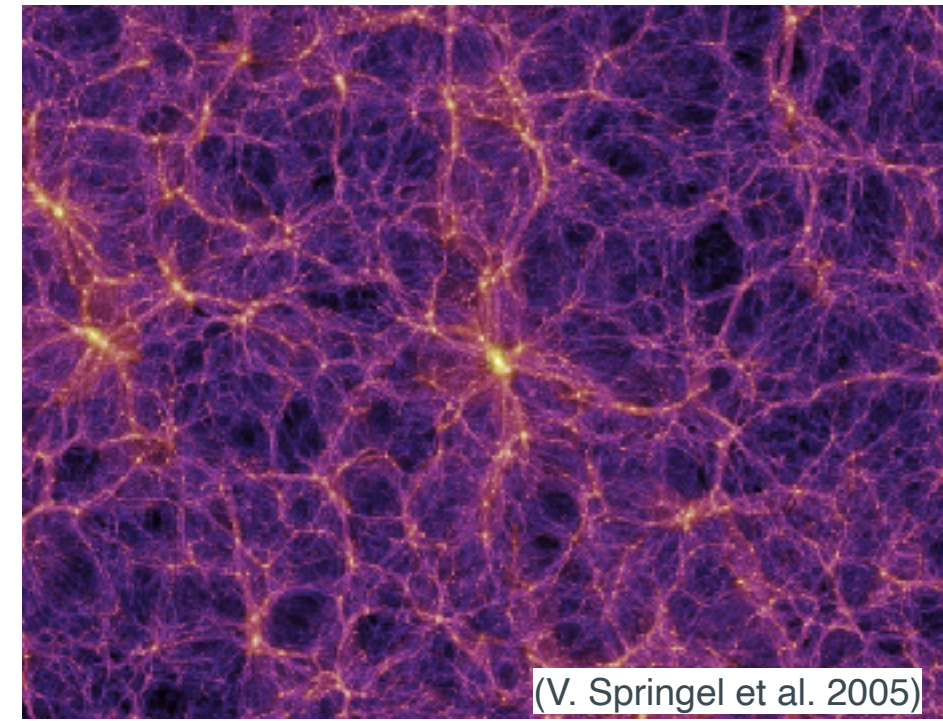


Grand Unification Theories



Dark Matter and Neutrinos

- What is Dark Matter?
- What is the nature of neutrinos?
- What is the mass of neutrinos?
- What is the neutrino mass ordering?
- Do neutrinos violate CP?
- Are there sterile neutrinos?



Dark Matter and Neutrinos

• What is Dark Matter?



Direct searches

• What is the nature of neutrinos?



$0\nu\beta\beta$ searches

• What is the mass of neutrinos?



End point of β decays

• What is the neutrino mass ordering?

• Do neutrinos violate CP?

• Are there sterile neutrinos?



Neutrino oscillation

Dark Matter and Neutrinos

• What is Dark Matter?

• What is the nature of neutrinos?

• What is the mass of neutrinos?

• What is the neutrino mass ordering?

• Do neutrinos violate CP?

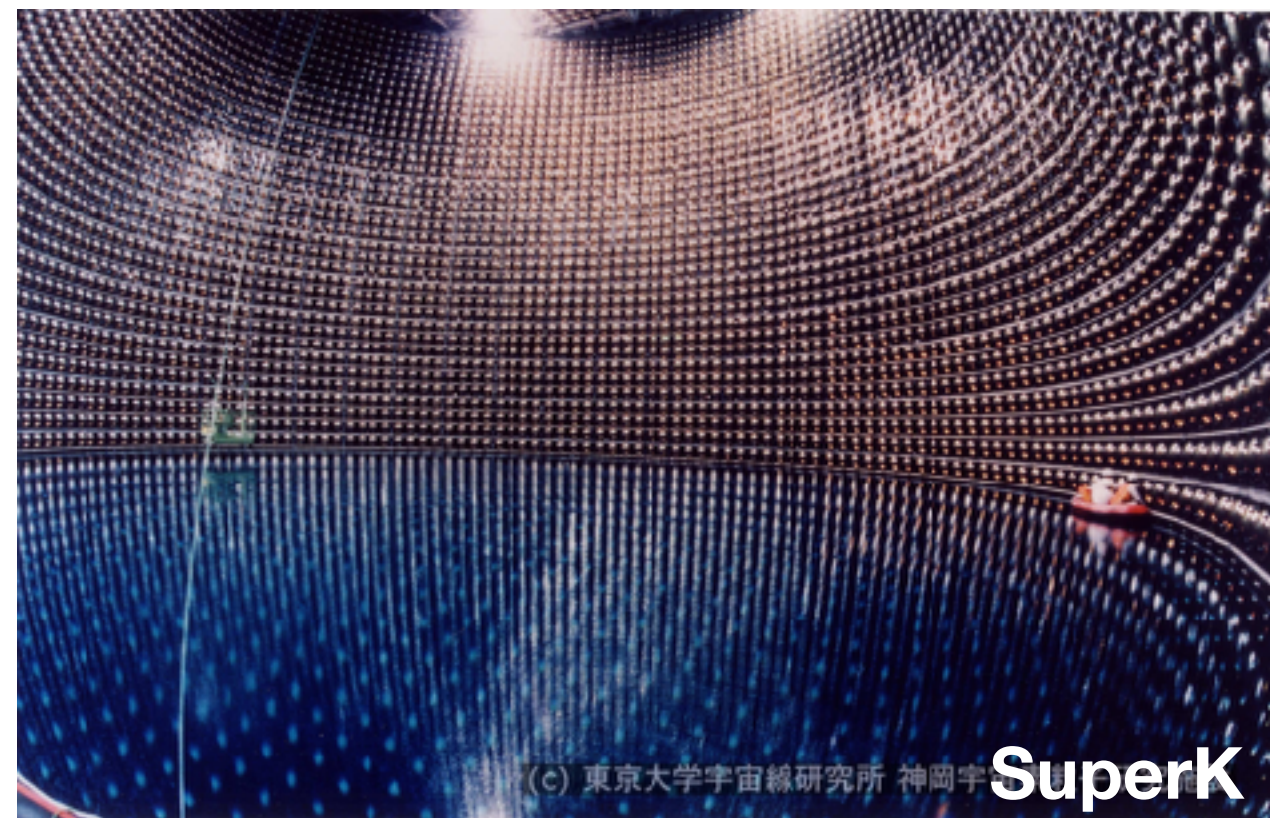
• Are there sterile neutrinos?

Large Liquid Detectors!

Liquid detector technologies

Water Cherenkov

- Established technology
- Cherenkov light
- Large-area photon detectors with some granularity (multiPMTs)
- Gd doping to increase sensitivity to neutrons



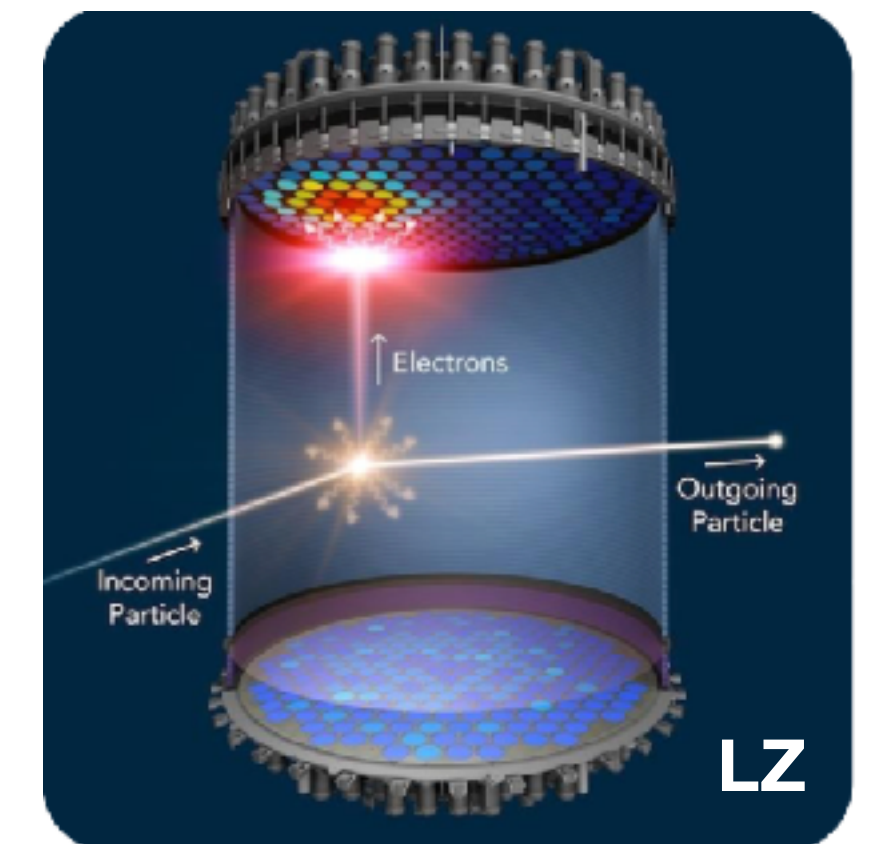
Liquid Scintillators

- Established technology
- Scintillation light (visible)
- Hybrid development for dual (Cherenkov + scintillation)
- Isotope loading ($0\nu\beta\beta$)



Noble Elements

- More recent technology
- Mostly Ar & Xe
- VUV scintillation light (Ar:128nm, Xe:175nm)
- Energy deposition shared between light (scintillation), charge (ionization) and heat (phonons)

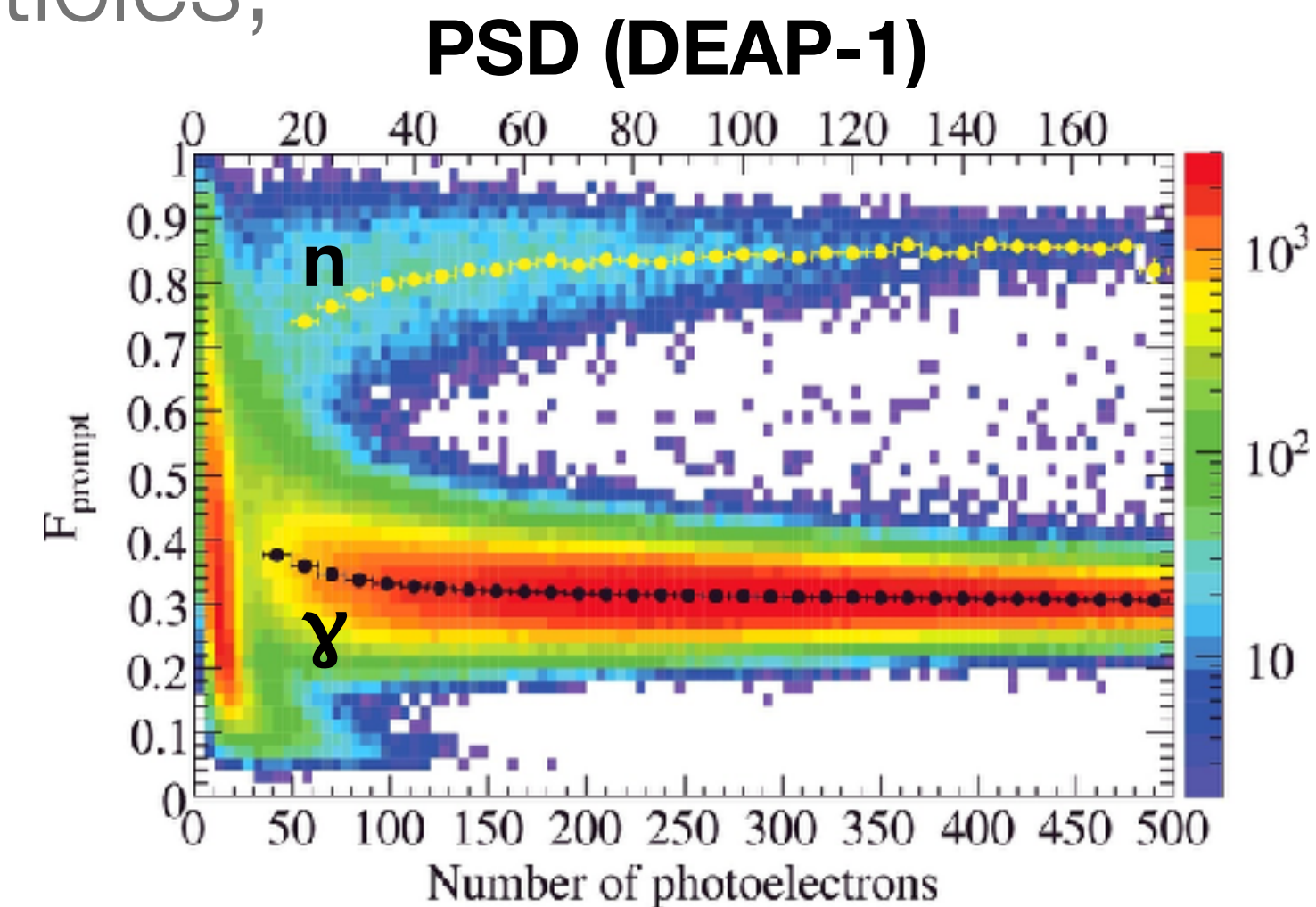


Light in detectors

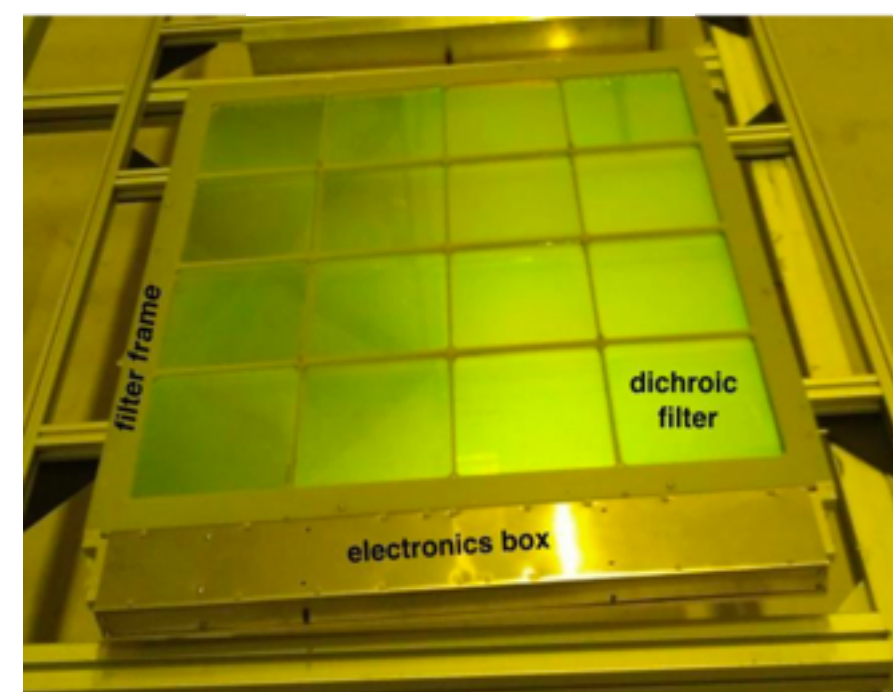
- Light is produced in detectors by charged particle energy deposition
- Light signals allow to reconstruct: time of arrival, energy of particles, types of particles, particle direction...

- To reconstruct all the light information we need to:

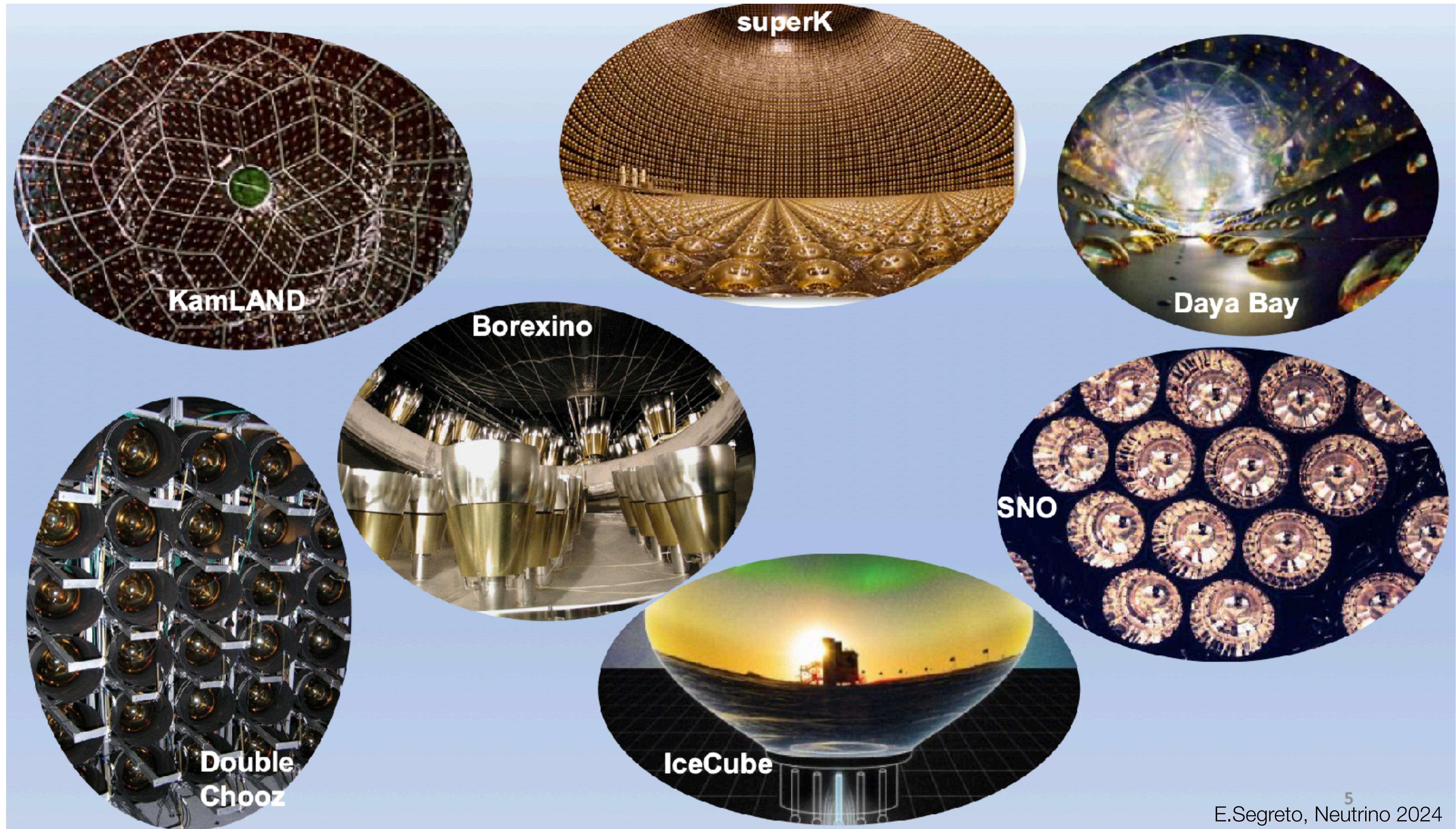
- ➔ Understand the light production mechanisms
- ➔ Understand and model light propagation
- ➔ Collect & Detect light



X-ARAPUCA

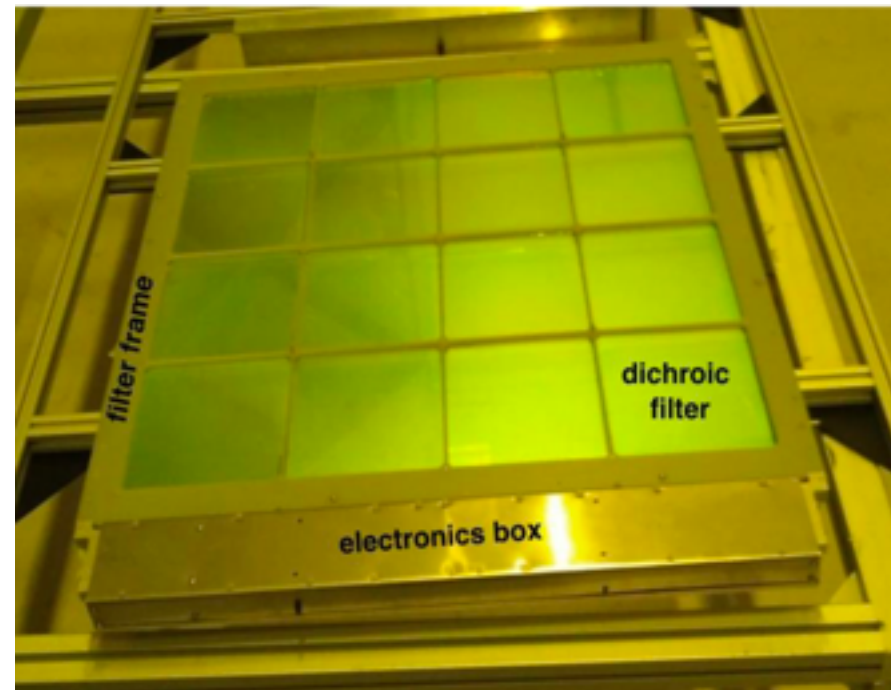
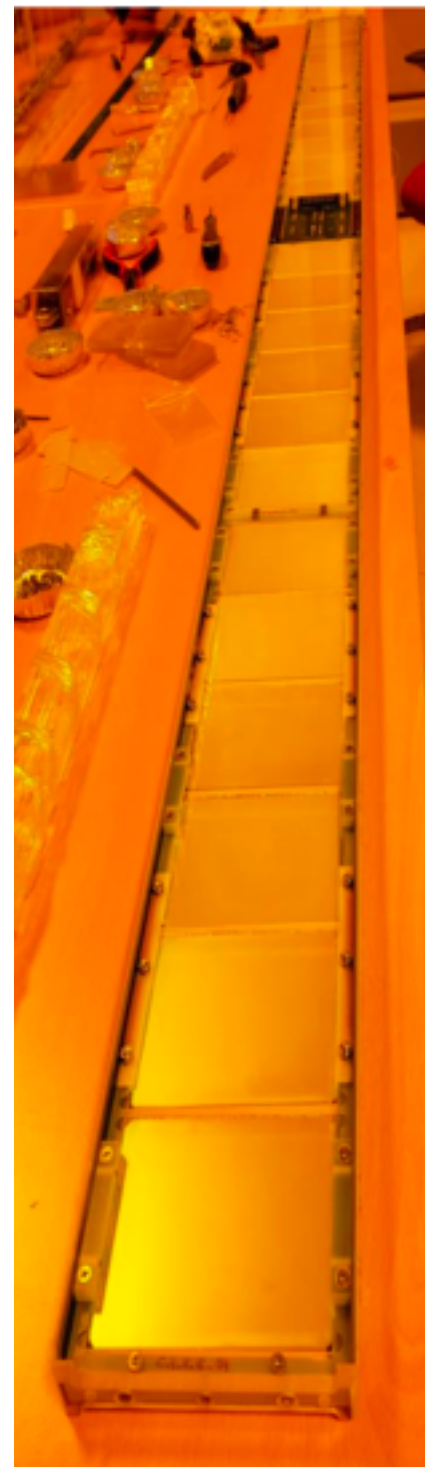


PMTs and Neutrinos (non exhaustive list)

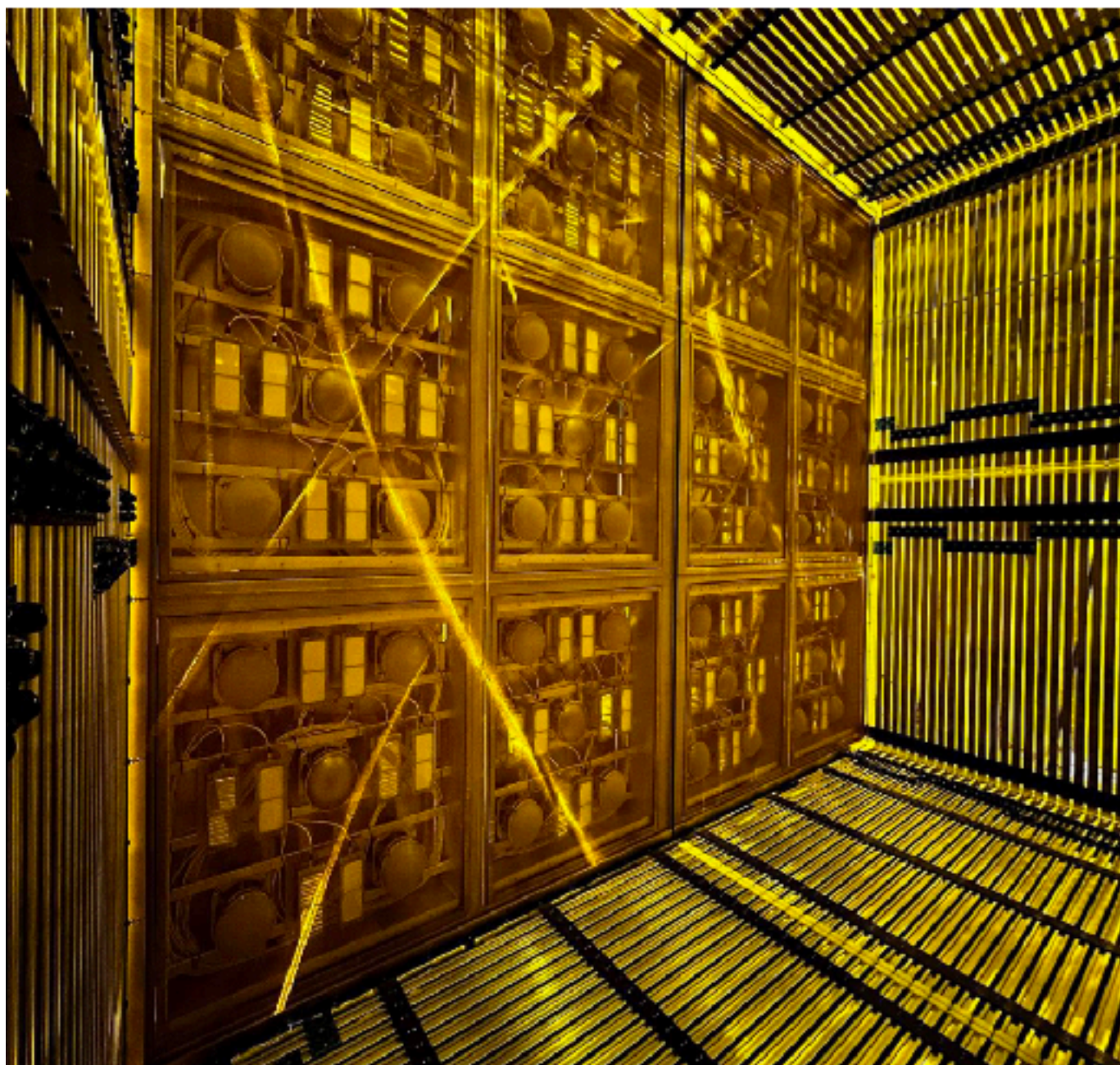


SiPMs and Neutrinos (non exhaustive list)

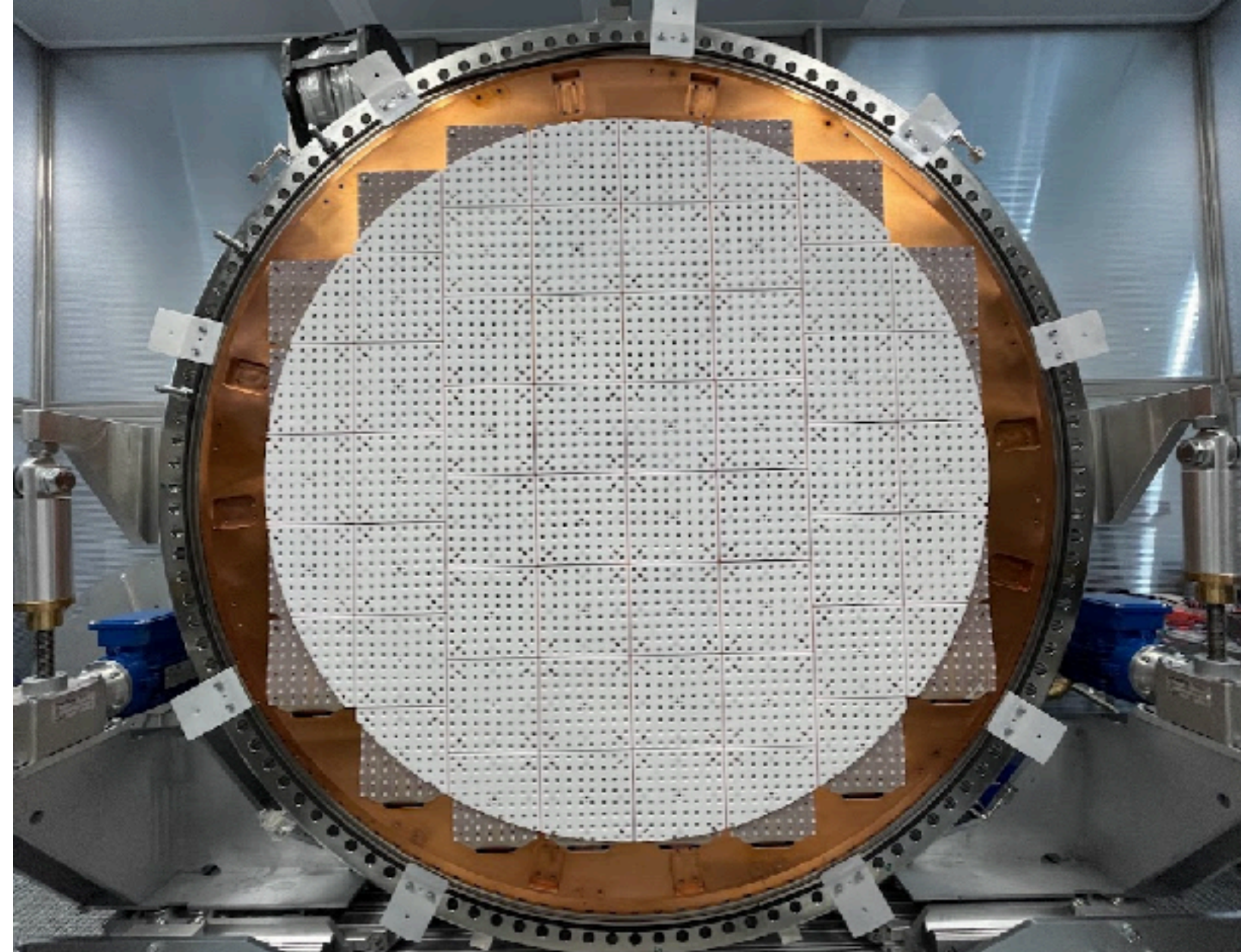
X-ARAPUCA



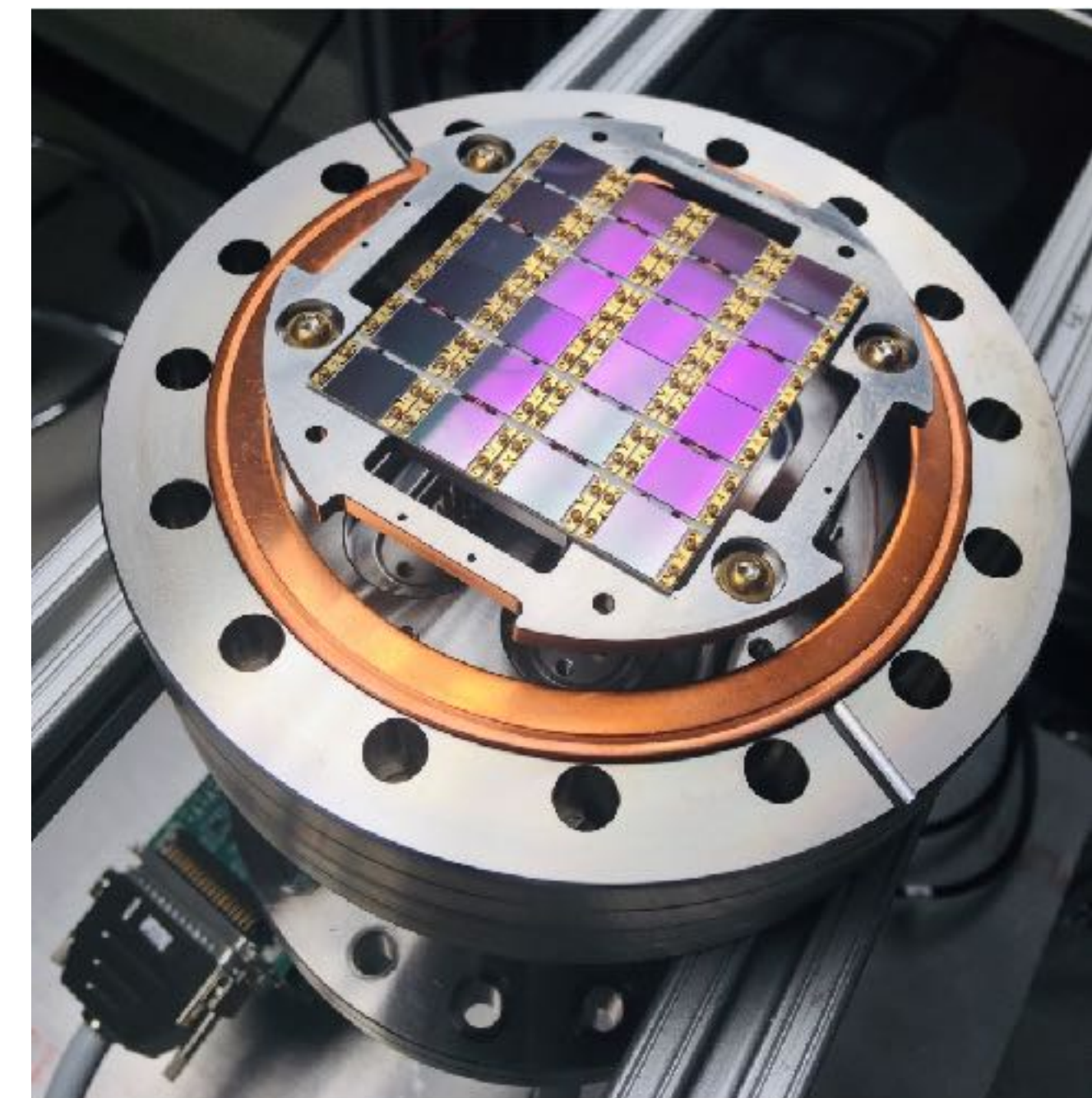
SBND



NEXT-100

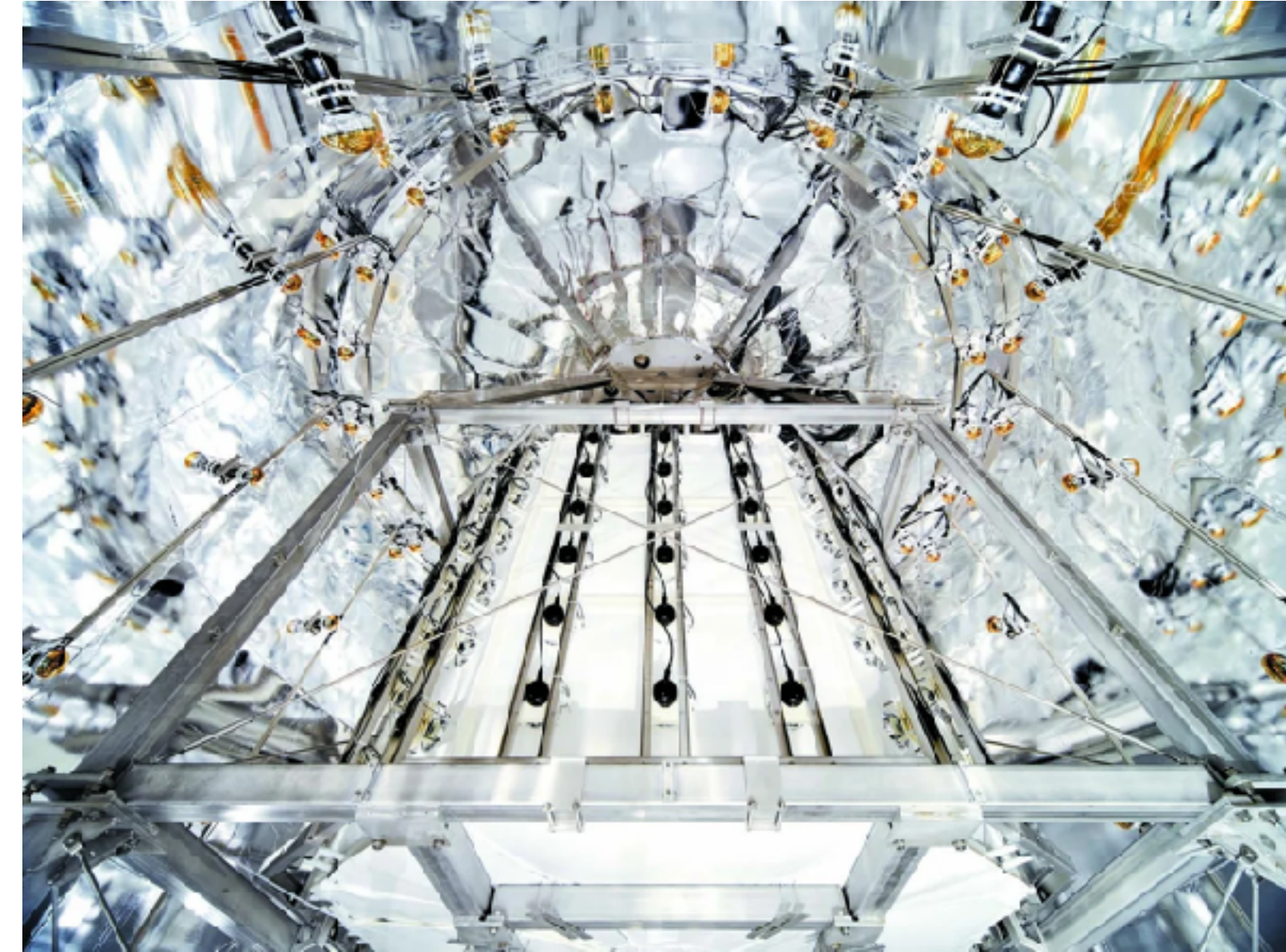
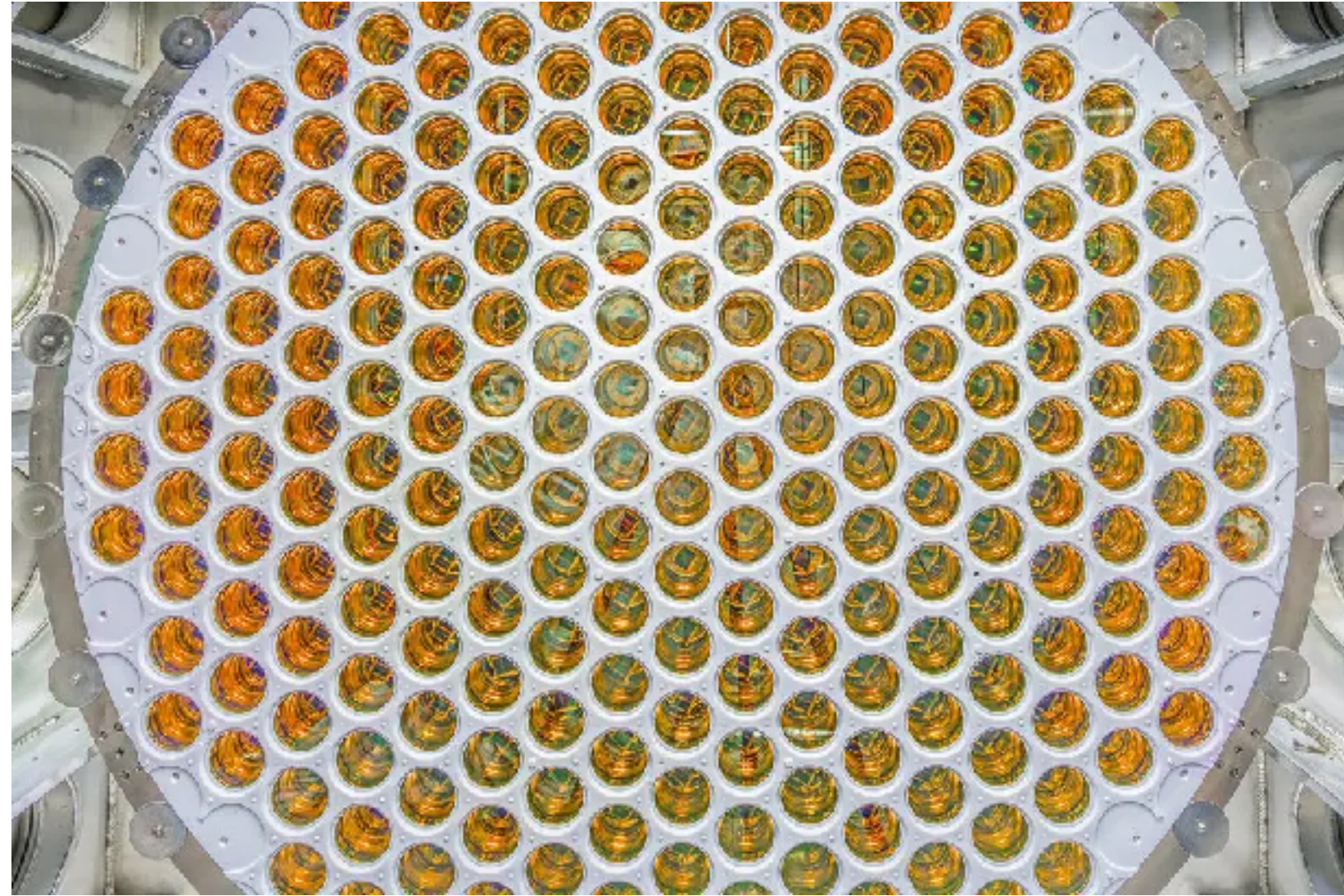


nEXO SiPM tile



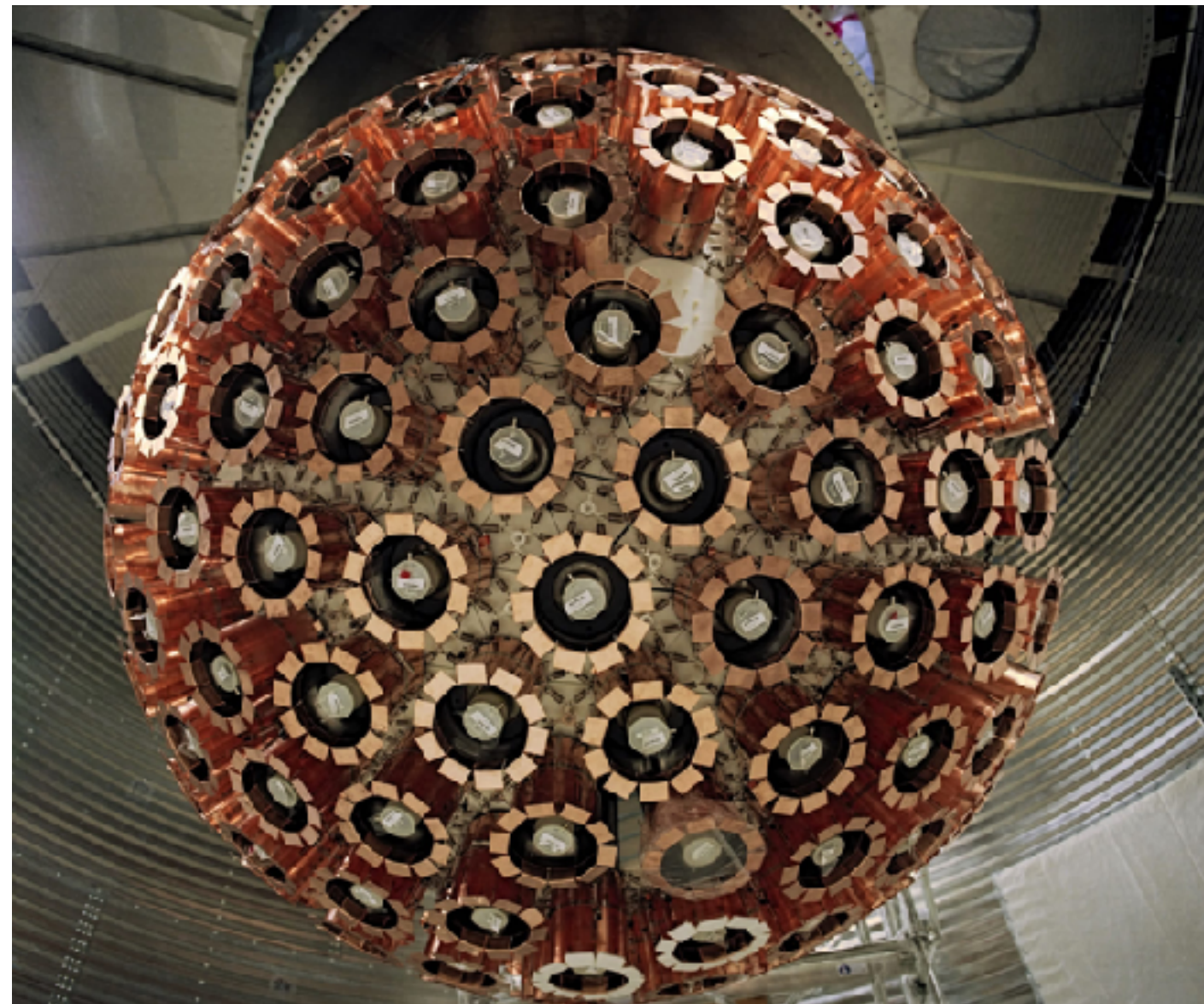
PMTs and Dark Matter (non exhaustive list)

LZ



XENON

DEAP-3600

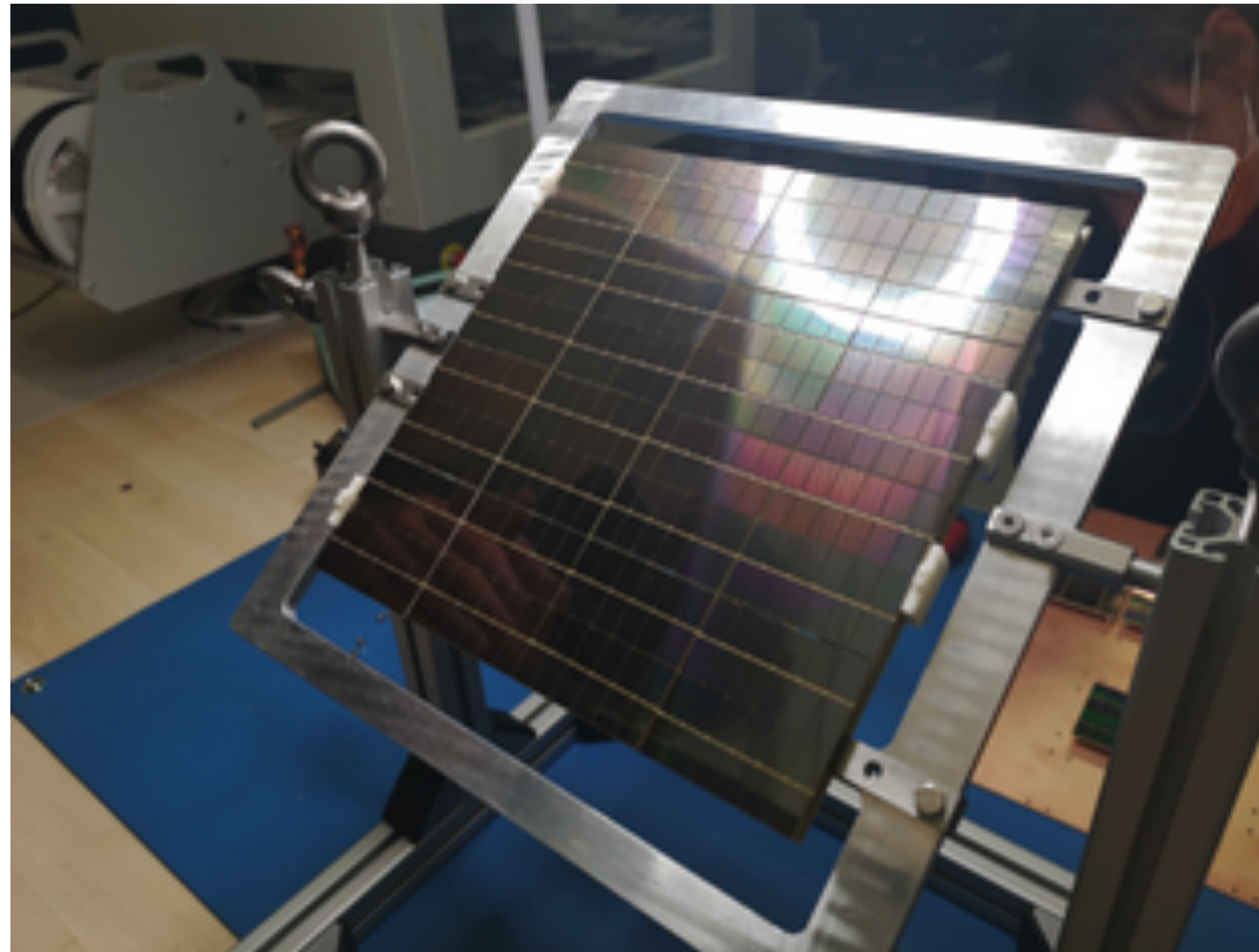


DarkSide-50

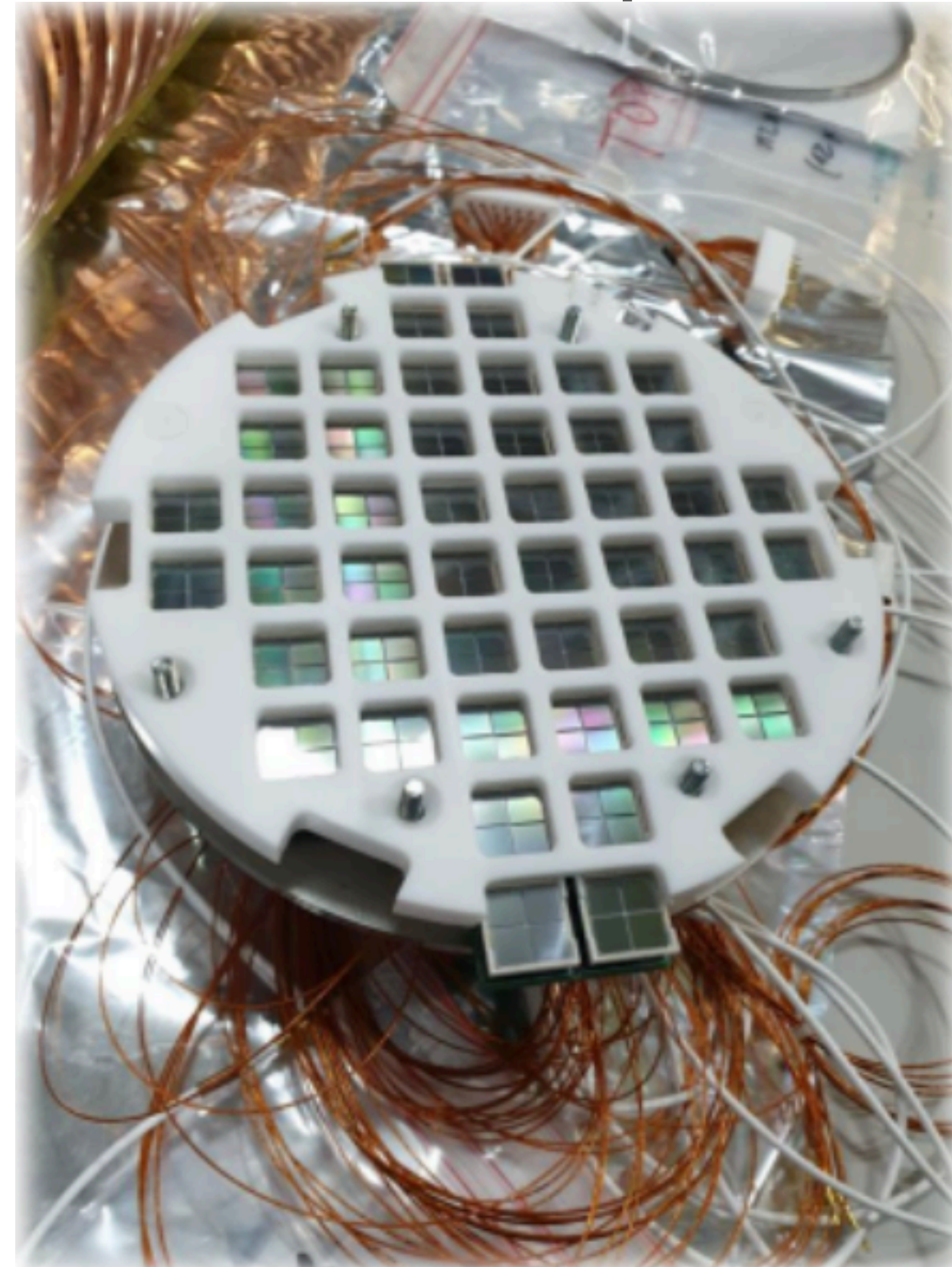


SiPMs and Dark Matter (non exhaustive list)

DarkSide-20k SiPM tile



Xenoscope



Physics needs for the future

Neutrinos

- **Push Energy thresholds down** to ~ 1 MeV to enhance oscillation physics, supernovae ν s study, to enable solar ν s ...
- **Unambiguous readout**
- **Scalability**

Dark Matter

- **Push Energy thresholds down** to 1 meV/10 eV/1 keV to enable low mass DM/1 GeV DM/WIMPs
- **Reduce background rates**
- **Scalability**

$0\nu\beta\beta$

- **Improve Energy Resolution** to sub-% FWHM
- **Reduce background rates**
- **Scalability**

Physics needs for the future

Neutrinos

- **Push Energy thresholds down** to ~ 1 MeV to enhance oscillation physics, supernovae ν s study, to enable solar ν s ...

Dark Matter

- **Push Energy thresholds down** to 1 meV/10 eV/1 keV to enable low mass DM/1 GeV DM/WIMPs

$0\nu\beta\beta$

- **Improve Energy Resolution** to sub-% FWHM

Improving light readout can address all of these!

- **Unambiguous readout**

- **Scalability**

- **Reduce background rates**

- **Scalability**

- **Reduce background rates**

- **Scalability**

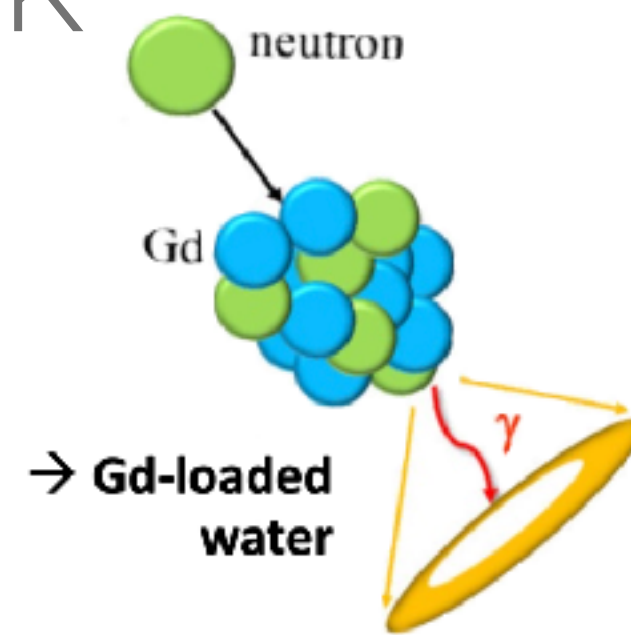
Current Goals (non exhaustive list)

Water Cherenkov

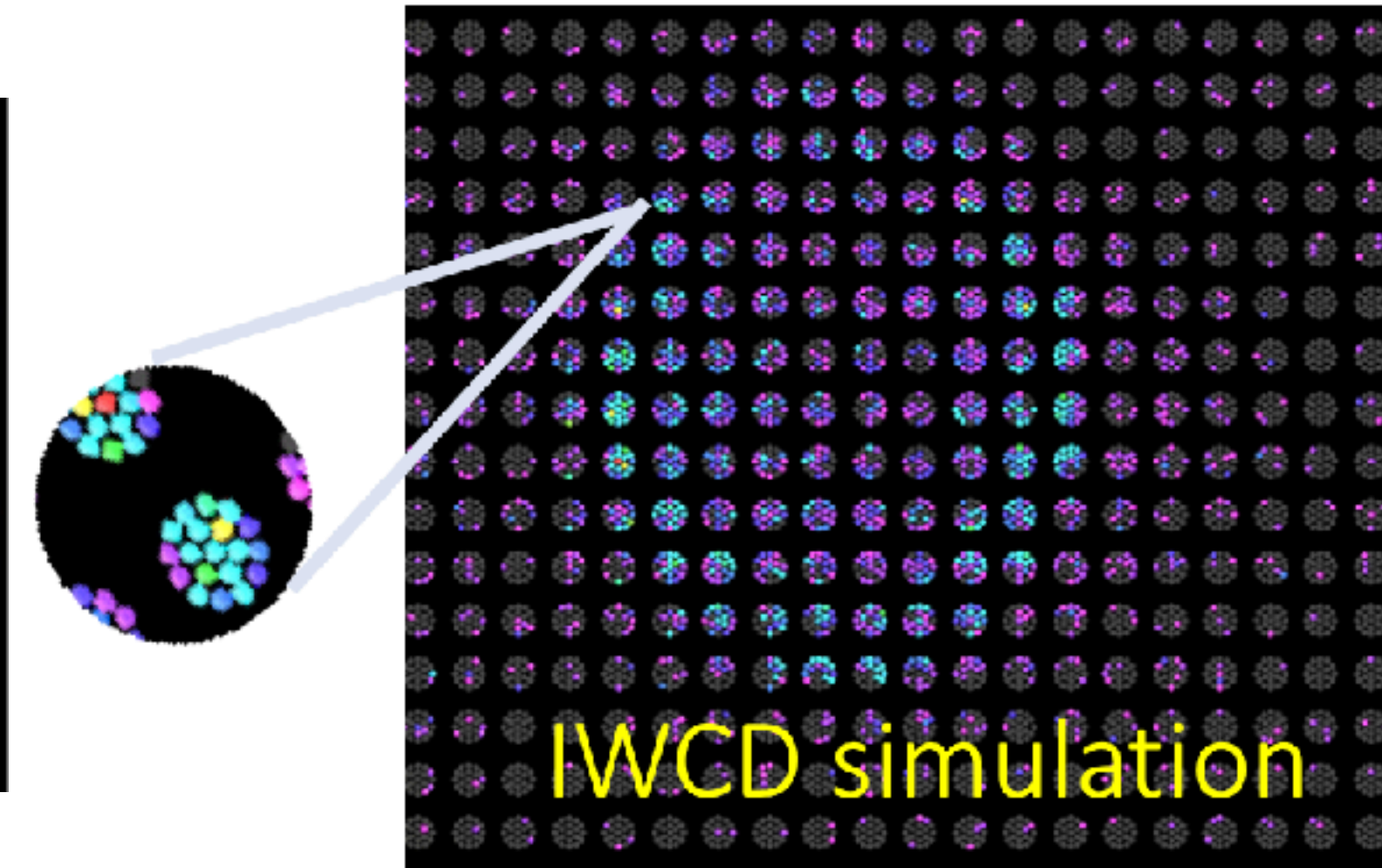
- Mainly targeted at neutrino physics:

→ **Neutrino oscillation:** HyperK

- Gd-doping ($>0.1\%$)
- Purification and recovery



HyperK



→ **Neutrino astronomy:**

KM3Net



TRIDENT



Current Goals (non exhaustive list)

Liquid Scintillators

- Mainly targeted at neutrino physics:

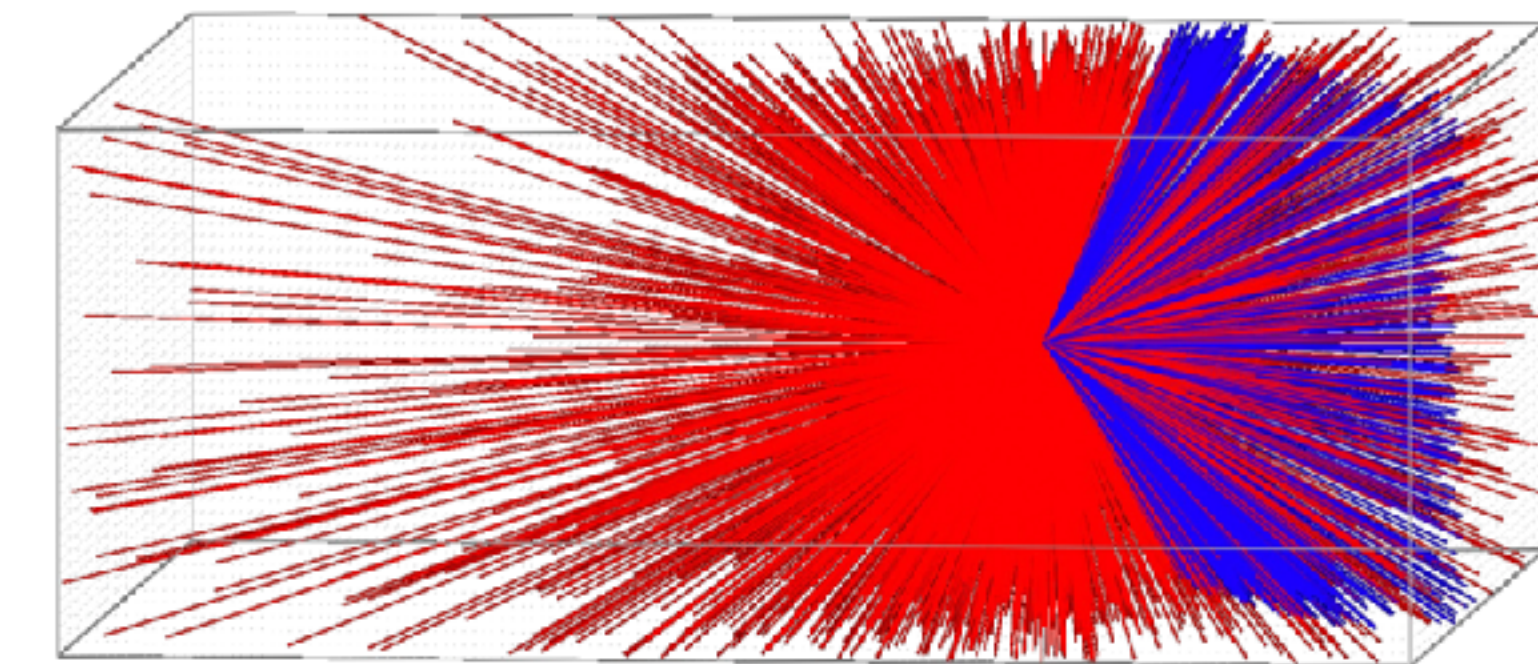
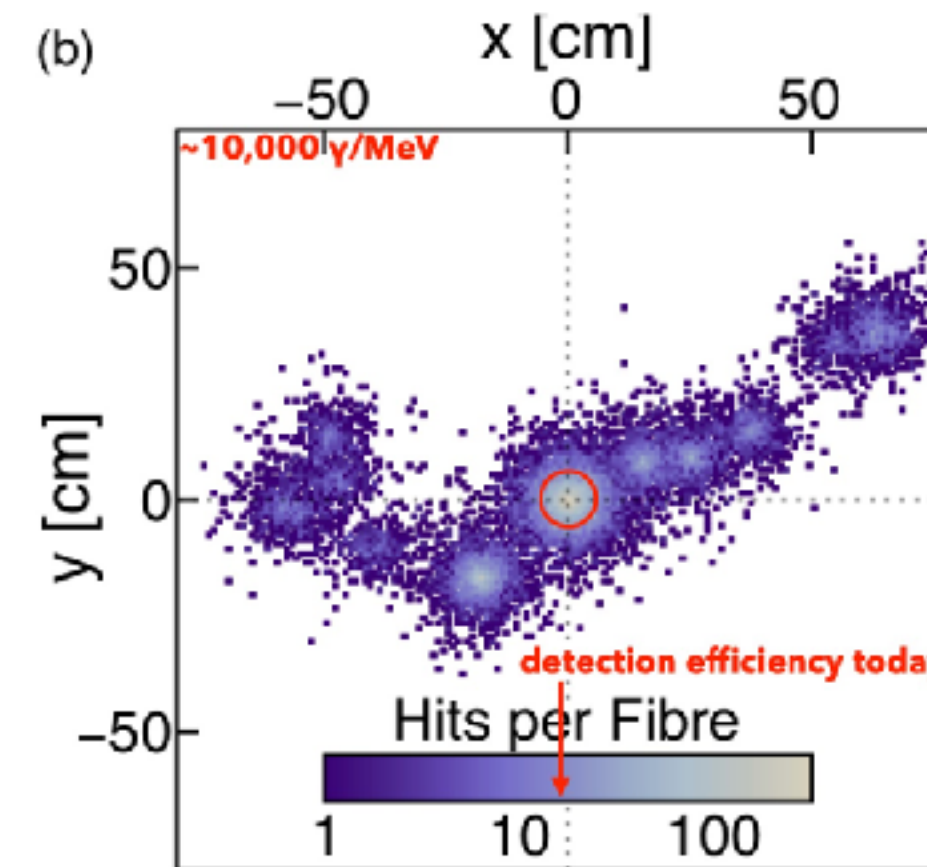
→ Neutrino studies: Hybrid LS, opaque LS

- Demonstration of novel LS
- Understanding of light propagation
- Demonstration at (ton)scale
- Gd-loading

→ $0\nu\beta\beta$: Te loading

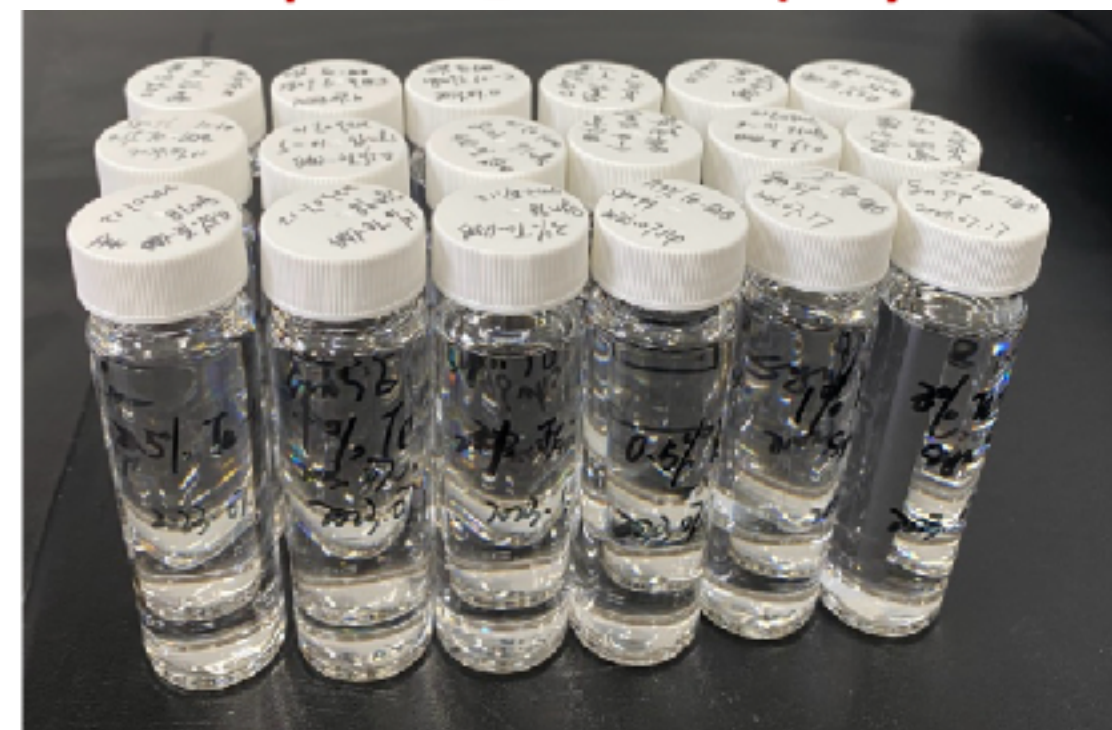
- Increasing loading (Te, Xe)

→ opaque scintillators

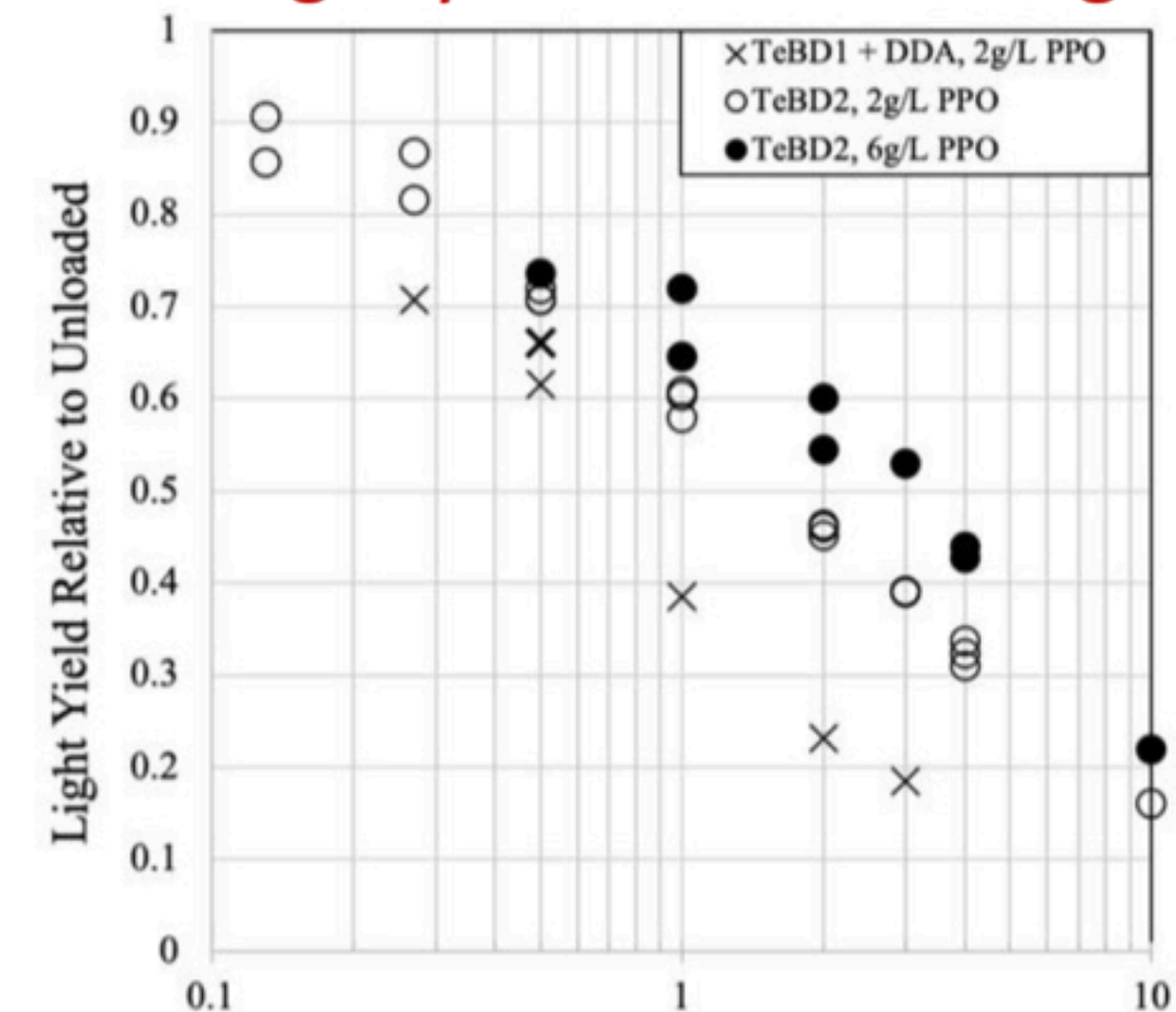


→ hybrid scintillators

Te-LS samples @ IHEP (Yayun Ding)



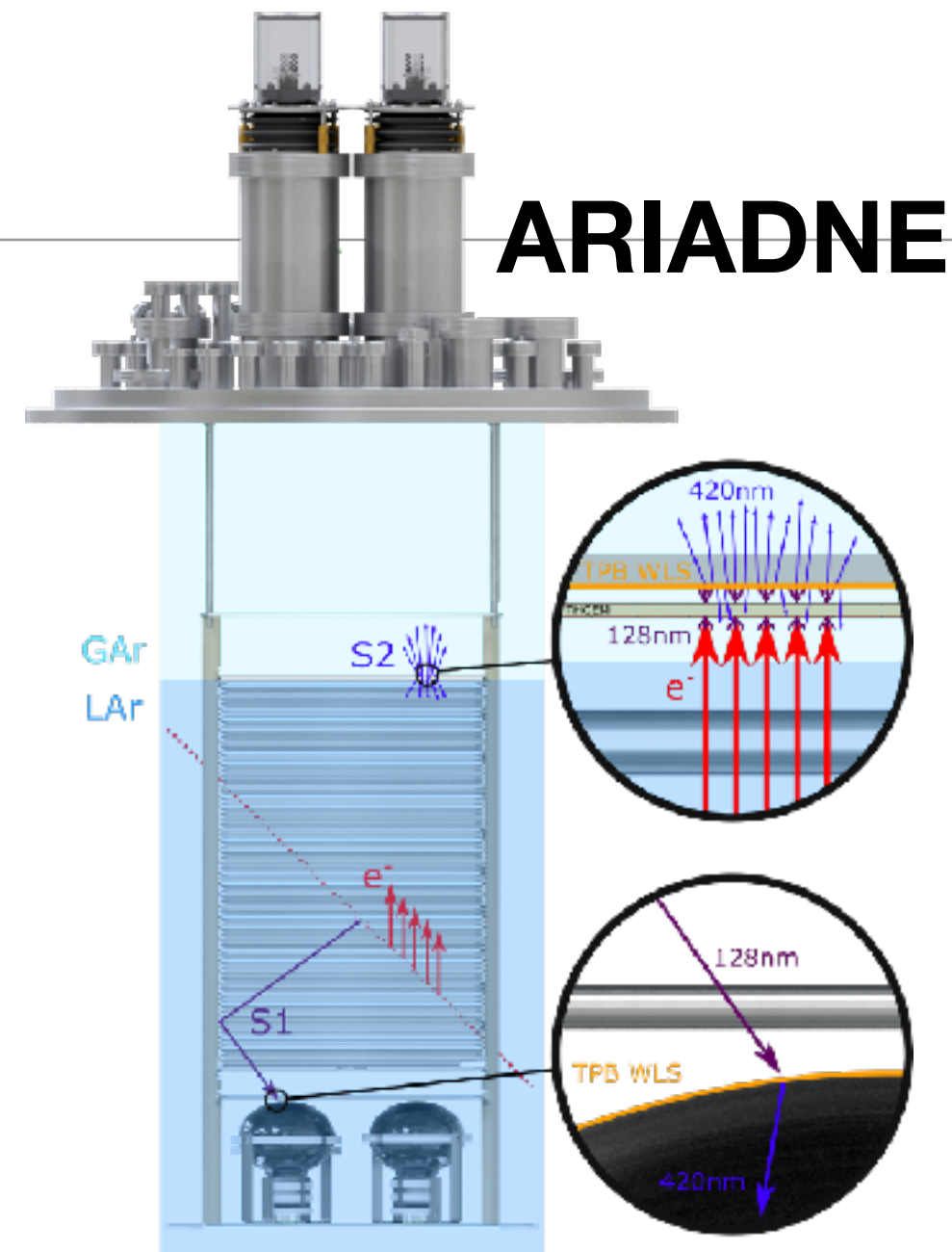
Light yield vs. Te-loading



Current Goals (non exhaustive list)

Noble Elements

- Mainly targeted at neutrino physics:
 - ➔ **Neutrino studies:** kilo-ton scale LAr
 - Amplification, combined charge-light
 - Microphysics (MeV-GeV)
 - Xe doping



- ➔ **$0\nu\beta\beta$:** ton-scale LXe
 - Microphysics (MeV)
 - Xe procurement



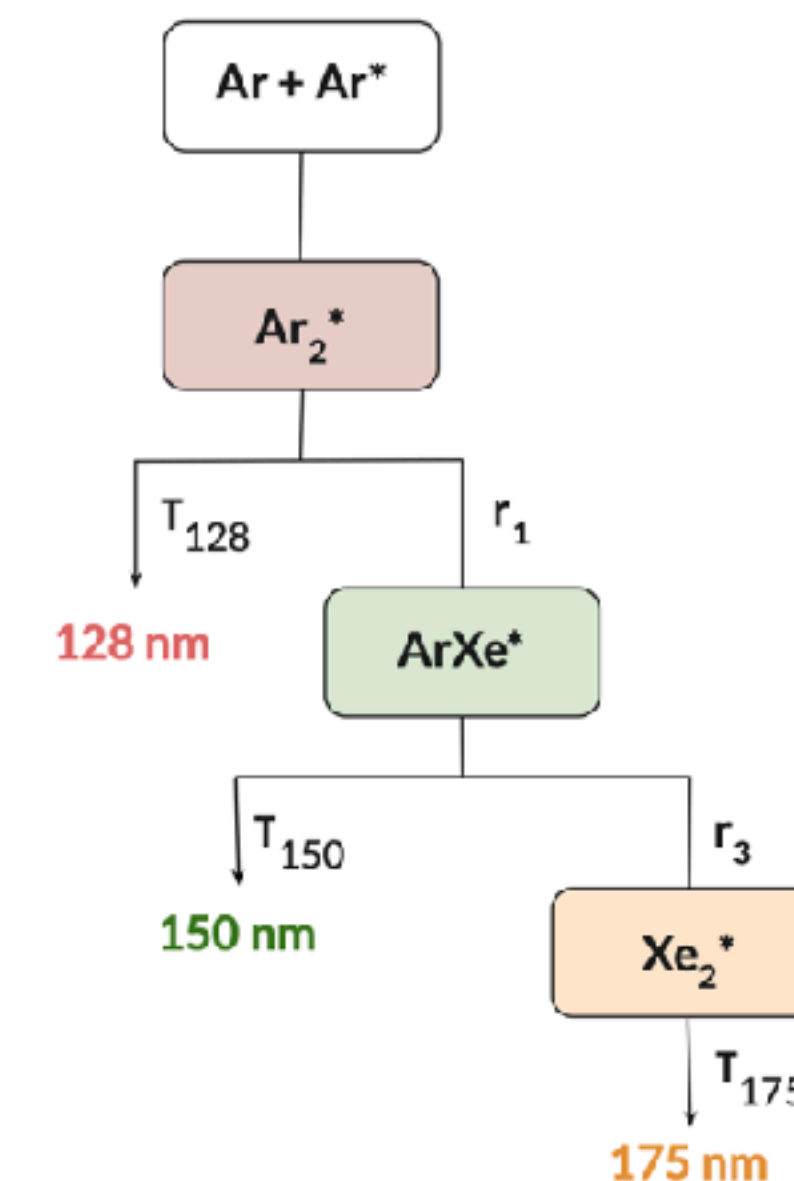
Linde Air Separation Unit in Hungary

- ➔ **Dark Matter:** (50-100) ton-scale LXe and LAr
 - Microphysics (keV) & light transport
 - UAr procurement, purification

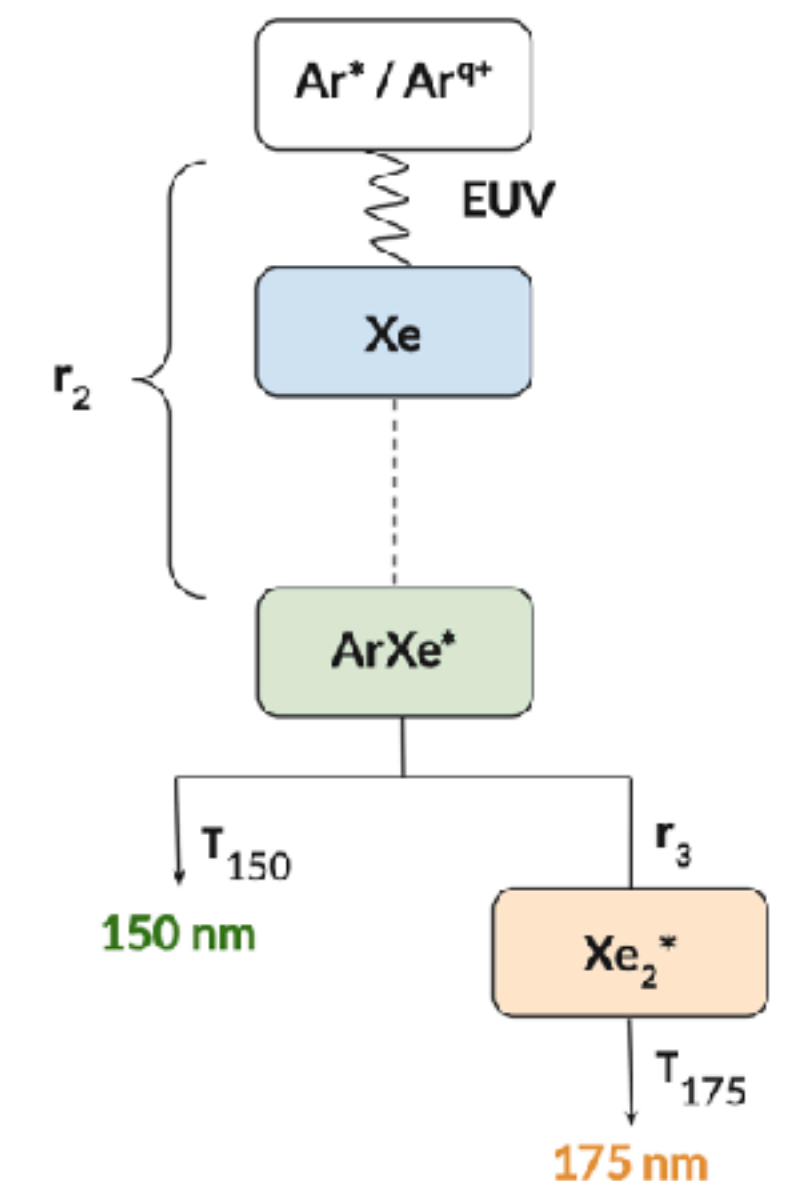
EUV scintillation from xe-doped LAr

P.Aanes et al., X-ArT : <https://arxiv.org/abs/2410.22863>

Collisional Branch



Radiative Branch



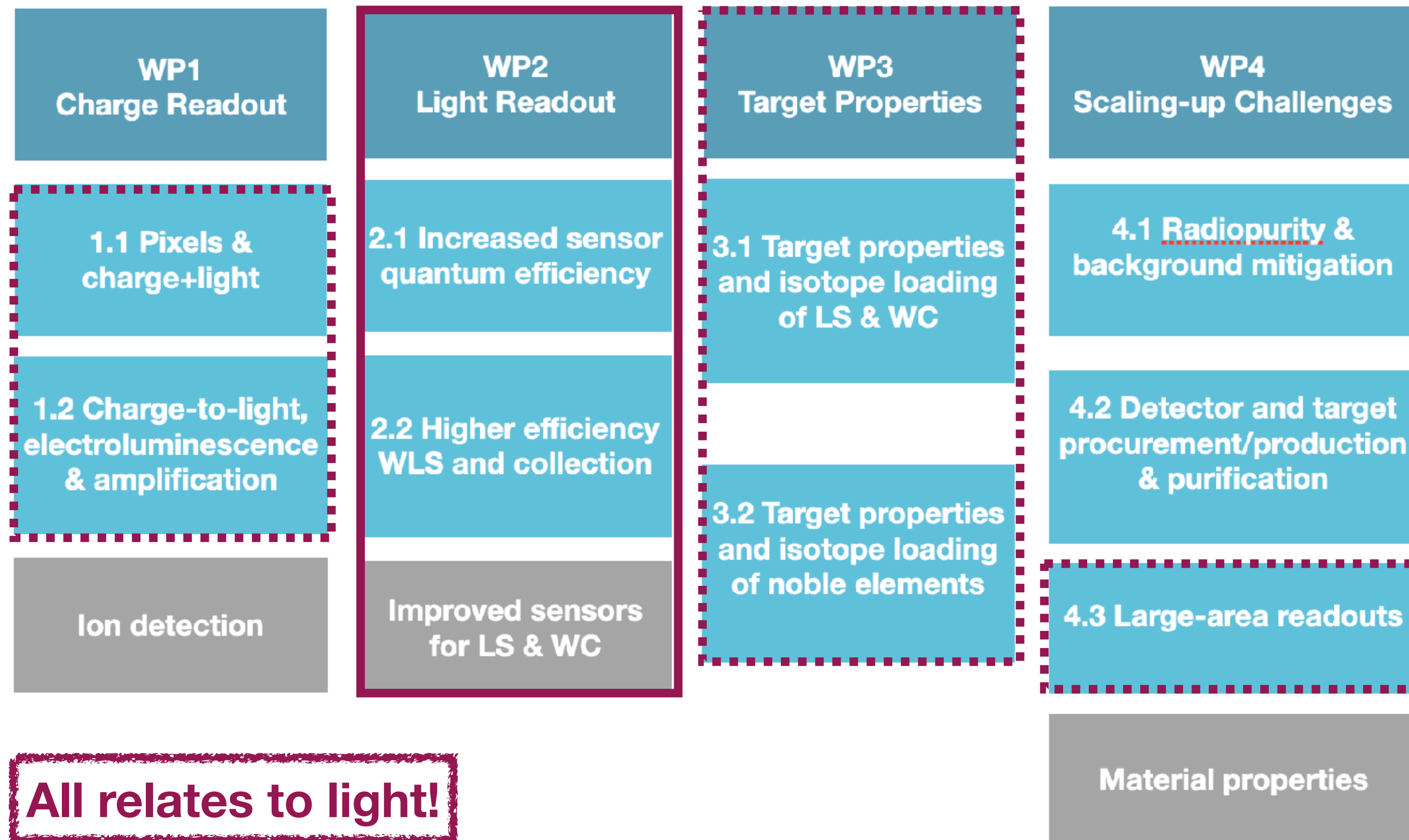
Targeted detector R&D

- DRD2 Collaboration: R&D for Liquid Detectors



Targeted detector R&D

- DRD2 Collaboration: R&D for Liquid Detectors



WP2.1: Increased sensor quantum efficiency Task Leaders: Paolo Agnes & Miguel Garcia Peris

- Targeted at **VUV** only (and for cryogenic operations) → Clear synergy with DRD4

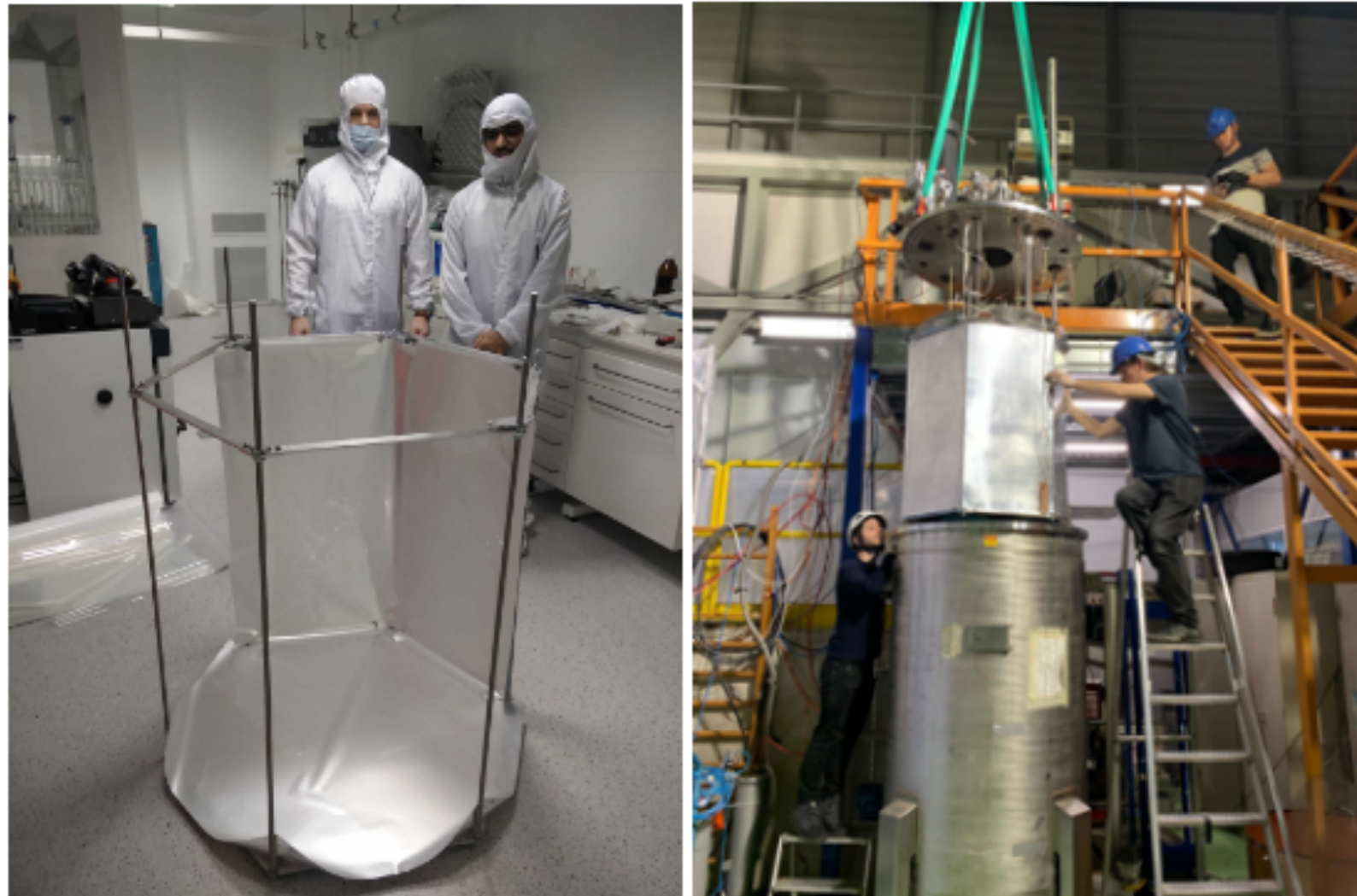
	Goal	Description
D2.1	Sensor development for VUV sensitivity	Development* and characterization of organic photosensors, coatings and passivation methods and of SPAD geometry for VUV detection
D2.2	Prototype SPAD arrays	Prototypes and characterization of new SPAD arrays for 3D-integrated FSI and BSI, analog BSI and monolithic arrays
D2.3	Report on VUV-optimized sensor	Report on the performance of new VUV-optimised sensors in term of PDE, noise and application to rare-event searches

* At Manchester, we will start testing 2D material coatings (e.g. graphene), potential collaboration possible!

WP2.2: Higher efficiency WLS and collection

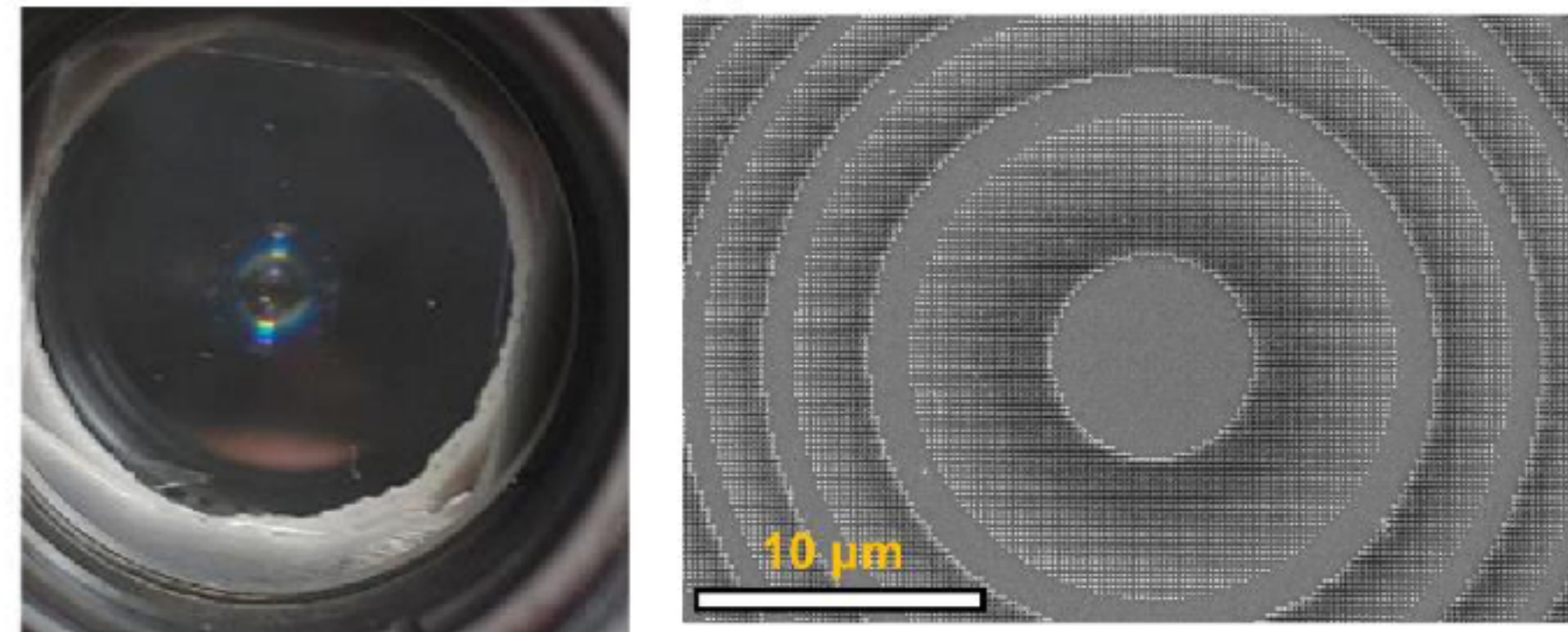
Title	Description
D2.4 Report on optimised WLS	Report on optimised WLS (VUV to visible) and evaporation systems
D2.5 Design report on light collection	Design report on VUV light collection in noble elements and light readout for liquid scintillators

Large-scale PEN WLS test at CERN



Led by Astrocent, CERN, U. Edinburgh, U. Hawaii, Nikhef, TUM, Uni. Zurich
Contact Marcin Kuzniak (Astrocent)

VUV metalenses as light collectors



A. Martins, T. Contreras, C. Stanford, M. Tuzi, J. M. Albo, C. O. Escobar, A. Para, A. Kish, J.-S. Park, T. F. Krauss, R. Guenette, (2024), High efficiency glass-based VUV metasurfaces, arXiv:2401.11315, submitted to ACS Nano.

A. Martins, A. F. da Mota, C. Stanford, T. Contreras, J. Martin-Albo, A. Kish, C. Escobar, A. Para, R. Guenette, (2023), Simple strategy for simulation of large area of axially symmetric metasurfaces, arXiv:2310.19121, accepted by JOSAB.

T. Contreras, A. Martins, C. Stanford, C. O. Escobar, R. Guenette, M. stancari, J. Martin-Albo, B. Lawrence-Sanderson, A. Para, A. Kish, F. Kellerer, (2023), A Method To Characterize Metalenses For Light Collection Applications, JINST 18 09, T09004.

A.A. Loya Villalpando, J. Martín-Albo, W.T. Chen, R. Guenette, C. Lego, J.S. Park² and F. Capasso, (2020), Improving the light collection efficiency of silicon photomultipliers through the use of metalenses, JINST 15 P11021.

WP 3: Target Properties

Water Cherenkov

- Increase of the Gd-concentration in water
- Synthesize nanoparticles-based water liquid scintillator and water-based nanocrystal liquid scintillators

Liquid Scintillators

- Development of slow WbLS with alternative fluors and quantum dots, and isotope loading in WbLS
- New class of slow WbLS, build demonstrator of WbLS purification
- Development of improved microphysics model for organic liquid scintillators
- Demonstration of $\beta\beta$ isotope loading in scintillators on several % level
- ...

Noble Elements

- LAr/LXe nuclear recoil quenching and its fluctuation down to sub-keV region
- LAr/LXe ionisation response to electron recoil/nuclear recoils vs electric field
- Low-energy calibration techniques
- Impact of contaminants on noble liquid responses
- Energy transfer and response in Xe-doped LAr
- ...

Summary

- Light detection is central to all neutrino and dark matter detectors
- Future generations would significantly benefit from increased and improved light detection
- Wide R&D activities already underway
- DRD2 collaboration aims to bring together the R&D community to strengthen the different R&D programs and to provide a platform for scientific exchange

Consider joining DRD2!

Email Roxanne Guenette (roxanne.guenette@manchester.ac.uk)
and/or Giuliana Fiorillo (giuliana.fiorillo@na.infn.it)