

BUBBLE, BUBBLE, PERTURB & TROUBLE

A NONPERTURBATIVE TEST OF NUCLEATION CALCULATIONS
FOR STRONG PHASE TRANSITIONS

With Oliver Gould & David J. Weir
Based on [arXiv:2404.01876](https://arxiv.org/abs/2404.01876)

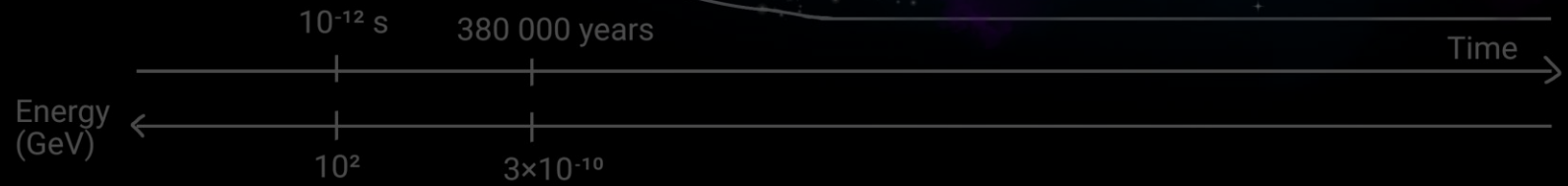
Anna Kormu (she/her)
CERN Cosmo Coffee
May 8th 2024
University of Helsinki & Helsinki Institute of Physics

1. Cosmological phase transitions

2. Gravitational waves

3. A short history of the bubble nucleation calculations

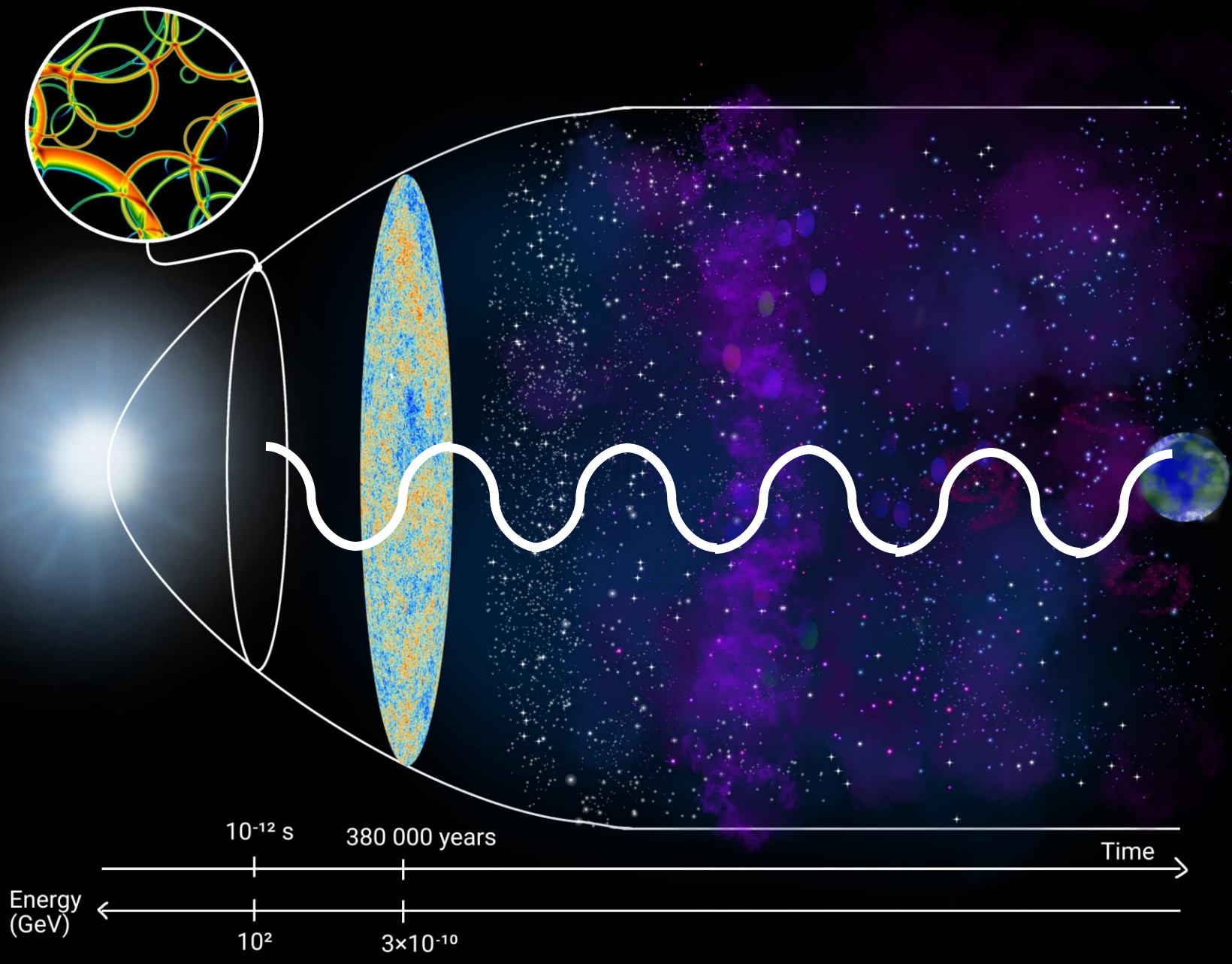
4. Bubble nucleation rate, nonperturbatively



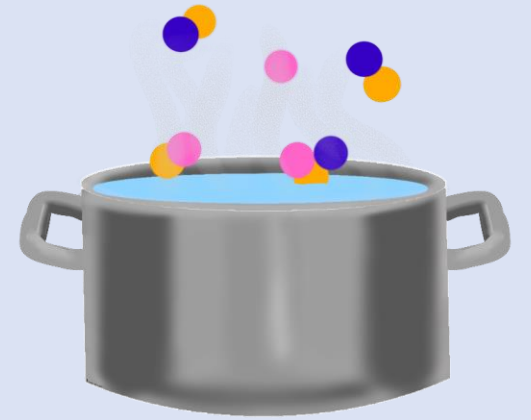
Phase transitions

- Grand Unified Theories, Electroweak, QCD...
- In the Standard Model (SM) the electroweak PT is a crossover
- SM is incomplete \rightarrow beyond SM (BSM) physics
- Things to look for: topological defects, bubbles from EWPT, ... ?

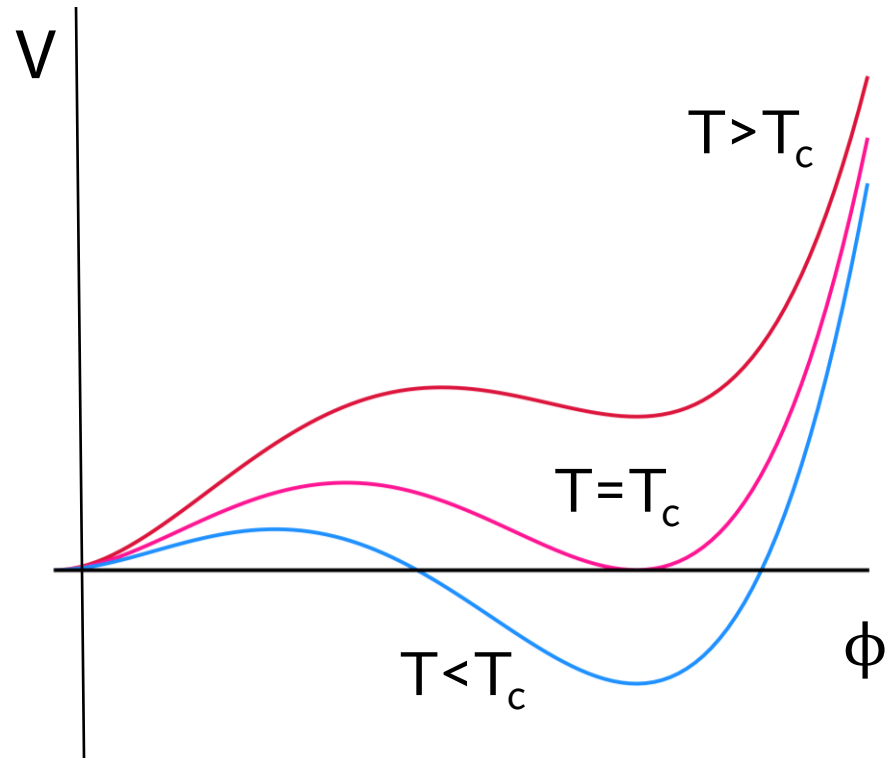
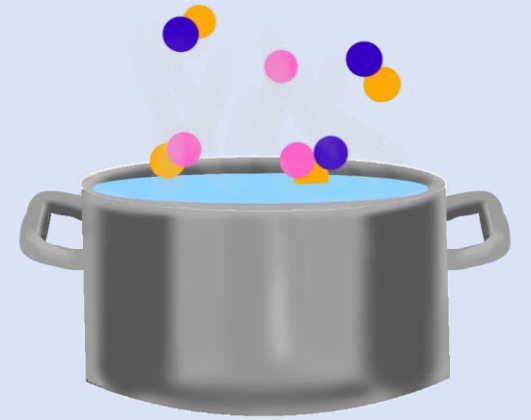




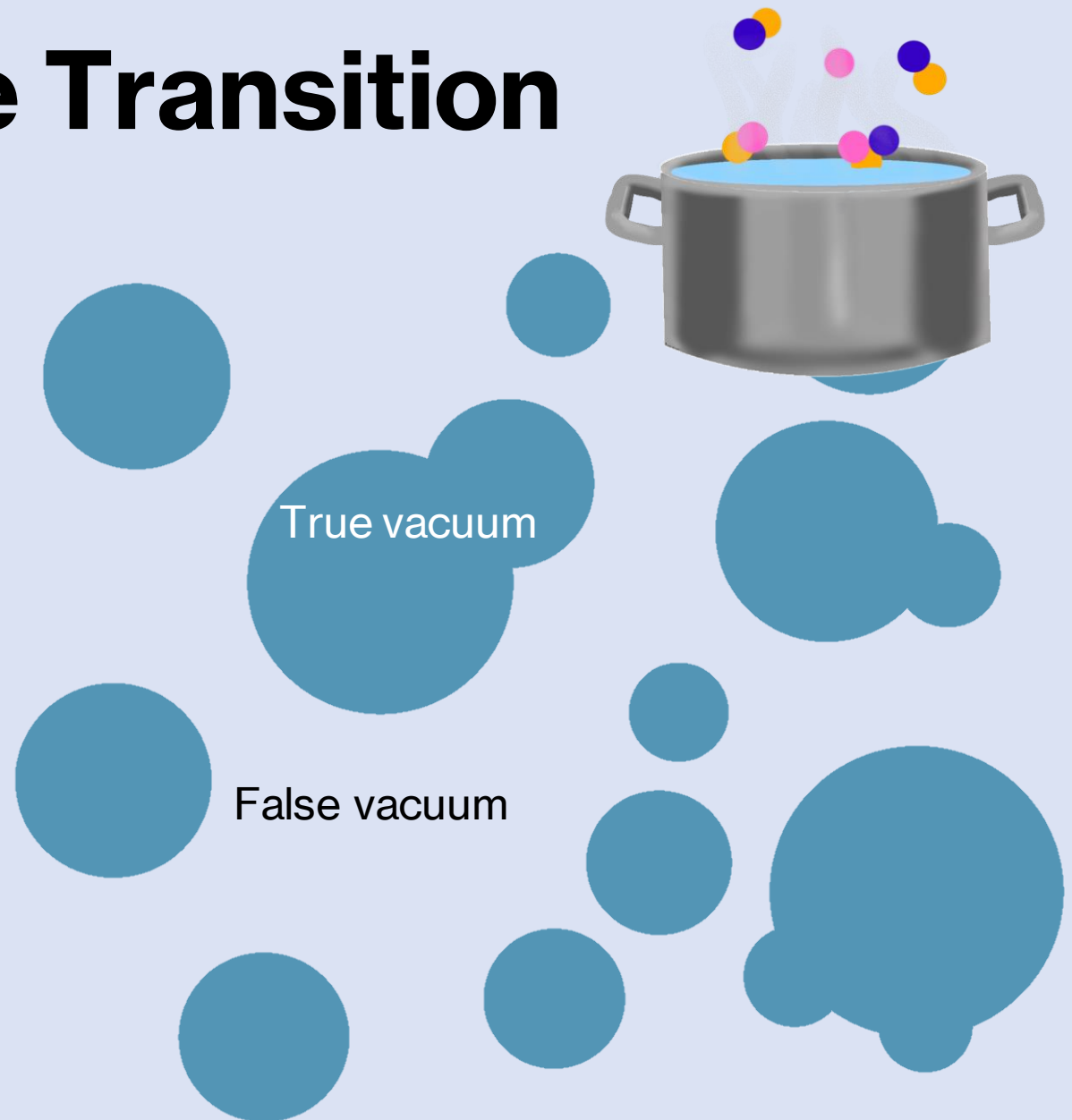
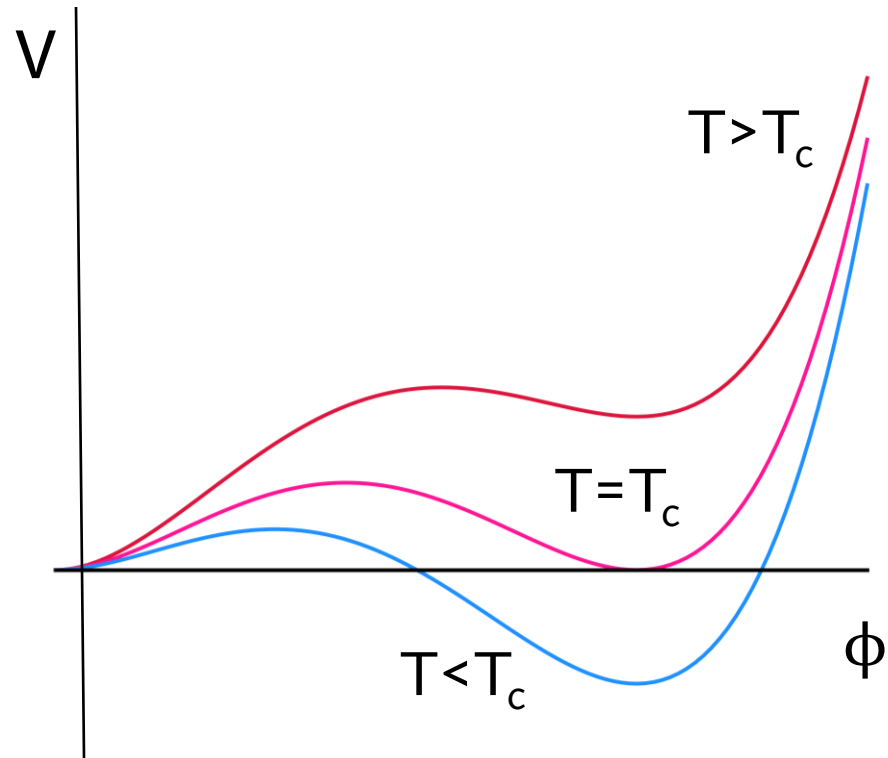
Electroweak Phase Transition



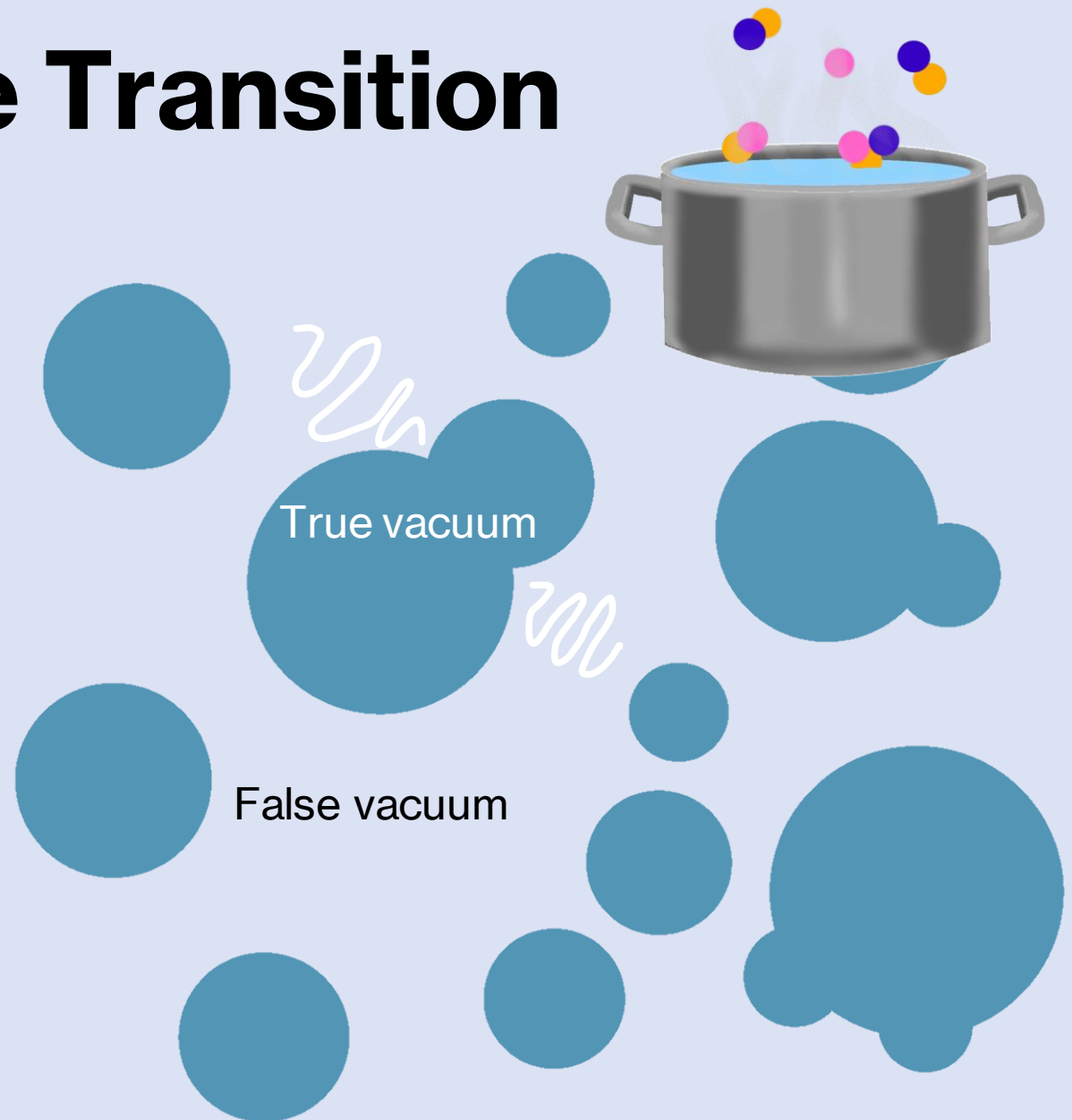
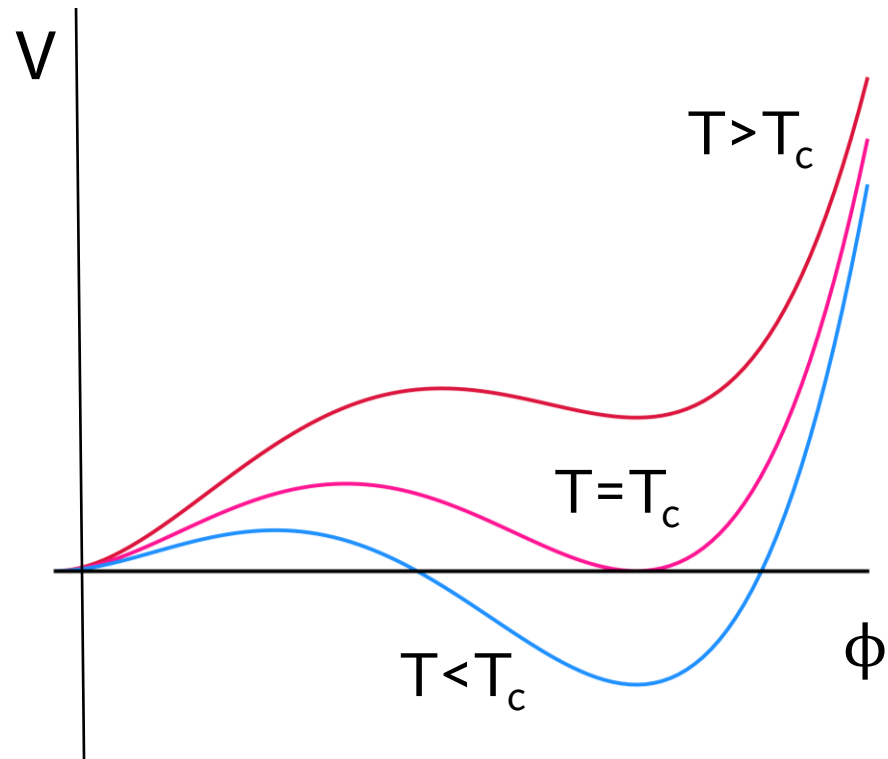
Electroweak Phase Transition



Electroweak Phase Transition



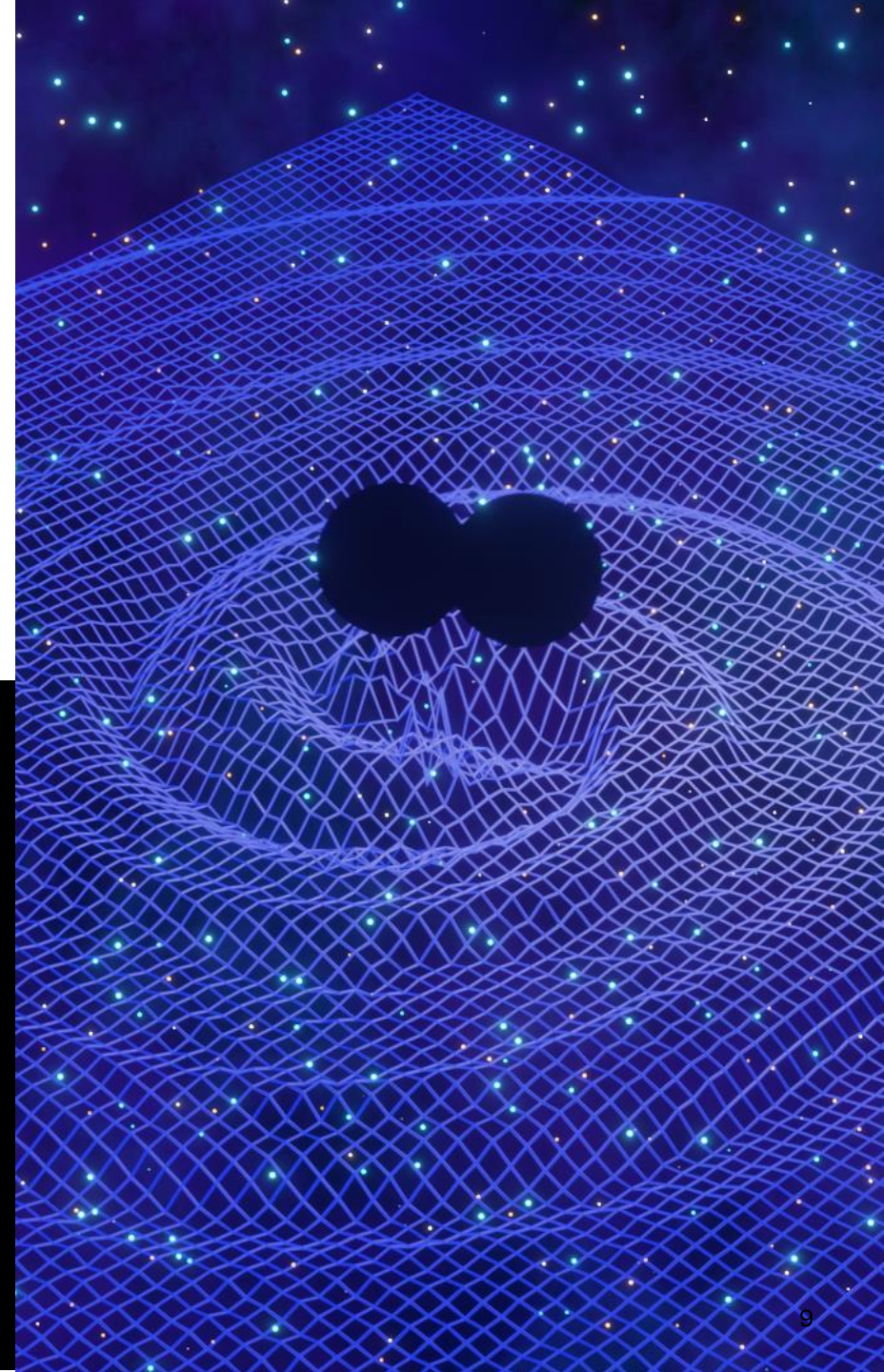
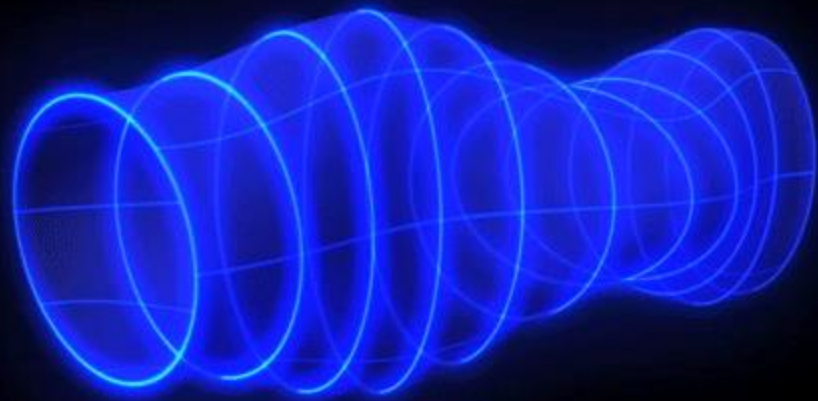
Electroweak Phase Transition



Gravitational waves

- Perturbations in the space-time geometry

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

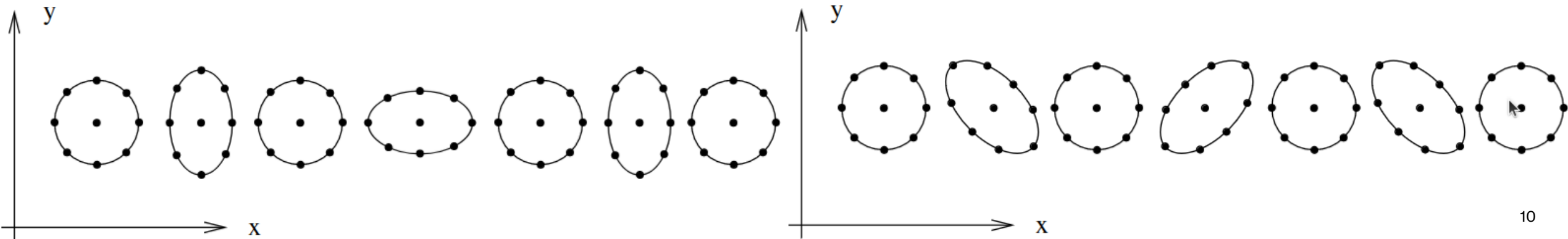


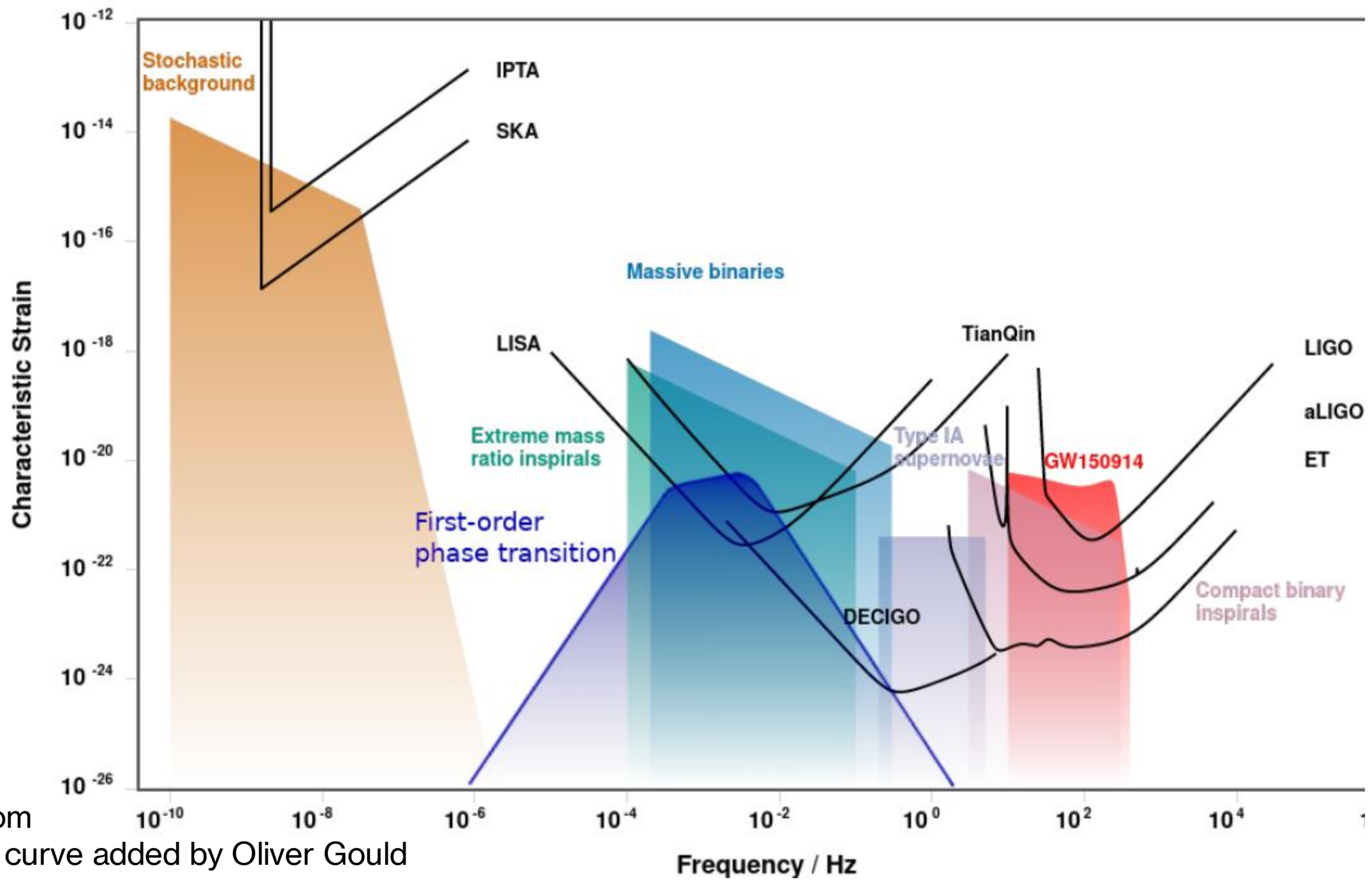
Gravitational waves

- Stochastic gravitational wave background

$$\sum \text{planar wave} \times \text{amplitude} \times \text{polarisation}$$

- GW spectrum usually modelled by broken power laws
- Main contributions: bubble wall collisions, sound waves in the surrounding plasma and turbulence





Gwplotter.com
 1st order PT curve added by Oliver Gould

Gravitational waves

$$\Omega_{\text{gw}} = F(T_*, R_*, \alpha_*, v_w)$$

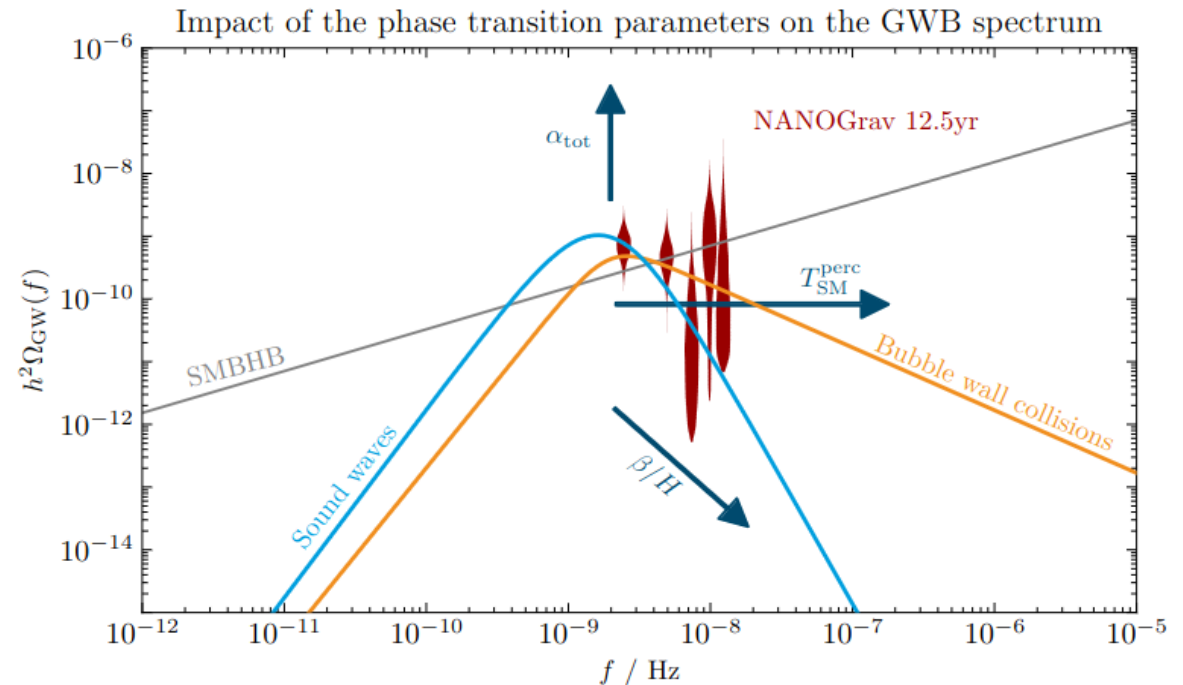
T_* : transition temperature

β_* : inverse duration of the transition

α_* : transition strength

v_w : bubble wall speed

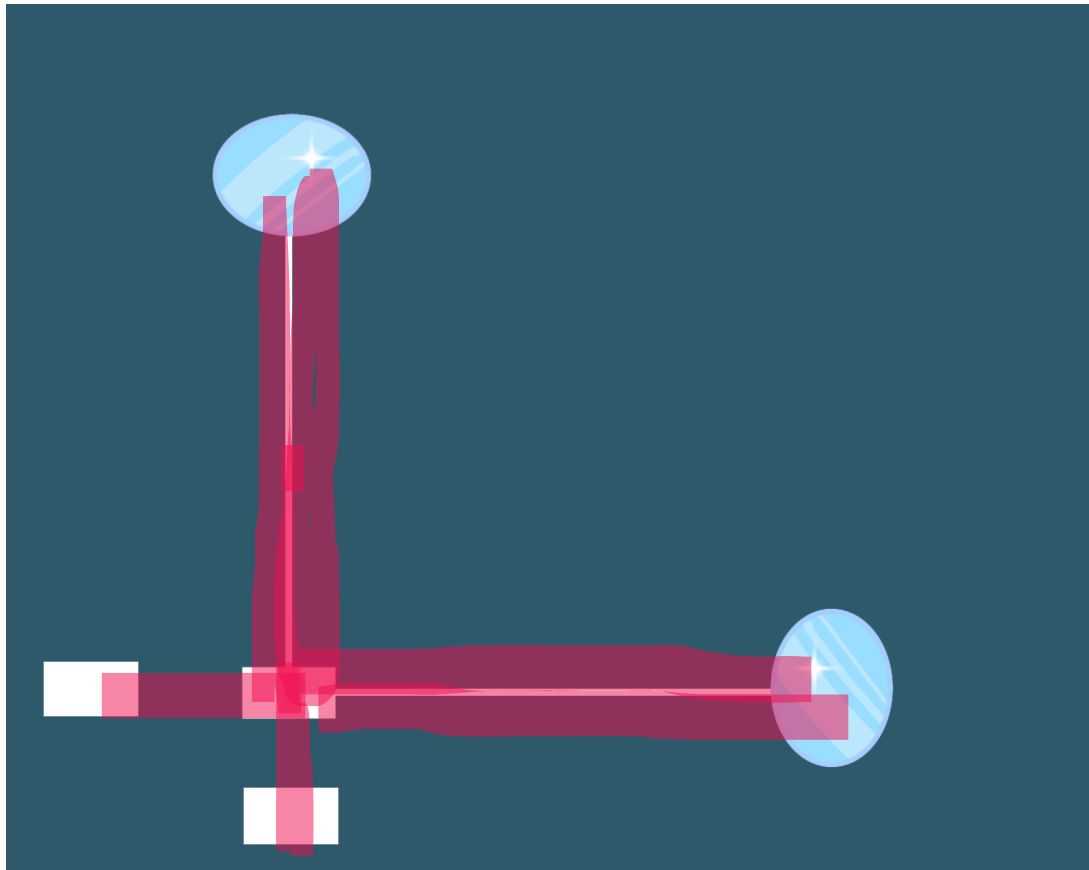
Depend on the nucleation rate!



Bringmann et al. [arXiv: 2306.09411](https://arxiv.org/abs/2306.09411)

GW DETECTION

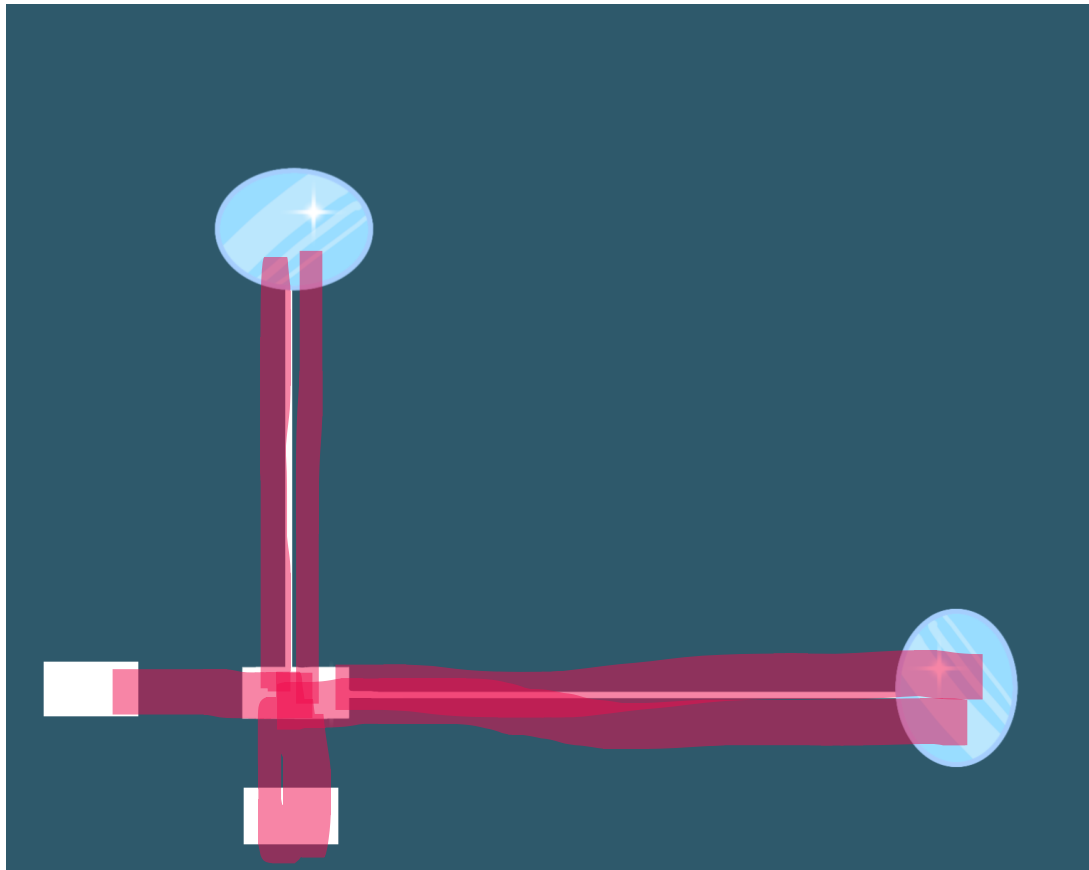
LIGO



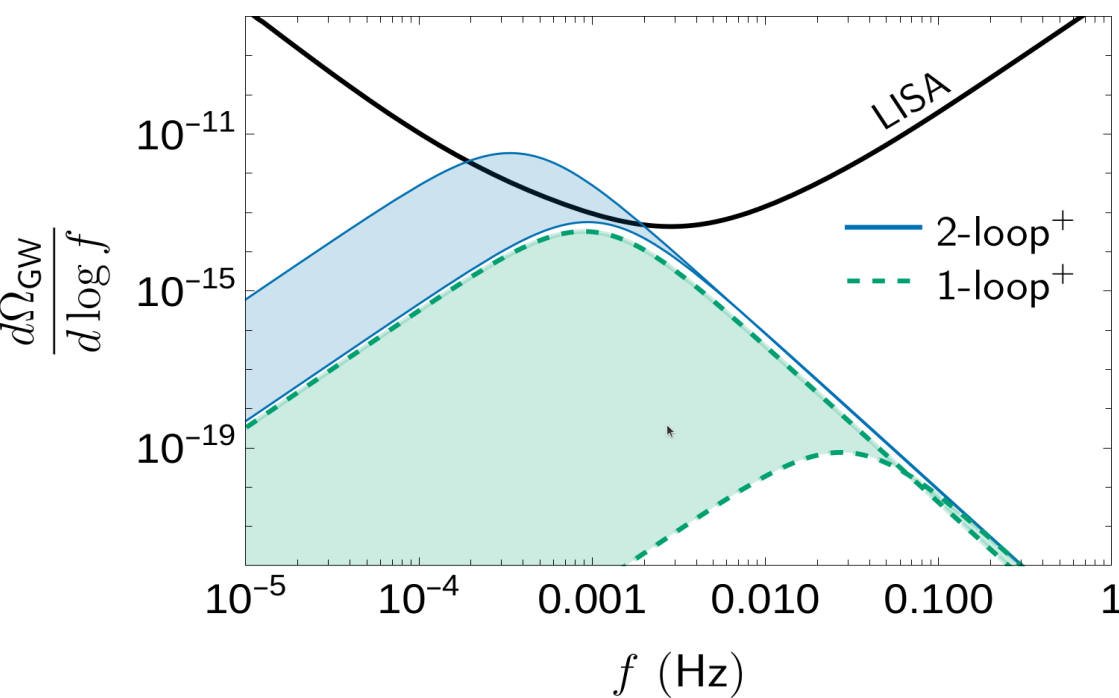
Caltech/MIT/LIGO Lab)

GW DETECTION

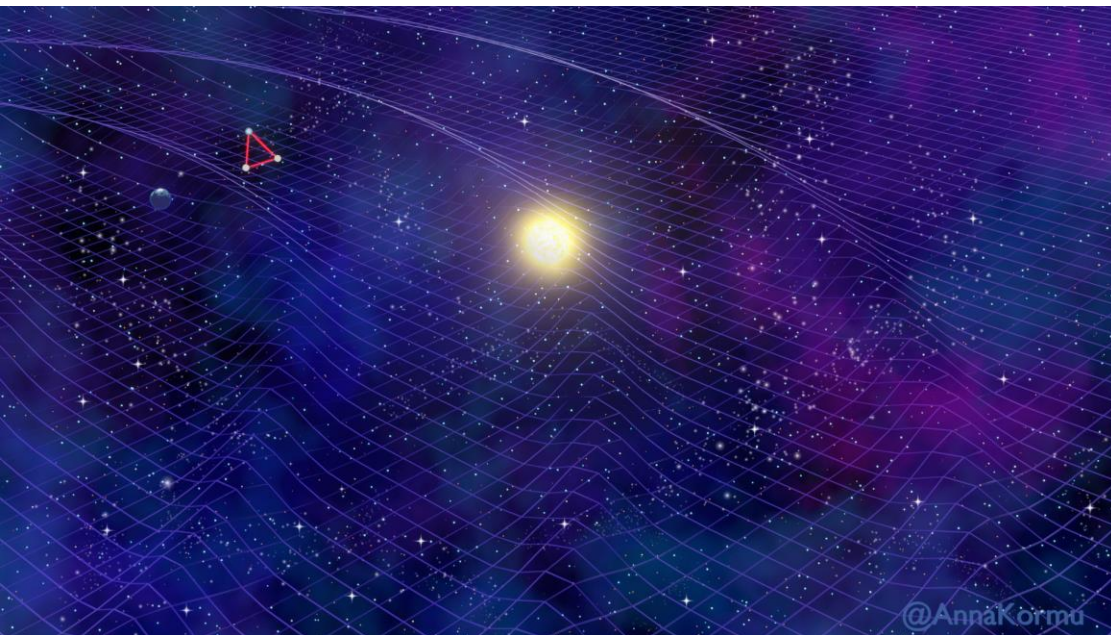
LIGO



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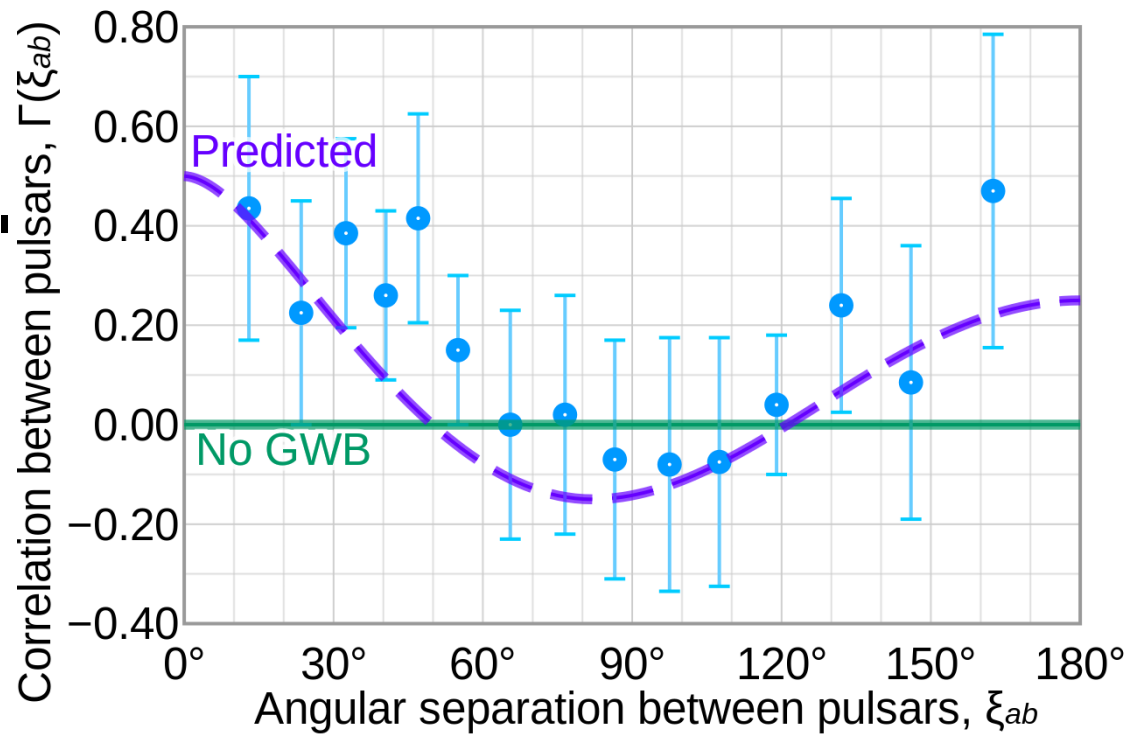
Gould, Tenkanen [arXiv:2104.04399](https://arxiv.org/abs/2104.04399)



GW detection

LISA (AND OTHER SPACE INTERFEROMETERS)

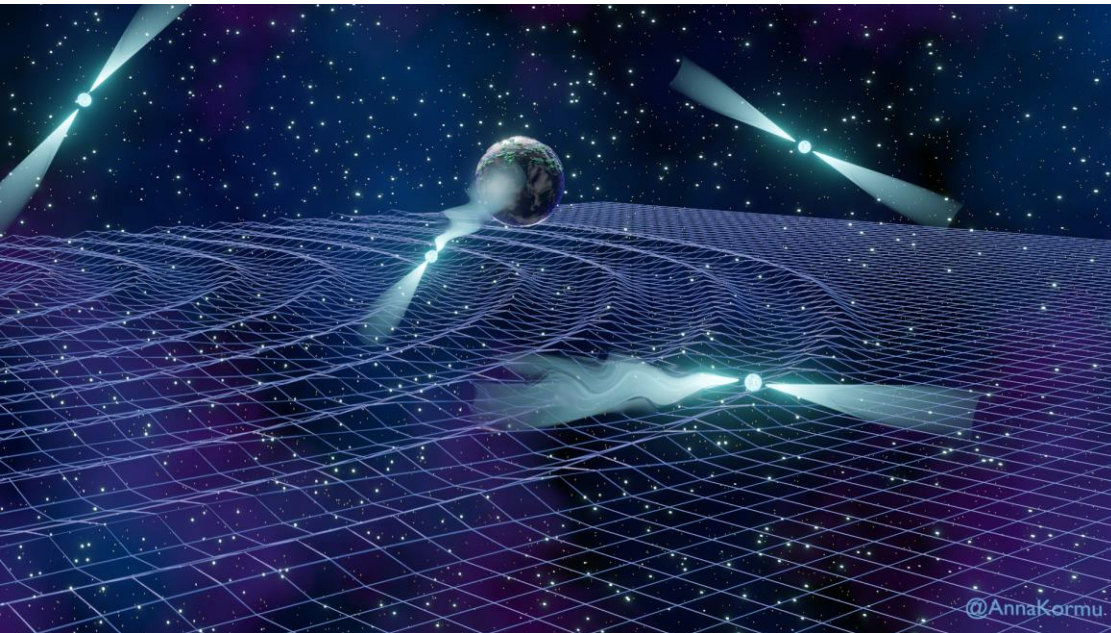
- Launch in 2036, mission adoption 27.1.2024
- Three spacecraft, laser arms 2.5 million km
- Measure changes in path length between spacecraft
- Taiji & TianQin launch in 2030s



GW detection

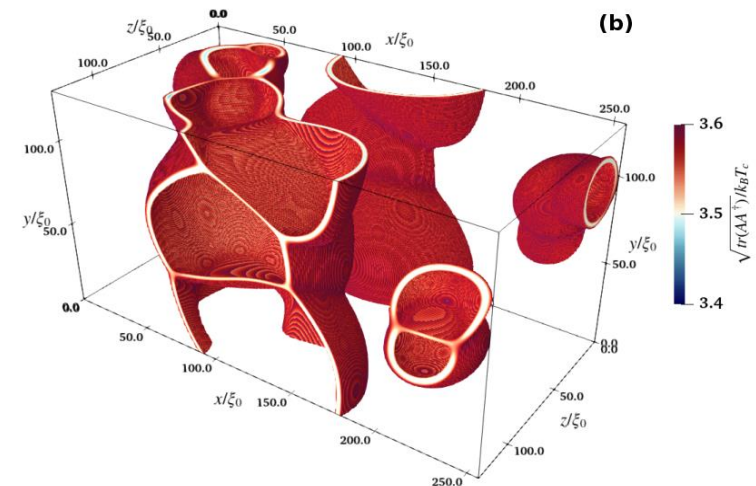
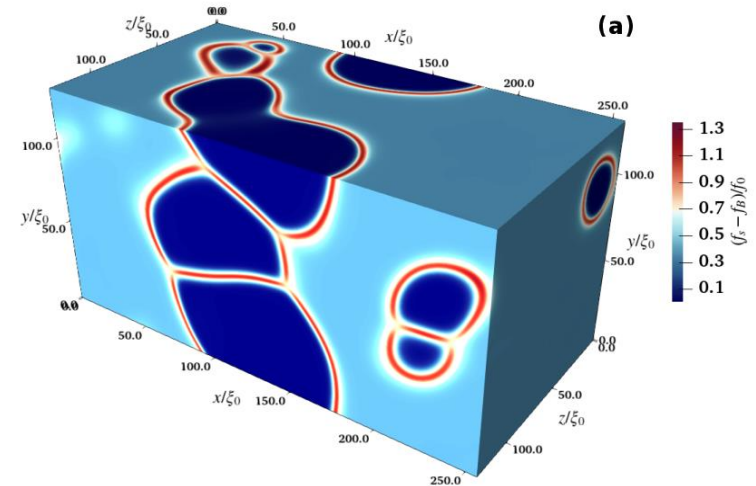
PULSAR TIMING ARRAYS

- Hints of stochastic GW background - June 2023
(European PTA, Indian PTA, NANOGrav, Parkes PTA '23)
- Mostly likely supermassive black holes, but new physics cannot be ruled out yet



Analogue experiments

- Testing cosmology in laboratory: nucleation theory essentially the same in laboratory and in cosmology
- Superfluid Helium-3 [Hindmarsh et al. arXiv:2401.07878](#)
- Ferromagnetic superfluids [Zenesini et al. arXiv:2305.05225](#)
- Proposals to test nucleation in (other) ultracold atomic gases
[arXiv:1408.1163](#) [arXiv:2212.03621](#) [arXiv:2307.02549](#)



[Hindmarsh et al. arXiv:2401.07878](#)

Key points – why the accurate estimation of bubble nucleation rate is important



Baryogenesis

GW
sources

Analogue
experiments

Fate of the False Vacuum

- Relativistic field theory generalisation Callan & Coleman ([Phys. Rev. D 16, 1762 \(1977\)](#))
- Finite temperature approach introduced later by Affleck & Linde ([Phys. Rev. Lett. 46, 388 \(1981\)](#), [Phys. Lett. B 100, 37 \(1981\)](#))

$$\Gamma = A_{\text{dyn}} \times \sqrt{\left| \frac{\det(S''[\phi_0]/2\pi)}{\det'(S''[\phi_b]/2\pi)} \right|} \left(\frac{\Delta S[\phi_b]}{2\pi} \right)^{3/2} e^{-\Delta S[\phi_b]}$$

$$\Gamma = A_{\text{dyn}} \times A_{\text{stat}}$$

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Usually not computed!

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- Determining the prefactor difficult, even in perturbation theory
- Perturbation theory suffers from the so-called infrared problem

$$\frac{g^2}{e^{E/T} - 1} \xrightarrow{E \ll T, \mathbf{p}=0} \frac{g^2 T}{m}$$

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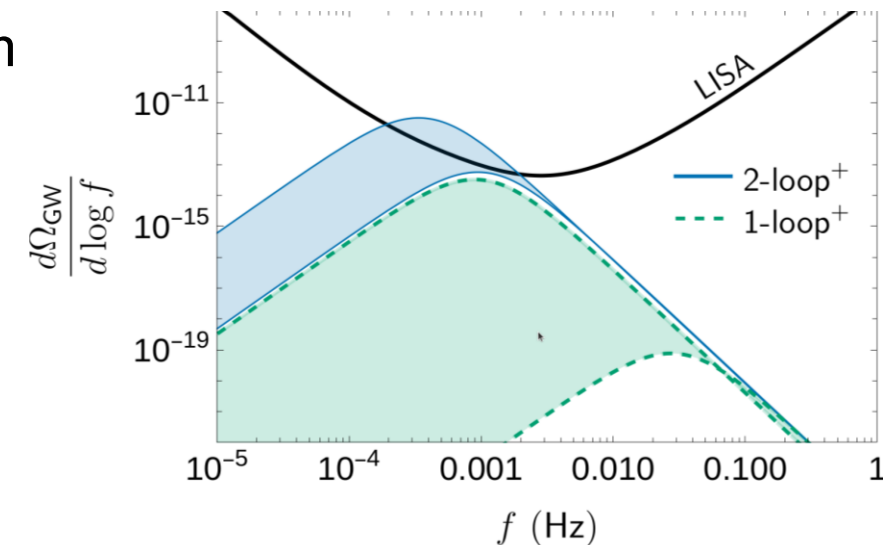
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- Determining the prefactor difficult, even in perturbation theory
- Perturbation theory suffers from the so-called infrared problem
- Introduces uncertainty! How accurate are our cosmological predictions?
- Moore, Rummukainen & Tranberg introduce a simulation method ([hep-lat/0103036](#), [hep-ph/0009132](#))

Key points – why the accurate estimation of bubble nucleation rate is important

Baryogenesis

GW
sources

Analogue
experiments

Uncertainty in
perturbative
calculations

Real scalar theory

For equilibrium dynamics,
see Gould,
[arXiv:2101.05528](https://arxiv.org/abs/2101.05528)

- Toy model possessing key features of BSM models
 - Potential has a tree-level barrier
 - Strong phase transition
 - Perturbative expansion simpler (we understand the dynamics)
- Dimensional reduction (imaginary time, high temp)

$$\mathcal{L} = -\frac{1}{2}\partial_\mu\varphi\partial^\mu\varphi - V(\varphi) - J_1\varphi - J_2\varphi^2$$

Interaction
terms

$$V(\varphi) = \sigma\varphi + \frac{1}{2}m^2\varphi^2 + \frac{1}{3!}g\varphi^3 + \frac{1}{4!}\lambda\varphi^4,$$

Model
parameters

Dimensional reduction

Kajantie et al. [hep-ph/9508379](https://arxiv.org/abs/hep-ph/9508379)

- At high temperatures system looks 3d
- Dimensional reduction 4d cont \rightarrow 3d cont
- Integrate out heavy modes, match correlation functions
- (3d cont \rightarrow 3d lattice)

\mathcal{L}_{4d}



\mathcal{L}_{3d}



$\mathcal{L}_{3d,\text{lattice}}$

Real scalar theory

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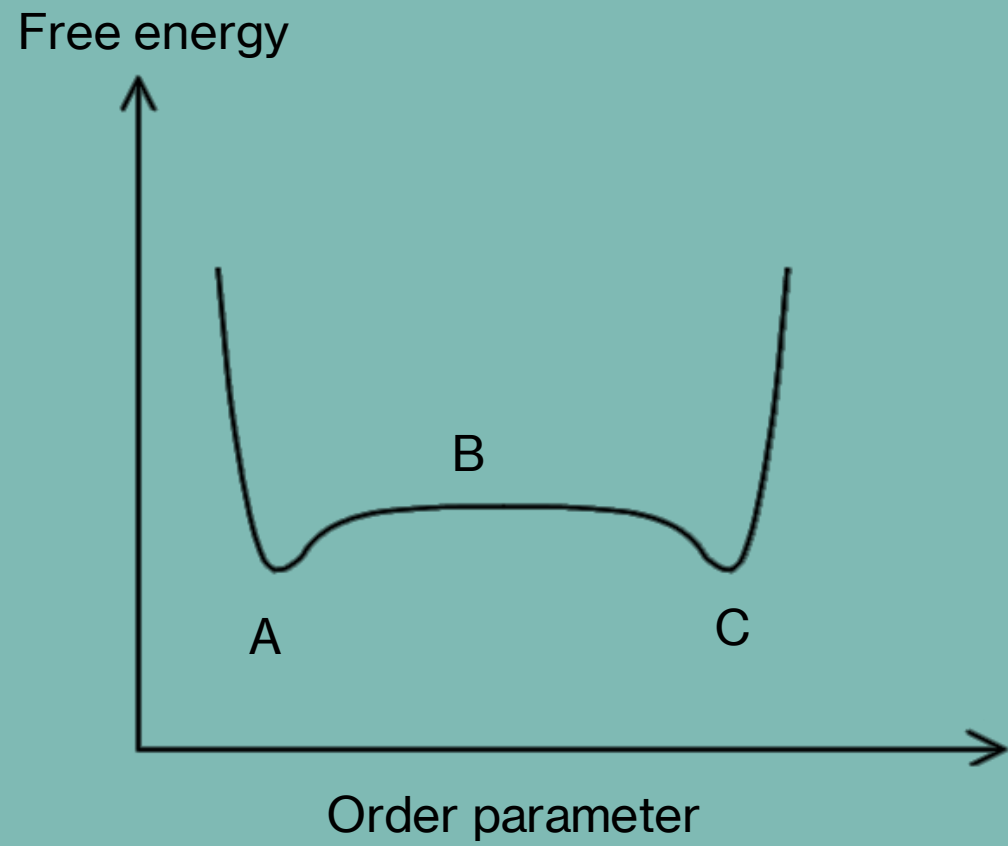
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Lattice spacing

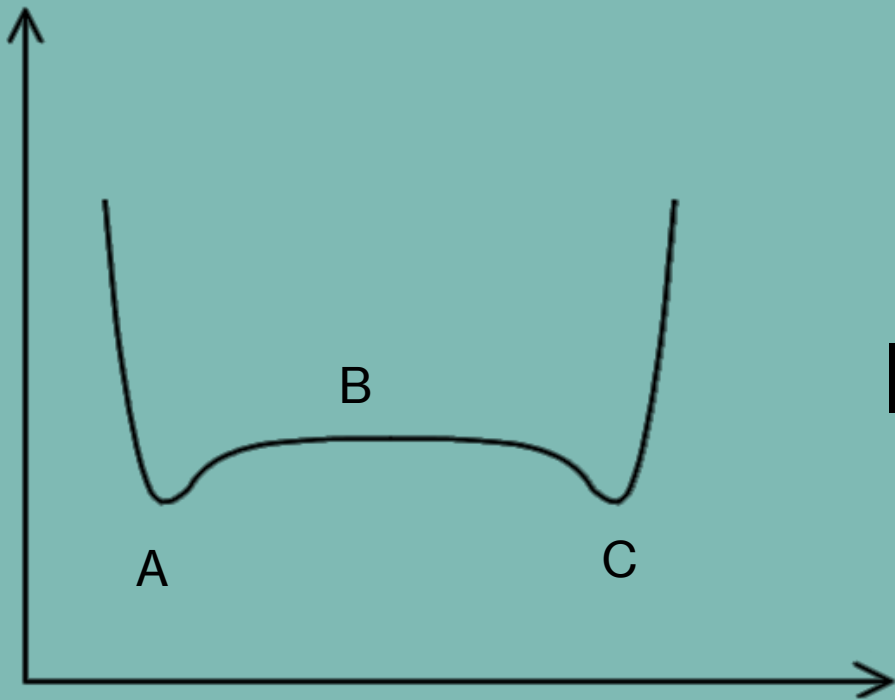
Renormalisation
counterterms

Lattice
parameters

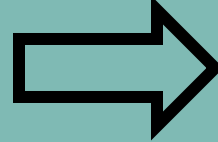
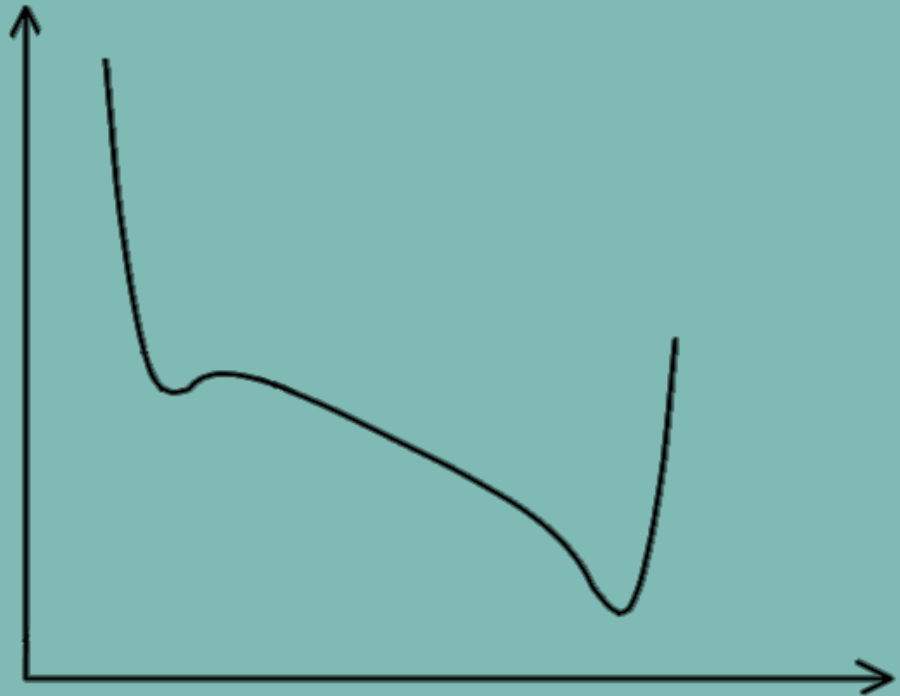
$$S_{\text{lat}} = \sum_x a^3 \left[-\frac{1}{2} Z_\phi \phi_x (\nabla_{\text{lat}}^2 \phi)_x + \sigma_{\text{lat}} \phi_x + \frac{1}{2} Z_\phi Z_m m_{\text{lat}}^2 \phi_x^2 + \frac{1}{3!} g_{\text{lat}} \phi_x^3 + \frac{1}{4!} Z_\phi^2 \lambda_{\text{lat}} \phi_x^4 \right]$$



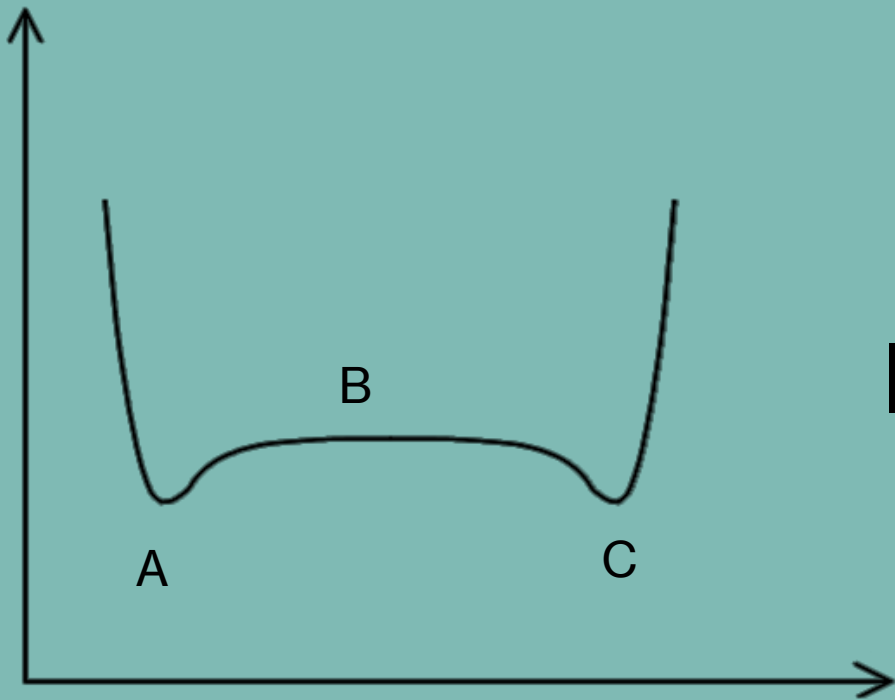
Free energy



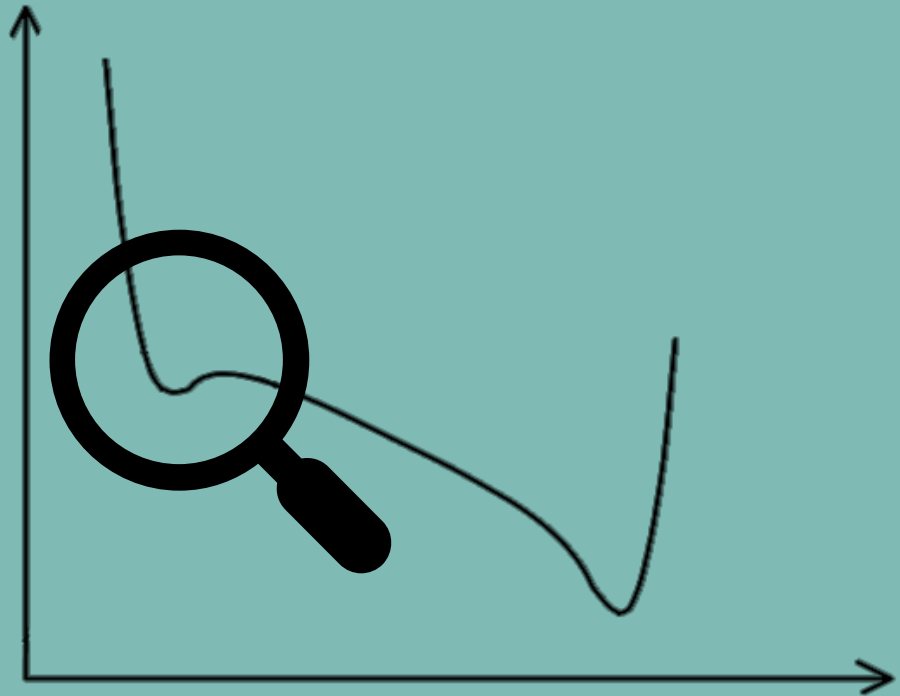
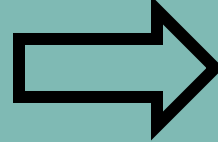
$T < T_c$

A large, hollow black arrow pointing from the left plot to the right plot, indicating a transition in the system's state.

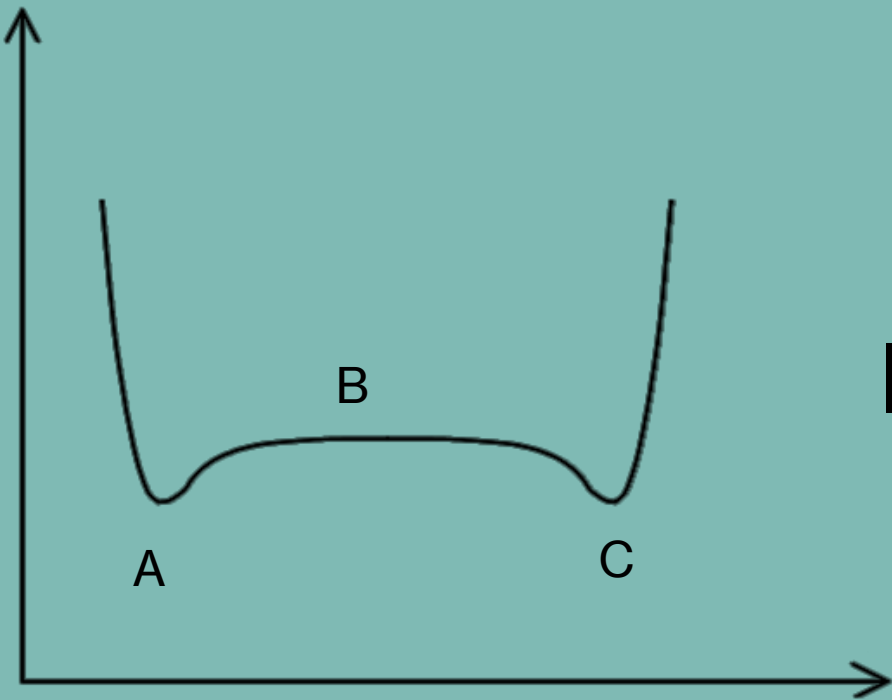
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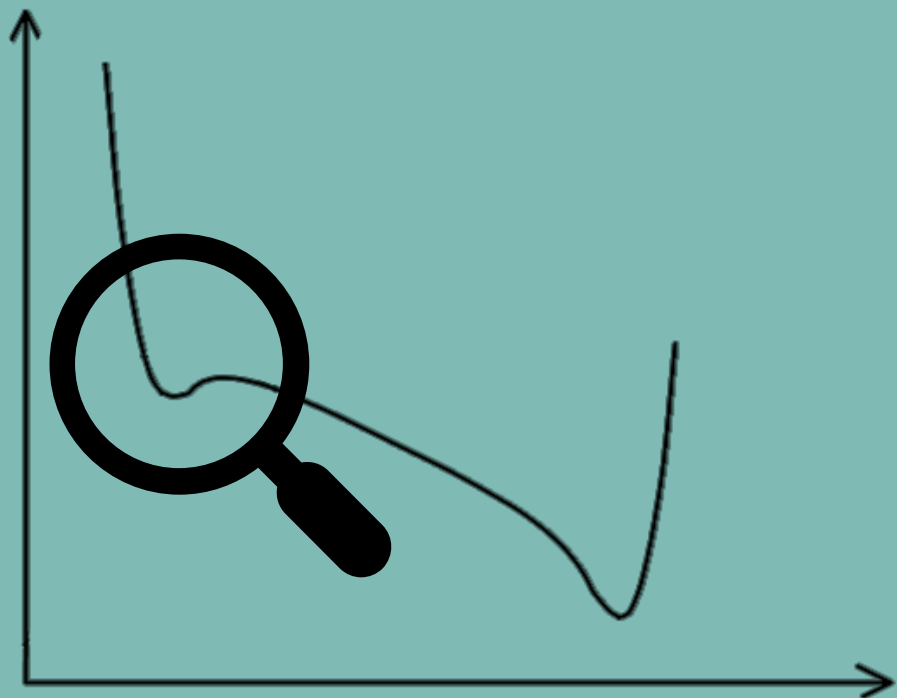
$T < T_c$



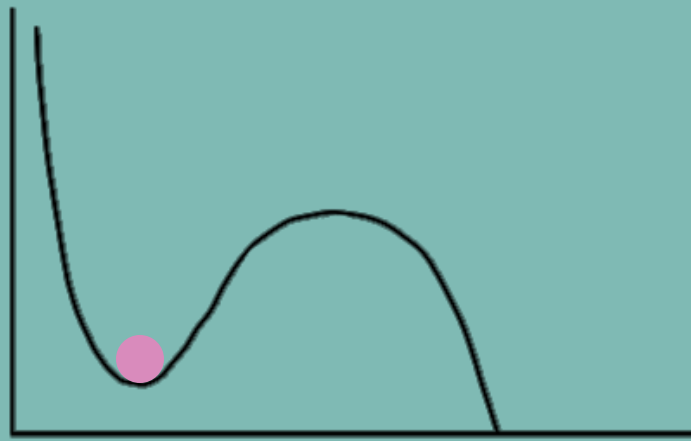
Free energy



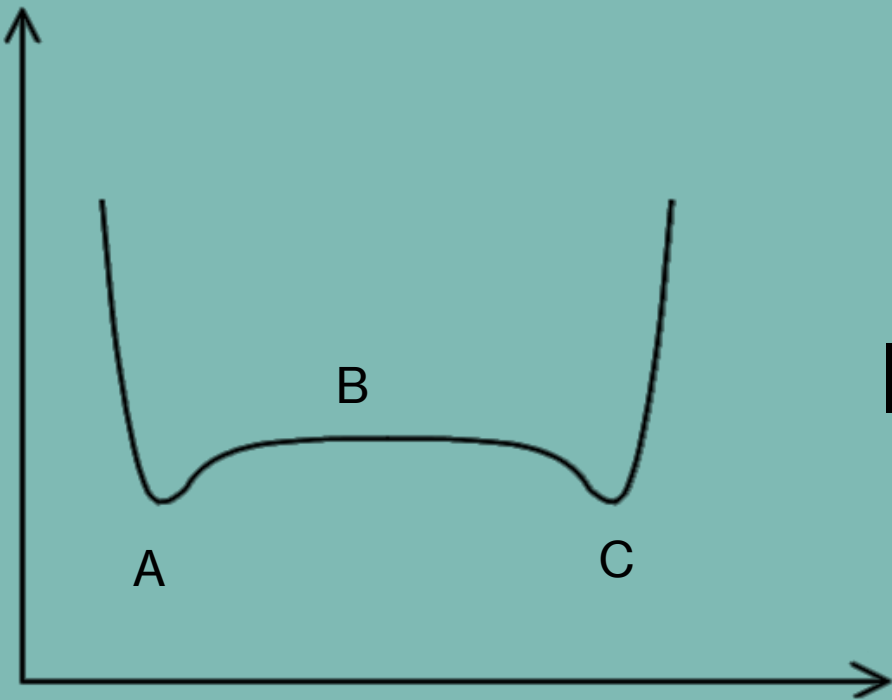
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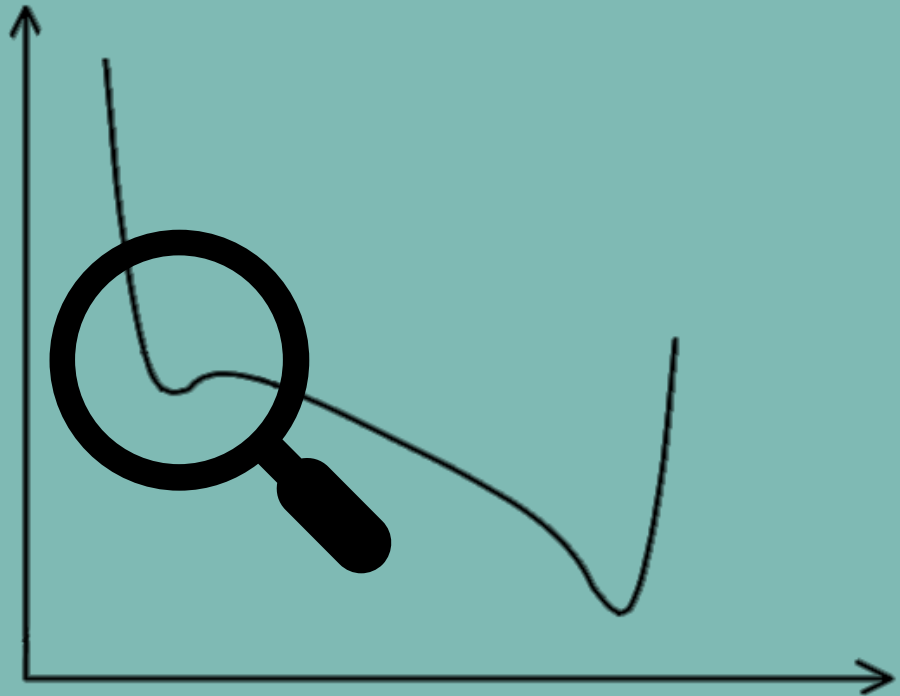
Order parameter



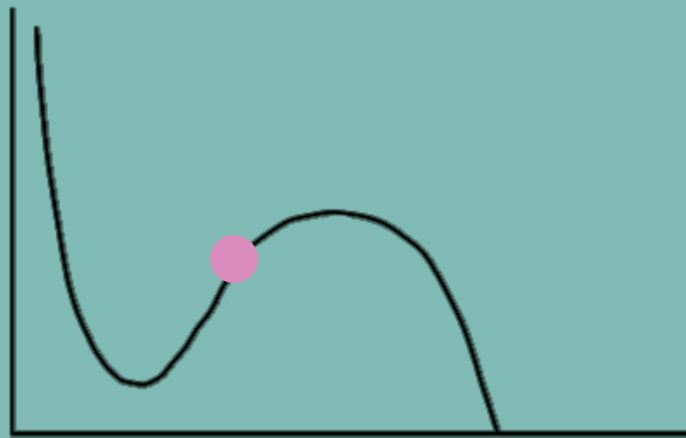
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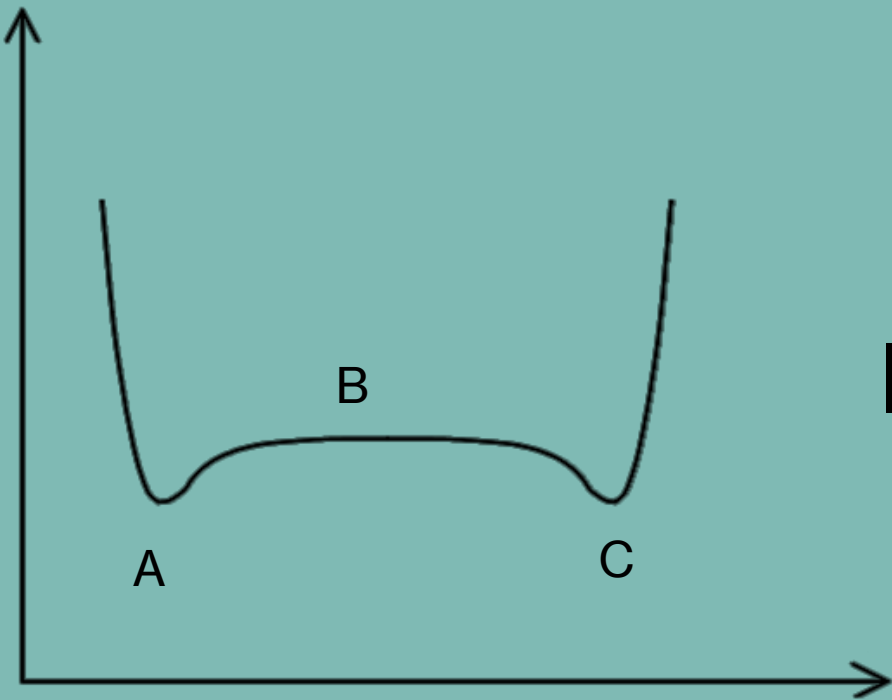
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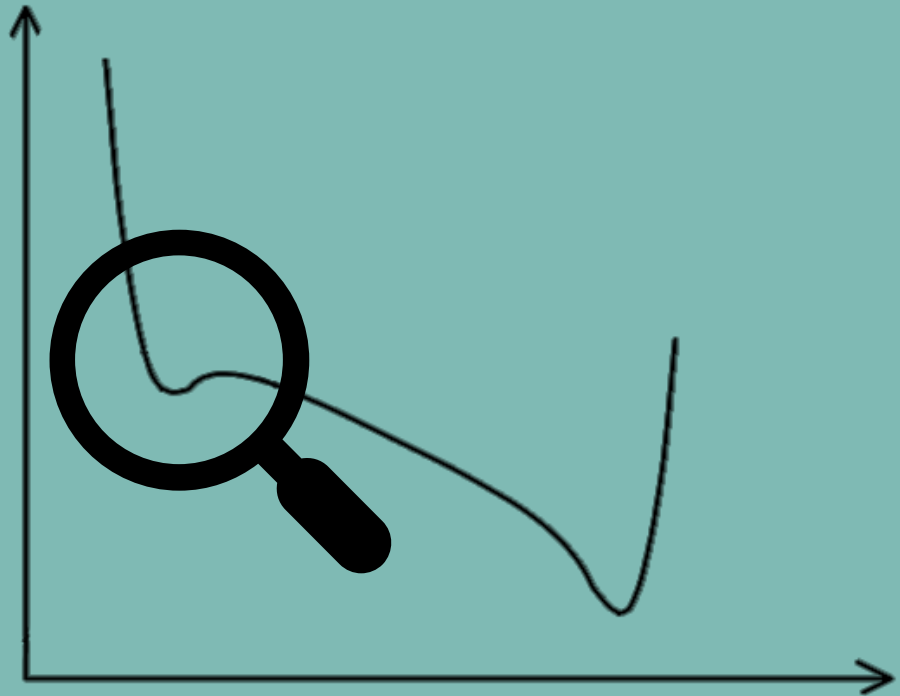
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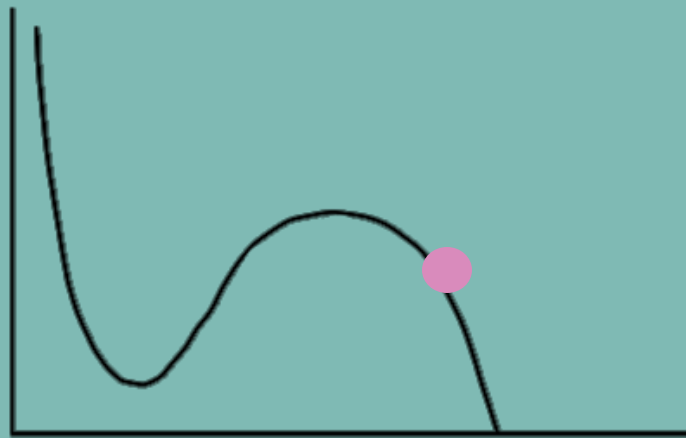
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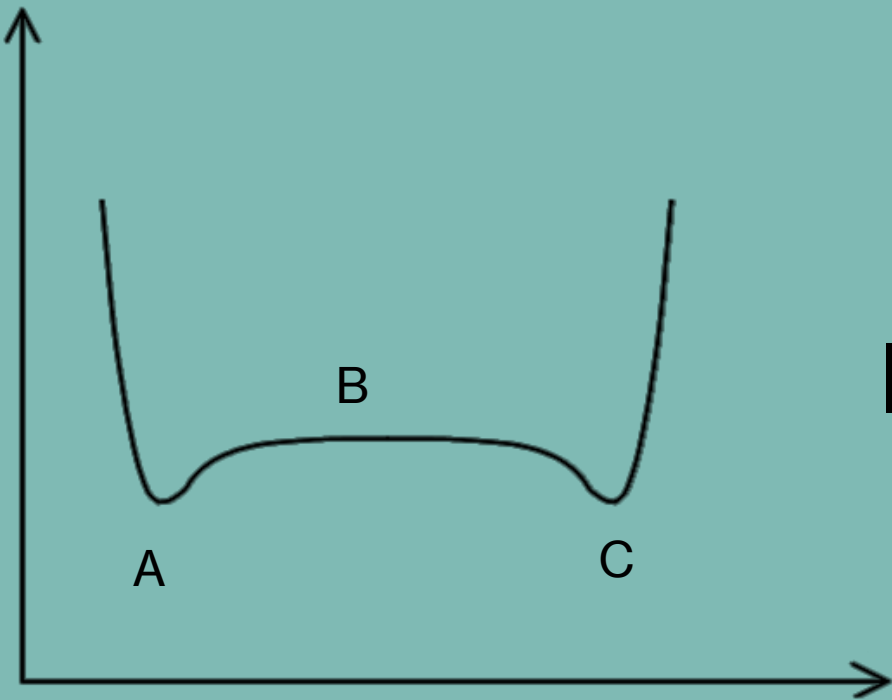
$T < T_c$



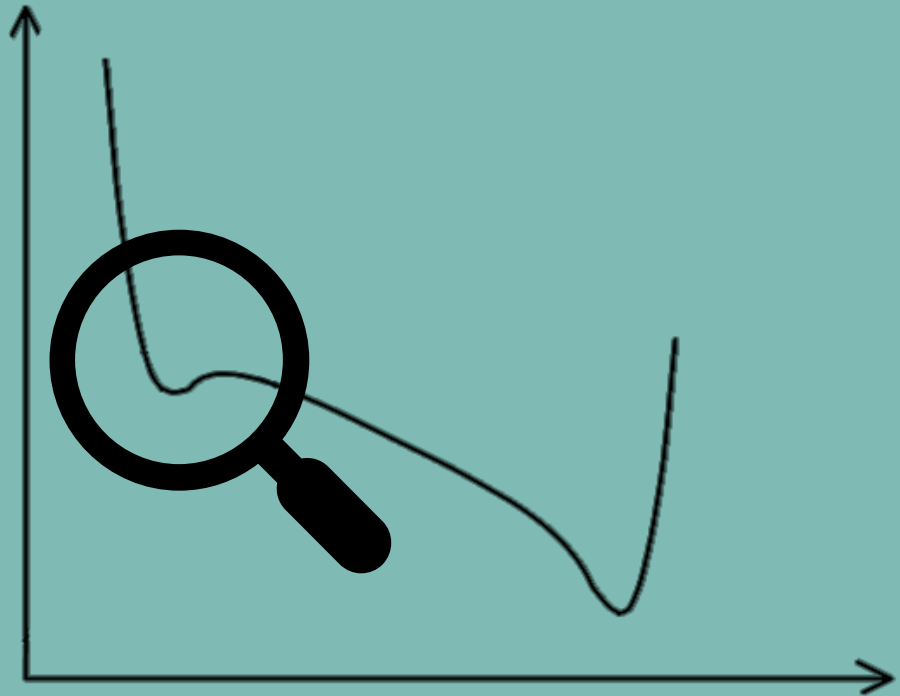
Order parameter



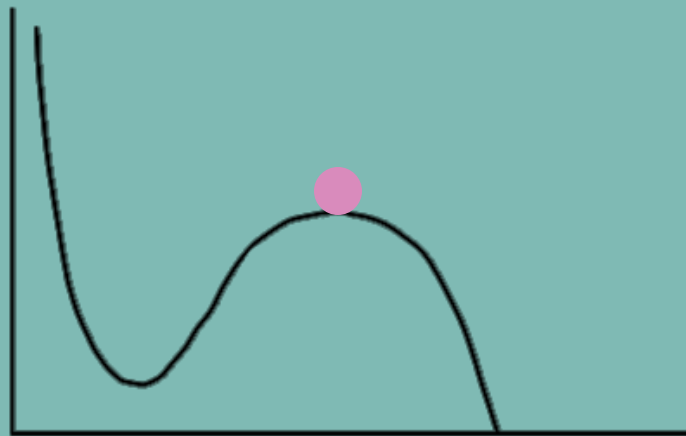
Free energy



$T < T_c$



Order parameter



Bubble nucleation, nonperturbatively

- Langevin equation

$$\partial_t \phi(t, \mathbf{x}) = \pi(t, \mathbf{x}),$$

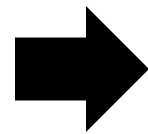
$$\partial_t \pi(t, \mathbf{x}) = -\frac{\delta H_{\text{eff}}}{\delta \phi} - \underbrace{\gamma}_{\text{Damping}} \pi(t, \mathbf{x}) + \underbrace{\xi(t, \mathbf{x})}_{\text{Gaussian noise term}}$$

1.

2.

3.

4.



$$\Gamma = A_{\text{dyn}} \times A_{\text{stat}}$$

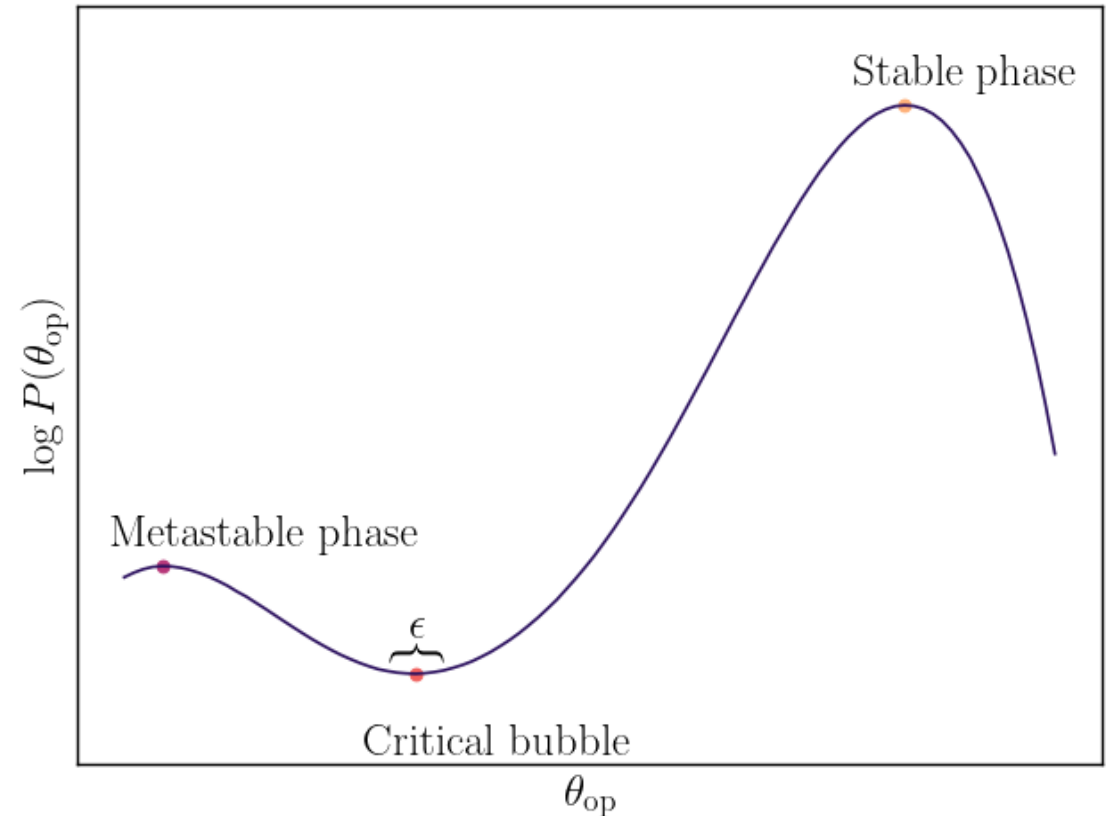
Bubble nucleation, nonperturbatively

1.

Pick an order parameter, simulate probability distribution

- Distinguishable behavior in the two phases

$$\theta'_{\text{op}} = \bar{\phi}$$



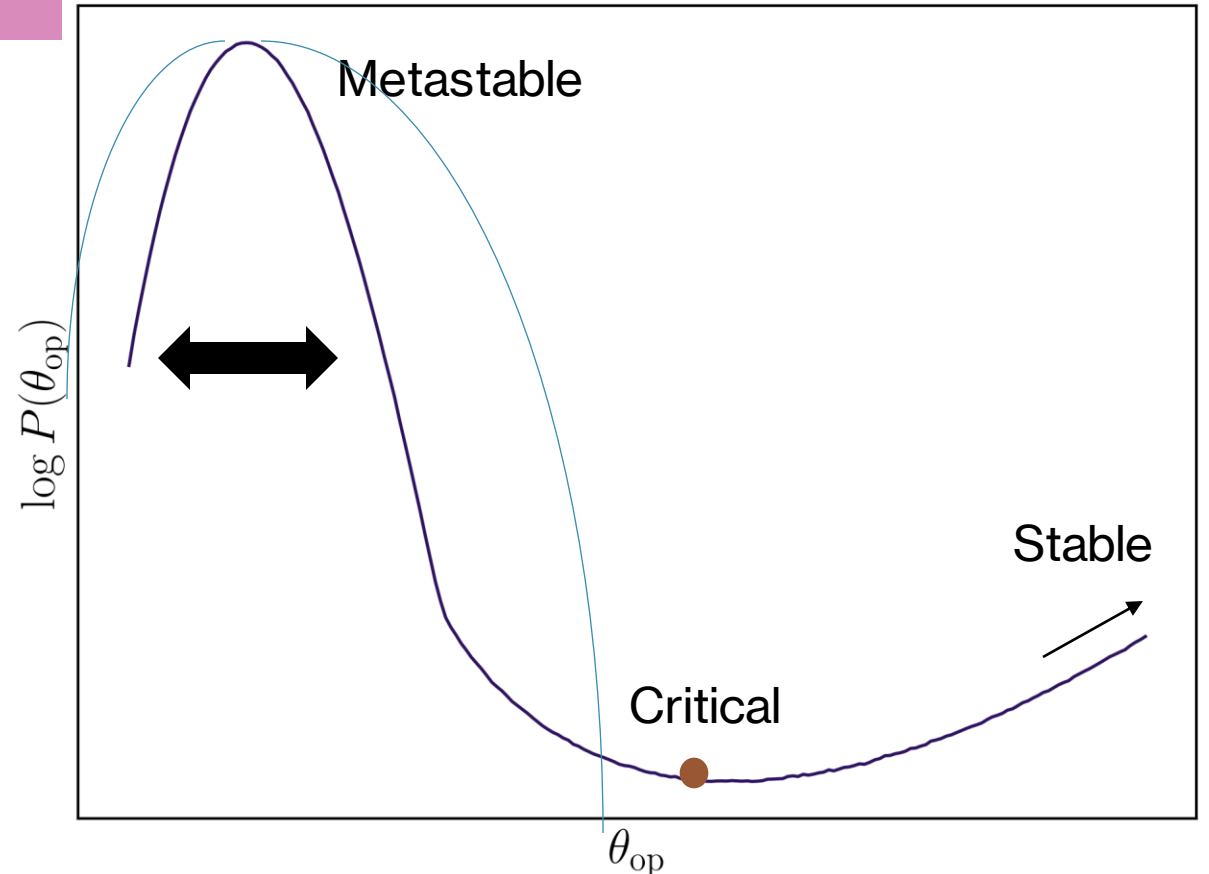
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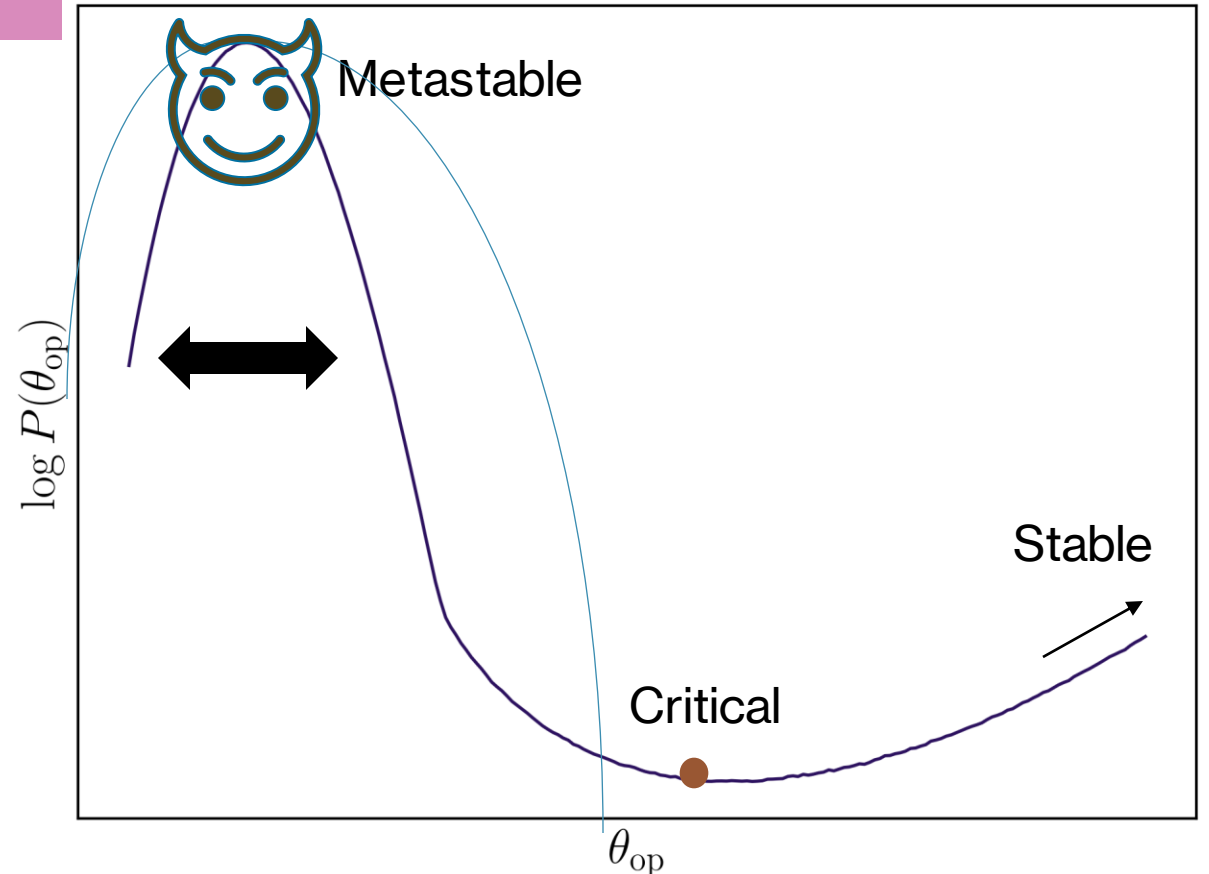
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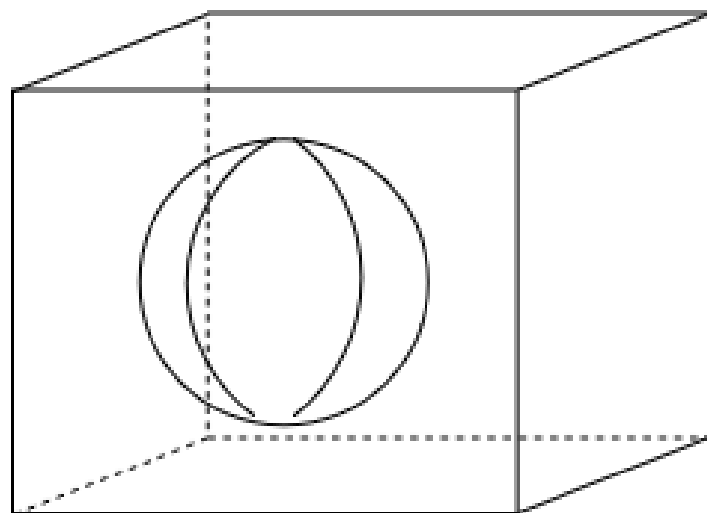
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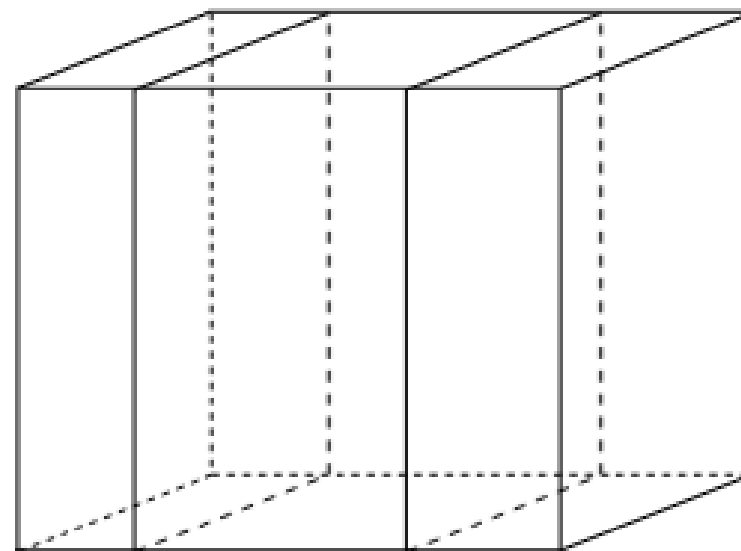
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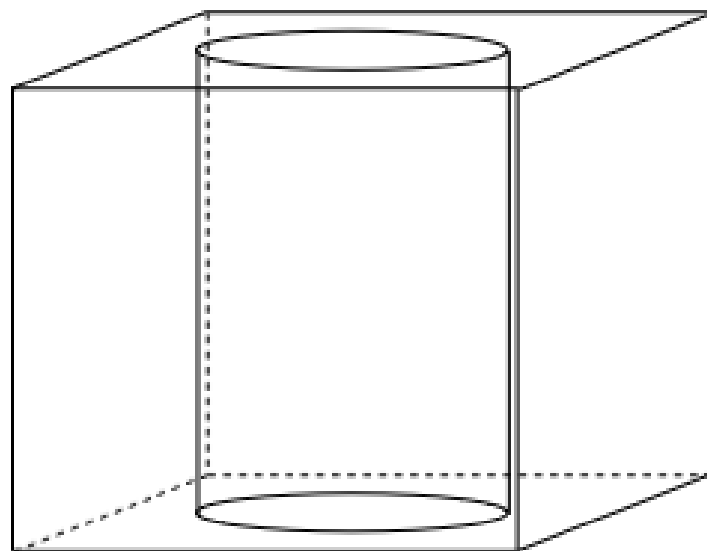




Droplet



Slab



Cylinder

Bubble nucleation, nonperturbatively

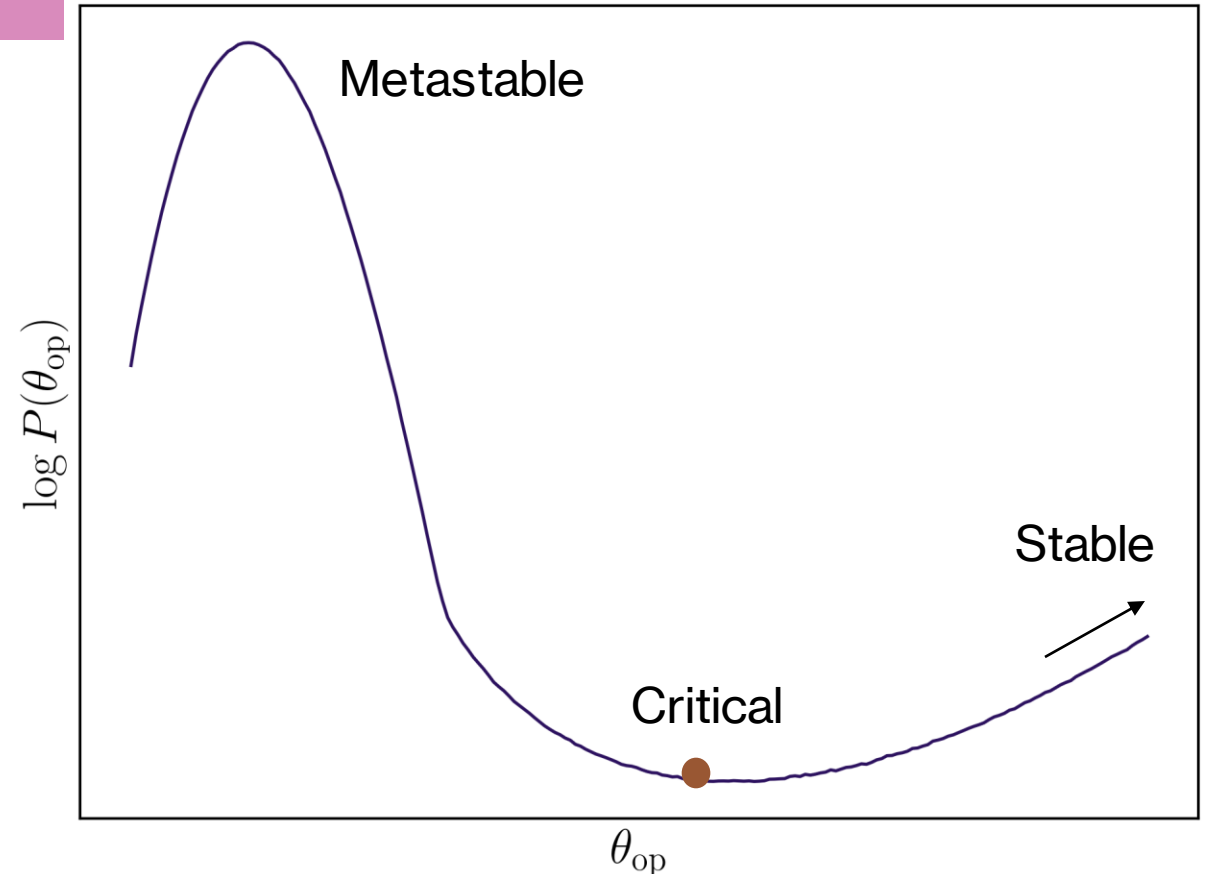
1.

Pick an order parameter, simulate probability distribution

- Distinguishable behavior in the two phases

$$\theta'_{\text{op}} = \bar{\phi} \quad \times$$

$$\theta_{\text{op}} = \bar{\phi}^2 - 2A\bar{\phi} \quad \checkmark$$



Bubble nucleation, nonperturbatively

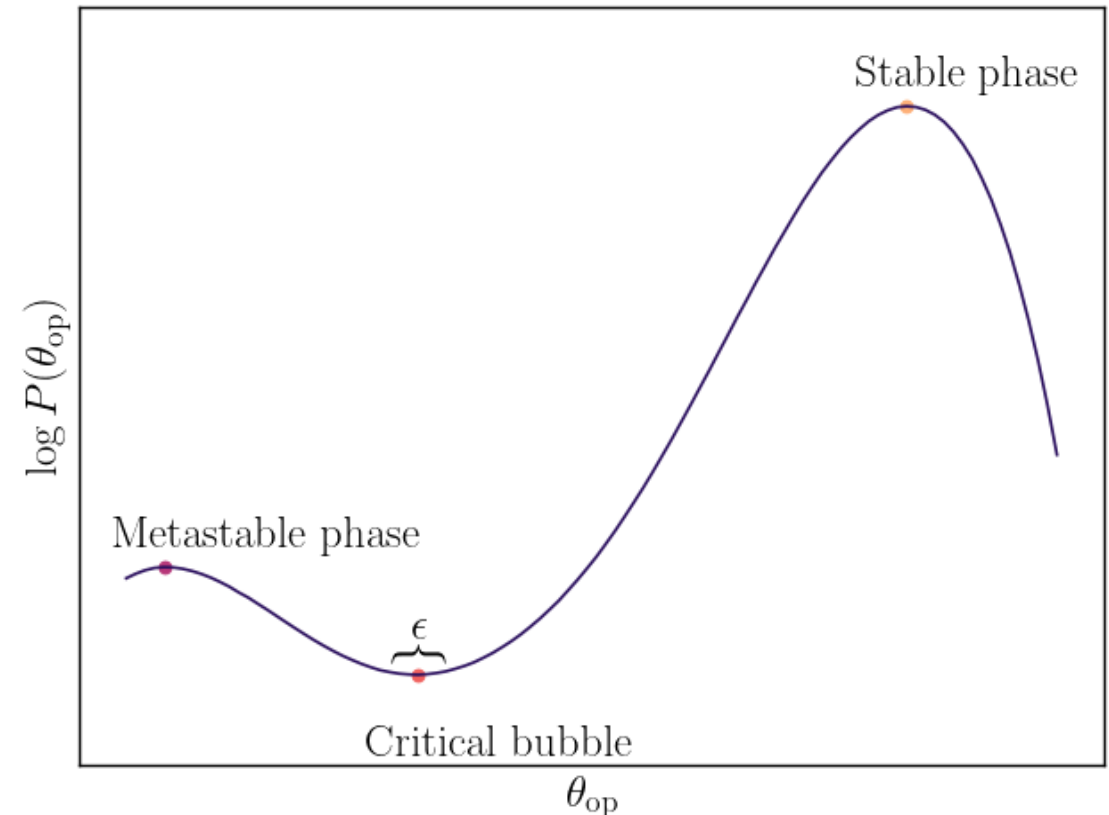
1.

Pick an order parameter, simulate probability distribution

- Distinguishable behavior in the two phases

$$\theta_{\text{op}} = \bar{\phi}^2 - 2A\bar{\phi}$$

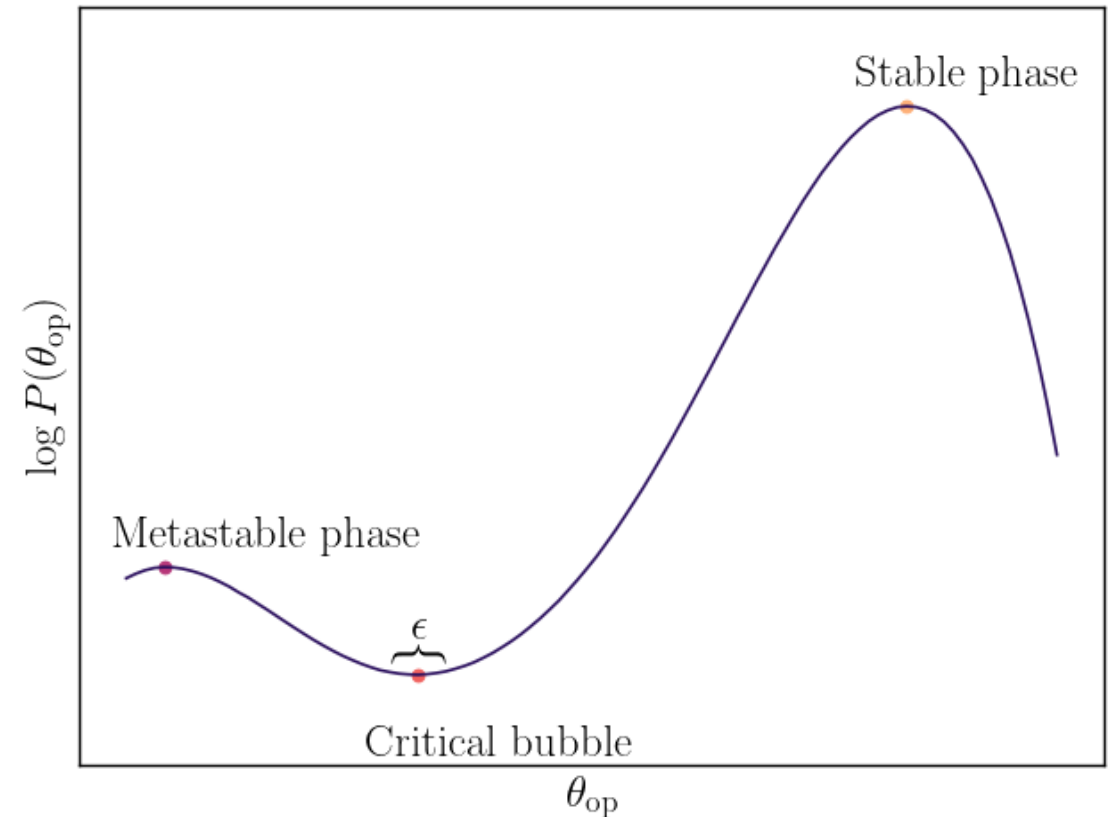
- Separatrix configurations suppressed by e^{-65}
→ Multicanonical Monte Carlo



Bubble nucleation, nonperturbatively

2. Draw separatrix configurations,
calculate the probability

- Draw field configurations from a narrow range $\theta_{\text{op}} \in [\theta_c - \frac{\epsilon}{2}, \theta_c + \frac{\epsilon}{2}]$
- These are initial conditions for the time evolution
- Calculate probability of the critical bubble and normalise it to the probability of being in the metastable phase

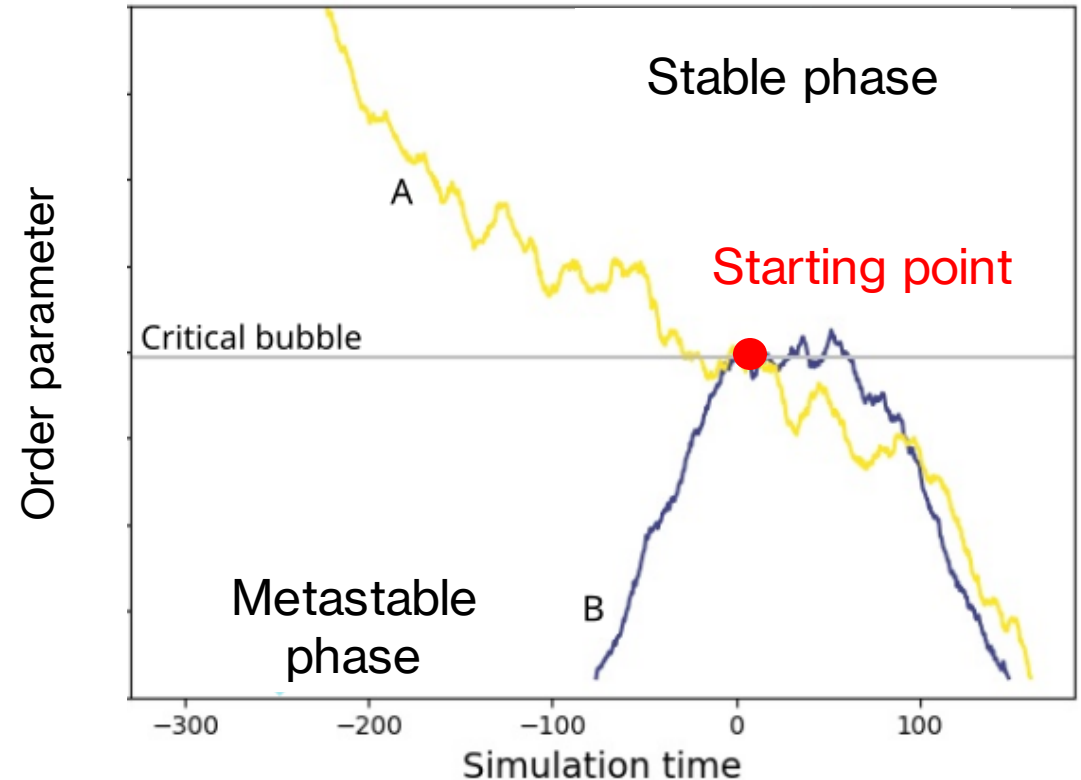


Bubble nucleation, nonperturbatively

3. Determine the tunnelling fraction

- Time evolve field configurations
- Determine if a time evolved trajectory has tunneled
- If metastable \rightarrow stable (or vice versa) then $\delta_{\text{tunnel}} = 1$

$$\mathbf{d} = \frac{\delta_{\text{tunnel}}}{N_{\text{crossings}}}$$

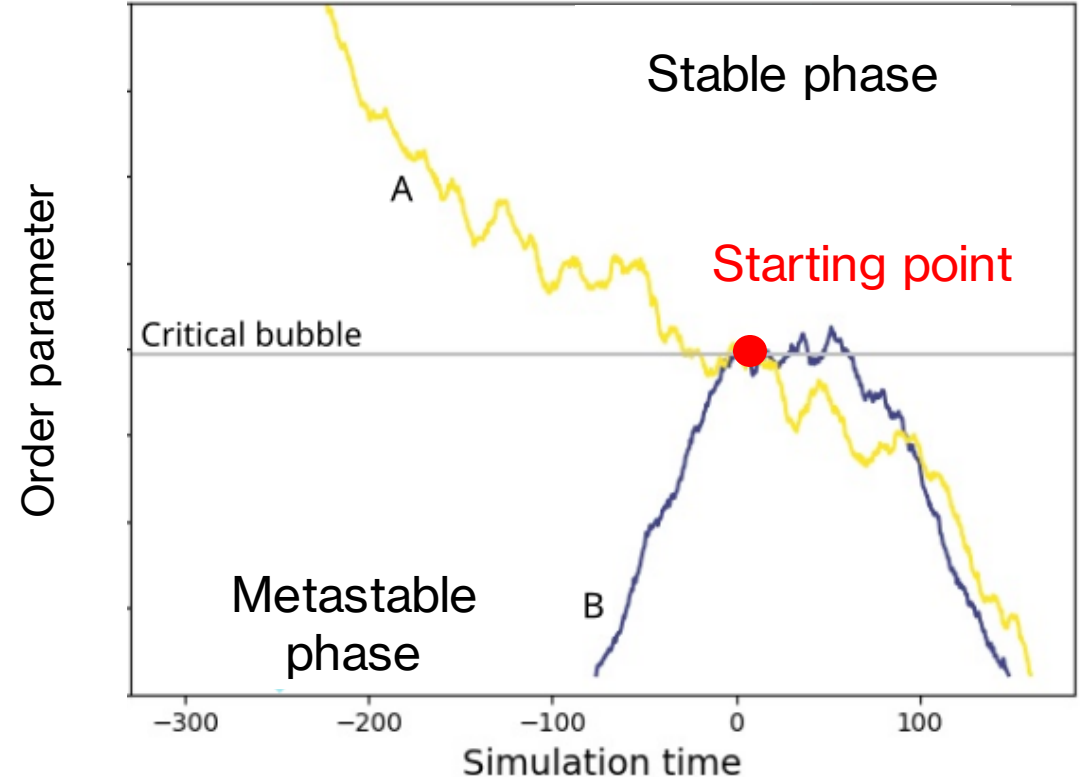


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Bubble nucleation, nonperturbatively

3. Determine the tunnelling fraction

- Forest-Ruth, 4th order accurate symplectic integrator built from Leapfrog

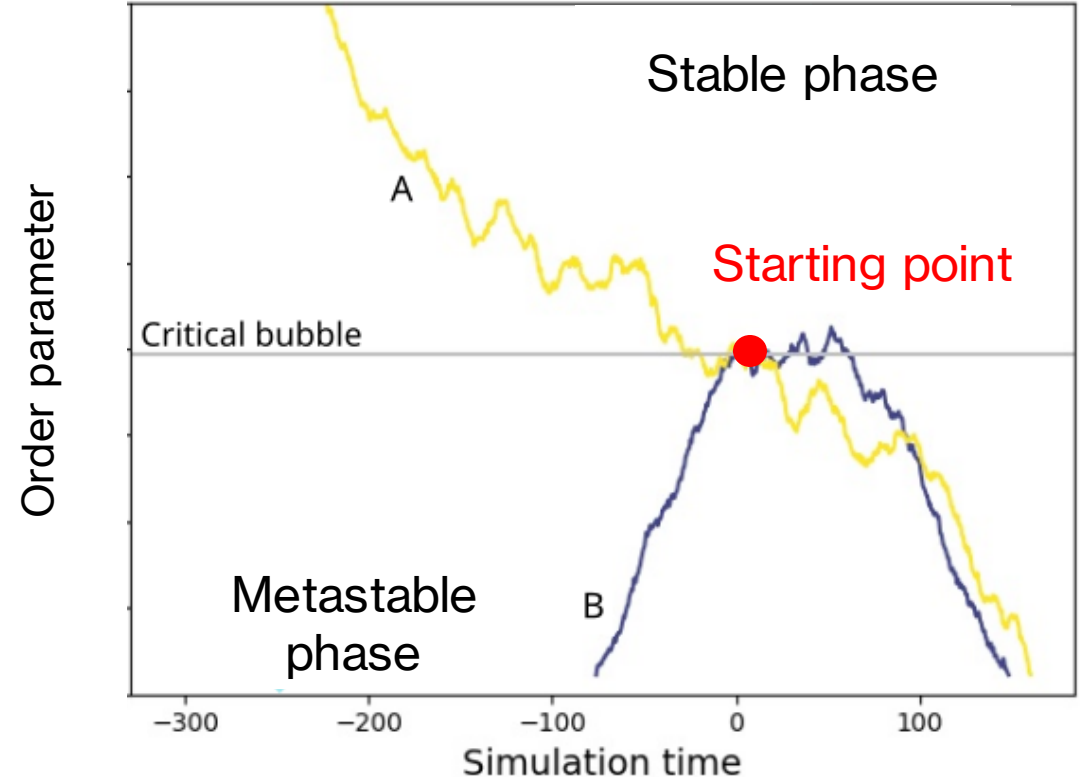
Leapfrog

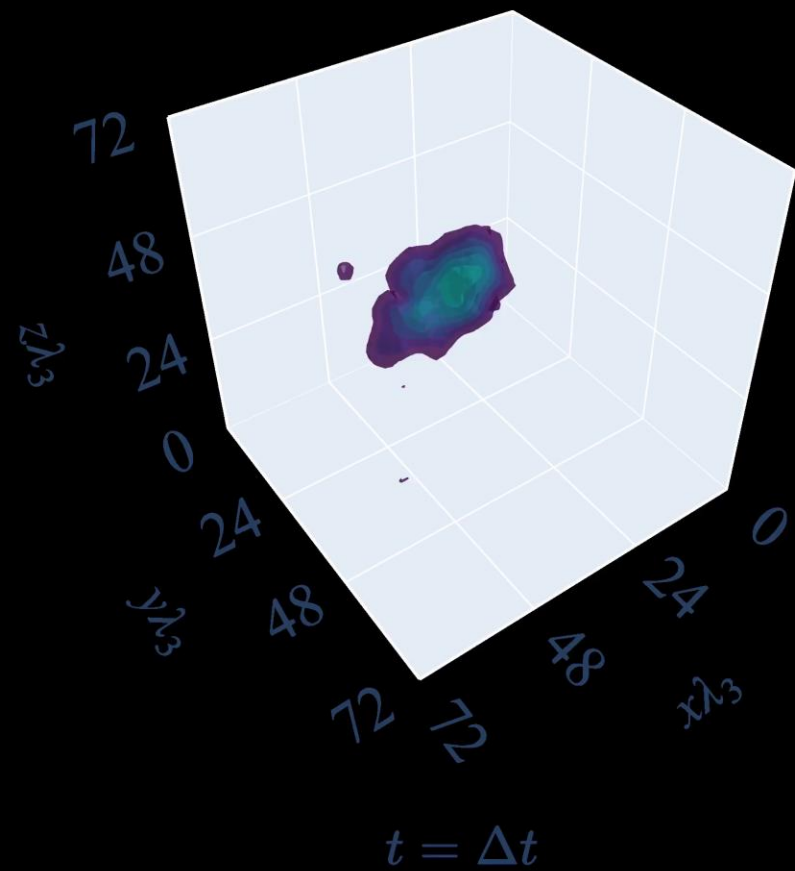
$$\pi_{t+\frac{1}{2},x} = \pi_{t,x} - \frac{\partial H_{\text{eff}}}{\partial \phi_{t,x}} \frac{\Delta t}{2},$$

$$\phi_{t+1,x} = \phi_{t,x} + a^3 \pi_{t+\frac{1}{2},x} \Delta t,$$

$$\pi_{t+1,x} = \pi_{t+\frac{1}{2},x} - \frac{\partial H_{\text{eff}}}{\partial \phi_{t+1,x}} \frac{\Delta t}{2}$$

- + noise and damping in momentum refresh





Bubble nucleation, nonperturbatively

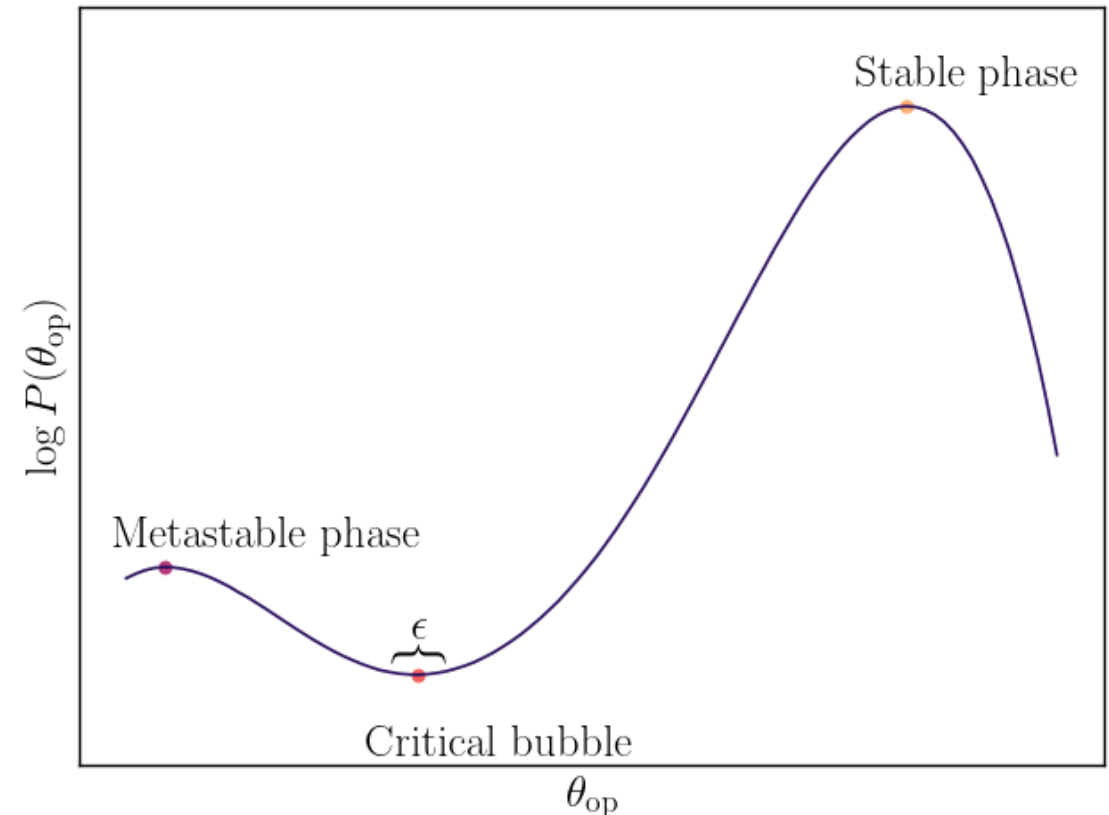
4.

Rate of change of the order parameter

- Flux - describes the rate of change of the order parameter as it crosses the separatrix

$$\langle \text{flux} \rangle = \left\langle \left| \frac{\Delta \theta_{\text{op}}}{\Delta t} \right|_{\theta_c} \right\rangle$$

- Can be solved analytically!
- Order parameter dependent!



Nucleation rate, non-perturbatively

1. Pick an order parameter, simulate probability distribution

2. Draw separatrix configurations, calculate the probability

3. Determine the tunnelling fraction

4. Rate of change of the order parameter

$$P_c^{\text{normalised}}$$

$$\frac{1}{2} \langle \mathbf{d} \rangle$$

$$\langle \text{flux} \rangle$$

Nucleation rate, non-perturbatively

1. Pick an order parameter, simulate probability distribution

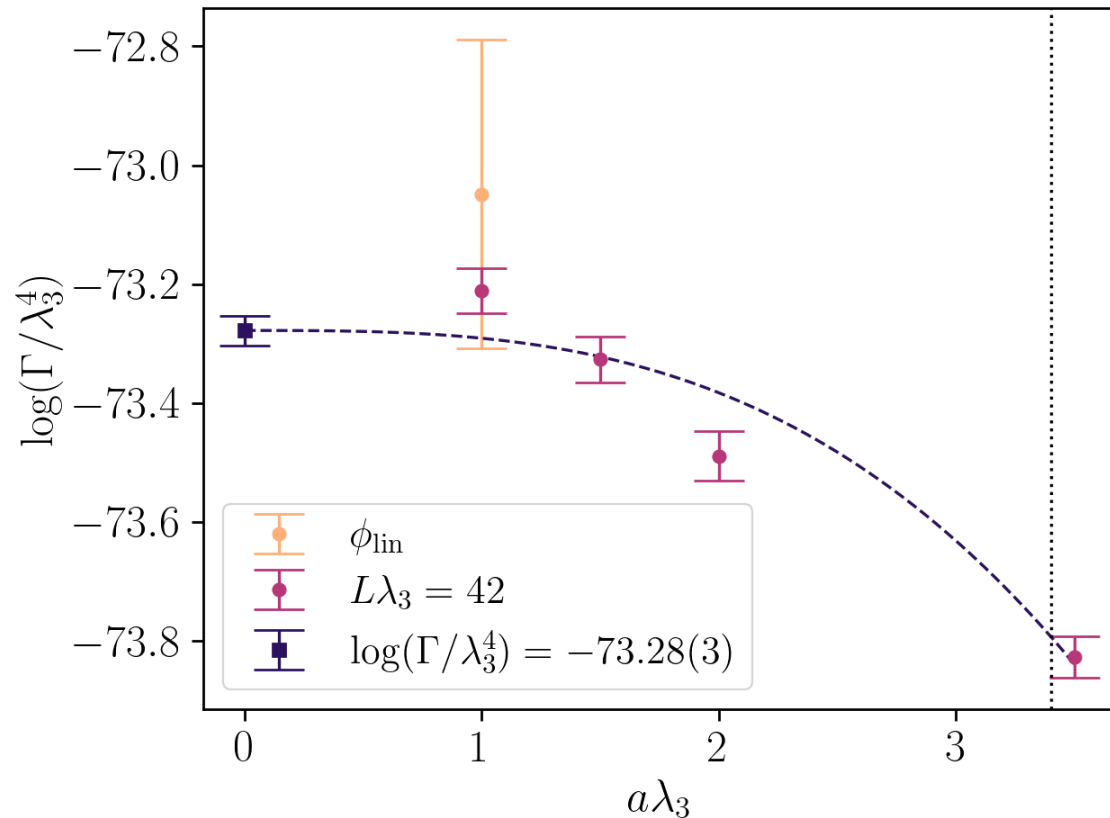
2. Draw separatrix configurations, calculate the probability

3. Determine the tunnelling fraction

4. Rate of change of the order parameter

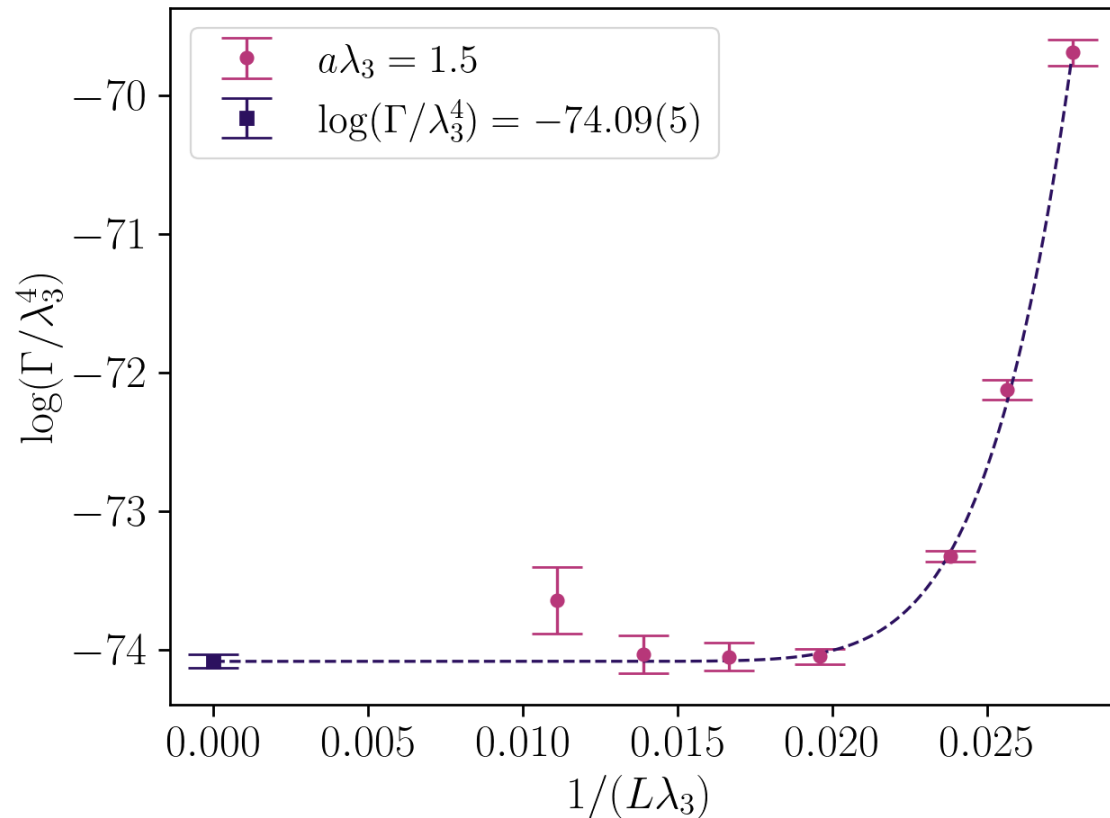
$$\Gamma \mathcal{V} \approx P_c^{\text{normalised}} \frac{1}{2} \langle \text{flux} \rangle \langle \mathbf{d} \rangle$$

Results – Continuum limit



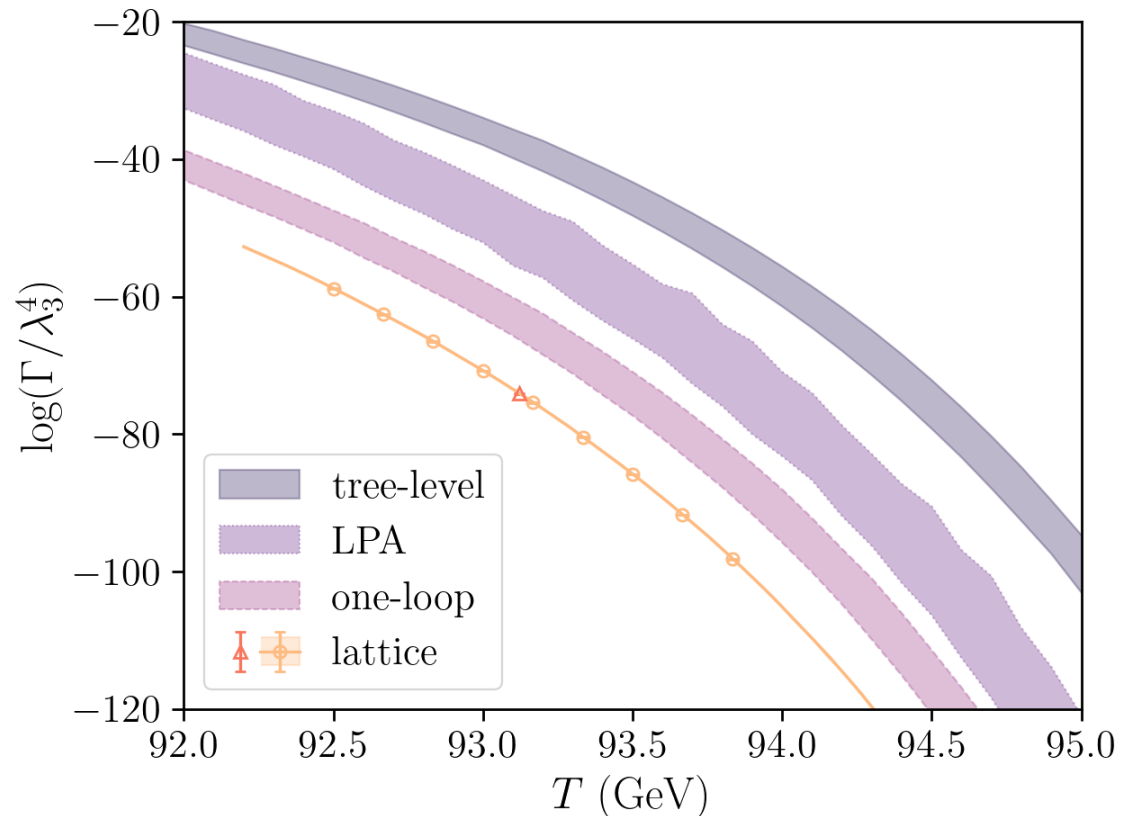
- Zero lattice spacing extrapolation
- We use $O(a^2)$ improved lattice discretisation
- One-loop approx for screening mass is 0.294
- Linear order parameter for reference, reduced errors in quadratic

Results – Volume limit



- Long range correlations die off exponentially with distance (3d model has no massless modes)
- \rightarrow we fit an exponential
- We find agreement with the perturbative screening mass
$$m_s/\lambda_3 = 0.287(12)$$

Results – Nucleation rate

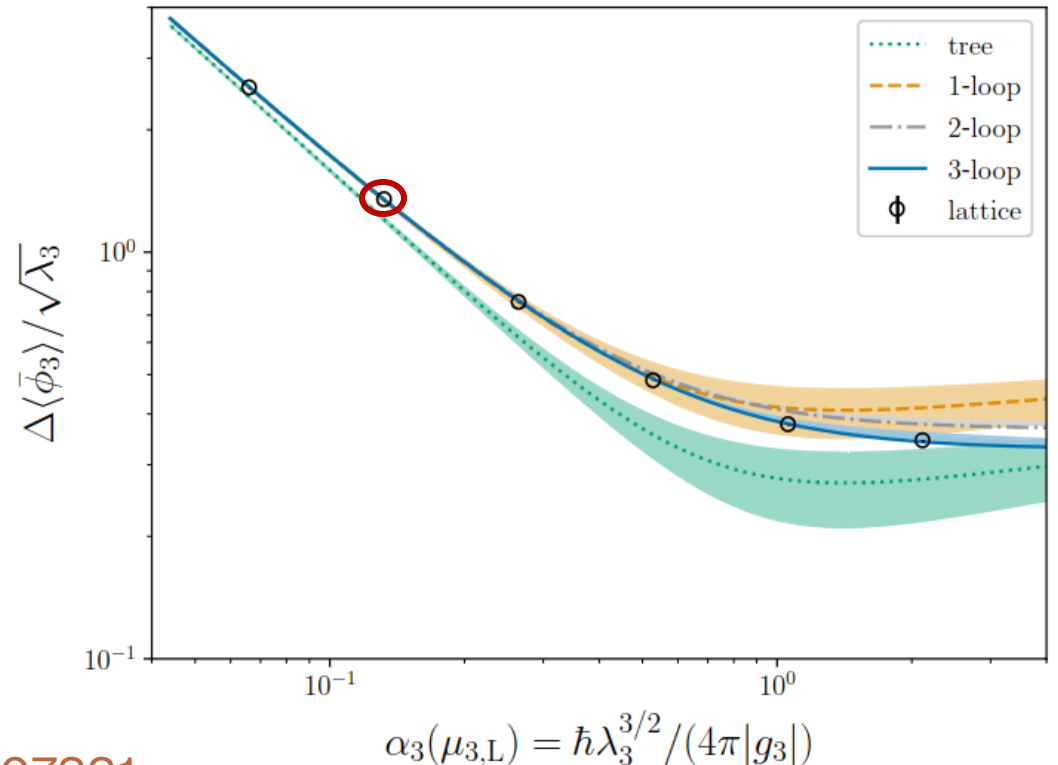


- Tree-level = functional det and dynamic prefactor approx as T^4
- LPA = Local potential approx
- Lattice = one simulated param point, reweighted to other temps

Key points – do we fully understand bubble nucleation?

Gould, [arXiv:2101.05528](https://arxiv.org/abs/2101.05528)

- Tree-level is 100% away, one-loop is 20% away in $\log \Gamma$, system should be well described by one-loop (in linear space 10¹⁷% and 10⁶%)
 - Do we understand the discrepancy between lattice and perturbative results?
- Latent heat on the lattice vs. one-loop agree to 1% (in linear space)
 - Semiclassical expansion breaks down? [arXiv:2201.07331](https://arxiv.org/abs/2201.07331)
 - Other saddle points? [arXiv:1604.06090](https://arxiv.org/abs/1604.06090), [arXiv:1806.06069](https://arxiv.org/abs/1806.06069)
 - Something else?



Conclusions

- Allows us to calibrate the uncertainty in PT parameters when obtained from perturbative results
- Accurate computations of the nucleation rate are crucial for calculating e.g. the GW power spectrum
- Our simulations show us a suppression of the nucleation rate by a factor of 20 compared to the one loop estimate
- Method and results can be applied to other theories

One-bubble takeaway

Nucleation rate
calculations are inaccurate
in perturbation theory,
lattice is significantly
better

BACKUP – DIMENSIONAL REDUCTION

- Fourier expansion

$$\int_{\tau} \int_x \left[\frac{1}{2} \psi(\tau, x) (-\nabla^2 - \partial_{\tau}^2 + m^2) \psi(\tau, x) \right] = \frac{1}{T} \sum_n \int_x \left[\frac{2}{2} \phi_n(x) (-\nabla^2 + (n\pi T)^2 + m^2) \phi_n(x) \right]$$

- Masses of the Fourier (Matsubara) modes are now

$$m_n^2 = (n\pi T)^2 + m^2$$

BACKUP – DIMENSIONAL REDUCTION

Hard:

$$E \sim \pi T$$



Soft:

$$E \sim gT$$



Supersoft:

$$E \sim g^{3/2}T$$



Ultrasoft:

$$E \sim g^2T$$

Backup - Reweighting

- Simulations are computationally expensive → use reweighting the order parameter histogram at different parameter points
- In our case we reweight in two parameters

