

# **Current Status and Future Prospects**

Vivek Singh, UC Berkeley

Physics in Collisions PIC-2024

Neutrinoless Double Beta Decay

# **Historical Introduction**

- 1930: Wolfgang Pauli proposes the neutrino to explain energy conservation in beta decay.
- **1935:** Maria Goeppert-Mayer predicts double beta decay ( $2\nu\beta\beta$ ), where two neutrons decay simultaneously, emitting two electrons and two neutrinos.
- **1937:** Ettore Majorana theorizes that neutrinos could be their own antiparticles (Majorana fermions).
- 1939: Wendell H. Furry proposes neutrinoless double beta decay (0vββ), which would violate lepton number conservation and prove the Majorana nature of neutrinos.
- **Ongoing (85 years on!):** Experimental searches for 0vββ continue, aiming to determine the nature of the neutrino and its mass.





Majorana

Pauli

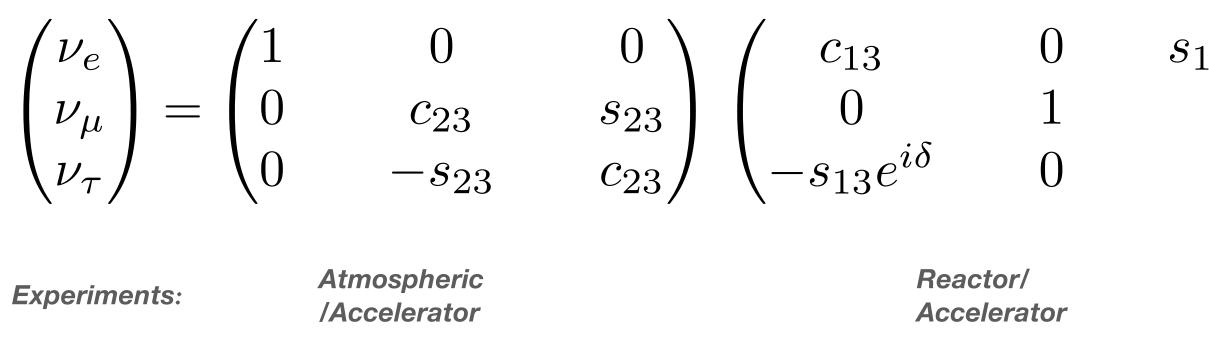


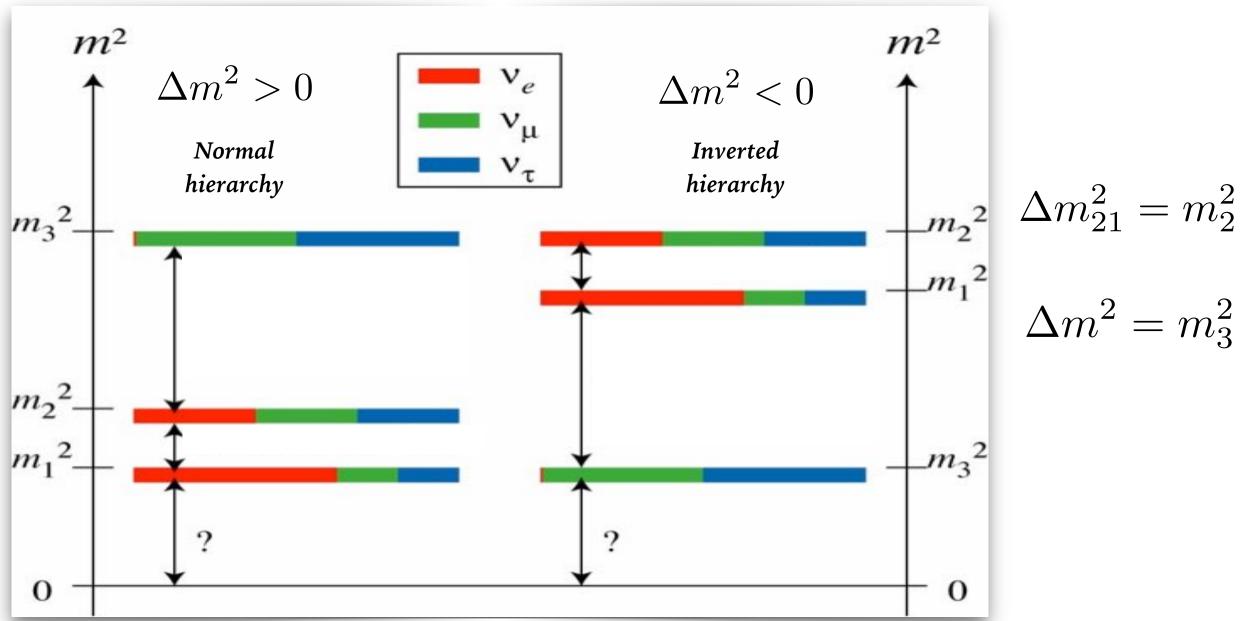


Furry



### Neutrinos





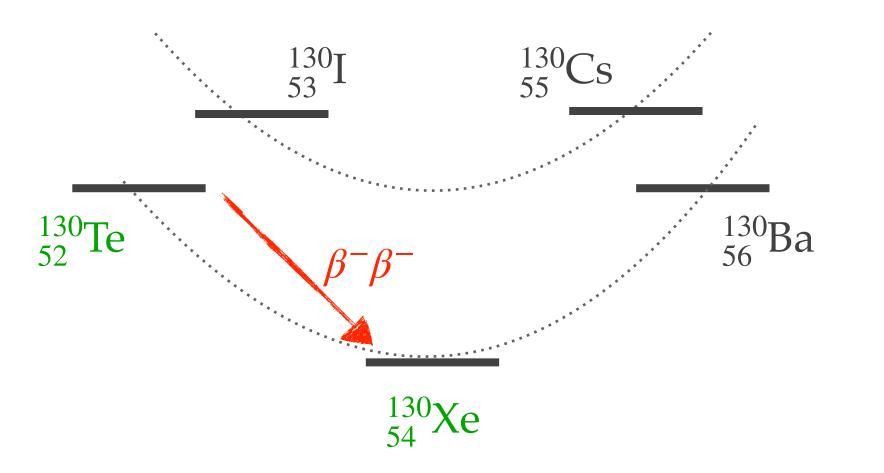
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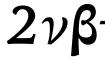
$$\begin{array}{cccc} 3e^{-i\delta}\\ 0\\ 0\\ c_{13} \end{array} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0\\ -s_{12} & c_{12} & 0\\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0\\ 0 & e^{i\frac{\alpha_2}{2}} & 0\\ 0 & 0 & e^{i\frac{\alpha_3}{2}} \end{pmatrix} \\ & & \\ \begin{array}{c} \text{Solar/}\\ \text{Reactor} & & \mathbf{0}\nu\boldsymbol{\beta}\cdot\boldsymbol{\beta} \end{array} \\ \hline \\ \hline \\ \begin{array}{c} \text{Open questions:}\\ \cdot & \text{Are there only three neutrinos?}\\ \cdot & \text{What is the mass scale?} \\ \hline \\ \text{Other properties that cannot be probed us oscillation data:}\\ \cdot & \text{Are neutrinos their own antiparticles?}\\ (\text{Majorana or Dirac?})\\ \cdot & \text{Can neutrinos responsible for matter a matter asymmetry in the universe?} \\ \hline \end{array}$$

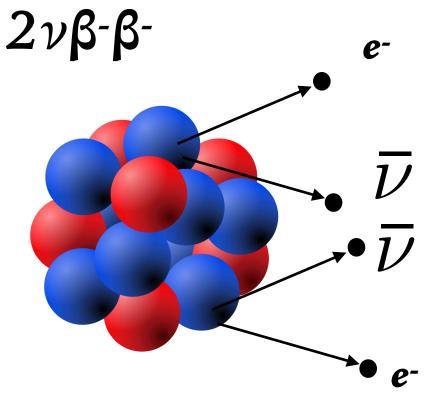


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# Neutrinoless Double Beta Decay (0vββ)







### Single beta decay forbidden.

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# Ονβ-β*e*e

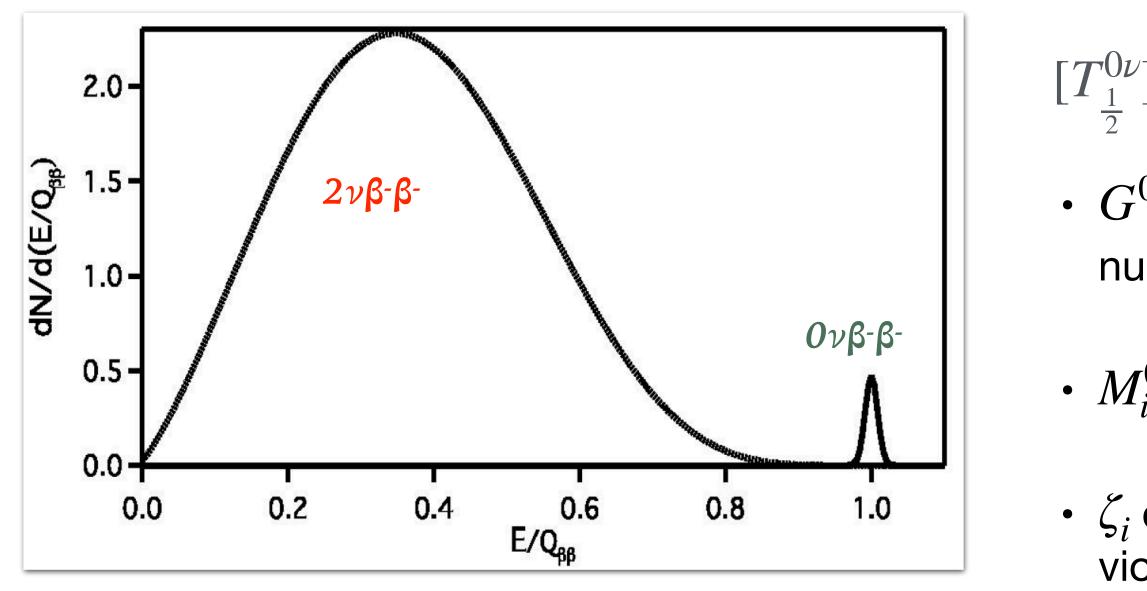
### ~35 nuclei candidates

- Can happen only if neutrinos are Majorana particles.
- Lepton number violation  $(\Delta L = 2)$
- **Requires physics beyond** standard model.
- Only 11 have  $Q_{\beta\beta} > 2 \text{MeV}$





### **Experimental Signature & the Decay Rate**



- Sum energy of emitted electrons: Peak at Q value of the decay.
- $G^{2\nu}(Z, Q)$  for  $(2\nu\beta\beta) \sim Q^{11}$
- $G^{0\nu}(Z, Q)$  for  $(0\nu\beta\beta) \propto Q^5$

$$]^{-1} = \sum_{i} G_{i}^{0\nu}(Z, Q) \cdot \left| M_{i}^{0\nu} \right|^{2} \cdot \zeta_{i}^{2}$$

•  $G^{0\nu}(Z,Q)$  is the phase-space factor that depends on the proton number (Z) of the decaying nucleus and the Q-value of the decay,

•  $M_i^{0\nu}$  is the nuclear matrix element (NME), and

•  $\zeta_i$  depends on the mechanism and mode of the lepton-numberviolating process.

In the scenario light-neutrino exchange  

$$[T_{\frac{1}{2}}^{0\nu}]^{-1} = G^{0\nu}(Z,Q) \cdot (g_A)^4 \cdot \left| M^{0\nu} \right|^2 \cdot \frac{m_{\beta\beta}^2}{m_e^2} \longrightarrow \begin{bmatrix} \text{Effective Majoral mass} \\ m_{\beta\beta} \\ |m_{\beta\beta}| = |\sum_{i=1}^{N} \frac{1}{2} \\ \text{Axial vector coupling (factored out of NME)} \end{bmatrix}$$

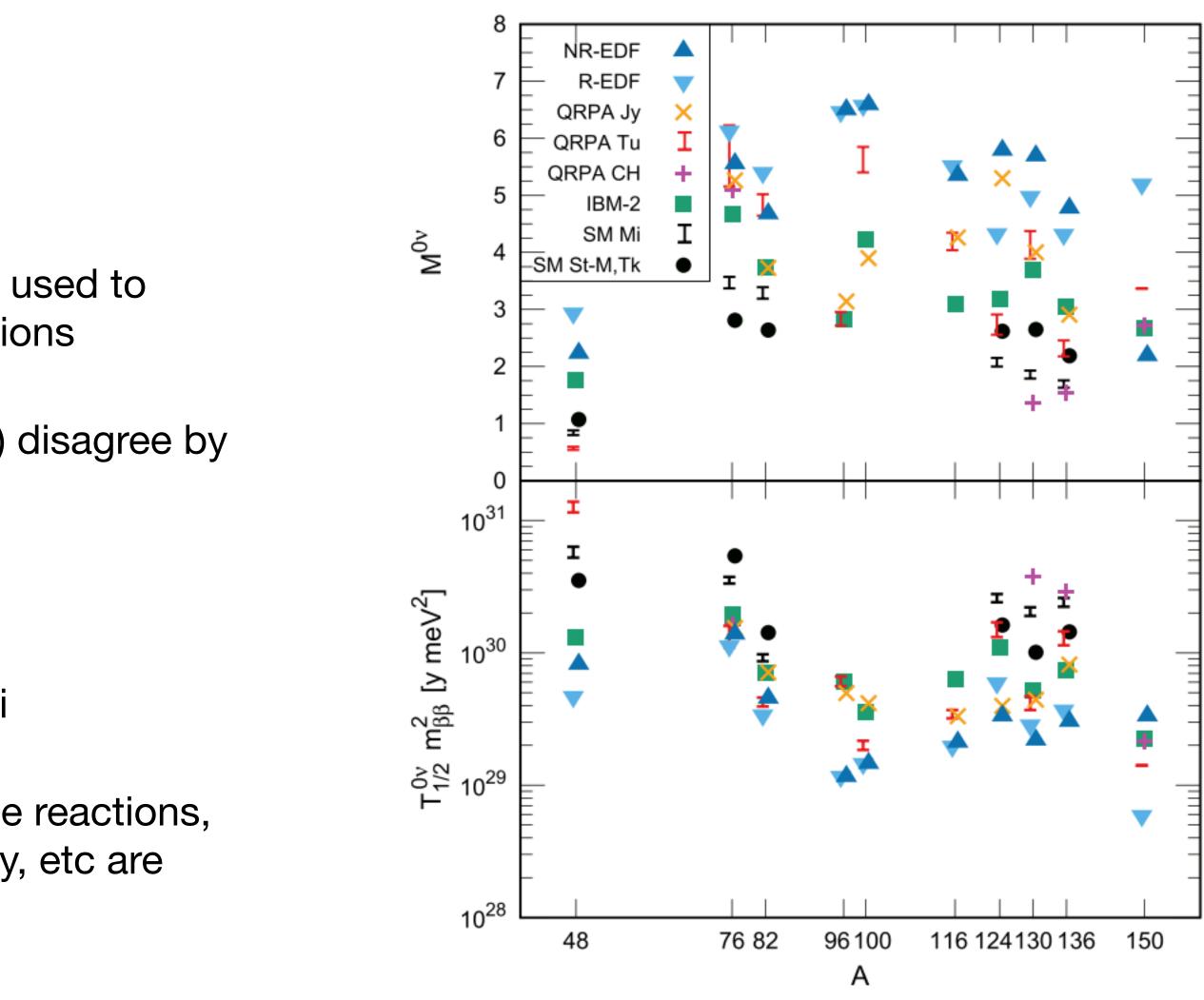
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# **Nuclear Matrix Elements Calculations**

- Extremely hard problem to solve
- Both microscopic and macroscopic nuclear models are used to calculate NMEs, each with its own strengths and limitations
- Different successful approaches (e.g., IBM, QRPA, EDF) disagree by a factor of 2-3
- Difficult to quantify errors in a reliable way
- Ab-initio methods but not yet applicable to heavy nuclei
- Various experimental probes, including charge exchange reactions, nucleon exchange, muon capture, double gamma decay, etc are used to test and constrain NME calculations



: Jonathan Engel and Javier Menéndez 2017 Rep. Prog. Phys. 80 046301



### **Experimental Sensitivity**

$$T_{1/2}^{0\nu} \propto \epsilon \cdot \sqrt{\frac{M \cdot t}{B \cdot \Delta E}}$$

(Background Limited)

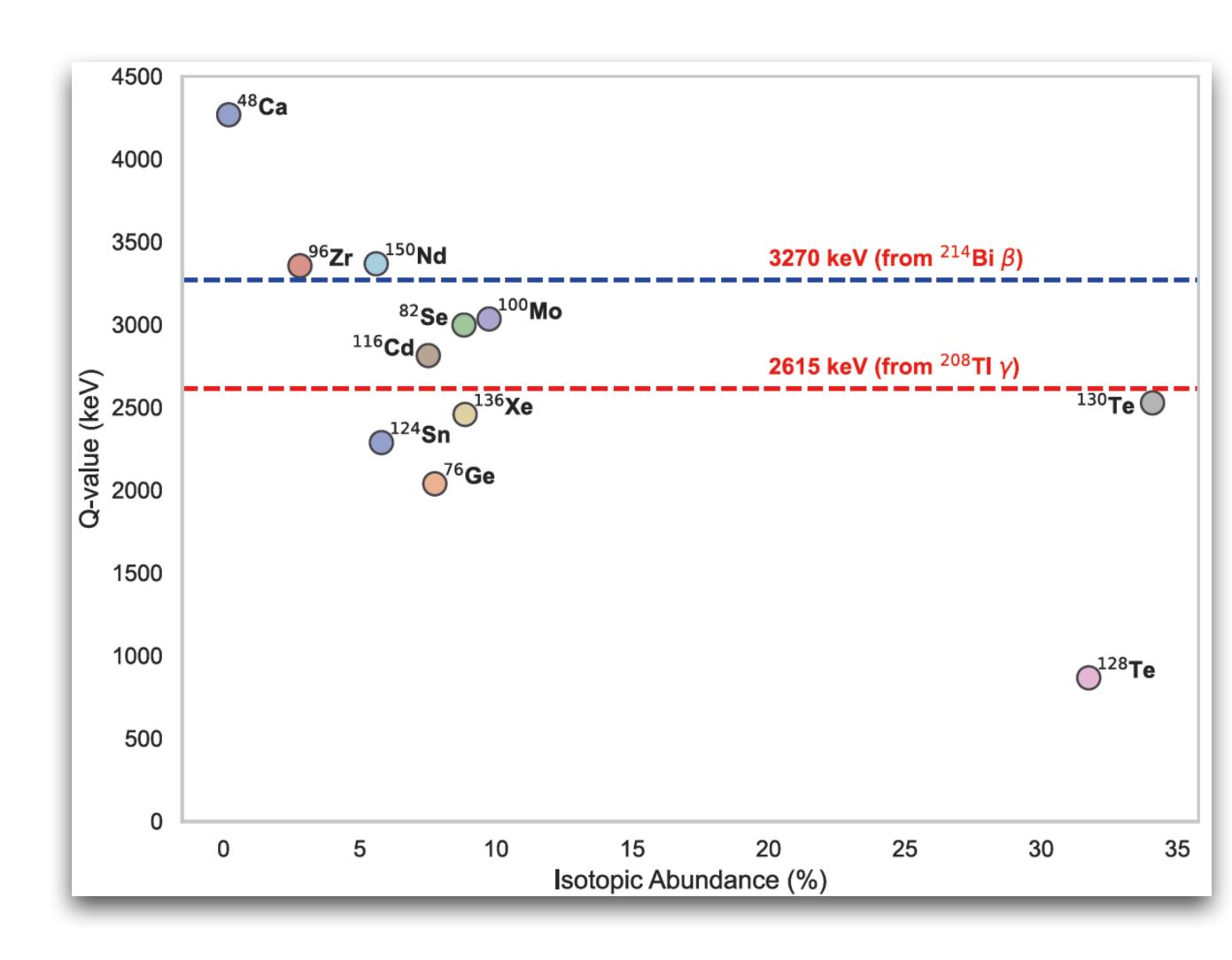
 $T_{1/2}^{0\nu} \propto \epsilon \cdot M \cdot t$ (Background "free")

M = mass of the isotope used in an experiemnt

 $\epsilon$  = Efficiency of detector

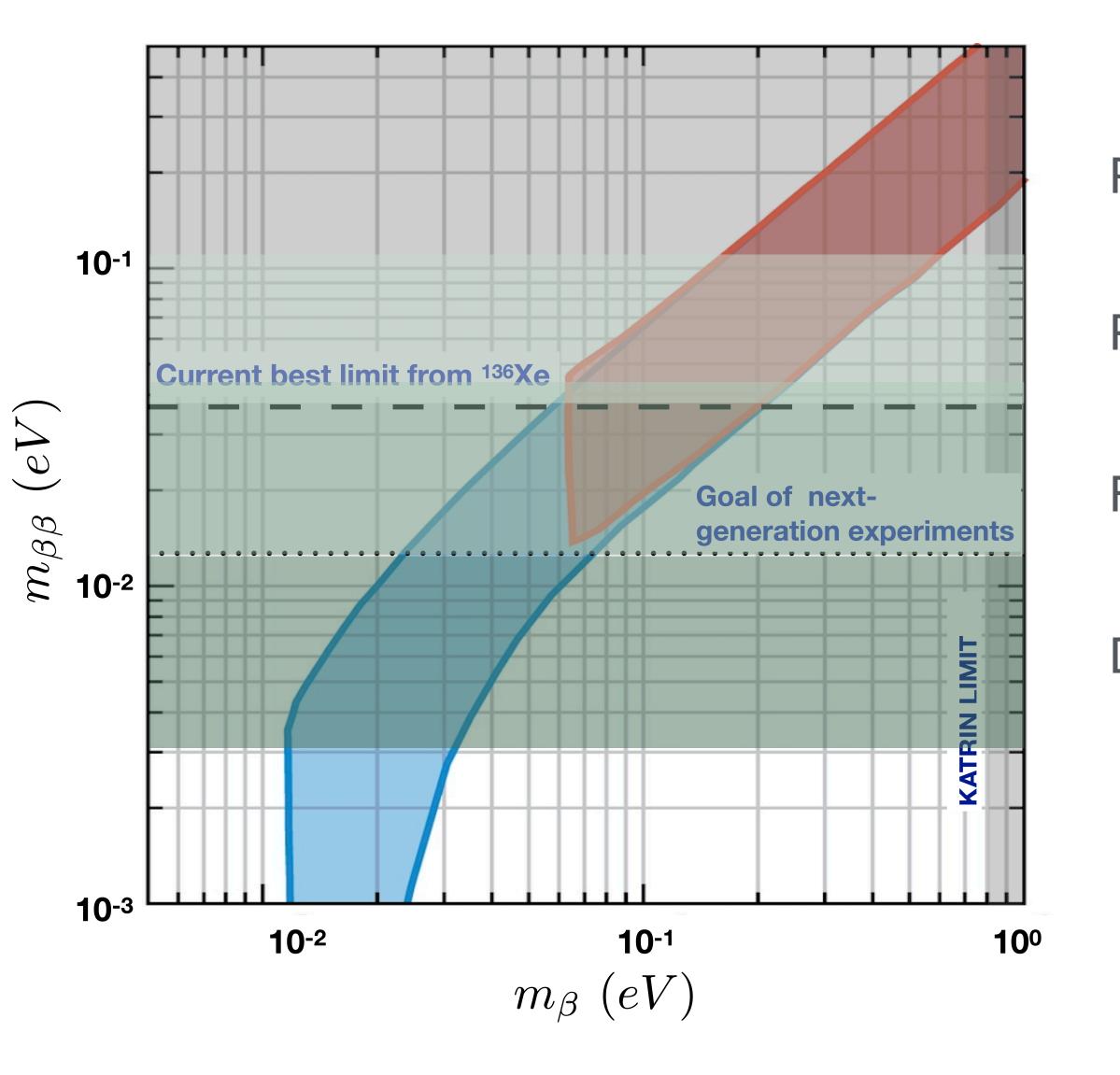
 $\Delta E$  = energy resolution in region-of-interest

B = background in counts per detector mass, energy, and time e.g, counts/(keV·kg·year)





### Implications



Past and Present (~10 kg)

Past and Near Future (~100 kg)

Future (~1000 kg)

Dreams (~10000 kg ?)

- Neutrinos are Majorana fermions.
  - Physics beyond the Standard Model.
- Constraints on absolute mass scale.
  - Probes the mass hierarchy of the neutrinos.
- Constraints on CP violating phases?





### Where do we stand right now?

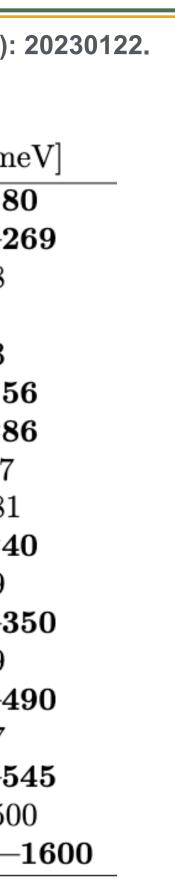
•	Diverse Research: Experiments use a	- E
	variety of isotopes to search for neutrinoless	M Ll
	double beta decay.	L C K
•	Active Field: Several experiments have	E
	already been completed, providing valuable	N C
	data.	SI A A
•	Future Focused: New, more sensitive	C C
	experiments are under construction and	C Su C
	planned, promising increased sensitivity and	
	potential for discovery.	No coi
		Ag

Adapted from Parno, D.S., Poon, A.W.P, Singh, V., Philosophical Transactions A 382.2275 (2024): 20230122.

Experiment	Status	Isotope	$T_{1/2}^{0\nu}$ [yr]	$m_{etaeta}$ [me
GERDA	Completed	$^{76}\mathrm{Ge}$	$1.8 imes 10^{26}$	79—18
MAJORANA	Completed	$^{76}\mathrm{Ge}$	$f 8.5 imes 10^{25}$	113 - 20
LEGEND-200	Taking Data	$^{76}\mathrm{Ge}$	$1.5  imes 10^{27}$	34 - 78
LEGEND-1000	Proposed	$^{76}\mathrm{Ge}$	$8.5 imes10^{28}$	9 - 21
$\mathrm{CDEX}\text{-}300 \nu$	Proposed	$^{76}\mathrm{Ge}$	$3.3 imes10^{27}$	18 - 43
KamLAND-Zen	Taking Data	$^{136}\mathrm{Xe}$	$f 2.3 imes 10^{26}$	36 - 15
EXO-200	Completed	$^{136}\mathrm{Xe}$	$3.5  imes \mathbf{10^{25}}$	93—28
nEXO	Proposed	$^{136}\mathrm{Xe}$	$1.3  imes 10^{28}$	6.1 - 27
NEXT-100	Construction	$^{136}\mathrm{Xe}$	$7.0 imes10^{25}$	66 - 281
CUORE	Taking Data	$^{130}\mathrm{Te}$	$3.8  imes \mathbf{10^{25}}$	70 - 24
SNO+	Construction	$^{130}\mathrm{Te}$	$2.1 imes10^{26}$	37 - 89
AMoRE-I	Completed	$^{100}\mathrm{Mo}$	$3.0 imes10^{24}$	210 - 3
AMoRE-II	Proposed	$^{100}\mathrm{Mo}$	$5.0 imes10^{26}$	17 - 29
CUPID-Mo	Completed	$^{100}\mathrm{Mo}$	$\mathbf{1.8  imes 10^{24}}$	280 - 42
CUPID	Proposed	$^{100}\mathrm{Mo}$	$1.5 imes10^{27}$	10 - 17
CUPID-0	Completed	$^{82}$ Se	$f 4.6 imes 10^{f 24}$	263-54
SuperNEMO-D	Construction	$^{82}$ Se	$4.0  imes 10^{24}$	260 - 50
CANDLES-III	Taking data	$^{48}Ca$	${f 5.6 imes 10^{22}}$	<b>2900</b> —

ot an exhaustive list. See the following for an excellent and mprehensive review

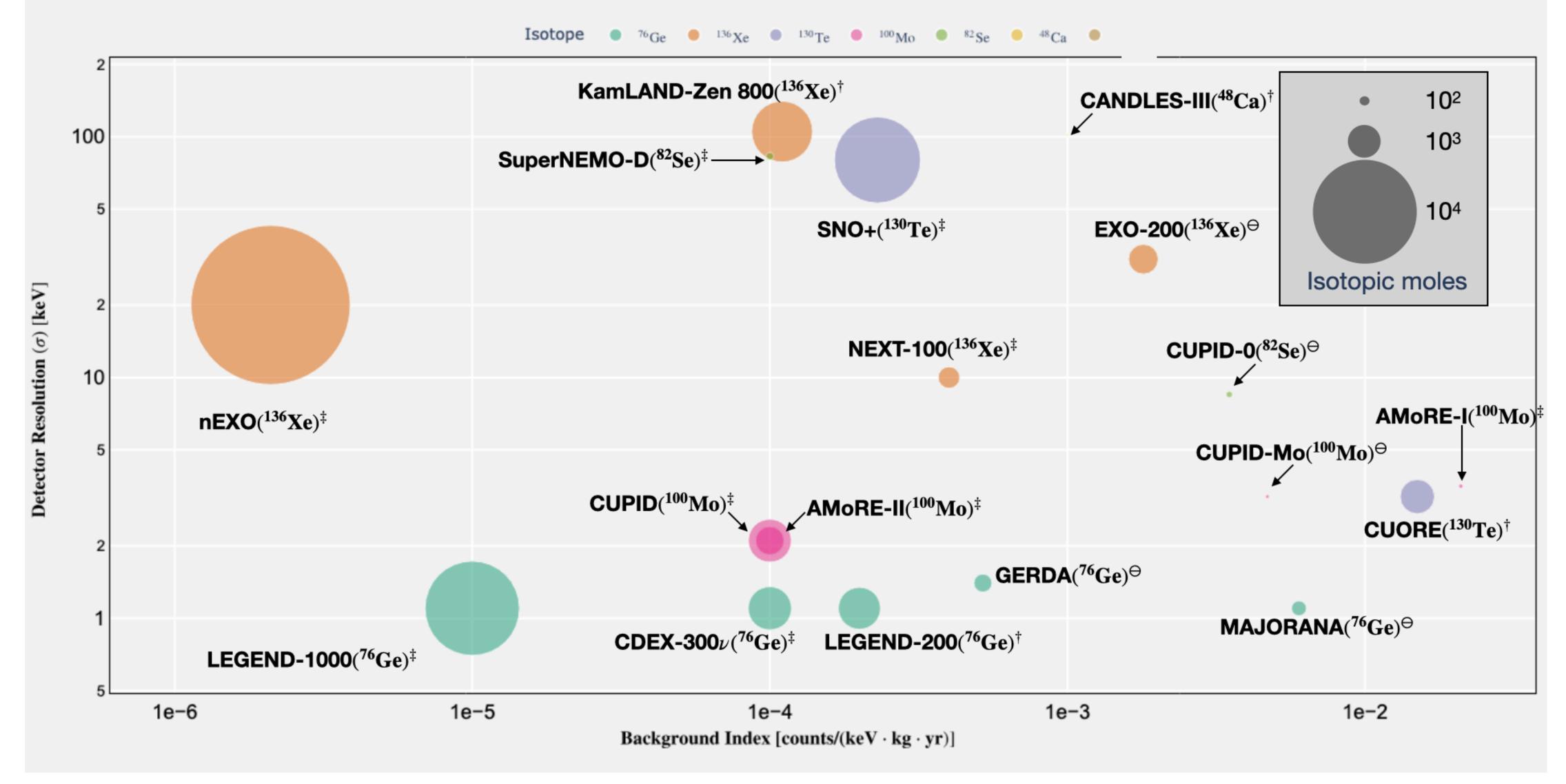
jostini, Matteo, et al. "Toward the discovery of matter creation with utrinoless ββ decay." Reviews of Modern Physics 95.2 (2023): 025002.







### Holy Trinity - Mass, Background, Resolution

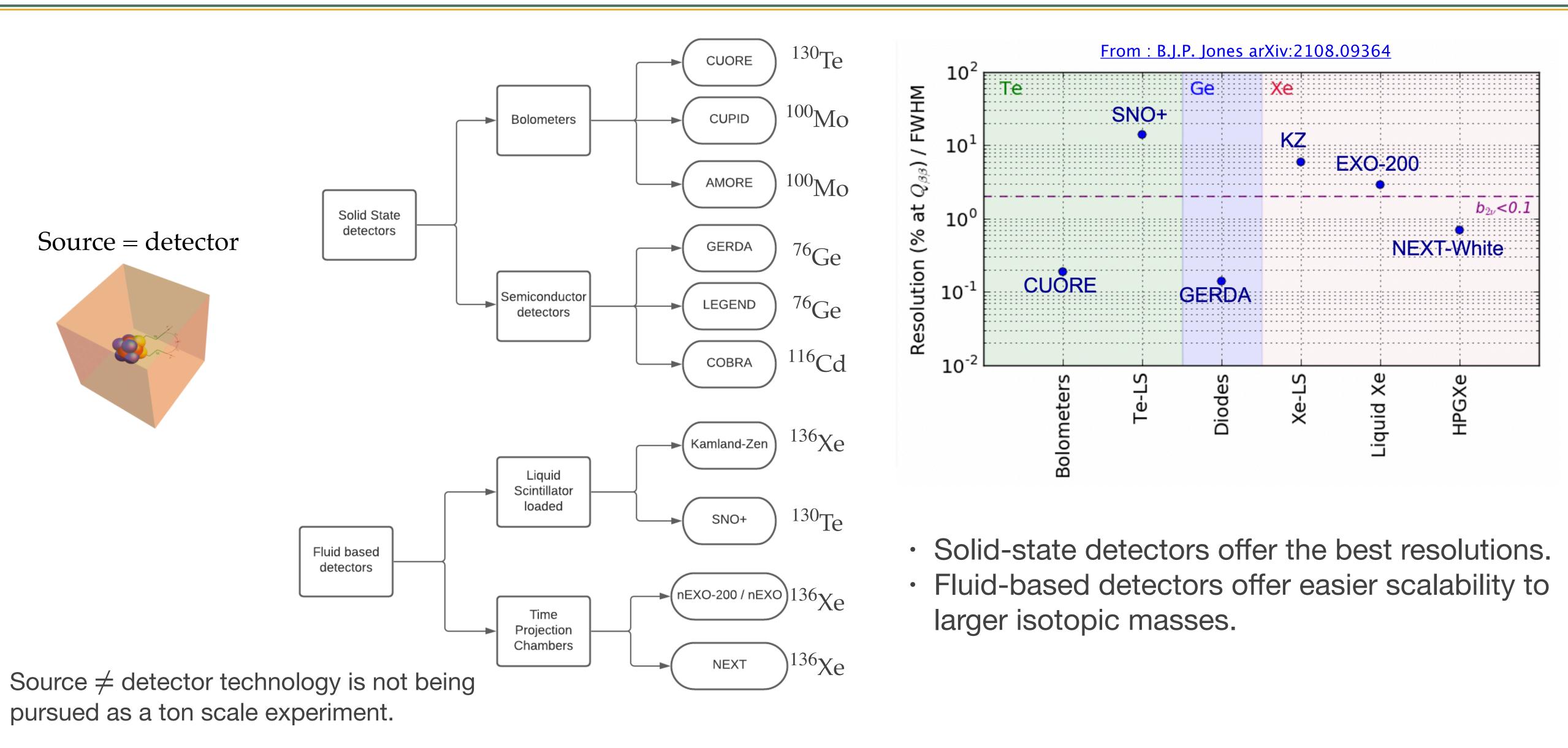


Adapted from Parno, D.S., Poon, A.W.P, Singh, V., Philosophical Transactions A 382.2275 (2024): 20230122.

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### **Detector Strategies**



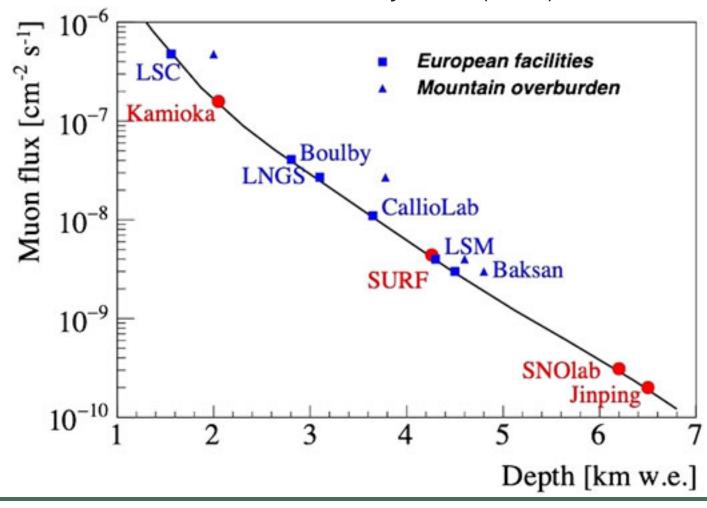
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# **Background Mitigation**

- Location, Location, Location: Experiments are conducted deep underground to shield from cosmic rays.
- Material Matters: Detectors are built from radioactively pure materials to minimize internal background.
- **Shielding Strategy:** Multiple layers of passive shielding (e.g., water, lead) are used to block external radiation.
- Active Veto: Some shielding materials double as • active detectors to identify and reject background events (e.g., cosmic muons).
- **Event Discrimination:** Sophisticated analysis techniques (timing, topology, particle ID) differentiate between signal events ( $0\nu\beta\beta$ ) and background.









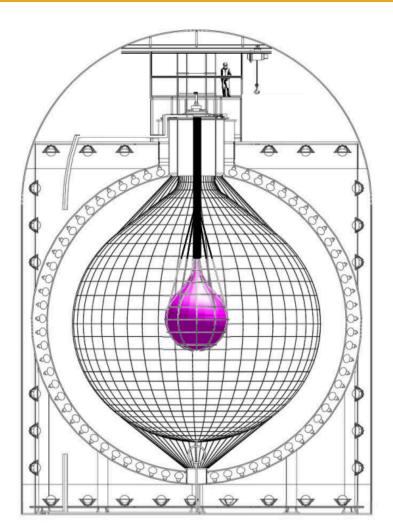
# The Hunt for 0vββ: A Tour of Experiments

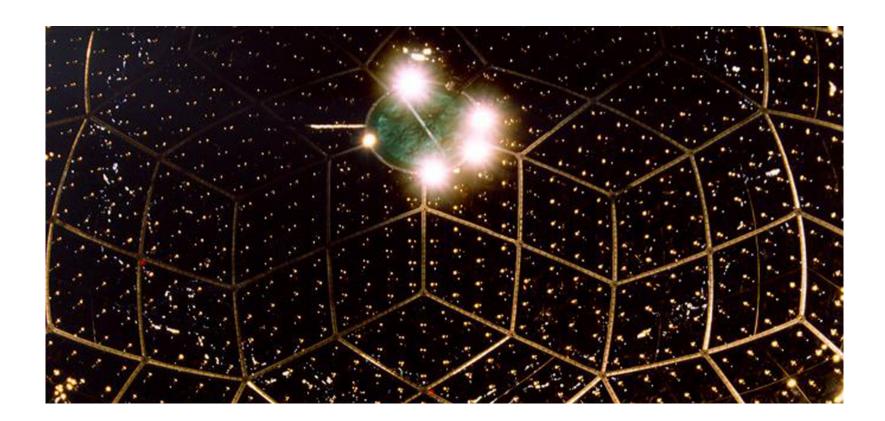
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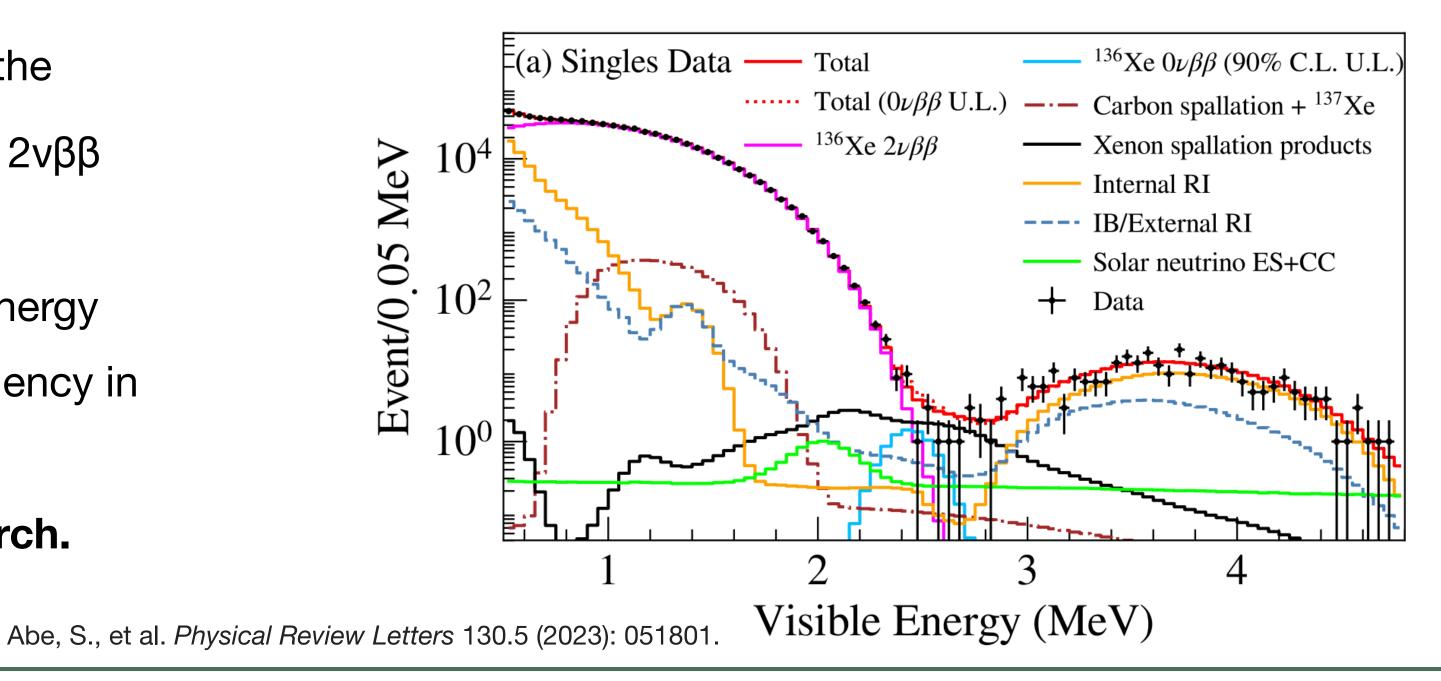


## KamLAND-Zen 800

- Kamioka Observatory in Japan data taking started in 2019.
- KamLAND-Zen utilizes a 745 kg Xenon gas (enriched to 90-91%) dissolved into 1 kiloton of liquid scintillator.
- Dual-phase design allows the liquid scintillator to act as both target and active shield.
- Major background sources include <sup>214</sup>Bi from the decay chain of  $^{238}$ U, cosmogenic  $^{10}$ C, and the  $2\nu\beta\beta$ decay itself.
- Despite excellent background reduction, the energy resolution is limited by the light collection efficiency in the large detector volume.
- Current world leader for most sensitive search.

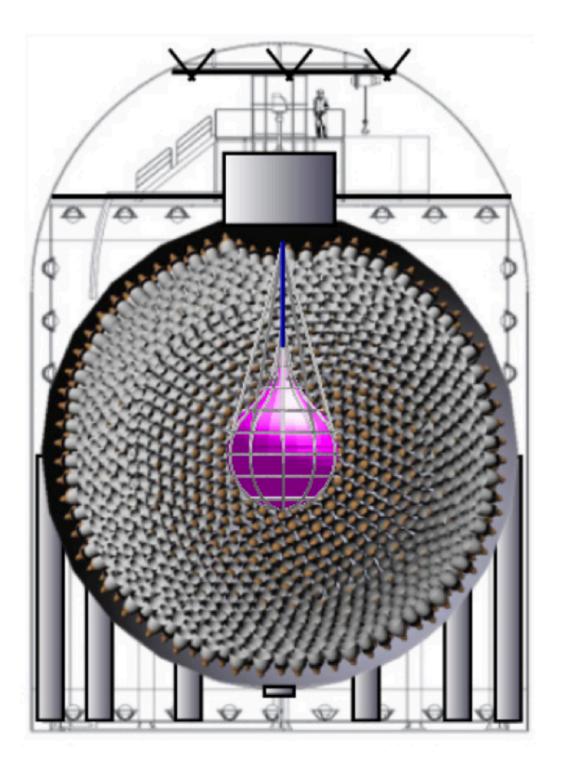








### KamLAND2-Zen



- Increase the Xenon mass to 1 ton.
- Improve energy resolution by using a brighter liquid
- This upgrade aims for a half-life sensitivity of >1.1 x
  - 10<sup>27</sup> years.

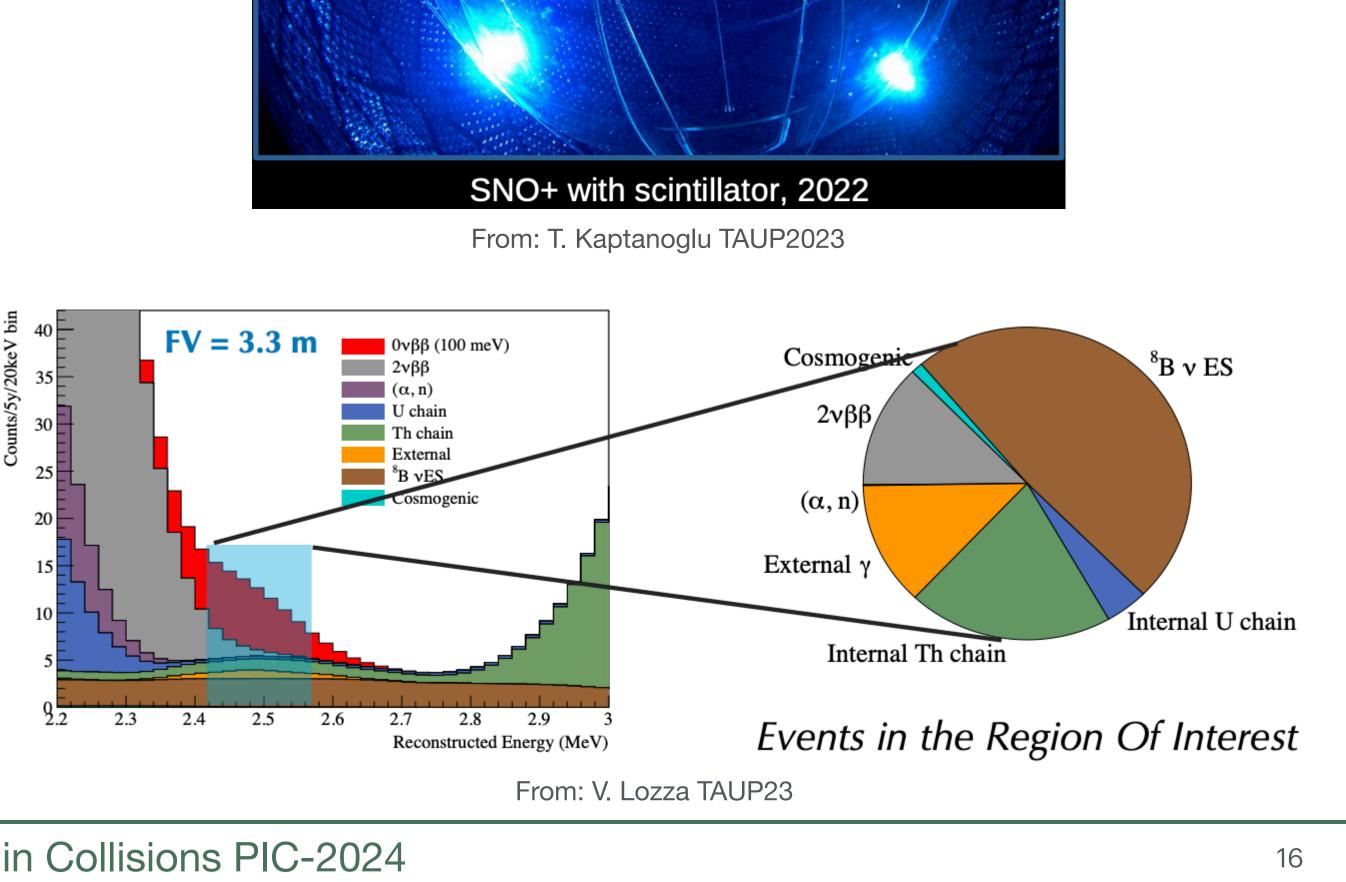
scintillator and more efficient photomultiplier tubes.



### SNO+

- SNO+ employs a novel technique to load Tellurium-130 into liquid scintillator, enhancing light yield and stability for improved detection.
- Situated in SNOLAB, Canada, providing significant depth for cosmic ray shielding.
- **Phased Approach:** SNO+ has undergone meticulous background characterization with water and pure scintillator phases, paving the way for Tellurium loading in 2025.
- With an initial 0.5% loading of natural Tellurium, SNO+ aims for a half-life sensitivity exceeding 2.1 x 10<sup>26</sup> years after 3 years of data taking.
- Future Upgrades: R&D efforts demonstrate the potential to increase the Tellurium-130 loading to 3%, promising a substantial boost in sensitivity for future searches.

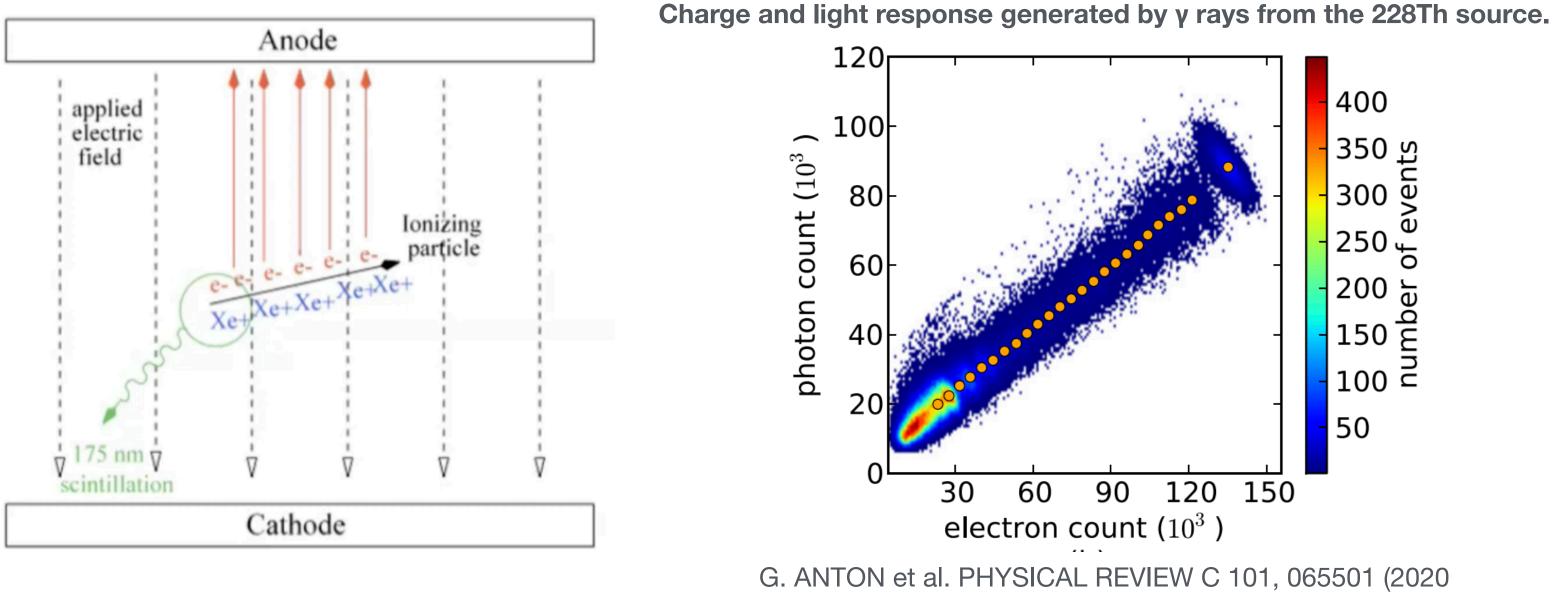




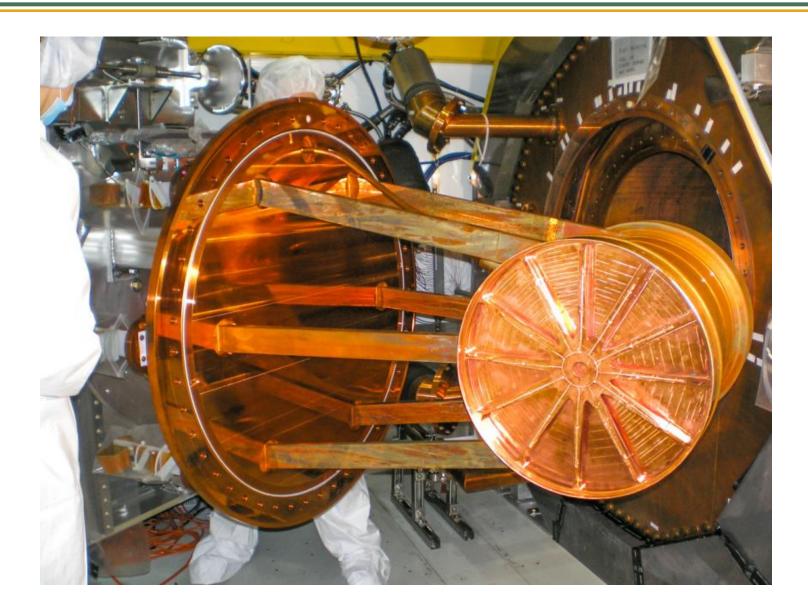
### EXC

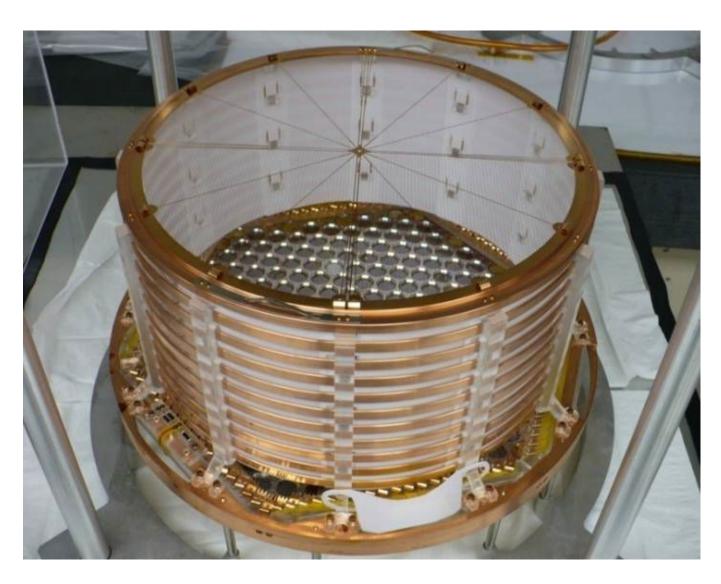
- Operated with enriched Xe at Waste Isolation Pilot Plant in two phases between 2011 and 2018.
- Liquid Xenon TPC. Readout plane made of LAAPDs + crossed wire grip.
- ~100 kg fiducial mass with 80% enriched 136-Xe.
- The TPC technology enabled detailed reconstruction of event topology, aiding in background rejection by differentiating signal events from background interactions.





### Vivek Singh, UC Berkeley



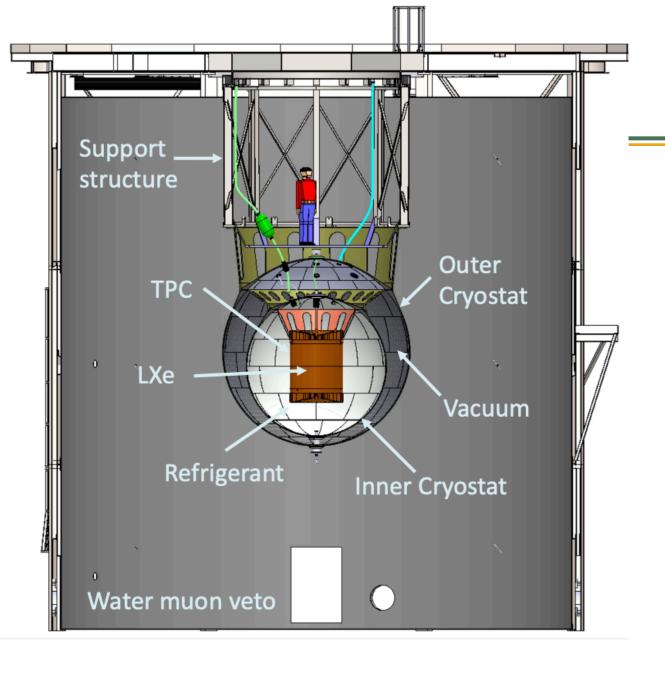




### nEXO

- A planned upgrade to EXO-200, featuring a larger liquid xenon TPC with 5 tons of enriched xenon.
- Aims for a half-life sensitivity exceeding 1.35 x 10<sup>28</sup> years
- Projected to reduce background by a factor of ~1000 compared to its predecessor, EXO-200.
- Targets an energy resolution of <1% to precisely</li> identify potential  $0\nu\beta\beta$  decay events.
- Exploring a revolutionary technique to tag the barium daughter atom, potentially eliminating almost all background in a future phase.

Images From: Samuele Sangiorgio TAUP 2023



### **nEXO TPC Conceptual Design**

Charge Tiles Charge Tiles Support tensioning spring SiPMs sapphire rods SiPM SiPM Staves staves Field Shaping Rings Cu field rings Support Rods and Spacers Cathode



Lawrence Livermore National Laboratory LLNL-PRES-85355



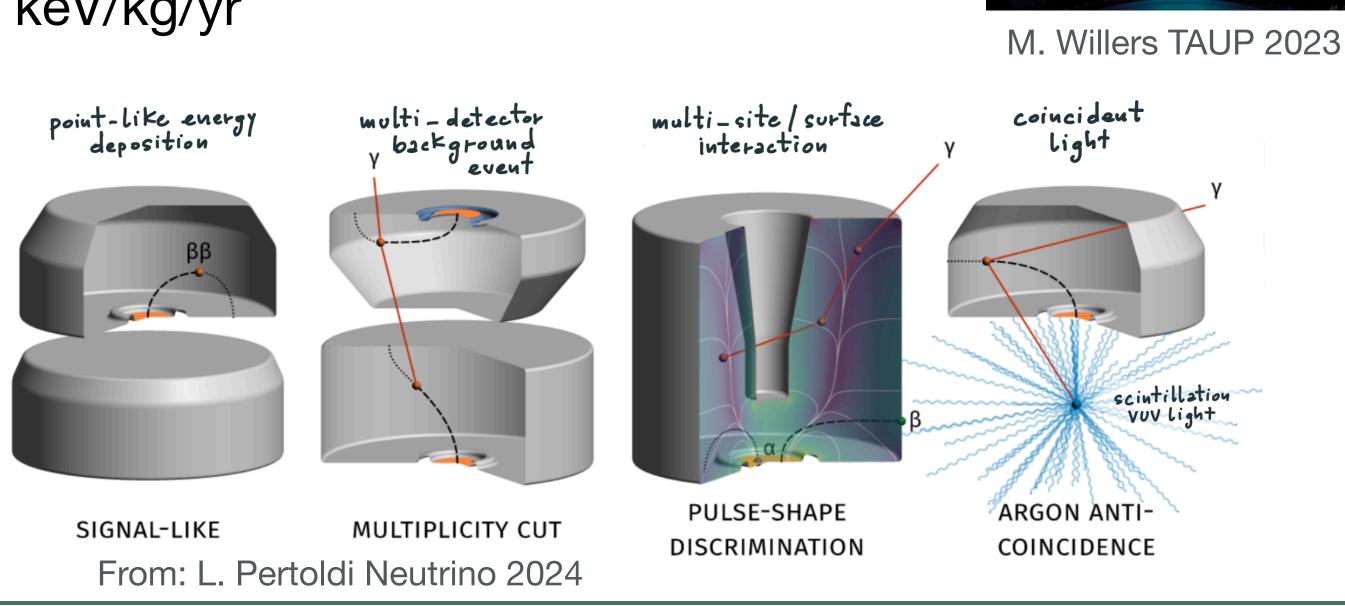
backbone								
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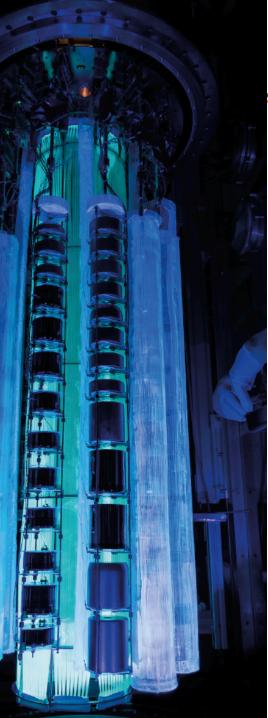


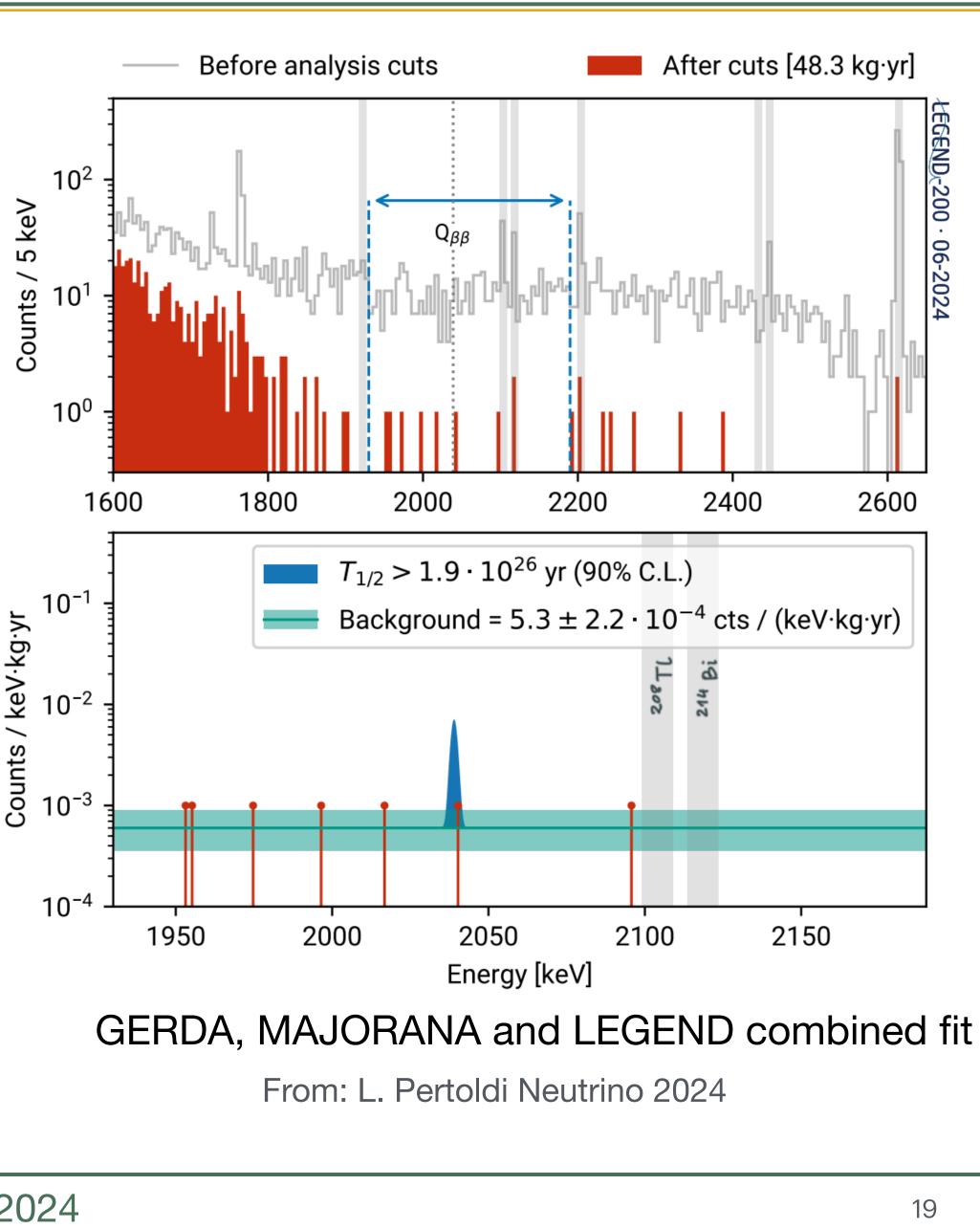
### LEGEND-200

- 142 kg of HPGe detectors
- Submerged in liquid Ar
- Mesh shroud to protect from 42K
- LEGEND-200 sees a low background
- 48.3 kg-yr of data
- Preliminary BI:  $(5.3 \pm 2.2) \times 10^{-4} \text{ cts/}$ keV/kg/yr



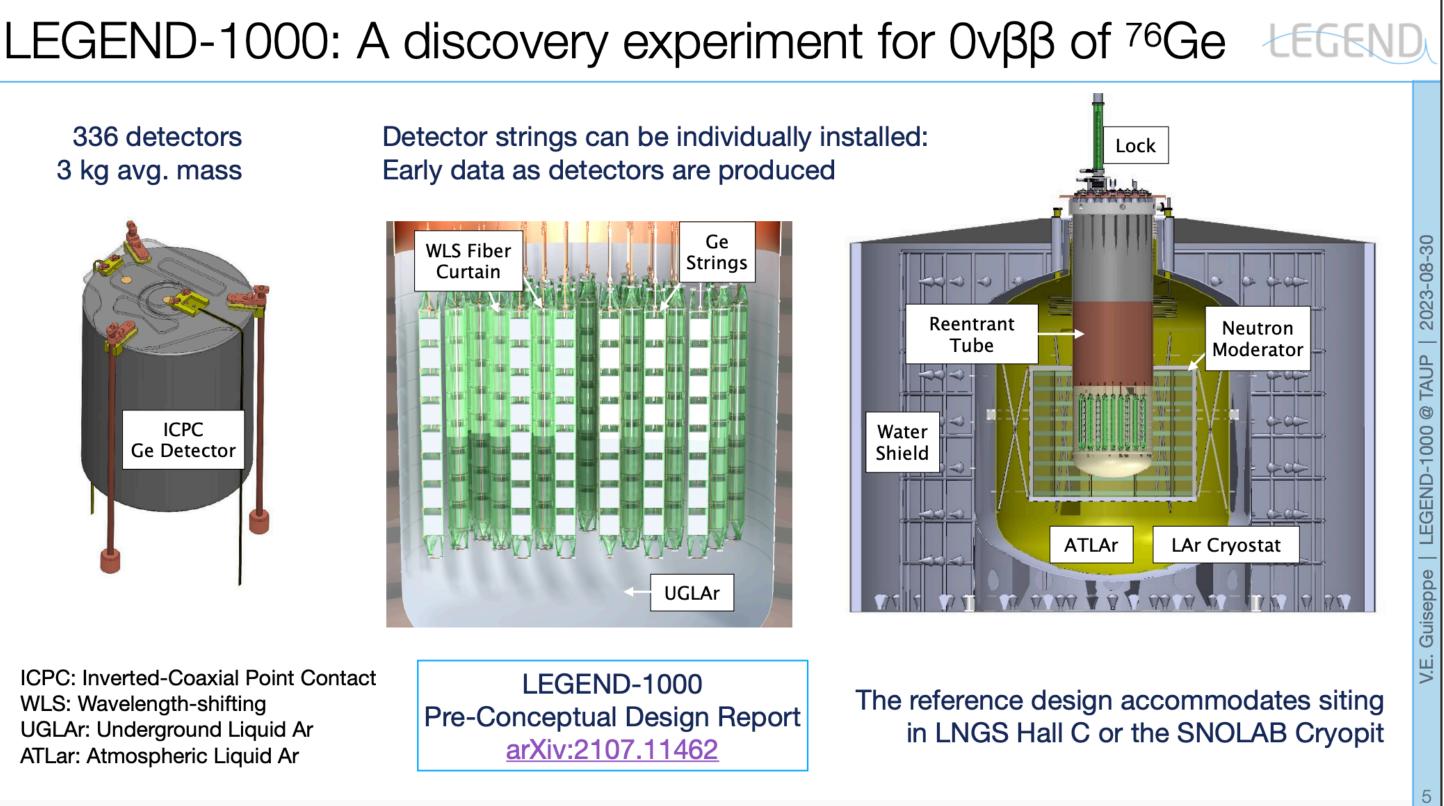
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# **LEGEND-1000**

- Builds on breakthrough developments by GERDA, MAJORANA, and LEGEND-200.
- Excellent energy resolution.
- Aims to be a quasi-background-free experiment
  - Larger volume/surface ratio of the detectors to reduce background.
  - Low-mass ASIC electronics
  - Reduction in 42Ar by procuring underground liquid Argon.
  - Deeper underground site (baseline design at SNOLAB)
- Aims for a half-life sensitivity exceeding 1.3 x 10<sup>28</sup> years



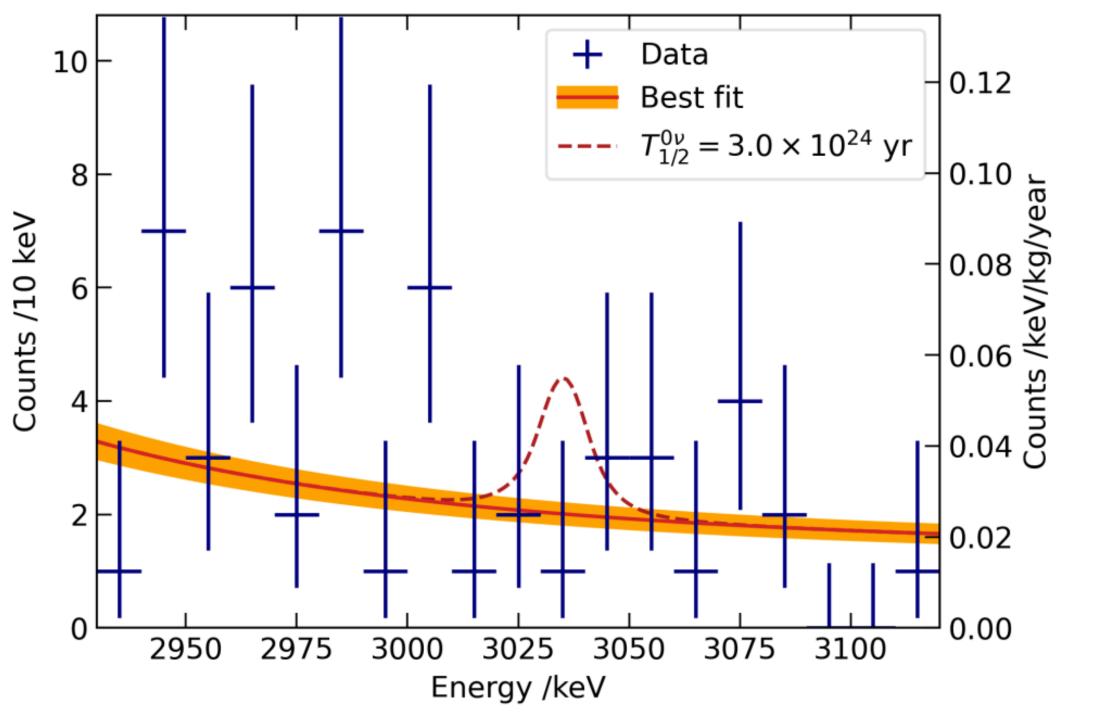
ICPC: Inverted-Coaxial Point Contact WLS: Wavelength-shifting UGLAr: Underground Liquid Ar ATLar: Atmospheric Liquid Ar

From: V. Guisippe TAUP 2023



### **AMoRE-I**

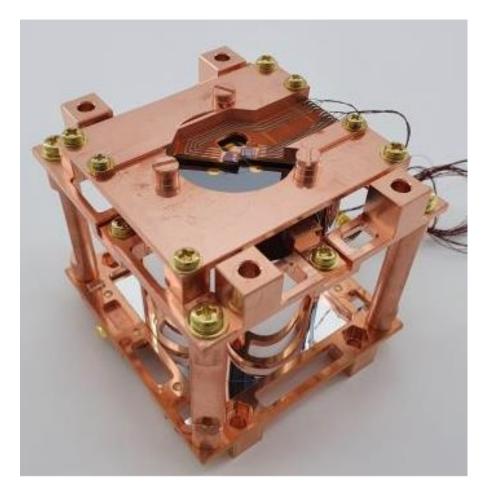
- At Y2L lab South Korea
- 13x CaMoO4 and 5x Li2MoO4 crystals
- Readout with metallic magnetic calorimeter (MMC) + SQUID sensors

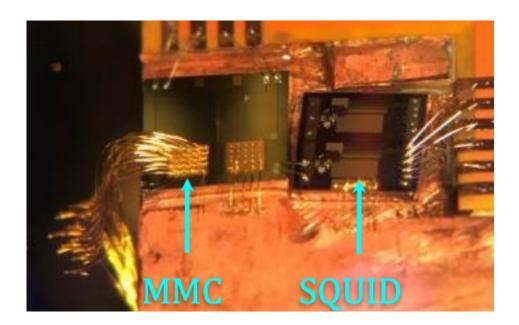


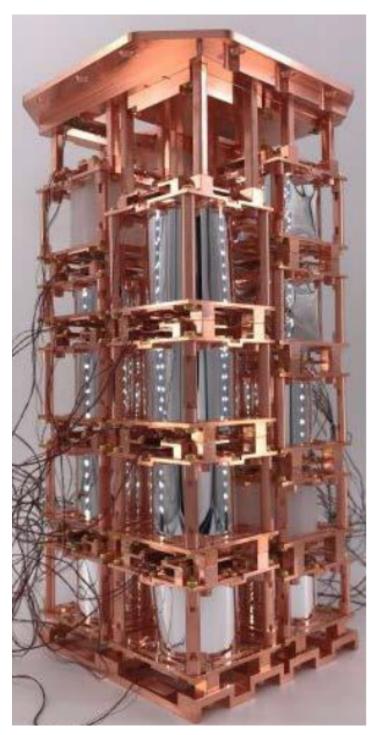


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- -0.08 Å/ - 0.06 Counts /ke/

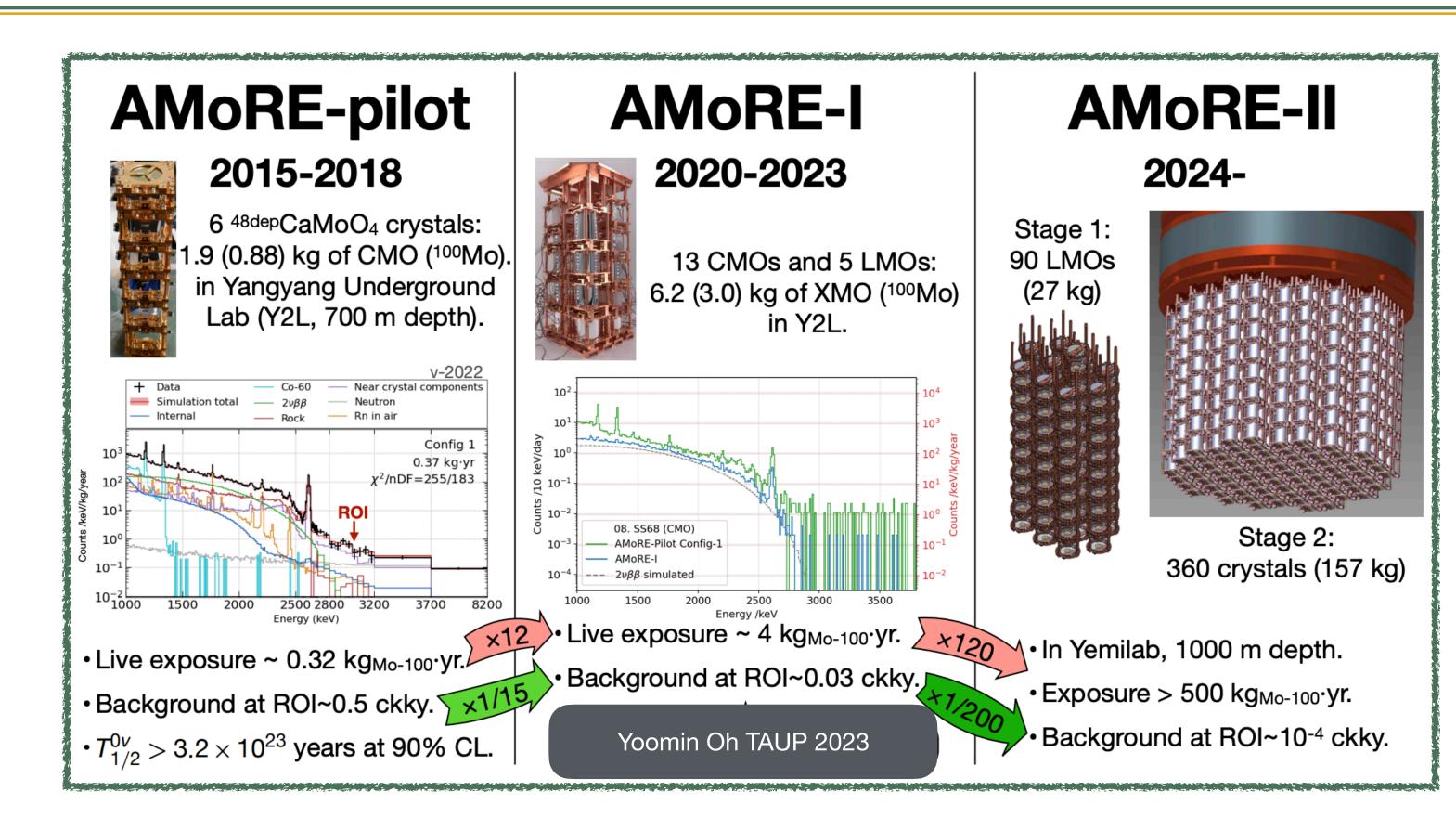








- Another next-generation <sup>100</sup>Mo experiement
- At a deeper Yemilab (1000 m rock overburden)
- Aim BI  $< 1 \times 10^{-4}$  cnts/ (keV. kg. y).
- Started crystal production.
- Aims for a half-life sensitivity exceeding 4.6 x 10<sup>26</sup> years



Agrawal, A., et al. arXiv preprint arXiv:2407.05618 (2024)

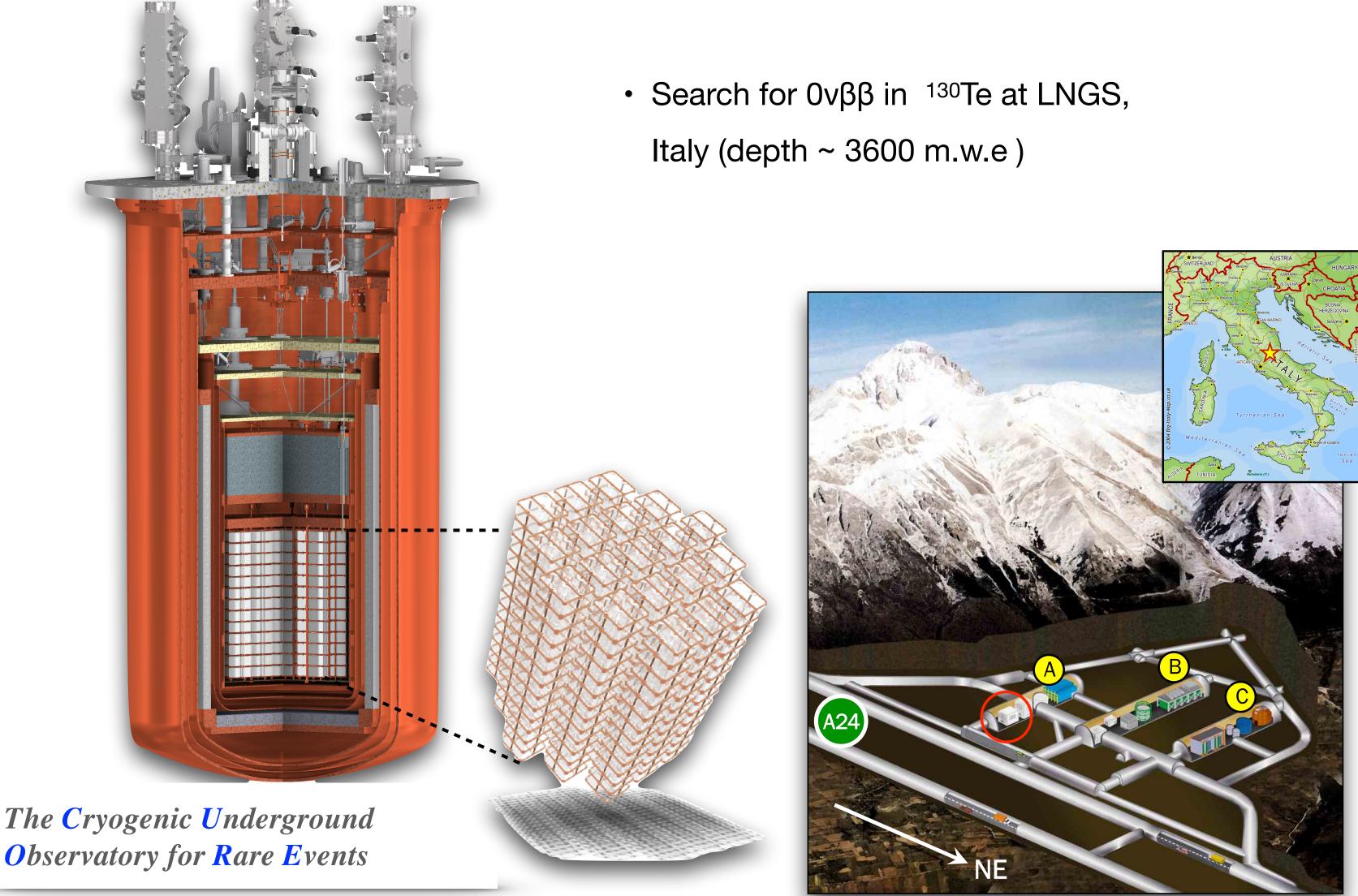
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### Discovery sensitivity for 500 kg. yr of exposure

 $m_{\beta\beta} < 20-35 \text{ meV}$ 

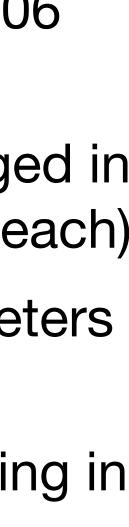


### CUORE



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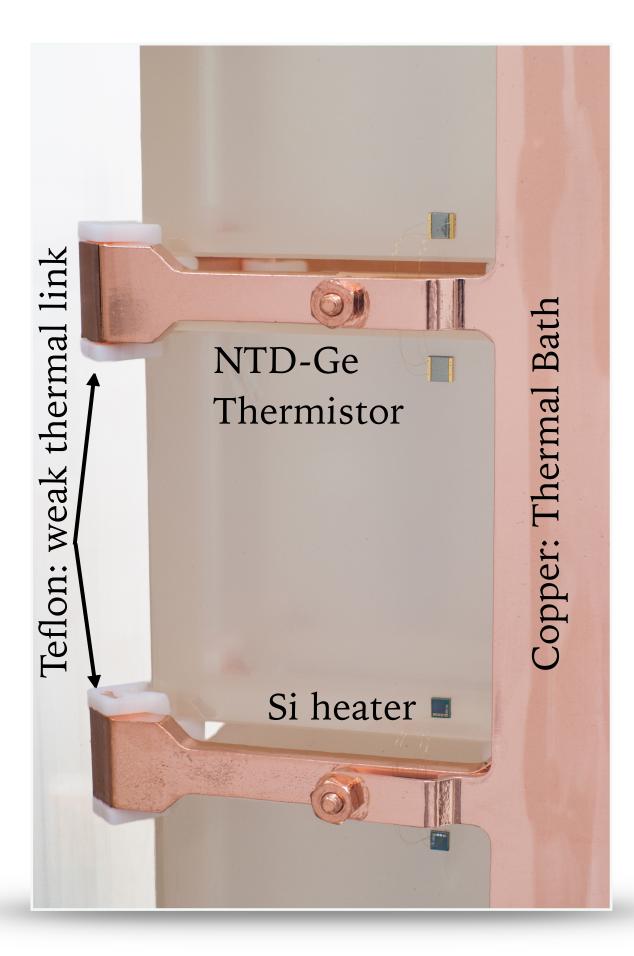
- $Q_{\beta\beta} = 2528 \text{ keV}$
- Isotopic mass of <sup>130</sup>Te : 206 kg
- 988 TeO<sub>2</sub> crystals (arranged in 19 towers with 13 floors each)
- Massive thermal calorimeters operated at ~10 mK
- CUORE started data-taking in 2017.
- Interruptions for cryogenic optimization until 2019 but have been steadily taking data ever since.





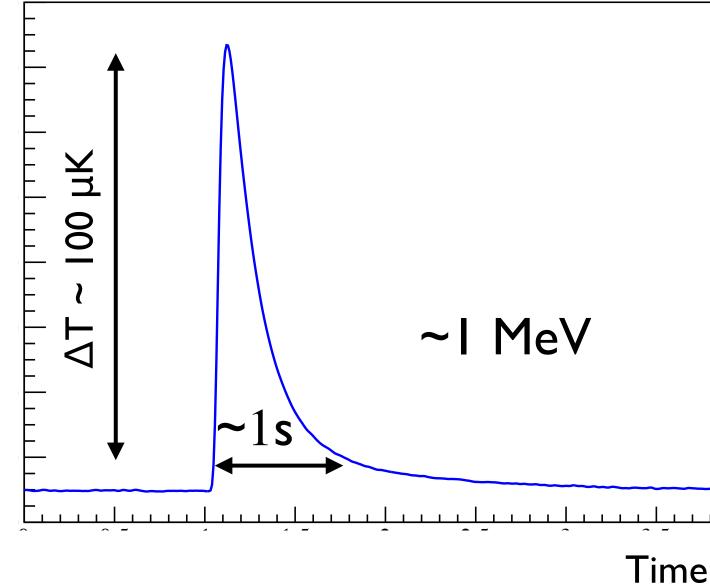


# **CUORE - DETECTOR PRINCIPLE**

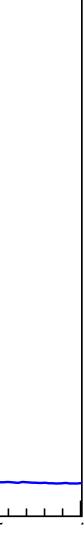


- 750 g (5x5x5 cm<sup>3</sup>) crystal
- 988 crystals installed in CUORE
- $\triangle T \sim 100 \ \mu K$  for 1 MeV energy deposit
- NTD-Ge thermistor read out
  - R(T) ~ R<sub>0</sub> exp [  $(T_0/T)^{1/2}$  ]
    - (large sensitivity at low T)
- Energy response calibrated using known ulletgamma sources
- Note:

  - Signal  $\rightarrow$  thermal channel only No active background rejection



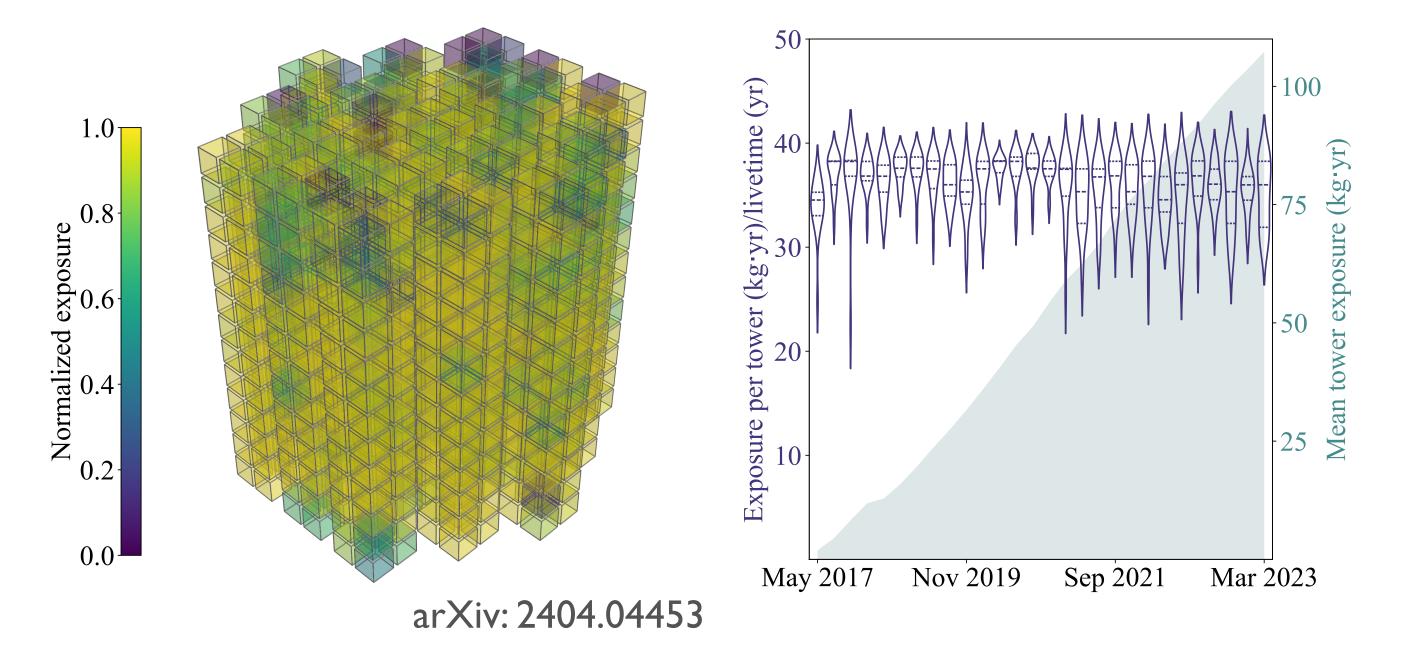
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# CUORE - DATA TAKING

- 34 datasets collected so far, each dataset ~1 month long (~50 kg.yr TeO2/ month)
- Published  $0\nu\beta\beta$  search results in 2022 for the data collected up to 2020
- Updated search with data collected up to April 2023 (~2 tonne.yrs of TeO<sub>2</sub>) exposure)



CUORE Preliminary 2039.0 kg.yr

Vivek Singh, UC Berkeley

### 365 days highlights from news & views 2022

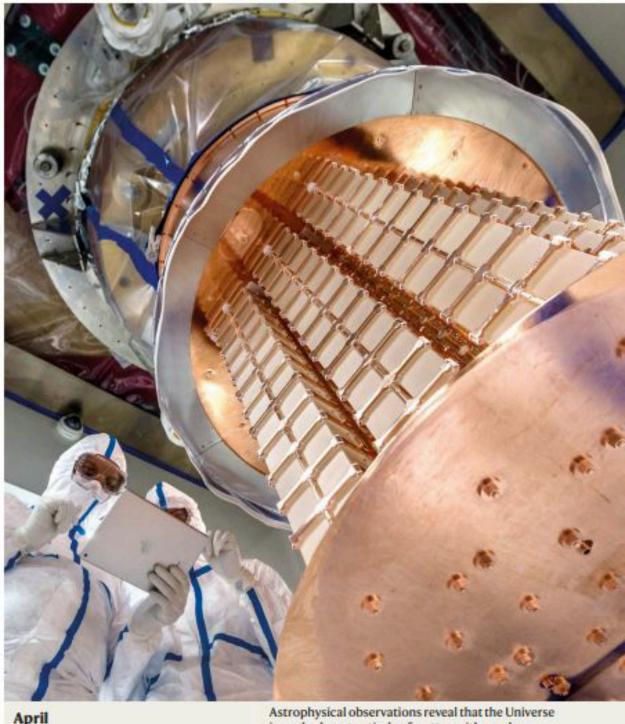
NeuroImaging

March

### **Brain changes after** COVID revealed by imaging

In 2020, the UK Biobank (a large scale biomedical database and research resource) launched a COVID-19 repeat-imaging study in which participants who had completed a medical-imaging session before the pandemic returned for an identical, second scan session. Douaud et al. explored these data, comparing scans pre- and post-pandemic. Participants who had tested positive for SARS-CoV-2 between the two scans exhibite changes in the brain cortex that are often associated with worsening brain health. This group also displayed increases in markers of tissue damage in brain regions connected to smell and taste. There is much more work to be done to extract all the useful information from this valuable data set. The UK Biobank's data sharing and Douaud and colleagues' release of their analysis code serve as an open invitation to join the effort.

Randy L. Gollub writing in Nature 604, 633-634 (2022). Original research: Nature 604, 697-707 (2022).



### April

### Nuclear physics

**Cryogenic mastery** aids bid to spot matter creation

is made almost entirely of matter, with nearly no antimatter in sight. However, laboratory and particlecollider experiments have so far observed the creation of matter and antimatter in equal parts. Big Bang theories that aim to explain the cosmic-matter imbalance predict that matter could be generated without antimatter in a 'little bang', during an ultra-rare nuclear process called neutrinoless double-ß decay. The CUORE Collaboration reports the most sensitive search yet for this type of decay using isotopes of tellurium. The decay was not observed, but the engineering feat was remarkable - requiring the stable operation of more than one tonne of experimental apparatus, at cryogenic temperatures close to 10 millikelvin over several years

### 2022 Nature Highlight!

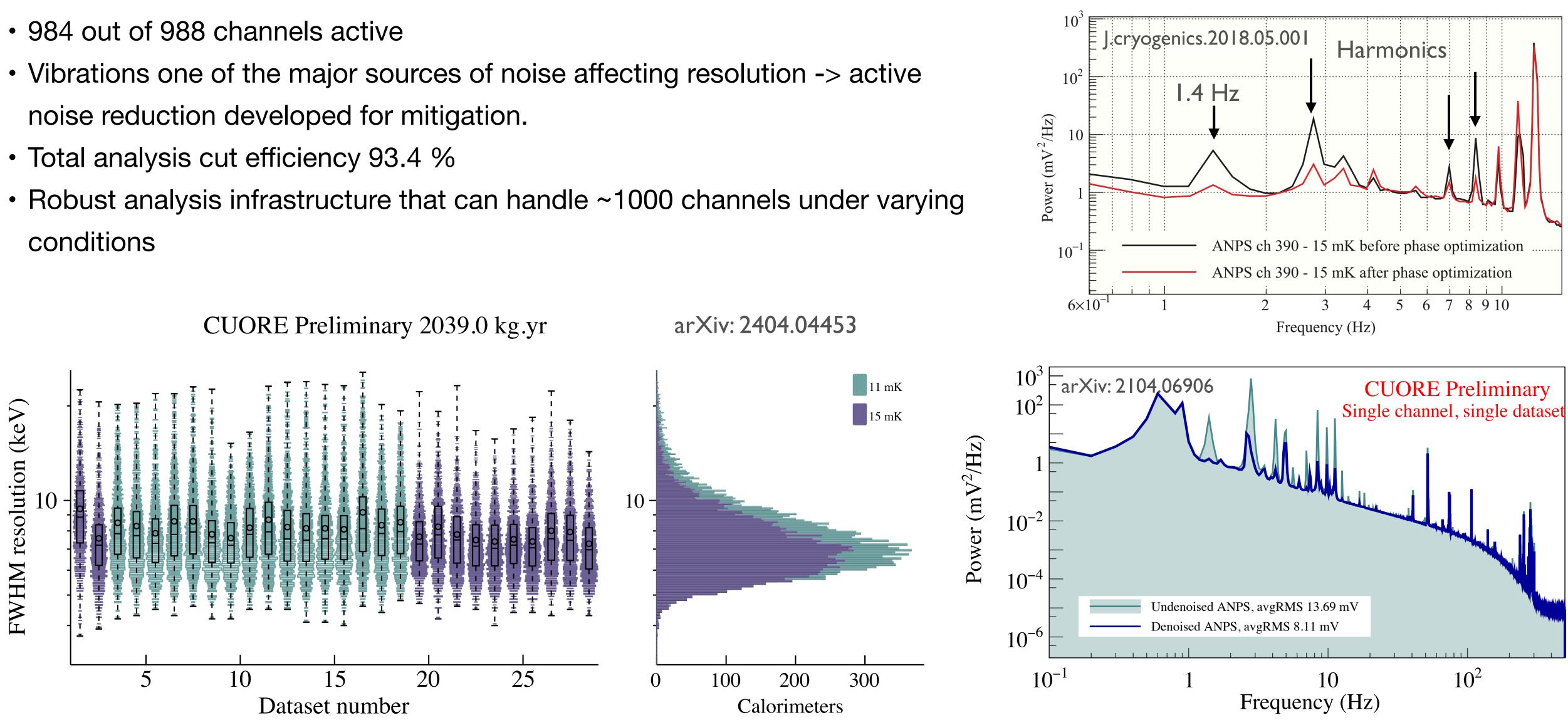
644 | Nature | Vol 612 | 22/29 December 2022

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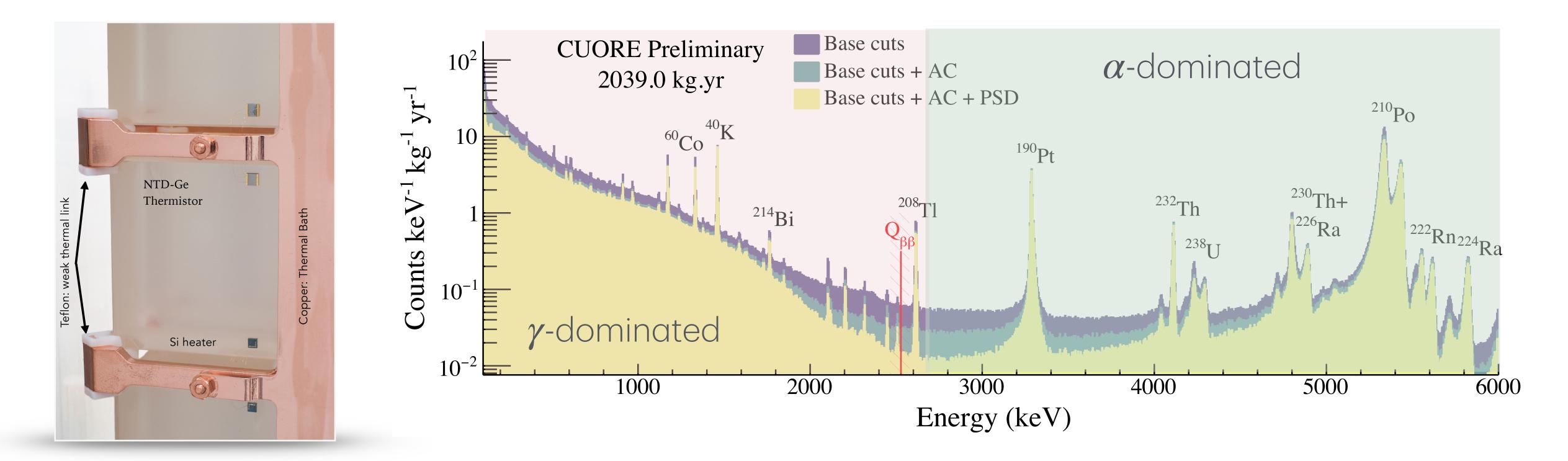
# **CUORE - Detector Performance**

- noise reduction developed for mitigation.
- conditions





### CUORE - Spectrum



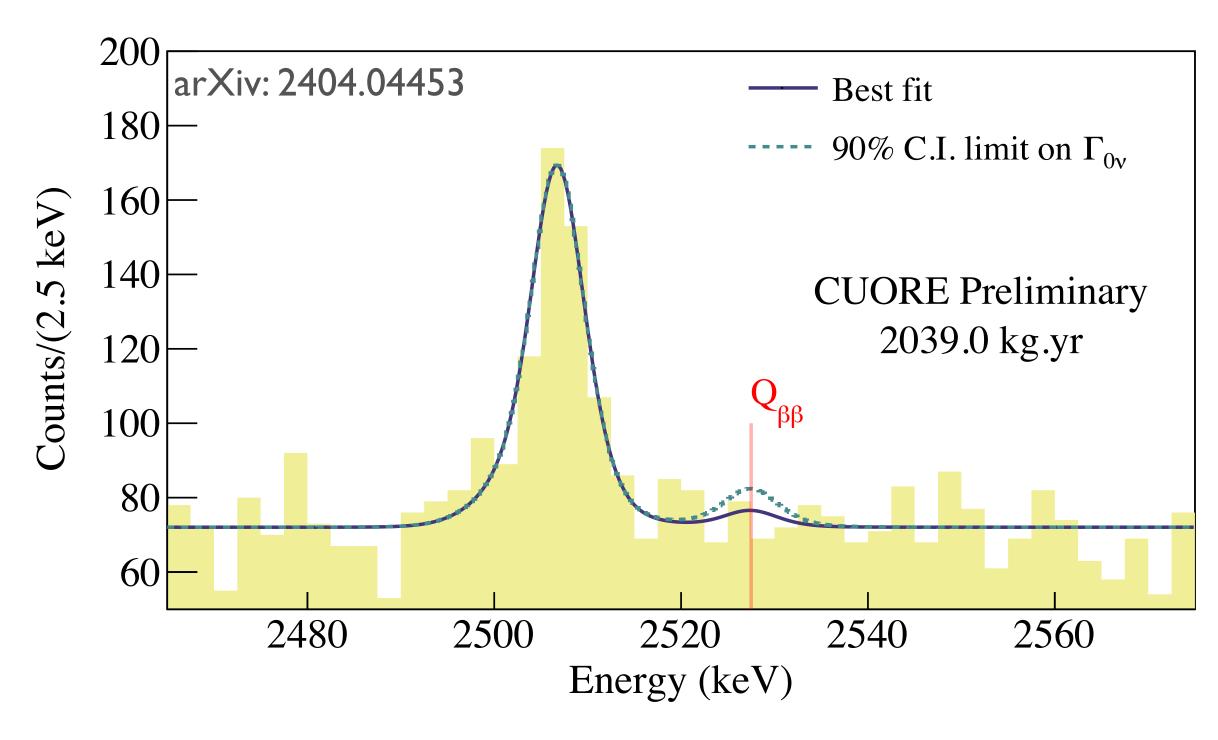
- Detector element detects only heat
- No particle ID

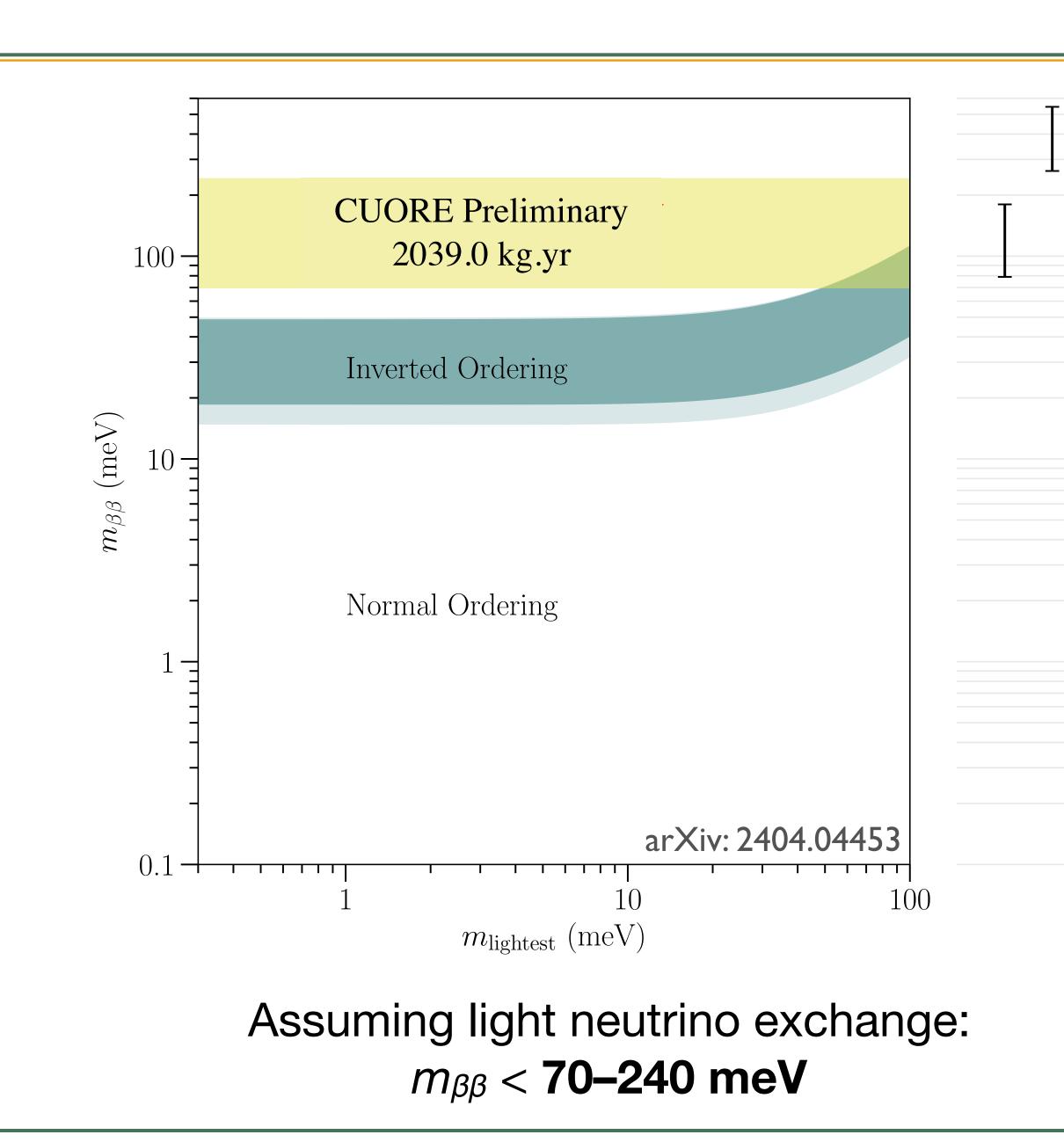
 ~ 90% of background in the ROI is from degraded alphas Muons are the next dominant background.



### CUORE - Spectrum

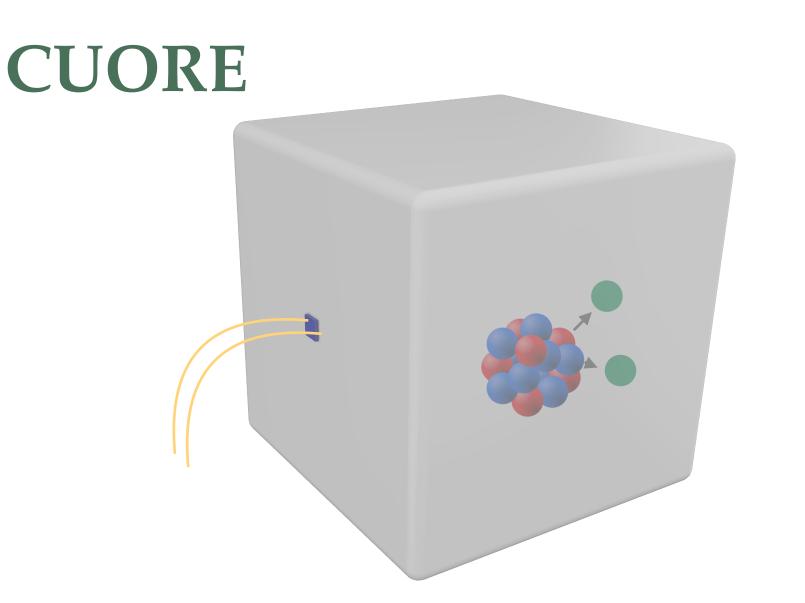
- Fit performed over [2465, 2575] keV with fit parameters:
  - $0\nu\beta\beta$  decay rate @  $Q_{\beta\beta}$
  - <sup>60</sup>Co sum peak amplitude
  - Background (flat)
- No evidence of  $0\nu\beta\beta$
- $T_{1/2}^{0\nu}$  > **3.8 x10**<sup>25</sup> yr (90% C.I.)







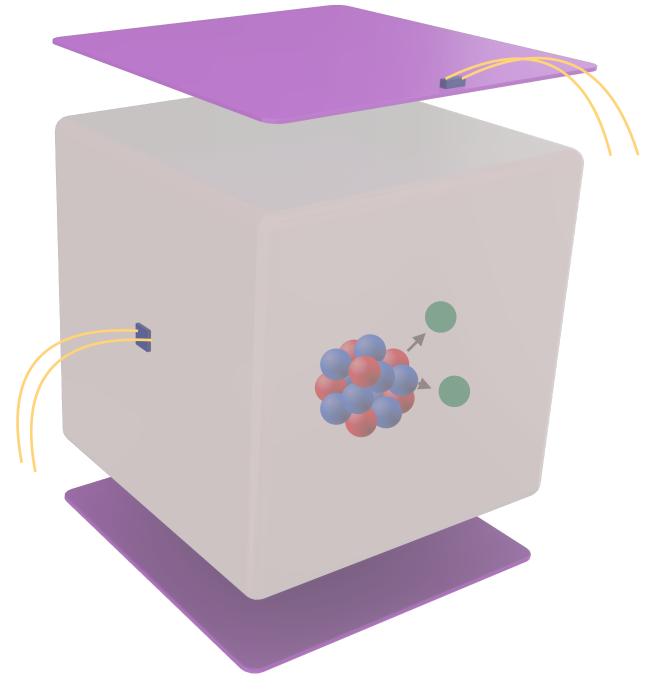
# CUPID - CUORE Upgrade with Particle ID



Operated at ~ 10 mK

- TeO<sub>2</sub> crystals (<sup>130</sup>Te isotope of interest)
- $Q_{\beta\beta} \sim 2528 \text{ keV}$
- Measures only heat (phonons) signals.
- Background Index :  $1.42 \times 10^{-2}$  cnts/ (keV. kg. y)

**CUPID** 



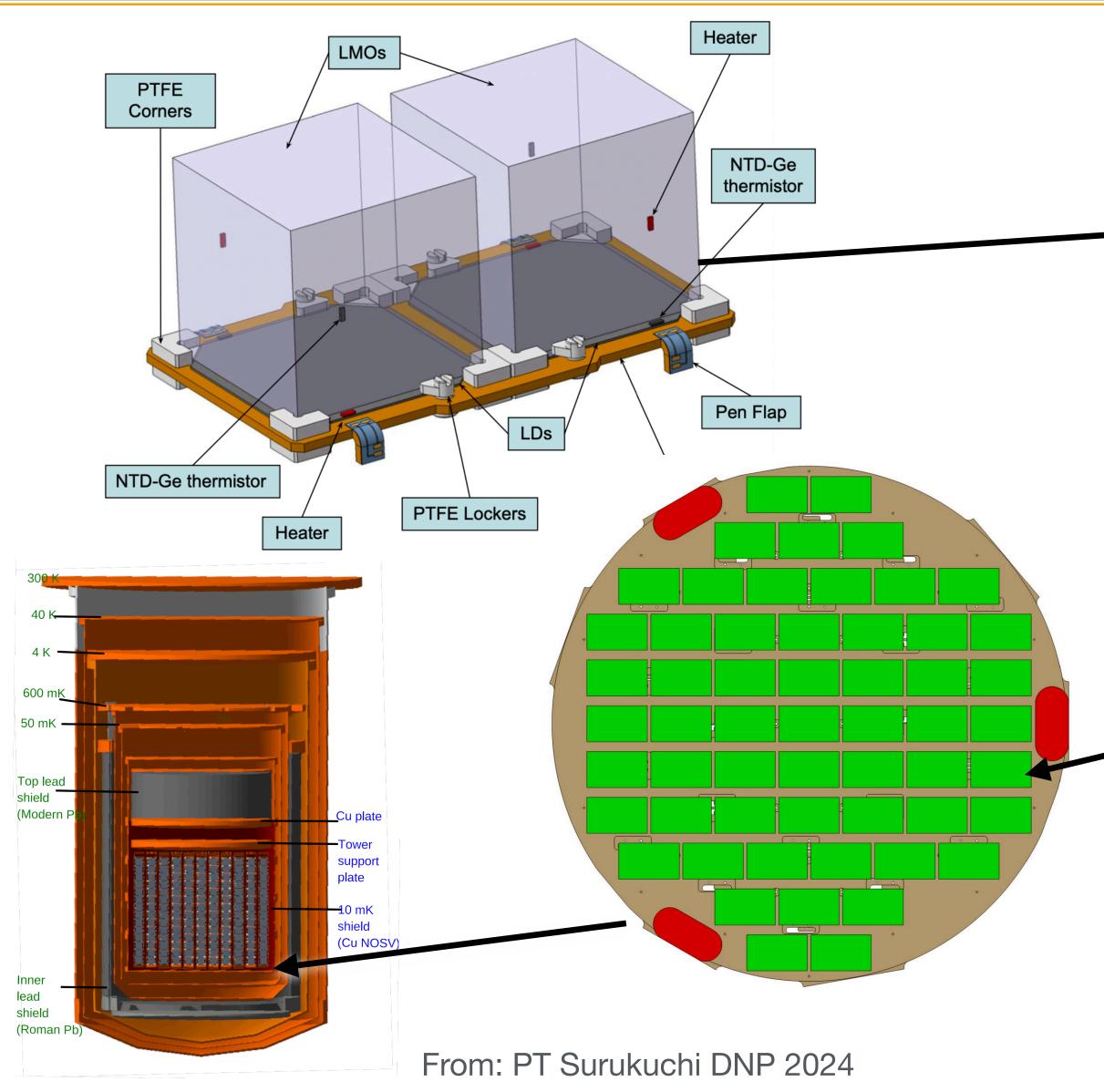
- Scintillating Li<sub>2</sub><sup>100</sup>MoO<sub>4</sub> crystals (<sup>100</sup>Mo isotope)
- $Q_{\beta\beta} \sim 3034 \text{ keV}$
- Simultaneous measurement of absorbed heat and emitted light.
- Background Index  $< 1 \times 10^{-4}$  cnts/ (keV. kg. y)

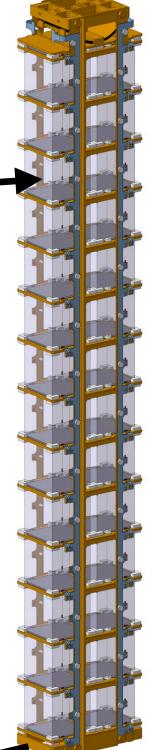




# CUPID - Design

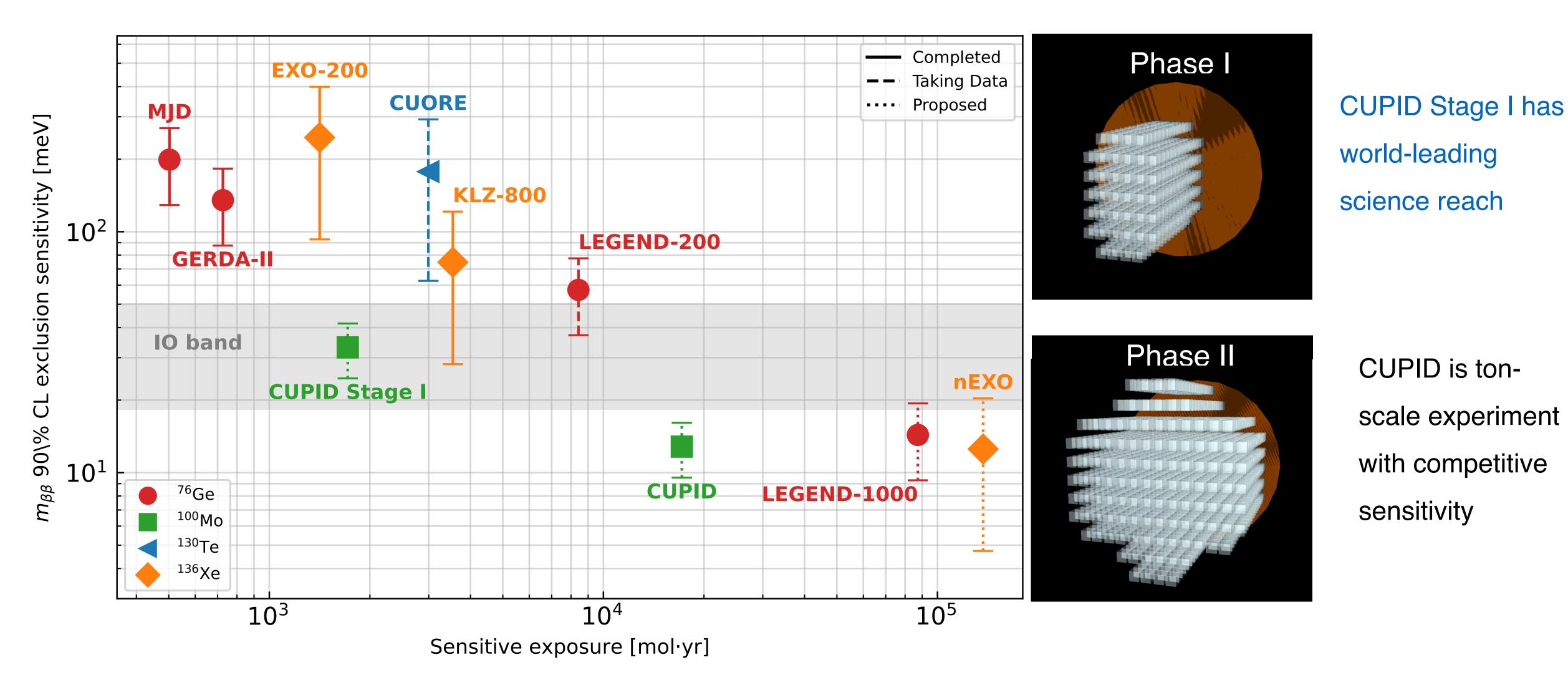
- Single detector:
  - Absorber: 4.5x4.5x4.5 cm<sup>3</sup> Li<sub>2</sub><sup>100</sup>MoO<sub>4</sub> crystals (> 95% enrichment)
  - Builds on the experience of CUORE and utilizes its cryogenic infrastructure.
  - Light detector: Ge light detector with Neganov-Trofimov-Luke amplification
- Detector array:
  - Modified detector array structure
  - 57 towers (total of 1596 crystals)
  - 240 kg of <sup>100</sup>Mo
- Addition of muon and neutron shields.







### **CUPID - Sensitivity**

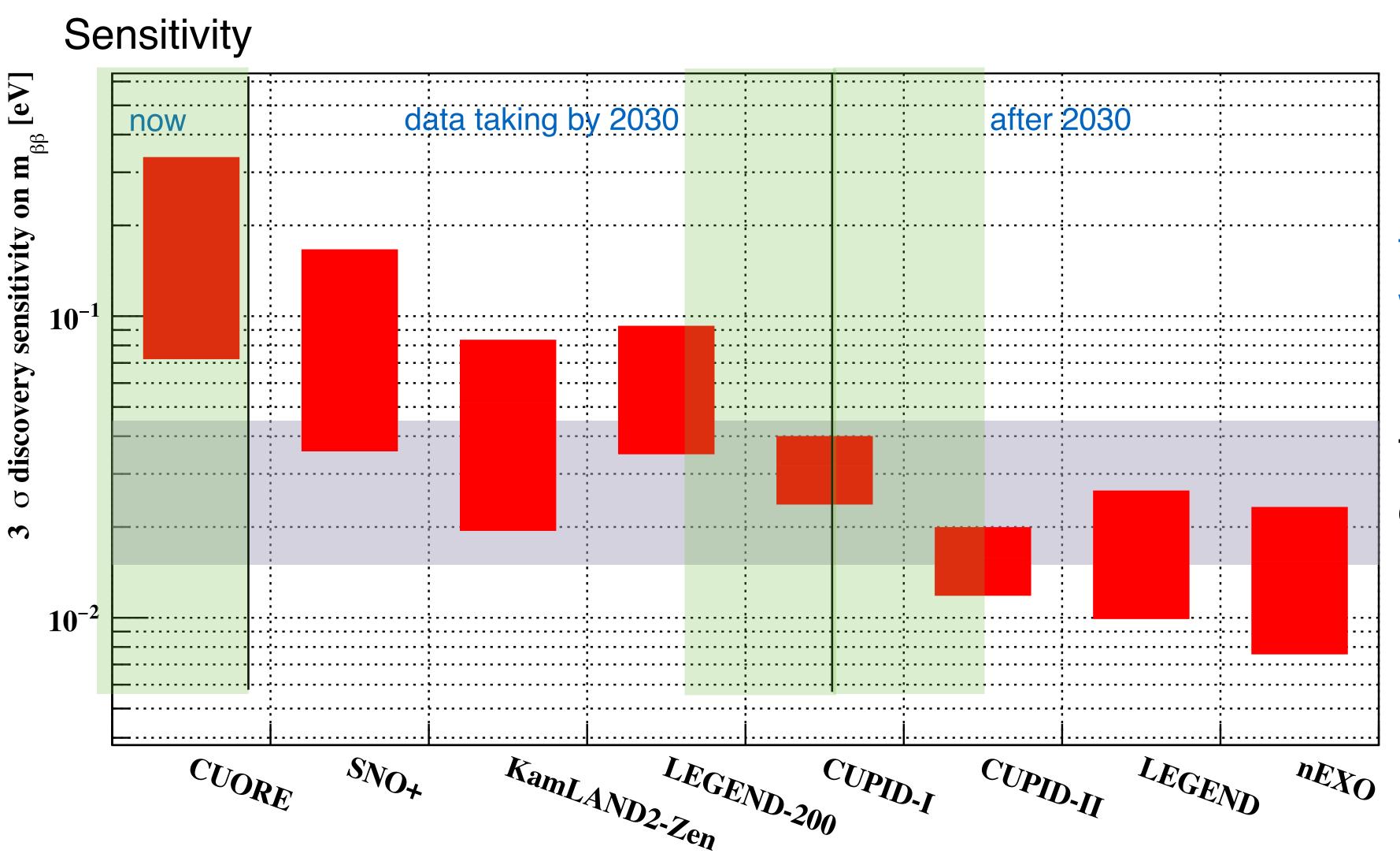








# **CUPID - Discovery Sensitivity**



Staged deployment enables first science data by 2030 with CUPID-I

Ton-scale experiment with competitive sensitivity







# What if there is a discovery?

Vivek Singh, UC Berkeley



# Searches in Multiple Isotopes Necesary

- **Confirmation and Understanding:** Studying 0vββ in multiple isotopes is crucial not only to confirm a discovery but also to pinpoint the specific mechanism driving the decay.
- **Diverse Sensitivities:** Different isotopes offer varying levels of sensitivity to 0vββ and background rejection capabilities, providing a more comprehensive search.
- Independent Uncertainties: Each isotope-based experiment has unique systematic uncertainties, allowing for crosschecks and improved reliability of results.
- Model Validation: Measurements across different isotopes help validate theoretical nuclear models and refine our understanding of the underlying physics.
- Complementary Approaches: Even experiments with limited scalability can provide valuable information, such as precise measurements of two-neutrino double beta decay and studies of specific isotopes like <sup>48</sup>Ca for benchmarking nuclear models.







### Angular Correlation, single electron spectra ...

### SuperNEMO

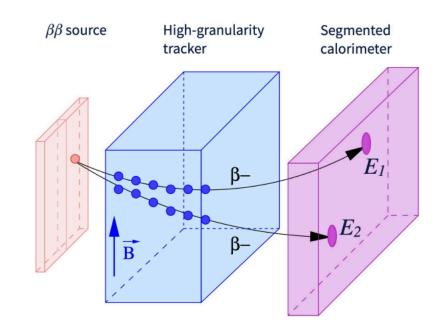


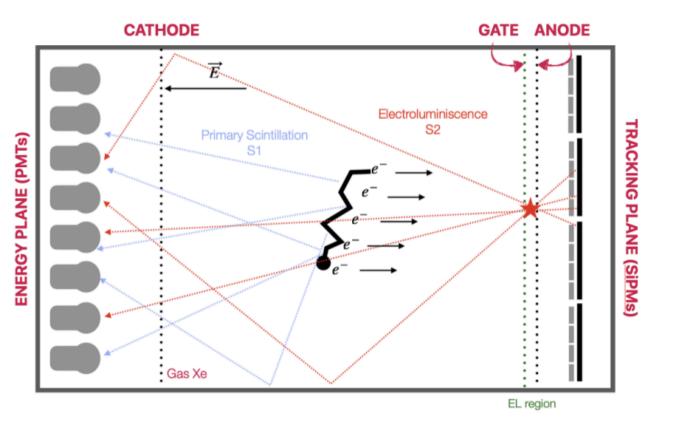
Figure 1: The NEMO technique

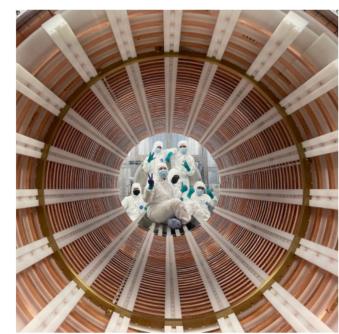
Patrick, C., and SuperNemo Collaboration. TAUP Proceedings 2024.



 SuperNEMO Demonstrator at LSM, France taking data with the full tracker and calorimeter from a 6.3kg Se-82 doublebeta source.









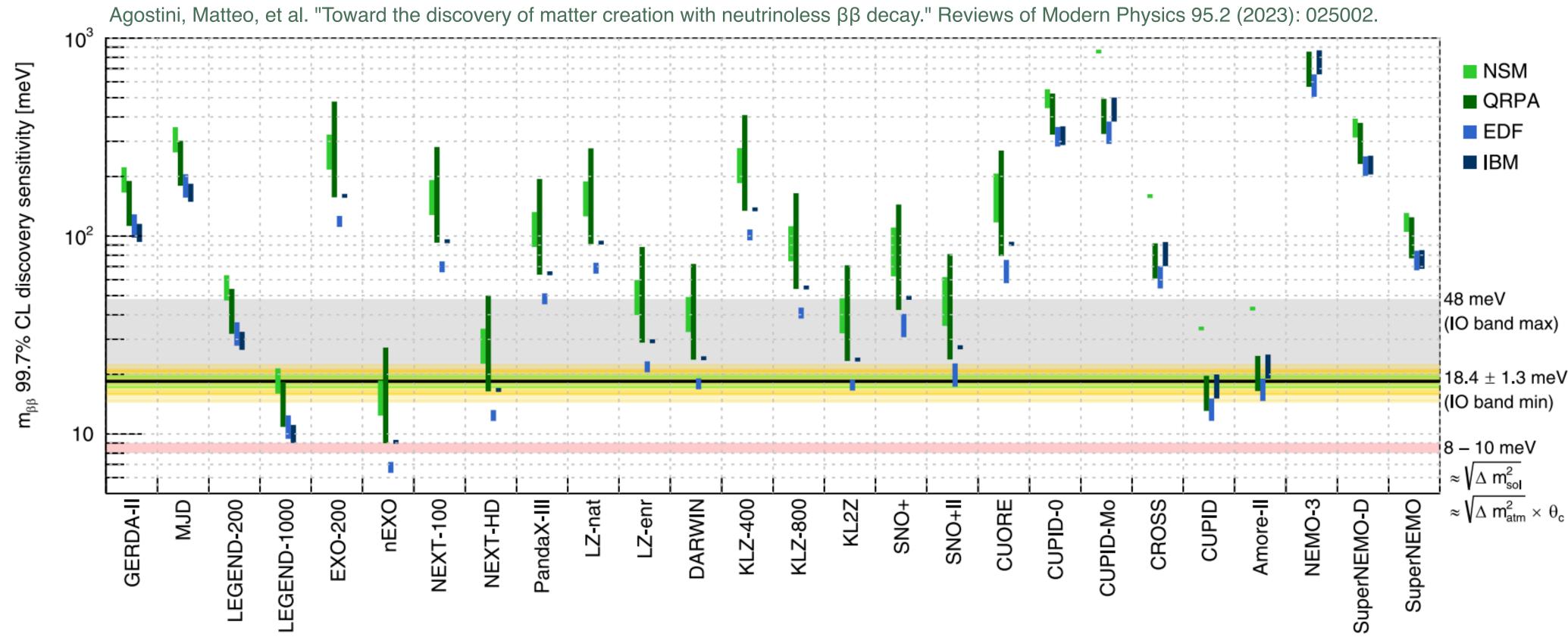
PoS EPS-HEP2023 (2024) 169

- NEXT is a high pressure gas TPC using enriched Xe
- Event topology through tracking
- NEXT-100 installation at Canfranc
- Plans to scale from 100 kg to 1 ton of Xenon
- R&D towards Ba tagging

Can shed light on the mechanism of the decay.



### Outlook



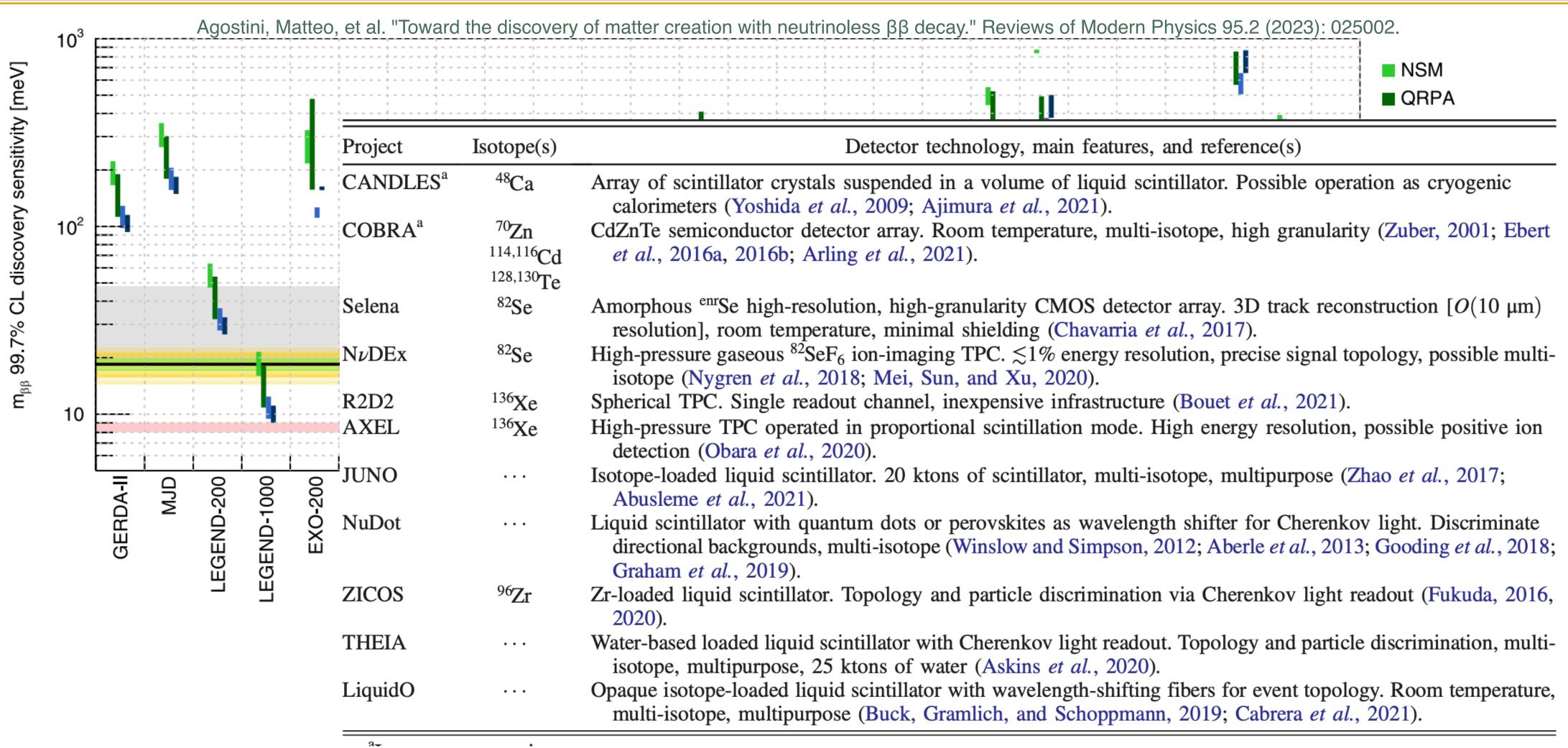
The next generation ~ton scale experiments aim for discovery (if IH) or push slightly beyond

Vivek Singh, UC Berkeley





### **Outlook**



### Plenty more experiment driving the development of cutting-edge detector technologies.



## Summary

- implications for particle physics.
- cosmology and astrophysics.
- This field requires expertise from both nuclear and particle physics
  - Vibrant experimental program worldwide.
  - experimental results.
  - decay can help refine these calculations.

0vββ would prove that neutrinos are their own antiparticles, a unique property with profound

• A 0vββ observation would allow us to finally measure the absolute mass of neutrinos, crucial for

• 0vββ is a lepton-number violating process, signaling new physics beyond our current understanding.

Accurate theoretical calculations of nuclear matrix elements are essential for interpreting

• Various experimental probes, like charge exchange reactions, muon capture, double gamma





