

High Brightness Electron Beams for Radiation Sources

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On behalf of BD & ICS Milano Group

Outline

❖ Basic concepts in Accelerator Physics & Beam Dynamics (BD) in the frame of **HBB** machines:

- Relativity – beam Rigidity
- Main beam optics: **Dipoles**, **Quadrupoles**, Sextupoles, **Solenoids**
- The transverse phase-space, **emittance** & **brightness**. The **longitudinal phase-space**
- Accelerators **cavities** basic concepts
- Sketch of a High Brightness Linear Accelerator (**LINAC**) for **Thomson/Compton sources**
- **STAR-II** upgrade, a real LINAC; some **BD** and machine images

❖ Useful codes for space charge dominated beam simulations

- The **Astra** code & examples of its use
- Examples of the Astra Use
- **Cain**, a Montecarlo quantum code to simulate electron-photon interactions & examples
- **Simulation** of Three different scattering cases, what will be done this coffee-break in the **laboratory**

Relativity basic concepts

Important:

When we speak of beam particle energy in an accelerator,
we refer to Kinetic Energy! (unless specified)

Mostly we use the I.S., with few exceptions:

- The Beam energy: eV (keV, MeV, ...)------[1eV = 1.6×10^{-19} J]
- Mass: eV/c² -----[Proton= 1.67×10^{-27} kg → 938 MeV/c²
-----[Electron = 9.11×10^{-31} kg → 0.511 MeV/c²]
- Momentum: eV/c -----[Proton @ $\beta=0.9$ → 1.94 GeV/c]

$$\beta \equiv \frac{v}{c} = \frac{pc}{E}$$
$$\gamma \equiv \frac{1}{\sqrt{1 - \beta^2}} \rightarrow \beta\gamma = \sqrt{\gamma^2 - 1}$$

$$\text{momentum } p = \gamma m v$$

$$\text{total energy } E = \gamma m c^2$$

$$\text{kinetic energy } K = E - m c^2 = m c^2 (\gamma - 1)$$

$$E = \sqrt{(m c^2)^2 + (p c)^2}$$

From an ATP Tennis player like Matteo Berrettini its service is about 90 J i.e. (5.6×10^{20} eV)

Extreme-energy cosmic ray (EECR) are typically protons with energy $\geq 10^{18}$ (0.8 J) (LHC beams: 7×10^{12})

Beam Rigidity

It's a relation between:
radius - magnetic field - momentum - charge
HOW HARD (or EASY) is it to deflect PARTICLES?

Centrifugal force

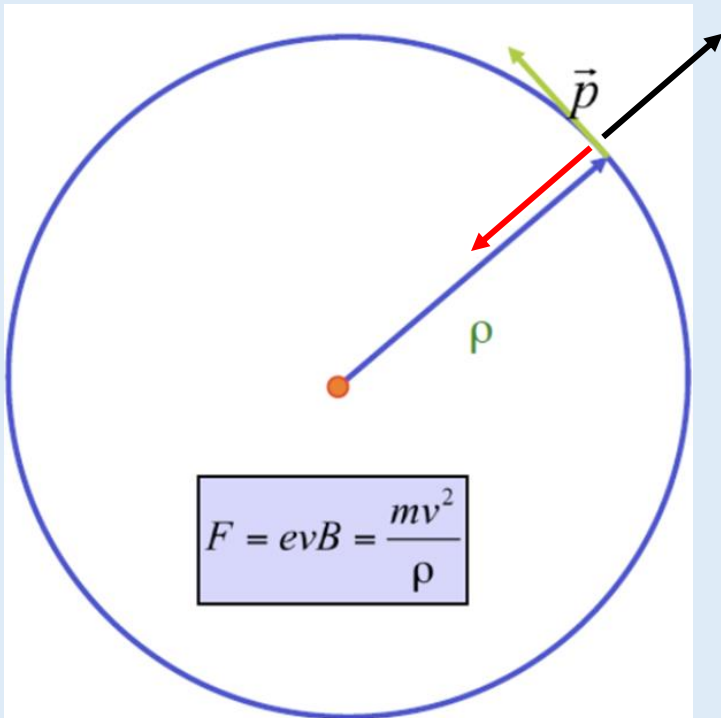
$$\mathbf{F}_{CF} = -m\mathbf{a}_{CF} = -m[\boldsymbol{\omega} \times (\boldsymbol{\omega} \times \mathbf{r})] = -m\omega(\omega^2 r)\hat{\mathbf{r}} = -m\frac{v_t^2}{r}\hat{\mathbf{r}}$$

Centripetal force (or generally **Lorentz Force**)

$$\mathbf{F}_{CP} = q(\mathbf{v} \times \mathbf{B})\hat{\mathbf{r}}$$

$$B\rho = \frac{p}{q}$$

$$B\rho [Tm] \approx 3.33 \frac{p \left[\frac{GeV}{c} \right]}{q [C]}$$



	LEP	LHC
ρ [m]	3096.175	2803.95
p_0 [GeV/c]	104	7000
B [T]	0.11	8.33

Room-temperature
coils

Superconducting
coils

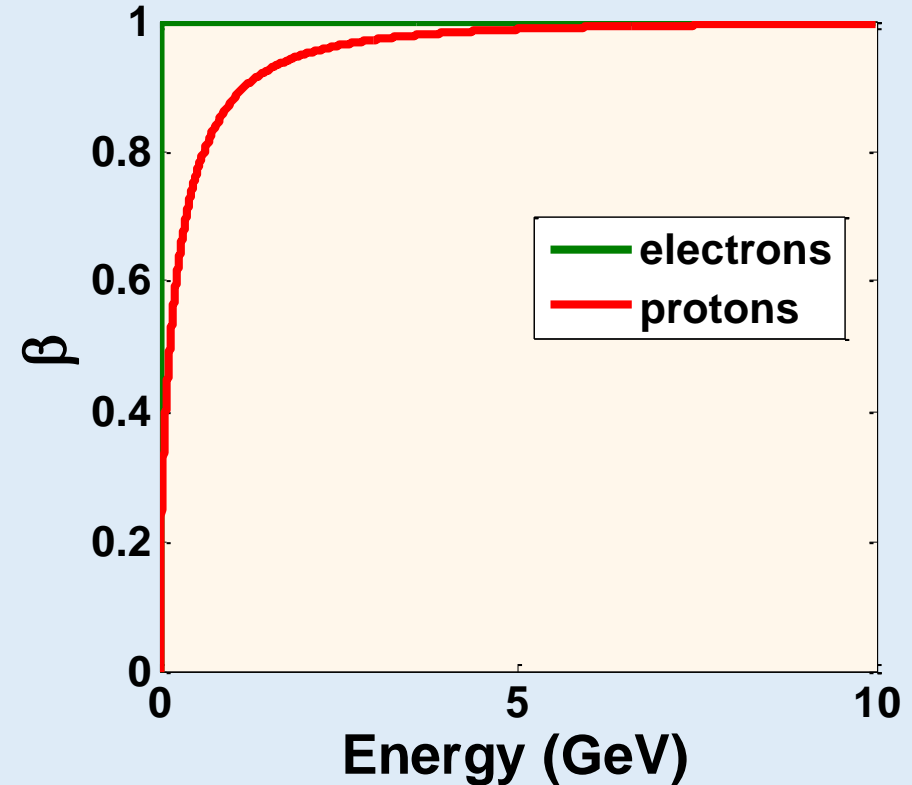
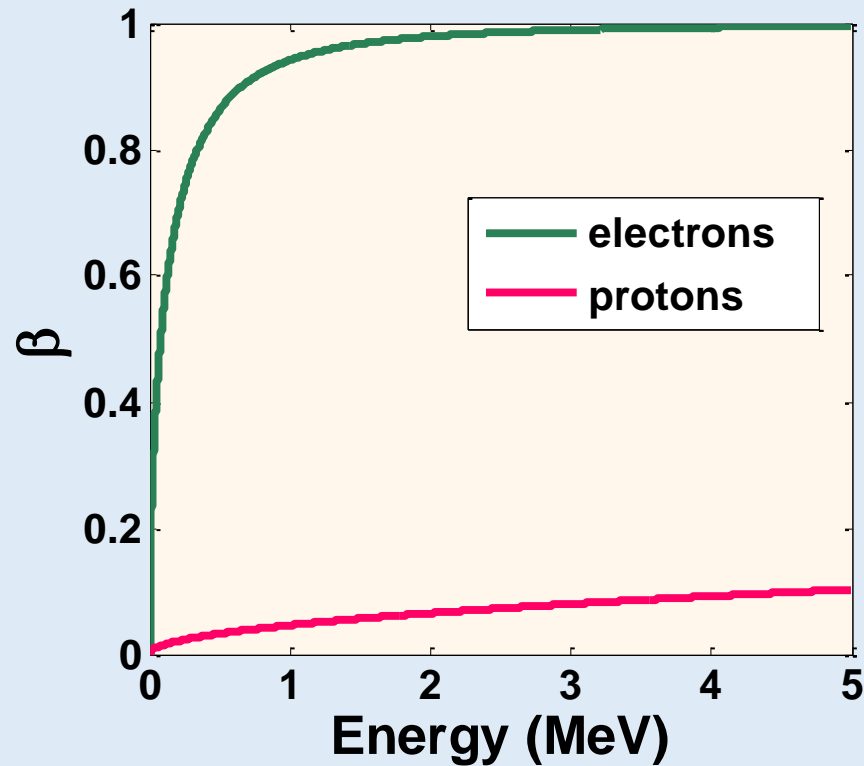
$$8 [T] \approx 0.11 [T] \cdot \frac{p_0 [\text{LHC}]}{p_0 [\text{LEP}]} \text{ Linear scaling from LEP}$$

Particle velocity as function of kinetic energy 1/2

$$\beta = \frac{v}{c}$$

$\beta = 1$ Particle at light velocity c

$$\beta\gamma = \sqrt{\gamma^2 - 1}$$
$$\beta = \sqrt{1 - \frac{1}{\gamma^2}}$$
$$\gamma = 1 + \frac{k}{mc^2}$$



Electrons are ultra-relativistic @ few MeV
Protons @ few GeV (mass ~ 2000 times electron mass)

Particle velocity as function of kinetic energy 2/2

$$E_0 = 0.511 \text{ MeV} \text{ or } 938.27 \text{ MeV}$$

$$E_{tot} = E_{kin} + E_0$$

$$\gamma = \frac{E_{tot}}{E_0}, \quad \beta = \sqrt{1 - \frac{1}{\gamma^2}}, \quad p = \beta E_{tot}$$

$$B\rho = 3.33 \cdot 10^{-3} p$$

@ Ultra High energy m_0 becomes negligible

Energia [MeV]	Rigidità $B\rho$ [Tm]	
	p	e^-
1	0,14	0,005
10	0,44	0,035
100	1,45	0,34
1.000	5,66	3,34
10.000	36,35	33,36
100.000	336	336
1.000.000	3335	3335

Energy lost per turn

$$\Delta U = \frac{e^2 \gamma^4}{3\epsilon_0 \rho}$$

For Proton:

$$\Delta U_p(\text{keV}) = 6.03 \frac{E(\text{TeV})^4}{\rho(\text{m})}$$

For e-beam:

$$\Delta U_e(\text{keV}) = 88.46 \frac{E(\text{GeV})^4}{\rho(\text{m})}$$

1 GeV

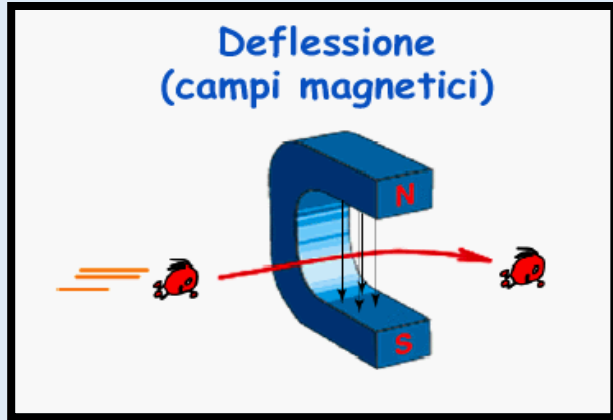
1 TeV

LHC= 7 TeV
 $\rho=5\text{km}, \Delta U=0.7 \text{ keV}$

FCC= 50 TeV
 $\rho=10\text{km}, \Delta U=5 \text{ MeV}$

Basic optics for accelerators – 1/3

Dipole fields to change beam direction

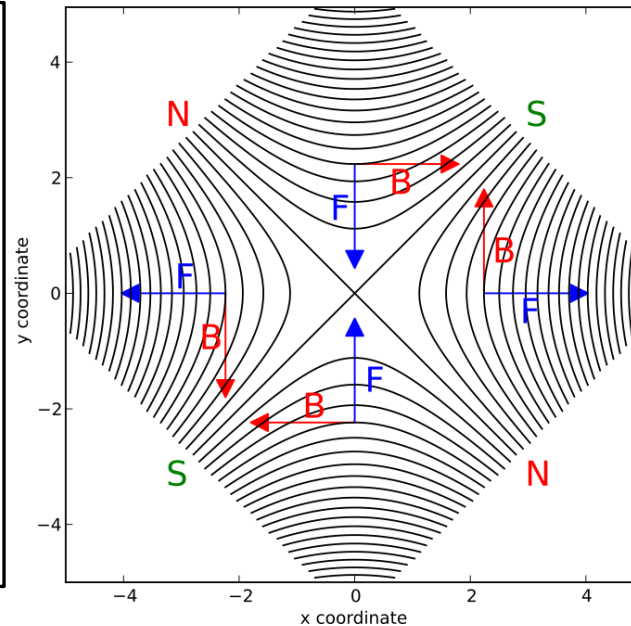
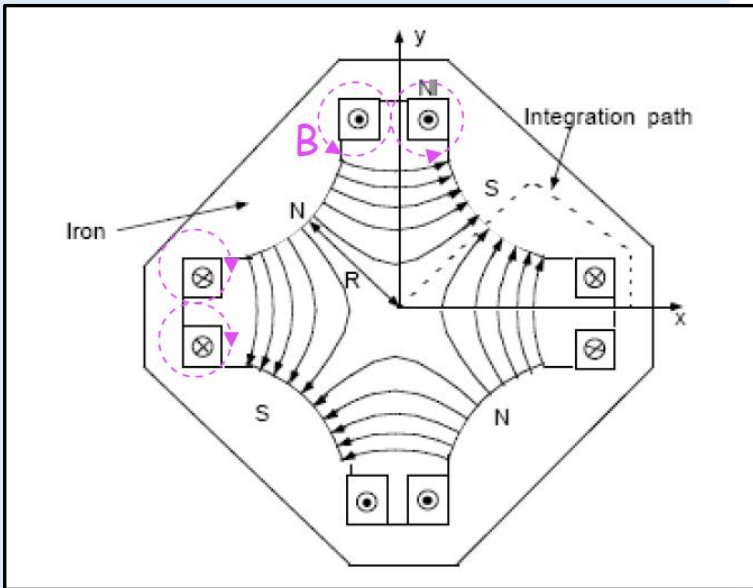


Bending Strength of a dipole

From rigidity $B\rho [Tm] \approx 3.33 \frac{p[\frac{GeV}{c}]}{q[C]}$

$$\frac{1}{\rho} \left[\frac{1}{m} \right] = \frac{0.2998 \cdot B [T]}{p[\frac{GeV}{c}]}$$

Quadrupole fields to focus or defocus the beam



The gradient is:

$$G \left[\frac{T}{m} \right] = \frac{dB_x}{dx} = \frac{dB_y}{dy}$$

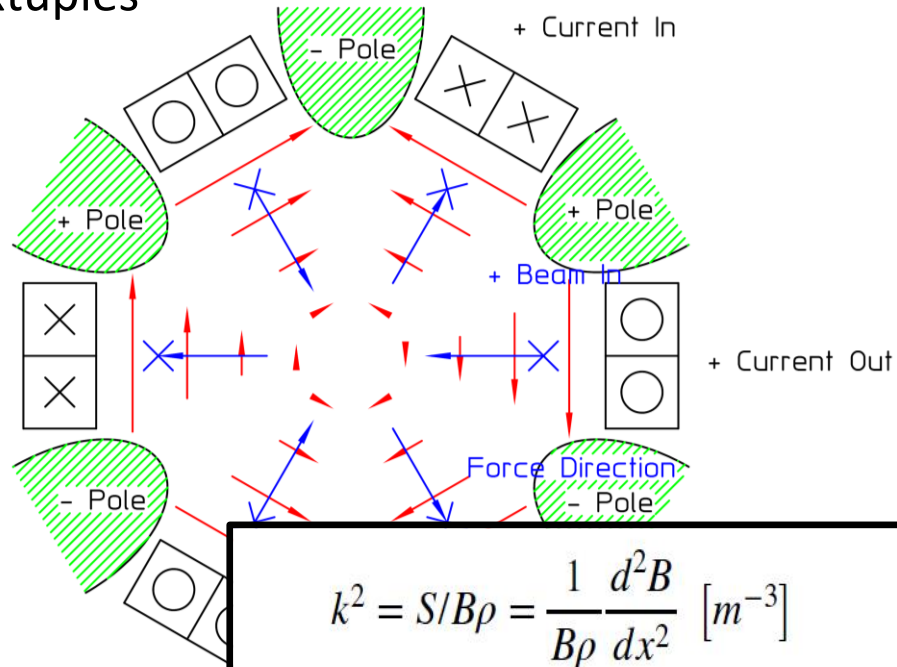
$$G \left[\frac{T}{m} \right] = \frac{B_{tip}}{R}$$

Linear Force with displacement (x):

$$F_x = qvGx, F_y = qvGy$$

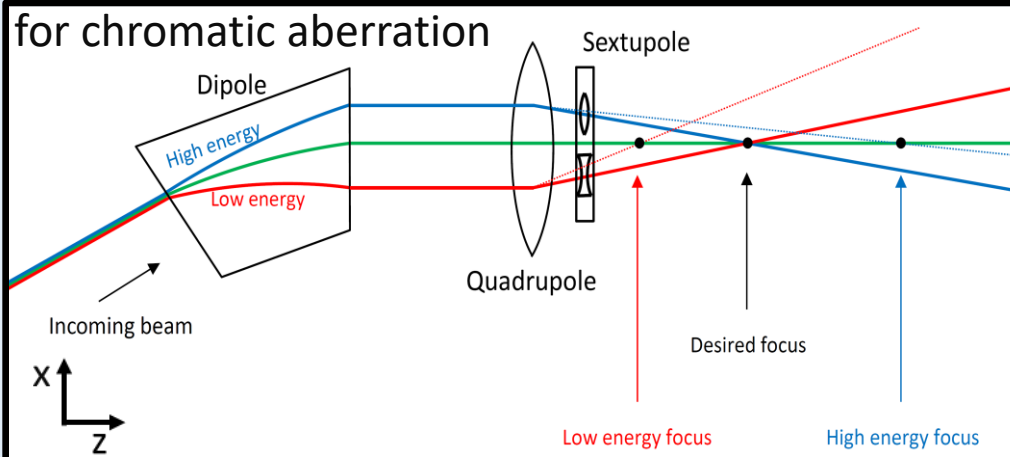
Basic optics for accelerators – 2/3

Sextuples



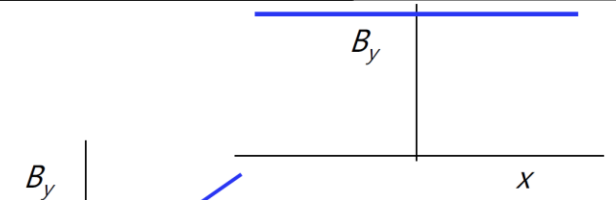
$$k^2 = S/B\rho = \frac{1}{B\rho} \frac{d^2B}{dx^2} [m^{-3}]$$

for chromatic aberration

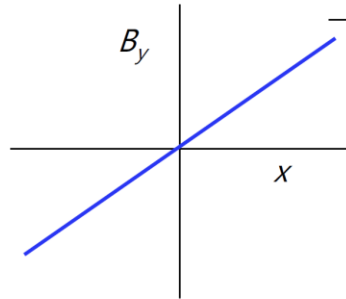


Fast recap: fields behaviour

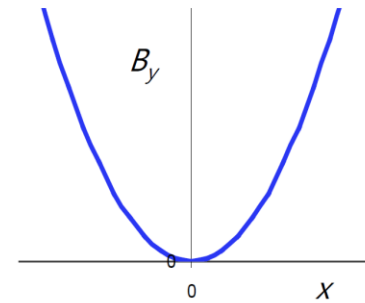
Dipole - constant field



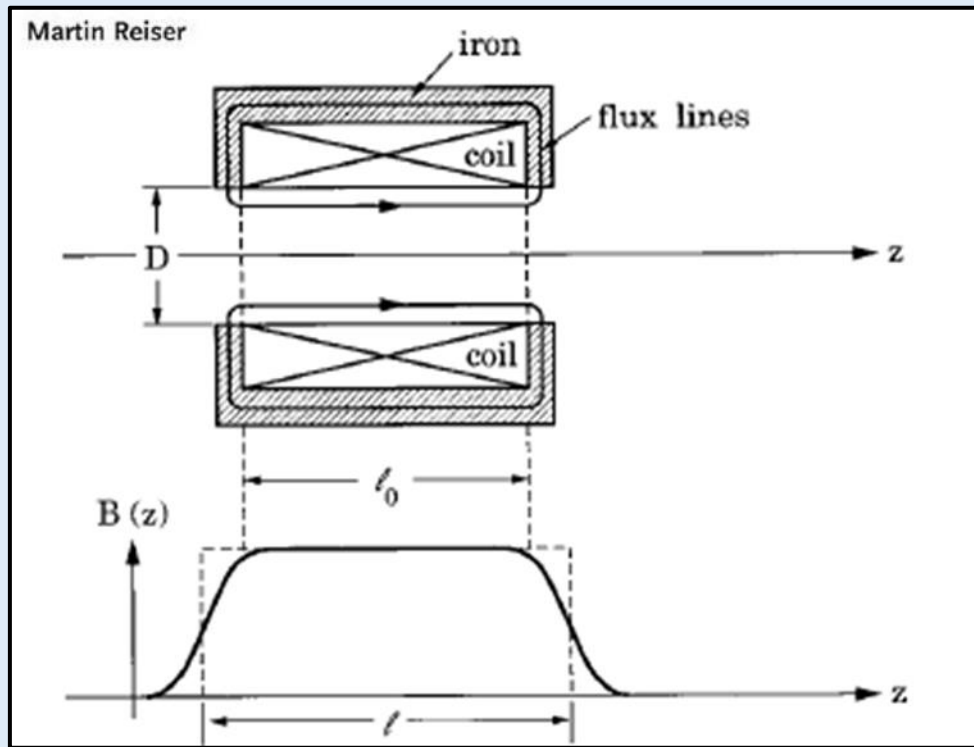
Quad - linear variation



Sextupole - quadratic variation

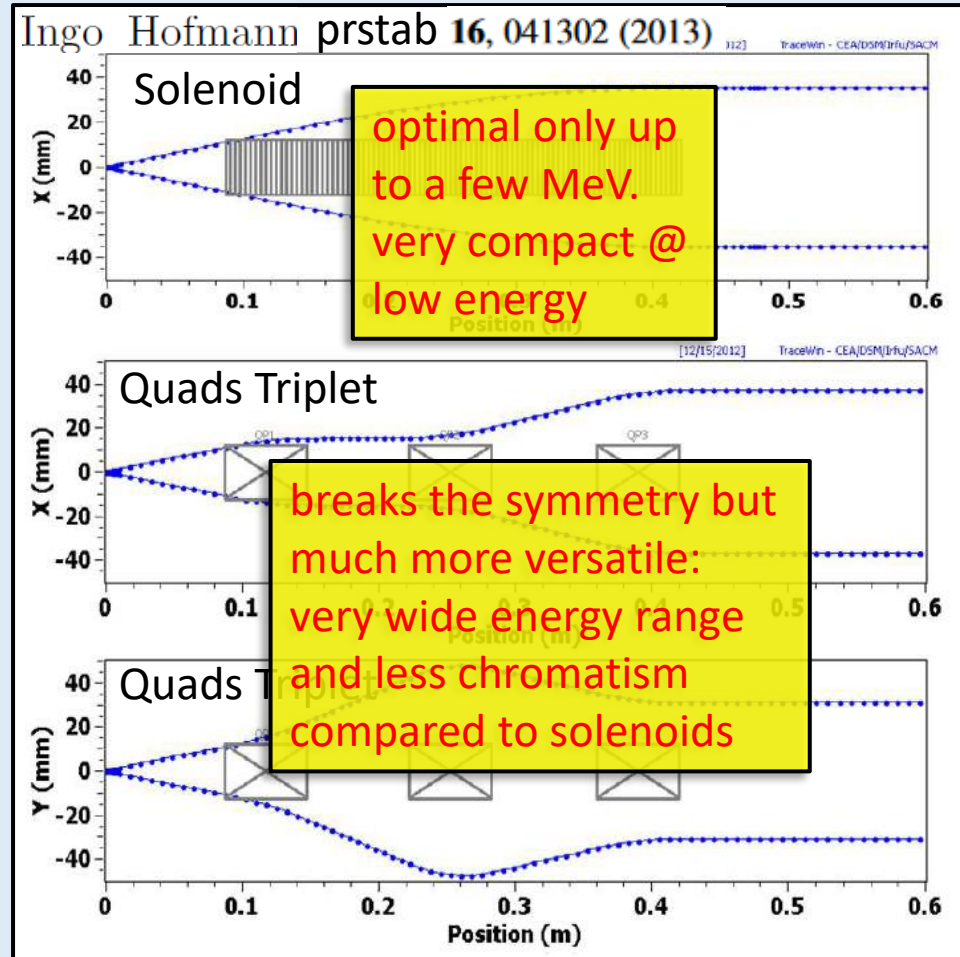


Basic optics for accelerators – 3/3



$$\frac{1}{f} = -\frac{r'}{r} = \left(\frac{q}{2mc\beta\gamma} \right)^2 \int_{z_1}^{z_2} B^2 dz$$

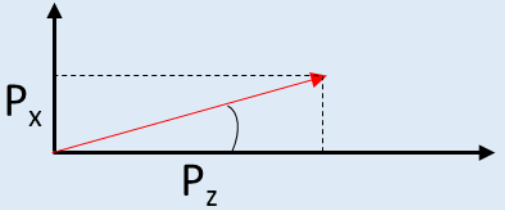
Solenoid VS Quads, Same solution



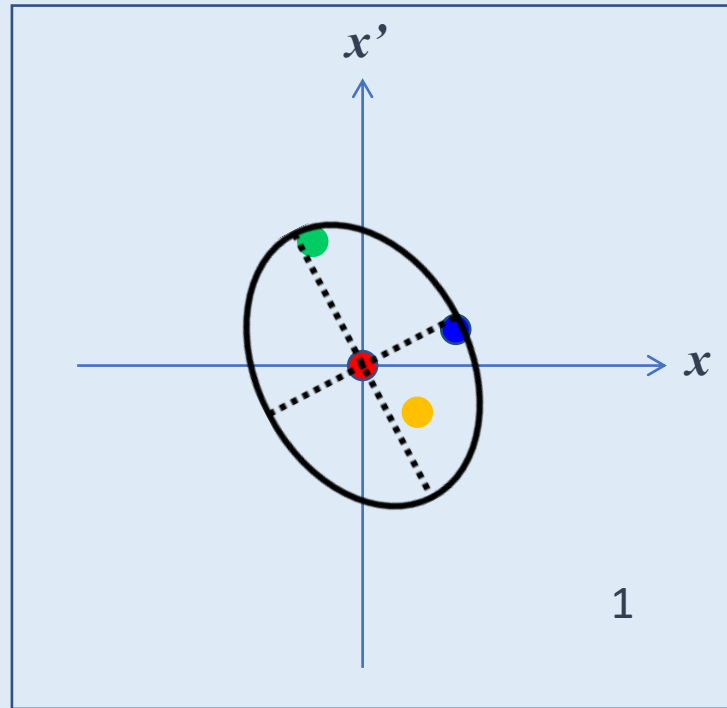
Important differences to remember:

- solenoids work well only at low beam energy: few MeV for electrons.
- solenoids give larger beam quality degradation if the beam is large (sig_x) into the optic (chromatism effects)

Bunch transverse Phase-space & emittance – 1/5

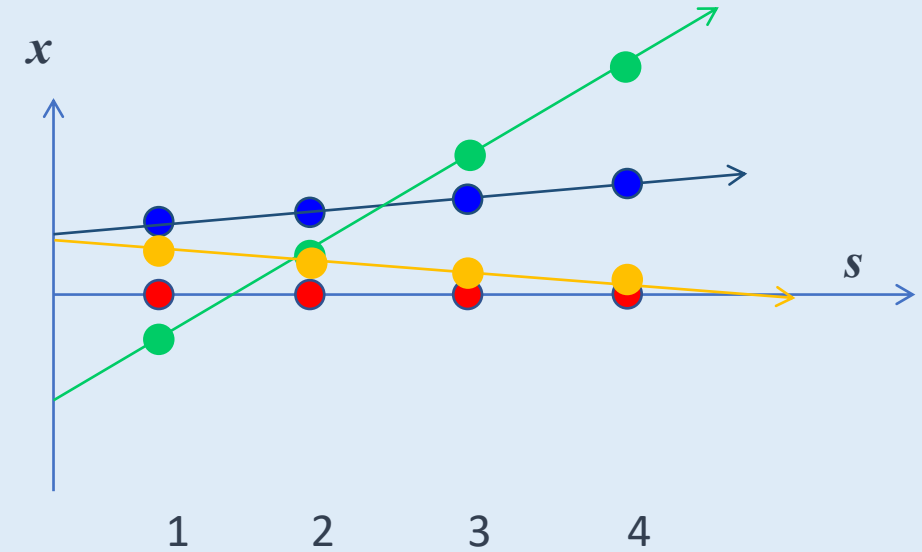


Paraxial approximation

$$x' = \frac{P_x}{P_z} \ll 1 \rightarrow \frac{P_x}{P_z} = \tan(\theta) \cong \theta$$


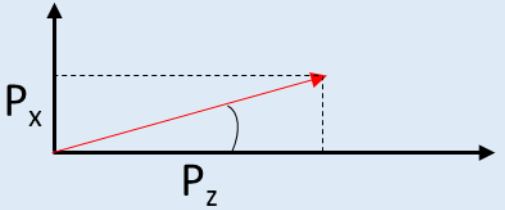
- Reference particle
- Other particles

Phase space

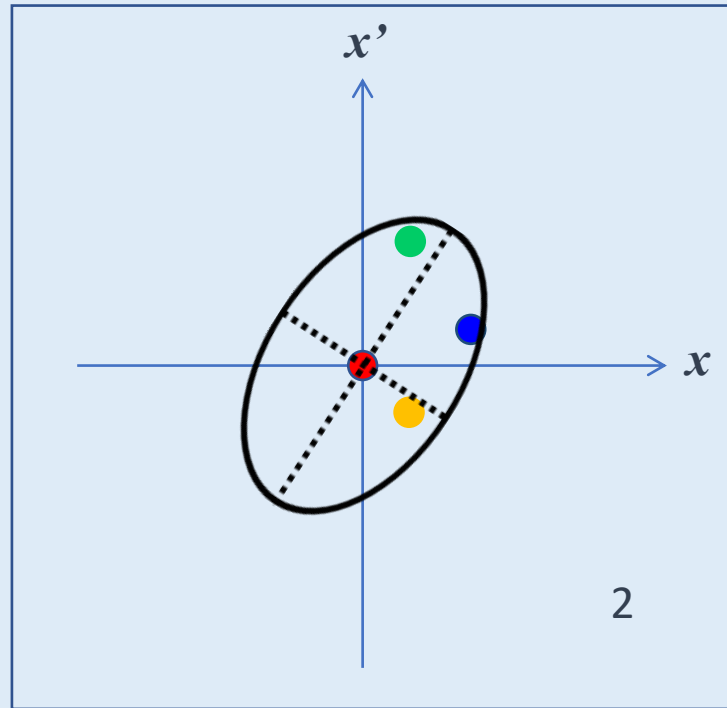


No external fields – Drift propagation

Bunch transverse Phase-space & emittance – 2/5

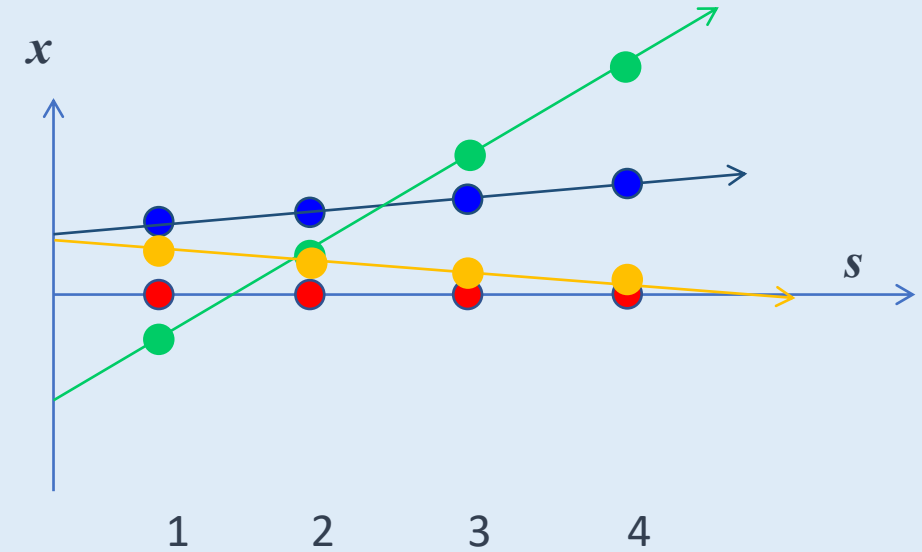


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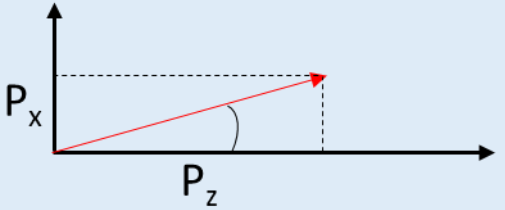
- Reference particle
- ● Other particles

Phase space

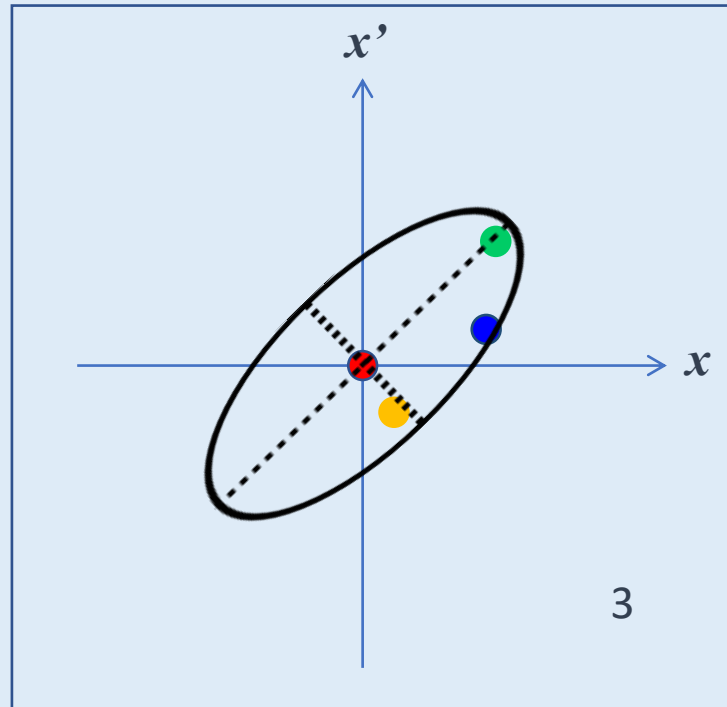


No external fields – Drift propagation

Bunch transverse Phase-space & emittance – 3/5

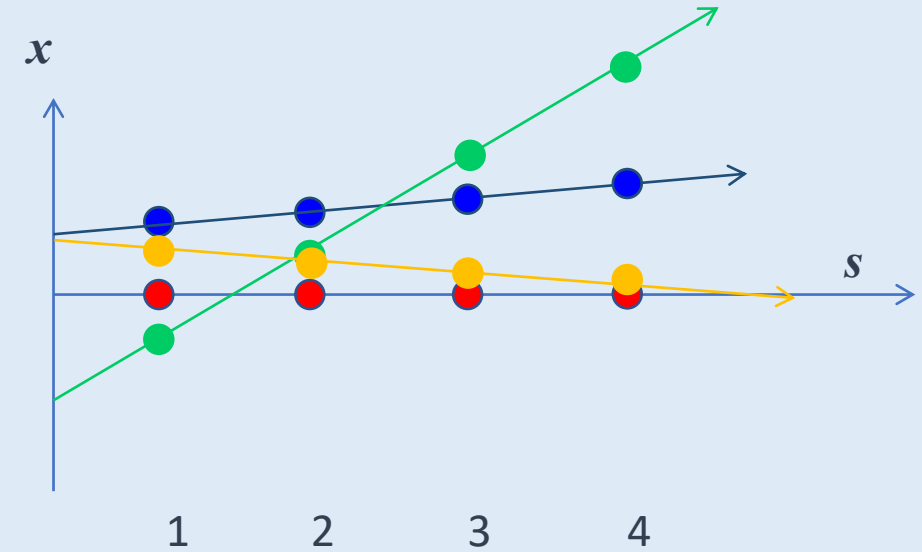


Paraxial approximation

$$x' = \frac{P_x}{P_z} \ll 1 \rightarrow \frac{P_x}{P_z} = \tan(\theta) \cong \theta$$


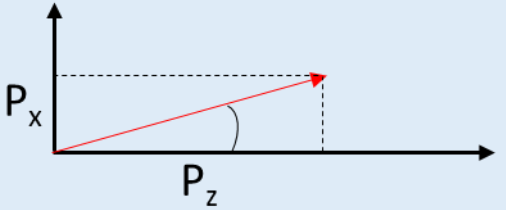
- Reference particle
- ● Other particles

Phase space

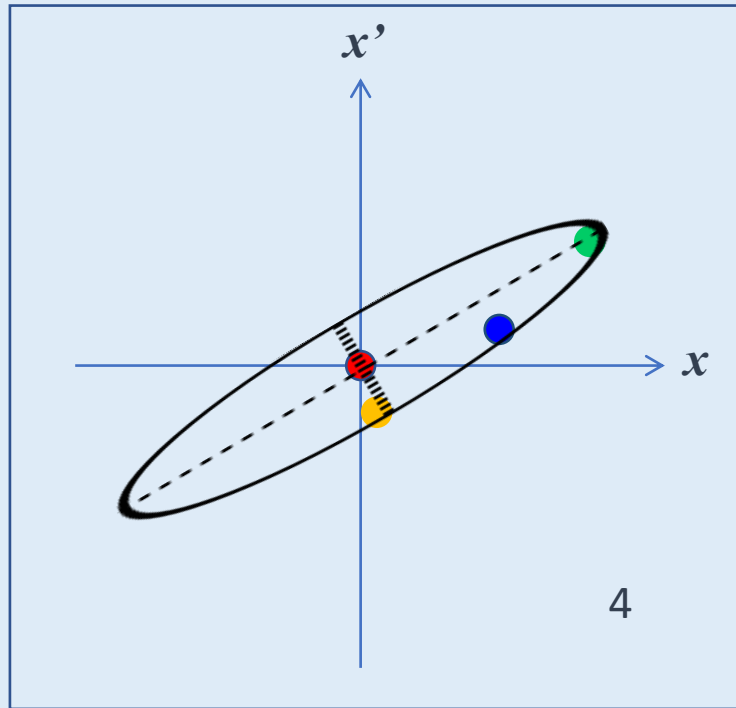


No external fields – Drift propagation

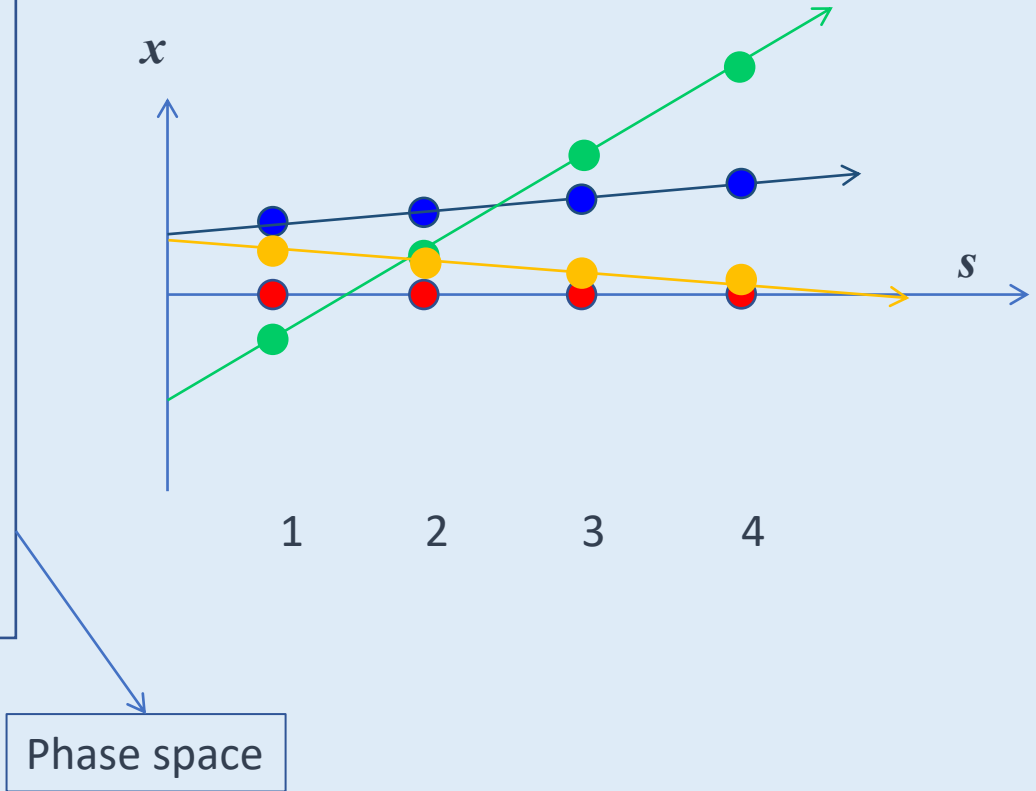
Bunch transverse Phase-space & emittance – 4/5



Paraxial approximation

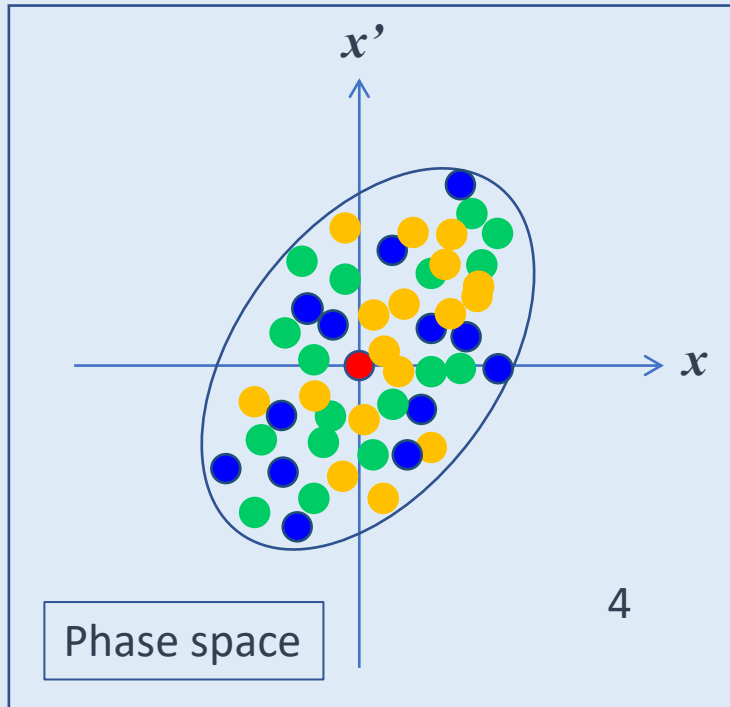
$$x' = \frac{P_x}{P_z} \ll 1 \rightarrow \frac{P_x}{P_z} = \tan(\theta) \cong \theta$$


- Reference particle
- Other particles



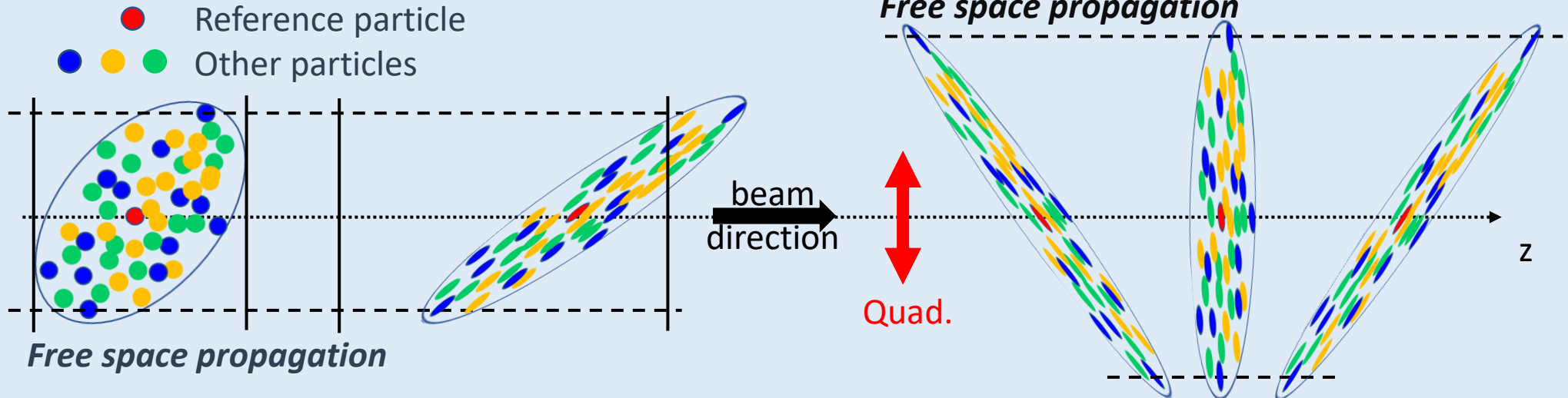
No external fields – Drift propagation

Bunch transverse Phase-space & emittance – 5/6



- Emittance = Area of phase space
- Horizontal phase space (x, x')
- Vertical phase space (y, y')
- Longitudinal phase space (Time-Energy)

For linear fields (↕) the emittance will be constant



Bunch transverse Phase-space & emittance – 6/6

rms-geometrical emittance

$$\epsilon_{n,x} = \beta\gamma \sqrt{\langle x^2 \rangle \langle x'^2 \rangle - \langle x \cdot x' \rangle^2}$$

↑ can be normalized

Ellipse equation (The Area is $\epsilon\pi$)

$$x'_j = \frac{p_{x,j}}{p_{z,j}}, y'_j = \frac{p_{y,j}}{p_{y,j}}$$

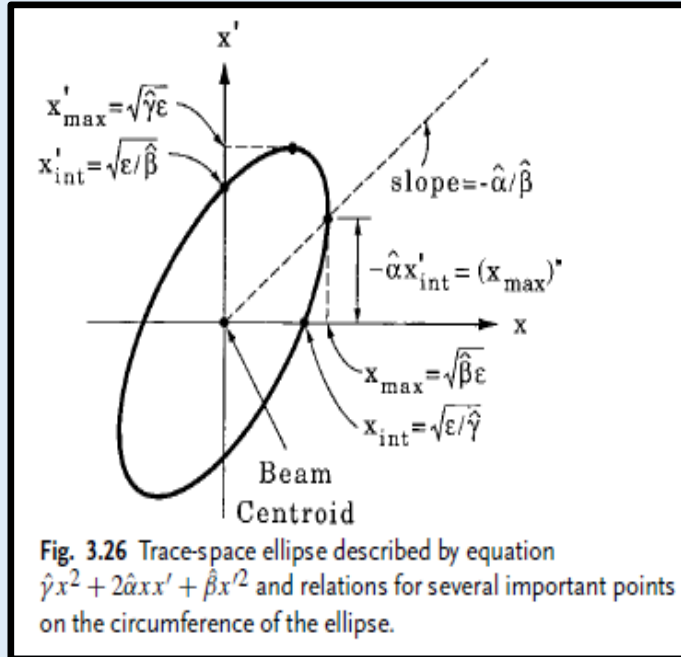
Twiss parameters

$$\beta_x = \frac{\langle x^2 \rangle}{\epsilon_x}$$

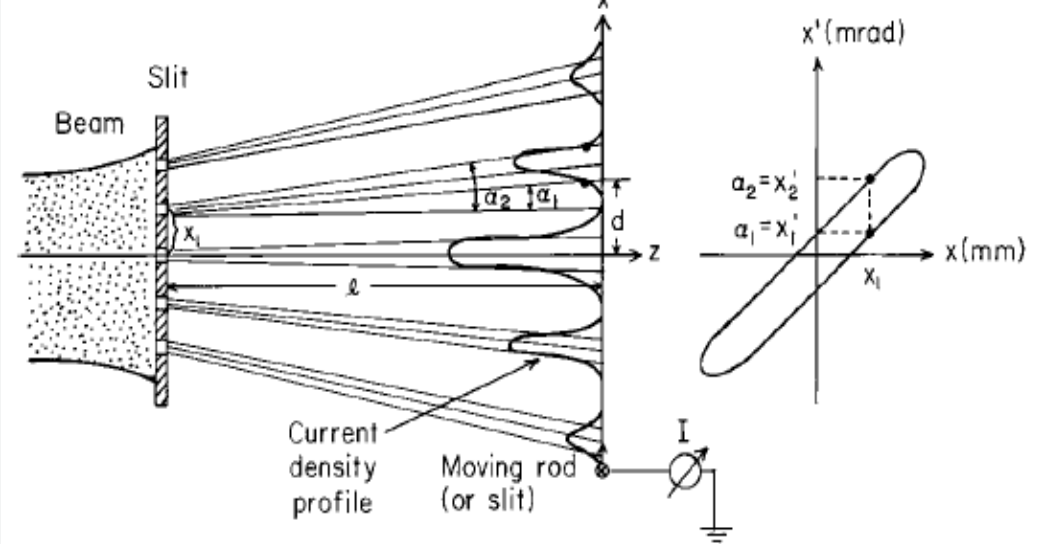
$$\gamma_x = \frac{\langle x'^2 \rangle}{\epsilon_x}$$

$$\alpha_x = -\frac{\langle xx' \rangle}{\epsilon_x}$$

$$\beta_x \gamma_x - \alpha_x^2 = 1$$



Martin Reiser – Theory and Design of Charged Particle Beams

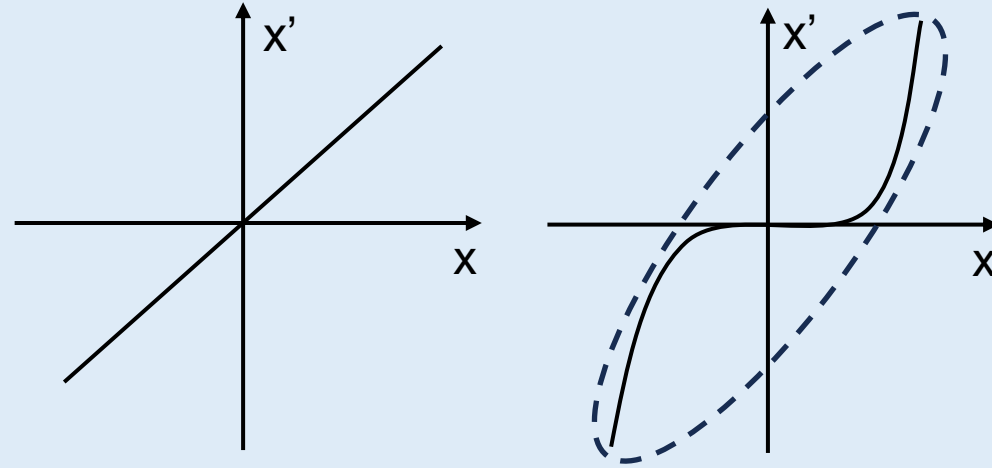
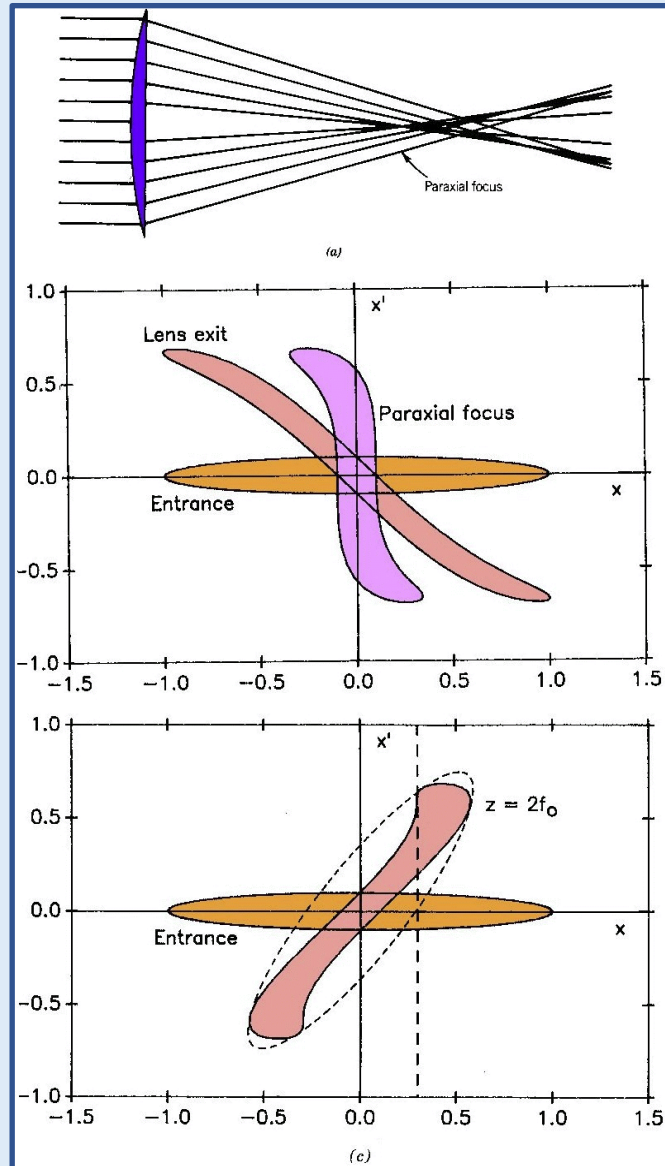


➤ Beam brightness definition:
Current density x unit solid angle

$$B = \frac{J}{d\Omega} = \frac{dI}{dS d\Omega} \quad \langle B_n \rangle = \frac{I}{\iint dS d\Omega} = \frac{2I}{\pi^2 \epsilon_{n,x} \epsilon_{n,y}} \left[\frac{A}{m^2} \right]$$

Brightness or Emittance Degradation

A focusing channel



Considering that for any position x the divergence of the particle is $x' = Cx^n$

$$\varepsilon_x^2 = \overline{x^2 x'^2} - \overline{xx'}^2 = C^2 (\overline{x^2 x^{2n}} - \overline{xx^{n+1}}^2)$$

with $n=1$ the straight line gives rms emittance equal 0. For $n \neq 1$ the emittance is not 0, also if the two distribution area are 0

❖ The Brightness

Brightness

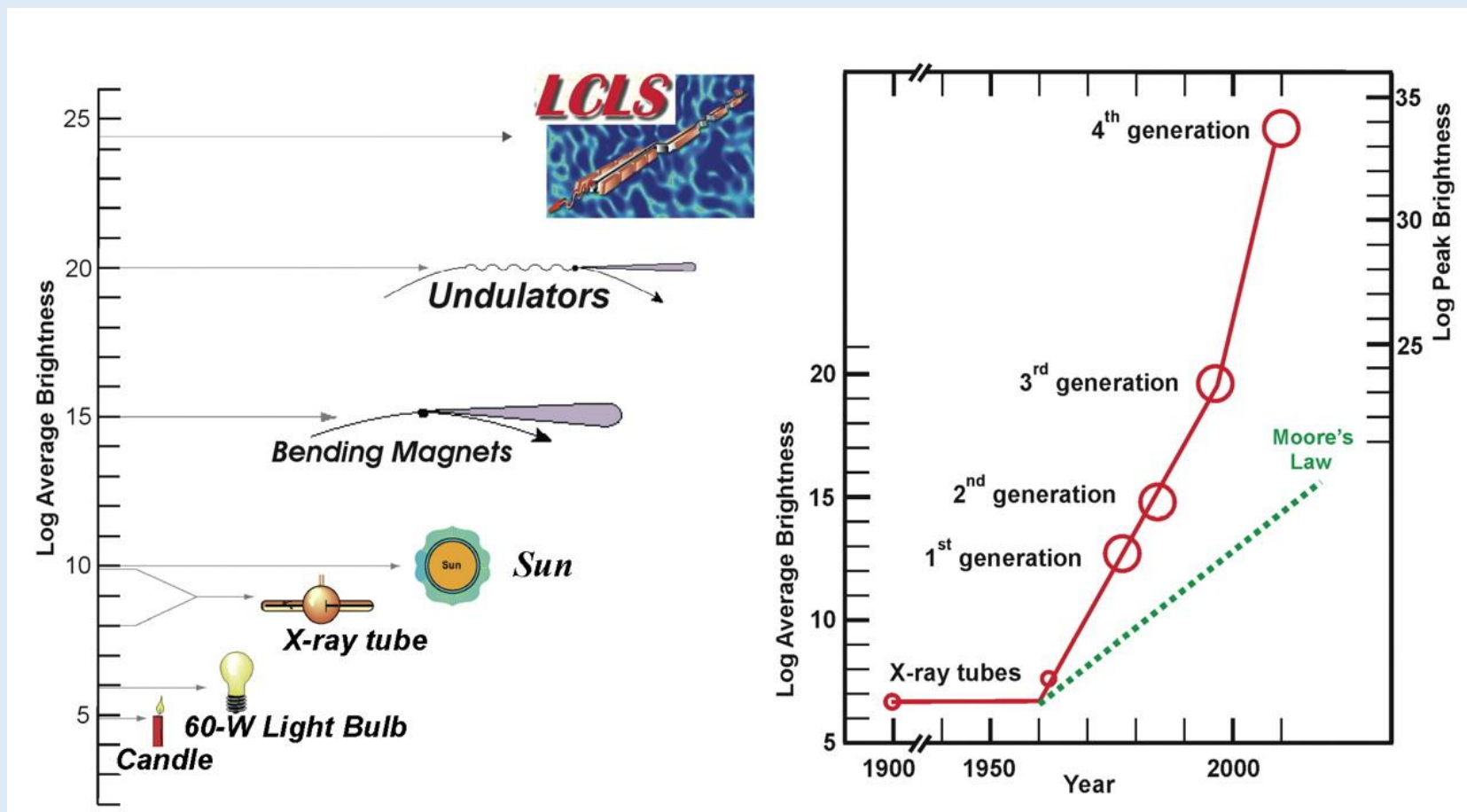
$$B = \frac{2I}{\pi^2 \varepsilon_{x,n} \varepsilon_{y,n}}$$

Current

$$I[A] = \frac{Q[C]c[\frac{m}{s}]}{\sigma_z[m]\sqrt{12}}$$

Emittance

$$\varepsilon_{n,x} = \beta\gamma\sqrt{\langle x^2 \rangle \langle x'^2 \rangle - \langle x \cdot x' \rangle^2}$$



Accelerators cavities basic concepts:

Resonant Modes, Phase velocity,
Standing & Traveling, low β cavities

RF CAVITIES

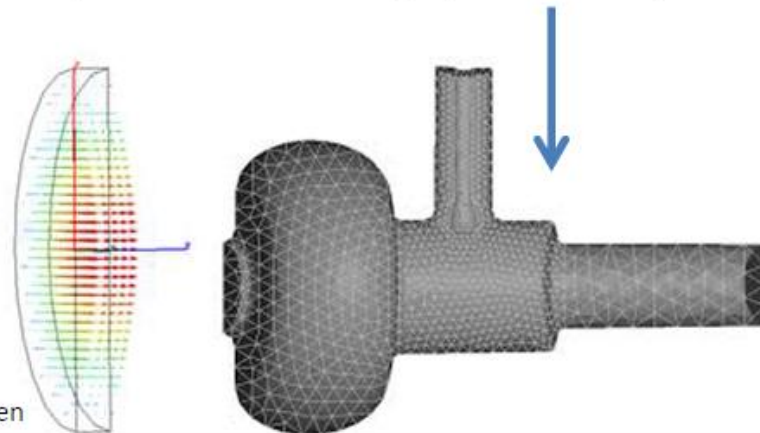
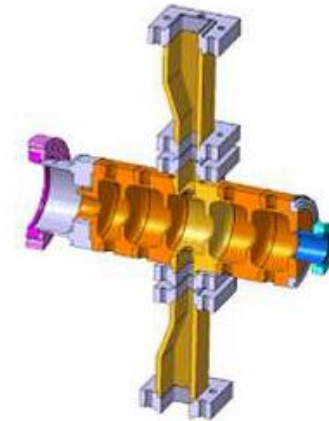
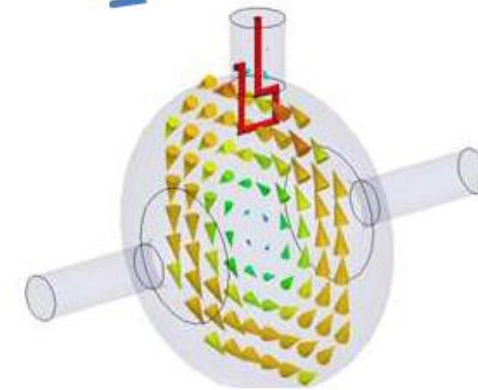
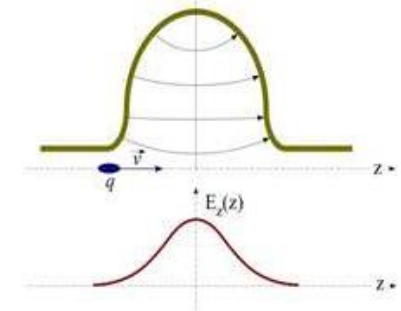
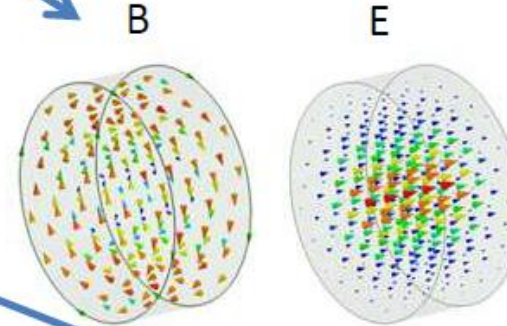
⇒ High frequency RF accelerating fields are confined in **cavities**.

⇒ The cavities are metallic closed volumes where the e.m. fields has a particular spatial configuration (**resonant modes**) whose components, including the accelerating field E_z , oscillate at some specific frequencies f_{RF} (resonant frequency) characteristic of the mode.

⇒ The modes are excited by **RF generators** that are coupled to the cavities through waveguides, coaxial cables, etc...

⇒ The resonant modes are called **Standing Wave (SW) modes**.

⇒ The spatial and temporal field profiles in a cavity have to be computed (analytically or numerically) **by solving the Maxwell equations** with the proper boundary conditions.



Wave Type	TM ₀₁	TM ₀₂	TM ₁₁	TE ₀₁	TE ₁₁
Field distributions in cross-sectional plane, at plane of maximum transverse fields					
Field distributions along guide					
Field components present	E_z, E_r, H_ϕ	E_z, E_r, H_ϕ	$E_z, E_r, E_\phi, H_r, H_\phi$	H_z, H_r, E_ϕ	$H_z, H_r, H_\phi, E_r, E_\phi$

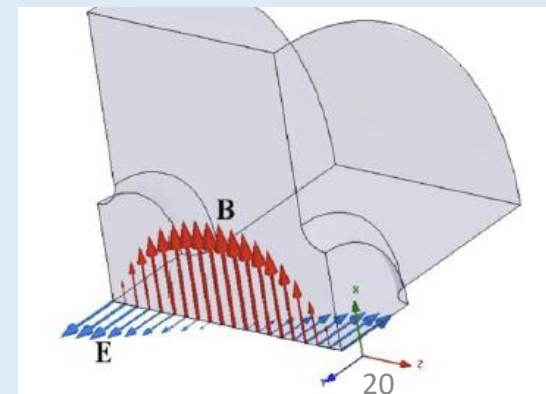
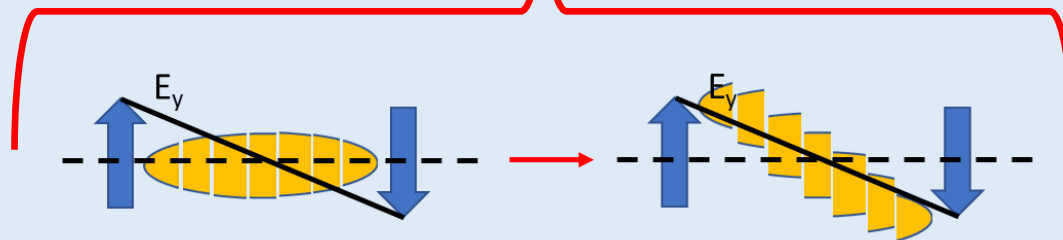
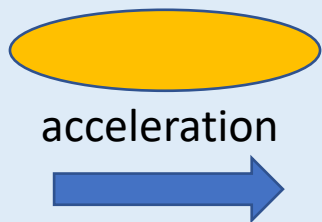


Fig. 13. Field profiles of the deflecting TM₁₁₀-like mode.

Dispersion curve and Phase Velocity v_p to accelerate $v_p = c\beta$

$$E_z = E_0 J_0(k_c r) \exp[j(\omega t - \beta z)]$$

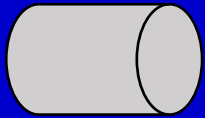
$$E_r = j\eta E_0 [1 - (\omega_c/\omega)^2]^{\frac{1}{2}} J_1(k_c r) \exp[j(\omega t - \beta z)]$$

$$H_\theta = jE_0 J_1(k_c r) \exp[j(\omega t - \beta z)]$$

TM₀₁

where J_0 and J_1 are Bessel functions, ω_c is the cut-off frequency, $k_c = \omega_c/c$, $\beta = \omega/v_p$, $\beta^2 = k^2 - k_c^2$, $k = \omega/c$ and η is the intrinsic impedance of the medium. For these solutions to exist, the propagation constant β must be real; this condition is fulfilled only if $v_p > c$.

Into a cylindrical guide:



$$\omega^2 = (k_c \cdot c)^2 + (\beta \cdot c)^2$$

$$v_p = \frac{\omega}{\beta} = \frac{c}{\sqrt{1 - (k_c \cdot c/\omega)^2}} > c$$

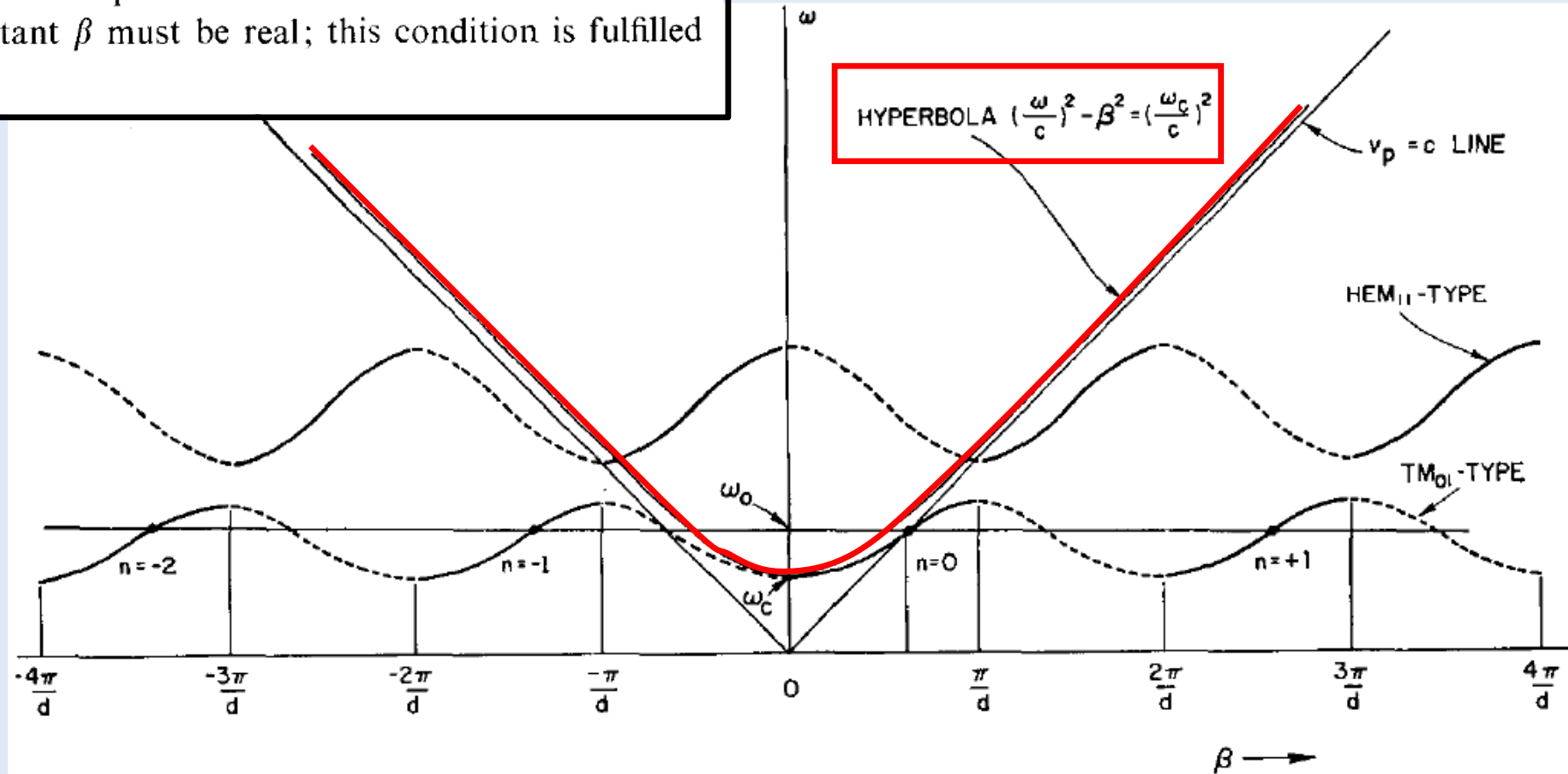


Fig. 2. Brillouin (or ω - β) diagram showing propagation characteristics for uniform and periodically loaded structures.

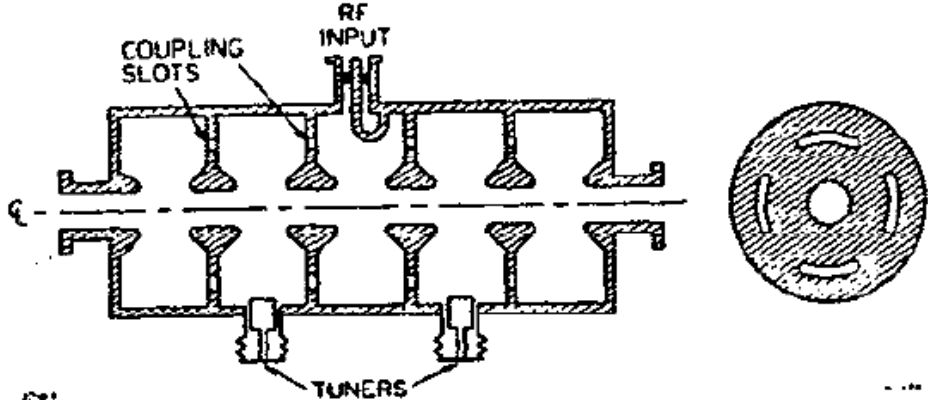


Fig. 3.3. Diagram showing the important features of a five-cell v-mode structure with magnetic field coupling.

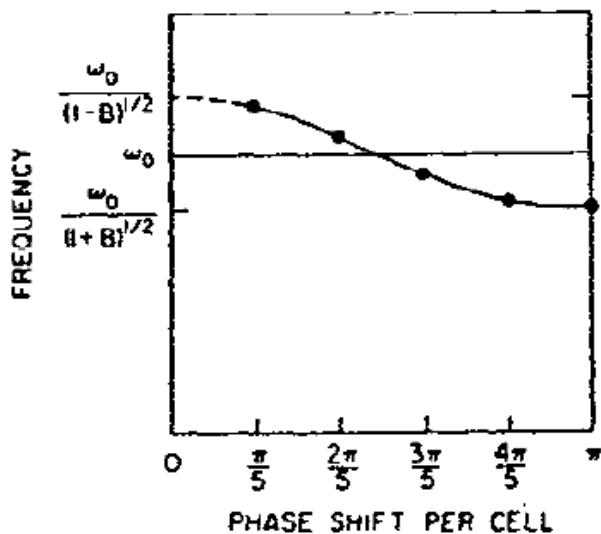
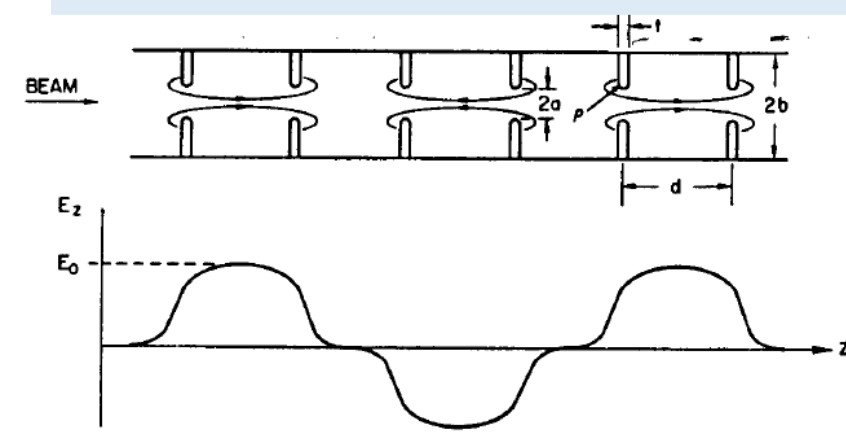
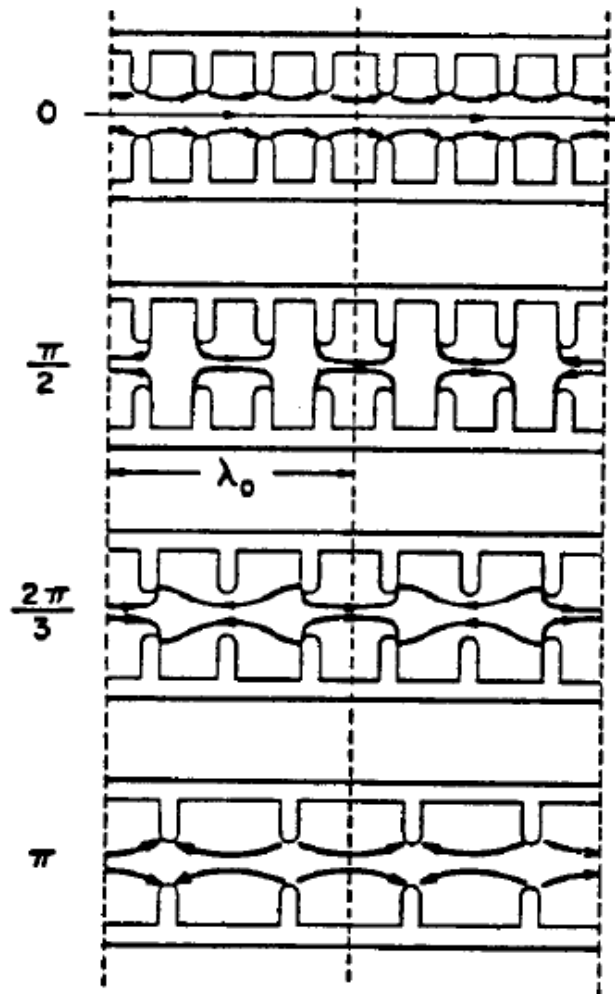


Fig. 3.4. Dispersion diagram for a five-cell structure with "flat" v-mode.



Low Beta cavities $\beta < 1$

DTL – Drift Tube Linear Accelerator

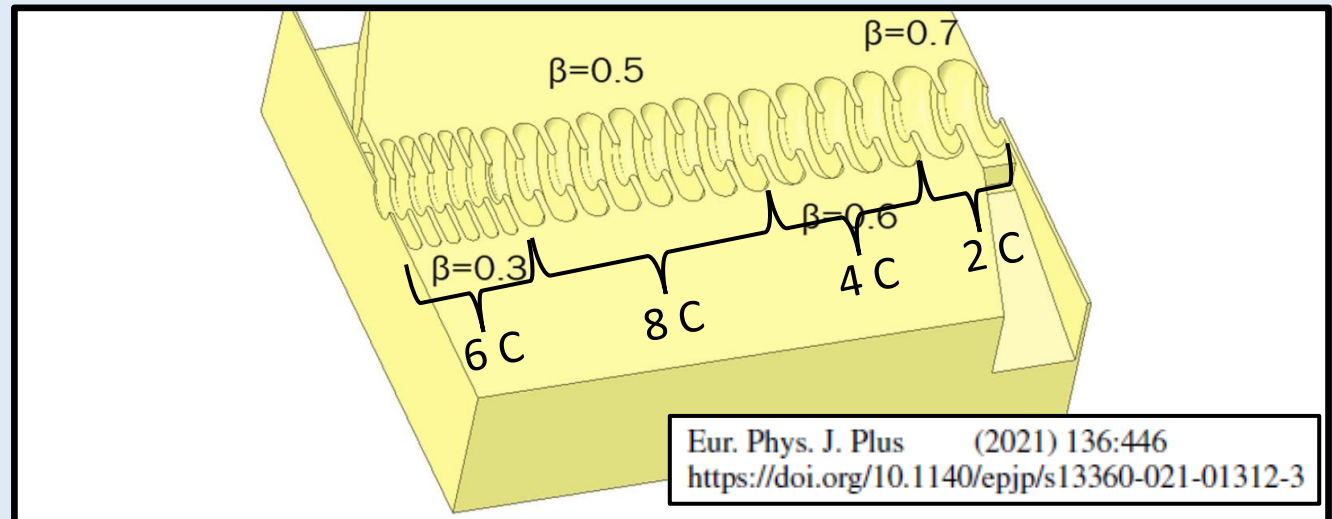
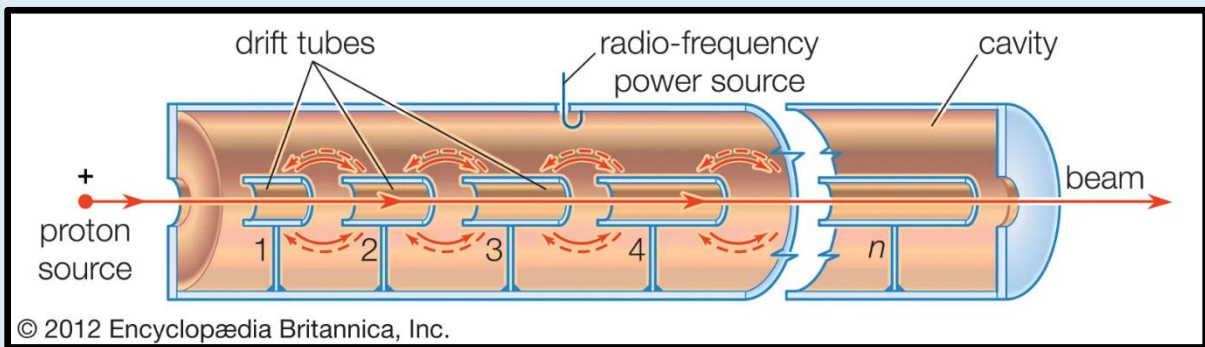


Fig. 31 Example of a buncher for a 40-kW 16.4 GHz linac, consisting of four $\beta_{ph} = \text{const}$ sections: six cells with $\beta_{ph} = 0.3$, eight cells with $\beta_{ph} = 0.5$, four cells with $\beta_{ph} = 0.6$, and two cells with $\beta_{ph} = 0.7$. The available power restricts the length due to low field amplitude A

Sergey V. Kutsaev (RadiaBeam Tech) – Electron bunchers for industrial RF linear accelerators: Theory and design guide²³

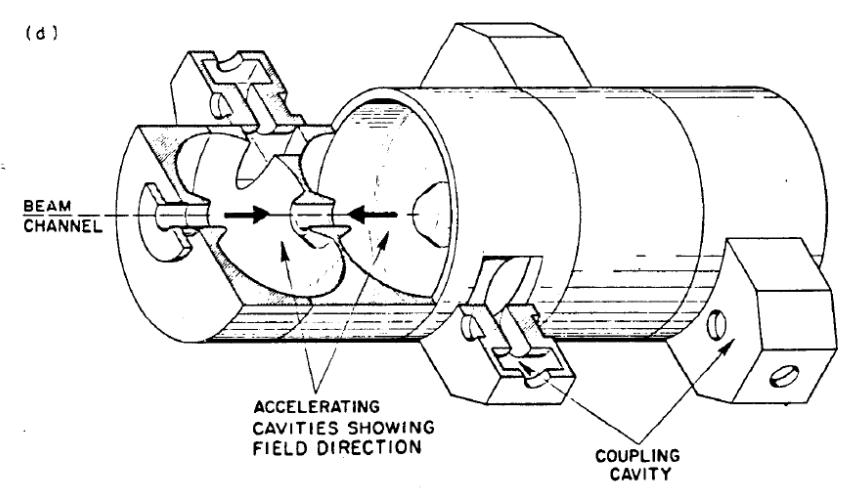
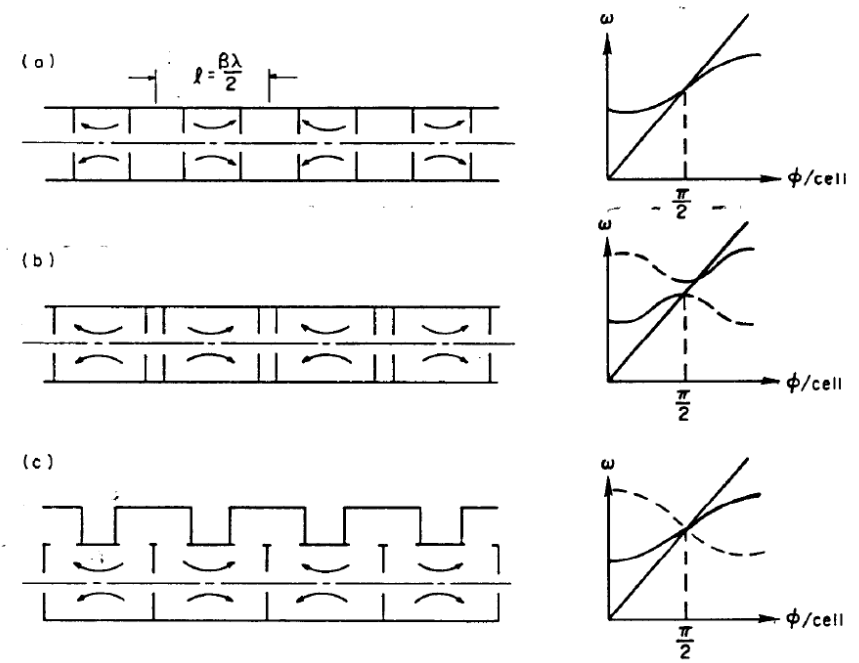
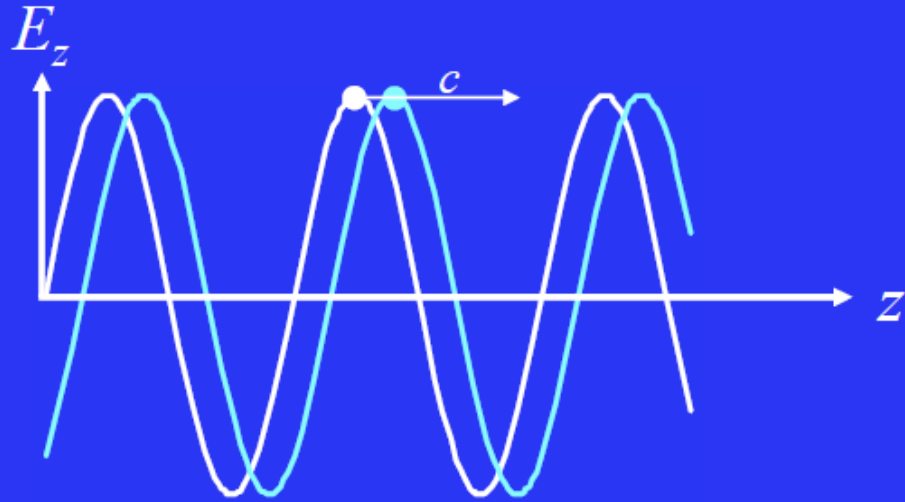


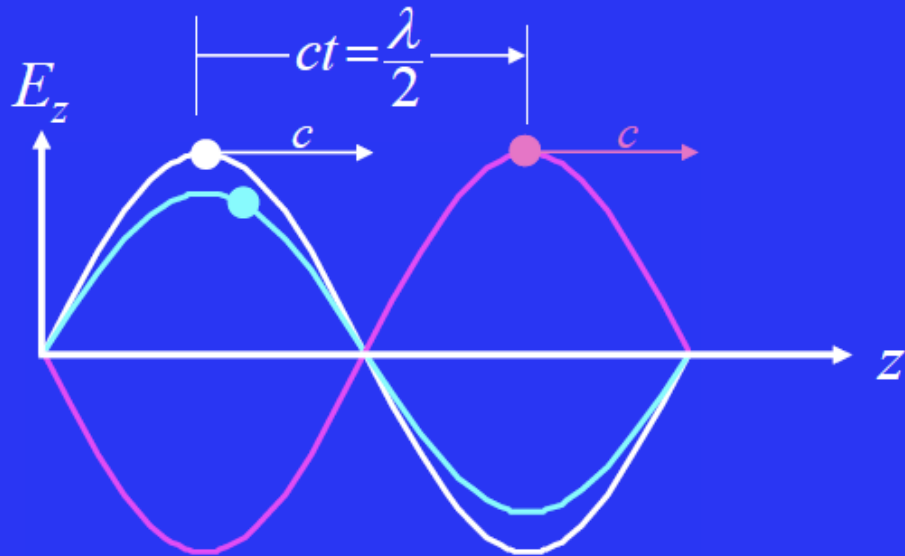
Fig. 31. Evolution of $\pi/2$ -mode iris-loaded structure to π -mode side-coupled structure.

TW & SW cavities – 1/4



travelling wave structure:
need *phase velocity* = c
(*disk-loaded structure*)

bunch sees constant field:
 $E_z = E_0 \cos(\phi)$

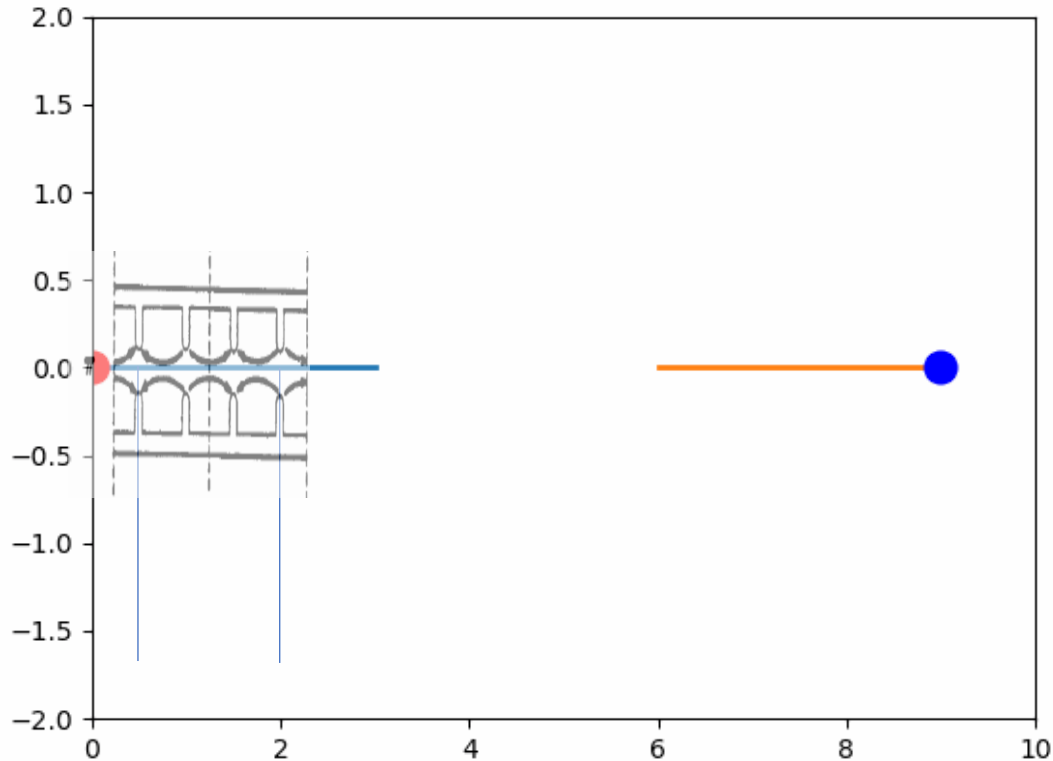


standing wave cavity:

bunch sees field:
 $E_z = E_0 \sin(\omega t + \phi) \sin(kz)$
 $= E_0 \sin(kz + \phi) \sin(kz)$

TW & SW cavities – 2/4

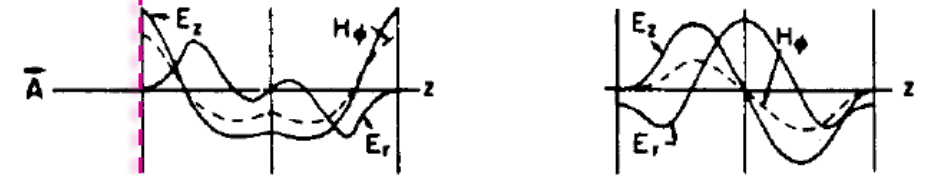
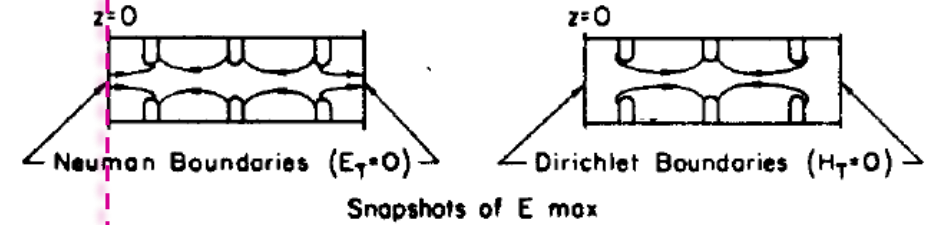
Standing Wave case; e.g. of an Energy Recovery Linac (ERL) case



$$TW = SW1 + SW2$$

SLAC-PUB-2295
March 1979

(c) Standing waves, $\beta_0 d = 2\pi/3$, for boundaries at $z=0$ and $z=3d$



Maximum amplitudes of E_z and E_r (in phase) and H_ϕ (leading in time quadrature)

(d) Standing waves, $\beta_0 d = 2\pi/3$, for boundaries shifted by $z=-d$

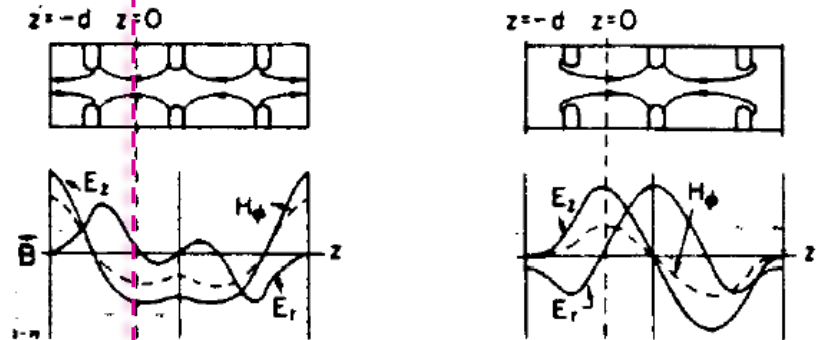


Figure 1

TW & SW cavities – 3/4

SLAC-PUB-2295
March 1979

$$\begin{aligned}\vec{B} &= e^{j\omega t + \pi} \sum_{-\infty}^{+\infty} 2a_n \cos \beta_n (z+d) \\ \vec{A} &= e^{j\omega t + \pi} \sum_{-\infty}^{+\infty} 2a_n \cos \beta_n z\end{aligned}\quad (6)$$

both of which are made up of one TW going left and one going right. The "trick" is to add them with the proper phases to have the TW's going left cancel and those going right add. This can be achieved by multiplying \vec{A} by $e^{j(\beta_0 d - \pi/2)}$ and \vec{B} by $e^{j\pi/2}$. Then:

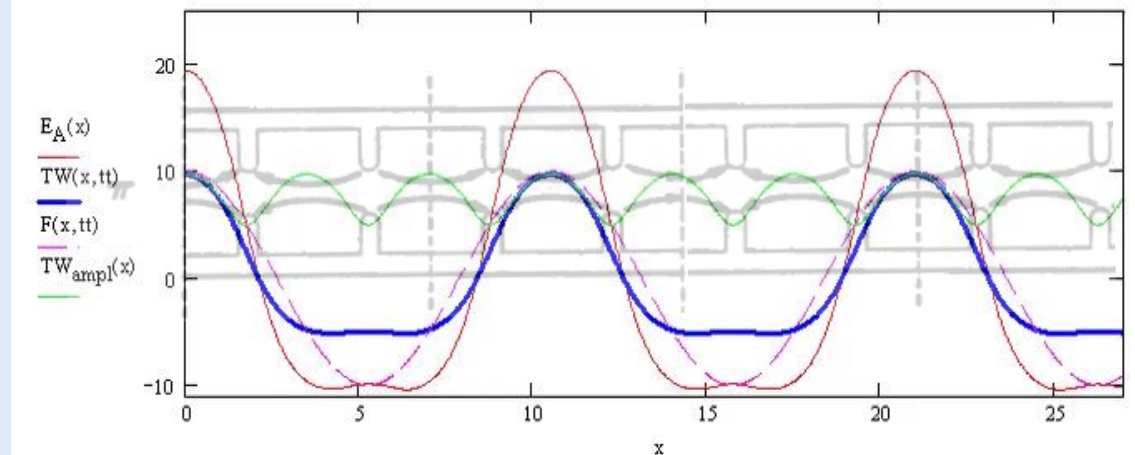
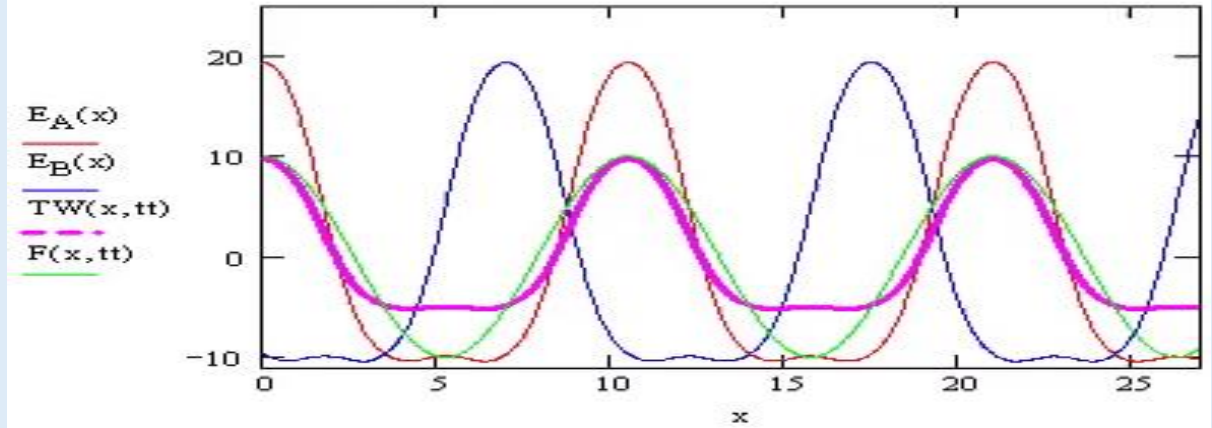
$$A e^{j(\beta_0 d - \frac{\pi}{2})} + B e^{j\frac{\pi}{2}} = 2 \sin \beta_0 d \underbrace{\sum_{-\infty}^{+\infty} a_n e^{j(\omega t - \beta_n z)}}_{\text{TW}}$$

and it follows that the amplitude and phase of the TW are:

$$|\text{TW}|^2 = \frac{A^2 + B^2 - 2AB \cos \beta_0 d}{4 \sin^2 \beta_0 d} \quad (7)$$

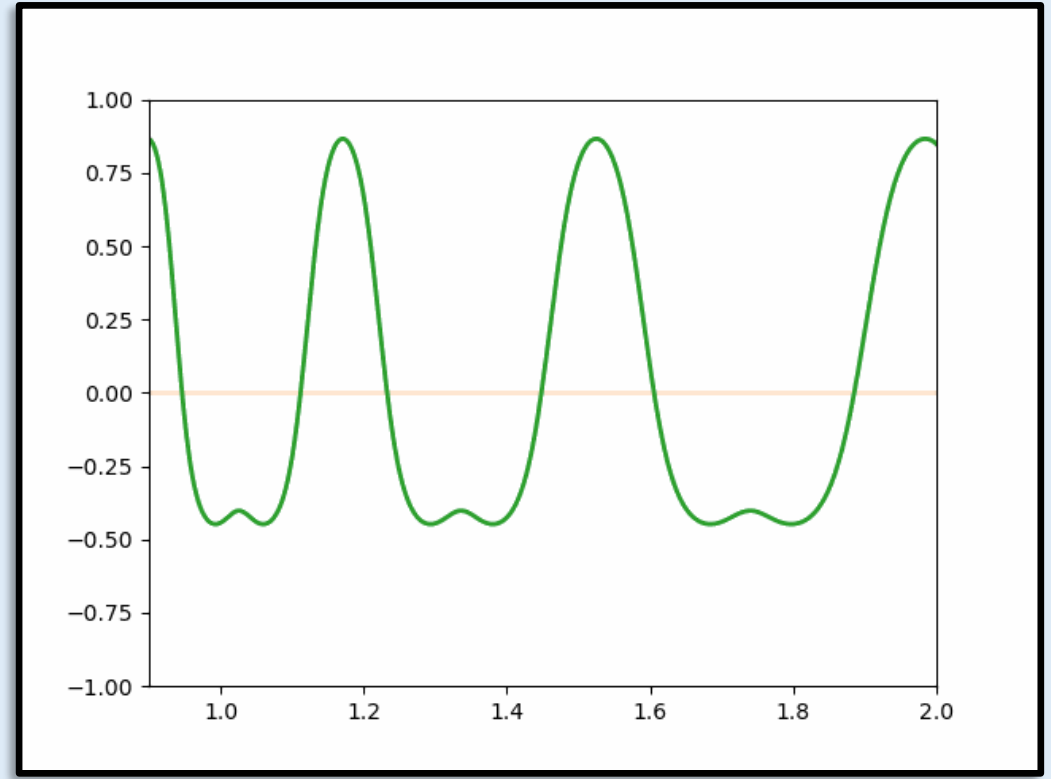
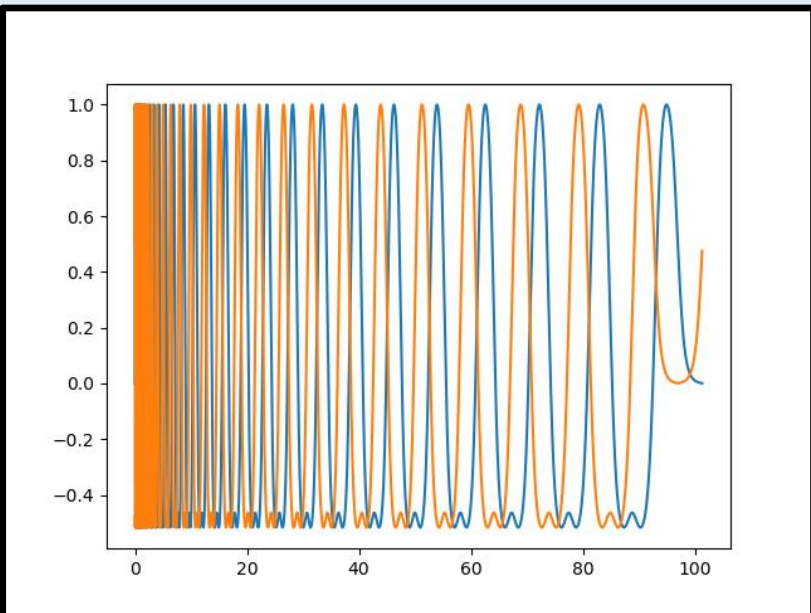
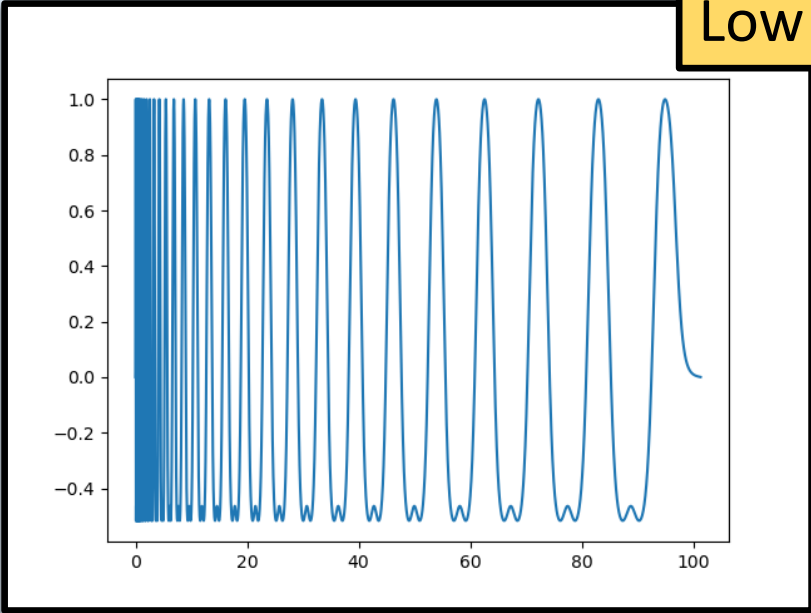
$$\tan \theta(z) = \frac{B - A \cos \beta_0 d}{A \sin \beta_0 d} \quad (8)$$

Traveling Wave case =
two SW's dephased



TW & SW cavities – 4/4

Low Beta cavities $\beta < 1$

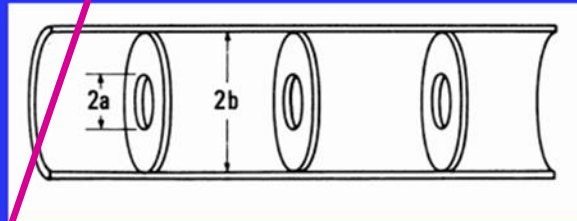
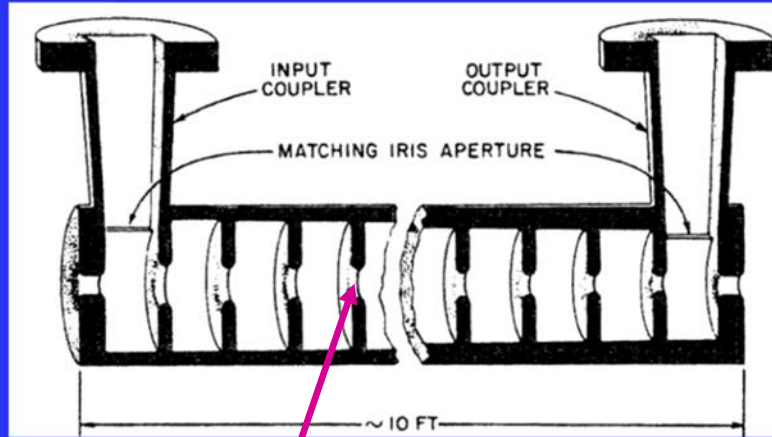


e.g.: 3 meters long SLAC type S-band module

Travelling wave structure

Circular waveguide mode TM_{01} has $v_p > c$

It needs to slow down the wave using irises



Parameters of the SLAC **constant-gradient** Travelling-Wave (TW) structure.

From G. A. Loew, R. B. Neal, "Accelerating Structures in Linear Accelerators" (1970)

$$G \left[\frac{MV}{m} \right] = \sqrt{r_l \left[\frac{M\Omega}{m} \right] \cdot \frac{P [MW]}{l_{cav}}}$$

$$r_l \propto \sqrt{f}$$

$$Grad(3 \text{ GHz}) \sim 35 \text{ MeV/m}$$

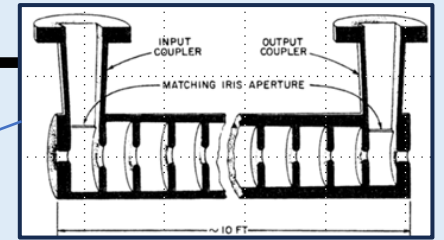
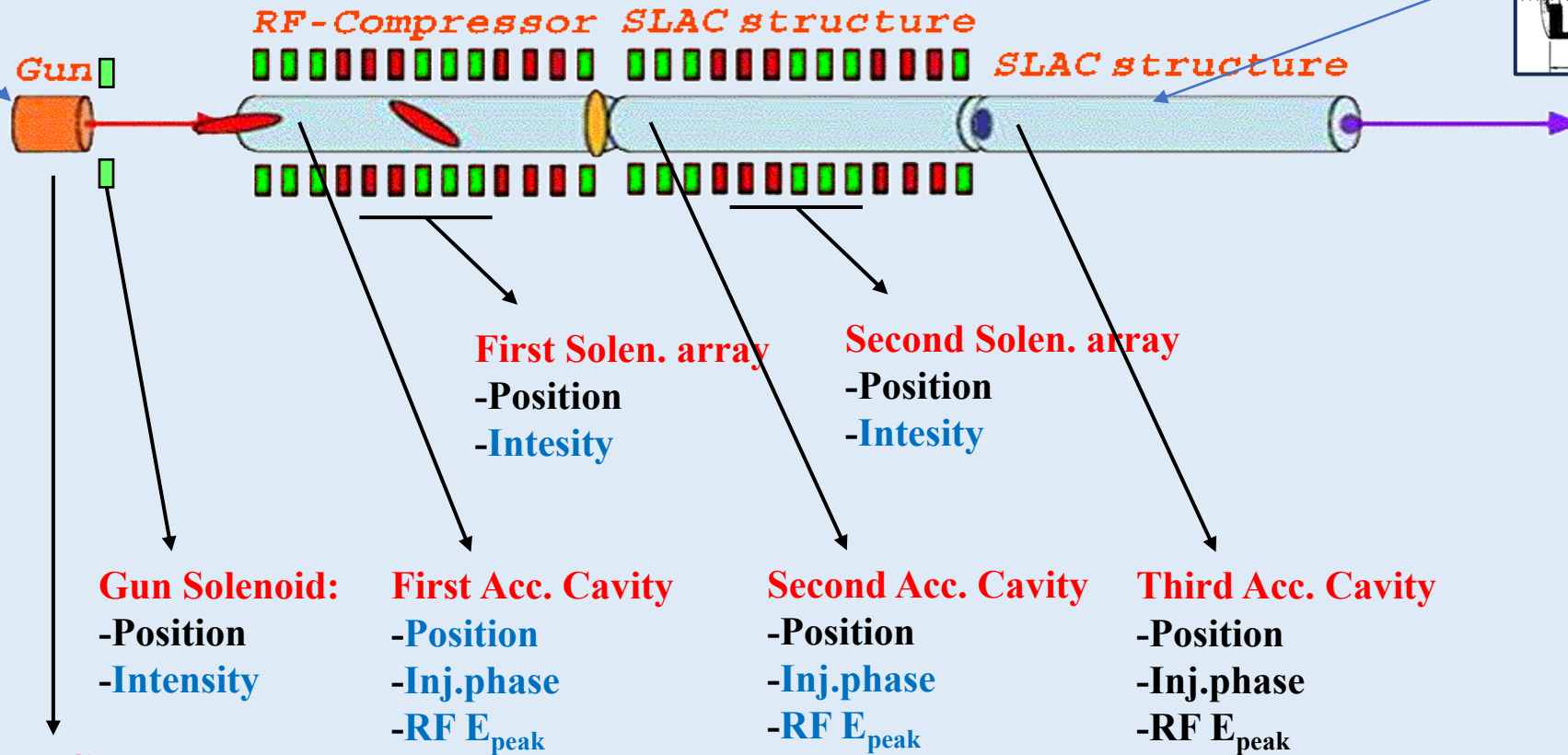
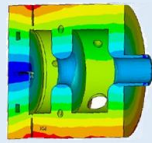
$$Grad(100 \text{ GHz}) \sim 200 \text{ MeV/m}$$

$$Grad(30 \text{ THz}) \sim 3.5 \text{ GeV/m}$$

Parameter	symbol	Units	Value
Frequency	$\omega/2\pi$	Hz	2856 MHz
Length	L	m	3.048
Cell radius	b	cm	4.17–4.09
Iris radius	a	cm	1.31–0.96
Cell length	d	cm	3.50
Phase shift per cell	ψ	-	$2\pi/3$
Disc thickness	h	cm	0.584
Quality Factor	Q	-	13,000
Shunt impedance per meter	r_l	$M\Omega/m$	52–60
Filling Time	t_f	nsec	830
Group Velocity	v_{gr}	%c	2.0–0.65
Attenuation	τ	"nepers"	0.57

Sketch of a typical High Brightness
electron LINAC

A typical ideal High Brightness electron BeamLine



The Gun:

- Laser Longitudinal profile
- Laser transverse dimension (relative uniformity)
- Inj. phase in RF Acc.Field
- RF E_{peak}

19 parameters, some strongly coupled (non linearly) to each other
In-blue harder parameters to be set (12/19)
In-black easier parameters to be set (7/19)

GIOTTO — Genetic Interface for Optimising Tracking with Optics

From 2007 up to day, the code is grew in power and versatility

What makes the difference:

Nowadays “quasi-classic” optimization techniques >> elitism; advanced mutation

operators; hill climbing; regeneration from best solutions; parallelization (Open-MPI, MS-mpi)
fitness function freely defined by the user, by using all the tracking code’s outputs (Astra)
or by a dedicated post processor for the Lcomb configuration:
En, Den, SigZ, Xemit, sigX, Yemit, SigY, emitY

Constraints freely defined by the user

NameList (nml) can be imported into a DB and each nml variables can be used as a Giotto variable to be optimized (genes) (ex. Phi(1)...Phi(50),maxe(1),maxb(1), sig_x,sig_clock ---
No limit in the number)

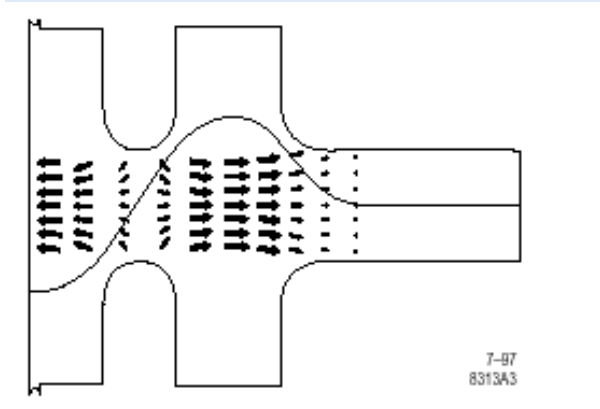
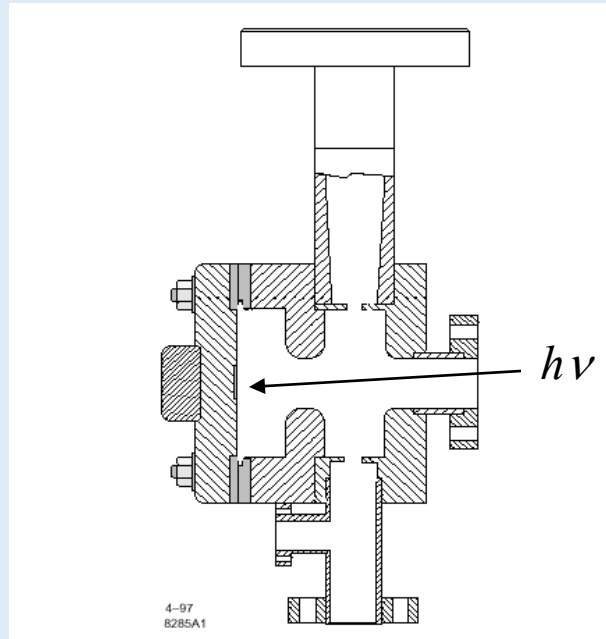
switches from Genetic Optimizations to Statistical Analysis. Each variable can be analyzed.
The **sampling interval** can be sampled in **uniform** or **Gaussian** way – very fast stat. analysis.

GIOTTO soon on the ASTRA repository -- <https://www.desy.de/~mpyflo/>
or write to alberto.bacci@mi.infn.it or marcello.rosseti@mi.infn.it

Radio-Frequency Photo-Injectors

UCLA/SLAC/BNL

S-band next gen. RF Gun



$$Q_{eff} = N_{electrons} / N_{laser-photon}$$

$$Q_{eff}(\text{Cu photo-cathode}) \cong 5 \cdot 10^{-5}$$

$$W_{Cu} = 4.2eV, h\nu = 4.6eV$$

$$Q = 1 \text{ nC needs } U_{las} = \frac{h\nu \cdot Q_{bunch}}{Q_{eff}} = 92 \mu J$$

Photo-Cathode Emissivity $J < 10 \text{ kA/cm}^2$

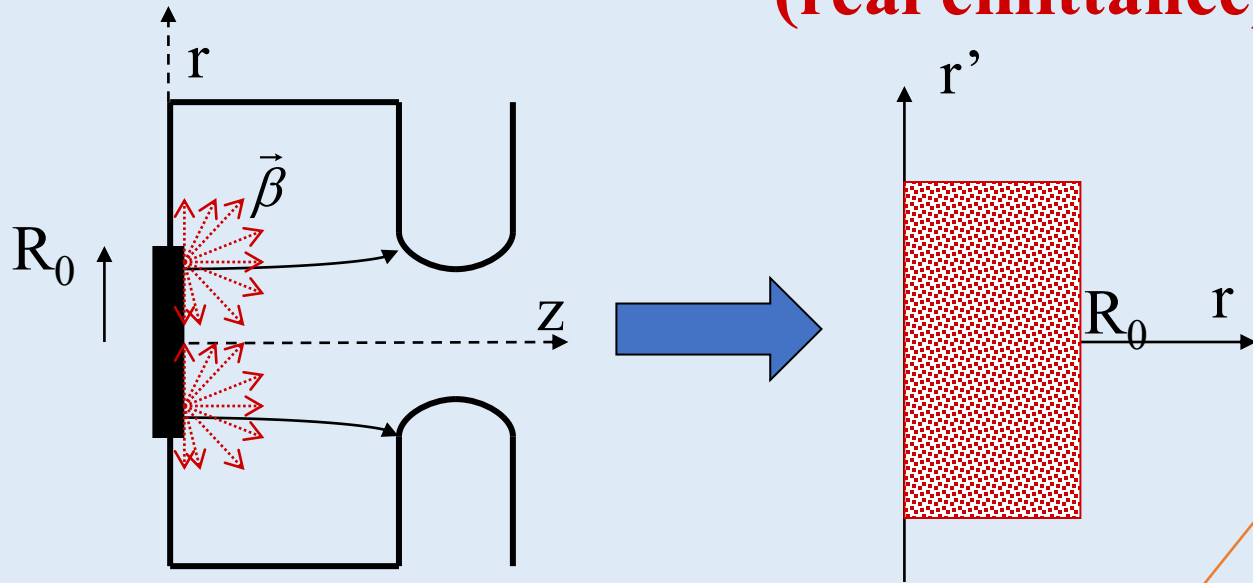
(t) Prompt emission on a ps time scale

Thermoionic Injectors

Cathode Emissivity $J < 20 \text{ A/cm}^2$

Beam Dynamics in Photo-Injectors

temperature emittance @ photo-cathode
(real emittance)



$$\varepsilon_n^{th} = \frac{\langle \beta \gamma \rangle}{2} \sqrt{\langle r^2 \rangle \langle r'^2 \rangle - \langle r r' \rangle^2}$$

$$\sigma_x = \frac{R_0}{2}$$

T_e (or r' is relative to the difference between
[Photo Energy - Photo-Electric working function])

$$\langle \beta \gamma \rangle \cong \langle \beta \rangle \cong \sqrt{2T_e / m_e c^2} = 2 \cdot 10^{-3} \sqrt{T_e [eV]}$$

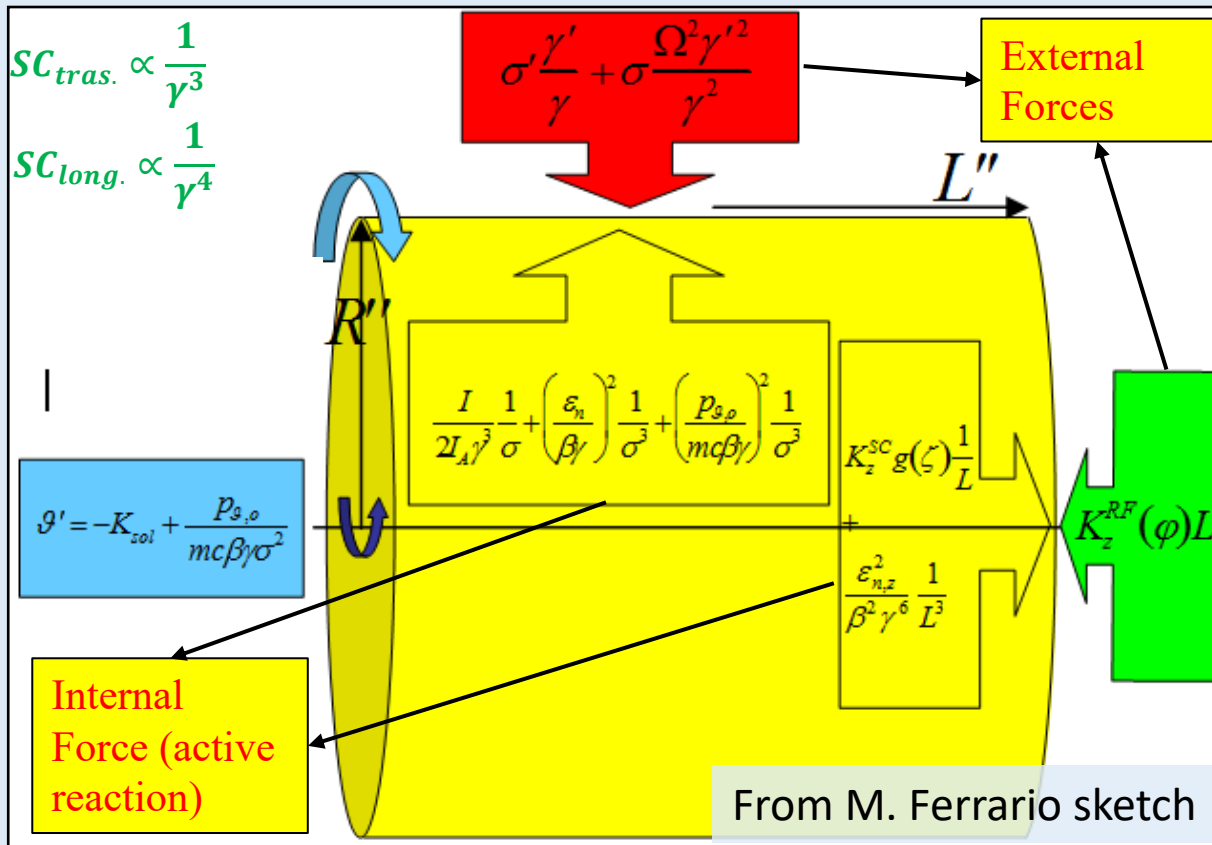


$$\varepsilon_n^{th} = \frac{\pi \langle \beta \rangle R_0}{4\sqrt{6}}$$

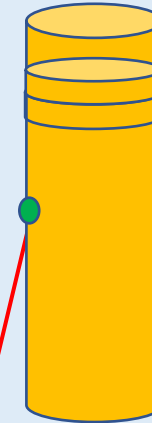
$$\varepsilon_n^{th} [mm \cdot mrad] = 0.64 R_0 [mm] \sqrt{T_e [eV]}$$

Optimize High Brightness BeamLines is challenging – 1

The electron beam is an 'reactive' distribution:



A Traveling Electron beam



e-beam velocity ↑

charged particle by Gauss-Amper F_s

$$\frac{\text{Coulomb rep. } F}{\text{Mag. attrac. } F} = \left(\frac{c}{v}\right)^2 = \beta^2$$

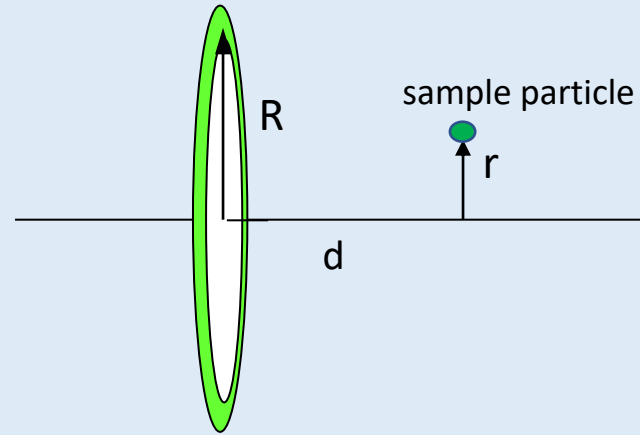
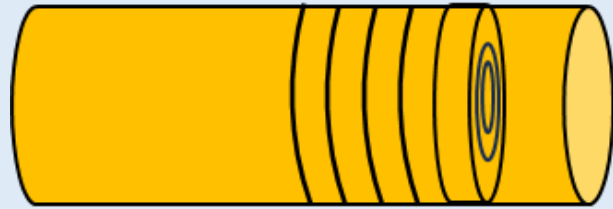
Space-Charge dominated bunches are hard to be tamed

CODES Model (2D cylindrical symmetry)

Space charge computation recipe:

Lorentz-transforming the particles position and field maps into the **average rest frame of the beam**.

It then applies static forces to the various rings of the cylindrical map assuming a constant charge density inside a ring. This algorithm requires to have some particles in each of the cell of the cylindrical grid.



In the rest frame of the circle, there is only an electrostatic field^[2]. This from the electrostatic potential (in polar coordinates):

$$V'_j(r, \theta) = \frac{\lambda_j R}{4\pi\epsilon_0} \int_0^{2\pi} \frac{1}{\sqrt{R^2 + r^2 - 2Rr \sin(\theta) \cos(\phi)}} d\phi$$

$$= \frac{Q_j}{2\pi^2 \epsilon_0} \frac{K\left(\frac{4Rr \sin(\theta)}{a^2 + r^2 + 2Rr \sin(\theta)}\right)}{\sqrt{R^2 + r^2 + 2Rr \sin(\theta)}}$$

in which $K(k)$ is the elliptic integral^[3]:

$$K(k) = \int_0^{\pi/2} \frac{1}{\sqrt{1 - k \sin(\theta)}} d\theta$$

$$E'_r = \frac{Q}{4\pi^2 \epsilon_0 r \sqrt{d^2 + 4Rr}} \left(K(\alpha) - \frac{R^2 - r^2 + z^2}{d^2} E(\alpha) \right)$$

$$E'_z = \frac{Qz E(\alpha)}{2\pi^2 \epsilon_0 d^2 \sqrt{d^2 + 4Rr}}$$

where

$$E(k) = \int_0^{\pi/2} \sqrt{1 - k \sin(\theta)} d\theta, \quad d^2 = (R - r)^2 + z^2 \quad \text{and} \quad \alpha = 4Rr / (d^2 + 4Rr)$$

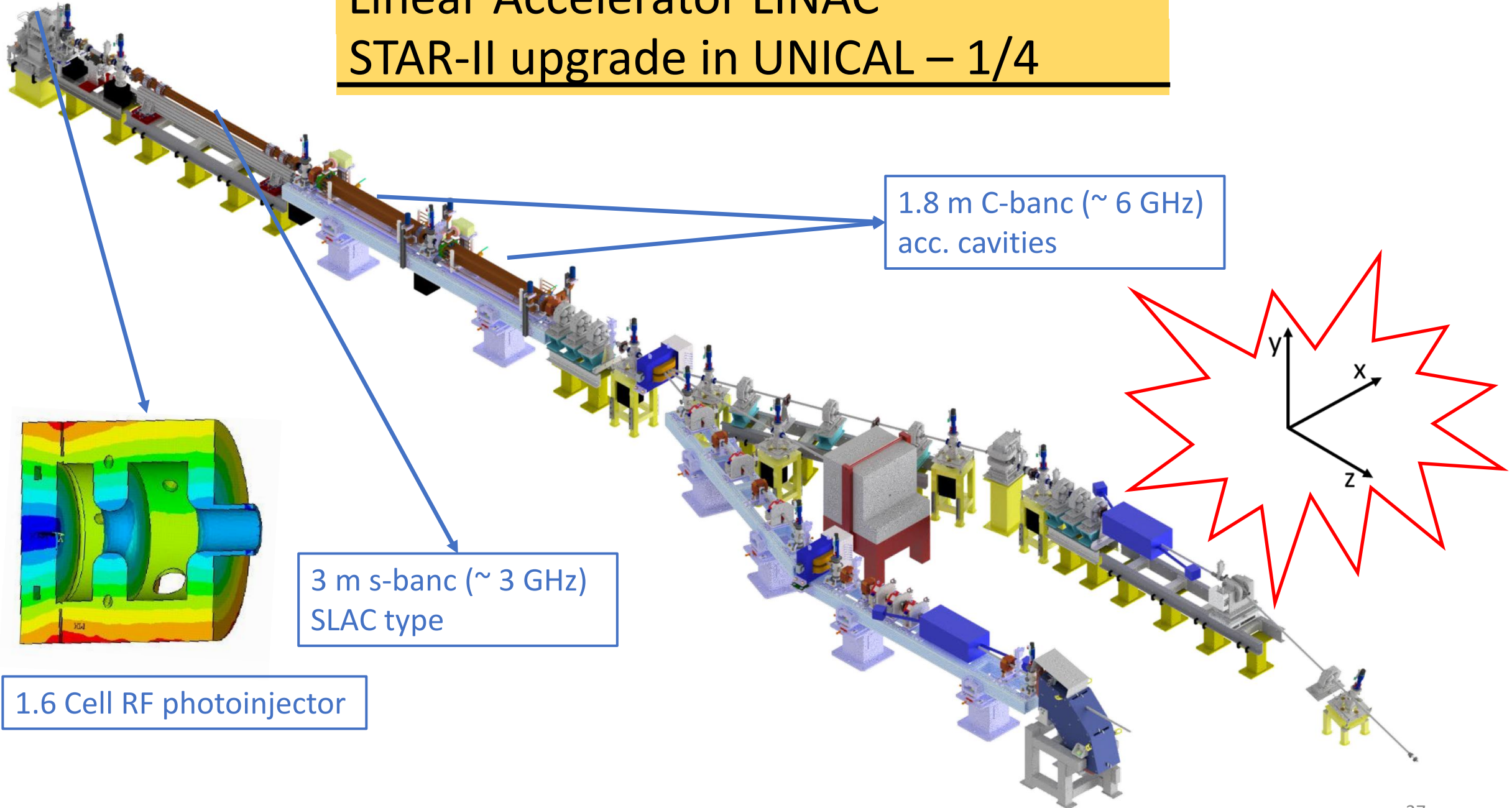
After seeing:

- **relativity and acceleration**
- **optics in accelerators**
- **Cavities basic concepts**

Let's see a real accelerator

- **Then come back to beam parameters**

Linear Accelerator LINAC STAR-II upgrade in UNICAL – 1/4

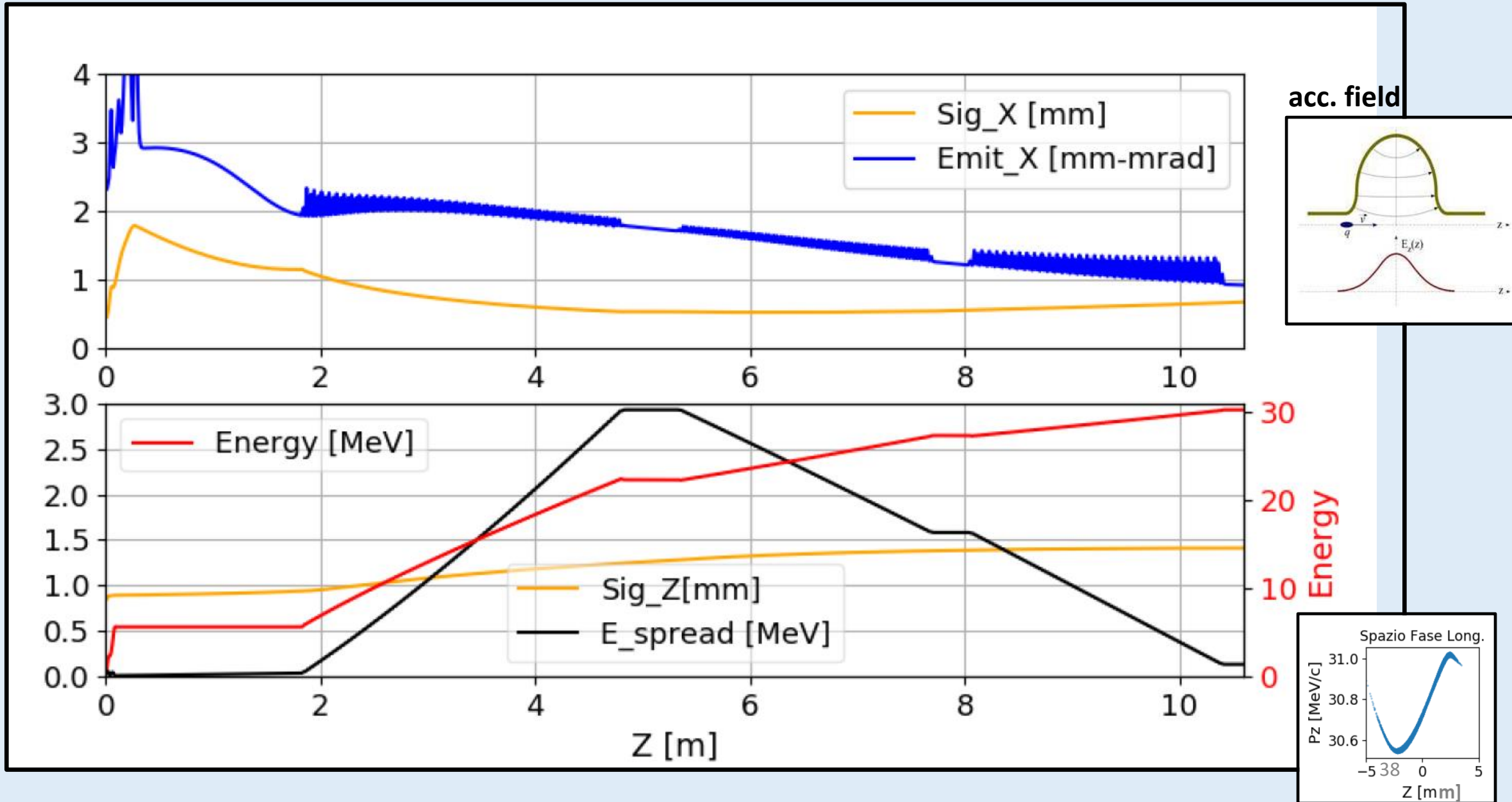


1.6 Cell RF photoinjector

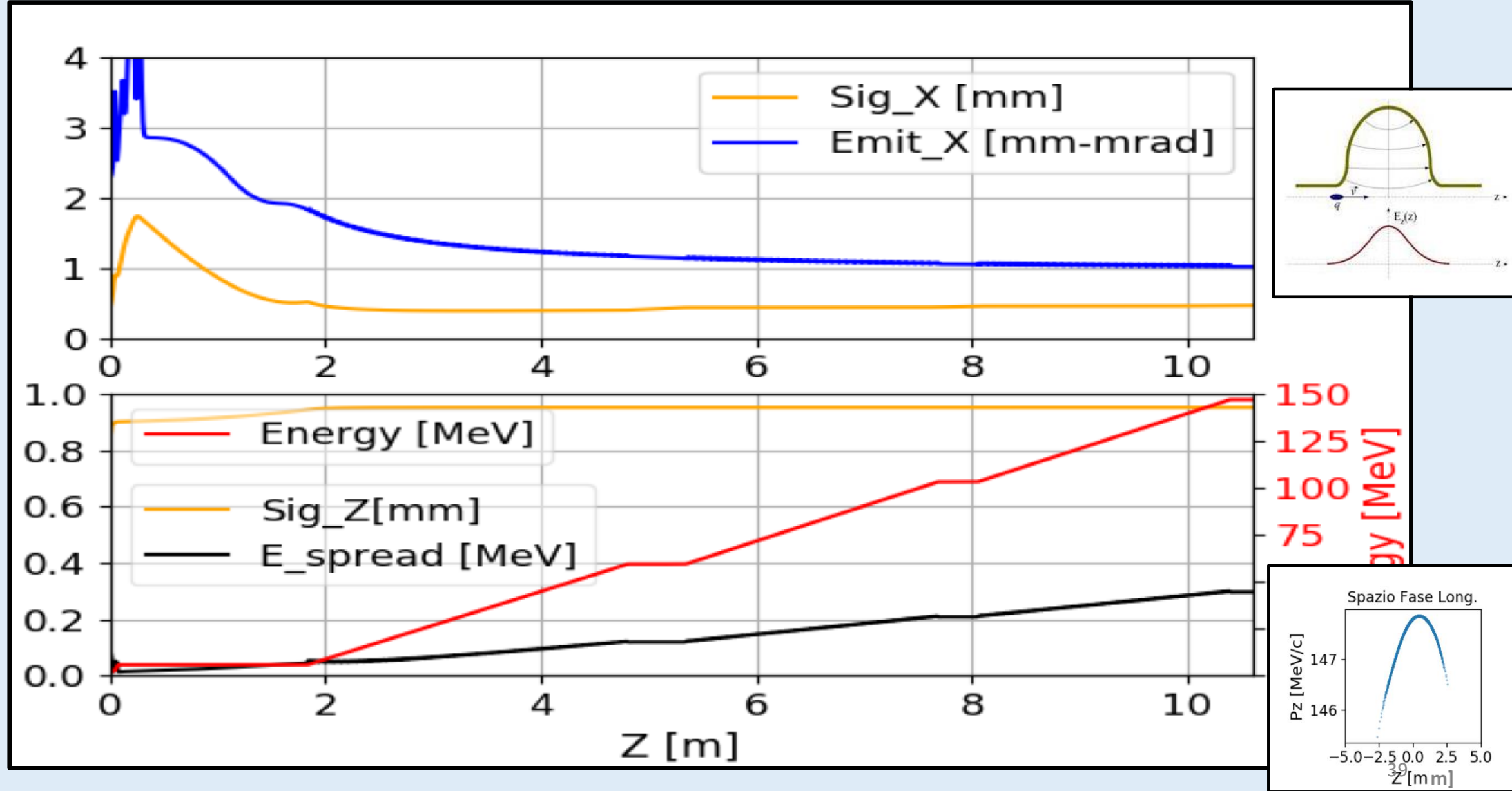
3 m s-banc (~ 3 GHz)
SLAC type

1.8 m C-banc (~ 6 GHz)
acc. cavities

Typical beam parameters for a LINAC (30 MeV – STAR case) – 2/4

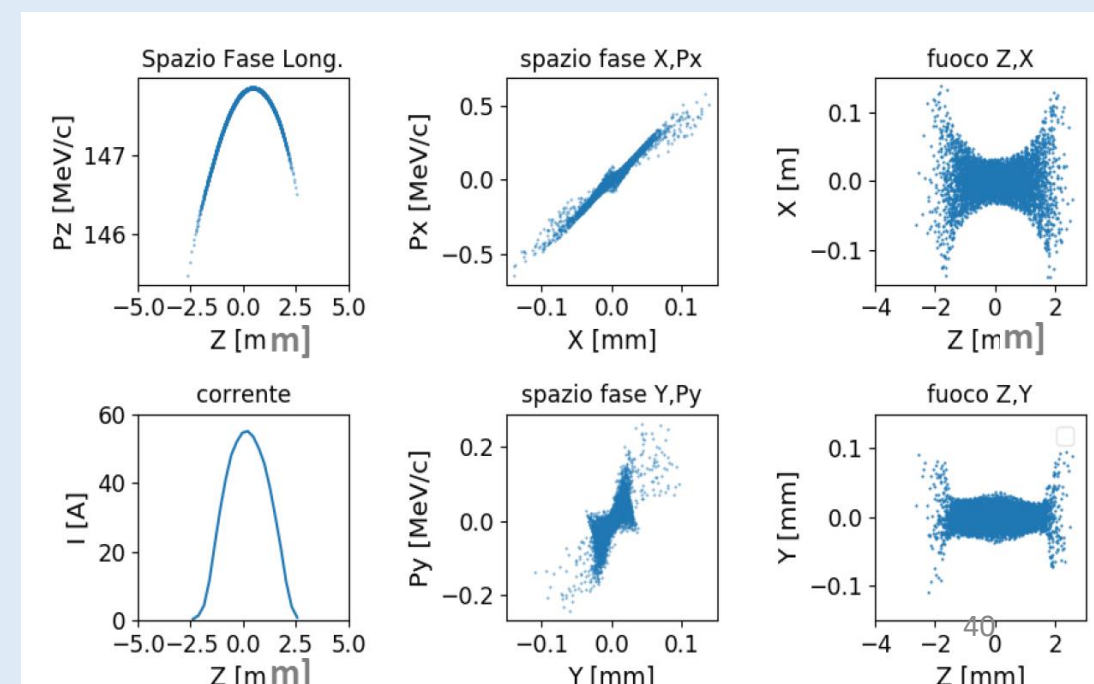
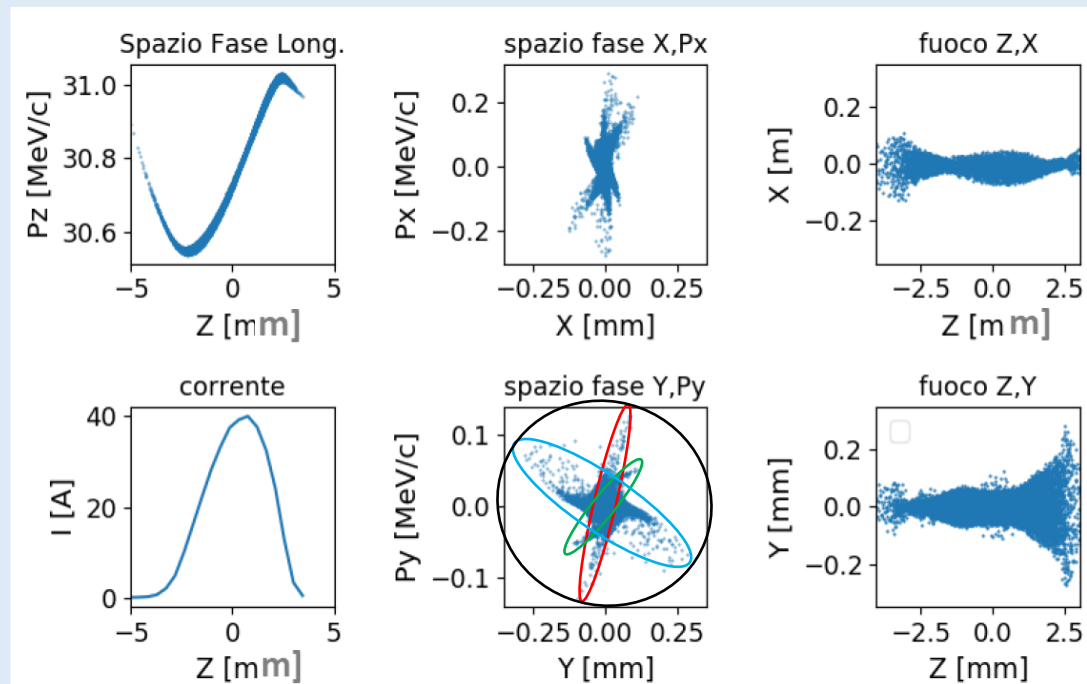
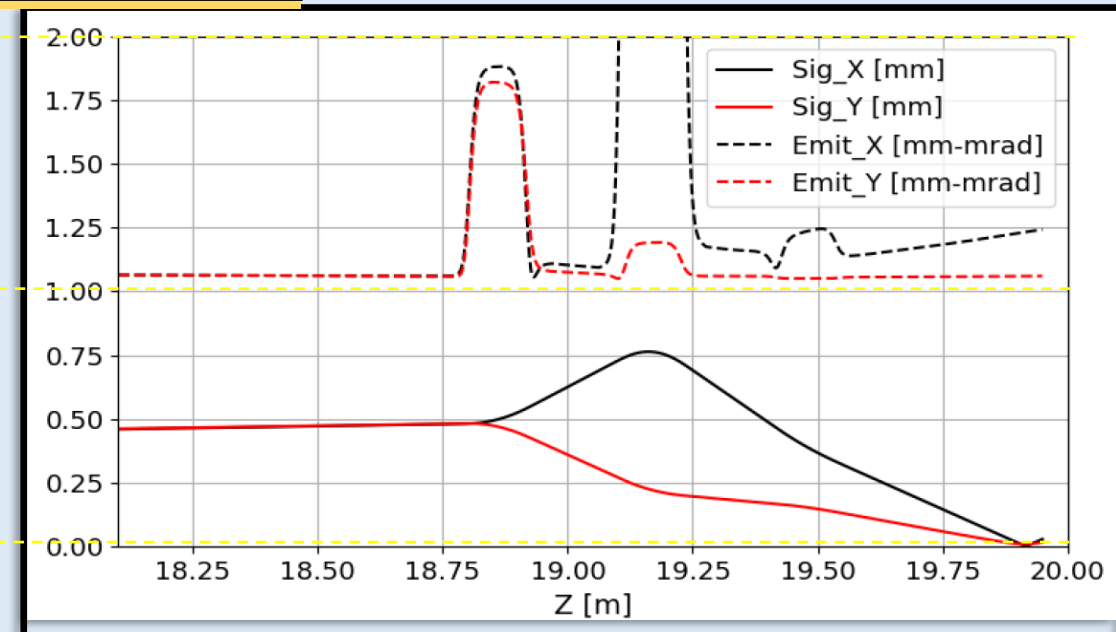
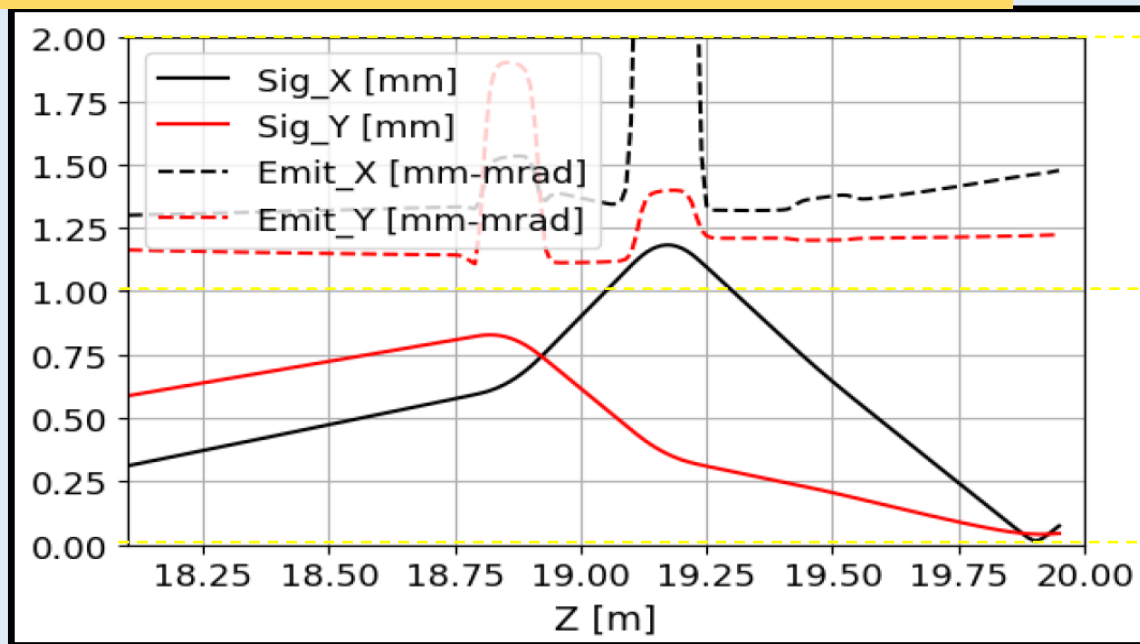


Typical beam parameters for a LINAC (150 MeV – STAR case) – 3/4



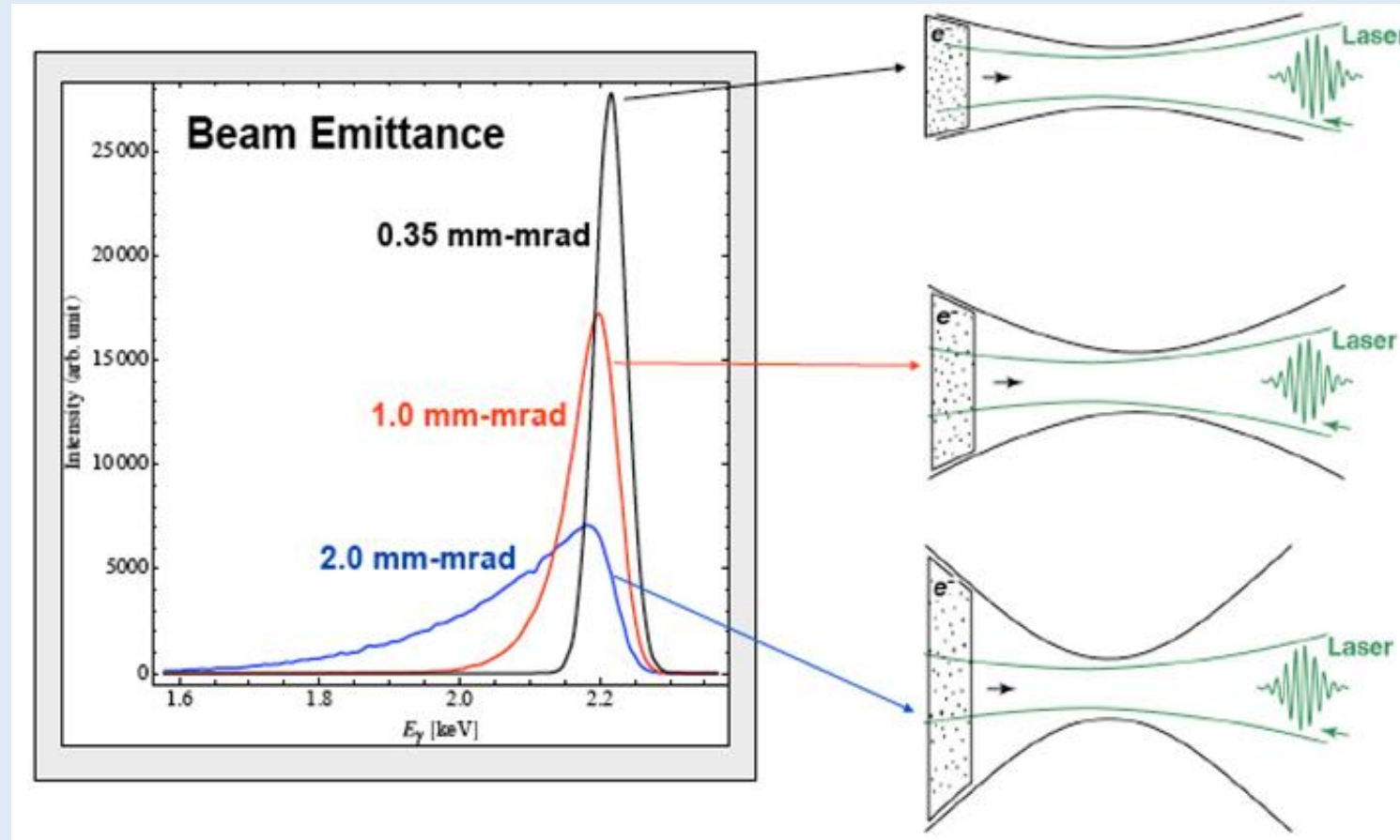
STAR – focusing channel – 30 MeV case

150 MeV case



Figurative affect of High Brightness in ICS

In Thomson/compton source – To drammmatically improve the Spectral Density

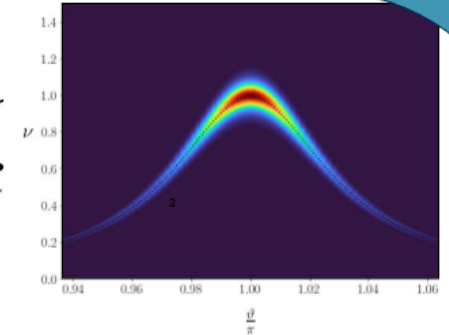
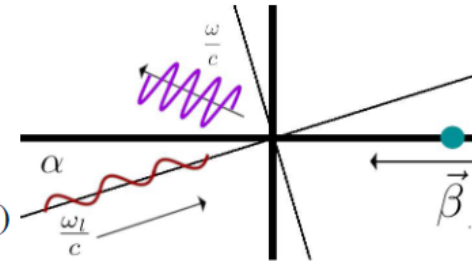
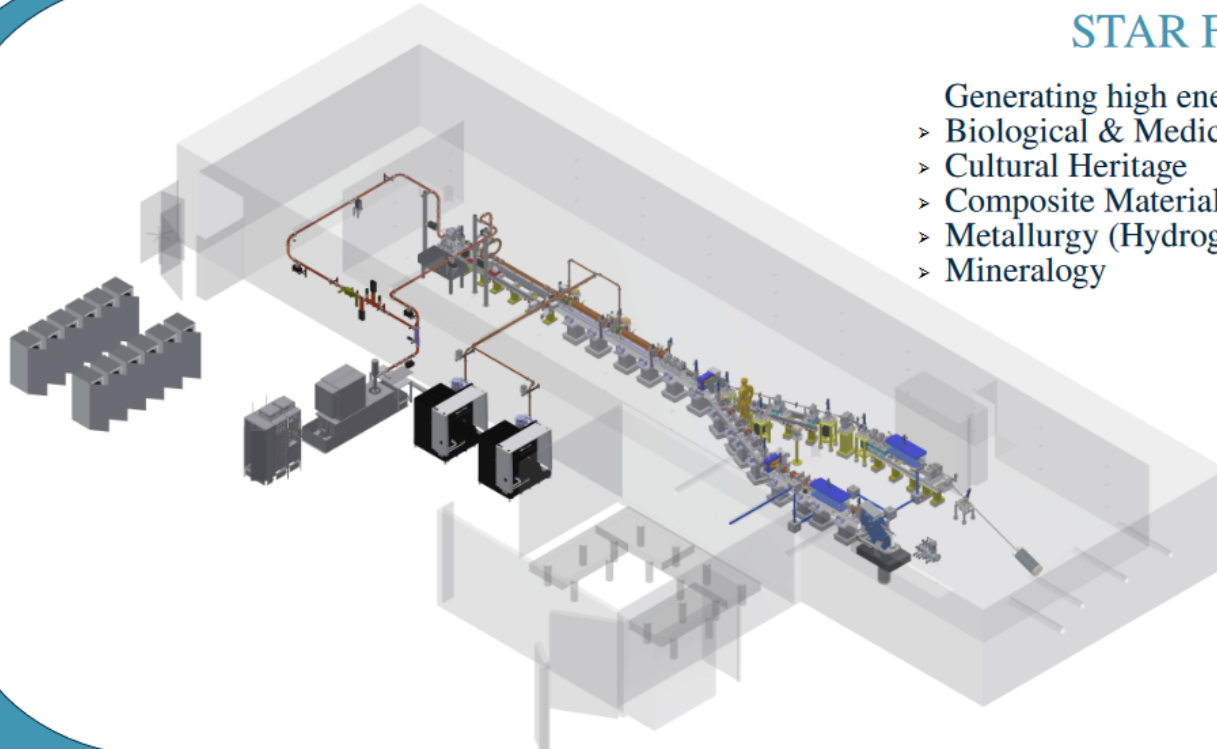


STAR HE-Linac



STAR Facility

- Generating high energy radiation for
- Biological & Medical Imaging
- Cultural Heritage
- Composite Materials
- Metallurgy (Hydrogen embrittlement)
- Mineralogy

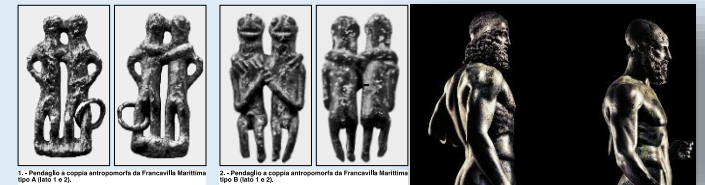


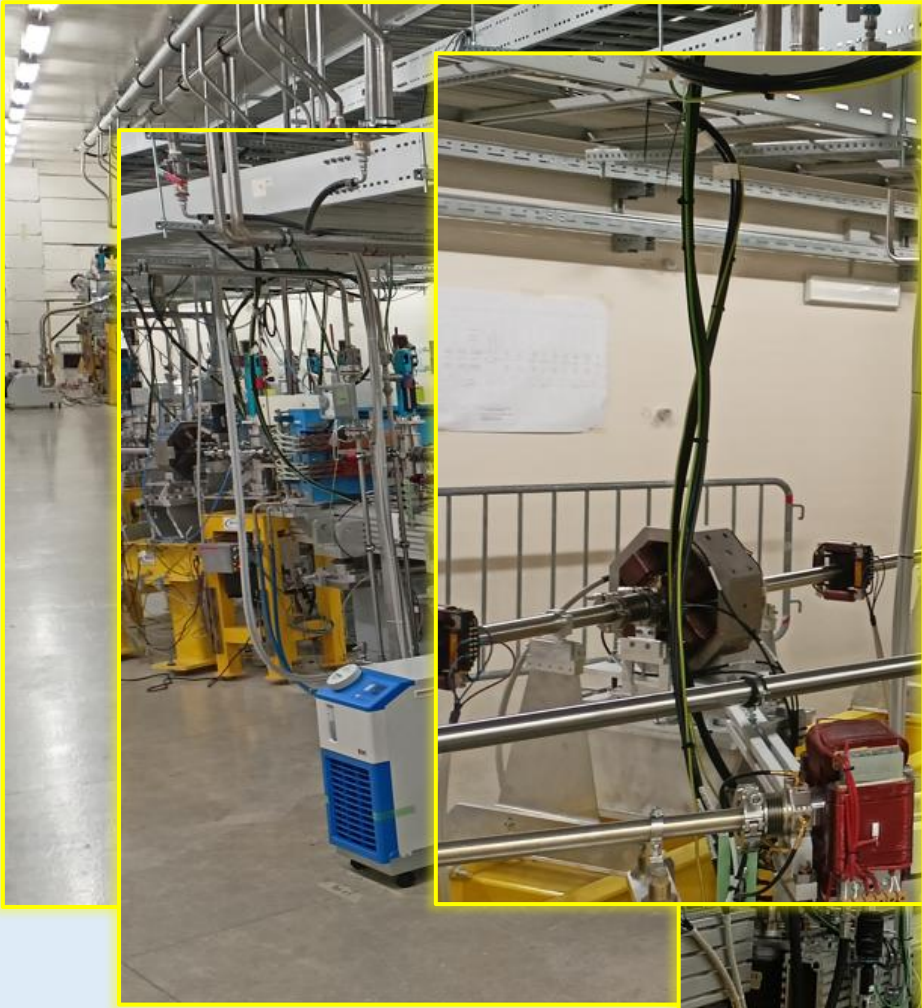
	Electron [Mev]	Photon [keV]
LE - line	23 -65	40 - 150
HE - line	40-150	25 - 350

- Electrons
- Emittance : 1 [mm mrad]
 - Charge : 100 - 500 [pC]
 - Bunch length : < 0.7 [mm]
 - Energy spread : 0.1 %, 0.05%

- (CPA) Laser
- Energy : > 0.5 [Joule]
 - Wavelength : 1030 [nm]
 - Bandwidth : 1 [nm]

REVISION	DATE	COMMENT	APPROVED
post-review	04/08/21	revised version after UniCal review/comments	A. Ghigo L. Serafini





The hangar



The Laser



Some of the more used Codes for LINAC in space-charge regime

A consistent simulation of the SC must be done by using **PIC or P-P codes**

Parmela	(Los Alamos National Lab. , L. Youg and J. Billen, PIC/P-P(?))
Tstep	(Parmela Son, from a Private Company, PIC)
Astra	(Desy, Klaus Floettmann , Free Code , PIC/P-P)
GPT	(Private company Pulsar Physics, Netherlands)
IMPACT-T and IMPACT-Z	(Berkeley Lab., Ji Qiang, Free Code)

Usually the Space-Charge is a main issue for Linac injectors up to 100 MeV. Typical applications: **FEL, Thomson/Compton sources or ultra short bunches for Plasma Wave Accelerators**)

Let's introduce:

ASTRA

A Space Charge Tracking Algorithm

Version 3.2

March 2017

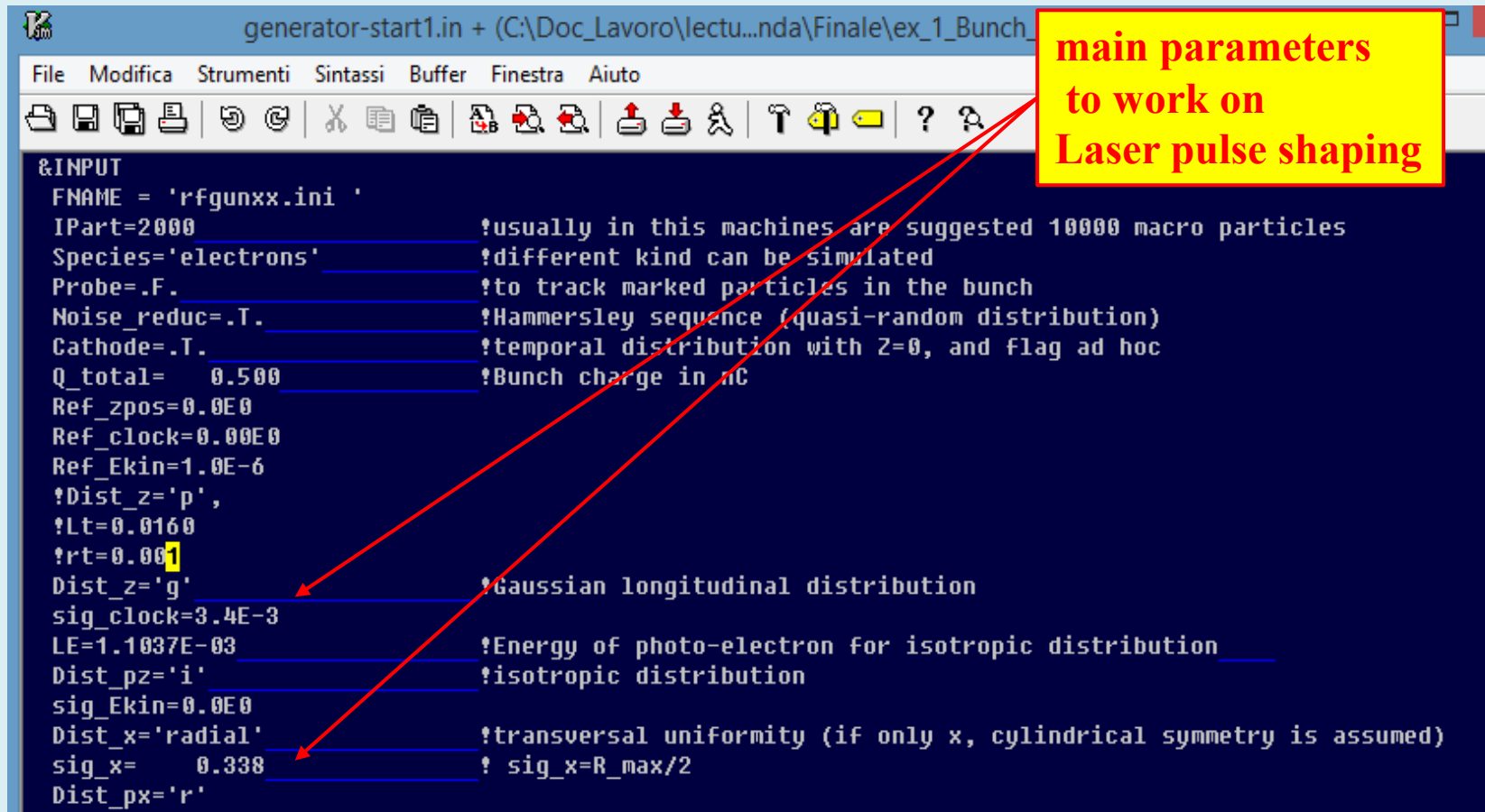
Astra input files and free parameters

Let see the **STAR-project** input files

All 3D (or 2D) pic tracking code have **two main algorithms**:

1) Bunch extraction from cathode or particles generation, 2) bunch tracker into the beam-line.

Input for e-bunch extraction:



```
&INPUT
FNAME = 'rfgunxx.ini '
IPart=2000          ?usually in this machines are suggested 10000 macro particles
Species='electrons' ?different kind can be simulated
Probe=.F.          ?to track marked particles in the bunch
Noise_reduc=.T.    ?Hammersley sequence (quasi-random distribution)
Cathode=.T.        ?temporal distribution with Z=0, and flag ad hoc
Q_total= 0.500     ?Bunch charge in nC
Ref_zpos=0.0E0
Ref_clock=0.00E0
Ref_Ekin=1.0E-6
?Dist_z='p',
?Lt=0.0160
?rt=0.001
Dist_z='g'          ?Gaussian longitudinal distribution
sig_clock=3.4E-3
LE=1.1037E-03     ?Energy of photo-electron for isotropic distribution
Dist_pz='i'        ?isotropic distribution
sig_Ekin=0.0E0
Dist_x='radial'    ?transversal uniformity (if only x, cylindrical symmetry is assumed)
sig_x= 0.338      ? sig_x=R_max/2
Dist_px='r'
```


Astra input files and free parameters

```
pls-start.in (C:\...Linac_0_8.8m) - GVIM
File Modifica Strumenti Sintassi Buffer Finestra Aiuto
&NEWRUN
Head='Gun120MV/m, 0.5nC, START_60MeV case'
RUN= 2 ,
Loop=F, Nloop=2
Distribution = 'rfgunxx.ini '
Xoff=0.0E0, Yoff=0.0E0
Lmagnetized=F
EmitS=T
PhaseS=T
TrackS=T
RefS=T
TcheckS=T
CathodeS=T
TRACK_ALL=T, PHASE_SCAN=F, AUTO_PHASE=T
check_ref_part=F,
ZSTART=0.0, ZSTOP=8.8
Zemit=2050
Zphase=2
Max_step=200000
H_max=0.0005
H_min=0.0000
/

&OUTPUT
/

&SCAN
LScan=F,
Scan_para='Phi(1)'
S_min=102, S_max=118, S_num=9
FOM(1)='bunch charge'
FOM(2)='hor emit'
FOM(3)='bunch length'
FOM(4)='hor spot size'
FOM(5)='phi end'
/

24,5 Cim
```

```
&CHARGE
Loop=F
LSPCH=T
!Nrad=15, Nlong_in=60
Nrad=15, Nlong_in=25
Cell_var=2.0
min_grid=0.0
Max_scale=0.1
Max_count=100
Lmirror=T
Linert=T
/

&Aperture
/

&FEM
/

&CAVITY
Loop=F,
LEfield=T
FILE_Efield(1) = 'new45.dat'
Nue(1)=2.856, MaxE(1)=120.0, Phi(1)=-3.741, c_pos(1)=0.0
FILE_Efield(2) = 'TWS_sparc.dat'
Nue(2)=2.856, MaxE(2)=24.1, Phi(2)=-1.339, c_pos(2)=1.75
c_numb(2)=84
/

&SOLENOID
Loop=F
LBfield=T,
FILE_Bfield(1)='GUNSOL_SPARC_+---.poi'
MaxB(1)=0.310647, S_pos(1)=0.19575
FILE_Bfield(2)='SOL1.txt', MaxB(2)=0.493, S_pos(2)=9.35
FILE_Bfield(3)='SOL1.txt', MaxB(3)=-0.769, S_pos(3)=10.0
/

&QUADRUPOLE
Questa è già l'ultima modifica 60,31 9
```

Astra main in/out files format

	1	2	3	4	5	6	7	8	9	10
Parameter	x	y	z	px	py	pz	clock	macro charge	particle index	status flag
Unit	m	m	m	eV/c	eV/c	eV/c	ns	nC		

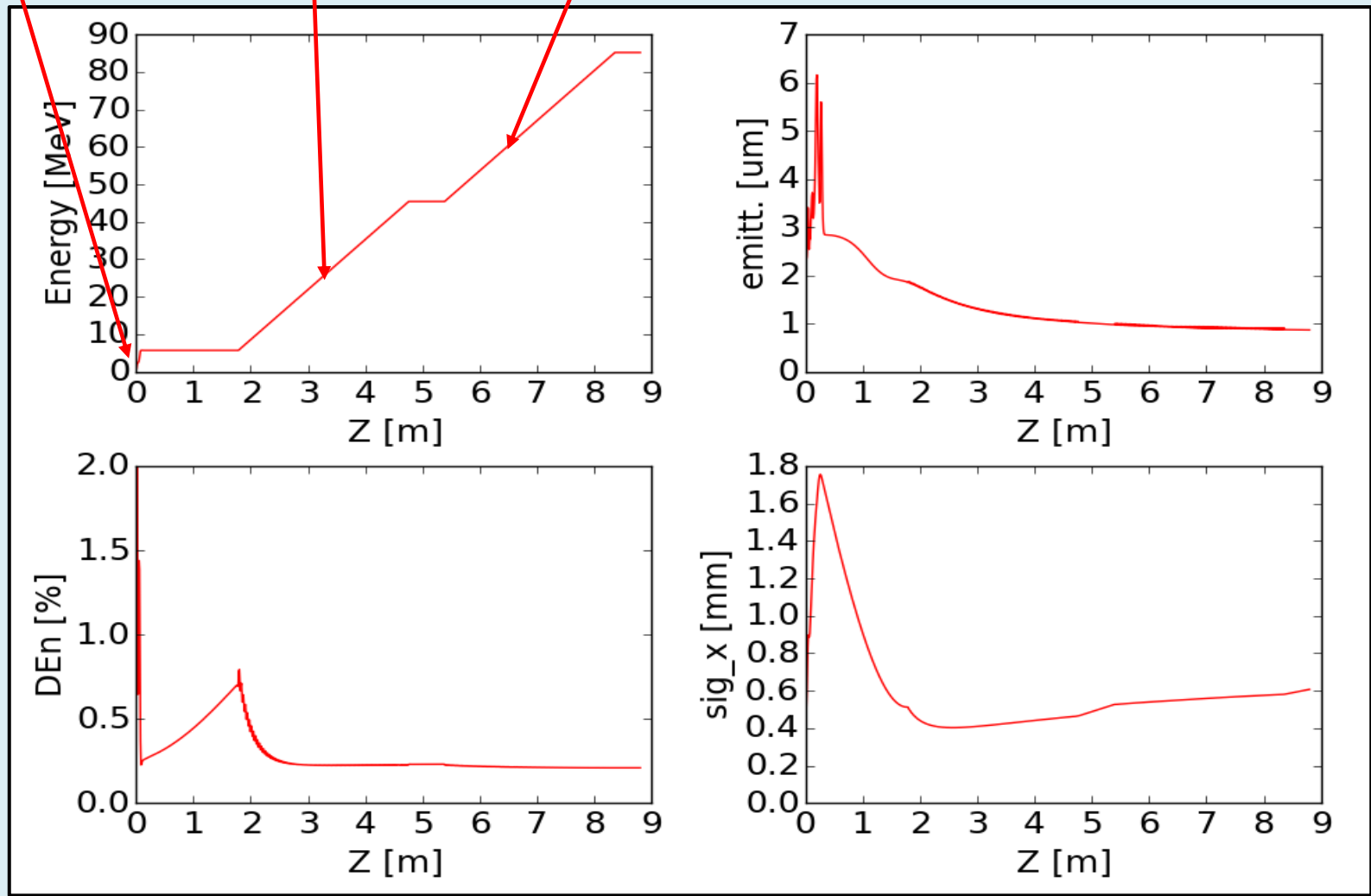
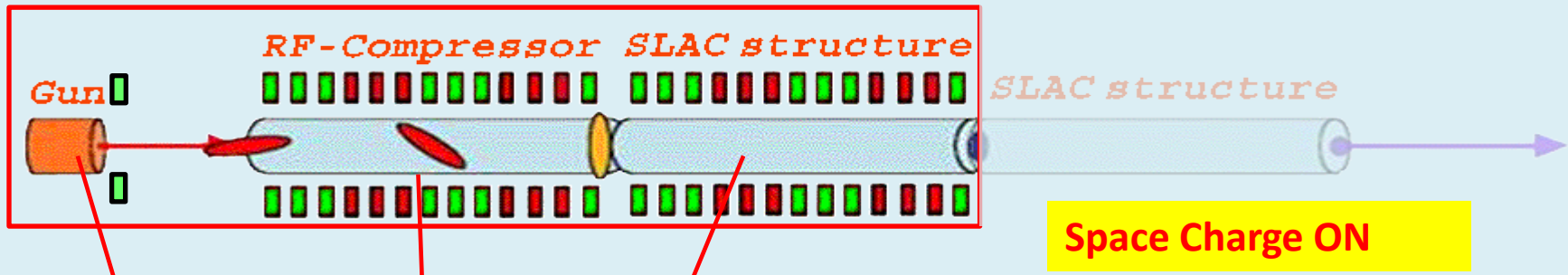
Table 1: Structure of particle distribution files.

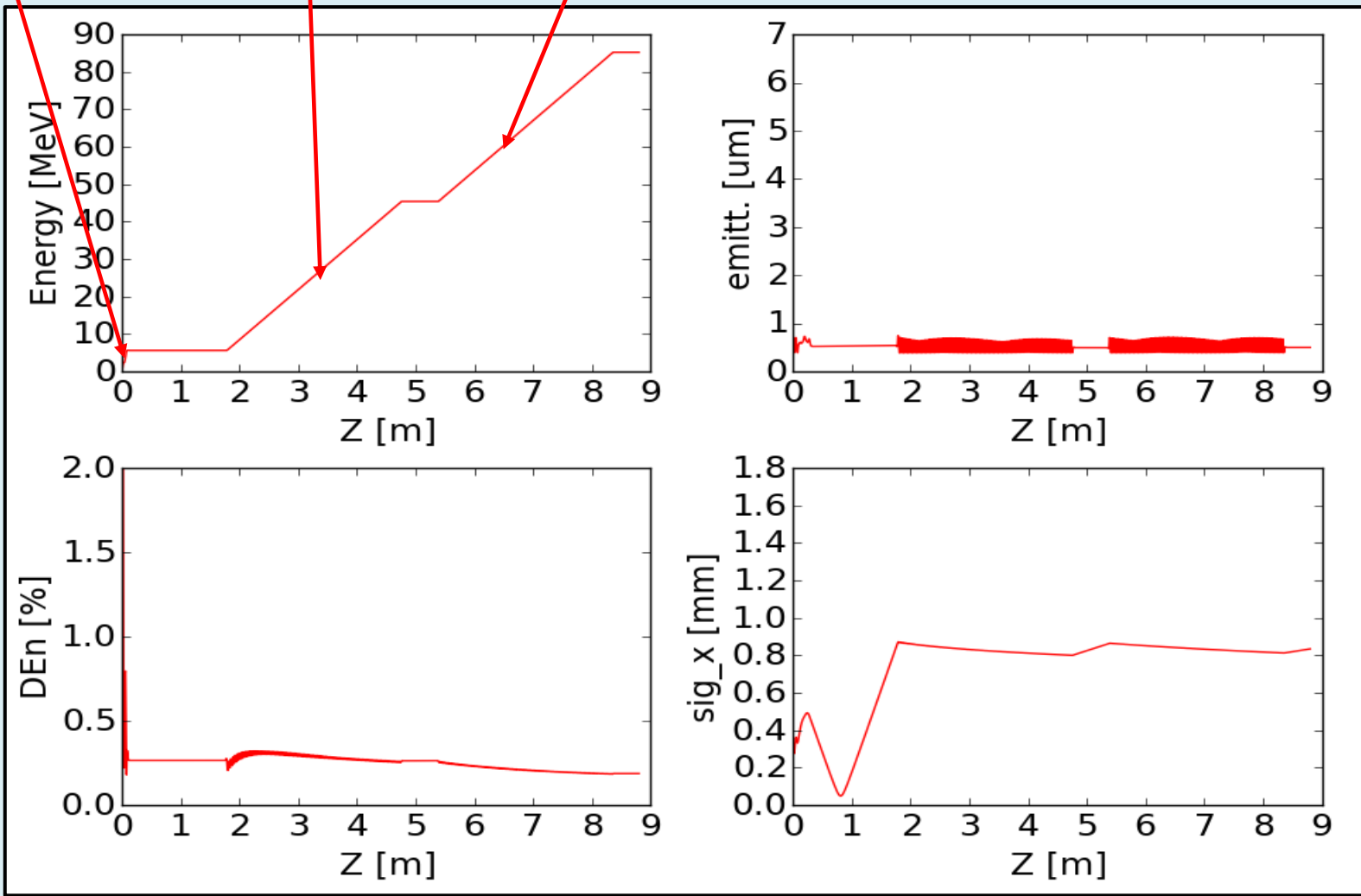
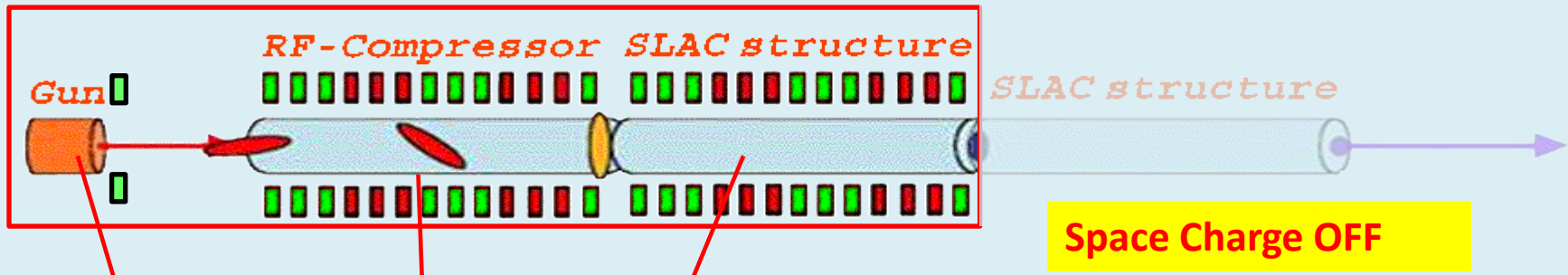
The first line of the file defines the coordinates of the reference particle in absolute coordinates. It is recommended to refer it to the bunch center. **Longitudinal particle coordinates, i.e. z, pz and t are given relative to the reference particle.** (If the

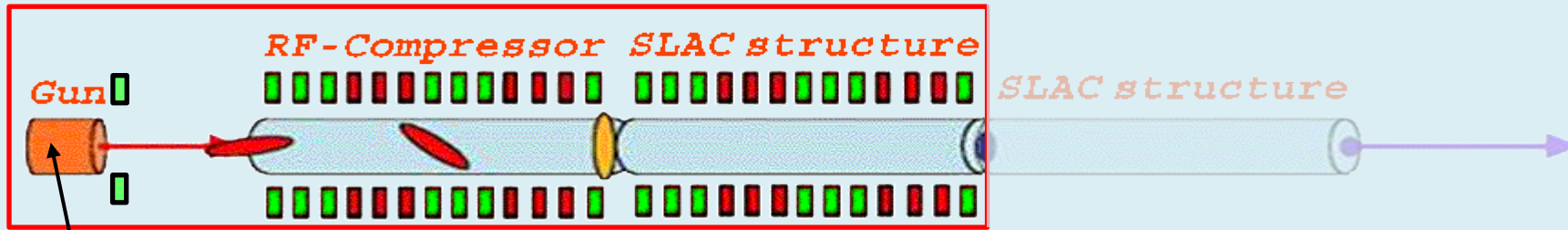
Status flag	Comment	Status
-99 ¹	average position of distribution	will not be tracked
-95	ref. particle only; $Z_0 > Z_{\text{Stop}}$	lost
-94	ref. particle only; more than Max_Step steps	lost
-92 ²	probe rejected by space charge at the cathode	lost
-91 ²	rejected by space charge at the cathode	lost
-90	probe particle before Z_{min}	lost
-89	particle before Z_{min}	lost
-86 ³	probe particle traveling backwards	lost
-85 ³	particle traveling backwards	lost
-31	particle discarded by user	lost
-30	particle preliminary discarded by user	lost
-22	probe secondary electron, lost on aperture	lost
-21	secondary electron, lost on aperture	lost
-20	passive probe particle, lost on aperture	lost

Name	1	2	3	4	5	6	7	8	9	Format	
ref	z m	t ns	pz MeV/c	$\frac{dE}{dz}$ MeV/m	Larmor angle rad	X_{off} mm	Y_{off} mm	px eV/c	py eV/c	1P,9E12.4 1P,9E20.12	
track	seq. numb	stat. flag	z m	x mm	y mm	Ez V/m	Er, or Ex V/m	0.0, or Ey V/m		2I5,1P,6E12.4	
Cathode	z m	t ns	long. sp. ch. field on cathode V/m	acc. field on cathode V/m	charge nC	min. grid position m	max grid position m	emission flag		1P,7E12.4,L3	
Fields	z m	t ns	Cavity gradient (i) (i = 1...number of cavities N_C) MV/m								1P, N_C E12.4
tcheck	z m	t ns	$\frac{\sigma_{r0}^{nr(r)}}{\sigma_r}$	$\frac{\sigma_{z0}^{nz(z)}}{\sigma_z}$	$\frac{\gamma^{nr(\gamma)}}{\gamma_0}$	$\frac{\sigma_{r0}^{nz(r)}}{\sigma_r}$	$\frac{\sigma_{z0} \cdot \gamma_0^{nz(\gamma)}}{\sigma_z \cdot \gamma}$	scaling counter		1P,7E12.4,I10	
Xemit	z m	t ns	X_{avr} mm	X_{rms} mm	X'_{rms} mrad	$\epsilon_{x,\text{norm}}$ π mrad mm	$X \cdot X'_{\text{avr}}$ mrad			1P,7E12.4	
Yemit	z m	t ns	Y_{avr} mm	Y_{rms} mm	Y'_{rms} mrad	$\epsilon_{y,\text{norm}}$ π mrad mm	$Y \cdot Y'_{\text{avr}}$ mrad			1P,7E12.4	
Zemit	z m	t ns	E_{kin} MeV	Z_{rms} mm	ΔE_{rms} keV	$\epsilon_{z,\text{norm}}$ π keV mm	$Z \cdot E'_{\text{avr}}$ keV			1P,7E12.4	

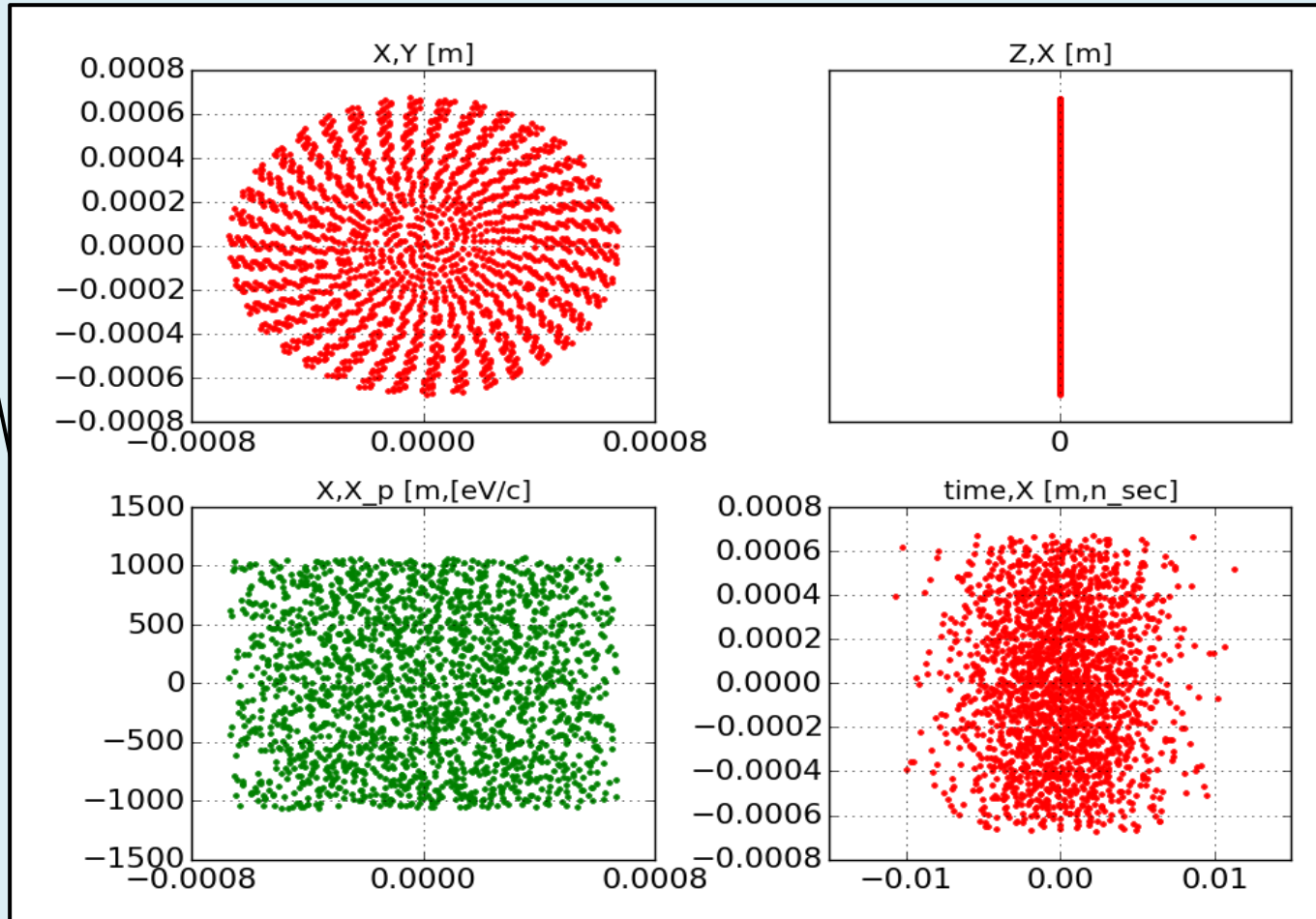
erture	lost
ost on aperture	lost
erture	lost
he cathode	not yet started
ode	not yet started
	not yet started
the cathode	not yet started
thode	not yet started
	tracking ⁴
	tracking ⁴
	tracking
	tracking
	tracking
s of generation 1,	tracking
eneration 1, 2...10	tracking

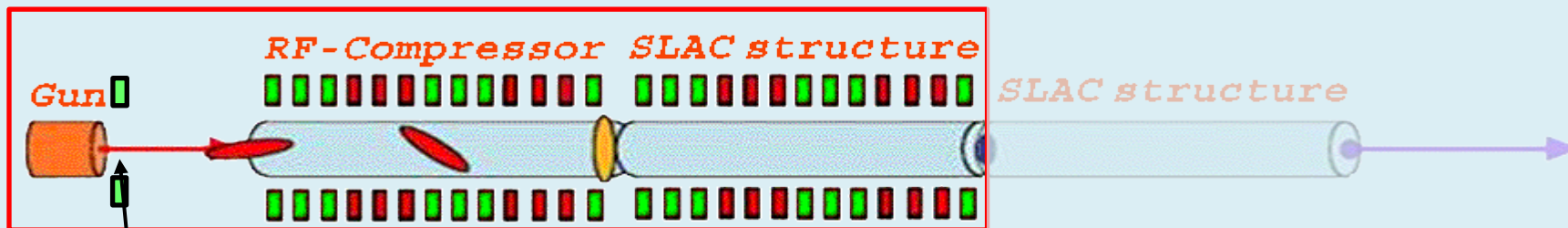




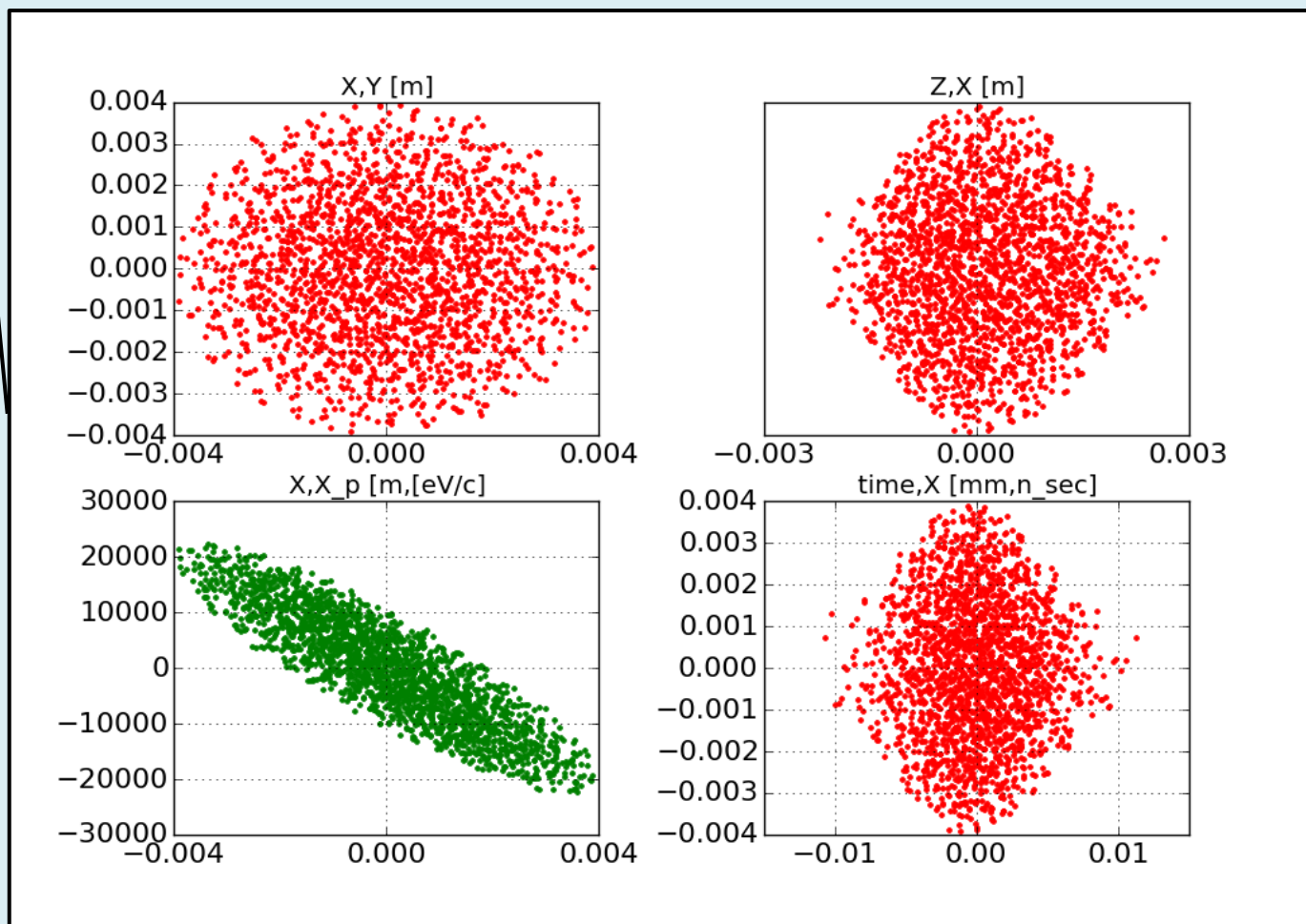


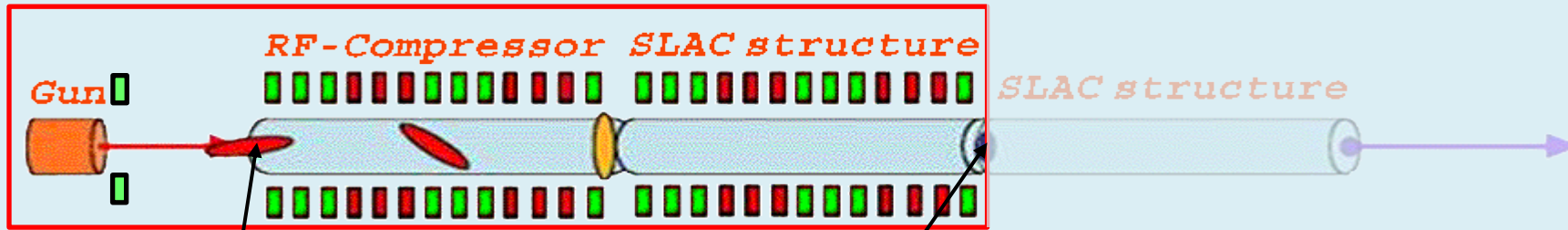
On the cathode



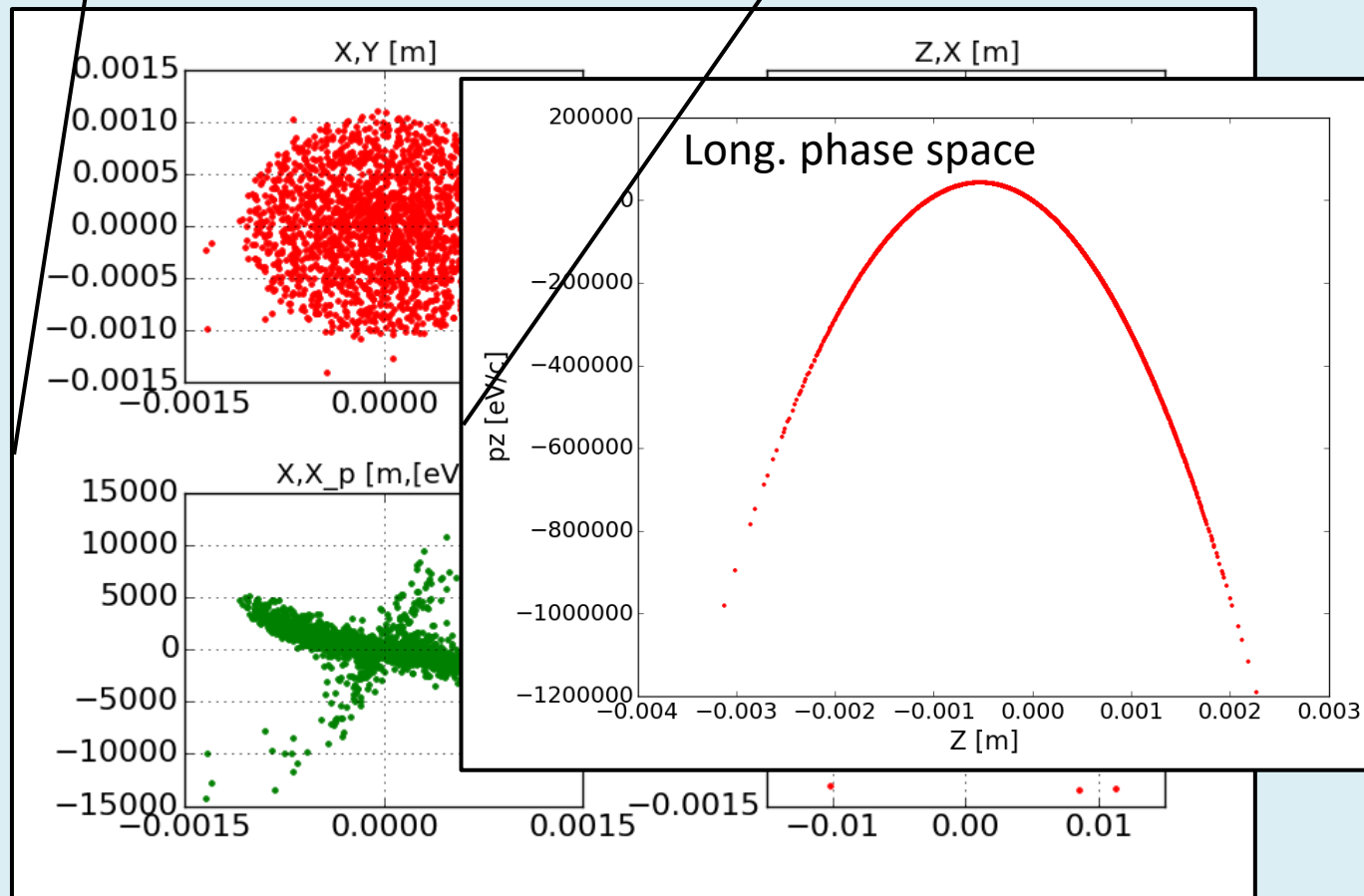


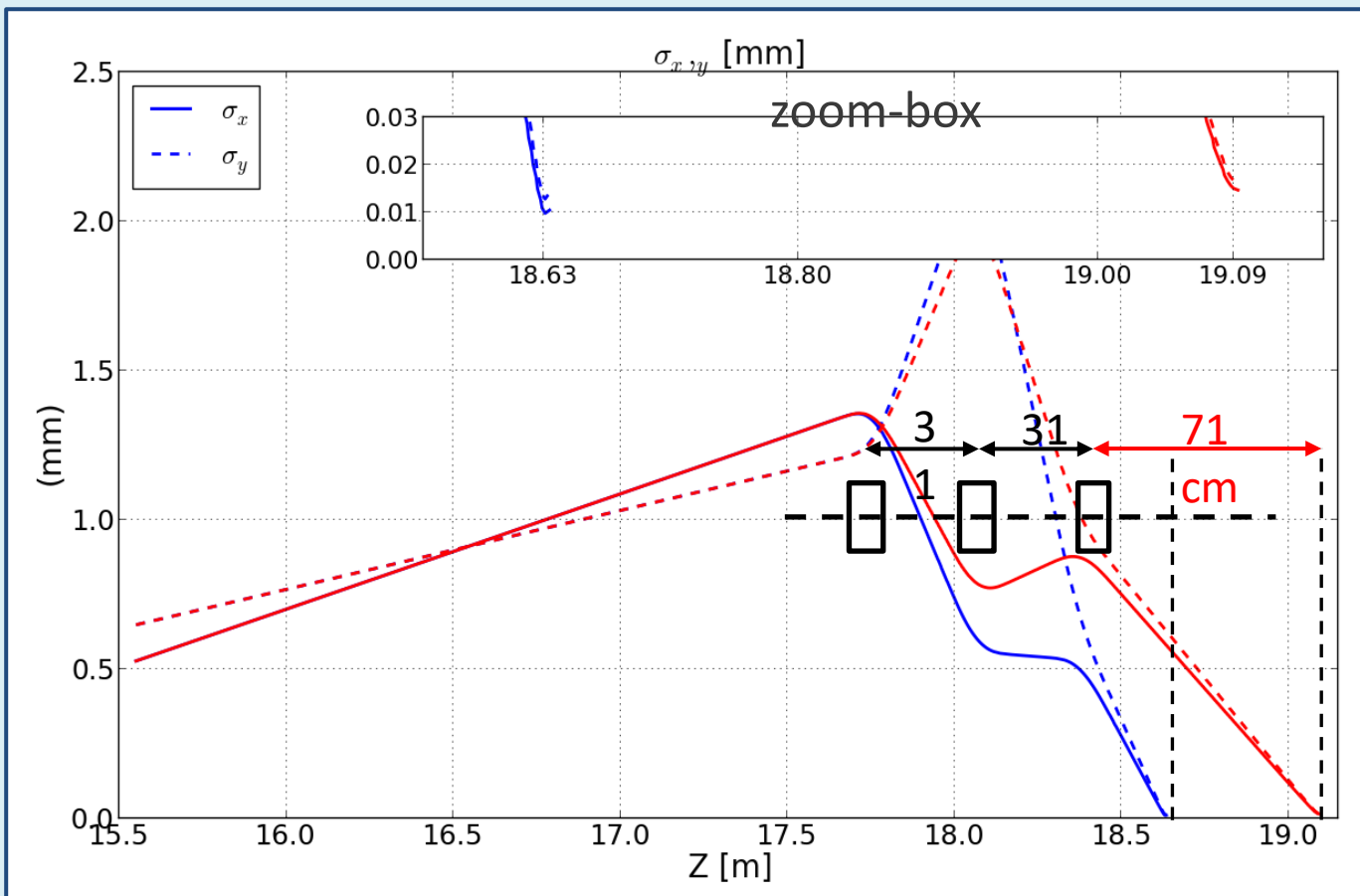
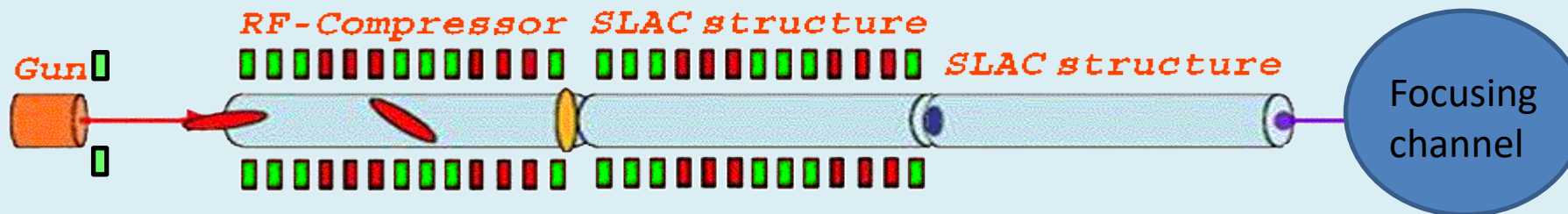
Gun Exit

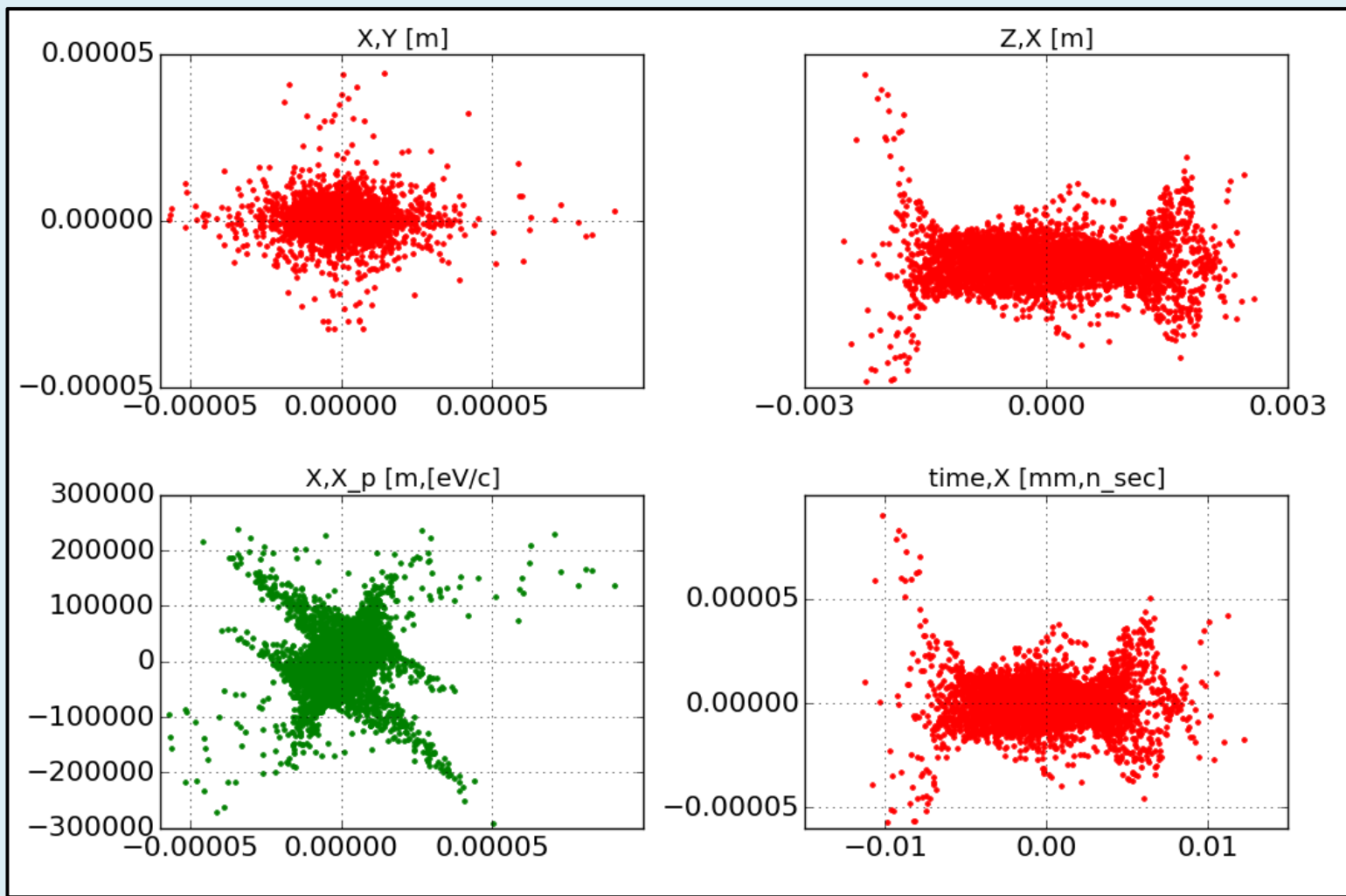
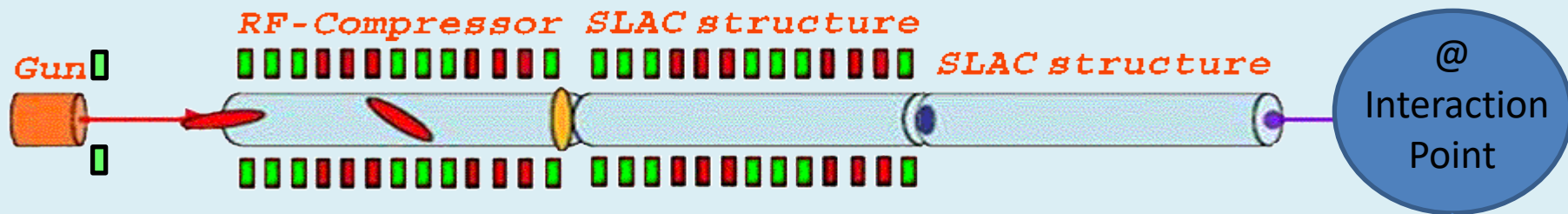




Linac entrance







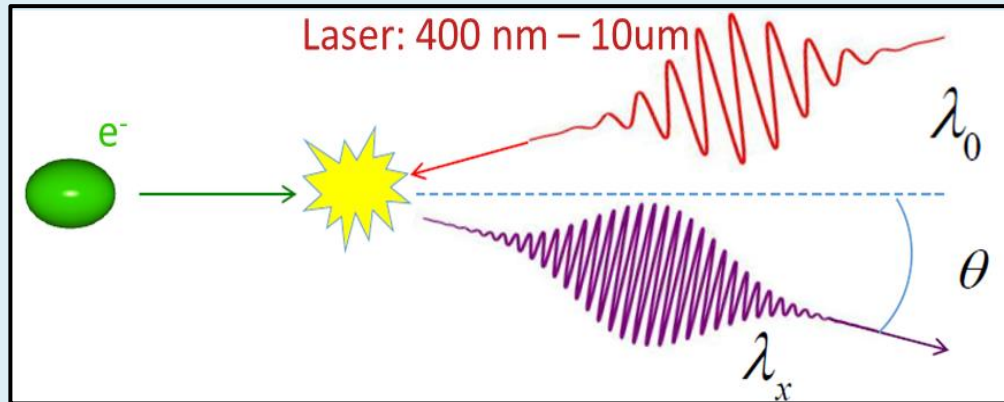
A Quantum Code, To simulate the Electron-Photon Bunches scattering

CAIN SIMULATION CODE

GAMMA BEAM

CAIN code
developed by K. Yokoya
Monte Carlo code
based on QED
Landau-Lifshitz approach

The Inverse Scattering Compton (ICS)

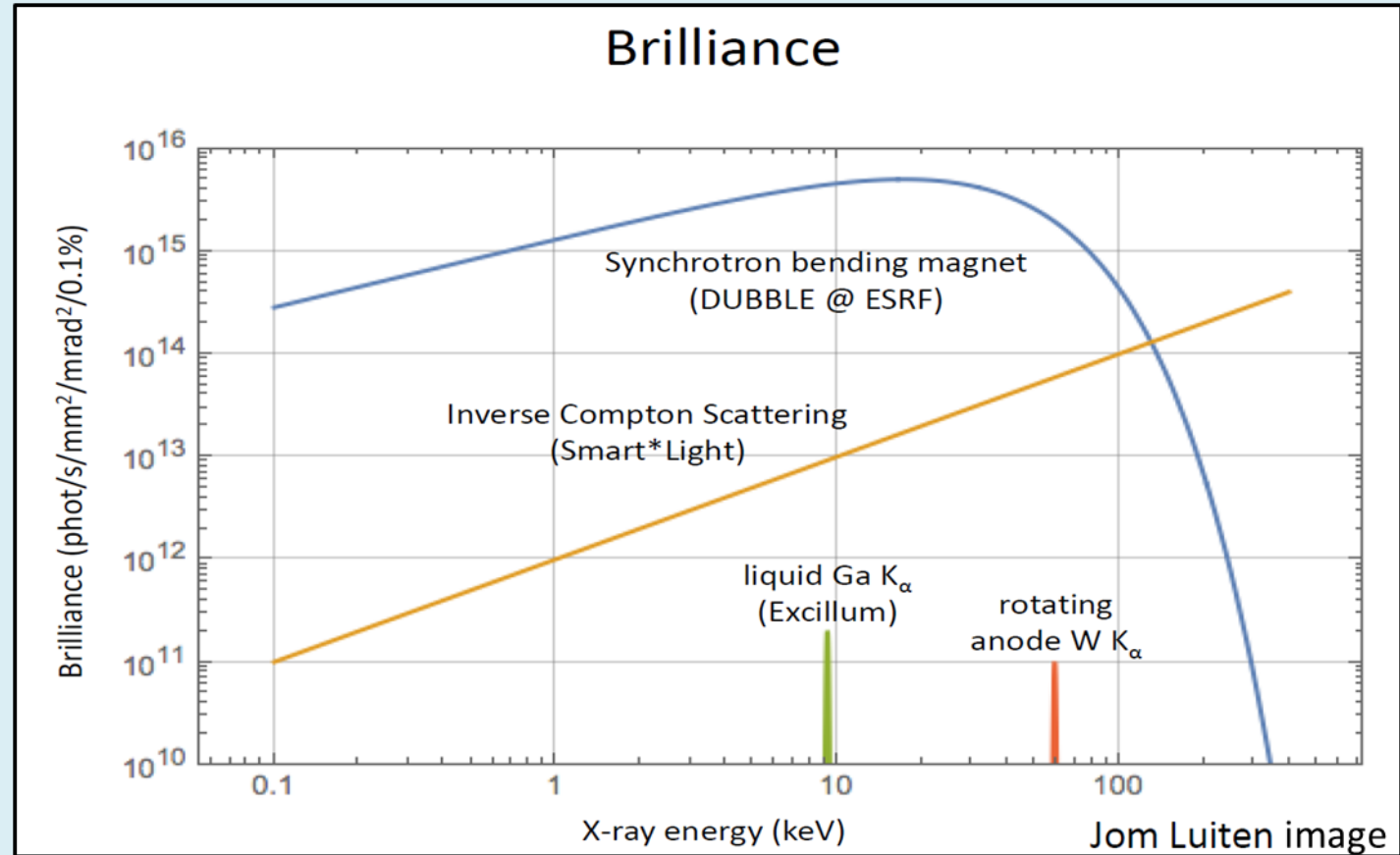


$$E'_{\text{ph}}(\theta) = \frac{4E_{\text{ph}}\gamma^2}{1 + X + \gamma^2\theta^2}$$

$$X = \frac{4E_e E_{\text{ph}}}{(m_0 c^2)^2} = \frac{4\gamma E_{\text{ph}}}{m_0 c^2} = 4\gamma^2 \frac{E_{\text{ph}}}{E_e}$$

$X \ll 1$

$$E'_{\text{ph}}(\theta = 0) = 4E_{\text{ph}}\gamma^2$$



Number of x-ray photons

$$N_{x,\text{total}} = N_e N_0 \frac{\sigma_T}{2\pi w_0^2}$$

Cross section

Lasere electron I-spot

Cain input file

```
ALLOCATE MP=50000;  
  
SET photon=1, electron=2, positron=3, mm=1D-3, micron=1D-6,  
nm=1D-9, mu0=4*Pi*1D-7, psec=1e-12*cvel,
```

```
sigz=0.000281,  
ntcut=5,  
laserwl=515.000000*nm,  
pulseE=0.400000,  
sigLr=14.000000*micron,  
w0=2*sigLr,  
rayl=Pi*w0^2/laserwl,  
sigt=1.500000*psec,  
angle=0.130900,  
tdl=1.0,  
powerd=(2*pulseE*cvel)/[Pi*sigt*sqrt(2*Pi)*w0^2],
```

laser parameters

```
SET MsgLevel=1;
```

```
SET Rand=5*40003.000000;
```

```
BEAM FILE='exp.dat';
```

incoming electron beam

```
LASER LEFT, WAVEL=laserwl, POWERD=powerd,  
  
TXYS=(0.000000, 0.000000, 0.000000, 0.000000),  
E3=(-Sin(angle), 0.0, -Cos(angle)), E1=(0,1,0),  
RAYLEIGH=(rayl,rayl), SIGT=sigt, GCUTT=ntcut,  
STOKES=(0.000000, 0.000000, 1.000000),  
TDL=(tdl,tdl);
```

**laser reference frame
and polarization**

```
LASERQED COMPTON, NPH=0;
```

linear compton scattering

```
SET MsgLevel=0; FLAG OFF ECHO;  
SET Smesh=sigt/3;  
SET it=0;
```

```
PUSH Time=(-ntcut*(sigt+sigz),ntcut*(sigt+sigz),250);  
IF Mod(it,20)=0;  
PRINT it, FORMAT=(F6.0, '-th time step');  
PRINT STAT, SHORT;  
ENDIF;  
SET it=it+1;  
ENDPUSH;
```

time evolution

```
DRIFT T=0;
```

```
WRITE BEAM, KIND=(electron), FILE='cain_output_electrons.dat';
```

```
WRITE BEAM, KIND=(photon), FILE='cain_output_photons_20000.dat';
```

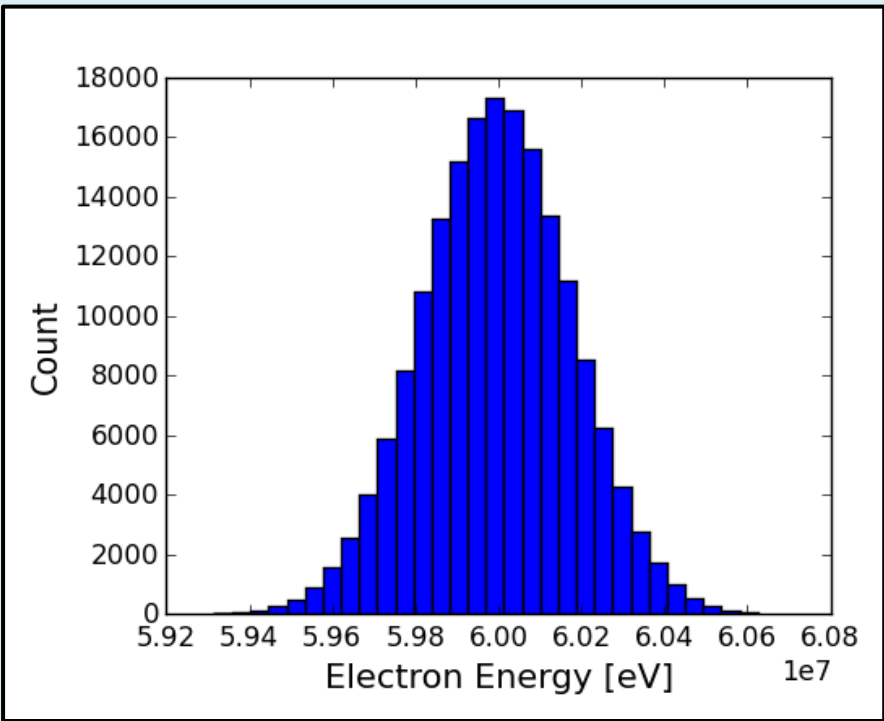
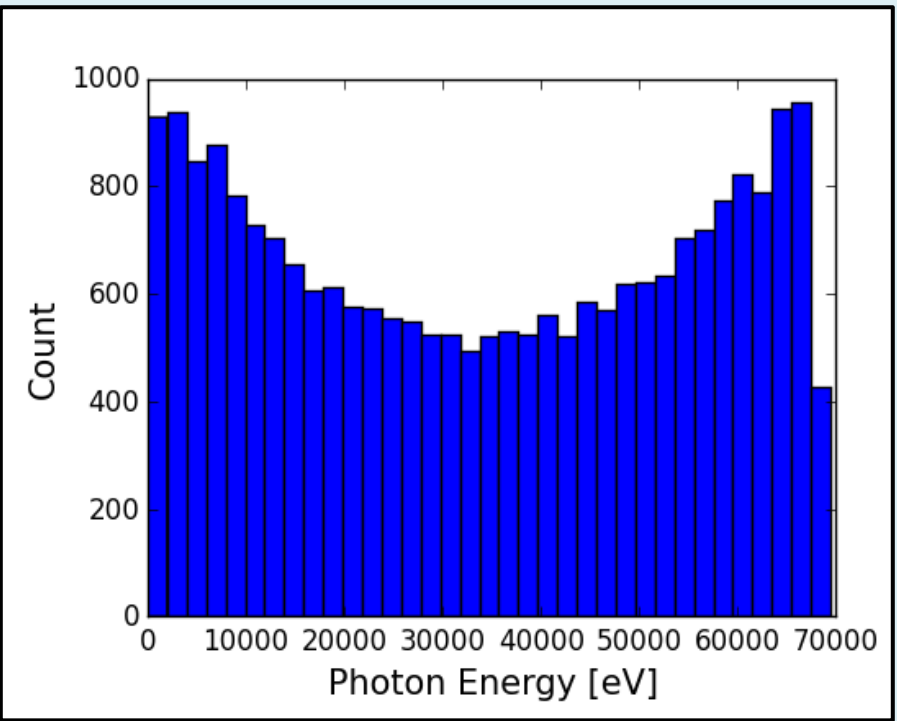
Cain input file: to define the bunch virtually

```
1 HEADER 'Compton Scattering';
2 !! Very high energy Compton scattering including nonlinear scattering.
3 ALLOCATE MP=350000;
4 SET photon=1, electron=2, positron=3,
5 mm=1e-3, micron=1e-6, nm=1e-9, mu0=4*Pi*1e-7, psec=1e-12*Cvel,
6 nsec=1e-9*Cvel;
7 ! define variables for electron beam
8 SET ee=60.0D6, gamma=ee/Emass, an=3.125D9, sigz=1.0*mm,
9 betax=2.7*mm, betay=2.7*mm, emitx=1.01D-6/gamma, emity=1.01D-6/gamma, sige=0.00003,
10 sigx=.Sqrt(emitx*betax), sigy=.Sqrt(emity*betay),
11 ntcut=3.0;
12 ! define variables for laser
13 SET laserwl=1.*micron, lambar=laserwl/(2*Pi), omegal=Hbarc/lambar,
14 rlx=1.5*mm, rly=1.5*mm, sigt=0.9*mm, !! Rayleigh length and pulse length
15 pulseE=0.4000, !! pulse energy in Joule
16 powerd=pulseE*Cvel/[Pi*lambar*sigt*.Sqrt(2*Pi*rlx*rly)],
17 xisq=powerd*mu0*Cvel*(lambar/Emass)^2, xi=.Sqrt(xisq),
18 eta=omegal*ee/Emass^2, lambda=4*eta,
19 angle=0.;
20 SET MsgLevel=1;
21 BEAM RIGHT, KIND=electron, NP=180000, AN=an, E0=ee,
22 TXYS=(0,0,0,0), GCUTT=ntcut,SIGE=sige,
23 BETA=(betax,betay), FMIT=(emitx,emity), SIGT=sigt, SPIN=(0,0,1);
24 LASER LEFT, WAVEl=laserwl, POWERD=powerd,
25 TXYS=(0,0,0,0),
26 E3=(0,-Sin(angle),-Cos(angle)), E1=(1,0,0),
27 RAYLEIGH=(rlx,rly), SIGT=sigt, GCUTT=ntcut, STOKES=(0,0,-1);
28 LASERQED COMPTON, NPH=0, XIMAX=1.1*xi, LAMB DAMAX=1.1*lambda,
29 ENHANCE=1, PMAX=0.5;
30 SET MsgLevel=0; FLAG OFF ECHO;
31 SET Smesh=sigt/3;
32 SET emax=1.001*ee, wmax=emax;
33 SET it=0;
34 PRINT CPUTIME;
35 PUSH Time=(-ntcut*(sigt+sigz),ntcut*(sigt+sigz),500);
36 IF Mod(it,50)=0;
37 PRINT it, FORMAT=(F6.0,'-th time step'); PRINT STAT, SHORT;
38 ENDIF;
39 SET it=it+1;
40 ENDPUSH;
41 PRINT CPUTIME;
42 ! Pull all particles to the back to the focal point
43 DRIFT S=0;
44 WRITE BEAM, KIND=(photon), FILE='temp_phot.dat';
45 WRITE BEAM, KIND=(electron), FILE='temp_el.dat';
46 PRINT STAT;
47 PLOT HIST, RIGHT, KIND=electron, H=En/1D9, HSCALE=(0,emax/1e9,50),
48 TITLE='Right-Going Electron Energy Spectrum';
```

STAR Linac,

X-ray source @ 60 MeV – recoil << of the energy spread $\Delta\gamma/\gamma$

Let consider an electron-bunch with an $\Delta\gamma/\gamma$ of 0.003 , $\lambda_{LASER}=1\mu m$ (0.4 J), $W_0=20 \mu m$



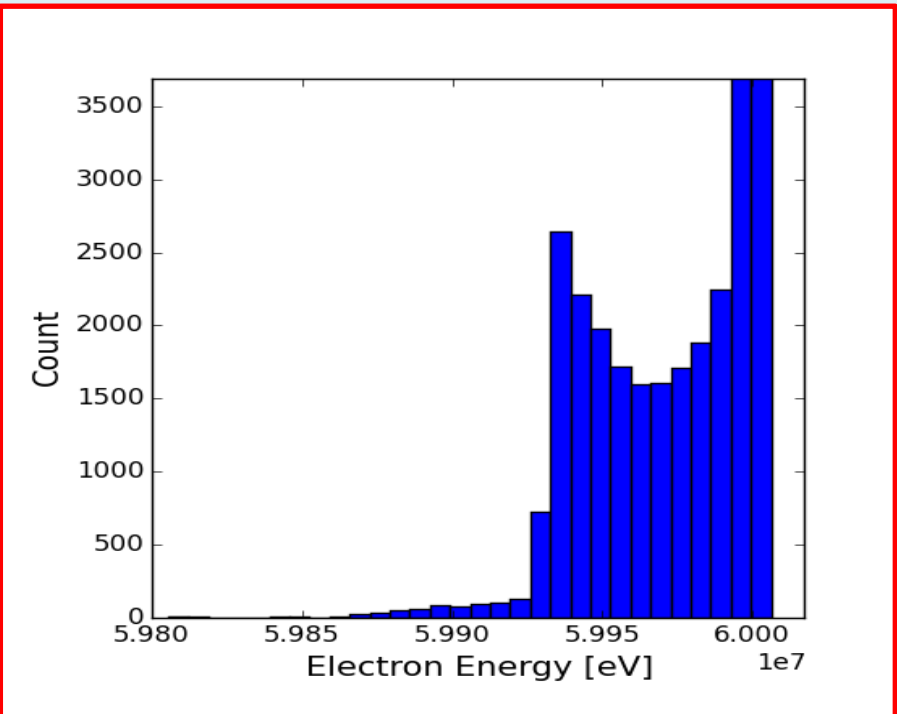
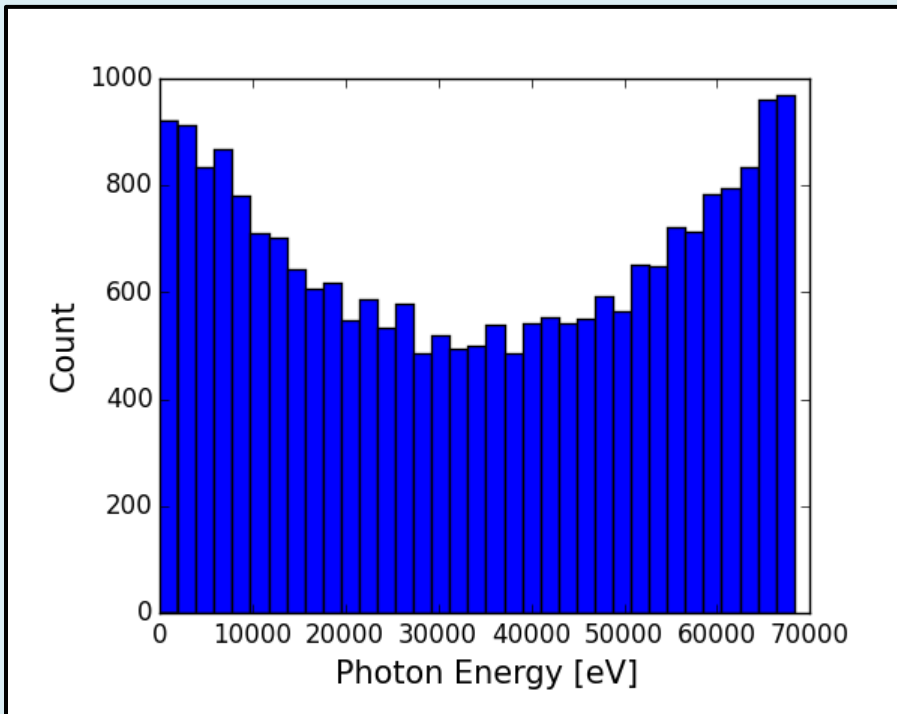
$$E_{max} = \frac{4\gamma^2 h\nu}{1+X}, \quad X = \frac{4\gamma h\nu}{511 \text{ keV}}$$

$$E_{max} = 71.340 \text{ keV}, \quad X = 0.00116 \text{ (recoil)}$$

$$E_{ph} (\lambda_{laser} = 1 \mu m) = 1.24 \text{ eV}$$

STAR Linac, X-ray source @ 60 MeV – recoil as before but $> \Delta\gamma/\gamma$

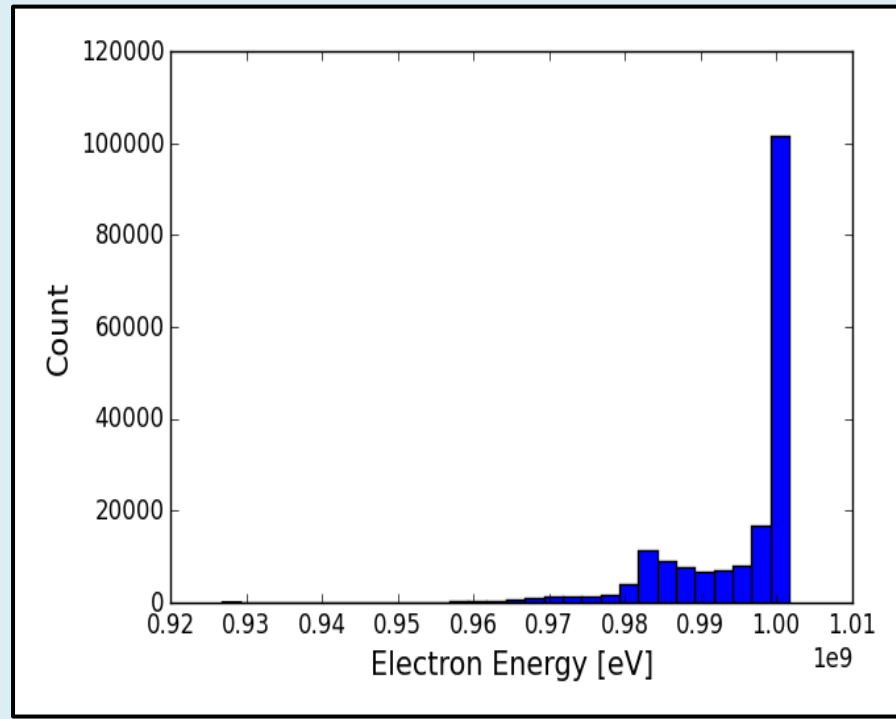
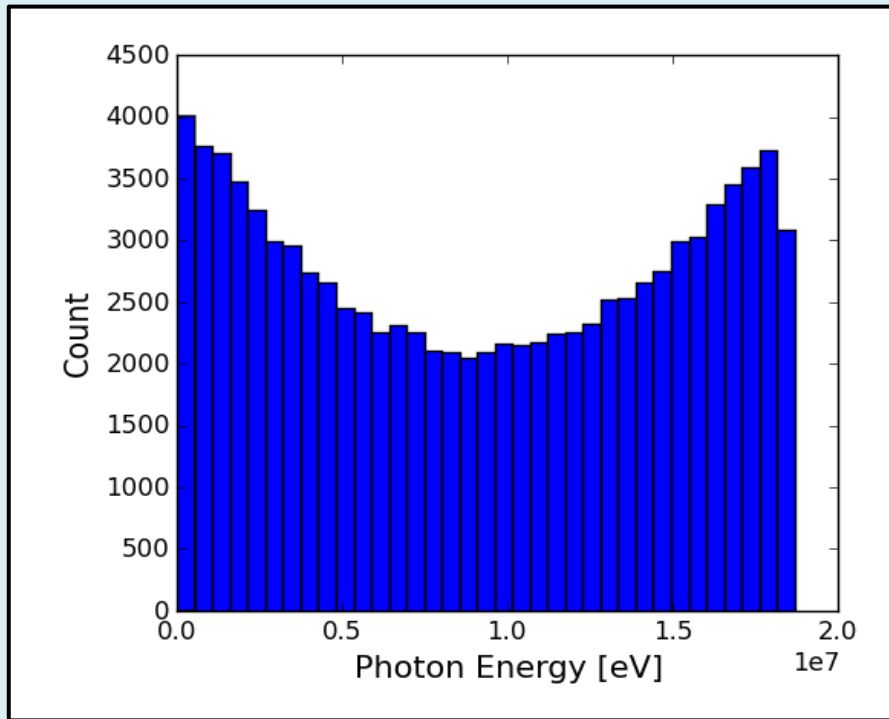
Let consider an electron-bunch with ultra low $\Delta\gamma/\gamma$ of 0.00003 and same condition as before



$E_{max} = 71.340keV, X=0.00116$

STAR Linac, X-ray source @ 1 GeV –not negligible recoil

Let consider an electron-bunch with an $\Delta\gamma/\gamma$ of 0.0005, scattering $\lambda=1\mu\text{m}$, 1 J, $W_0=20\mu\text{m}$ laser



$$E_{max} = \frac{4\gamma^2 h\nu}{1+X}, \quad X = \frac{4\gamma h\nu}{511\text{keV}} = \frac{8.4 \cdot 10^8 \cdot \text{Laser } E_{ph} [\text{eV}] \cdot W_0 [\mu\text{m}]}{\text{Phot } E_{en} [\text{eV}] \cdot (\sigma_x^2 [\mu\text{m}] + 0.25 \cdot W_0 [\mu\text{m}])}, \quad E_{max} = 19.46\text{MeV}, \quad X=0.019$$

$$E_{ph} (\lambda_{laser} = 1 \mu\text{m}) W_0 = \frac{1.24 \sqrt{P_L} \cdot \lambda_L / \pi}{\sqrt{R_L}}$$

$$E_{max} = \frac{4\gamma^2 h\nu}{1} \sim 20 \text{ MeV}$$

Suggested readings 1/2

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R.L. Sheffield et al., Proc. 1988 Linear Accelerator Conf., Williamsburg, VA, Oct. 1988, CEBAF rep. 89-001 (1989), p.520

C. Travier, Particle Accelerators 36 (1991), p.33

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J.B.Rosenzweig and L.Serafini, Phys. Rev. E-**49** (1994), p.1599

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W.K.H. Panofsky and W.A. Wenzel, Rev. Sci. Instr. 27 (1956), p.967

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J.B. Rosenzweig and E. Colby, AIP CP 335 (1995), p.724

L.Serafini, Particle Accelerators 49 (1995), p.253

L. Serafini et al., NIM A387 (1997), p.305

L.Serafini and J.B.Rosenzweig, Phys. Rev. E-55 (1997), p.7565

Proceedings of the ICFA 1999 Workshop on The Physics of High Brightness Beams, Los Angeles, 1999, Published on World Sci. ISBN 981-02-4422-3, June 2000

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S. G. Anderson and J. B. Rosenzweig, PRSTAB 3 (2000), p. 094201-1

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INFN-SPARC Project Web Site <http://pcfasci.fisica.unimi.it/Homepage.html>

End