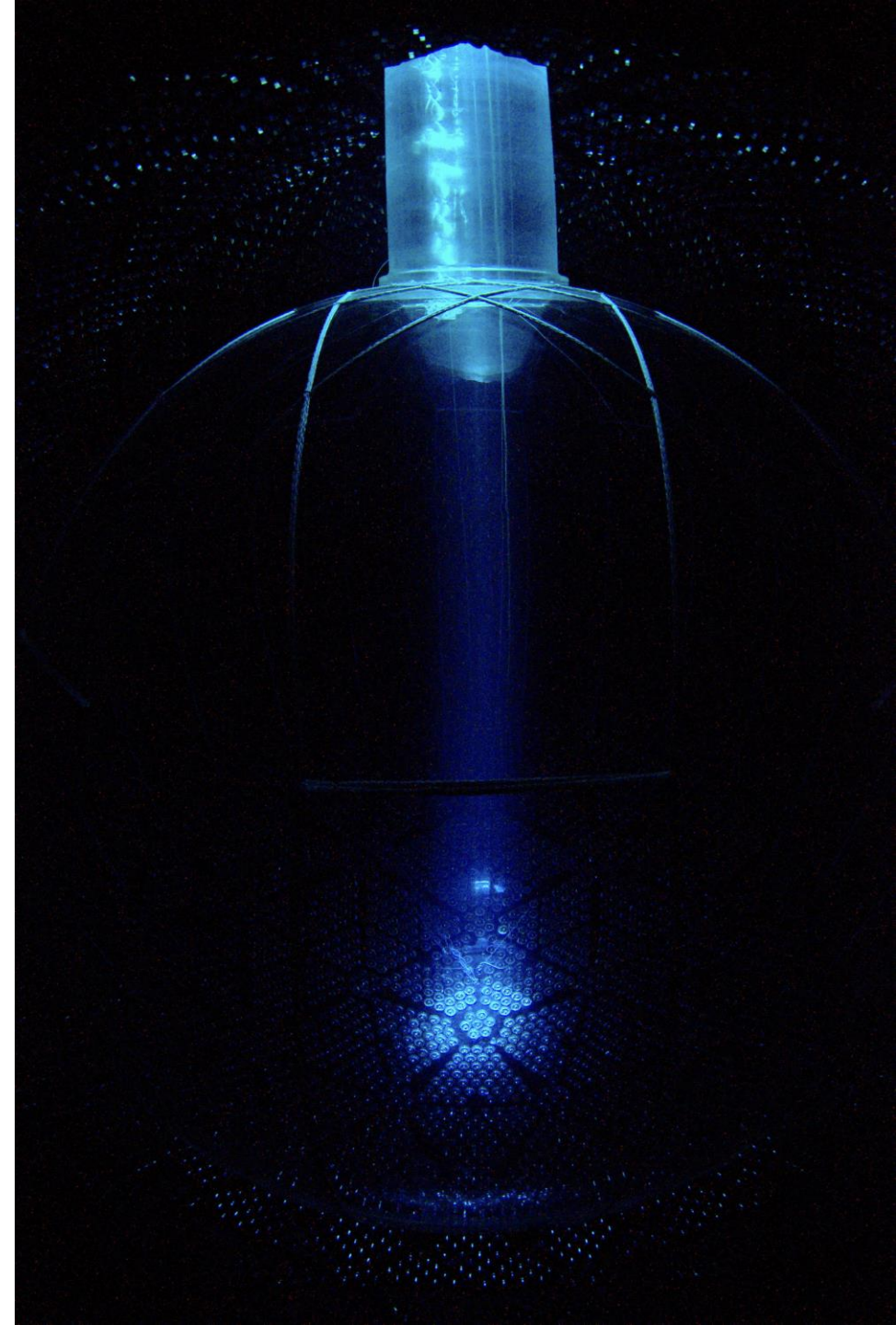




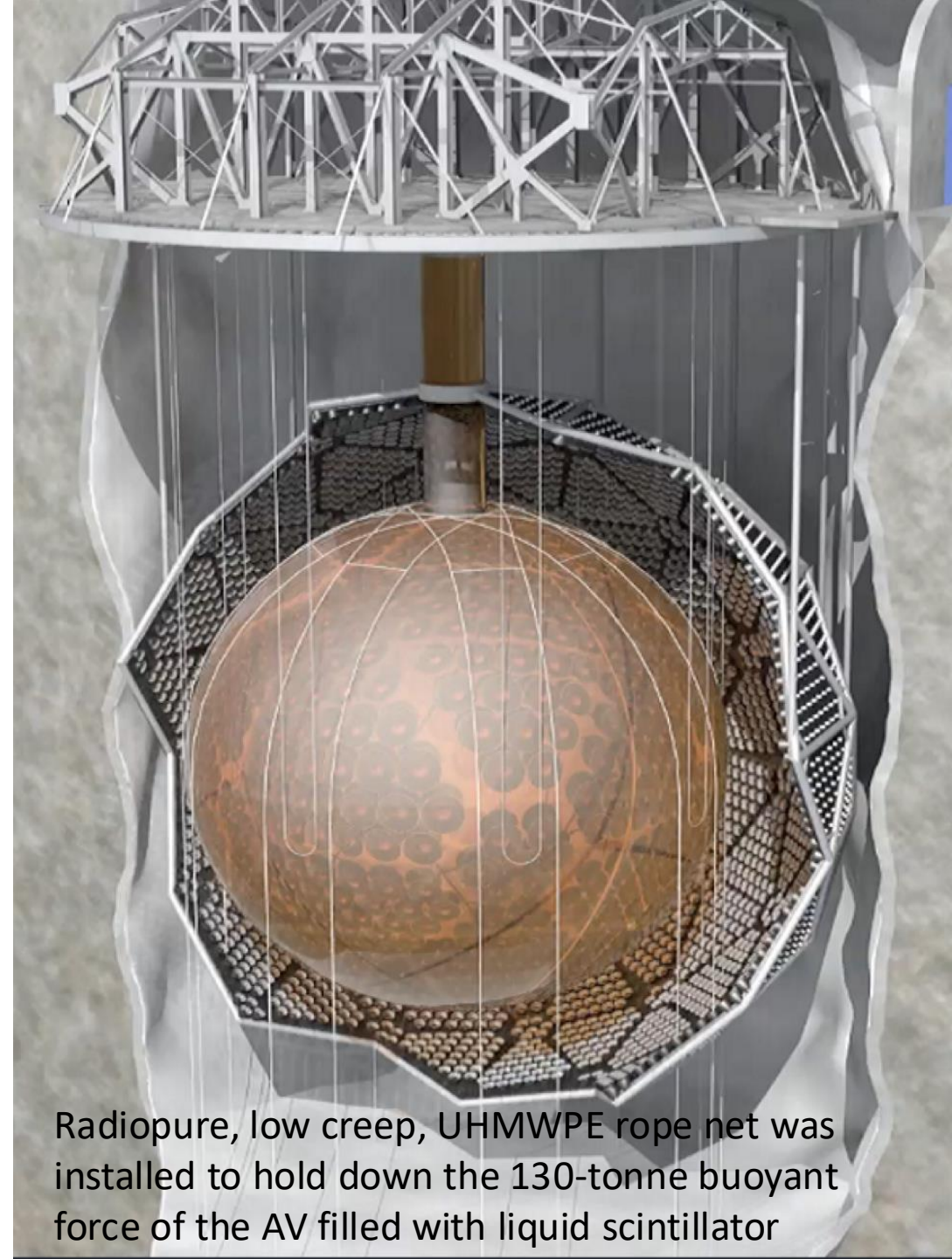
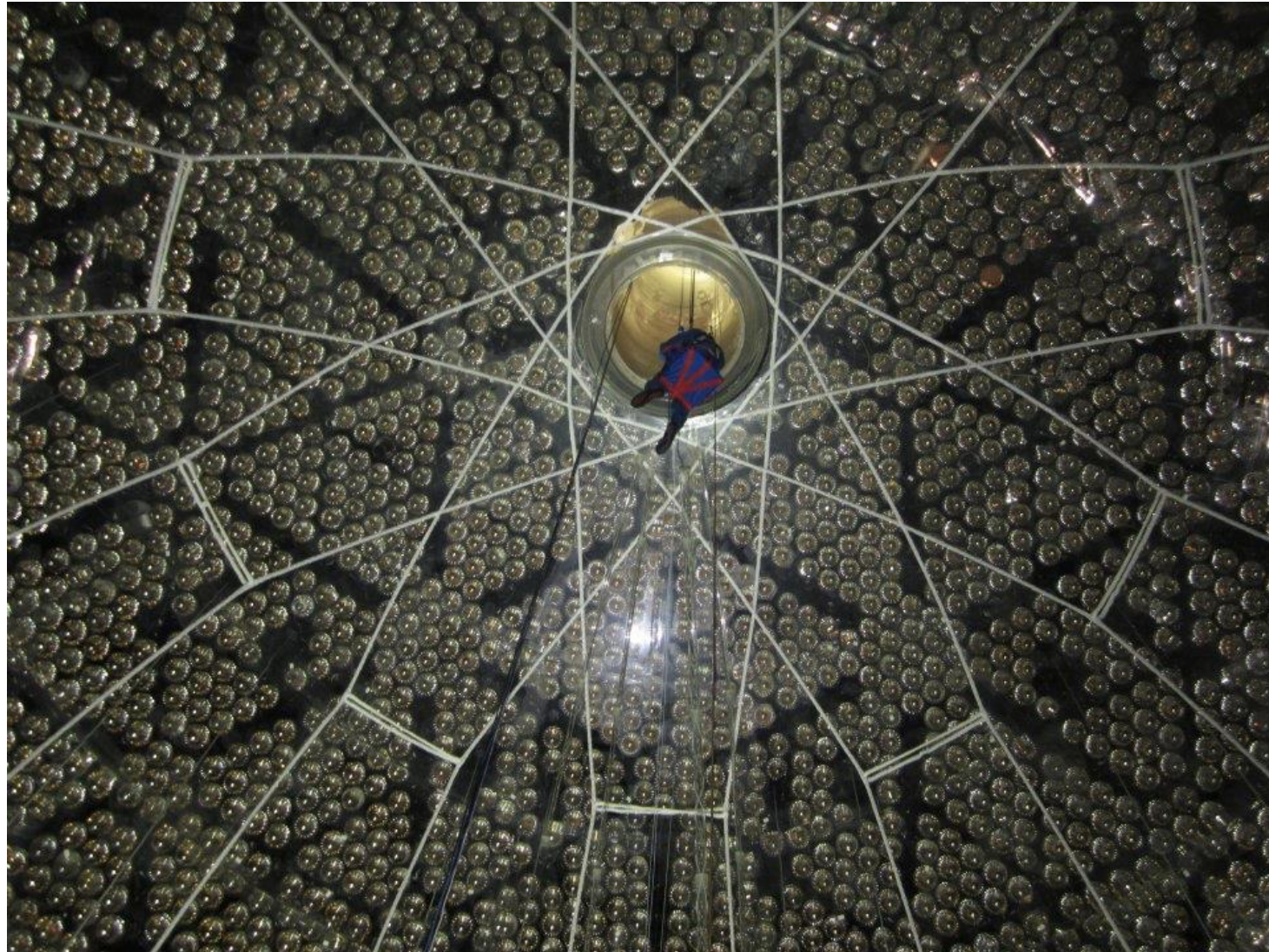
Recent Results

McDonald Institute 2024 Community Meeting

Mark Chen
Queen's University and CIFAR
August 8, 2024

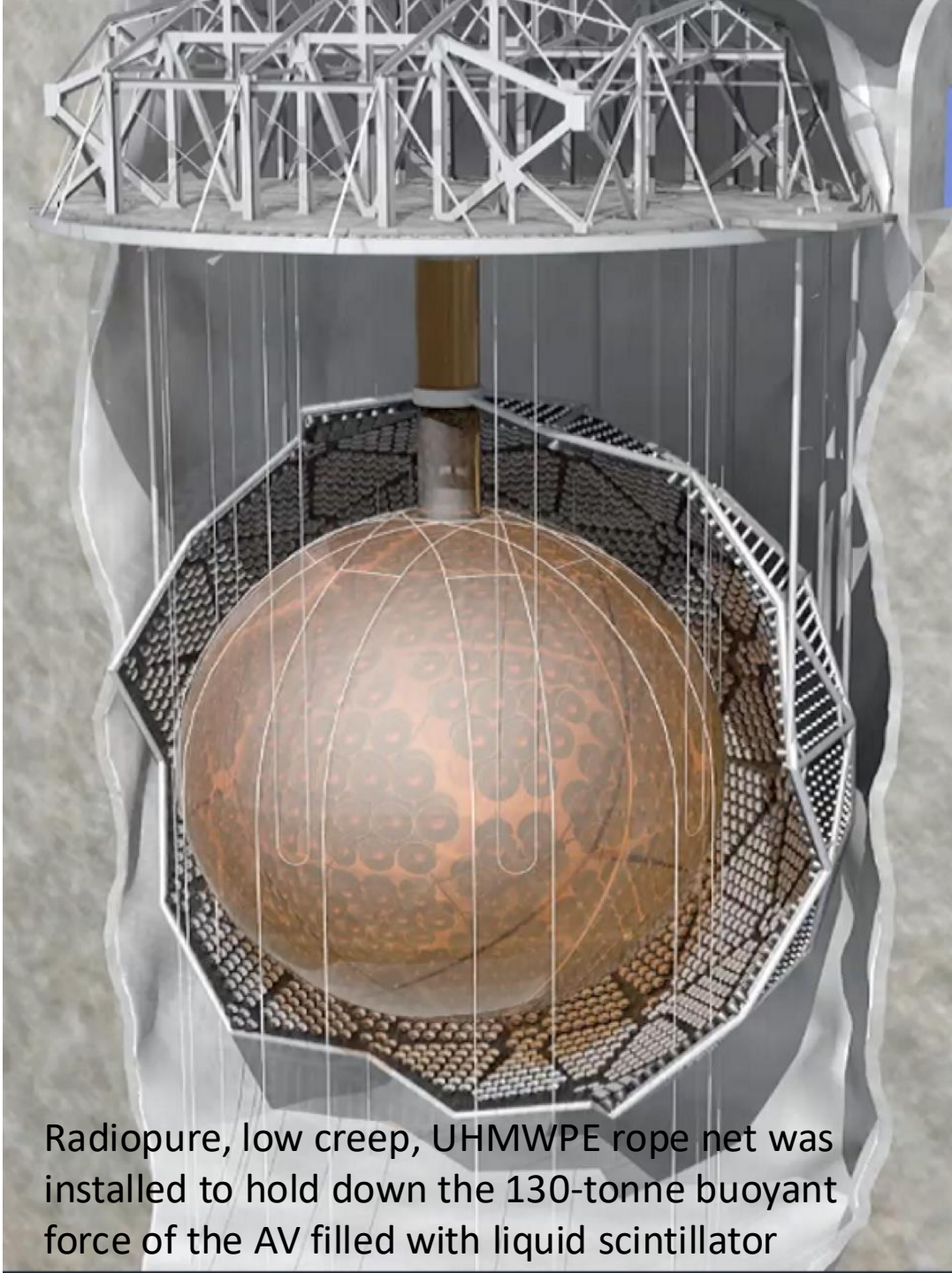
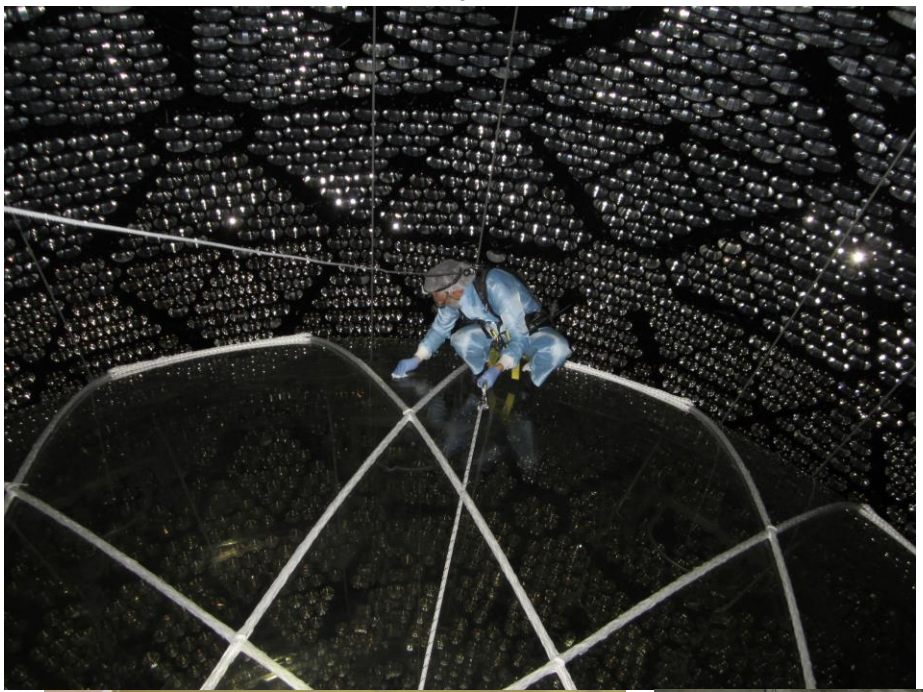


SNO+ is the Sudbury Neutrino Observatory Filled with Liquid Scintillator



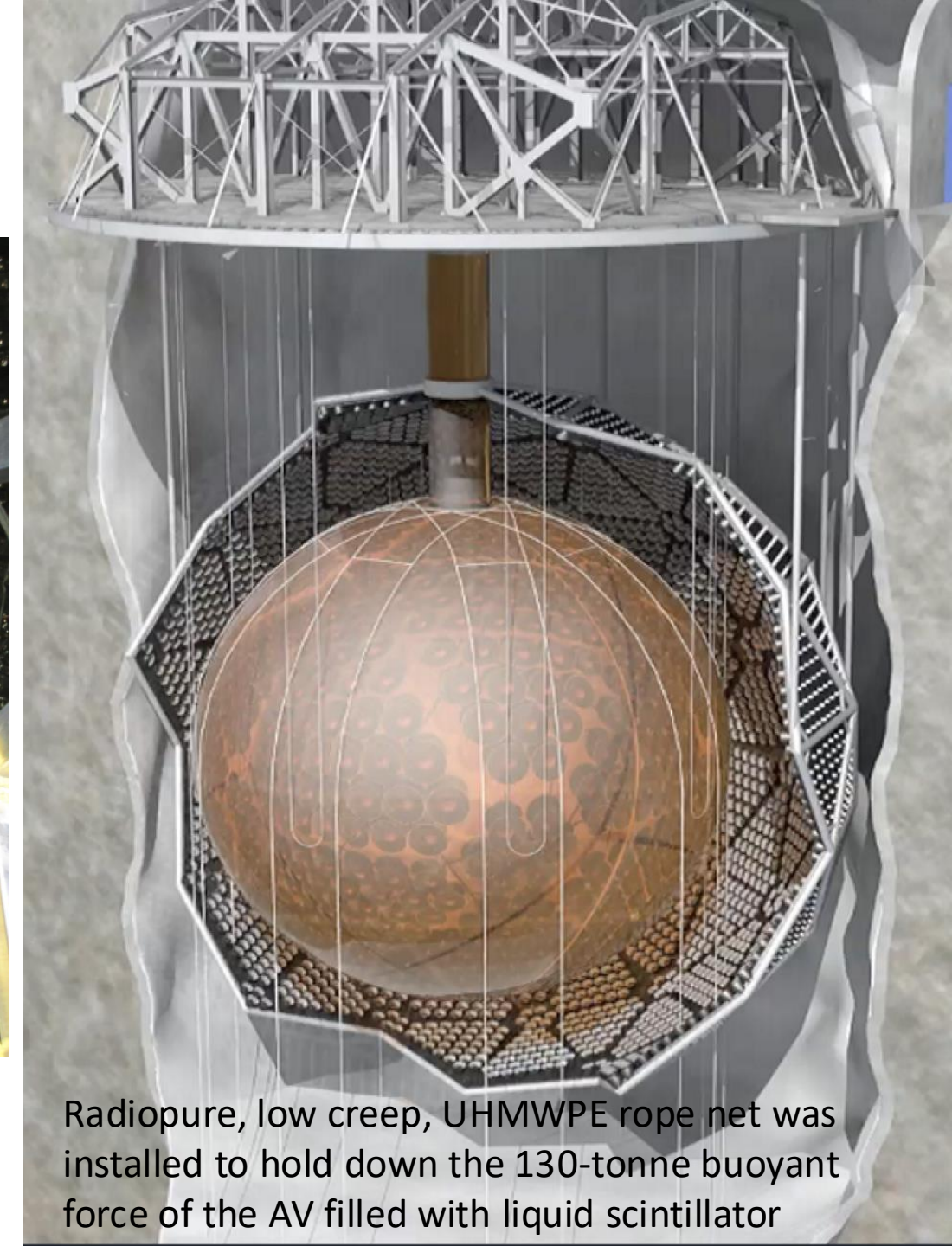
Radiopure, low creep, UHMWPE rope net was installed to hold down the 130-tonne buoyant force of the AV filled with liquid scintillator

SNO+ is the Sudbury Neutrino Observatory Filled with Liquid Scintillator

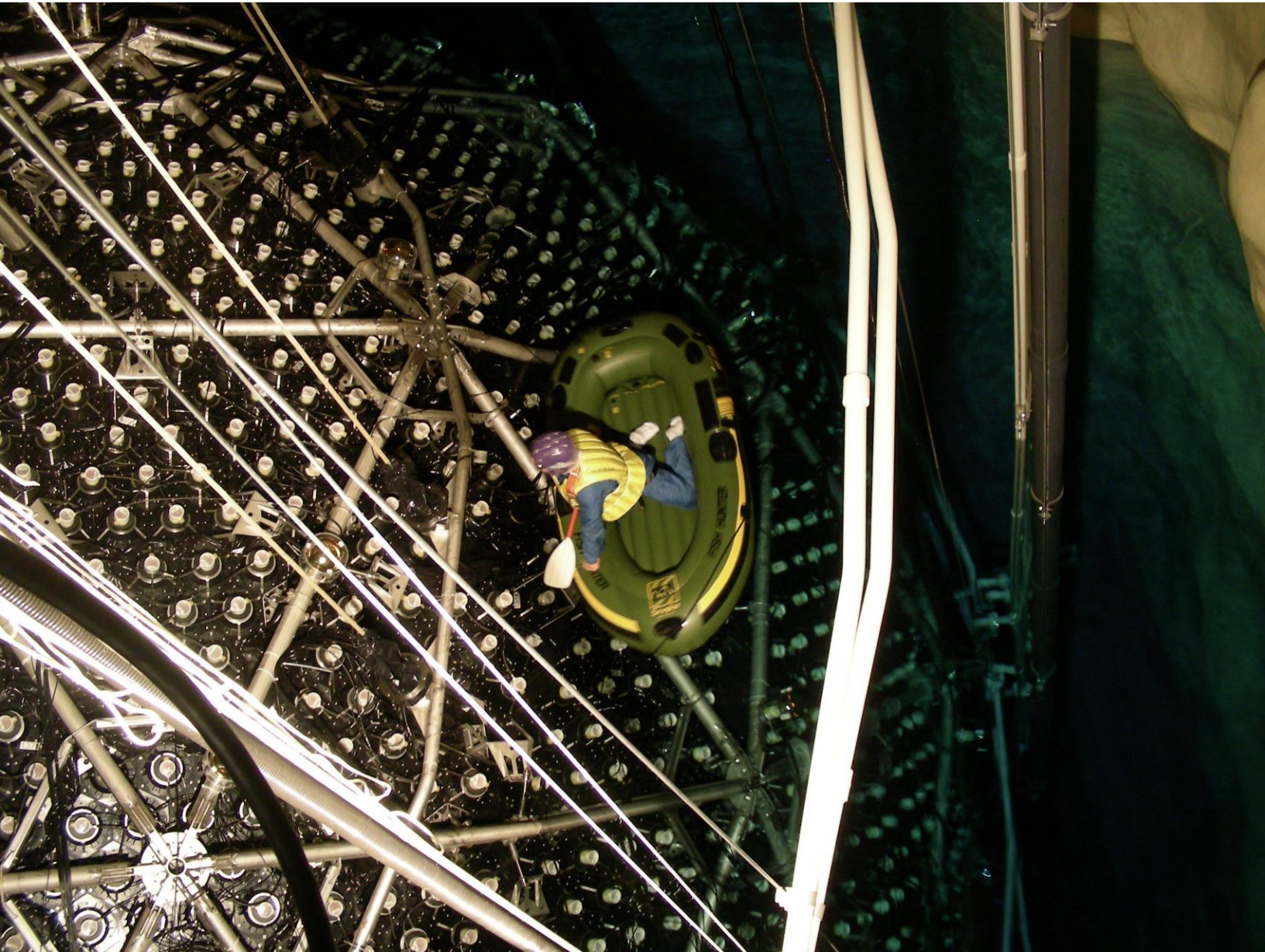


Radiopure, low creep, UHMWPE rope net was installed to hold down the 130-tonne buoyant force of the AV filled with liquid scintillator

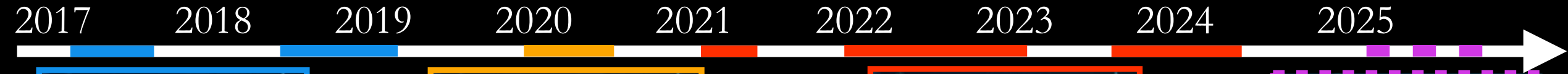
SNO+ is the Sudbury Neutrino Observatory Filled with Liquid Scintillator



Radiopure, low creep, UHMWPE rope net was installed to hold down the 130-tonne buoyant force of the AV filled with liquid scintillator



SNO+ Data-taking Timeline



Water phase

- High Rn
- Low Rn



Partial-fill phase

- Scintillator over water
- Stop in fill due to Covid



Scintillator phase

- Low PPO
- Nominal PPO
- Added bis-MSB



Next:

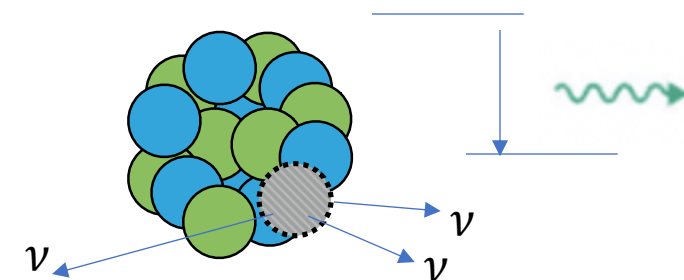
Tellurium-loaded phase

The background is a close-up, high-angle shot of water with a strong blue color cast. It features several concentric circular ripples, with the most prominent one in the upper center. The water surface is textured with smaller, irregular ripples and reflections of light, creating a shimmering effect. The overall composition is centered and balanced.

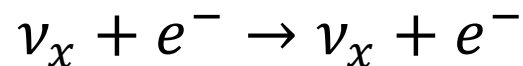
SNO+ Water Phase

SNO+ Water Phase Physics Results

- World's best limits on invisible modes of nucleon decay
 - 2022 update, published in *Phys. Rev. D*



- Solar neutrinos
 - detected via neutrino-electron elastic scattering



PHYSICAL REVIEW D **99**, 012012 (2019)

Measurement of the ^8B solar neutrino flux
in SNO+ with very low backgrounds

- now with *even lower backgrounds*

- First observation of reactor $\bar{\nu}_e + p \rightarrow e^+ + n$ events using *pure water* (undoped)
 - published in *Phys. Rev. Lett.*
 - made possible by $\sim 50\%$ neutron detection efficiency (highest in a water Cherenkov detector)

Decay Mode	Partial Lifetime Limit	Existing Limits
n	9.0×10^{29} y	5.8×10^{29} y [5]
p	9.6×10^{29} y	3.6×10^{29} y [6]
pp	1.1×10^{29} y	4.7×10^{28} y [6]
np	6.0×10^{28} y	2.6×10^{28} y [6]
nn	1.5×10^{28} y	1.4×10^{30} y [5]

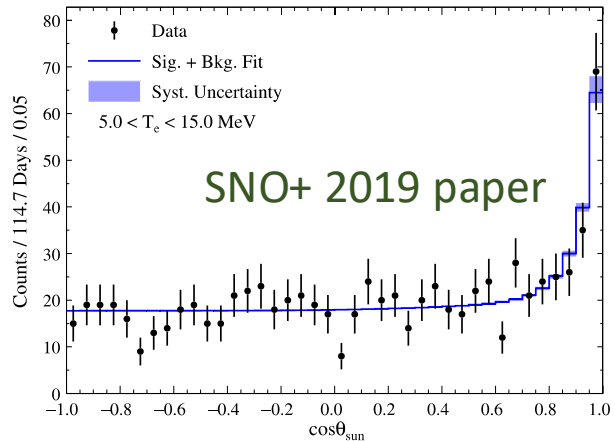
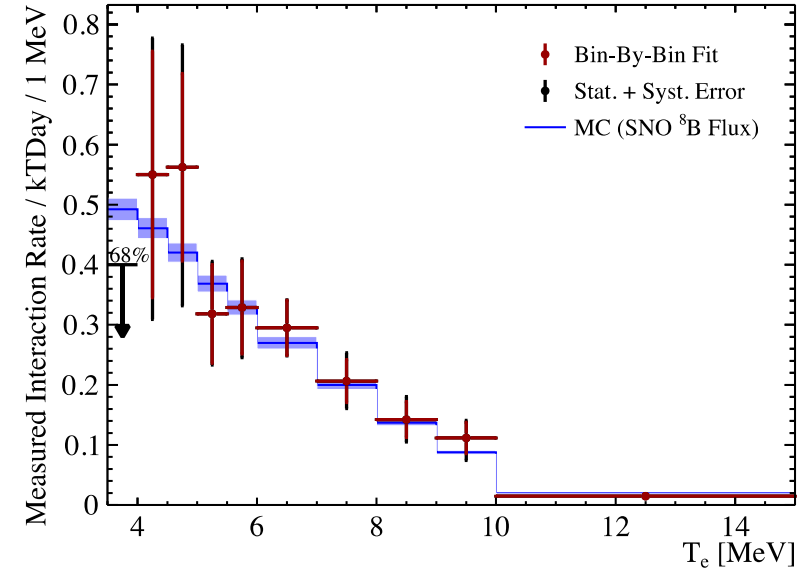
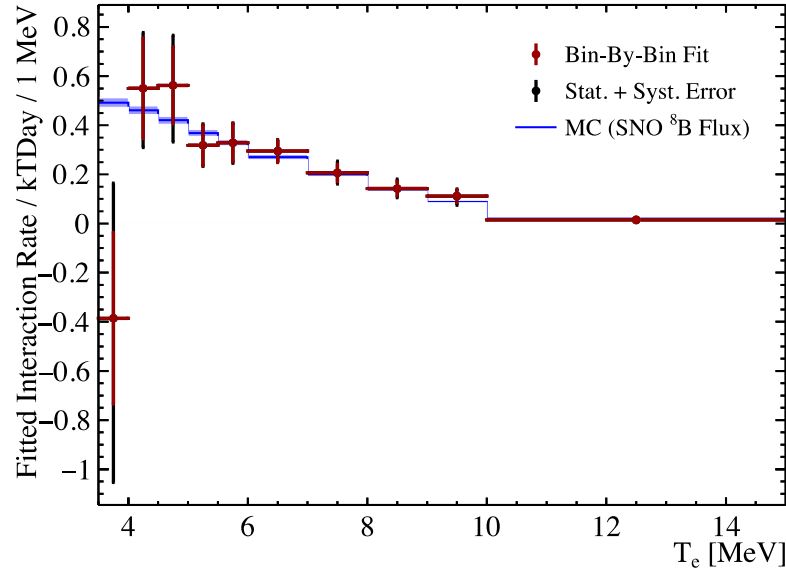
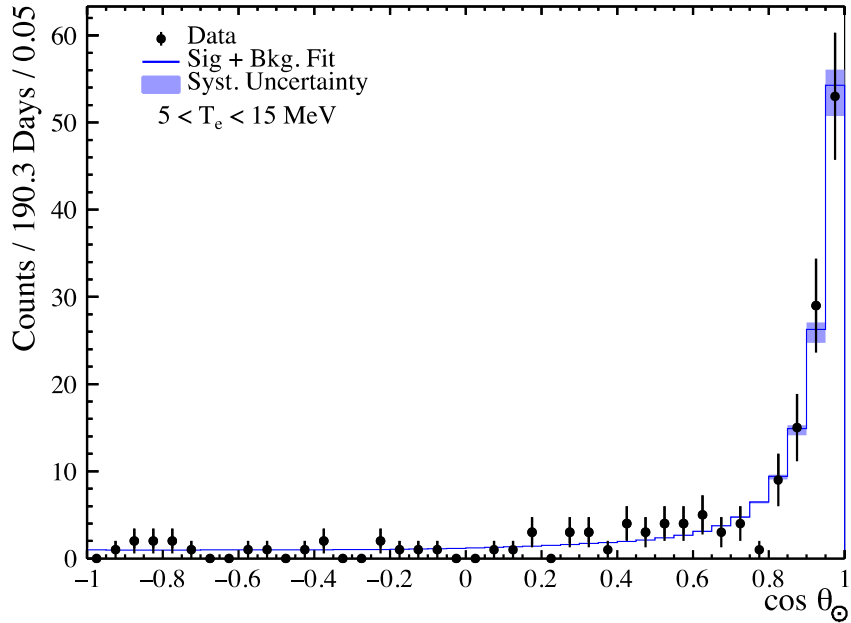
Best limits are from SNO+

Recent Result

Submitted to PRD, arXiv:2407.17595

^8B solar neutrinos in SNO+ (Full Water Phase)

Measured flux: $5.36^{+0.41}_{-0.39}(\text{stat.})^{+0.17}_{-0.16}(\text{syst.}) \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$



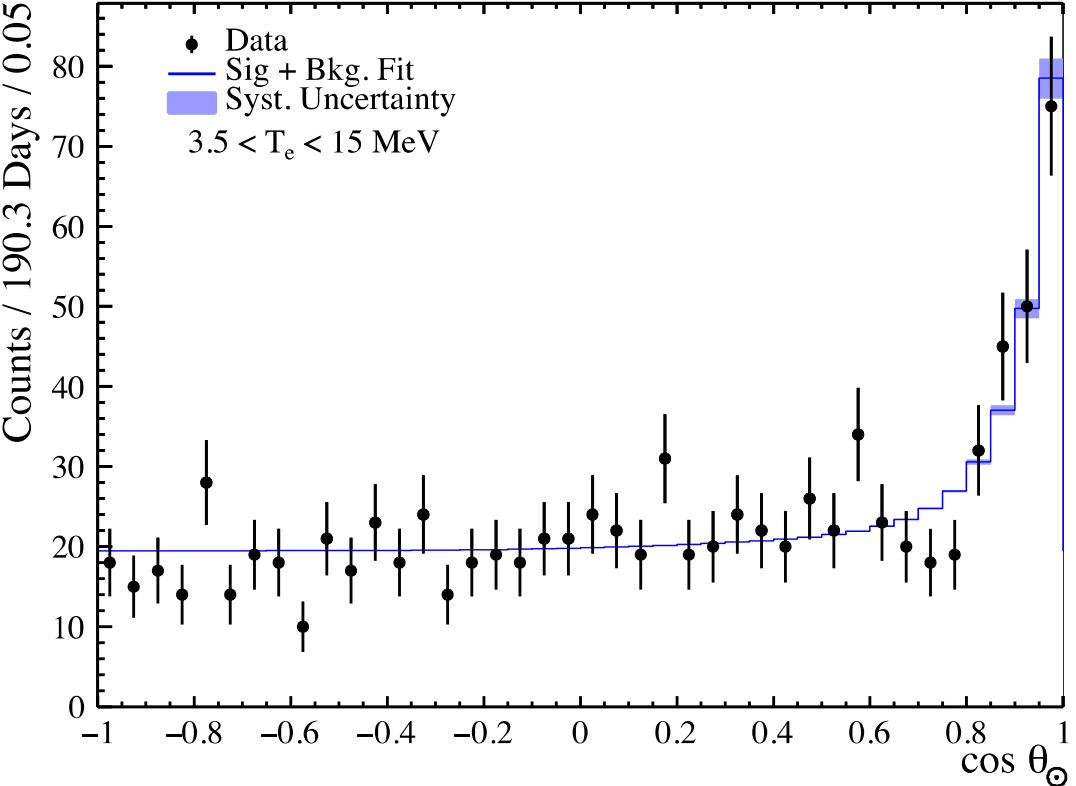
same 5.0 MeV threshold

Recent result includes 190.3 live-days data from the extended water phase, with $\sim 1/10$ Rn levels

Recent Result

^8B solar neutrinos in SNO+ (Full Water Phase)

Measured flux: $5.36^{+0.41}_{-0.39}(\text{stat.})^{+0.17}_{-0.16}(\text{syst.}) \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$



Even lower backgrounds!

Low energy threshold of 3.5 MeV (on par with the lowest achieved so far in a water Cherenkov detector)

Recent result includes 190.3 live-days data from the extended water phase, with $\sim 1/10$ Rn levels

SNO+ Water Phase list of physics publications

- Set world-leading limits on invisible modes of nucleon decay, [PRD 99, 032008 \(2019\)](#); [PRD 105 112012 \(2022\)](#)
- “Measurement of the ^8B solar neutrino flux in SNO+ with very low backgrounds”, [PRD 99, 012012 \(2019\)](#)
- Highest efficiency (~50%) for neutron detection in a water Cherenkov detector, [PRC 102, 014002 \(2020\)](#)
- Detection of antineutrinos from distant reactors using only pure water, [PRL 130, 091801 \(2023\)](#)

PHYSICAL REVIEW LETTERS **130**, 091801 (2023)

Editors' Suggestion

Featured in Physics

Evidence of Antineutrinos from Distant Reactors Using Pure Water at SNO +

240 and 340 km away

- “Measurement of the ^8B solar neutrino flux using the full SNO+ Water Phase”, [arXiv:2407.17595 \(2024\)](#)

technical papers

- SNO+ “Detector Paper” [JINST 16, P08059 \(2021\)](#)
- SNO+ Scintillator Paper “Development, characterization and deployment of the SNO+ liquid scintillator” [JINST 16, P05009 \(2021\)](#)
- Water Phase optical calibration [JINST 16, P10021 \(2021\)](#)

SNO+ Scintillator Fill



Started in mid-late 2019 and was proceeding smoothly (post-commissioning) when the pandemic struck, halting all activities for >6 months. At 365 tonnes filled (~45%), SNO+ **partial-fill** benefited from a quiet period with no operations, allowing radon backgrounds to decay and background levels in the LS to be measured.

SNO+ Partial Fill

- LS backgrounds measured at

$^{214}\text{BiPo}$ delayed coincidences for U chain

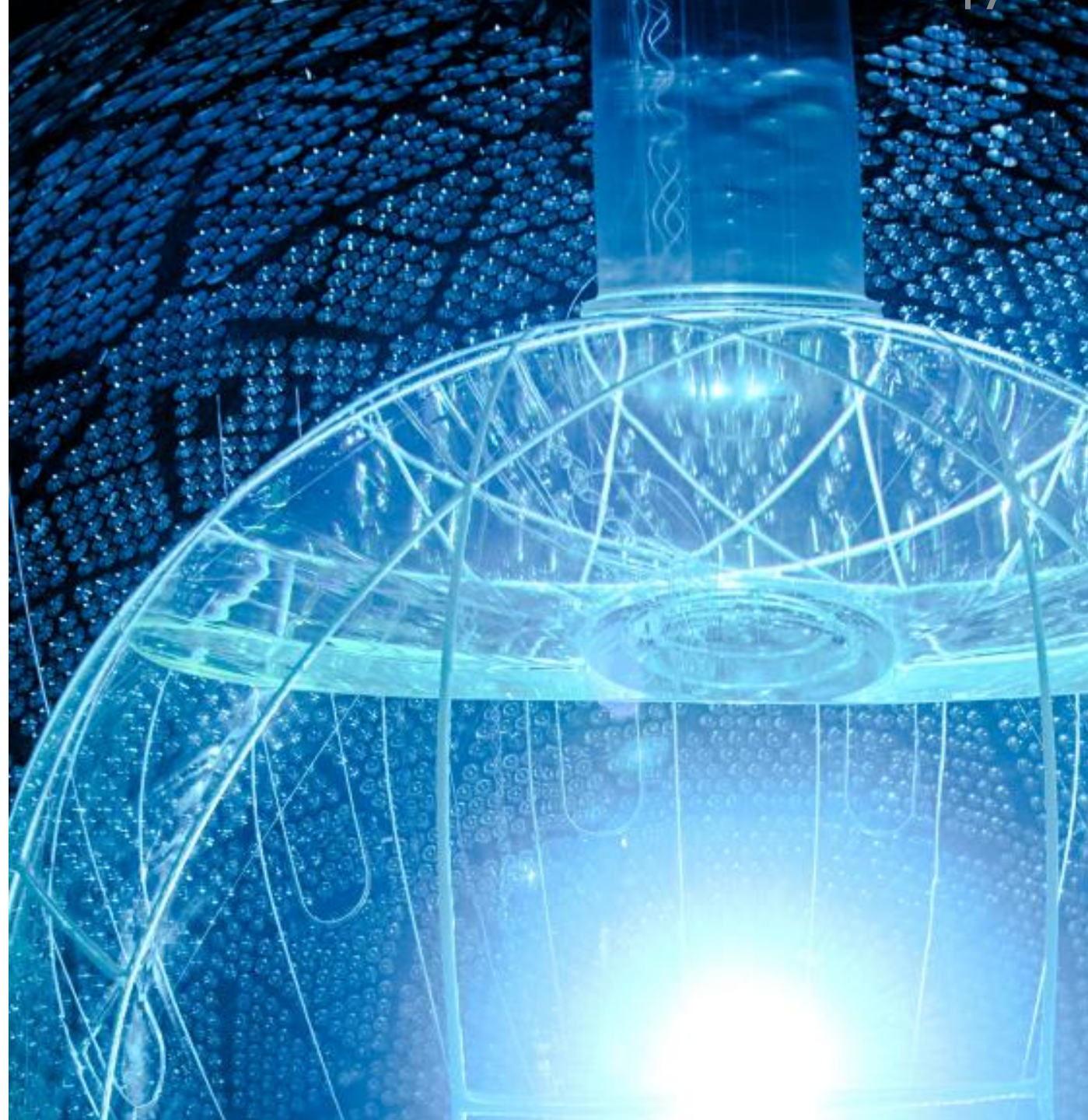
$$(4.7 \pm 1.2) \times 10^{-17} g_{\text{U}}/g_{\text{LAB}}$$

$^{212}\text{BiPo}$ delayed coincidences for Th chain

$$(5.3 \pm 1.5) \times 10^{-17} g_{\text{Th}}/g_{\text{LAB}}$$

meeting SNO+ background targets for double beta decay

- Optical properties of LS 👍
- Also physics from SNO+ partial fill...

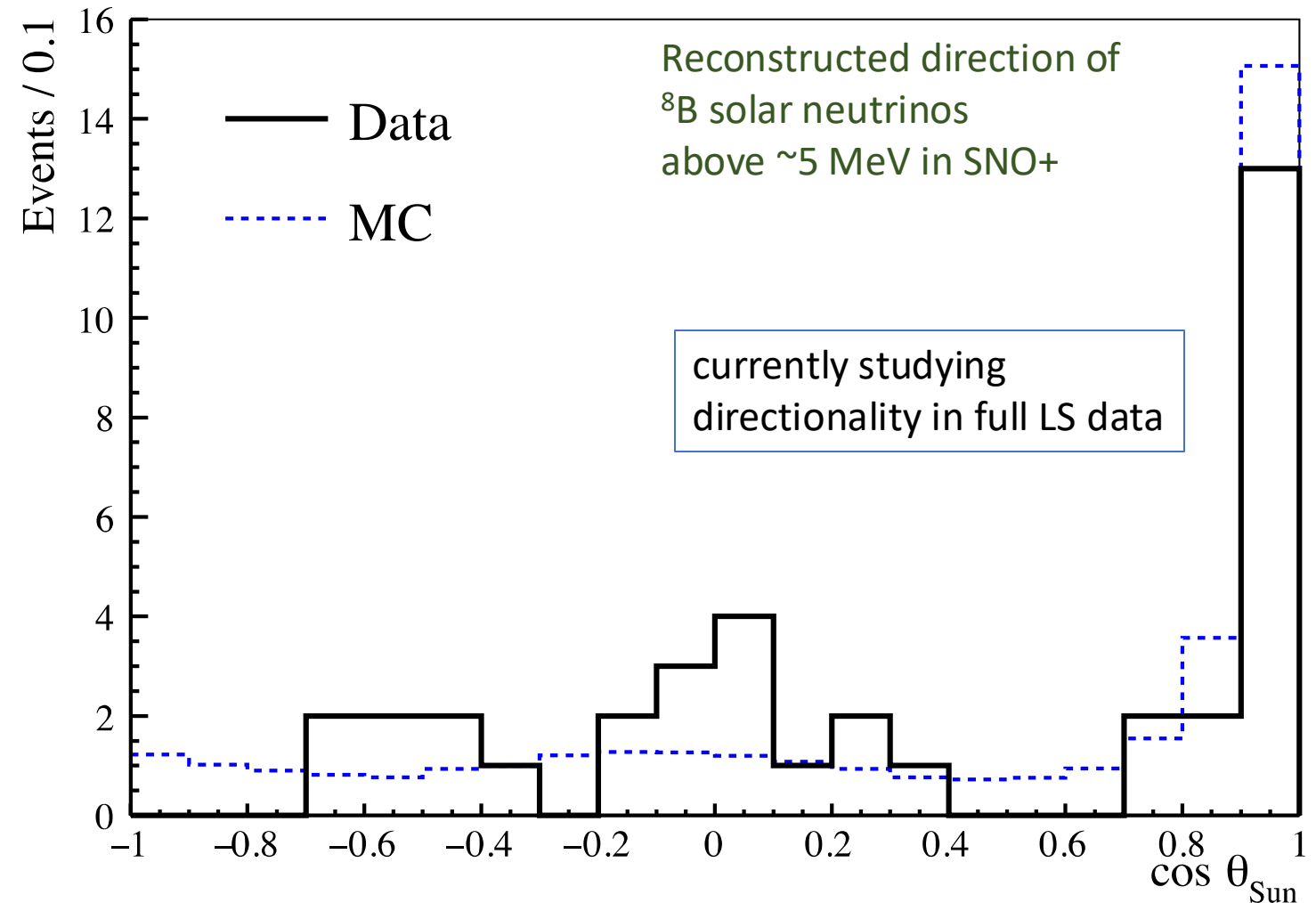


Event-by-Event Direction Reconstruction of Solar Neutrinos in SNO+ Liquid Scintillator

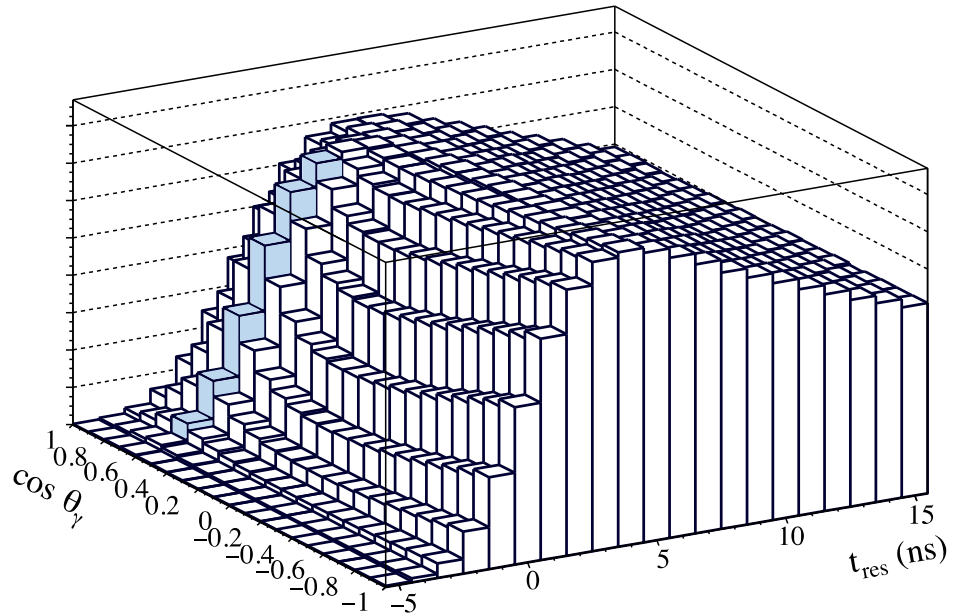
- Borexino had published the observation of a correlation between early PMT hits in the forward direction caused by the Cherenkov light produced by ${}^7\text{Be}$ and CNO solar neutrinos in liquid scintillator
“Correlated and Integrated Directionality”
- SNO+ result: each recoil electron event’s direction can be reconstructed by fitting with the combined Cherenkov+scintillation pdf

This is a first – **event-by-event direction reconstruction** of MeV events **in liquid scintillator**!

“Event-by-event direction reconstruction of solar neutrinos in a high light yield liquid scintillator”, *Phys. Rev. D* **109**, 072002 (2024)

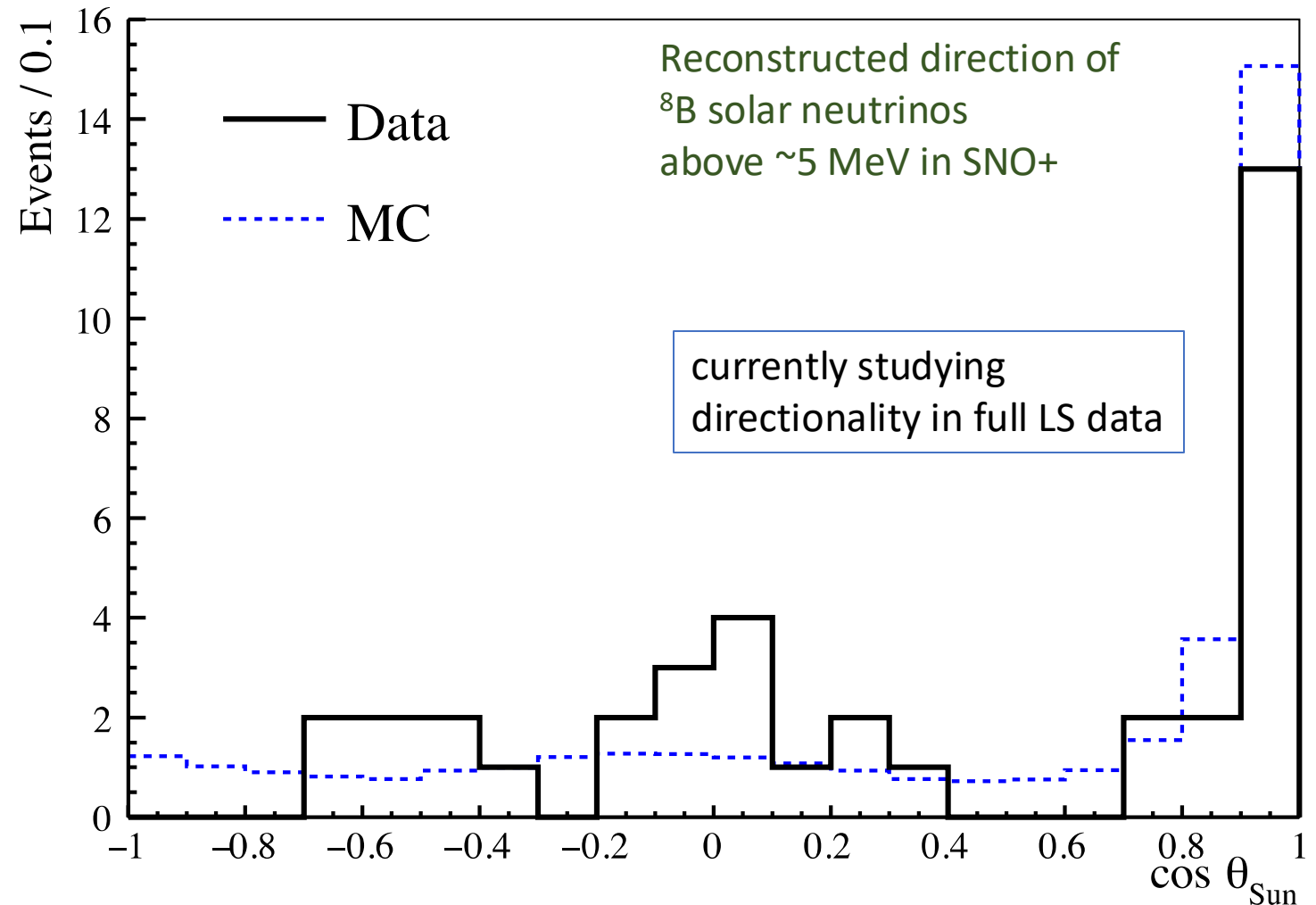


Event-by-Event Direction Reconstruction of Solar Neutrinos in SNO+ Liquid Scintillator



This is a first – **event-by-event direction reconstruction** of MeV events in liquid scintillator!

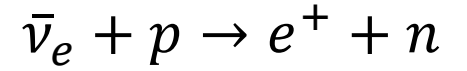
“Event-by-event direction reconstruction of solar neutrinos in a high light yield liquid scintillator”, *Phys. Rev. D* **109**, 072002 (2024)



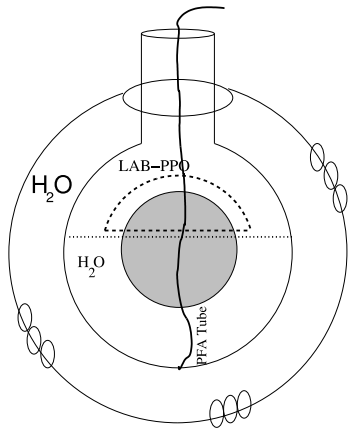
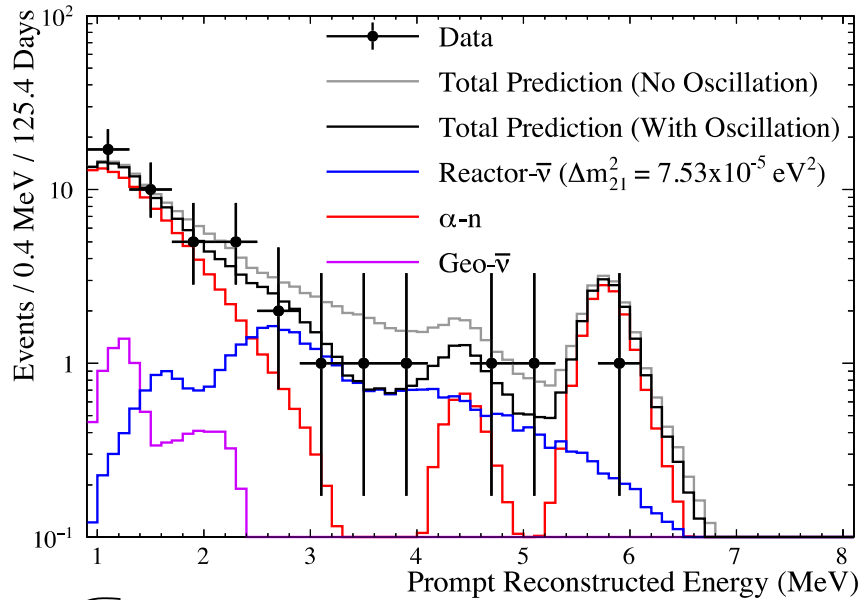
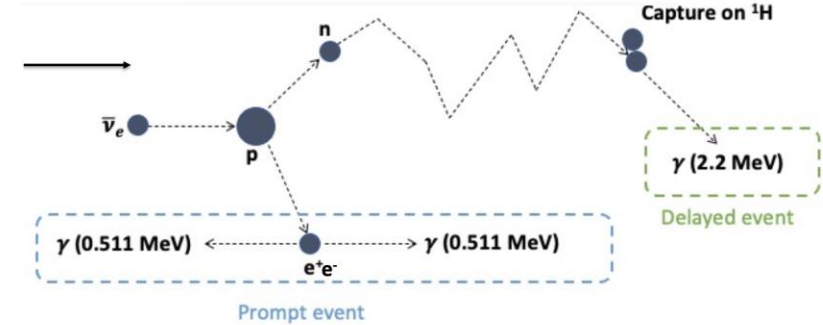
Recent Result

Submitted to PRD, arXiv:2405.19700

Antineutrino Physics with Partial-Fill Scintillator

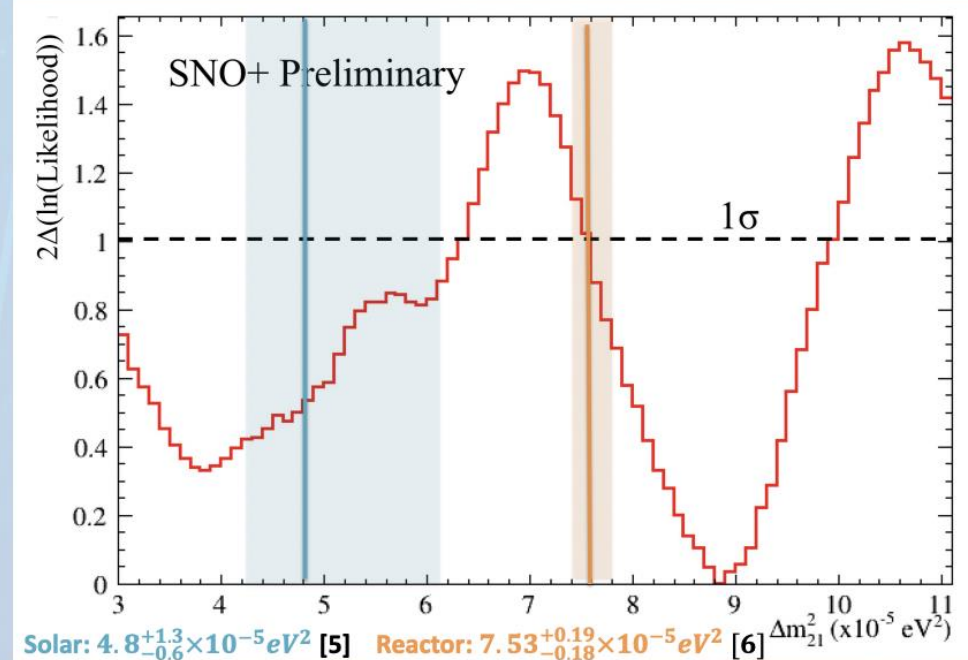


$$\text{Measured } \Delta m_{21}^2 = 8.85_{-1.33}^{+1.10} \times 10^{-5} \text{ eV}^2$$



Δm_{21}^2 likelihood space for observed IBD candidates

Δm_{21}^2 values corresponding to **solar** neutrino measurements and **reactor** neutrino measurements allowed within 1σ .



The background of the slide is a complex, abstract visualization of a spherical structure, likely representing the SNO+ detector. It features a dense network of thin, light-colored lines forming a grid-like pattern over a dark blue and green background. There are several bright, glowing spots scattered throughout, particularly in the lower half, which could represent individual scintillator crystals or data points. The overall aesthetic is technical and scientific.

SNO+ Scintillator Phase

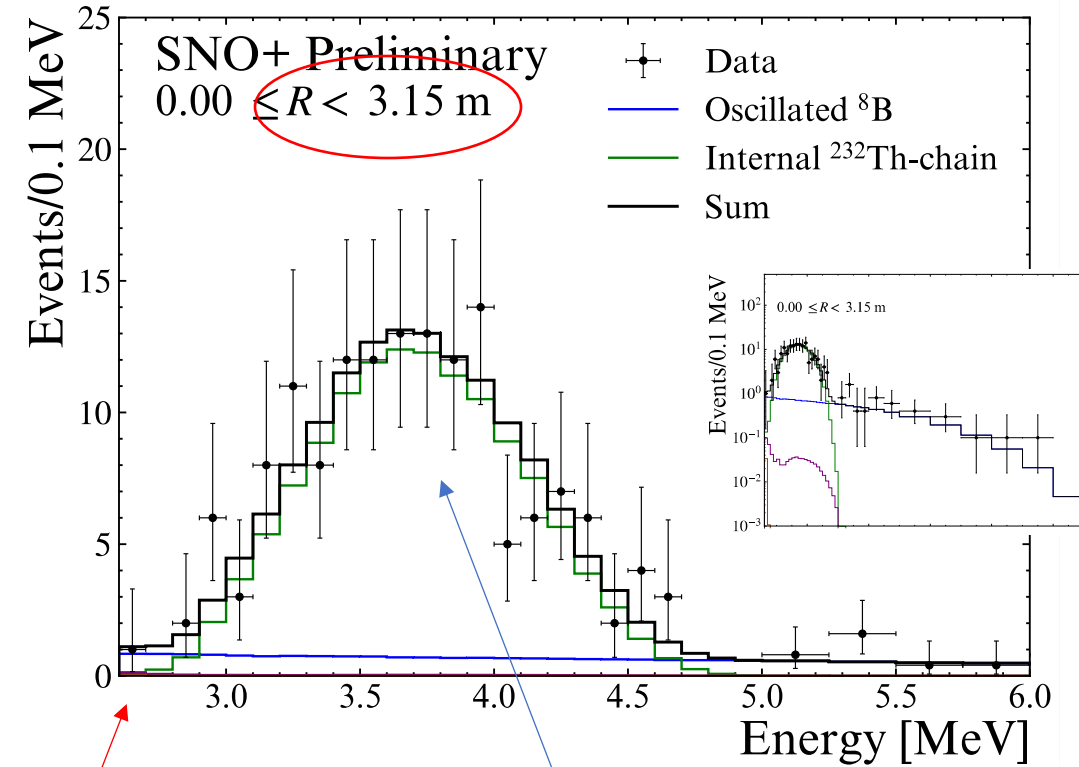
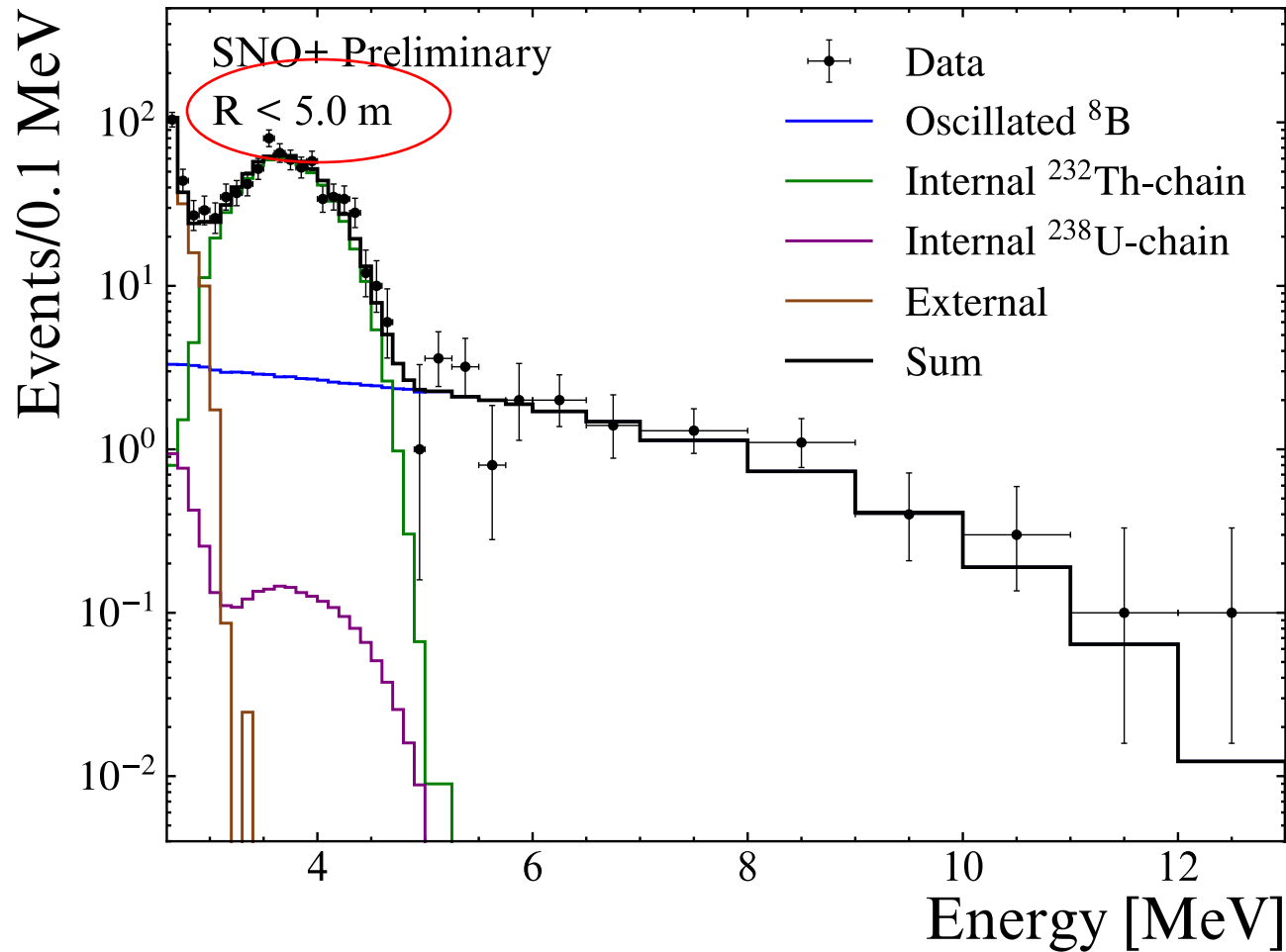
Physics Goals in the Scintillator Phase

- Solar neutrinos (e.g. ${}^8\text{B}$) at lower energies
- Reactor antineutrinos flux, spectrum, oscillations (Δm_{12}^2 , in particular)
- Geo neutrinos
- SNO+ is supernova neutrino live
- and other “exotic” physics (e.g. MIMP dark matter searches, fermionic dark matter, DSNB – diffuse supernova neutrino background, nucleon decay)

Recent Result

SNO+ Scintillator Phase: ^8B Solar Neutrinos

Preliminary data as shown at Neutrino 2024

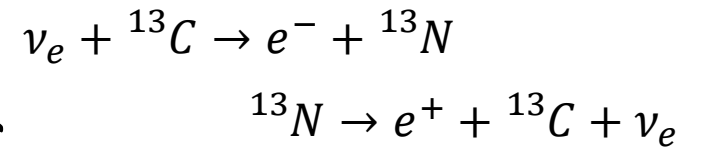


These energies for ^8B solar ν 's haven't been studied before!

Internal Th backgrounds between 3 and 4 MeV – background rejection using directionality/multi-site being studied

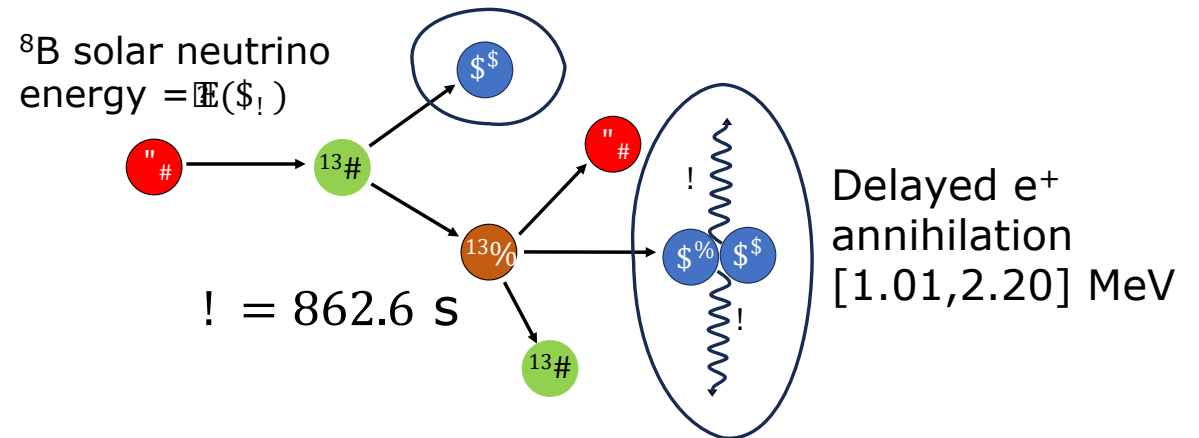
Recent Result

Charged-current ν_e capture on ^{13}C

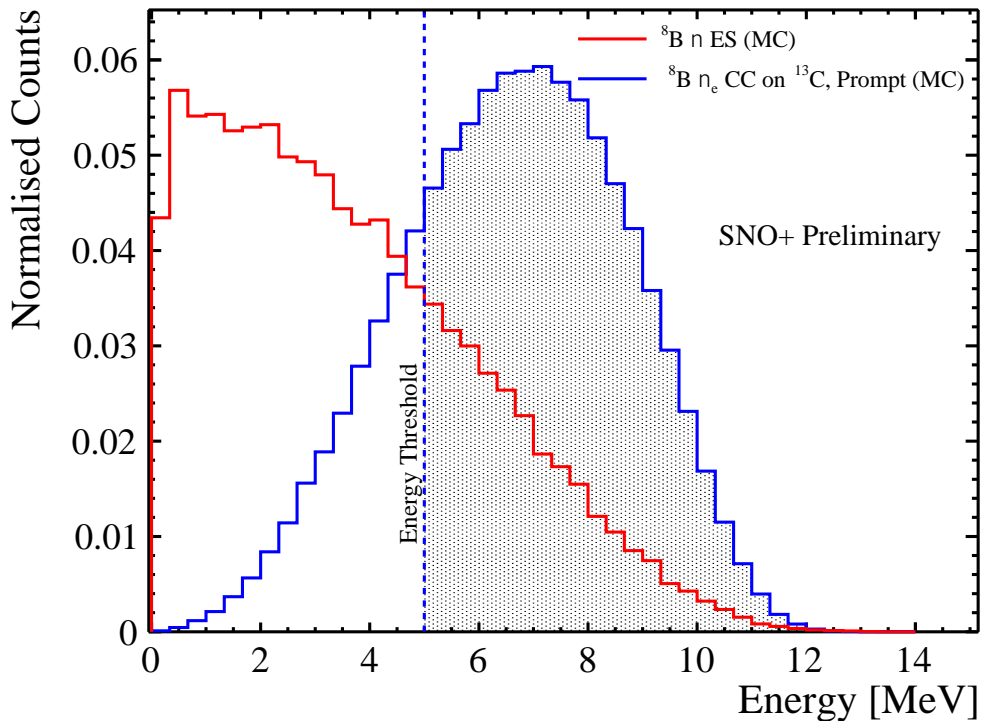


Up-to-now unobserved CC
reaction for electron neutrinos
(solar ν_e)

$$\text{Prompt } e^- \text{ energy} = E(\nu_e) - 2.2 \text{ MeV}$$



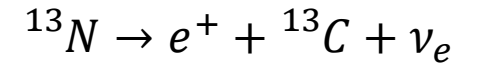
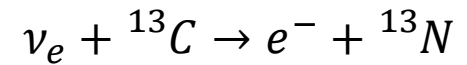
^{13}C only 1.1% isotopic abundance; but $\sim 12\times$ higher cross section per target than ${}^8\text{B}$ solar neutrino ES on electrons



Cosmogenic backgrounds (e.g. ${}^{11}\text{Be}$) obscured this reaction in previous experiments but at SNOLAB depth, the main background is an accidental coincidence with a ${}^8\text{B}$ solar neutrino ES event

Recent Result

Charged-current ν_e capture on ^{13}C



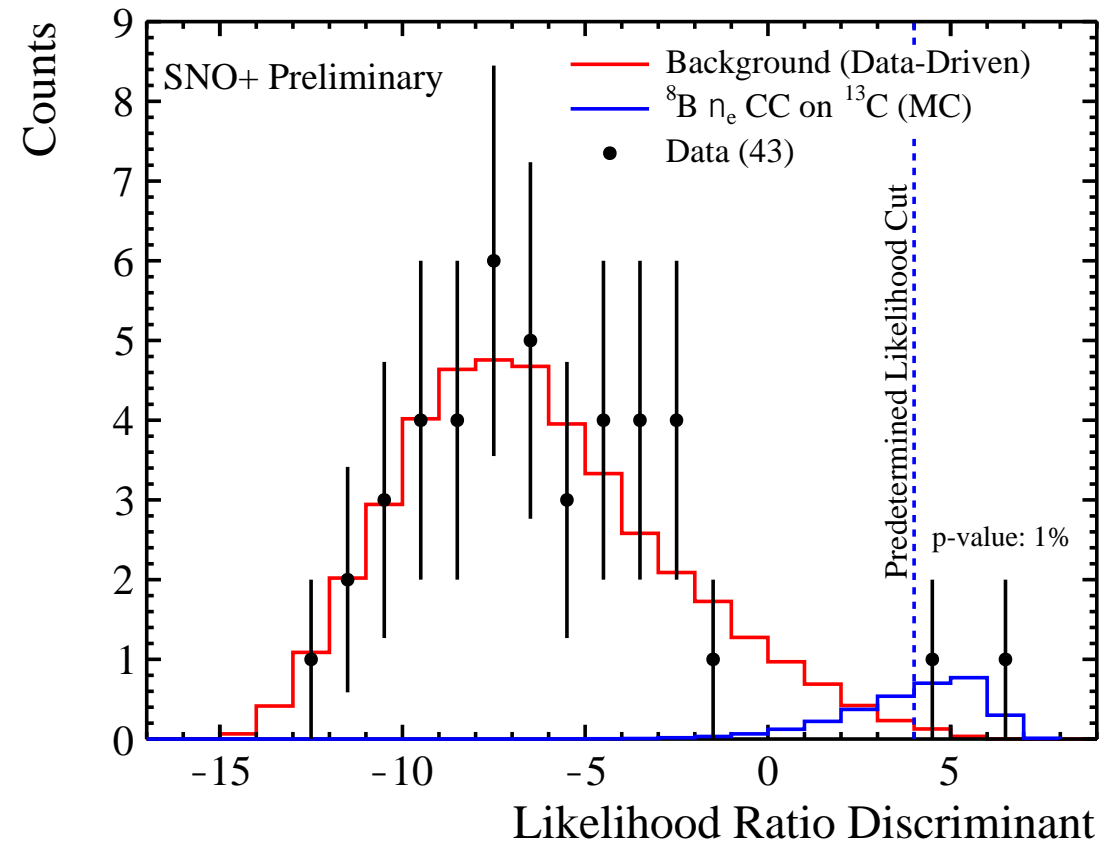
150.51 live-days in SNO+ analyzed

2 events found!

(0.1-0.3 backgrounds expected)

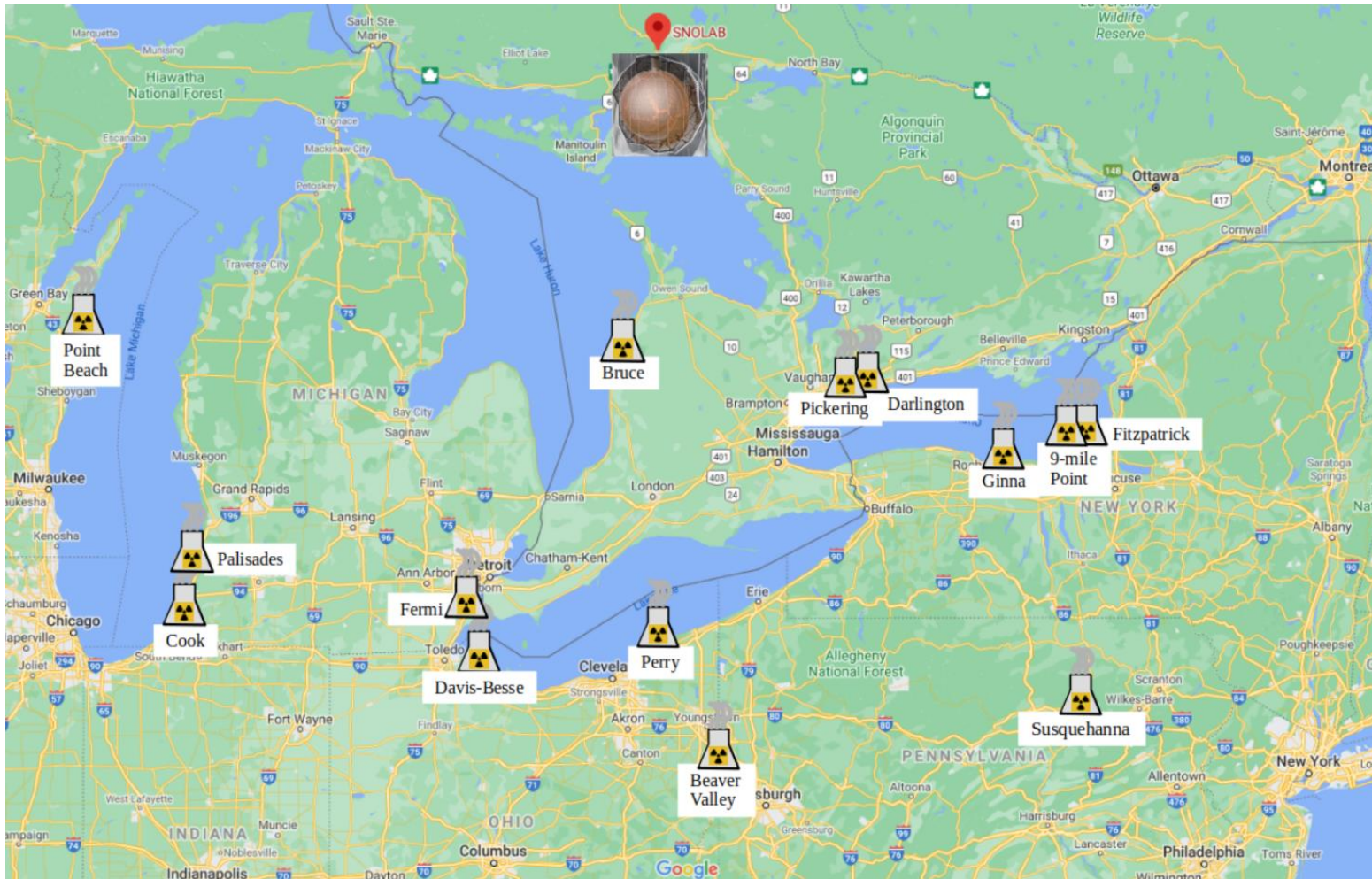
Indication of a signal!

Accumulating more statistics;
publication in preparation...

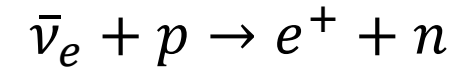
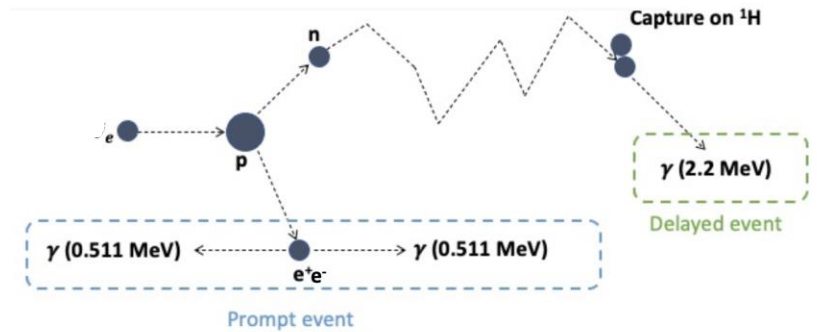


EXPECTED	BOX	LIKELIHOOD
BACKGROUND	0.31	0.17
SIGNAL	1.83	1.79

Reactor Antineutrinos in SNO+



Inverse Beta Decay (IBD)



Coincidence event

Prompt – positron kinetic energy (several MeV)

plus 1.022 MeV from annihilation γ 's

Delayed – neutron capture 2.2 MeV γ

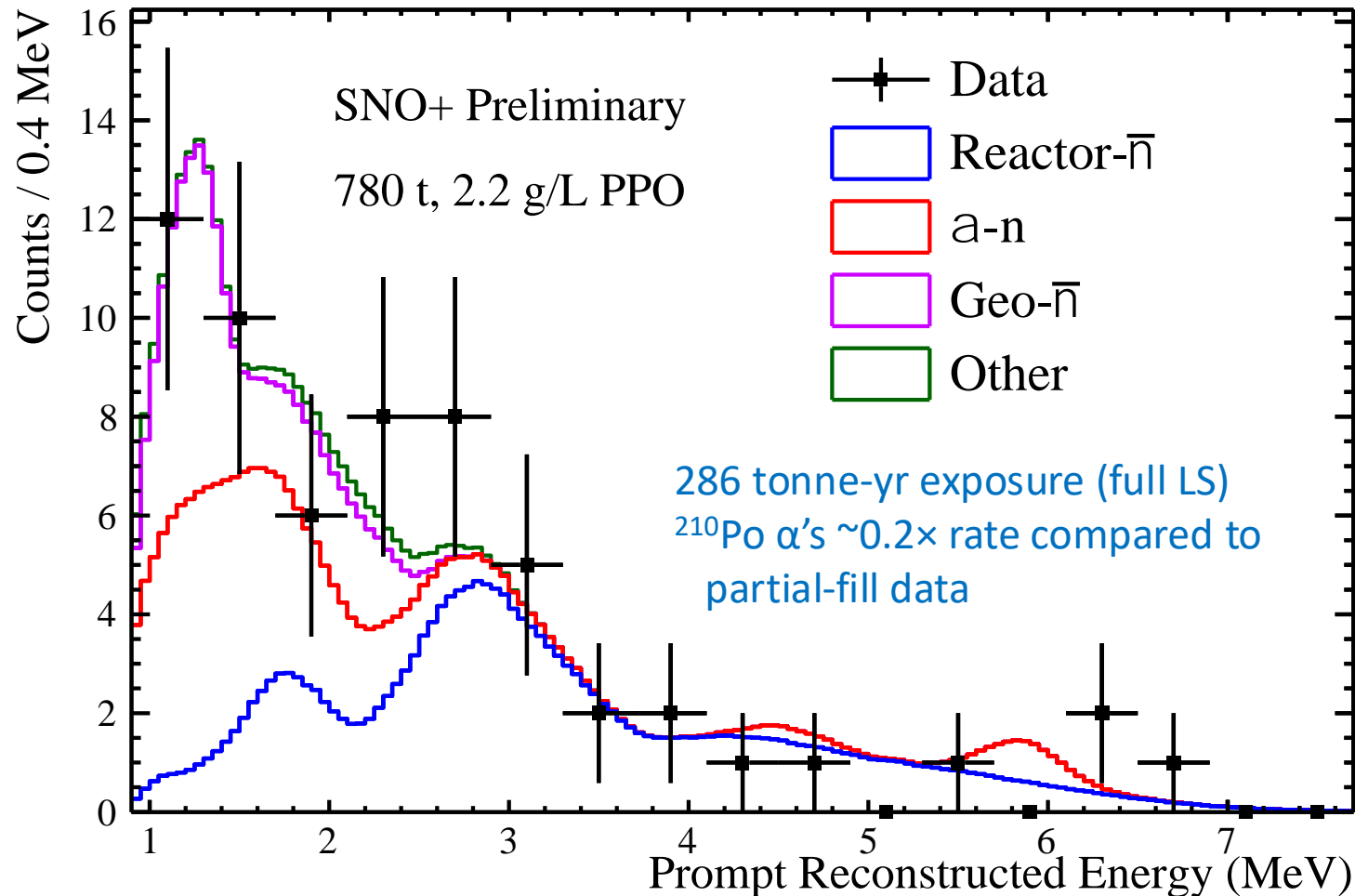
Recent Result

SNO+ Scintillator Phase: Antineutrinos in Full LS

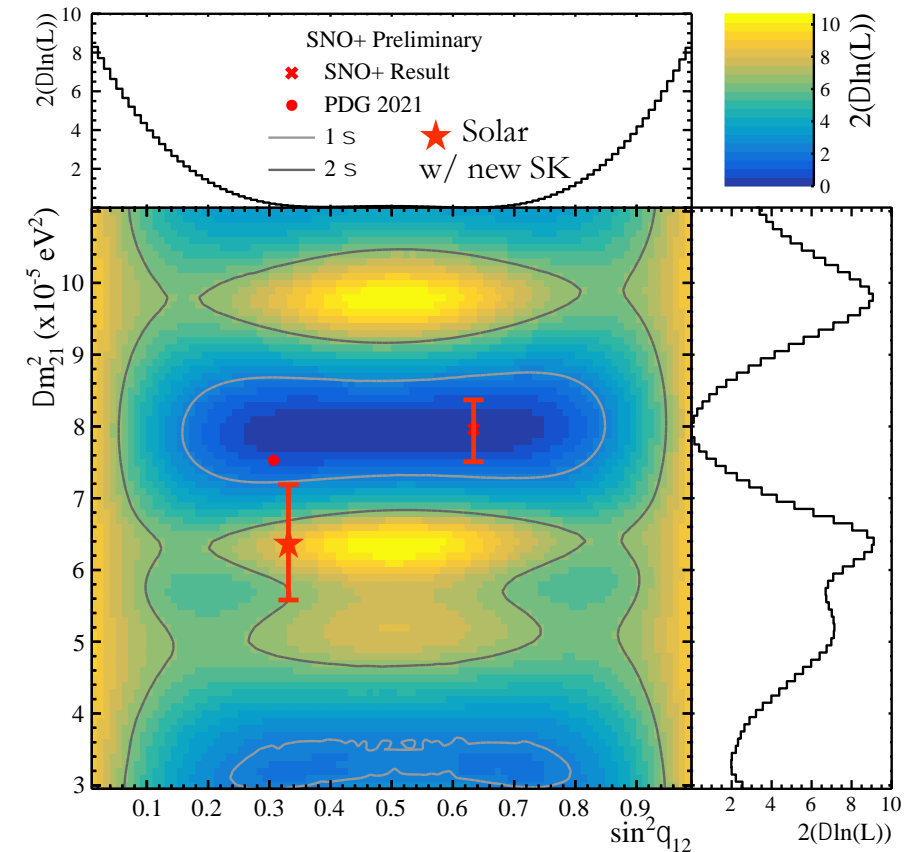
Preliminary data as shown at Neutrino 2024

SNO+ is detecting geo neutrinos!!

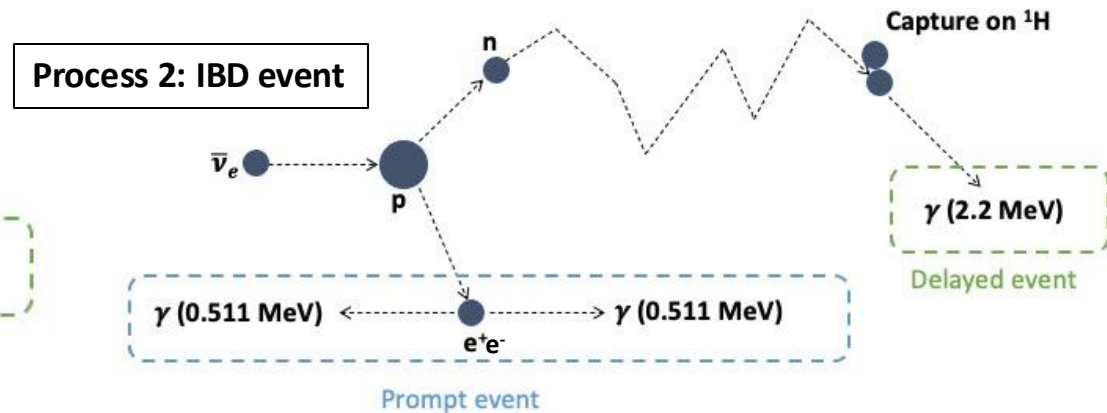
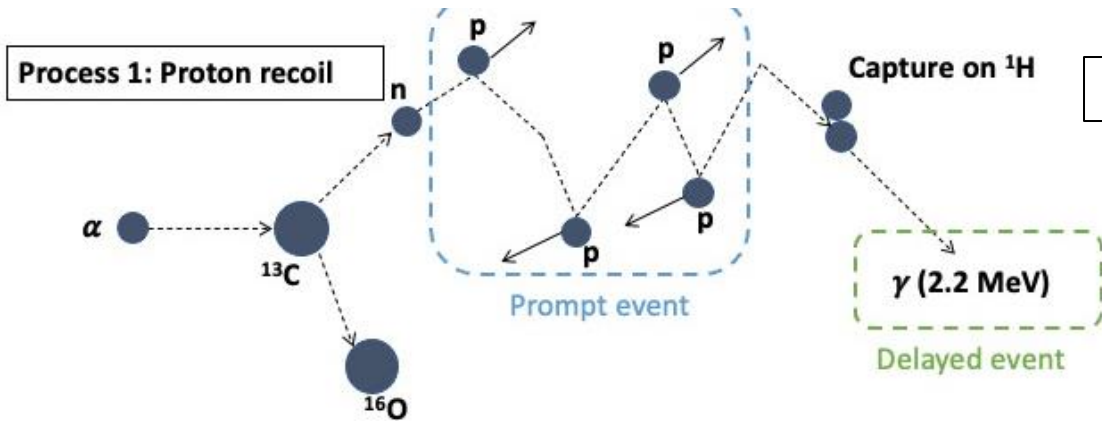
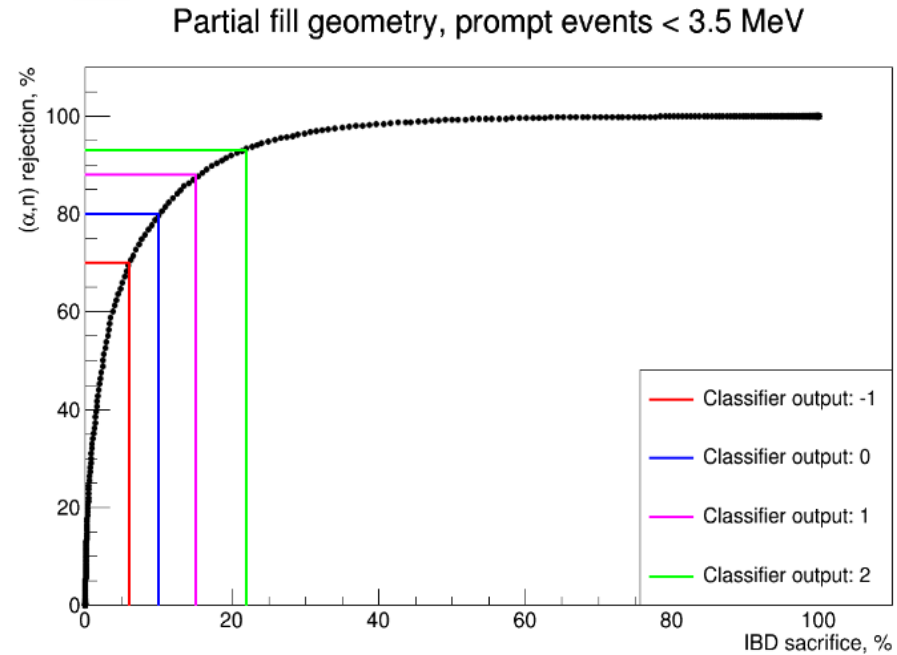
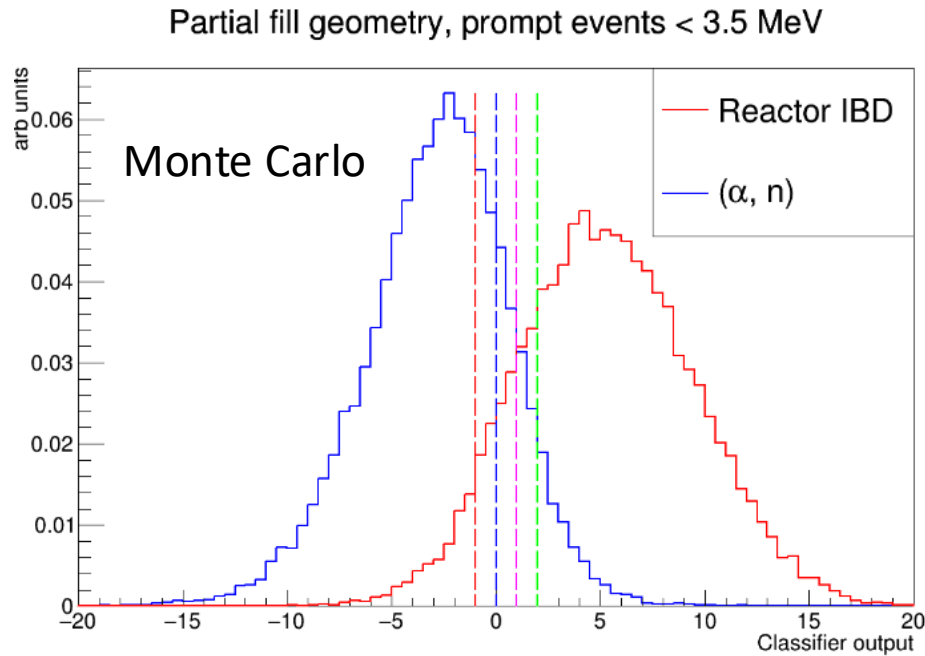
Ultra-prelim geo neutrino flux: 64 ± 44 TNU



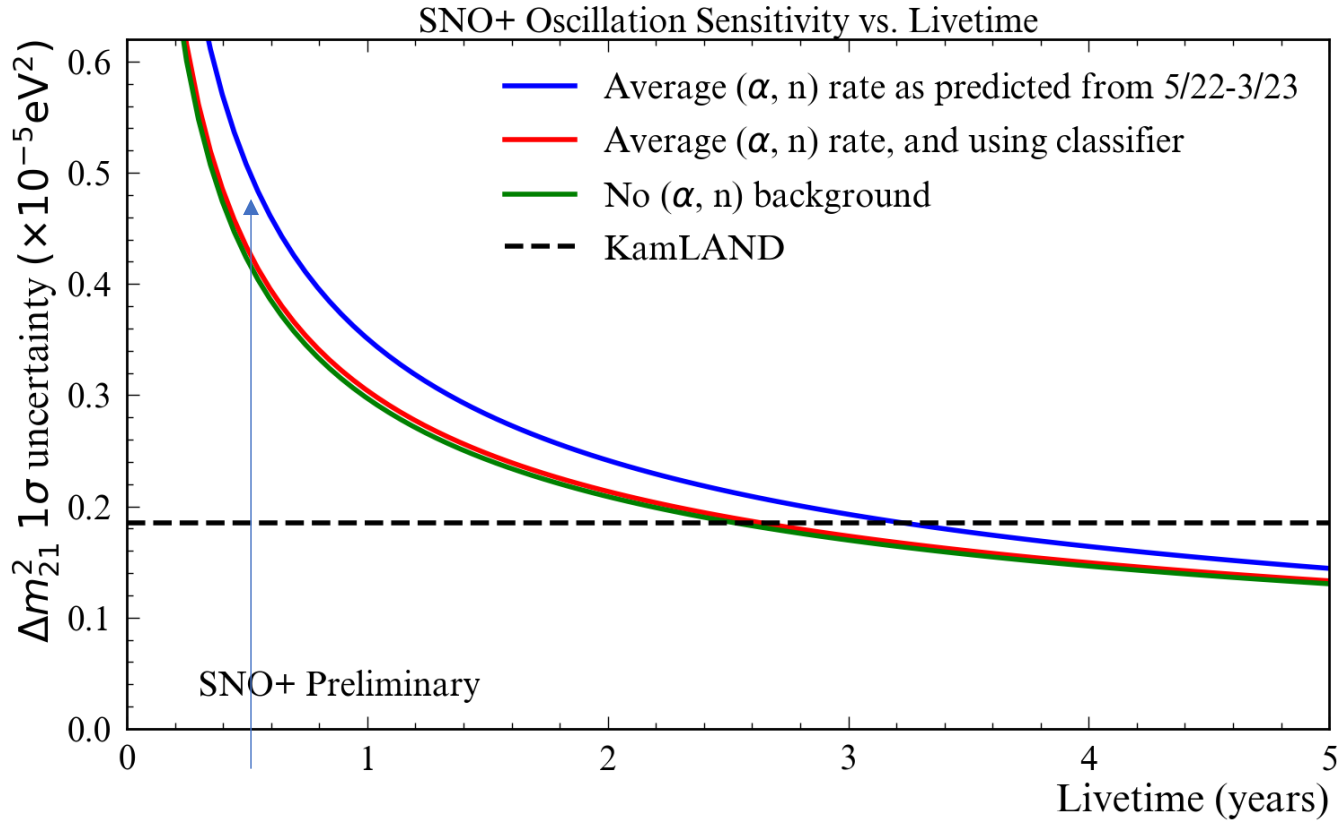
Measured $\Delta m_{21}^2 = 7.96_{-0.41}^{+0.48} \times 10^{-5} \text{ eV}^2$



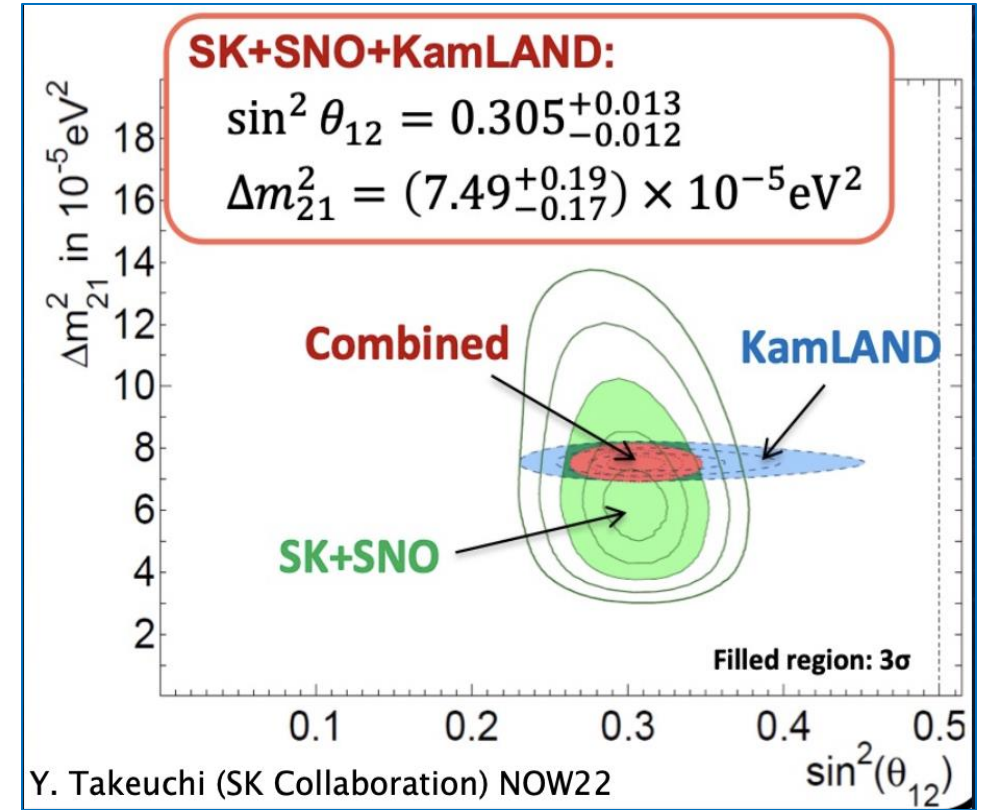
(α, n) Classifier (tested in partial-fill) – New SNO+ analysis technique



Objectives for SNO+ Scintillator Phase Antineutrino: Reactor Antineutrinos Δm_{12}^2

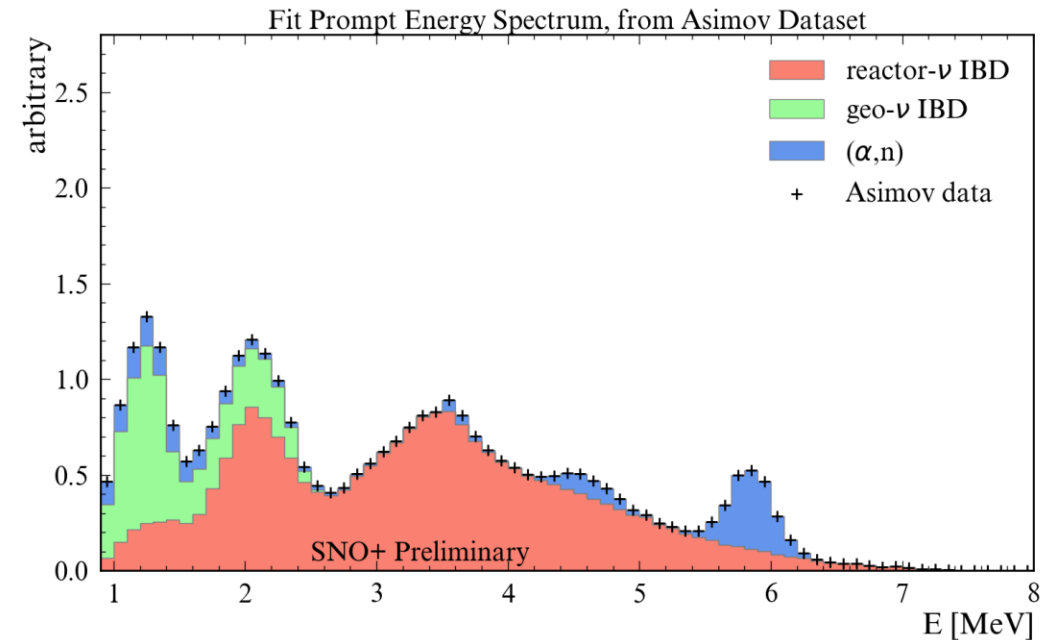
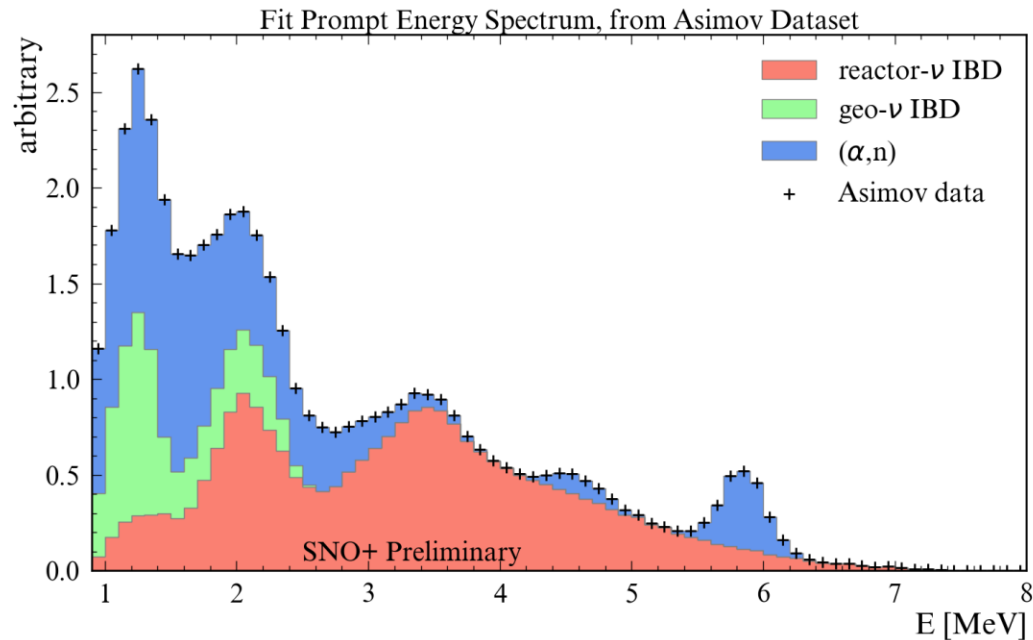


test Δm_{12}^2 with ν_e (solar) and $\bar{\nu}_e$ (reactor)
 can easily distinguish 5.0 from 7.5 ($\times 10^{-5} \text{eV}^2$)



Objectives for SNO+ Scintillator Phase Antineutrino: Geo Neutrinos

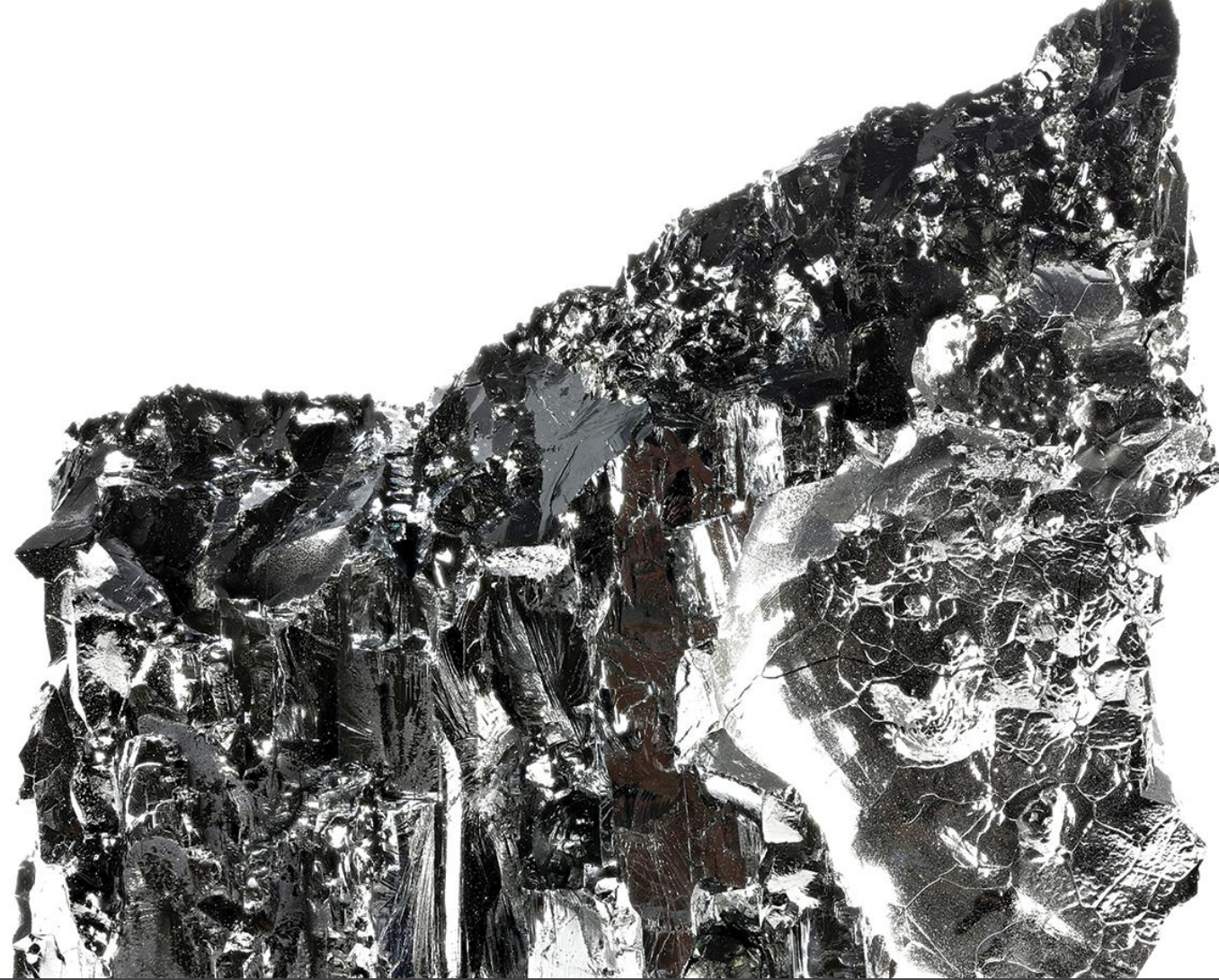
Simulated impact on prompt energy spectrum from (α,n) classifier:



(α,n) classifier being retuned for 2.2 g/L PPO; will improve geo neutrino signal extraction – first results by end of 2024

The advantages of a well-understood detector with very low backgrounds...

- are being demonstrated!
- enabling a wide variety of interesting physics topics to be studied – great for HQP training!
- With the **detector performing well**; with **all background components being measured and constrained** (most coming in at or below target levels), it looks promising for the final phase of SNO+...



SNO+ Tellurium Double Beta Decay Phase

Summary

- SNO+ is an **operating liquid scintillator neutrino detector** filled with LAB + 2.2 g/L PPO (and now also 2.2 mg/L bis-MSB) and taking data
- **Diverse** program of neutrino physics is underway as the experiment prepares for the next phase – double beta decay search by adding natural tellurium to the scintillator
- Recent results:
 - Water Phase ^8B solar neutrinos with even lower backgrounds – paper submitted
 - Partial-fill reactor antineutrino oscillations – paper submitted
 - Full-fill reactor antineutrino oscillations measurements – presented
 - Geo neutrino flux (ultra-preliminary measurement) – presented
 - ^8B solar neutrino measurements: flux, spectrum, oscillation analysis – presented
 - CC ν_e capture on ^{13}C , never observed before – first indications presented

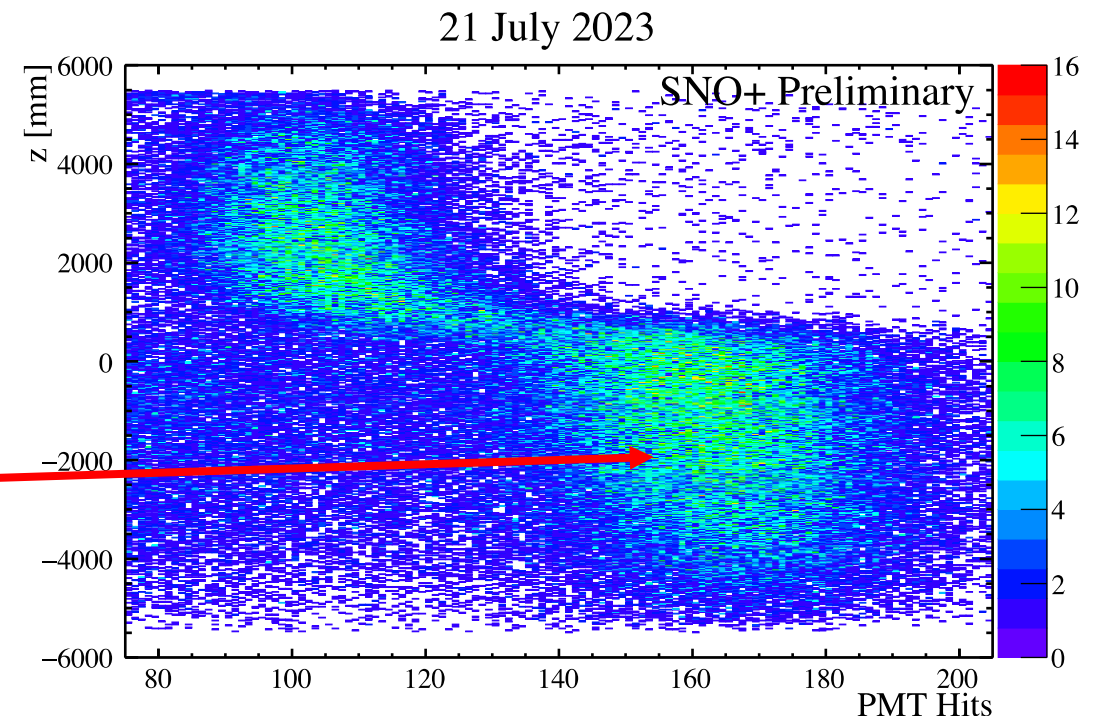


BACKUP

Recent addition of bis-MSB (and BHT)

- Recently added 2.2 mg/L bis-MSB to the liquid scintillator cocktail (plus a small amount of BHT, antioxidant) in preparation for SNO+ Te Double Beta Decay Phase
- Boosted light yield in the detector by $\sim 1.7\times$

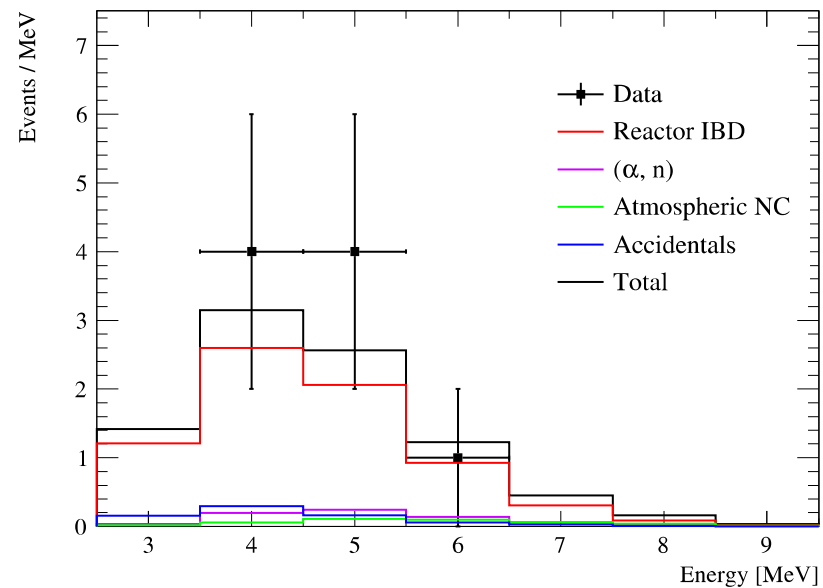
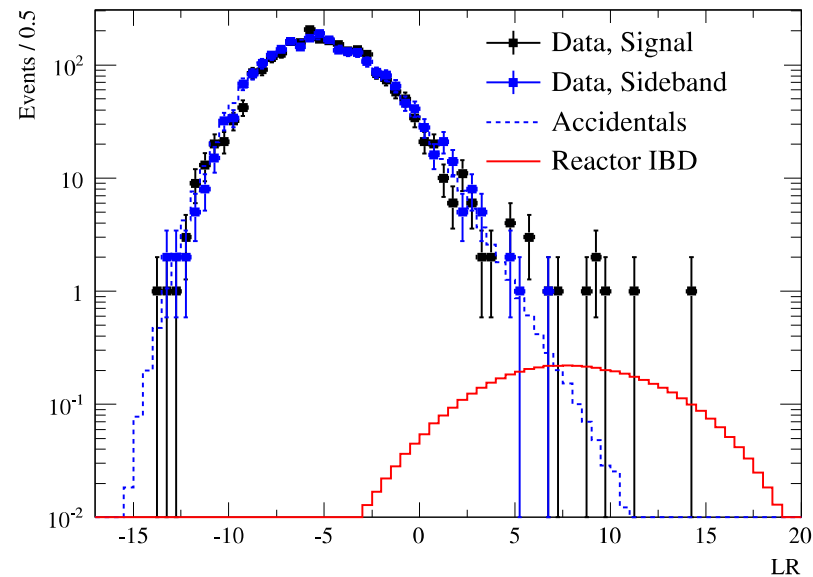
Plot on the right shows ^{210}Po events in detector versus z position after adding ~ 0.5 mg/L bis-MSB to the bottom of the detector



Antineutrino IBD Events in SNO+ Water? Yes!

Water Phase – Detection of IBD events (reactor antineutrinos) using pure water → **this is a first**

Two independent analyses – likelihood ratio and Boosted Decision Tree – both with 3σ detection significance; using event selection overlap + non-overlap, calculated combined discovery significance of 3.5σ



SNO+ Scintillator Purification Plant

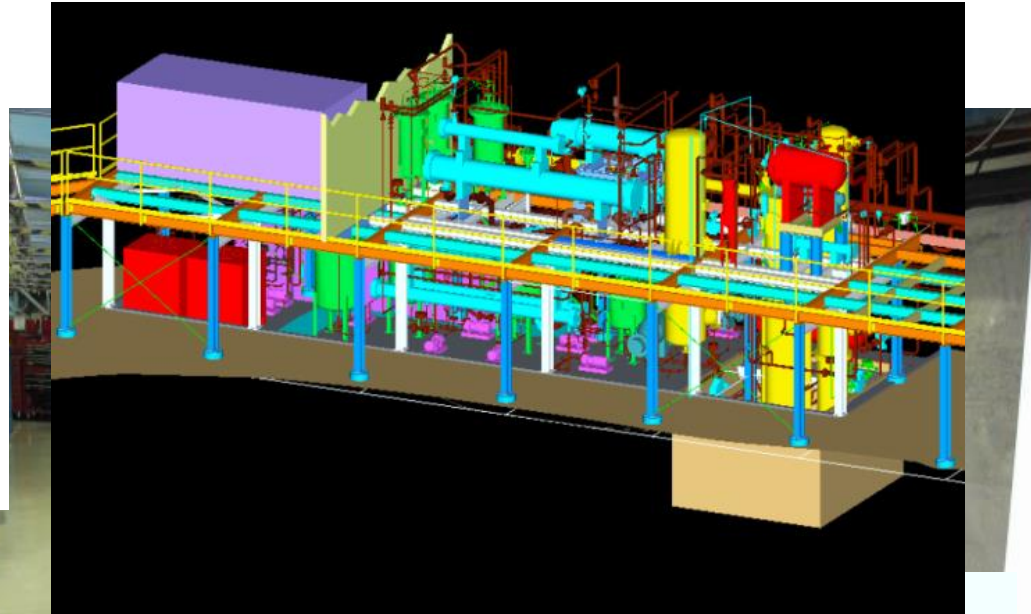
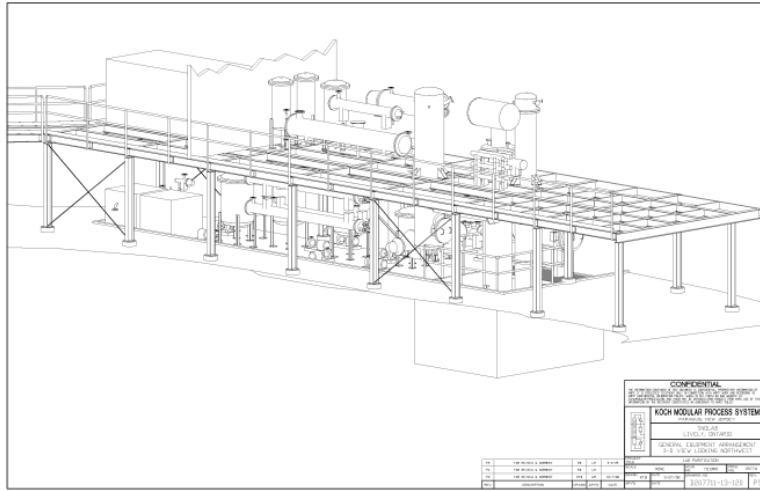
- reinforced mezzanine steel
- enlarged D2O pit “mining in a clean room”
- installed columns, vessels, heat exchangers, tank, pumps, valves, high-grade sanitary piping (orbital-welded, electropolished stainless steel tubing)
- utility plumbing (cooling water, compressed air, vent, boil-off nitrogen)
- process control, wiring, instrumentation, electrical
- firewalls, fire detection and suppression



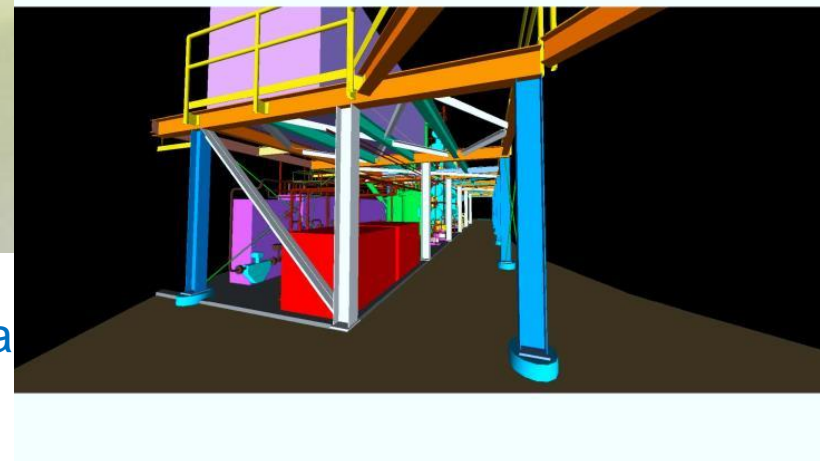
SNO heavy water purification system was here

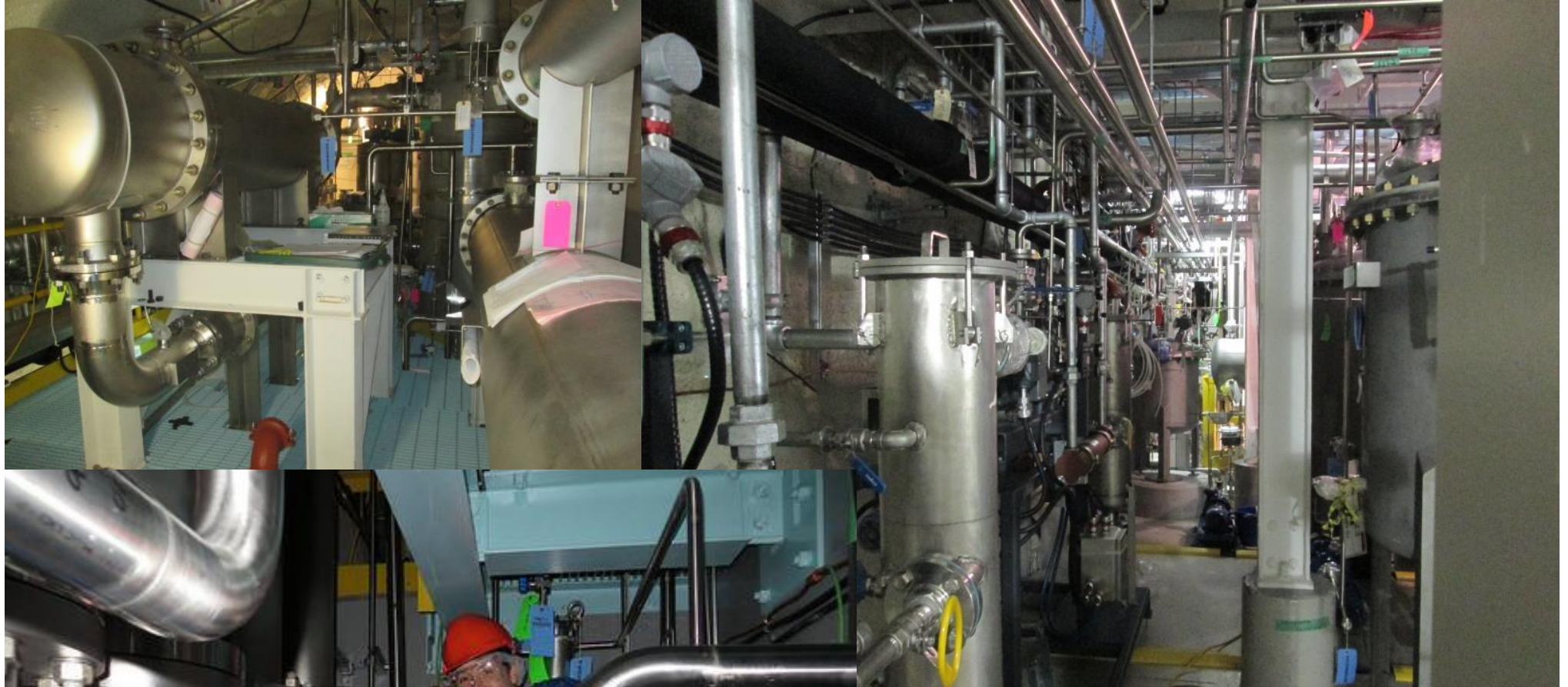
SNO+ Scintillator Purification Plant

- rein
- enla
clea
- inst
hea
pur
san
welca, electropolished
stainless steel tubing)
- utility plumbing (cooling
water, compressed air, vent,
boil-off nitrogen)
- process control, wiring,
instrumentation, electrical
- firewalls, fire detection and
suppression



SNO hea





SNO+ upgrades also included

- Refurbishing the electronics
- Repair of many “dead” PMT bases
- All-new DAQ
- New cover gas system
- New calibration systems capable of deploying in LAB scintillator
- New *in-situ* injected LED/laser light calibration system
- Calibration system cameras (for photogrammetry)

...in addition to the hold-down ropes and the scintillator plant

