

**Production of light nuclei in Au+Au collisions at
 $\sqrt{s_{NN}} = 7.7$ and 27 GeV at RHIC-STAR**

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ALICE-STAR India Collaboration Meeting

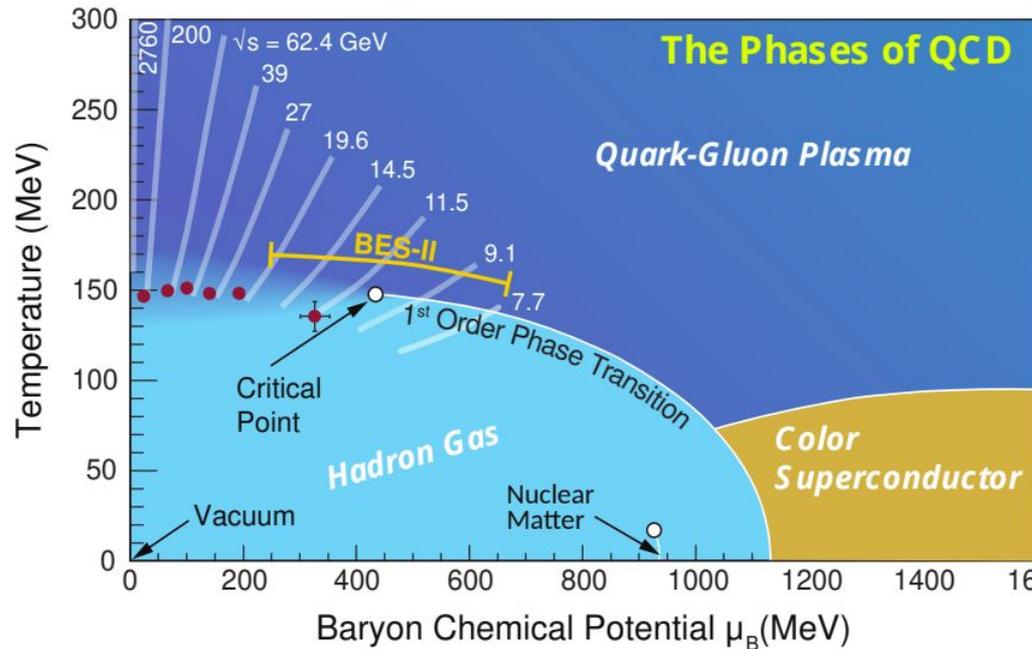
24th - 27th June, 2024

Institute of Physics, Bhubaneswar

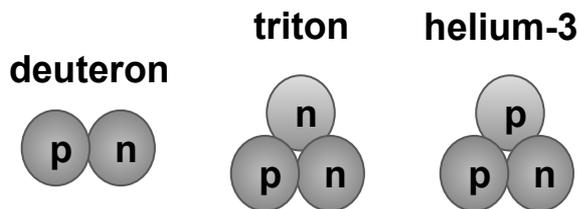
Previously presented:

https://indico.cern.ch/event/1329196/contributions/5665364/attachments/2756905/4803916/Updated_sibaram_ALICE_STAR_India_Colla_Nov_2023.pdf

- ❖ Introduction
- ❖ The STAR experiment
- ❖ Analysis details
 - Data set, event, and track selection
 - Particle identification
 - Signal extraction
- ❖ Corrections
- ❖ Results
- ❖ Summary
- ❖ Outlook



Nucl. Phys. A **1017** 122343 (2022)



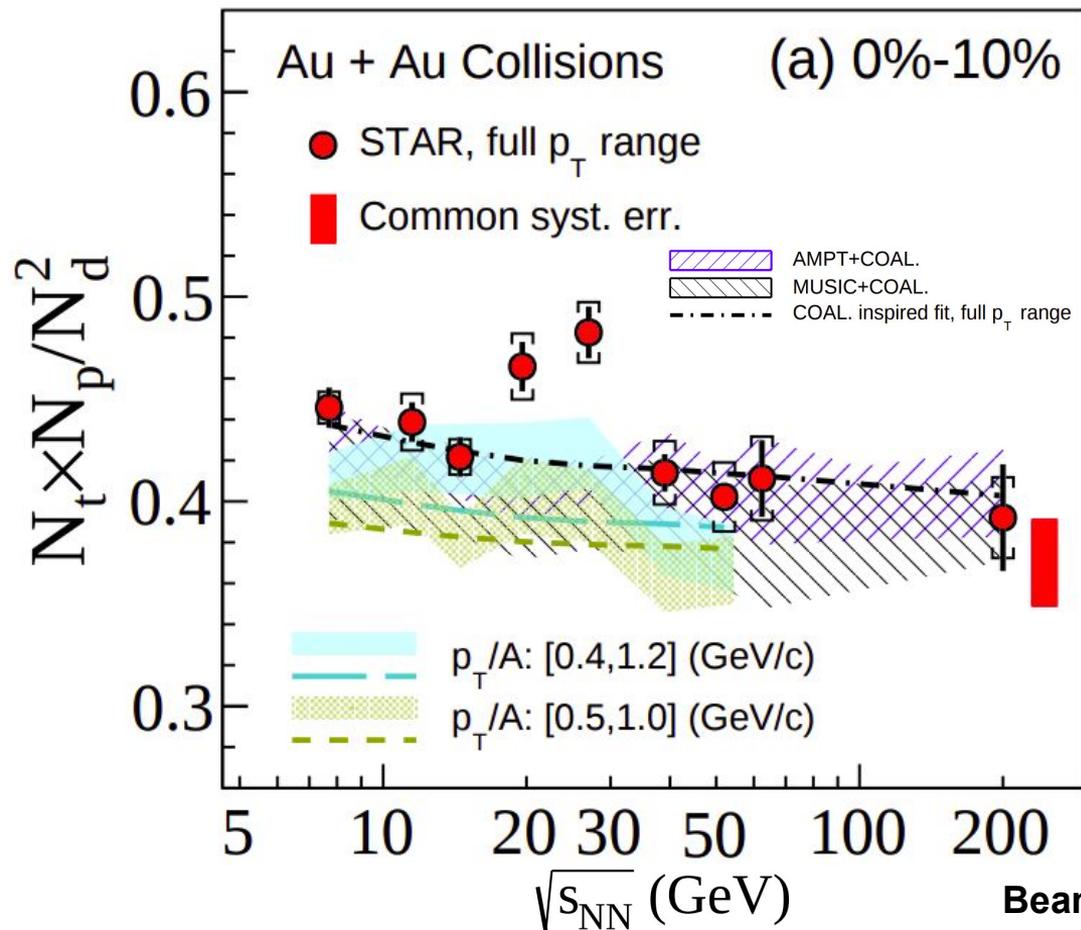
- ❑ Goal of heavy-ion collision experiment is to explore **QCD phase diagram**
- ❑ **QGP to Hadronic matter** -> continuous at low μ_B and 1st order at higher μ_B (by QCD)
- ❑ Search for **Critical Point (CP)** in phase diagram

❑ Choice of observable : **Compound light**

$$\text{nuclei ratio } \frac{N_t \times N_p}{N_d^2} \propto (1 + \Delta n)$$

Phys. Lett. B **774** 103-107 (2017)
Eur. Phys. J. A **57** 11313 (2021)

Beam Energy Scan program (BES-I) collider mode :
7.7, 9.1, 11.5, 14.5, 19.6, 27, 39, 62.4, 200 GeV

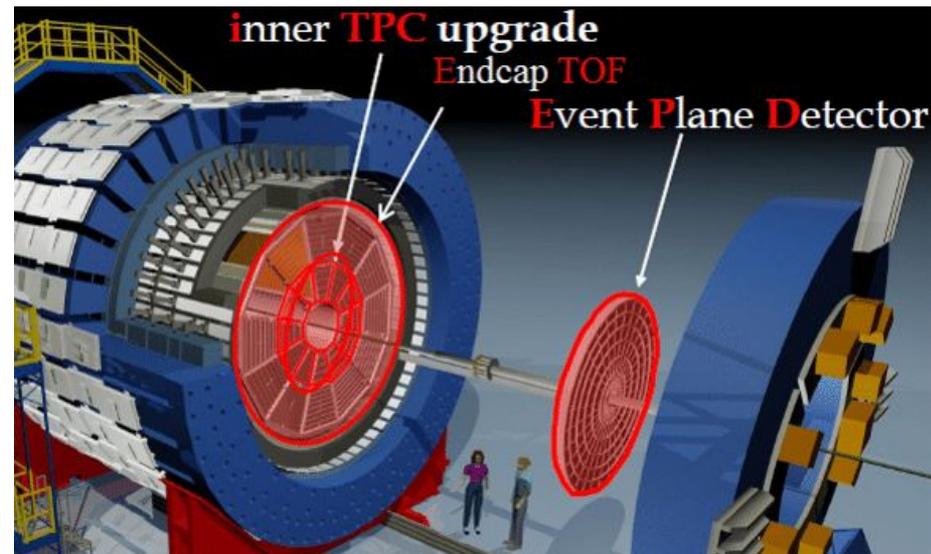


Deviation from baseline is 4.1 sigma

- Light nuclei yields and ratios could provide a probe to search for signature of critical phenomena

Beam Energy Scan program (BES-II) collider mode
: 7.7, 9.2, 11.5, 14.6, 17.3, 19.6 GeV

- ❑ **Solenoidal Tracker at RHIC (STAR)** is one of the large detector systems at RHIC consisting of several sub-detectors
- ❑ **Time Projection Chamber (TPC)** and **Time of Flight (TOF)** are two sub-detectors used for particle identification at STAR
- ❑ The upgrade to **iTPC** provides better momentum resolution, better dE/dx resolution, and improved acceptance at high rapidity ($|\eta| < 1.5$)



J. Phys. Conf. Ser. **742** 012022 (2016)

$$\text{Pseudo-rapidity : } \eta = -\ln \tan \frac{\theta}{2}$$

Au+Au, $\sqrt{s_{NN}} = 7.7$ GeV

Trigger setup:

production_7p7GeV_2021

Data file: picoDst

Stream: st_physics

Trigger selection: 810010,

810020, 810030, 810040

Production tag: P22ib

Number of Events ~ 52.5 M

Au+Au, $\sqrt{s_{NN}} = 27$ GeV

Trigger setup:

27GeV_production_2018

Data file: picoDst

Stream: st_physics

Trigger selection:

610001, 610011, 610021, 610031,

610041

Production tag: P19ib

Number of Events ~ 267 M

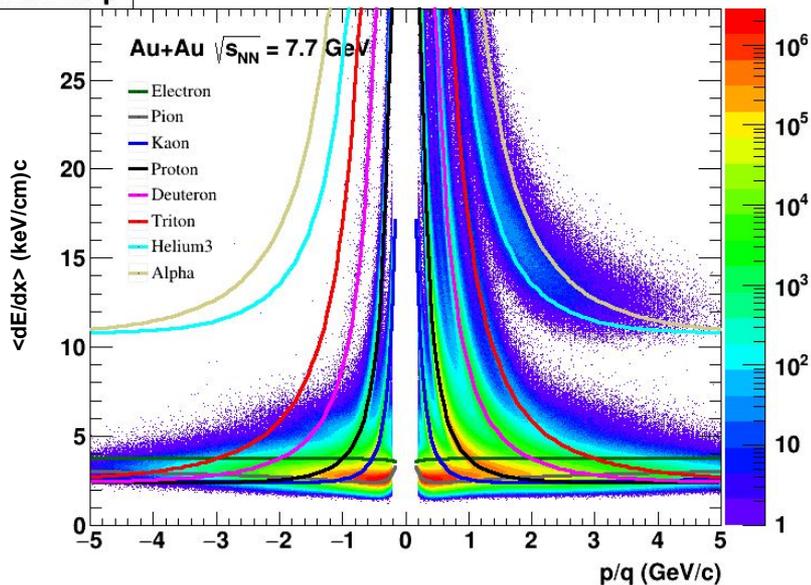
Event Selection:

Cuts (27 GeV)	Value
$ v_z $	≤ 70 (40) cm
v_r	≤ 2 cm

Track Selection:

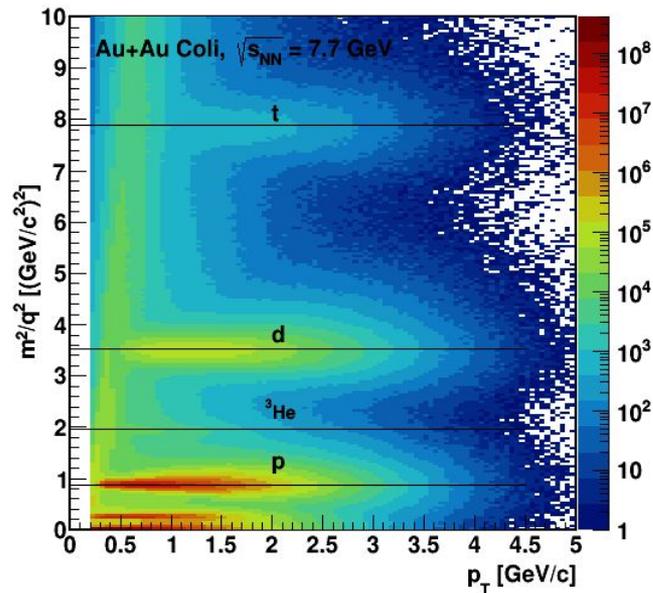
Cuts (27 GeV)	Value
nHitsFit	≥ 25 (20)
nHitsdEdx	≥ 15 (10)
nHitsFit/nHitsPoss	≥ 0.52
DCA	< 1 cm
p_T (GeV/c)	≥ 0.2
$ y $ (d)(t)(³ He)	$< (0.3) (0.5) (0.5)$

dEdx vs p



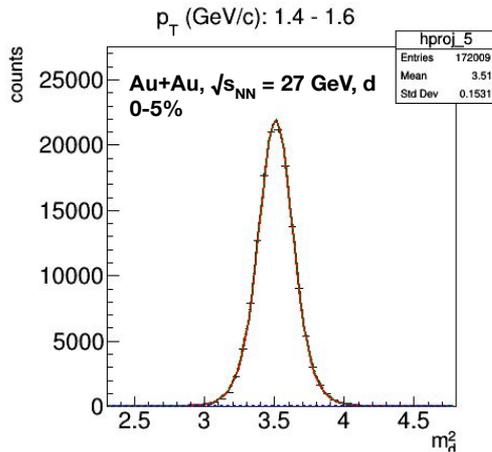
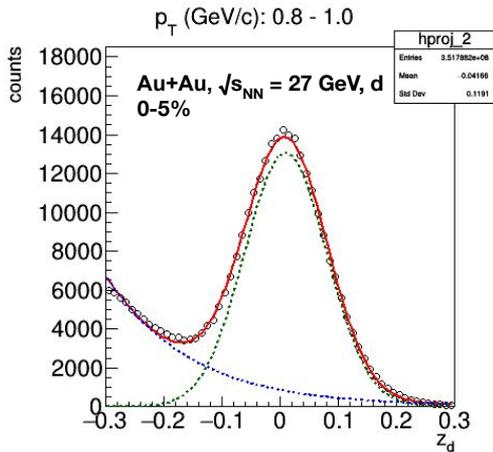
- PID using dE/dx information from TPC at low p_T

$$Z = \ln\left(\frac{\langle dE/dx \rangle}{\langle dE/dx \rangle_{theory}}\right)$$

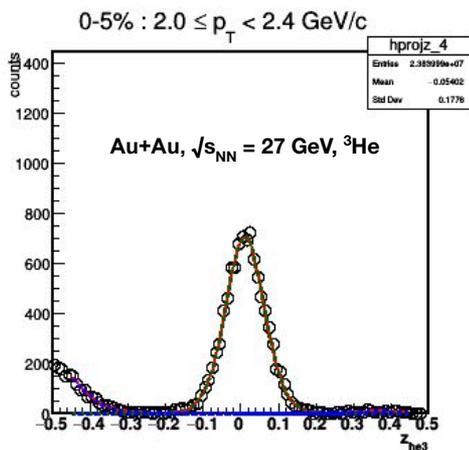


- PID using m^2/q^2 information from TOF at intermediate p_T

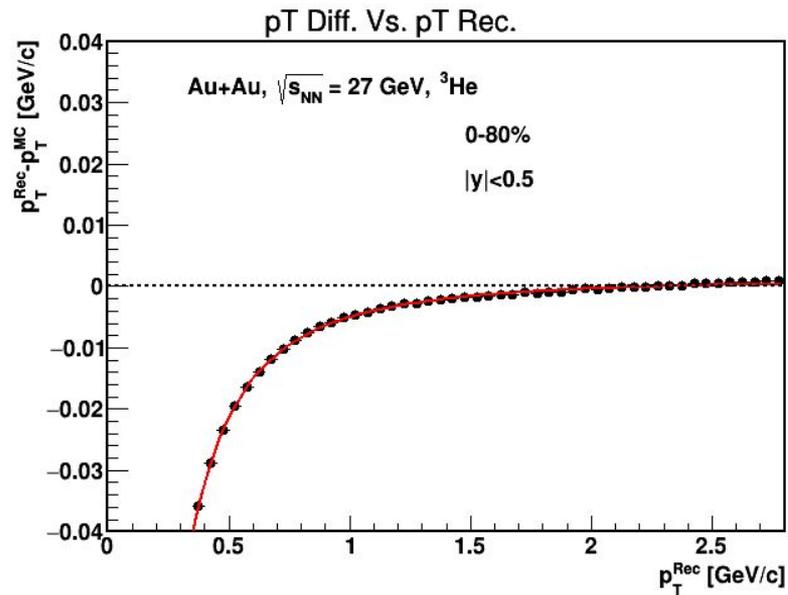
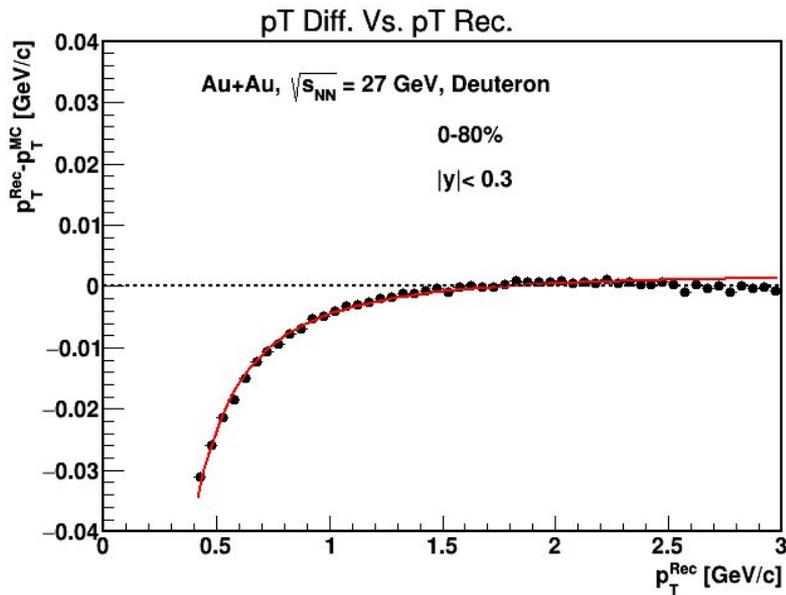
$$\frac{m^2}{q^2} = p^2(1/\beta^2 - 1)$$



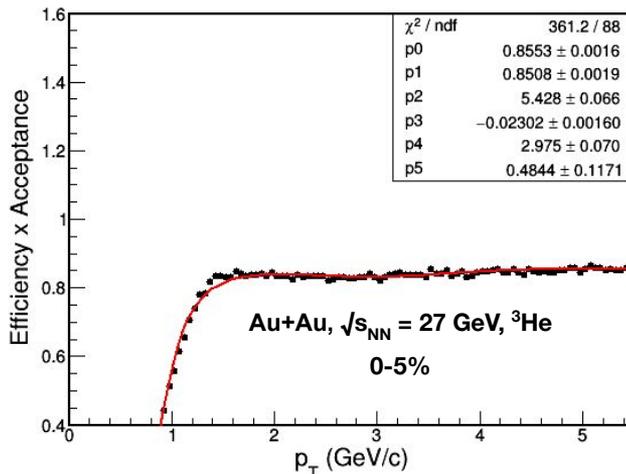
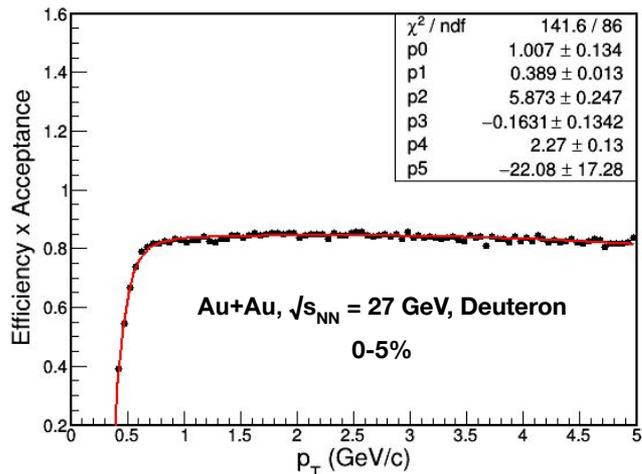
- Total function
- - - Background
- . . . Signal



- At low p_T : Gaussian + exponential
- At intermediate p_T : student_t + exponential/gaussian



$$\text{Fit function: } f(p_T) = a_1 + a_2 \left(1 + \frac{a_3}{p_T^2}\right)^{a_4}$$



Fit function:

$$f(p_T) = a_1 \exp\left(-\left(\frac{a_2}{p_T}\right)^{a_3}\right) + a_4 \exp\left(-\frac{(p_T - a_5)^2}{2a_6}\right)$$

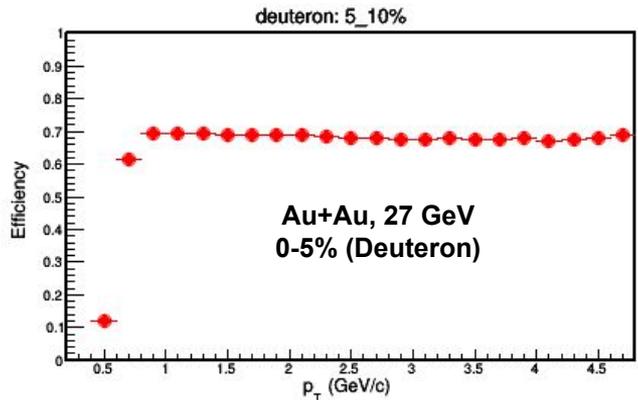
Absorption correction factor:

$$f(p_T) = p_0 + p_1 p_T + p_2 p_T^2 + p_3 p_T^3 + p_4 p_T^4 + p_5 p_T^5 + p_6 p_T^6 + p_7 p_T^7 +$$

Particle	p_0	p_1	p_2	p_3	p_4	p_5	$p_6 p_T^8$	p_7	p_8
Deuteron	1.17269	-0.342489	0.287993	-0.126426	0.0325183	-0.00508311	0.000475232	-2.44273e-05	5.30683e-07
^3He	1.24114	-0.325563	0.190857	-0.0591357	0.0108598	-0.00122826	8.46123e-05	-3.29416e-06	5.65162e-08

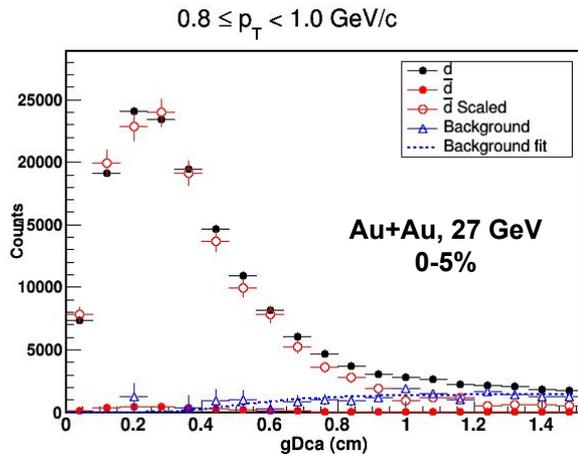
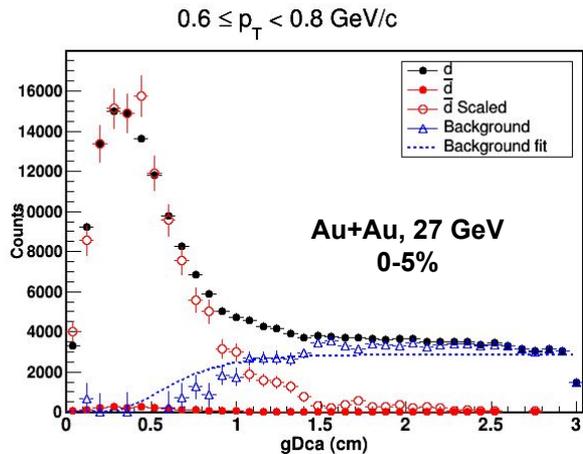
Reference:

[Link1](#),
[Link2](#),
[Link3](#)



→ TOF matching efficiency is defined as

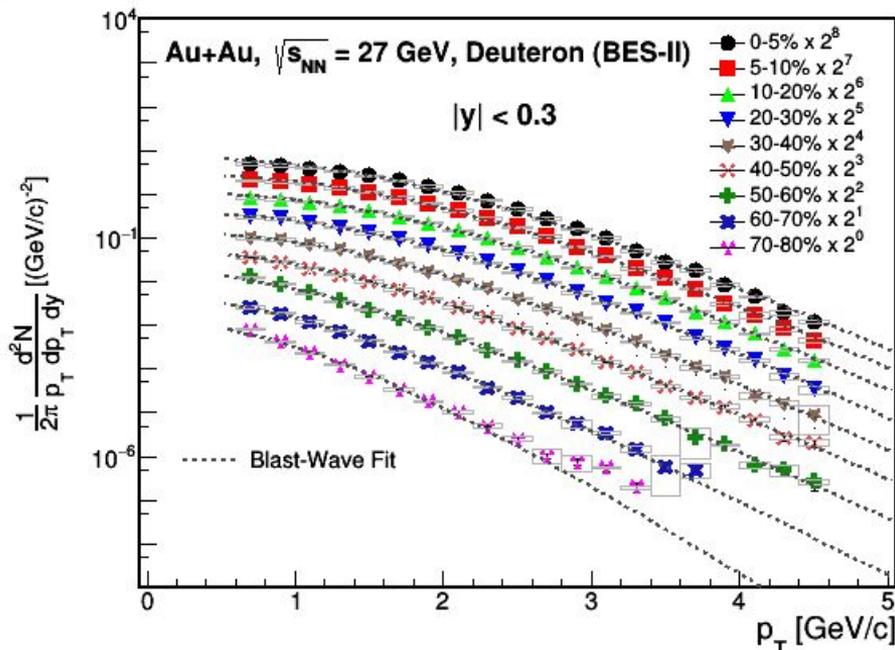
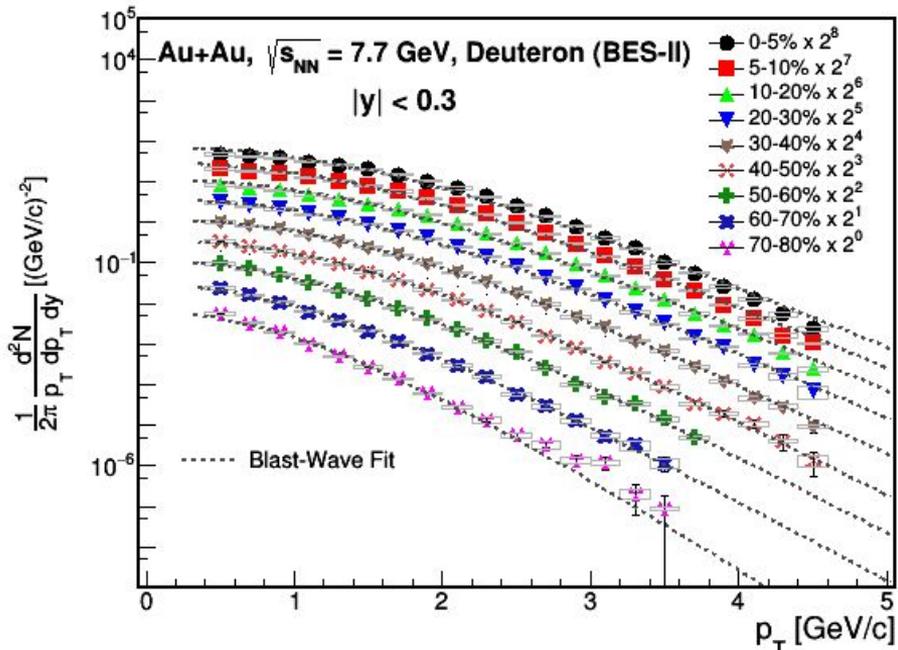
$$\epsilon_{TOF} = \frac{N_{TOF}}{N_{TPC}}$$



Fit function:

$$f(x) = a_1 \bar{d}(x) + d_{bkg}(x)$$

$$d_{bkg}(x) = a_2 \left(1 - \exp\left(\frac{x}{a_3}\right)\right)^{a_4}$$



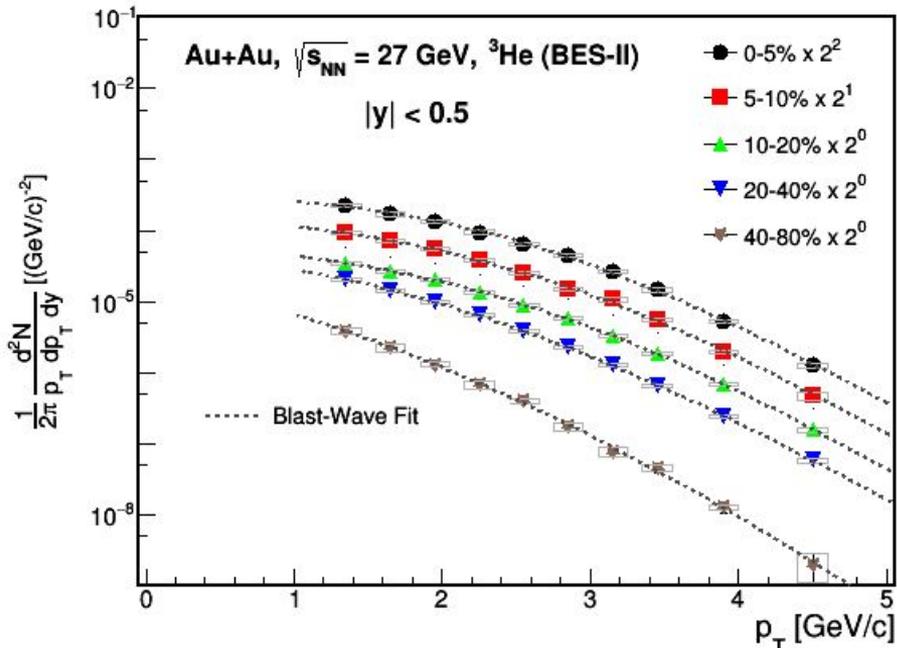
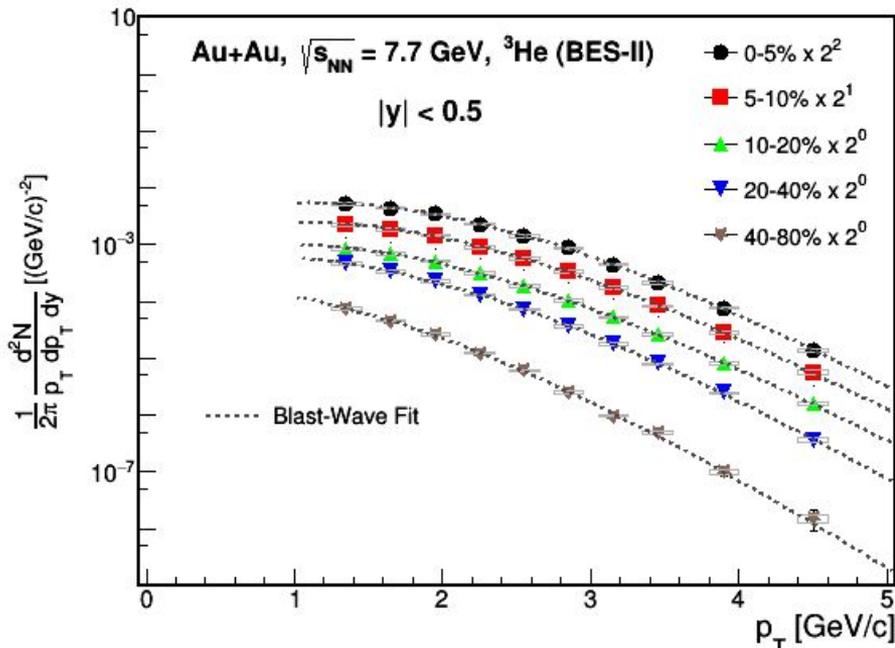
$$\frac{1}{2\pi} \frac{d^2N}{p_T dp_T dy} \propto \int_0^R r dr m_T I_0\left(\frac{p_T \sinh \rho}{T}\right) K_1\left(\frac{m_T \cosh \rho}{T}\right)$$

$$\rho = \tanh^{-1} \beta_r, \beta_r(r) = \beta_s \left(\frac{r}{R}\right)^n$$

Phys. Rev. C **99**, 064905 (2019)

□ A clear p_T and centrality dependence of light nuclei spectra is observed

- Energy loss correction (✓)
- TPC efficiency and acceptance (✓)
- TOF efficiency (✓)
- Absorption correction (✓)
- Background correction (✓)

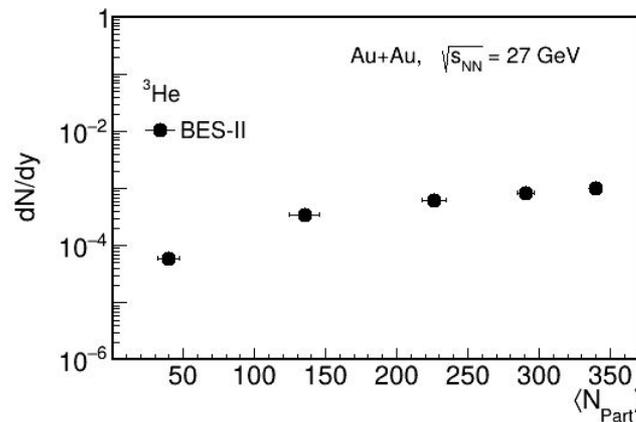
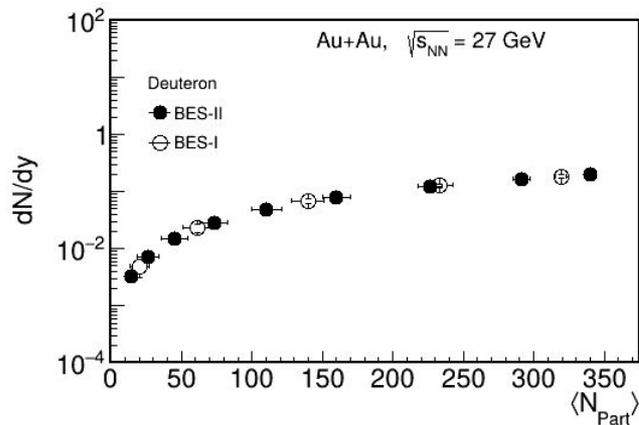
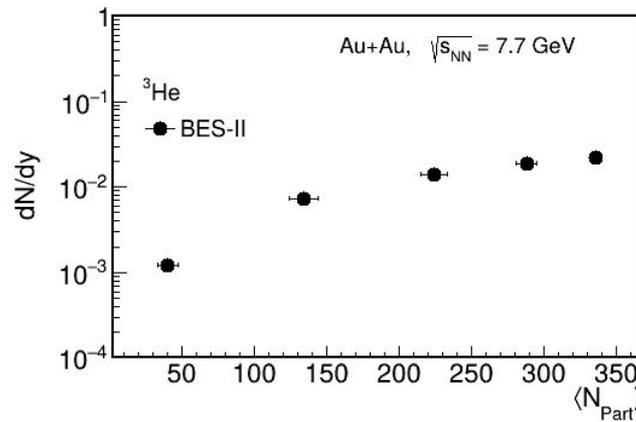
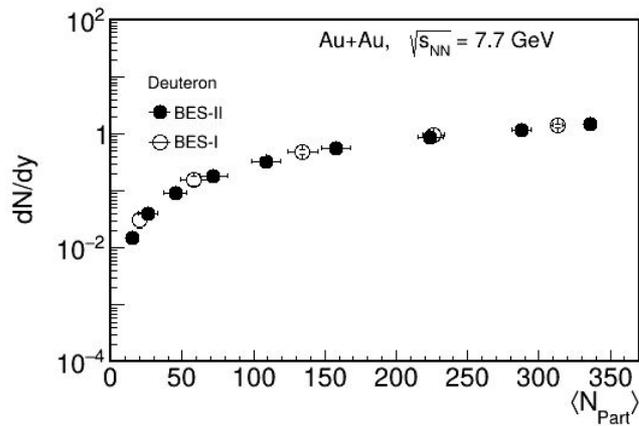


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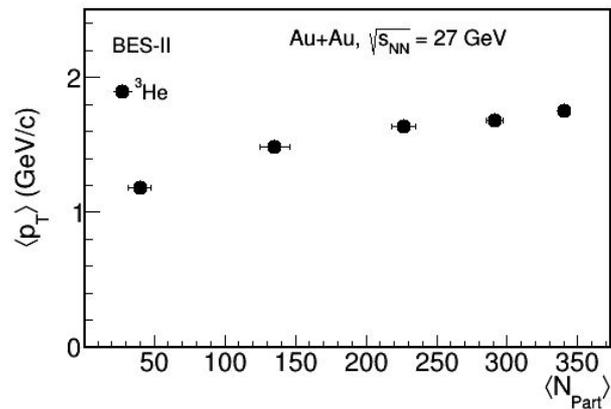
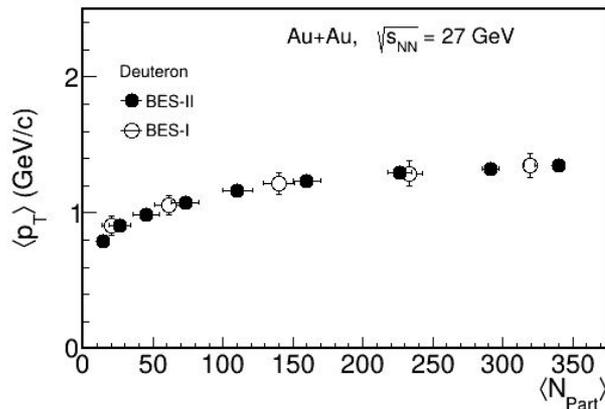
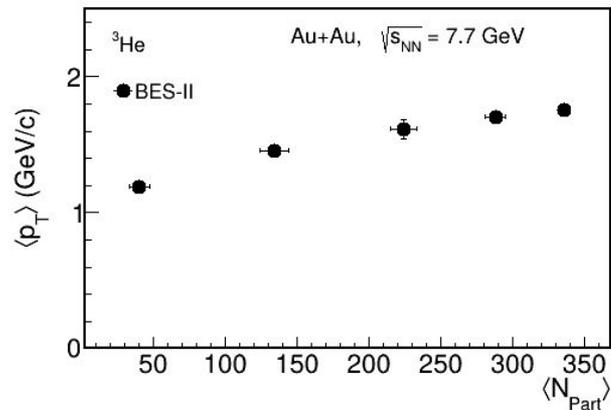
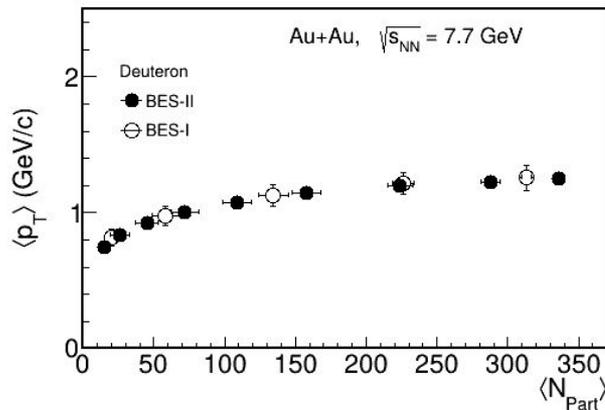
$$\rho = \tanh^{-1} \beta_r, \beta_r(r) = \beta_s \left(\frac{r}{R}\right)^n$$

☐ A clear p_T and centrality dependence of light nuclei spectra is observed

- Energy loss correction (✓)
- TPC efficiency and acceptance (✓)
- Absorption correction (✓)



- Yield Increases with increasing centrality \rightarrow effect of baryon stopping is stronger in central collisions



- $\langle p_T \rangle$ increases with increasing centrality \rightarrow radial flow increases with centrality

- ❑ Corrected p_T spectra of deuteron is shown in Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ and 27 GeV
- ❑ 1st measurement of ${}^3\text{He}$ p_T spectra is presented in Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ and 27 GeV
- ❑ dN/dy and $\langle p_T \rangle$ is calculated for light nuclei in Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ and 27 GeV.

Outlook

- ❑ p_T spectra for proton in BES-II energies
- ❑ Estimate the compound light nuclei ratios
- ❑ Complete the analysis for other BES-II energies

Thank You All !

Backup

→ At low p_T : $\frac{1}{\sqrt{2\pi\sigma^2}} e^{(x-\mu)^2/2\sigma^2} + ae^{bx}$

→ At intermediate p_T : $\frac{\Gamma(\frac{\nu+1}{2})}{\Gamma(\frac{\nu}{2}\sqrt{\pi\nu\sigma^2})} \left[1 + \frac{1}{\nu} \left(\frac{x-\mu}{\sigma}\right)^2\right]^{-\frac{\nu+1}{2}} + ae^{bx}$