

This will be a board talk, so the slides are not comprehensive





Intro to Machine Learning in Astrophysics & Cosmology

(with inescapable personal bias)

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Making the impossible, possible



Astrophysics and Cosmology: questions

Astrophysics and Cosmology: questions

AGN feedback

 $\Lambda?$

 H_0 ?

Dust modelling

Galaxy-halo connection



Astrophysics: data



DECam















Data x, Model M

Model parameters $p(\theta | x, M)$ Data x, Model M

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Data x, Model M $p(\theta | x, M)$ $p(M_1 | x)$ vs $p(M_0 | x)$



Some examples....

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Galaxy cluster mass inference - Ho et al. 2402.05137



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Mass of the local group (Lemos, NJ, et al. 2010.08537)





CAMELS: Villaescusa-Navarro et al. 2109.10915

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Using Moment Networks (NJ & Wandelt 2011.05991)

1.2

Truth

1.4

1.6

1.8

2.0

1.0

CAMELS: Villaescusa-Navarro et al. 2109.10915





Model 1: galaxies are intrinsically aligned



Simple "Evidence Network" result: $\log_{10} K = -0.8 \ (\pm 0.3)$

Galaxy alignment model comparison (NJ & Wandelt 2403.0231)

Generative AI



GPT4 - "Can you draw me a scientist looking at a galaxy from London?"

Generative AI: Diffusion models

Initial conditions (density field):

Late Universe, i.e. now (density field):

Legin et al. 2304.03788





Generative AI: Diffusion models for inverse problems



Remy et al. 2201.05561



Mutual information between latents and halo assembly history



Cosmological simulation

Lucie-Smith et al. 2305.03077

Interpretation...



"Rediscovering" Newton's gravity equation from data with graph networks: Lemos, NJ, Cranmer, Ho, Battaglia 2202.02306





Back-up slides...

 $p(x \mid \theta)$?



Normalising Flow

Simulated data

 $p(x \mid \theta)$?

Image credit: Eric Jang



Normalising Flow

Simulated data

How do I know this is right?



How do I know this is right?



How do I know this is right?

UC





The challenge of model comparison

$p(M_1 | x)$ vs $p(M_0 | x)$



Bayes factor: $K = \frac{p(x_O|M_1)}{p(x_O|M_0)}$

What do people usually do?

Marginal likelihood:

$p(x_O|M_1) = \int p(x_O|\theta, M_1) \ p(\theta|M_1) \ \mathrm{d}\theta$



Evidence Networks

1. Generate/collect data for each model: $x_i \sim p(x | M_1)$

2. Bespoke loss function: $\mathcal{V}(f(x), m)$

3. Train networks to estimate Bayes factor: $f^*(x_0) = \log K$



What does training a neural network do?

What does training a neural network do?

$$I[f] = \sum_{m \in \{0,1\}} \int \mathcal{V}(f(x),m) \ p(x,m) \ dx$$
neural network

What does training a neural network do?

$$I[f] = \sum_{\substack{m \in \{0,1\}}} \int \begin{array}{c} \mathcal{V}(f(x),m) \ p(x,m) \ dx \\ \uparrow & \uparrow \\ \text{model label} \end{array} \right|$$

 $I[f] = \sum_{m \in \{0,1\}} \int \mathcal{V}(f(x),m) \ p(x,m) \ dx$

 $\mathcal{V}(f(x),m) = e^{\left(\frac{1}{2} - m\right)f(x)}$

Exponential loss

 $\mathcal{V}(f(x),m) = e^{\left(\frac{1}{2} - m\right)f(x)}$



How does this work?



 $\mathcal{V}(f(x),m) = e^{\left(\frac{1}{2} - m\right)f(x)}$ $f^*(x_O) = \log K$

How does this work?



How does this work?





leaky parity odd power (l-POP) transform



$$\tilde{\mathcal{J}}_{\alpha}(x) = \operatorname{sgn}(x) |x|^{\alpha}$$

 $\tilde{\mathcal{J}}_{\alpha}(x) \coloneqq x |x|^{\alpha - 1} \quad \forall x \in \mathbb{R} \text{ where } \alpha \in \mathbb{R} \text{ and } \alpha \geq 1$



Demonstration: does this work with data?

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Model 1: Linear growth term Model 2: No linear growth term





How do we know it worked?



Model posterior blind coverage test



Considering alternative methods...



"Likelihood-free"?

$$p(x_O|M_1) = \int p(x_O|\theta, M_1) \ p(\theta|M_1) \ \mathrm{d}\theta$$

Considering alternative methods...



"Likelihood-free"?

 $p(x_O|M_1)$

Considering alternative methods...

UCL



Neural density estimation: $p(M_1|x) \& p(M_0|x)$



PolyChord (assumed likelihood)



