RGE effects in global EFT fits Eleni Vryonidou University of Manchester









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LHC EFT WG meeting, 2/12/24













Processes and observables







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Processes and observables

$$c_i^6(\mu)a_{n,i}^6(\mu) + \mathcal{O}\left(\frac{1}{\Lambda^4}\right) \longrightarrow \text{Constraints on WC}$$













Huge effort to improve each one of these steps!

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For the best results: Do this in a global fit!

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Ingredients of global fits **SMEFiT** as a global fit example

Flavour assumption:

 $U(2)_q \times U(3)_d \times U(2)_u \times (U(1)_\ell \times U(1)_e)^3$ + Yukawa of bottom, charm and tau

50 degrees of freedom: 2F, 2L2H, 4H, Bosonic SMEFiT3.0 Celada, Giani, Mantani, Rojo, Rossia, Thomas, EV, ter Hoeve arXiv:2404.12809

See also: FitMaker arXiv:2012.02779, HEPfit arXiv:1910.14012

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Experimental data

Category	Processes	^{at} SMEF
Top quark production	$t\bar{t} + X$, 11
	$tar{t}Z,tar{t}W$	21
	$tar{t}\gamma$	2
	single top (inclusive)	28
	tZ,tW	13
	$tar{t}tar{t}$, $tar{t}bar{b}$	12
	Total	19
	Run I signal strengths	22
Higgs production and decay	Run II signal strengths	36 (
	Run II, differential distributions & STXS	71
	Total	12
Diboson production	LEP-2	40
	LHC	41
	Total	81
EWPOs	LEP-2	44
Baseline dataset	Total	44

- Bounds varying between operators •
- Most Wilson coefficient bounds below 1 for Λ =1 TeV •
- Quadratic terms important especially for 4F operators
- Least constrained coefficients are 4-top operators •

SMEFiT3.0 Celada, Giani, Mantani, Rojo, Rossia, Thomas, EV, ter Hoeve arXiv:2404.12809

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Future of global fits

How can we improve fits?

More observables:

- Particle level observables
- New final states
- Better description: EFT in backgrounds

Better EFT predictions

More/less/different operators: • Different flavour assumptions • UV inspired scenarios

EFT is a QFT, renormalisable order-by order in $1/\Lambda^2$

SMEFT computations at dimension-6

 $\Delta Obs_n = Obs_n^{\mathsf{EXP}} - Obs_n^{\mathsf{SM}} = \sum_{i} \frac{c_i^6(\mu)}{\Lambda^2} \left| a_{n,i}^6(\mu) \right| + \mathcal{O}\left(\frac{1}{\Lambda^4}\right)$

NLO QCD & loop-induced: Done (SMEFT@NLO) Degrande, Durieux, Maltoni, Mimasu, EV, Zhang arXiv:2008.11743 http://feynrules.irmp.ucl.ac.be/wiki/SMEFTatNLO

NLO EW: Some examples available, progress towards automating these as well (see next talk)

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(see next talk)

How about this μ ? When should we worry about it?

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How about this μ ? When should we worry about it?

- Observables with different natural scales (flavour, EW, Higgs, top, 4-tops)
- Differential distributions: e.g. in a typical top fit bins of reaching 2 TeV
- Eventually we want to match to the UV

- NLO EW: Some examples available, progress towards automating these as well

What is next for global fits?

Need for RGE running and mixing:

Different scales and Dynamical scales Anomalous dimension matrix known at 1-loop:

Jenkins et al arXiv:1308.2627, 1310.4838, Alonso et al 1312.2014

- 2499x2499 anomalous dimension matrix available
- Including both QCD, weak, Yukawa terms
- Implemented in various tools: e.g. Wilson: Aebischer et al <u>1712.05298</u>, RGESolver: Di Noi and Silvestrini arXiv:2210.06838

**2-loop RGE: not known in general (some pieces known, typically from flavour physics and recent computations: see e.g. arXiv:2410.07320)

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Example 1: top physics

$$\frac{dc_i(\mu)}{d\log\mu} = \gamma_{ij} \, c_j(\mu)$$

Example: Turn on 1 operator at high-scale

Compute effect on top pair cross-section

 $c_{Ou}^{1} = 1$ at 2 TeV

Aoude, Maltoni, Mattelaer, Severi, EV arXiv:2212.05067

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Implementation of RGE in Monte Carlo How to get RGE improved predictions?

Anomalous dimension included in UFO model **Mathebra Running code extracted automatically by MG5_aMC** Output Define coefficients at given scale, and run to preferred scale Options in run_card.dat

CONTROL The additional running scale (not QCD) Such running is NOT include in systematics computation True = fixed_other_scale ! False means dynamical scale 1000 = muo ref fixed ! scale to use if fixed scale mode = muo over ref ! ratio to mur if dynamical scale 1.0

Aoude, Maltoni, Mattelaer, Severi, EV arXiv:2212.05067

Fixed scale or a dynamical (as for alphas)

Dynamical scale a function of usual μ_{R}

Example 2: Higgs physics

$$\frac{dc_i(\mu)}{d\log\mu} = \gamma_{ij} c_j(\mu) \qquad \mathcal{O}_{\varphi G} = \left(\varphi^{\dagger}\varphi - \frac{v^2}{2}\right) G^a_{\mu\nu} G^{\mu\nu}_a \quad \bullet \quad \mathcal{O}_{t\varphi} = \left(\varphi^{\dagger}\varphi - \frac{v^2}{2}\right) \bar{Q}\tilde{\varphi}t + \text{h.c.}$$

$$\mathcal{O}_{tG} = ig_s(\bar{Q}\tau^{\mu\nu}T^a\tilde{\varphi}t)G^a_{\mu\nu} + h.c$$

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 $C_{t\varphi}$

QCD induced running

Impact of RGE on constraints

How does running and mixing impacts the constraints?

Top sector fit:

Aoude, Maltoni, Mattelaer, Severi, EV arXiv:2212.05067 More important for differential distributions & measurements with very different scales

RGE evolution within MC:

PS by PS point computation of coefficients: dynamical scale e.g. $H_T/2$

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Impact of RGE on constraints How does running and mixing impacts the constraints? Higgs sector fit

Maltoni, Ventura, EV arXiv:2406.06670

See also Battaglia, Grazzini, Spira, Wiesemann arXiv: 2109.02987 Di Noi, Grober arXiv:2312.11327 Di Noi, Grober, Mandal arXiv: 2408.03252

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Incl. ATLAS+CMS	ttH CMS	HH Proj
$- \mu = \sqrt{M_H^2 + P_T^2} -$	$\mu = \sqrt{M_H^2 + P_T^2}$	μ= <i>Μ</i> _F
μ= M _H	μ= <i>M</i> _H	— μ= M _F
μ= 1 TeV	μ= 1 TeV	μ= 1 T

Eventually need to be taken into account in a global fit!

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How about a real global fit? How to practically include these effects in a global fit?

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- Data points cover a large range of scales from $M_{\rm Z}$ to 2-3 TeV
- Ideally we would like a fully dynamical scale setup as often employed in SM predictions
- That requires rerunning the MC for every single theoretical prediction
- **Reasonable approximation?**
- Assign one scale to each data point

How about a real global fit? How to practically include these effects in a global fit?

Practical SMEFiT implementation:

- Associate one scale to each observable (445 datapoints)
- Use Wilson to do the running between chosen starting and _... the scale of each datapoint

$$\Gamma(\mu,\mu_0;lpha_s,lpha) = \exp\left(\int_{\mu_0}^{\mu} d\log(\mu')\gamma(lpha_s,lpha)
ight)$$
 Evo

$$T_{\rm EFT}(\boldsymbol{c}(\mu)/\Lambda^2) = T_{\rm SM} + \sum_{i=1}^{n_{\rm op}} \kappa_i \frac{c_i(\mu)}{\Lambda^2} \qquad \qquad T_{\rm EFT}(\boldsymbol{c}(\mu_0)/\Lambda^2) = T_{\rm SM}$$

 $= T_{\rm SM} +$

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Validation with arXiv:2212.05067

ter Hoeve, Mantani, Rojo, Rossia, EV in preparation

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Correlation patterns change

When can the RGE matter?

For unconstrained operators RGE can offer new sensitivity Example: 4-heavy 4-fermion coefficients Significant improvement of bounds due to impact on EWPO at FCC

Marginalised 95 % C.L. intervals

ter Hoeve, Mantani, Rojo, Rossia, EV in preparation

How about particular UV complete models?

2HDM in decoupling limit

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2-scalar EW quadruplet model

 $\mathscr{L}_{UV} \supset -\lambda_{\Phi} H^* H^* (\varepsilon H) \Phi - \lambda_{\widetilde{\Phi}} H^* H^* H^* \widetilde{\Phi} / \sqrt{3} + \text{h.c.},$

Durieux, McCullough, Salvioni arXiv:2209.00666

Impact of RGE and 1-loop matching Significant improvement of reach due to impact on EWPO at FCC

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Summary

Global fit results affected by the precision of EFT predictions

Aim to include more and more precise theory predictions in the fits

constraints

UV complete models

- Inclusion of RGE effects in predictions is necessary, and can significantly affect
- Particularly important for poorly constrained operators and in the constraints on

