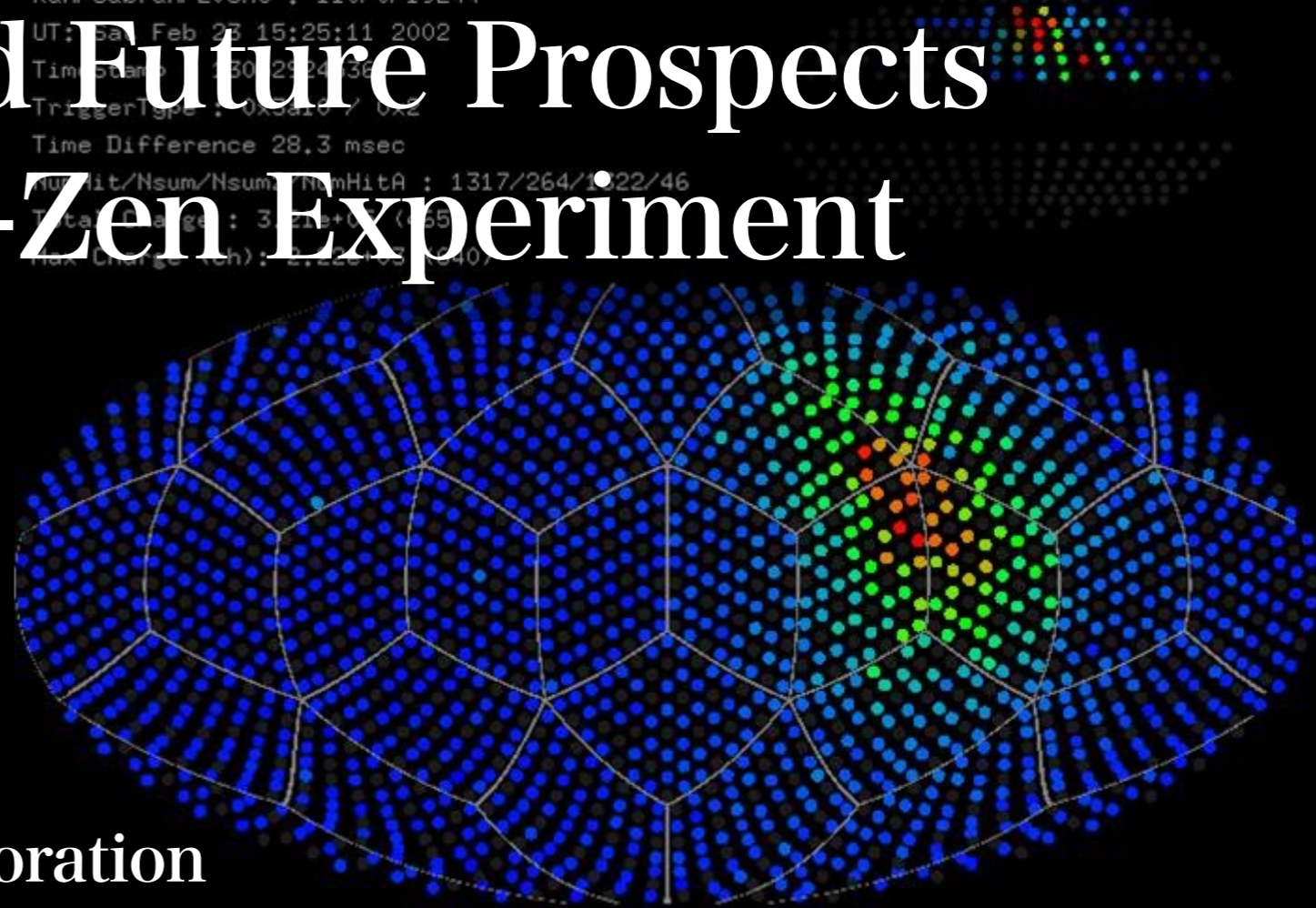


Latest Results and Future Prospects of the KamLAND-Zen Experiment

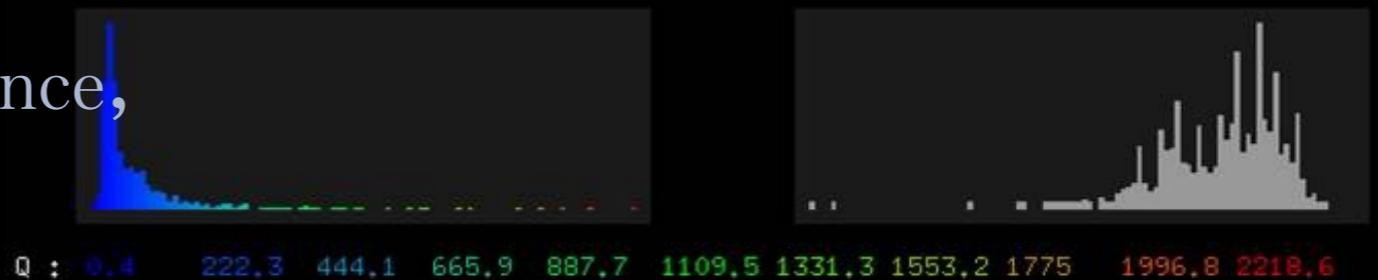
```
KamLAND Event Display
Run/Subrun/Event : 110/0/19244
UT: 5a, Feb 23 15:25:11 2002
TimeStamp : 30292436
TriggerType : 0x54167 0x2
Time Difference 28.3 msec
NHit/Nsum/NsumA/NHitA : 1317/264/122/46
Total Charge : 3.21e+07 (455)
Max Charge (ch) : 2.22e+05 (040)
```



Haruhiko MIYAKE

for the KamLAND-Zen collaboration

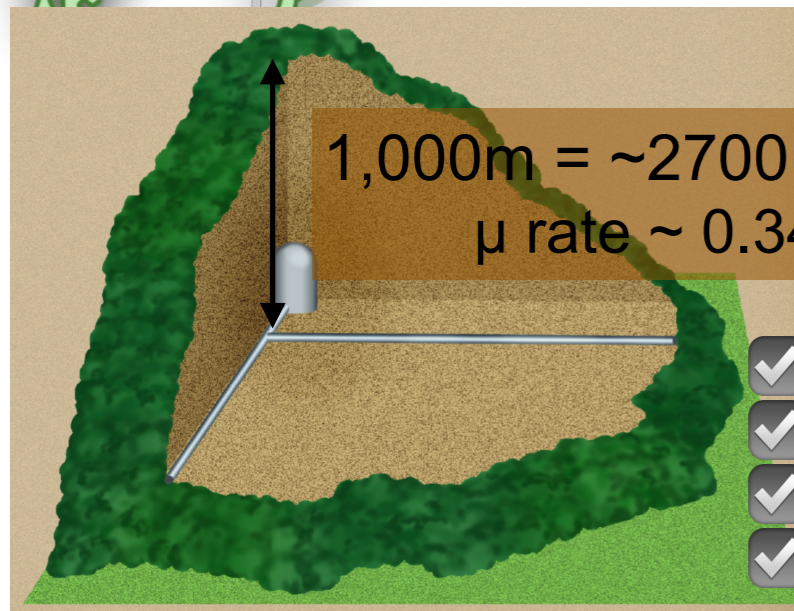
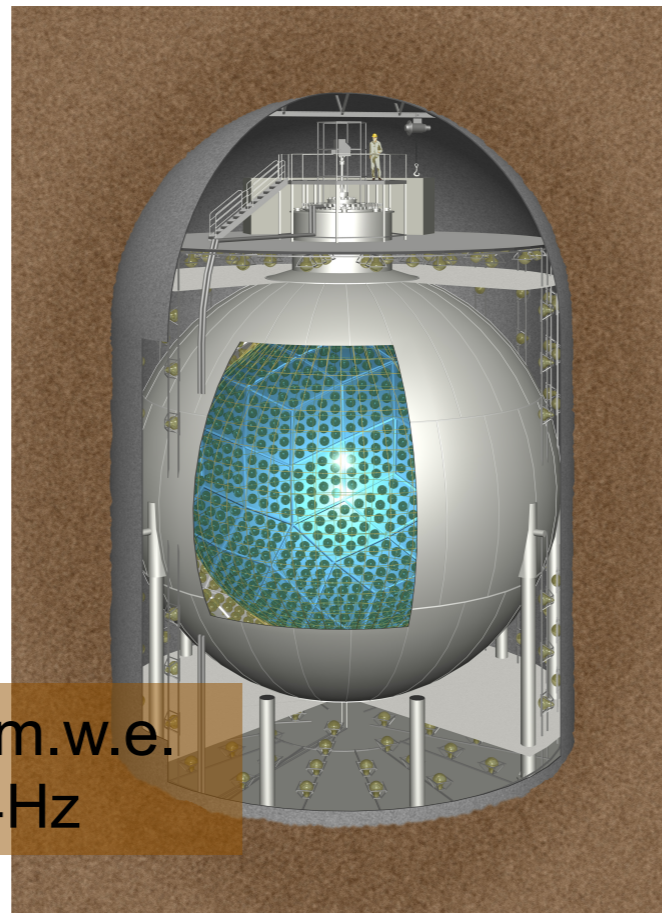
Research Center for Neutrino Science,
Tohoku University



43rd International Symposium on Physics in Collision
NCSR “Demokritos”, October 23, 2024

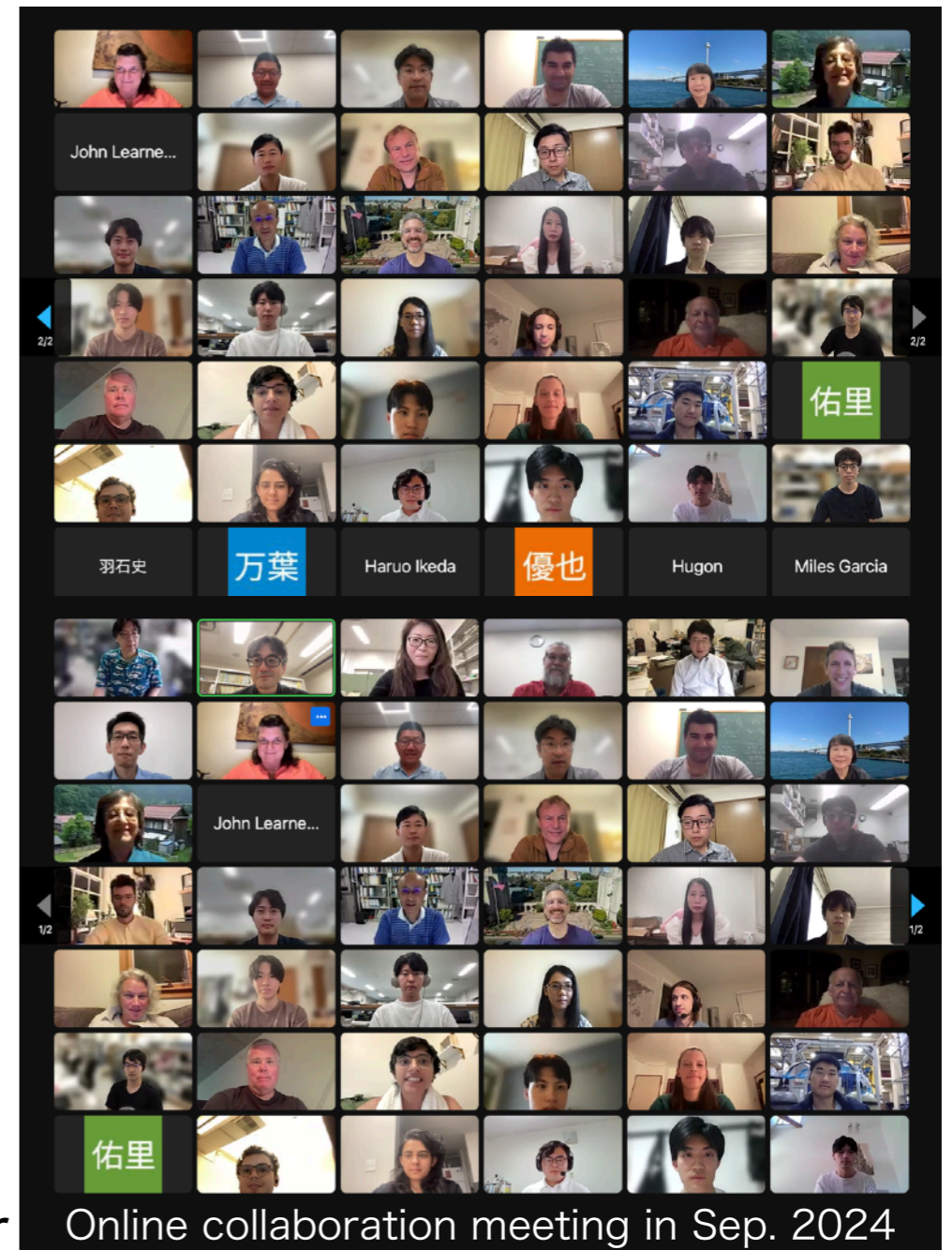
KamLAND/KamLAND-Zen Collaboration

~50 physicists work on this project



1,000m = ~2700 m.w.e.
 μ rate ~ 0.34Hz

- ✓ 1 kton liquid scintillator
- ✓ 2000 PMTs
- ✓ ϕ 13m balloon vessel
- ✓ Water Cherenkov outer detector



Online collaboration meeting in Sep. 2024



Neutrinoless Double beta decays

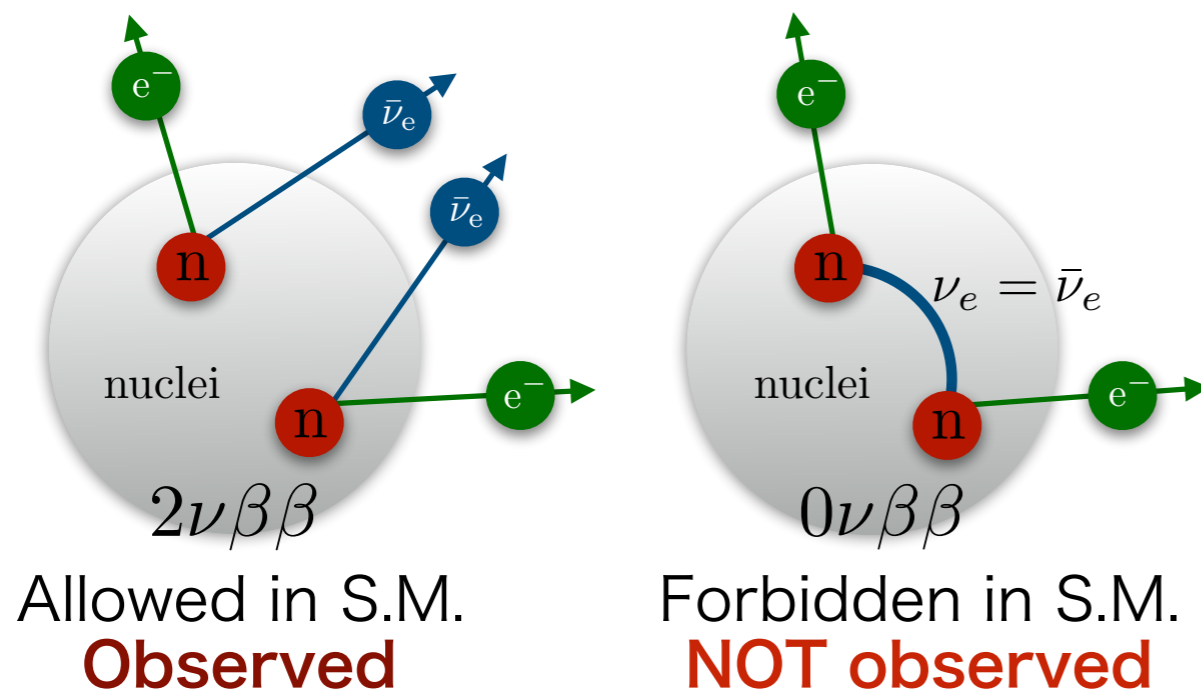
$0\nu\beta\beta$ decay search is the direct test of

Majorana nature ($\nu = \bar{\nu}$)

Needed by promising scenarios to explain the **matter dominant universe**

- ▶ see-saw mechanism
- ▶ Leptogenesis

▶ **Two dominant double beta decay mode**

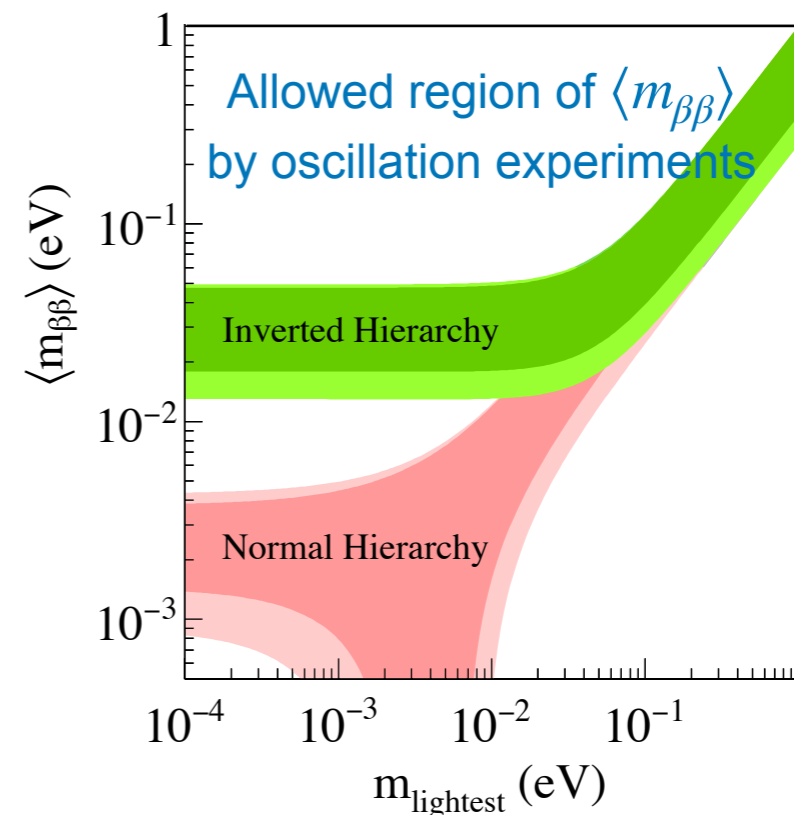


▶ **Access neutrino mass**

$$\text{Half life } (T_{1/2}^{0\nu})^{-1} = G^{0\nu} |M^{0\nu}|^2 \langle m_{\beta\beta} \rangle$$

Effective Majorana mass

Effective Majorana mass is one of the channels to access neutrino mass



Very long $0\nu\beta\beta$ decay half-life ($> 10^{18}$ yr) = Ultra rare events

→ **Large amount of isotopes & low BG environment** is needed.

Requirements for Experiments

- ▶ $0\nu\beta\beta$ decays are extremely rare events requiring special experimental design

1. Large amount of targets Merit of KamLAND-Zen

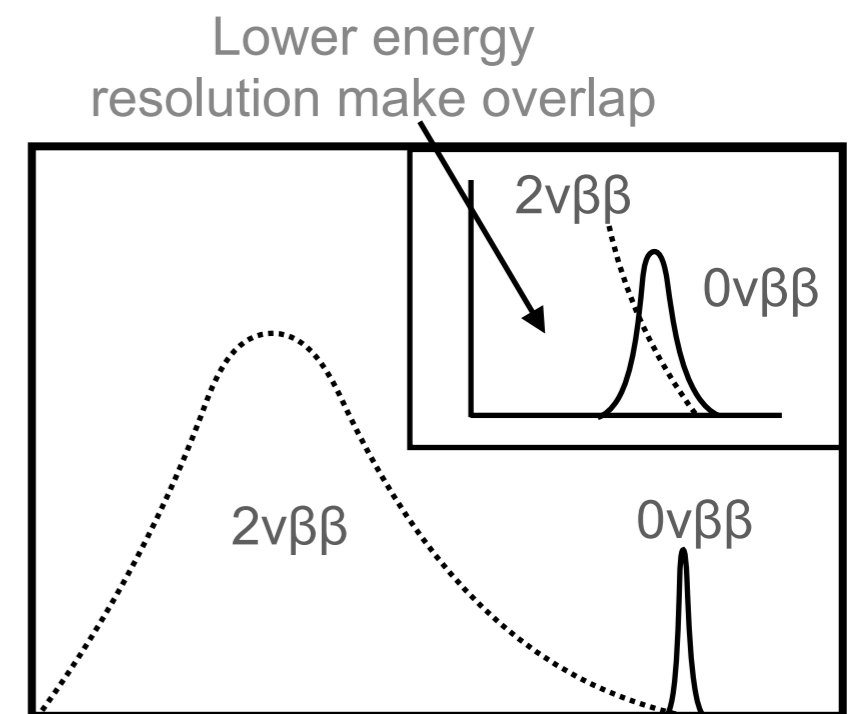
- Larger detector
- Isotopic enrichment

2. Super clean detector Merit of KamLAND-Zen

- Environmental radioactivity (^{238}U , ^{232}Th) in detector material should be reduced.
- Underground experiment to reduce cosmic muon rate

3. Energy resolution Demerit of KamLAND-Zen

- Observe total kinetic energy of 2 electrons.
- The larger overlap of $2\nu\beta\beta$ and $0\nu\beta\beta$ makes $2\nu\beta\beta$ a serious background.
- Ge detector has excellent energy resolution



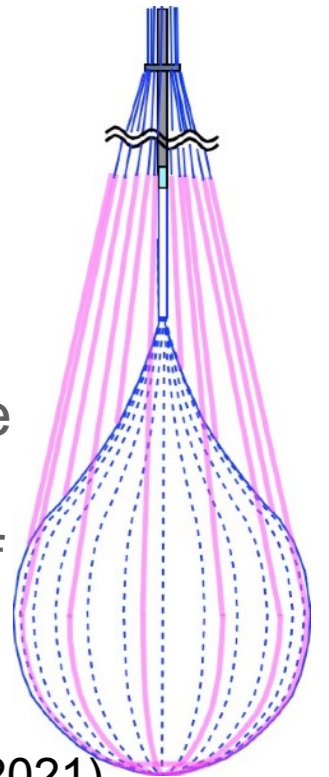
KamLAND-Zen

Double beta decay isotope: ^{136}Xe

- Q-value 2.458 MeV
- Enrichment ~90%
- Dissolved into LS ~3% by weight
 - Xe loaded liquid scintillator (Xe-LS)

Xe-LS container

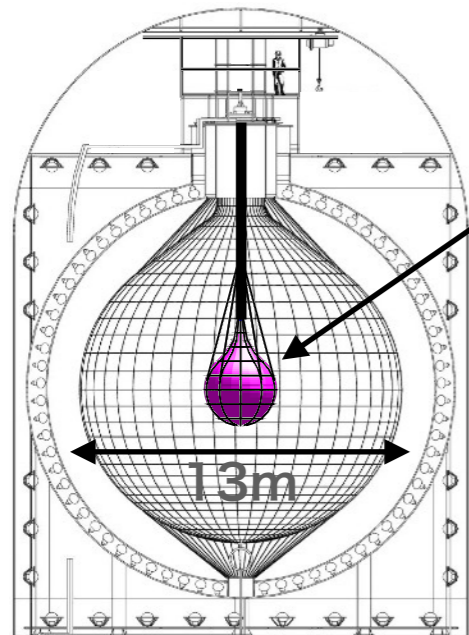
- ✓ thin(25 μm) nylon film balloon
- ϕ 3m for KLZ400
- ϕ 4m for KLZ800
- ▶ Xe is installed only within the container
 - Utilize cleanest volume of the large detector
- ➔ Ultra low BG is achieved



JINST 16 P08023 (2021)

Past KamLAND-Zen 400

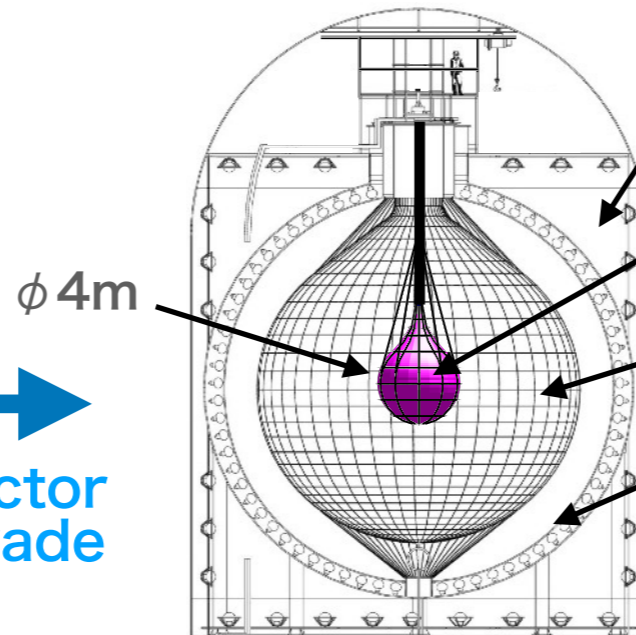
320-380 kg of Xenon
Data taking in 2011 - 2015



Present KamLAND-Zen 800

~750 kg of Xenon
DAQ started in 2019

➔
detector
upgrade



Water Cherenkov detector for active cosmic muon veto

Xe loaded liquid scintillator (Xe-LS)

Liquid scintillator (no Xe)

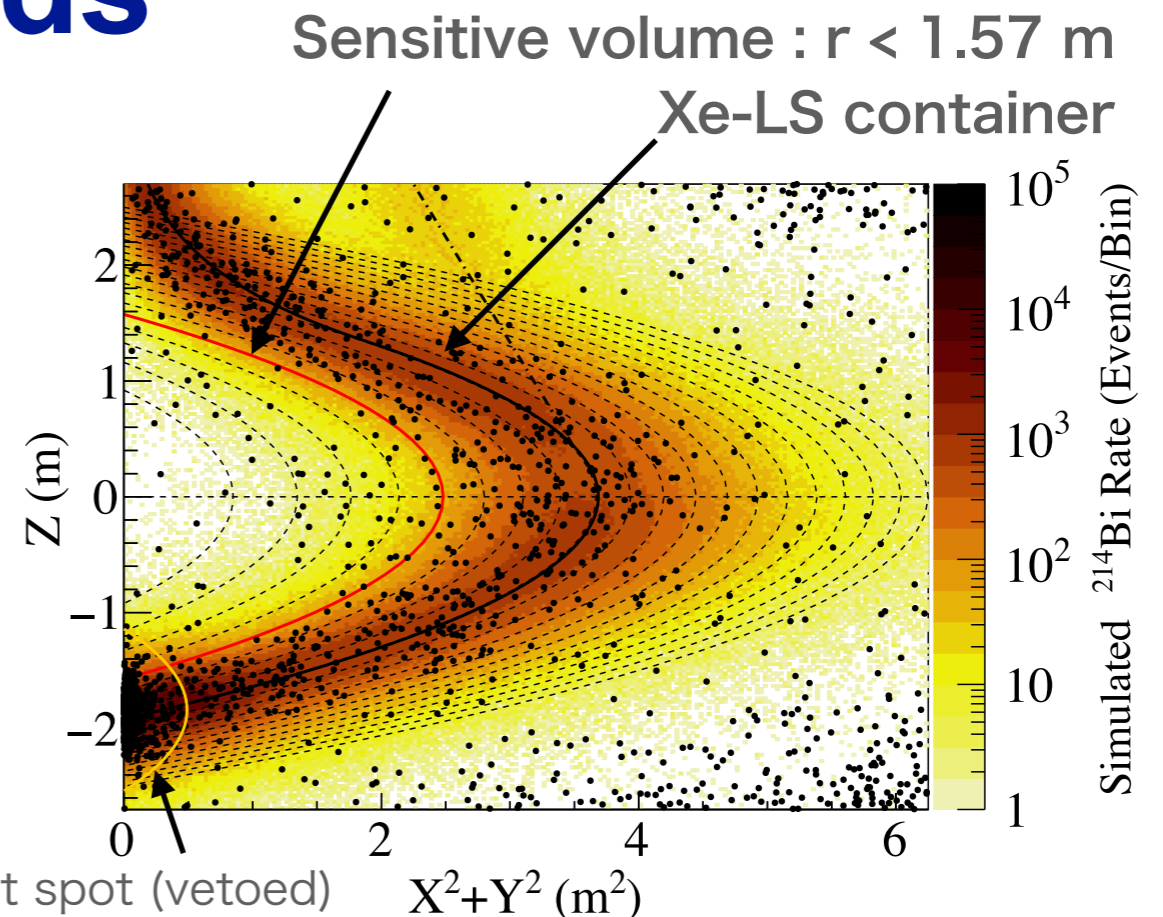
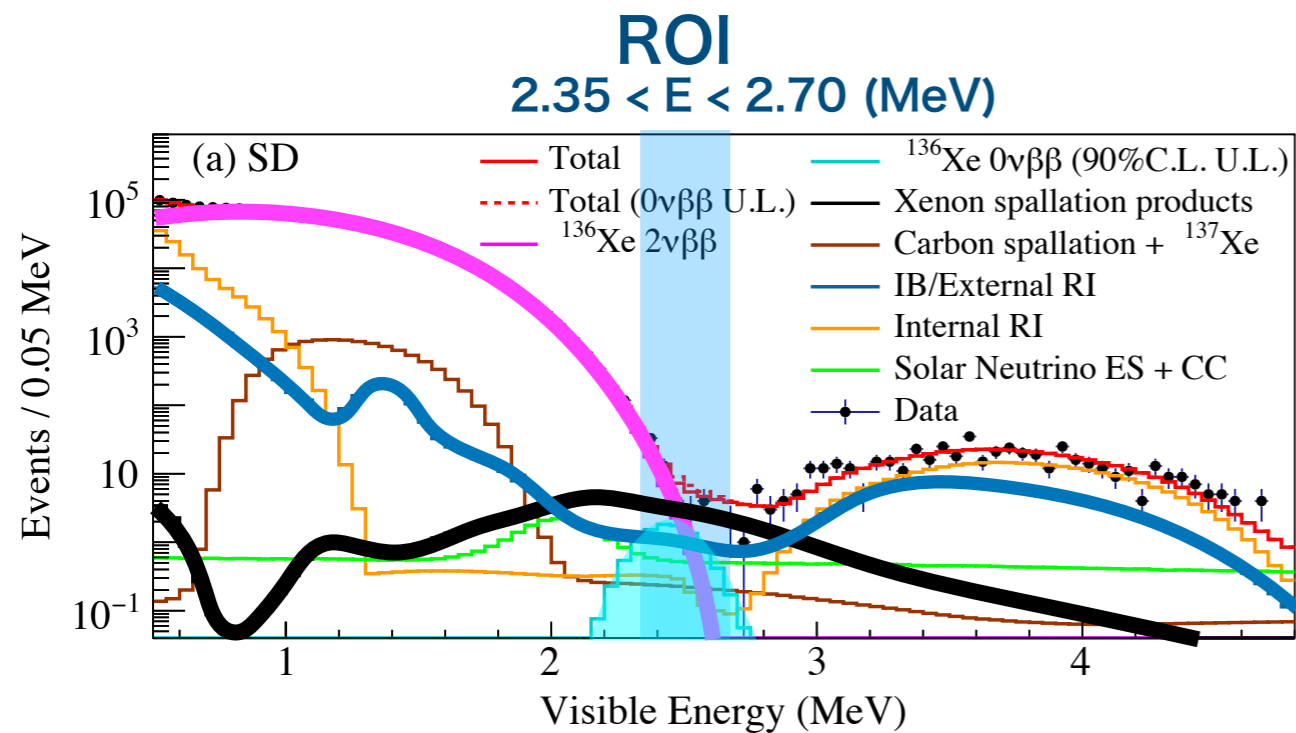
Geo-neutrinos, solar-neutrinos etc..

Event vertex and energy are reconstructed from time and charge of 1879 PMTs
 $6.7\% / \sqrt{E(\text{MeV})}, 13.7\text{cm} / \sqrt{E(\text{MeV})}$

KamLAND-Zen 800 completed DAQ on Jan. 11, 2024

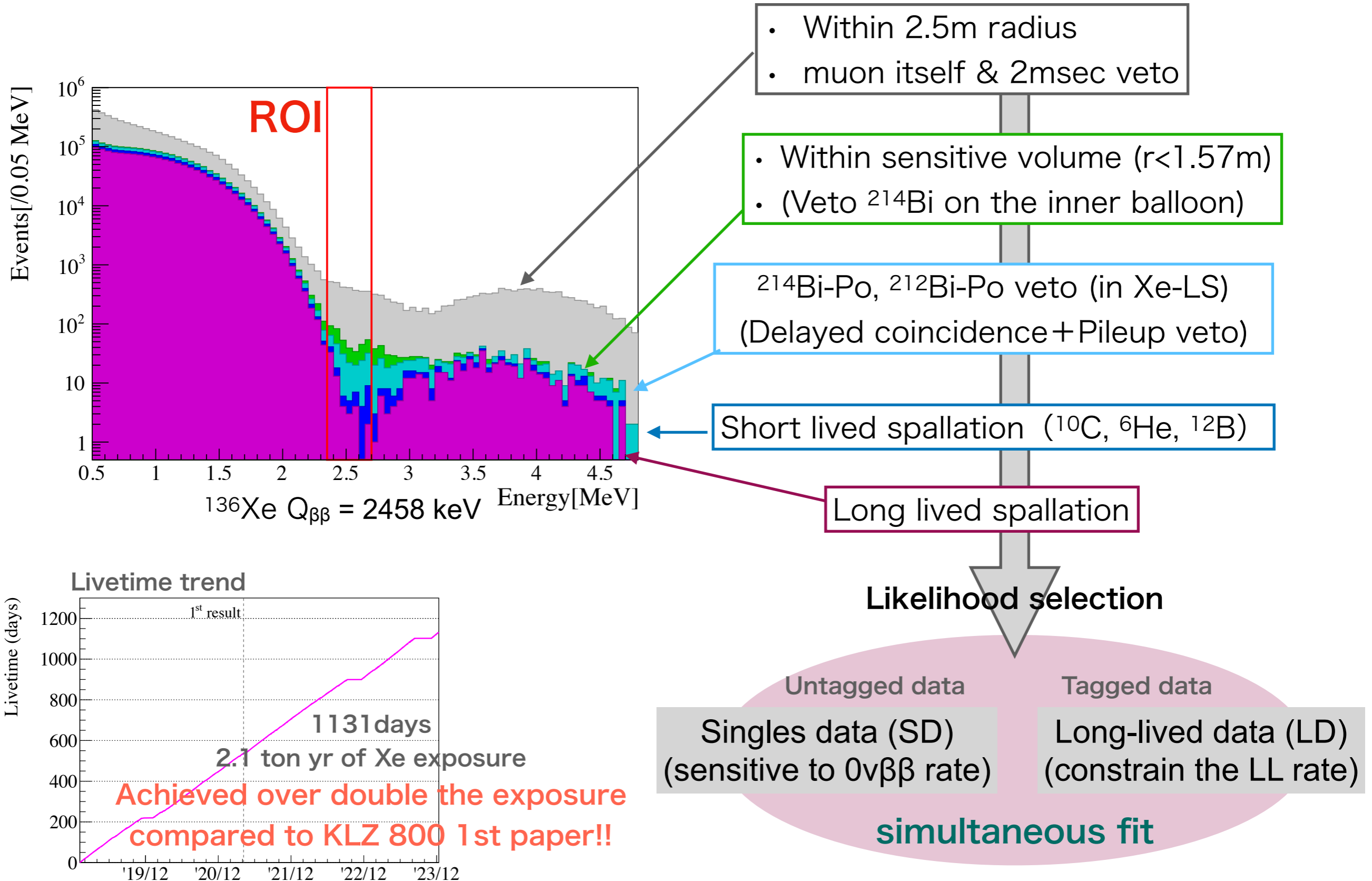
Dominant Backgrounds

2.35 < E < 2.70 MeV,
livetime 523.4 days



- **^{238}U , ^{232}Th on the balloon (XeLS container) film**
 - The delayed coincidence, $^{214}\text{Bi}(\beta) \rightarrow ^{214}\text{Po}(\alpha)$, does not effectively work on the film due to a quenching of α decays (delayed signal).
- **$2\nu\beta\beta$ decays**
 - Inevitable background source
 - Enhancing detector energy resolution will reduce $2\nu\beta\beta$ in ROI.
- **Cosmic muon induced Xe spallation products (Long-lived products)**
 - A few hours or a few days life-time isotopes are difficult to tag by the delayed coincidence or simple volume cut. Phys. Rev. C **107**, 054612 (2023)

Event Selection, livetime



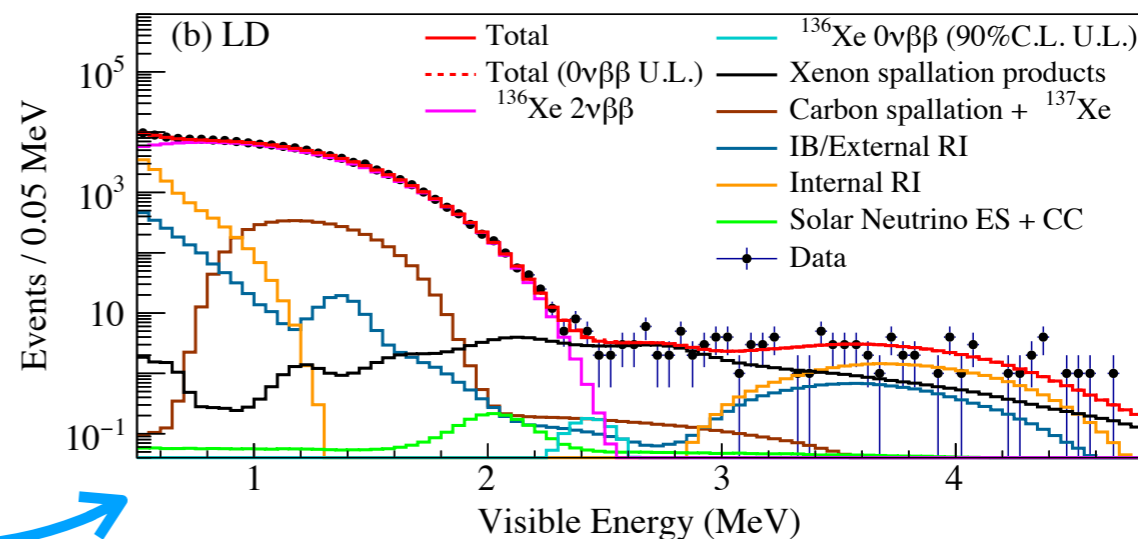
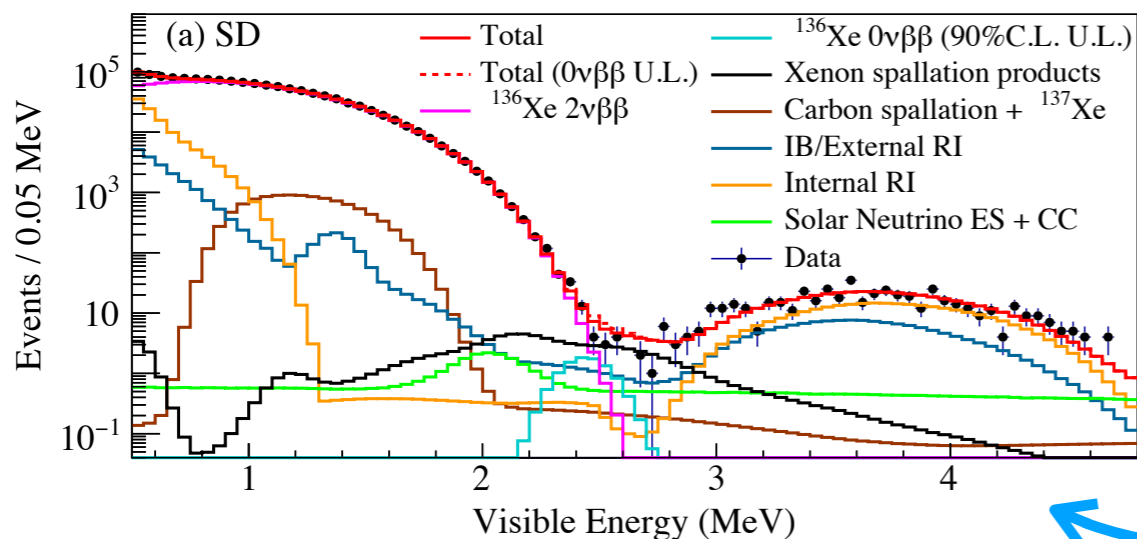
^{136}Xe $0\nu\beta\beta$ decay Half-life limit arXiv:2406.11438[hep-ex]

Singles data (SD)
(sensitive to $0\nu\beta\beta$ rate)

Livetime = 1131 days

Long-lived product data (LD)
(used to constrain the LL rate)

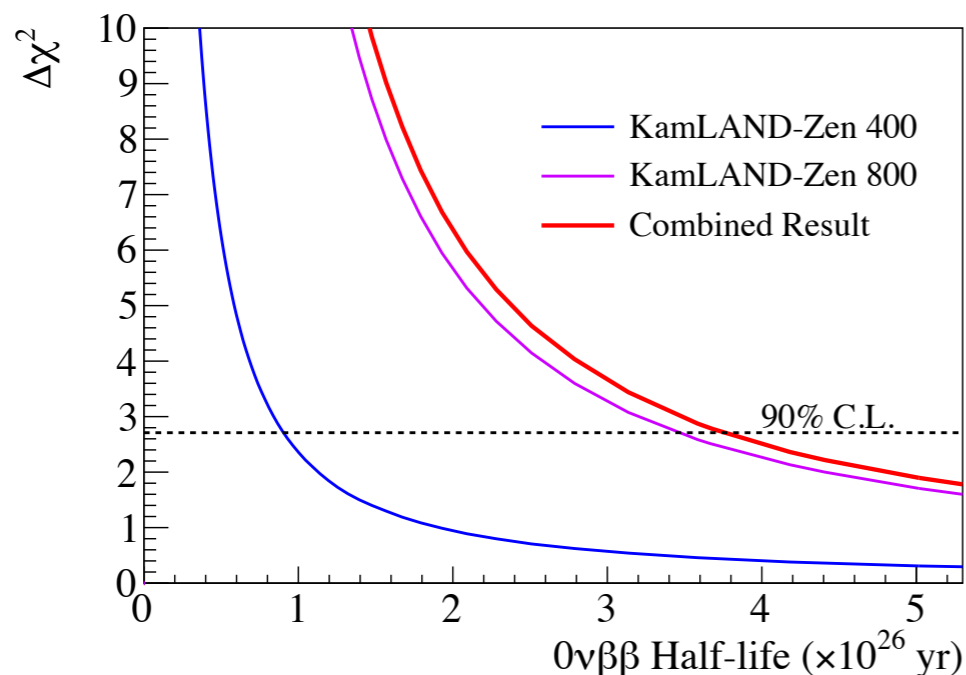
Livetime = 111.0 days



simultaneous fit

KamLAND-Zen 800 only result : $T_{1/2}^{0\nu} > 3.4 \times 10^{26}$ yr (90% C.L.)

► Combined analysis (KamLAND-Zen 400 + 800)



KLZ 400 only : 0.9×10^{26} yr

Combined result (90% C.L.):

$$T_{1/2}^{0\nu} > 3.8 \times 10^{26} \text{ yr}$$

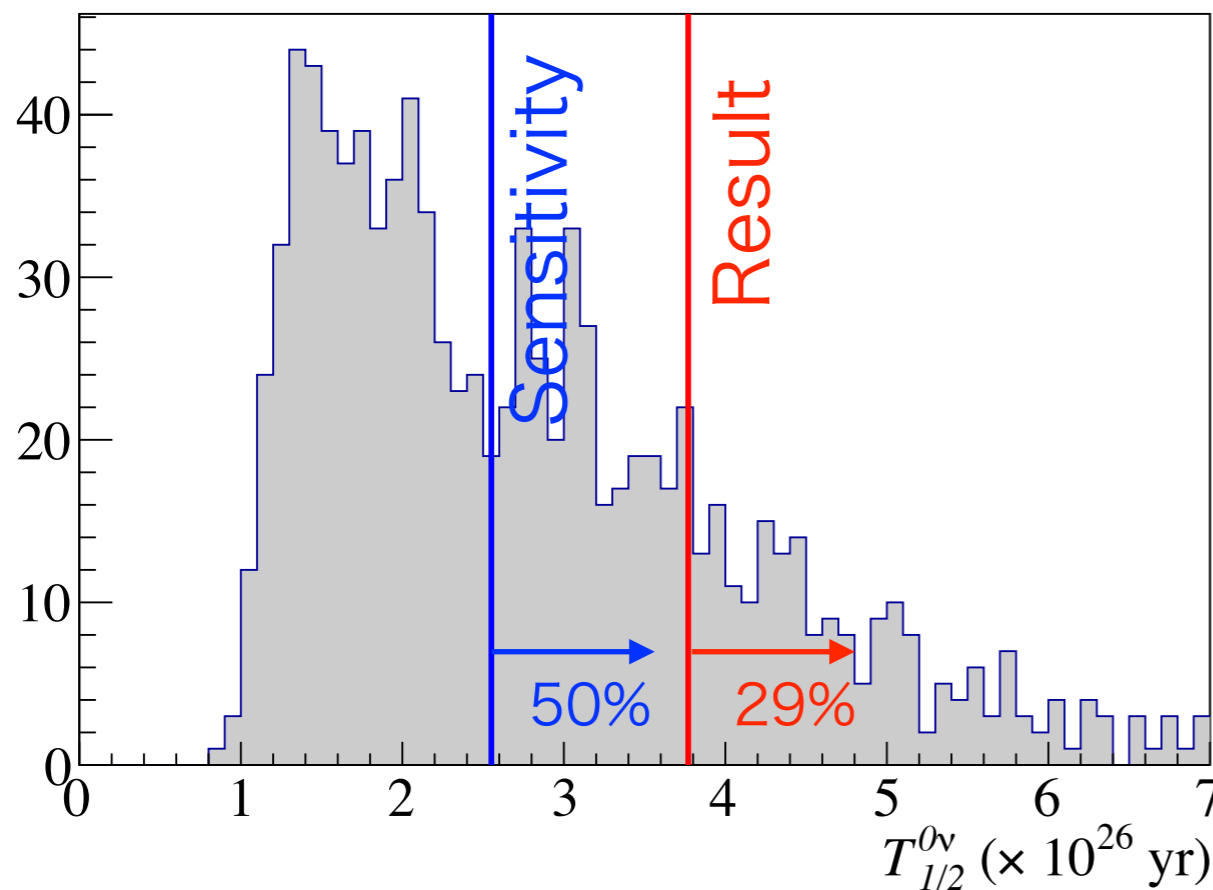
Sensitivity (90% C.L.)

$$T_{1/2}^{0\nu} = 2.6 \times 10^{26} \text{ yr}$$

Upper limits from Toy MC

To evaluate the validity of the fit results, the experimental sensitivity was assessed.

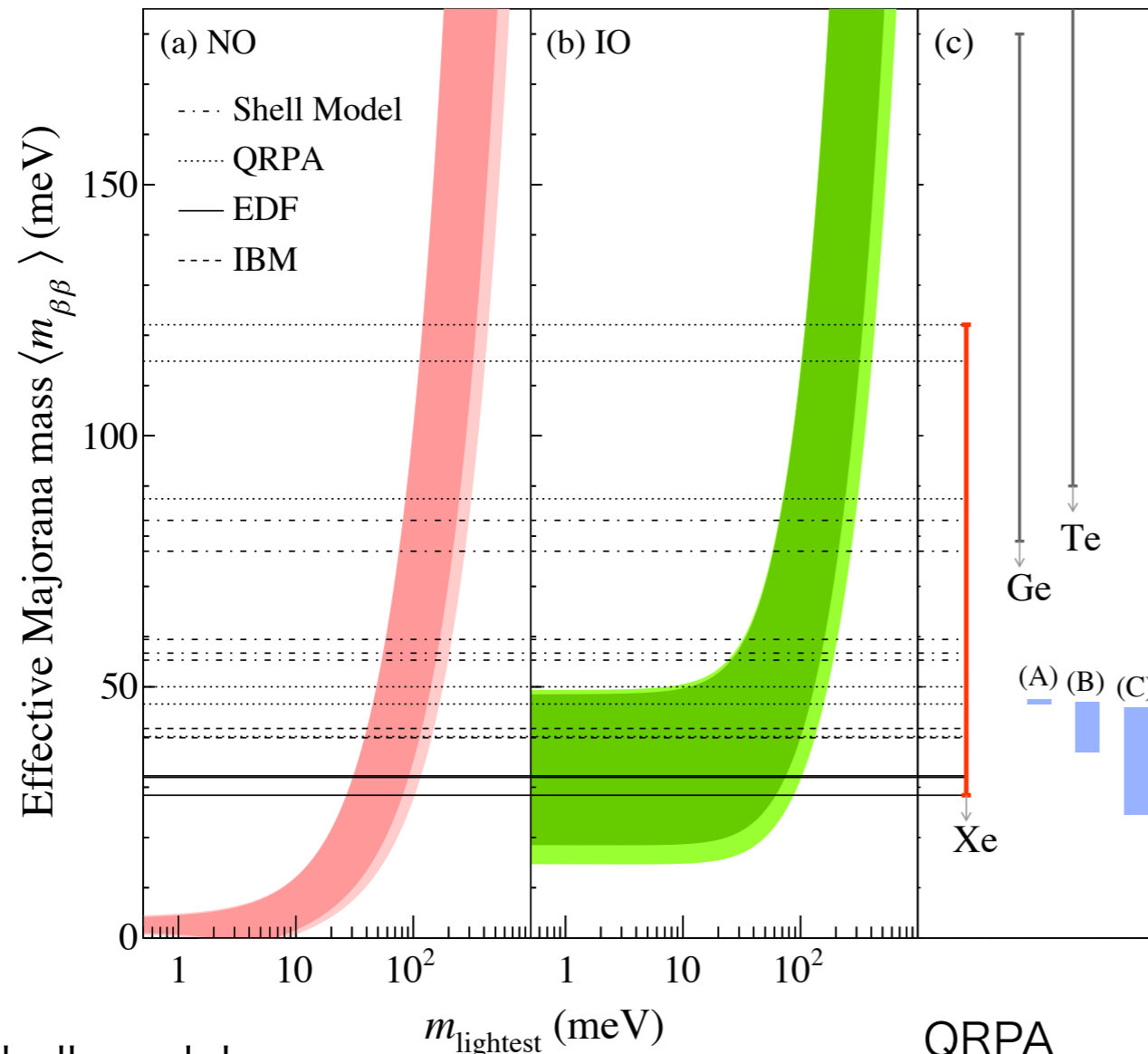
- Generated 1,500 ToyMC datasets assuming a Poisson distribution based on the best-fit model.
- Calculated the 90% confidence level limit for each of the 1,500 MC datasets.
- Defined the median of the 1,500 trials as the experimental sensitivity (Median Sensitivity).



Sensitivity: $T_{1/2}^{0\nu} > 2.6 \times 10^{26} \text{ yr}$ (90% C.L.)

Result: $T_{1/2}^{0\nu} > 3.8 \times 10^{26} \text{ yr}$ (90% C.L.)

Limit on the effective Majorana mass



Lower limit of half life

$$T_{1/2}^{0\nu} > 3.8 \times 10^{26} \text{ yr}$$

$$\left(T_{1/2}^{0\nu}\right)^{-1} = G^{0\nu} |M^{0\nu}|^2 \langle m_{\nu} \rangle^2$$

NME($M^{0\nu}$): 1.11-4.77
assuming $g_A \sim 1.27$

Upper limit of Majorana mass

$$\langle m_{\beta\beta} \rangle < 28-122 \text{ meV}$$

Most stringent test of the
neutrino mass in the IO

Shell model

- [1] J. Menéndez, *J. of Phys. G* **45**, 014003 (2018).
- [2] M. Horoi and A. Neacsu, *Phys. Rev. C* **93**, 024308 (2016).
- [3] L. Coraggio, A. Gargano, N. Itaco, R. Mancino, and F. Nowacki, *Phys. Rev. C* **101**, 044315 (2020).
- [4] L. Coraggio *et al.*, *Phys. Rev. C* **105**, 034312 (2022).

EDF

- [10] T. R. Rodríguez and G. Martínez-Pinedo, *Phys. Rev. Lett.* **105**, 252503 (2010).
- [11] N. L. Vaquero, T. R. Rodríguez, and J. L. Egido, *Phys. Rev. Lett.* **111**, 142501 (2013).
- [12] L. S. Song, J. M. Yao, P. Ring, and J. Meng, *Phys. Rev. C* **95**, 024305 (2017).

QRPA

- [5] M. T. Mustonen and J. Engel, *Phys. Rev. C* **87**, 064302 (2013).
- [6] J. Hyvärinen and J. Suhonen, *Phys. Rev. C* **91**, 024613 (2015).
- [7] F. Šimkovic, A. Smetana, and P. Vogel, *Phys. Rev. C* **98**, 064325 (2018).
- [8] D.-L. Fang, A. Faessler, and F. Šimkovic, *Phys. Rev. C* **97**, 045503 (2018).
- [9] J. Terasaki, *Phys. Rev. C* **102**, 044303 (2020).

IBM

- [13] J. Barea, J. Kotila, and F. Iachello, *Phys. Rev. C* **91**, 034304 (2015).
- [14] F. F. Deppisch, L. Graf, F. Iachello, and J. Kotila, *Phys. Rev. D* **102**, 095016 (2020).

Theoretical model

- (A) K. Harigaya, M. Ibe, and T. T. Yanagida, *Phys. Rev. D* **86**, 013002 (2012)
- (B) T. Asaka, Y. Heo, and T. Yoshida, *Phys. Lett. B* **811**, 135956 (2020).
- (C) K. Asai, *Eur. Phys. J. C* **80**, 76 (2020)

Limit on the effective Majorana mass

	Ref.	$M^{0\nu}$	$\langle m_{\beta\beta} \rangle$ (meV)
Shell model	[1]	2.28, 2.45	59.4, 55.3
	[2]	1.63, 1.76	83.1, 77.0
	[3, 4]	2.39	56.7
QRPA	[5]	1.55	87.4
	[6]	2.91	46.6
	[7]	2.71	50.0
	[8]	1.11, 1.18	122, 115
	[9]	3.38	40.1
EDF theory	[10]	4.20	32.3
	[11]	4.77	28.4
	[12]	4.24	32.0
IBM	[13]	3.25	41.7
	[14]	3.40	39.9

 : NMEs in IO (< 50 meV)

Lower limit of half life

$$T_{1/2}^{0\nu} > 3.8 \times 10^{26} \text{ yr}$$

$$\left(T_{1/2}^{0\nu}\right)^{-1} = G^{0\nu} |M^{0\nu}|^2 \langle m_{\nu} \rangle^2$$

NME($M^{0\nu}$): 1.11-4.77
assuming $g_A \sim 1.27$

Upper limit of Majorana mass

$$\langle m_{\beta\beta} \rangle < 28-122 \text{ meV}$$

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- [3] L. Coraggio, A. Gargano, N. Itaco, R. Mancino, and F. Nowacki, *Phys. Rev. C* **101**, 044315 (2020).
- [4] L. Coraggio *et al.*, *Phys. Rev. C* **105**, 034312 (2022).

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- [11] N. L. Vaquero, T. R. Rodríguez, and J. L. Egido, *Phys. Rev. Lett.* **111**, 142501 (2013).
- [12] L. S. Song, J. M. Yao, P. Ring, and J. Meng, *Phys. Rev. C* **95**, 024305 (2017).

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- [6] J. Hyvärinen and J. Suhonen, *Phys. Rev. C* **91**, 024613 (2015).
- [7] F. Šimkovic, A. Smetana, and P. Vogel, *Phys. Rev. C* **98**, 064325 (2018).
- [8] D.-L. Fang, A. Faessler, and F. Šimkovic, *Phys. Rev. C* **97**, 045503 (2018).
- [9] J. Terasaki, *Phys. Rev. C* **102**, 044303 (2020).

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- [13] J. Barea, J. Kotila, and F. Iachello, *Phys. Rev. C* **91**, 034304 (2015).
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- (B) T. Asaka, Y. Heo, and T. Yoshida, *Phys. Lett. B* **811**, 135956 (2020).
- (C) K. Asai, *Eur. Phys. J. C* **80**, 76 (2020)

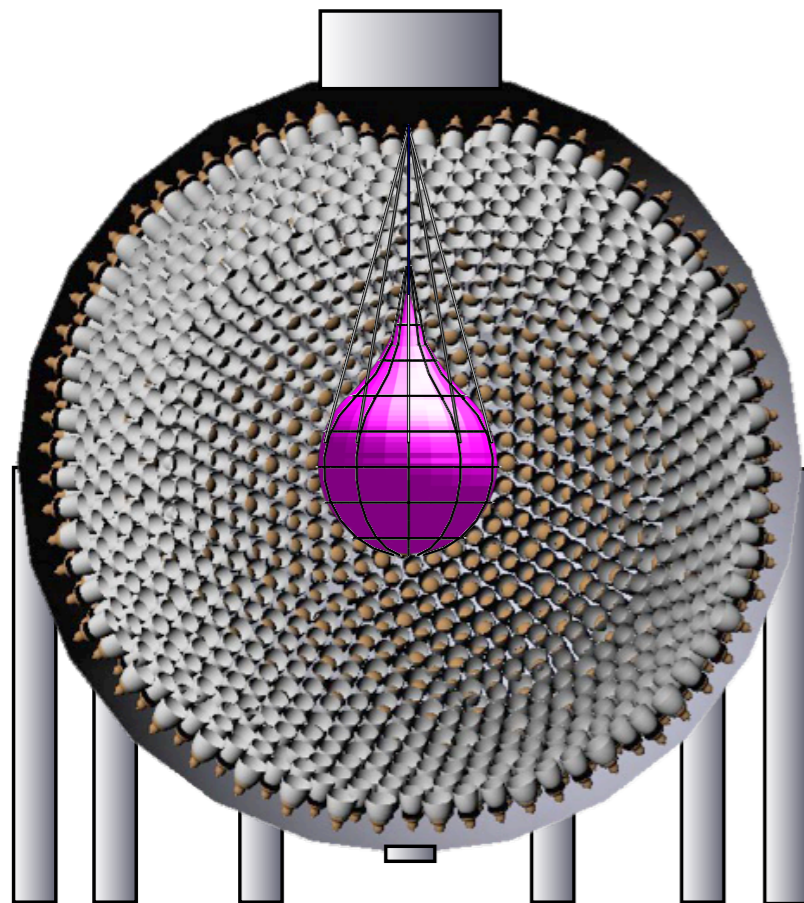
KamLAND2-Zen

Next generation detector R&D is now in progress

Film BG reduction

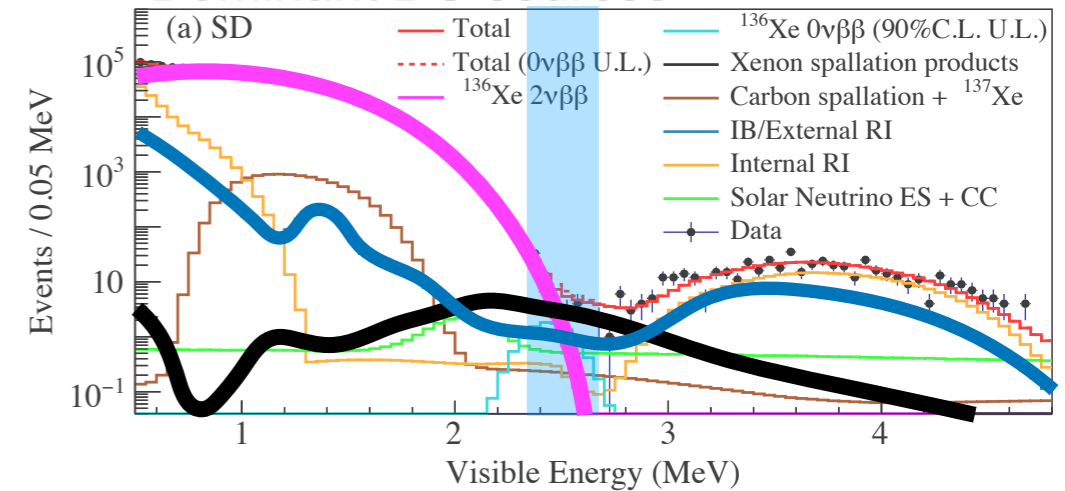
Scintillation balloon film

Enlarge sensitive volume by tagging $^{214}\text{Bi}(\beta) \rightarrow ^{214}\text{Po}(\alpha)$ sequential decay on the film.



1000 kg of Xe
Increase isotope mass

Dominant BG sources



$2\nu\beta\beta$ decay

Enhance detector energy resolution ($\sigma(2.6\text{MeV})4\% \rightarrow \sim 2\%$)

High QE PMT & Light guide mirror

Improve light collection efficiency and photo coverage

Brighter LS

Higher light yield and transmittance

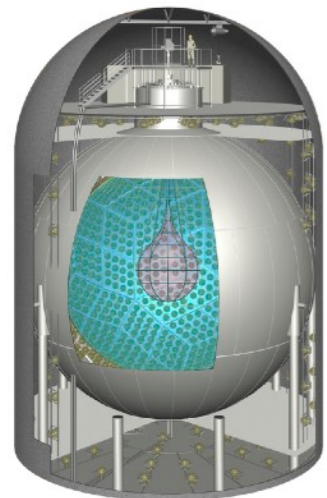
Cosmic muon induced Xe spallation products

New Dead-time free electronics

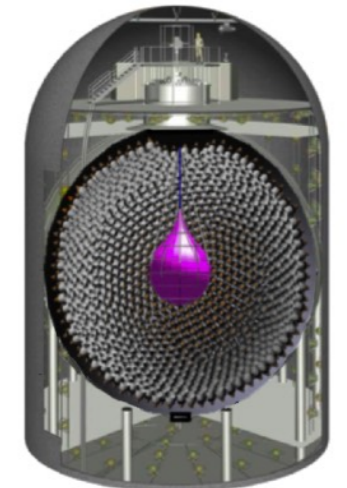
Collect all the neutron information (multiplicity, vertices) from the noisy period

Target $\langle m_{\beta\beta} \rangle \sim 20 \text{ meV}$ (5 yrs)

Plan of Detector Upgrades



KamLAND & KamLAND-Zen



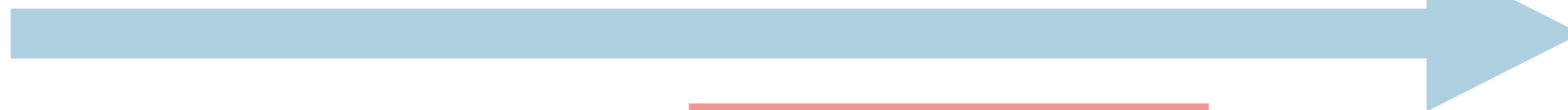
KamLAND2 & KamLAND2-Zen

2024

2025

2026

2027



High QE PMT & Mirror installation

Cleaner Outer balloon

High light yield LS

New electronics

Calibration system

Xe & Inner Balloon extraction

LS extraction

Super clean facility

Outer balloon & PMT dismantling

Scintillation balloon

Xe LS

KamLAND2-Zen

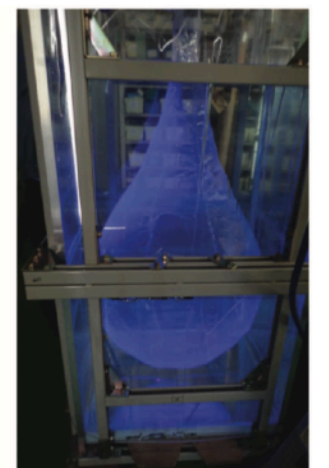
Done!!

here!!

KamLAND Dismantling

KamLAND2 construction

We started detector upgrades in this year and plan to launch KamLAND2 in 2027!!

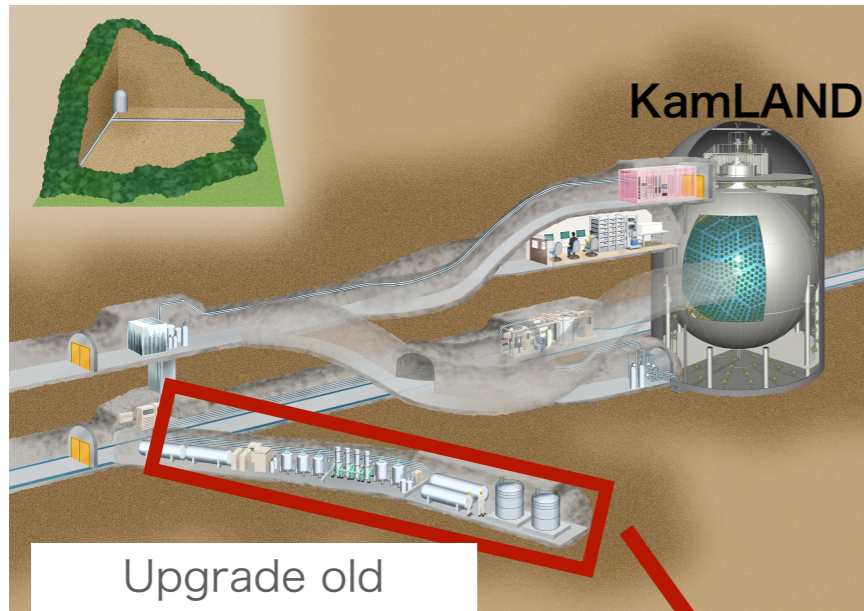


KERNEL

神岡極稀現象研究拠点

Kamioka Extremely Rare phenomena and NEutrino research Lab

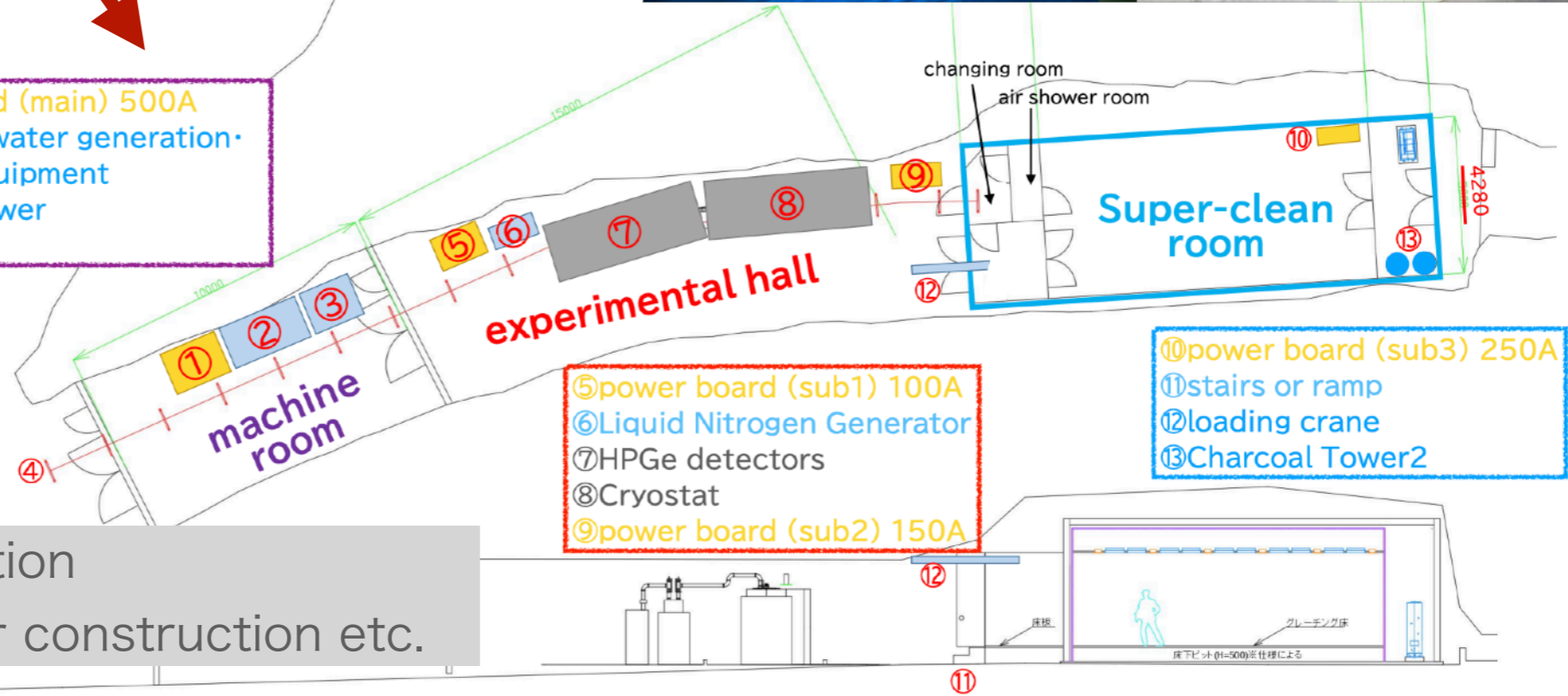
We are now constructing super clean facility in Kamioka mine



Upgrade old purification area!!



- ①power board (main) 500A
- ②Ultra-pure water generation-degassing equipment
- ③Charcoal tower
- ④Cable tray



- ⑤power board (sub1) 100A
- ⑥Liquid Nitrogen Generator
- ⑦HPGe detectors
- ⑧Cryostat
- ⑨power board (sub2) 150A

- ⑩power board (sub3) 250A
- ⑪stairs or ramp
- ⑫loading crane
- ⑬Charcoal Tower2

- Inner balloon construction
- crystallization, detector construction etc.

Summary

- ✓ KamLAND-Zen searches for neutrinoless double beta decay with ^{136}Xe loaded liquid scintillator.

- ▶ Combined result for KamLAND-Zen 400 + 800

$$T_{1/2}^{0\nu} > 3.8 \times 10^{26} \text{ yr}, \quad \langle m_{\beta\beta} \rangle < 28 - 122 \text{ meV}$$

- ✓ We started detector upgrade toward KamLAND2/KamLAND2-Zen.

- Expand sensitive volume (scintillation balloon)
- Enhanced energy resolution (HQEPMT, Light guide)
- Enhanced tagging efficiency of Xe fission products (New electronics)

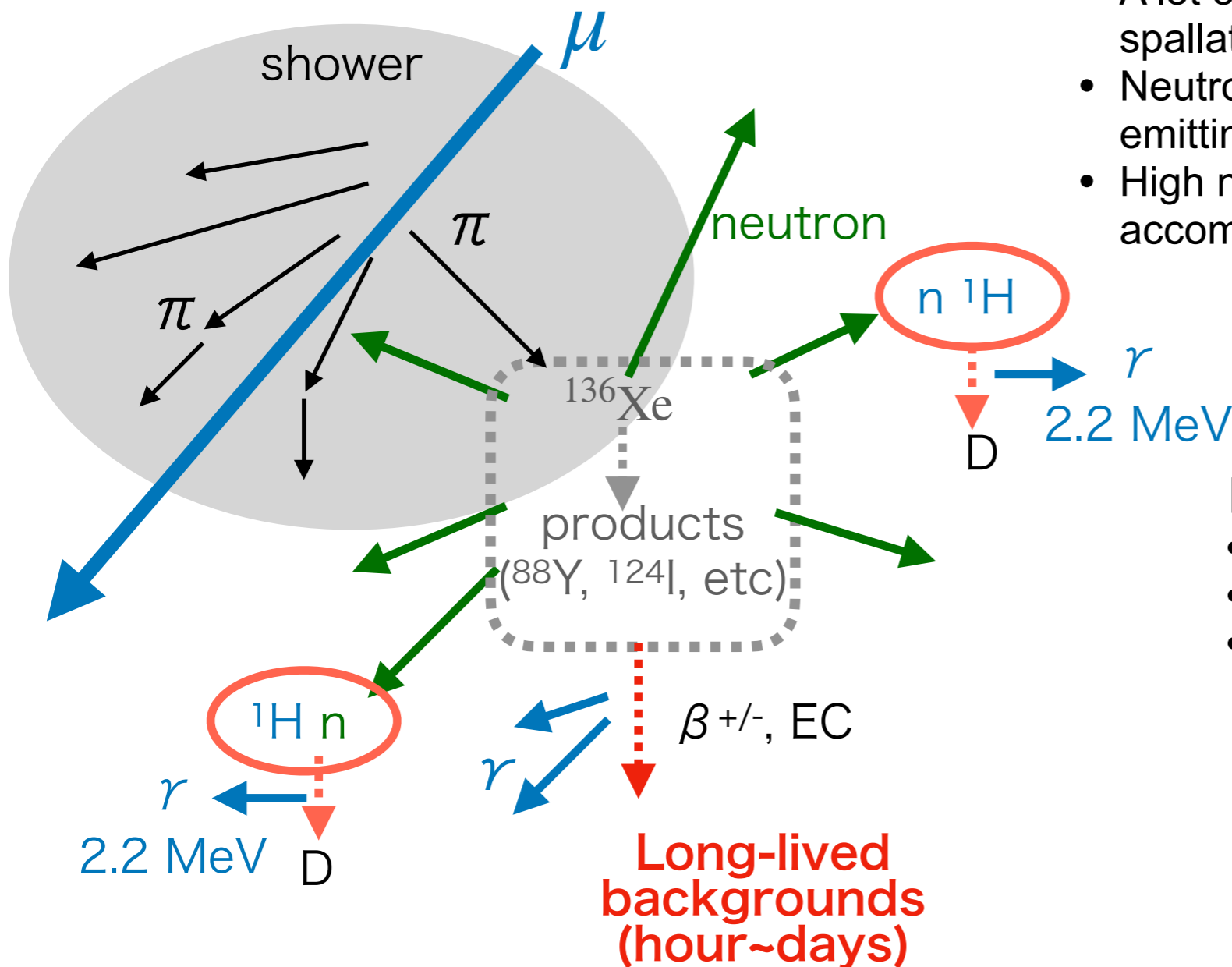
- ✓ The target of KL2-Zen is $\langle m_{\beta\beta} \rangle = 20 \text{ meV}$

Backup

Cosmic muon induced backgrounds

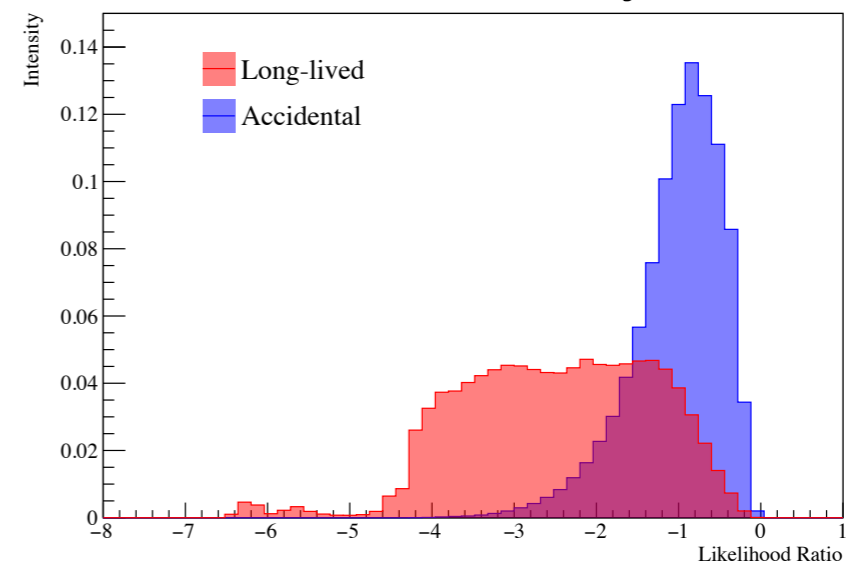
Neutrons give us information

- A lot of neutrons are produced at the same time of spallation products.
- Neutrons are immediately ($207\mu\text{s}$) captured by ^1H emitting 2.2 MeV gamma in LS.
- High neutron multiplicity events are likely to be accompanied by spallation products.



Likelihood selection based on,

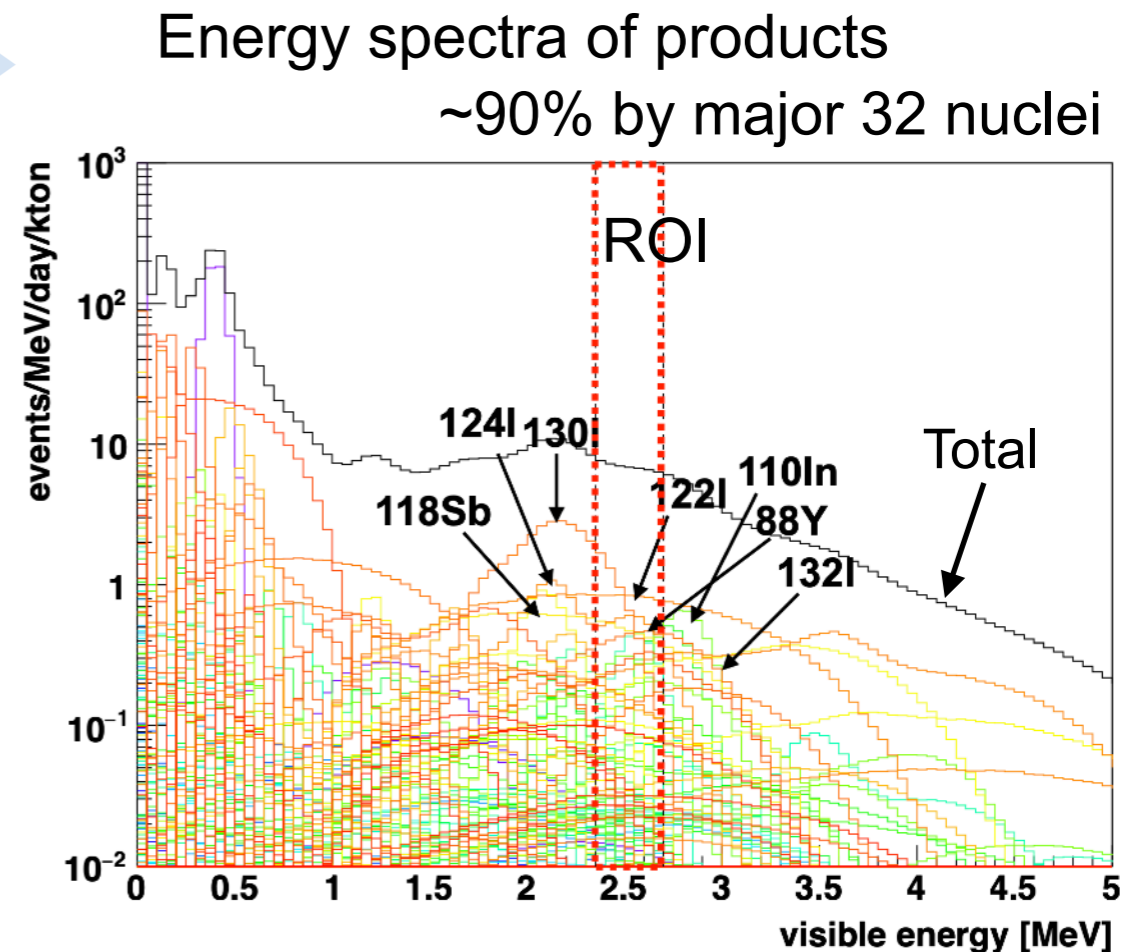
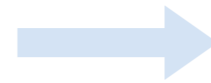
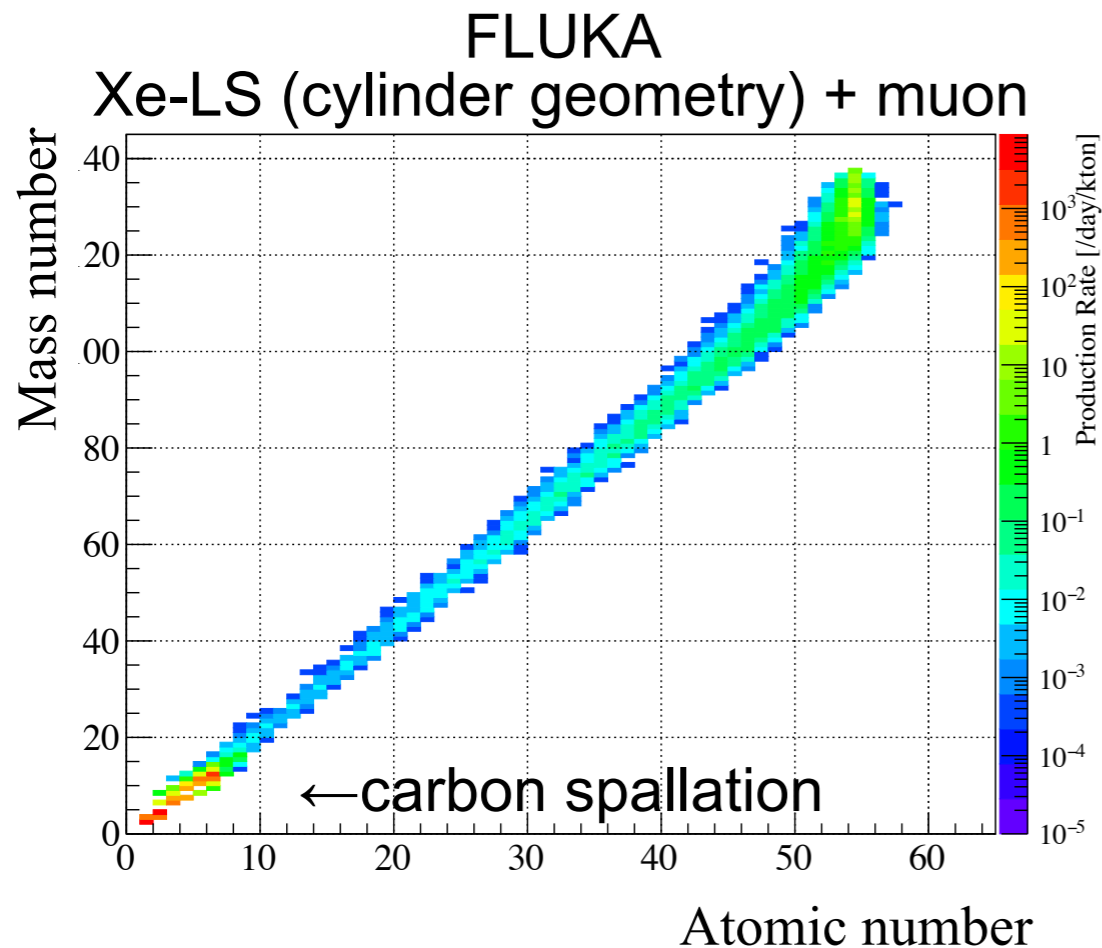
- Neutron multiplicity
- Distance from neutron capture
- Delta-time from muon injection



Decay of this long-lived spallation products make serious background in KamLAND-Zen

Xenon spallation products (Long-lived products)

- ✓ Individual yields are small but many candidates are produced
- ✓ Total yield become one of the main background → new major background



- Longer half-lives (~hours to ~days)
- Neutron multiplicity is higher than carbon's

A likelihood method is developed

Parameters: Time difference from muon, distance between Xe-spallation and neutron capture gamma, effective number of neutron

Rejection efficiency
42.0±8.8%

This rejected data-set is also used
for simultaneous fitting (next page)

Scintillation inner balloon

- BG(^{214}Bi) reduction from Xe-LS container

Tag this decay with scintillation inner balloon

PTEP. Volume 2019, Issue 7, 073H01, S. Obara et al.

- Current background level

$$^{238}\text{U} \sim 3 \times 10^{-12} \text{ g/g}_{\text{film}}$$

ref. initial film (after washed)

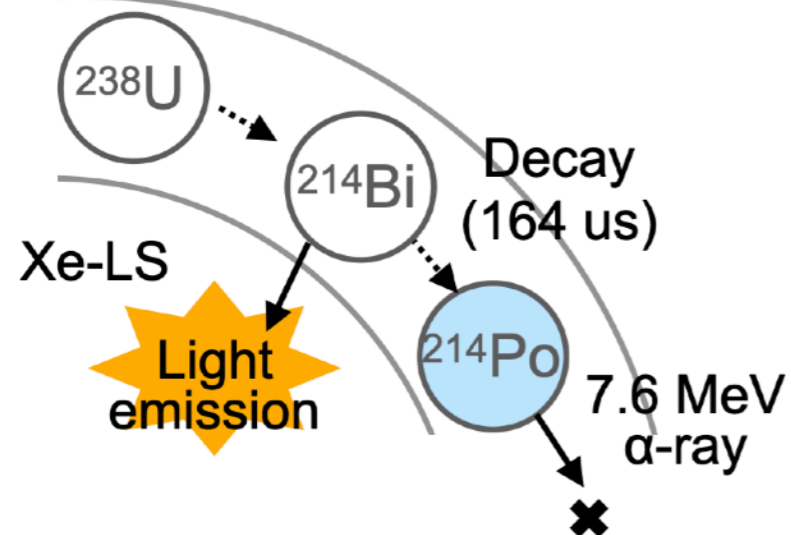
$$^{238}\text{U} \sim 2 \times 10^{-12} \text{ g/g}_{\text{film}}$$

Almost same level

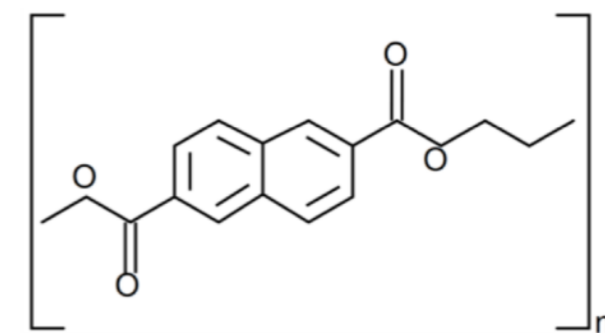
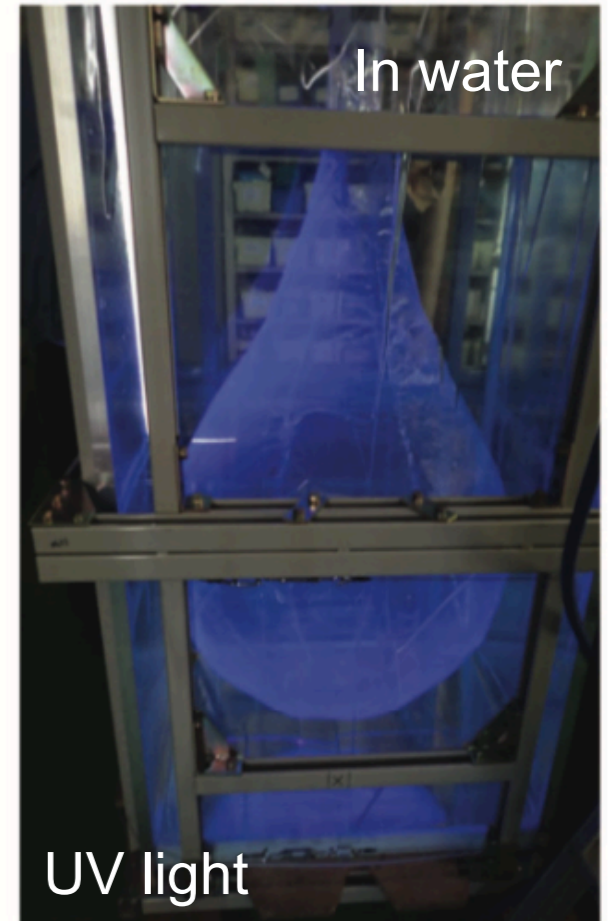
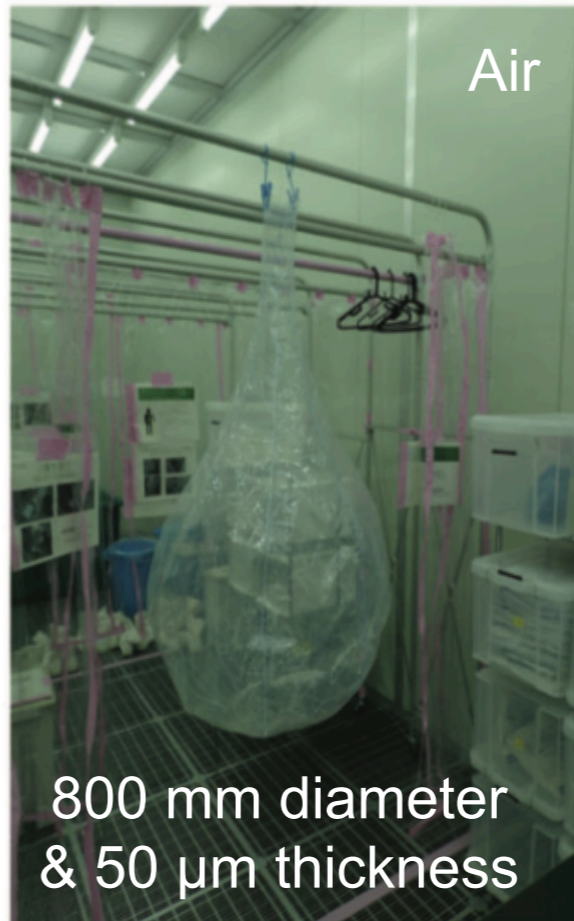
Difficult for further improvement

- Current tagging efficiency
~50% due to ^{214}Po alpha decay

Xe-LS container



cf) Xe-LS tagging eff. 99.97%



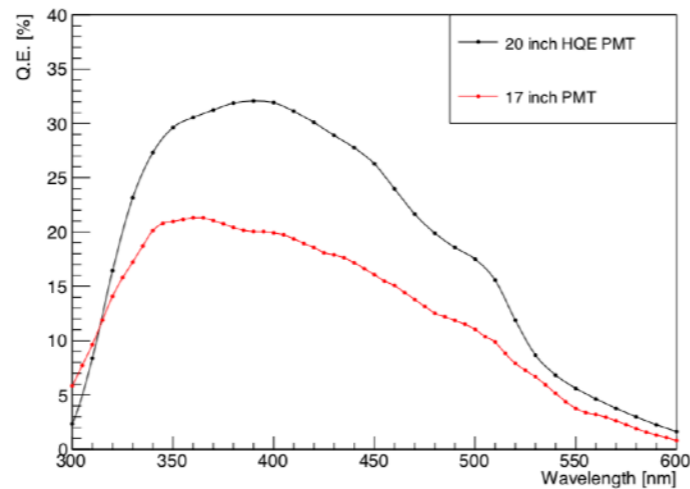
PolyEthylene Naphthalate (PEN)

High QE PMT & light guide

20 inch PMT (R12860-03LXA)



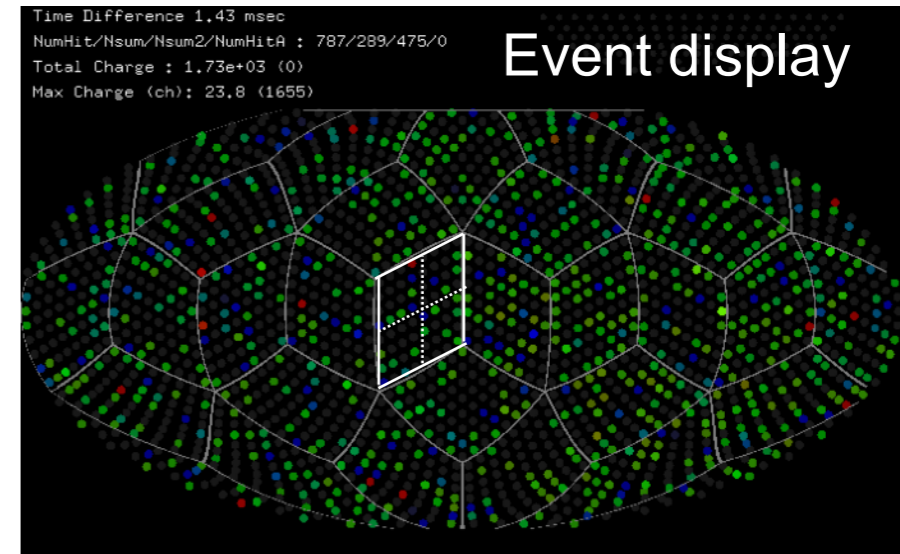
High quantum efficiency



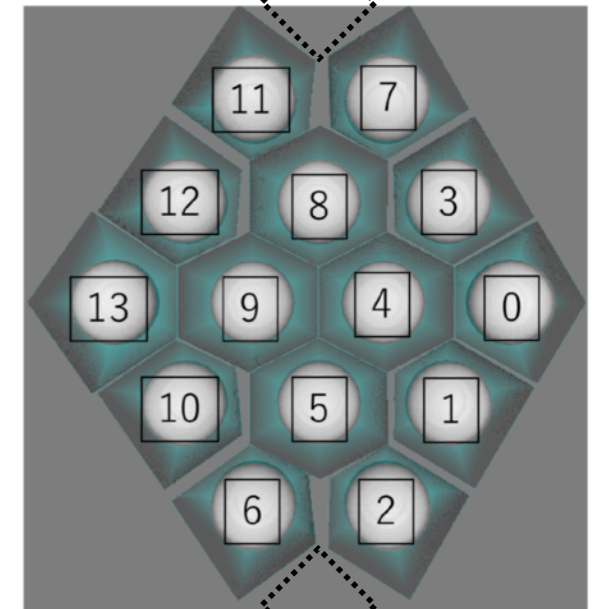
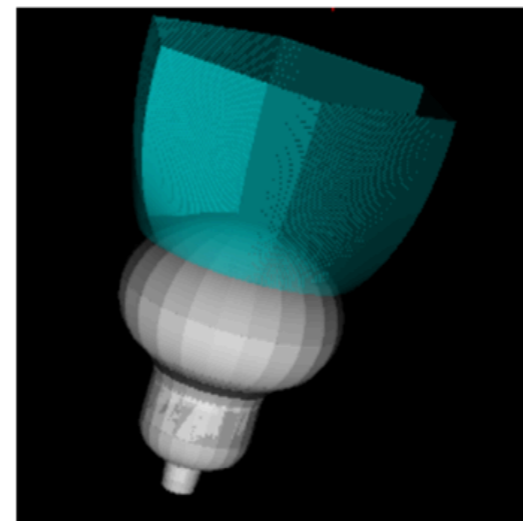
PMT spec:

	New 20"	17"
Dynode	Box & line	Line-focus
Q. E. @400nm	31.9%	23% @390 nm
P/V ratio	3	3.4
Raise time [ns]	6	
Time Transit Spread [ns]	2.4	3.5
Dark pulse rate (ave.) [kHz]	8	22.1

PMT placed at rhombic triacontahedron
(Current KL: 1325 17 inch & 554 20 inch PMTs)



Fill in the gaps with light collecting mirror



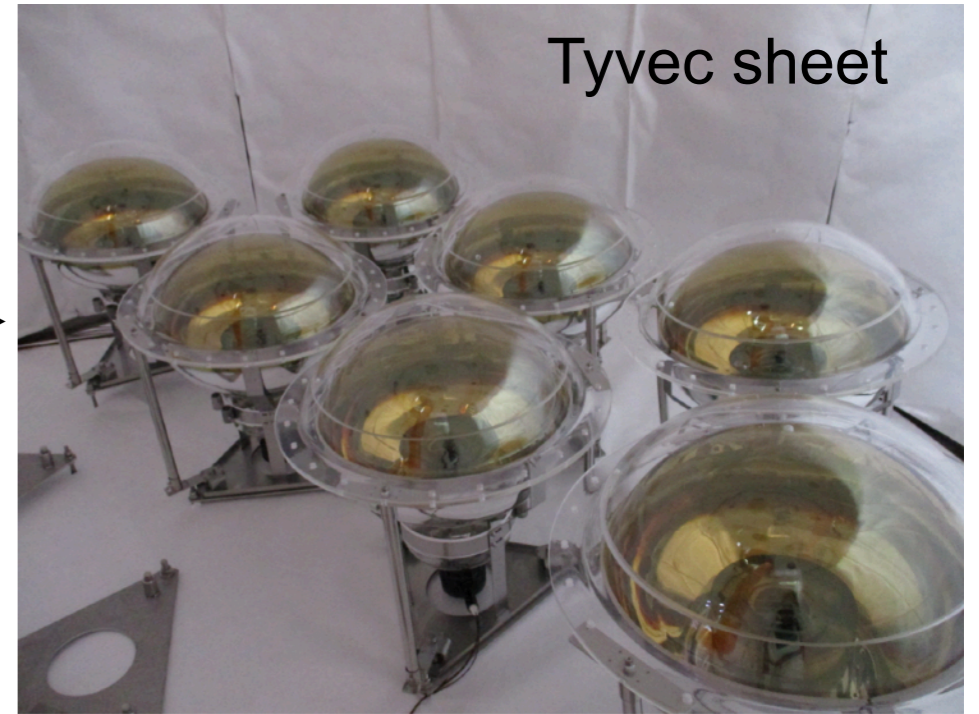
4 hexagons + 10 pentagons + 2 squares
Photo coverage ~34% → ~100%

Prototype detector for KL2

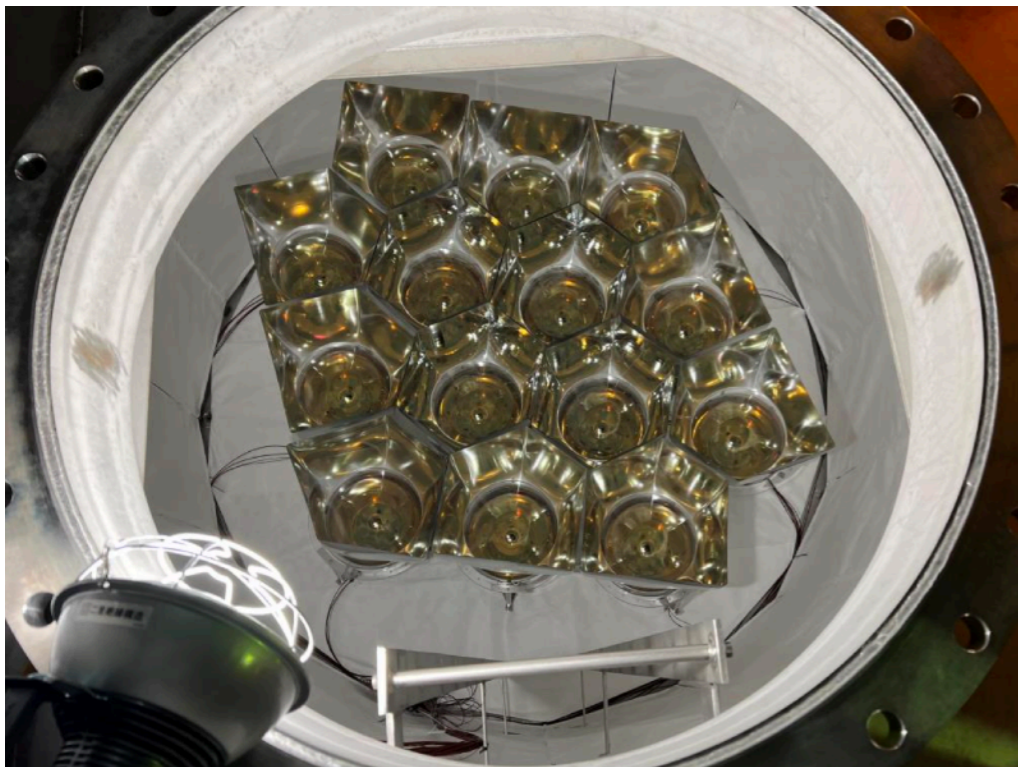
Cleaning is important work!



Inside of the detector
PMTs w/o light collection mirror



All PMTs & mirrors installed



Pure water injection

