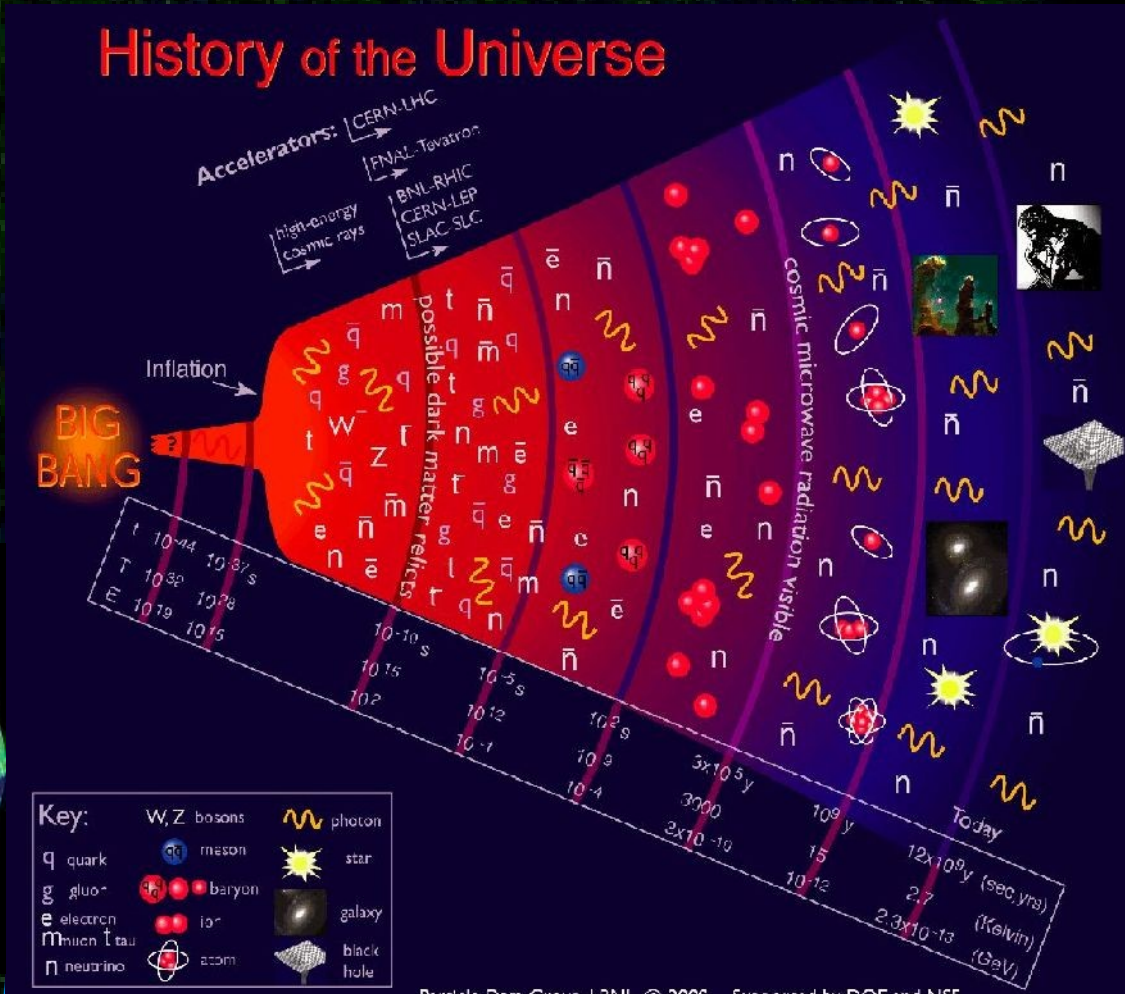
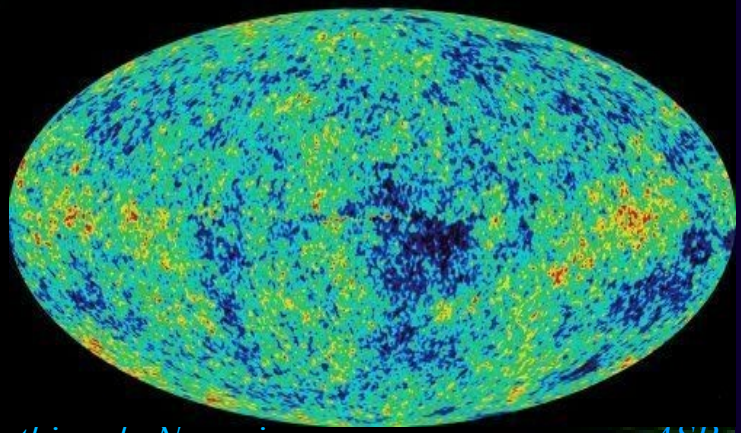
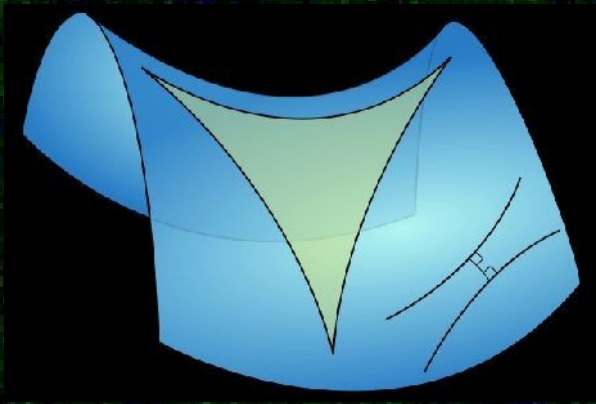


Cosmology – Lecture I – Theory

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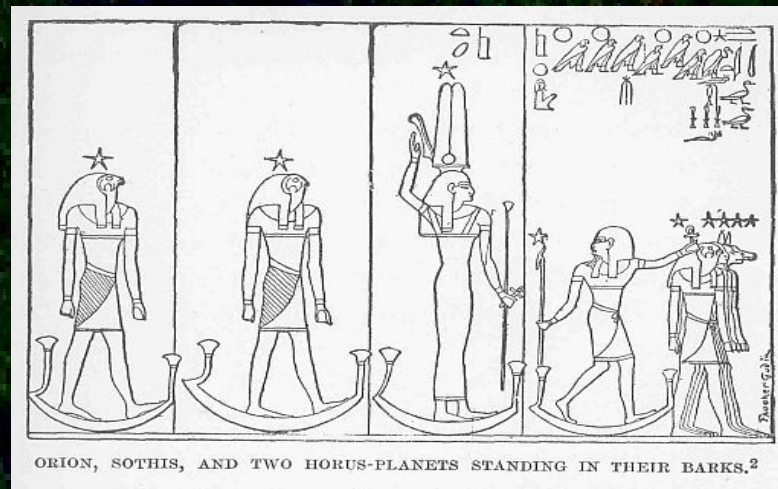
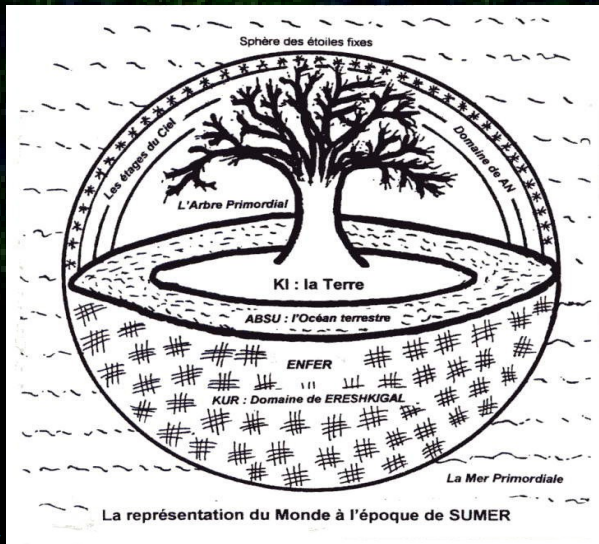




General Introduction Theory

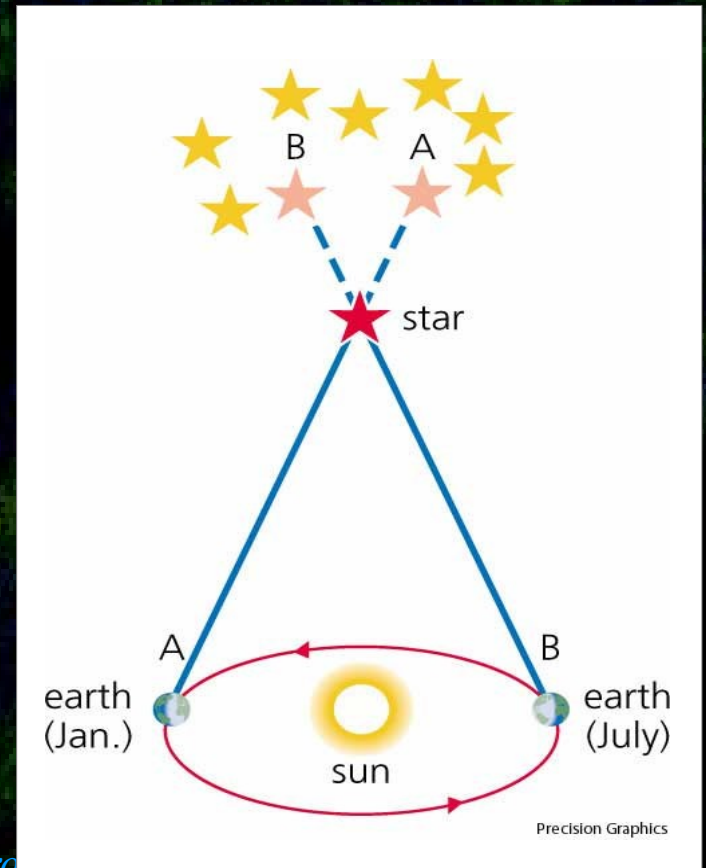
What is Cosmology?

- ❑ Fundamental questions about the origin and destiny of the Universe:
 - ❑ What is the Universe made up of ?
 - ❑ How did the matter and structures form in the Universe ?
 - ❑ Why is the Universe as we see it ?
 - ❑ What is our place in the Universe ?
 - ❑ Did the Universe always exist, and if not, what is its age ?
 - ❑ How will the Universe evolve / possibly end ?
- ❑ Questions that appear in all cultures/religions
- ❑ Many different answers across history



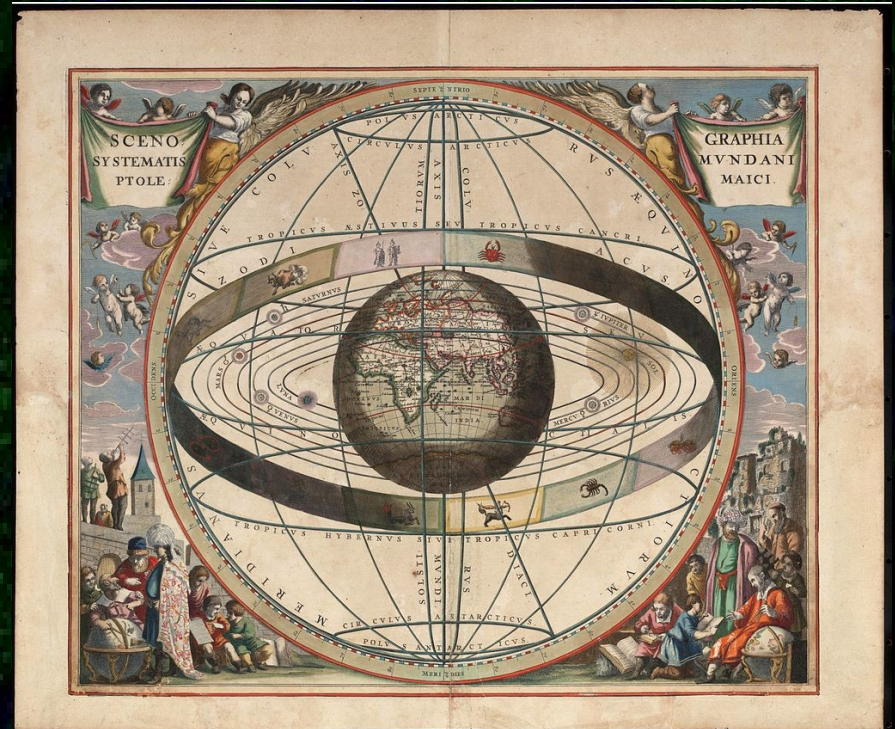
Historical Cosmology

- Movement of the planets & stars:
 - During one night
 - From one night to the other: movement of planets
 - From one year to the other: apparent movement of stars
 - From different places on the earth
- Try to explain all movement in the sky with a consistent model



Geocentric Model

- ❑ Ancient Greece (Aristotle)
- ❑ Earth at centre, fixed stars
- ❑ “Perfect” orbits : circles around Earth
- ❑ 10 embedded spheres, last one for stars
- ❑ Was not able to describe the “retrograde motion of planets”



Retrograde motion of planets

- ❑ Caused by differential orbital velocity, innermost planets orbiting fastest
- ❑ Was important events in astrology!
 - ❑ Avoid signature of contracts during retrograde motion of Mercury
 - ❑ Do not buy metallic object during retrograde motion of Mars



Mercury in "retrograde"

S Sun

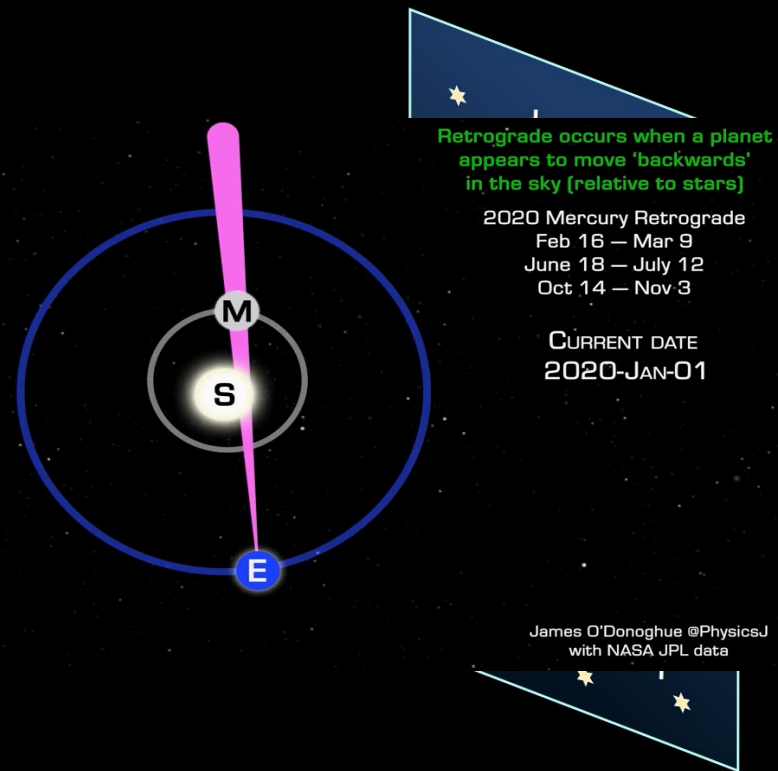
M Mercury

E Earth

Path of Mercury
in Earth's sky



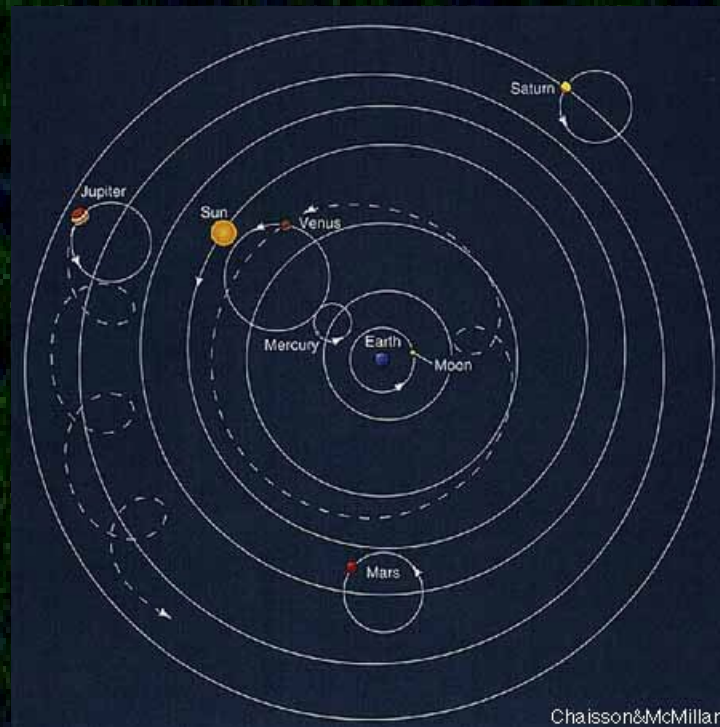
A retrograde cannot
affect human affairs



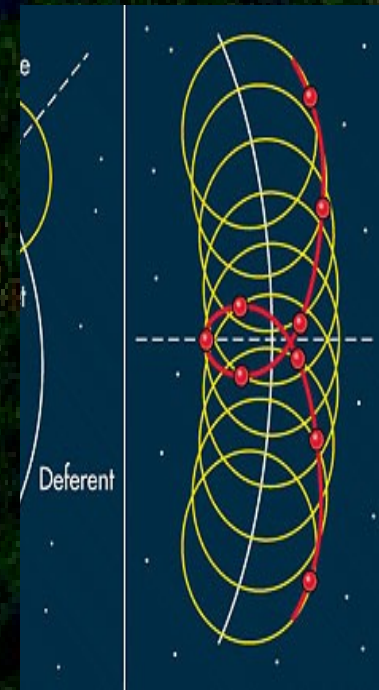
Theory of epi-cycles

- ❑ Initiated by Hipparchus (190 – 120 BC)
- ❑ Complicated movement of planets explained by epi-cycles:
 - ❑ Planets orbit on a circle, whose centre orbits the Earth
 - ❑ Able to describe this retrograde motion

Schema huius præmissæ diuisionis Sphærarum .

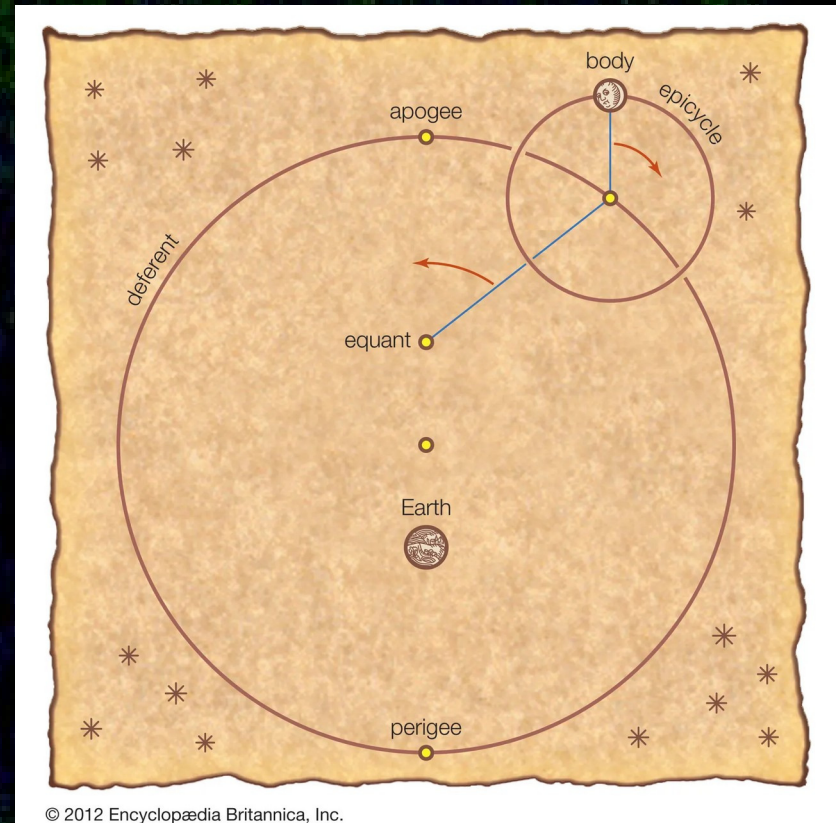


Chaisson&McMillan



Model of Ptolemy

- ❑ Refinement of the theory of epi-cycles:
- ❑ Centre of rotation (deferent circle) shifted with respect to the Earth
 - ❑ Accounts for variation of distances (apogee and perigee)
 - ❑ But not for elliptical orbit
- ❑ Order of spheres from Earth outward is:
 - 1) Moon
 - 2) Mercury
 - 3) Venus
 - 4) Sun
 - 5) Mars
 - 6) Jupiter
 - 7) Saturn
 - 8) Fixed Stars

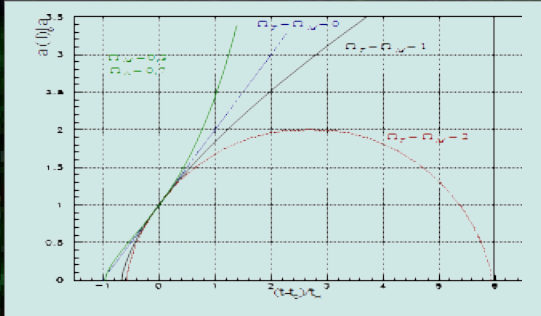


Major Steps in History

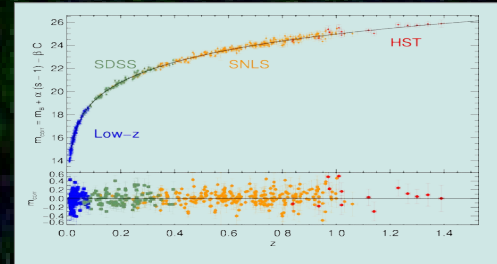
- ❑ -3000 : Flat earth, mythological Cosmology (Egypt, ...)
- ❑ ~100 : Earth at centre (Ptolemy)
- ❑ 1520 – 1680 : Sun at centre (Copernic, Newton)
- ❑ 1917 : Universe is infinite (Einstein)
- ❑ 1922 : Evolving Universe (Friedman – Lemaître)
- ❑ 1964 : Discovery of Cosmological Background. Big Bang model (Penzias & Wilson)
- ❑ > 2000 : Accelerated expansion (Supernova Ia, ...), modern cosmology

Open questions, observables

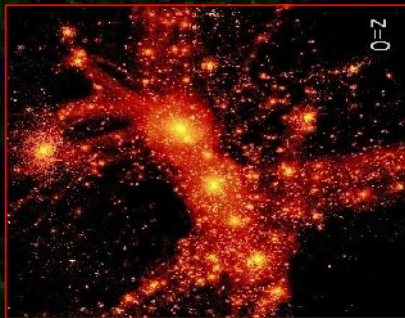
□ Evolution of the Universe



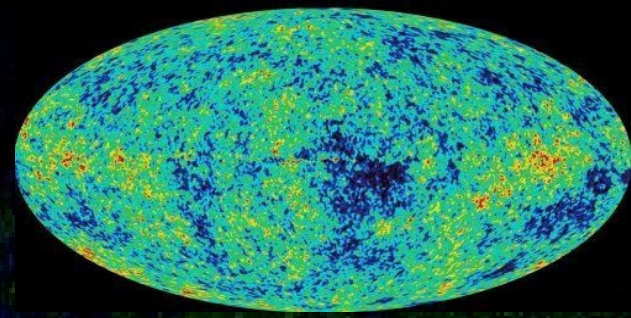
□ Supernova 1a: distance versus recession velocity



□ Formation of structures



□ Cosmological Background



□ Big bang Nucleo-synthesis

□ Abundances of light elements

□ ...

A Cosmic Microwave Background (CMB) fluctuation map, showing temperature variations across the sky. The map is a circular, textured pattern of green, blue, and yellow, representing the distribution of matter and energy in the early universe. The colors indicate temperature fluctuations, with warmer regions in yellow and cooler regions in blue.

Cosmology without General relativity (!)

Is a static Universe possible ?



- Take a Universe with many galaxies isotropically distributed
- Gravity force between each pair of galaxies is **attractive**
- Calculate the evolution in a **mean gravitational field**

Is a static Universe possible?

□ Consider only one Galaxy at distance $R(t)$

□ Forces:

□ Radial by symmetry

□ Isotropic pressure \rightarrow no net force

□ Radial force due to inner matter
(Gauss theorem)

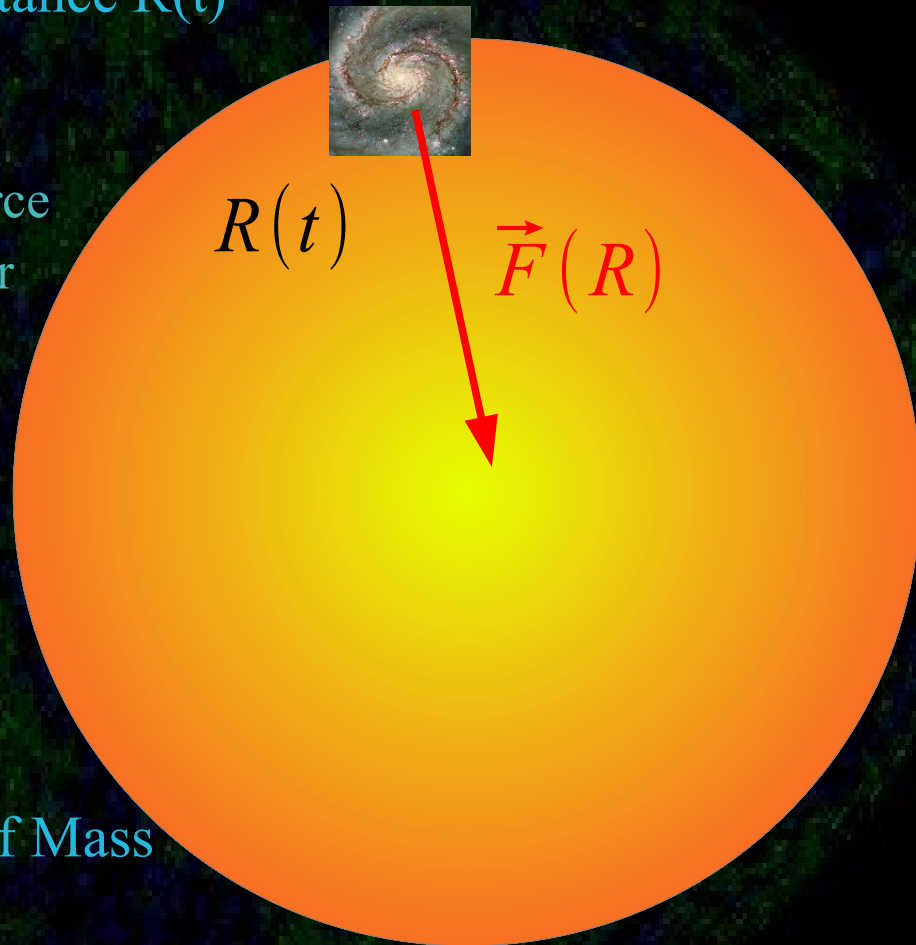
$$\vec{F}(R) = -\frac{G M(R) m}{R^2} \vec{u}_R$$

□ Evolution of a “bubble”:

$$\frac{d^2 R}{dt^2} = -\frac{G M(R)}{R^2}$$

□ Matter Universe, conservation of Mass

$$M(R) = \frac{4}{3} \rho_m(t) R^3 = C_{\text{ste}}$$



Evolution of a matter Universe

- Gravitational force

$$\vec{F}(R) = -\frac{GM(R)m}{R^2} \vec{u}_R$$

- Fundamental principle

$$\frac{d^2 R}{dt^2} = -\frac{GM(R)}{R^2}$$

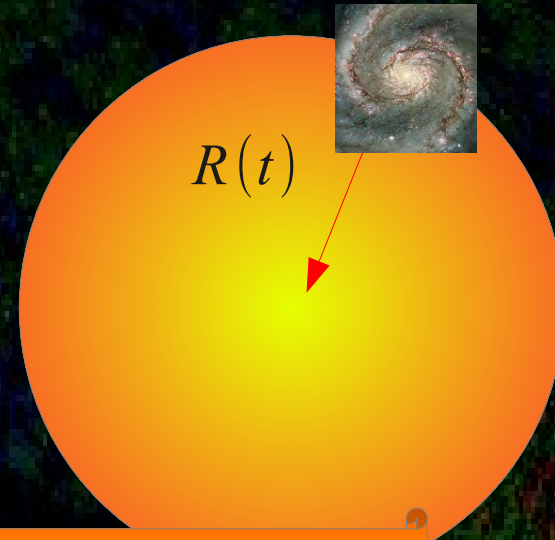
- Conservation of mass

$$M(R) = \frac{4}{3} \rho_m(t) R^3 = C_{\text{ste}}$$

- Evolution Equation:

$$\left(\frac{\ddot{R}}{R}\right) = -\frac{4\pi}{3} \rho_m G \Rightarrow \dot{R} \ddot{R} = -\frac{4\pi}{3} (\rho_m R^3) G \frac{\dot{R}}{R^2}$$

$$\Rightarrow \left(\frac{\dot{R}}{R}\right)^2 = \frac{8\pi}{3} (\rho_m R^3) \frac{G}{R^3} + \frac{C}{R^2}$$



Only for Matter!

Evolution of a matter Universe

- Evolution Equations:

Velocity

$$\left(\frac{\dot{R}}{R}\right)^2 = M_0 \frac{G}{R^3} + \frac{C}{R^2}$$

Acceleration

$$q = \left(\frac{\ddot{R}}{R}\right) = -\frac{8\pi G \rho_m}{3} \leq 0$$

- Expansion of the Universe is decelerated by matter content
- C is a constant specific to the Universe (Curvature! – see later)
- This does **NOT** require general relativity, pure classical mechanics!

**NO static matter
Universe is possible !!**

Interlude – Why no static Universe? Olber's paradox (1758-1840)

- ❑ Imagine a infinite, static Universe existing since ever.
- ❑ Isotropic distribution of Galaxies
- ❑ Light received by a galaxy at distance R scales as $1/R^2$
- ❑ Number of galaxies at distance $[R, R+dR]$ scales a $R^2 dR$
- ❑ Each slice contribute to \sim same value, integration leads to infinity



**The night sky must
be White!**



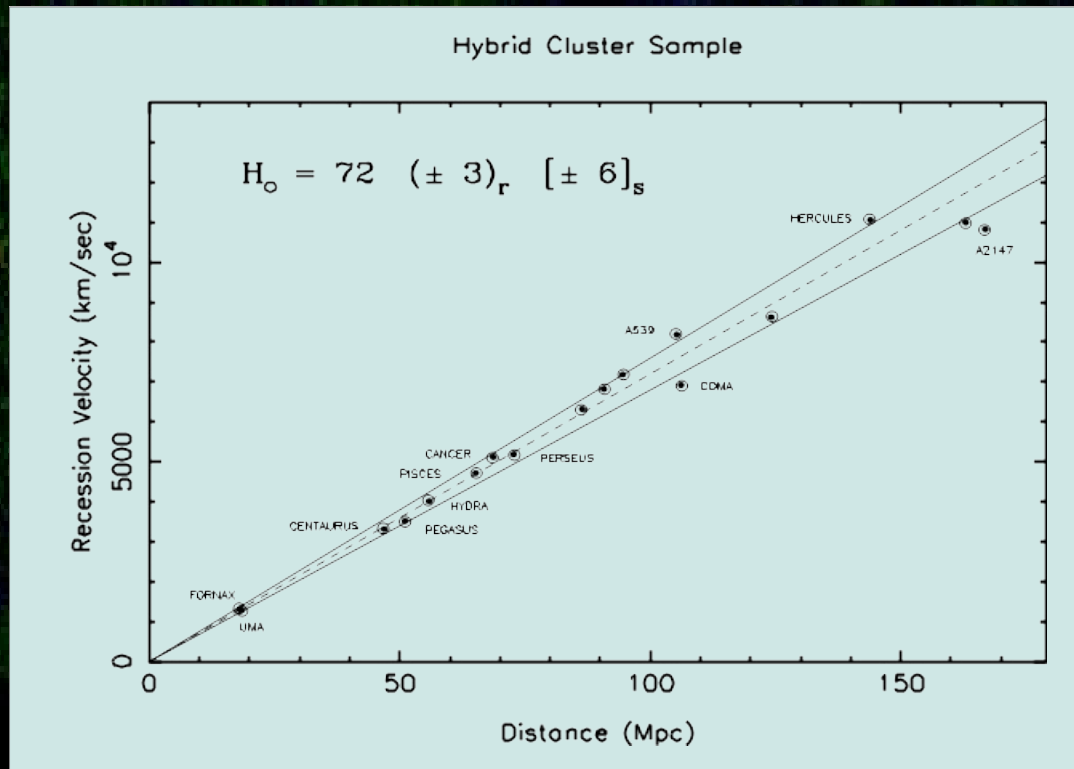
Observation – Hubble Law

- Galaxies are separating apart at a speed proportional to their distance

$$\frac{dR}{dt} = H_0 R + v_p \Rightarrow H_0 = \left\langle \frac{\dot{R}}{R} \right\rangle_{t=t_0}$$

Hubble flow

Proper Motion



Evolution of a matter Universe

- Rewriting the evolution equations with current value

$$\left(\frac{\dot{R}}{R}\right)^2 = \frac{8\pi}{3}(\rho_m R^3) \frac{G}{R^3} + \frac{C}{R^2} \Rightarrow H_0^2 = \frac{8\pi}{3}(\rho_m^0 G) + \frac{C}{R_0^2}$$

- Critical density $\rho_c = \frac{3 H_0^2}{8 \pi G}$, $\Omega_m = \frac{\rho}{\rho_c}$

Matter

(Curvature)

- Dimensionless evolution equation:

$$\frac{1}{H_0^2} \left(\frac{\dot{R}}{R}\right)^2 = \left(\Omega_m \left(\frac{R_0}{R}\right)^3 + (1 - \Omega_m) \left(\frac{R_0}{R}\right)^2\right)$$

- Slowdown of expansion driven by matter:

$$\left(\frac{\ddot{R}}{R}\right) = -\frac{4\pi G \rho_m}{3} = -\frac{\Omega_m}{2} H_0^2$$

Evolution of a matter Universe

- $\Omega_m = 0$, monotonic expansion

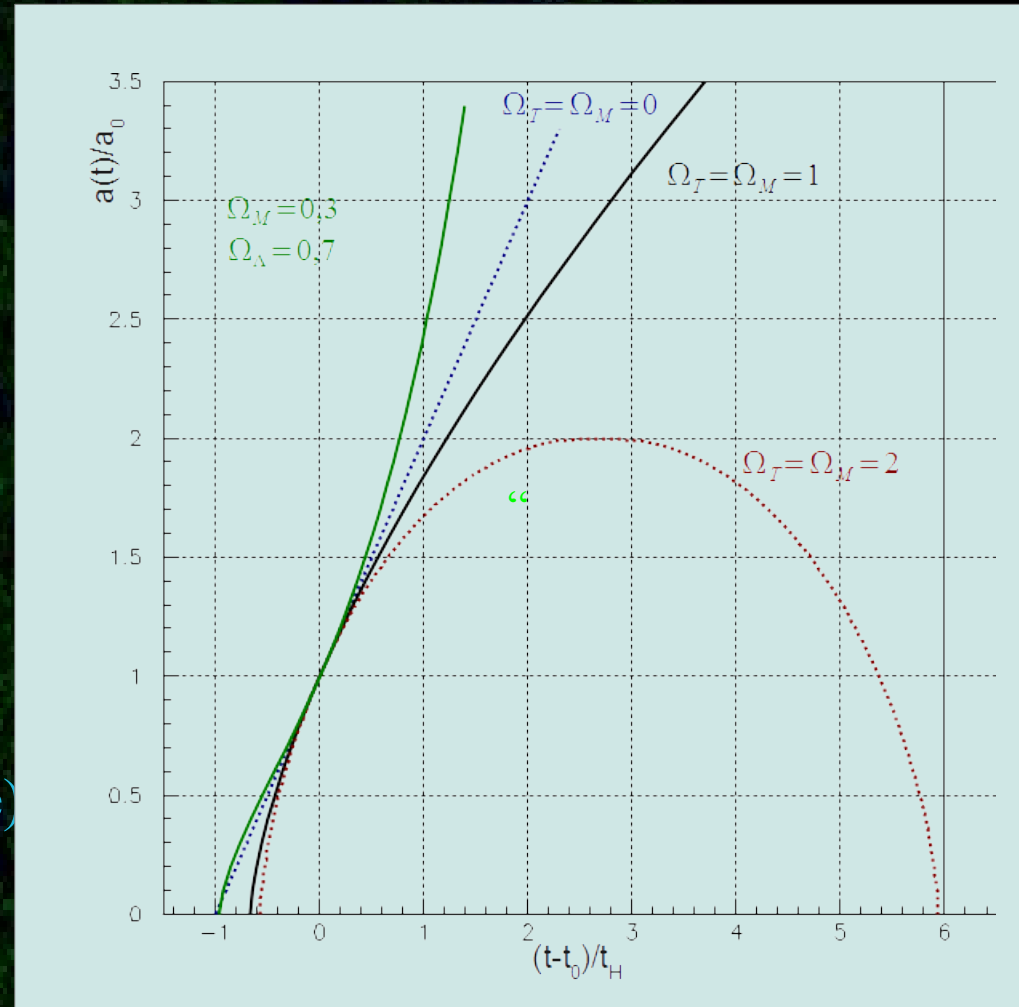
$$R(t) = R_0 H_0 \times t$$

- $\Omega_m = 1$ (critical Universe)
Decelerating expansion

$$R(t) = R_0 \left(\frac{3}{2} H_0 \times t \right)^{2/3}$$

- $\Omega_m > 1$ (over-critical Universe)
Collapsing Universe

$$R_{max} = R_0 \frac{\Omega_m}{(\Omega_m - 1)}$$



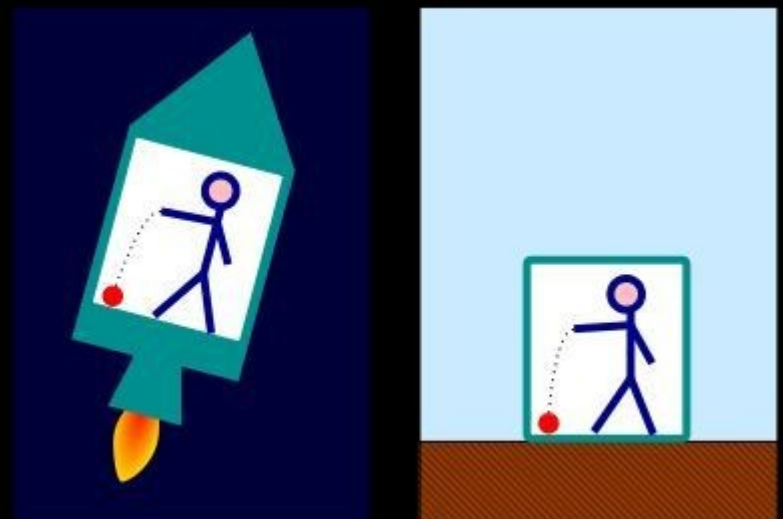
$$\frac{1}{H_0^2} \left(\frac{\dot{R}}{R} \right)^2 = \left(\Omega_m \left(\frac{R_0}{R} \right)^3 + (1 - \Omega_m) \left(\frac{R_0}{R} \right)^2 \right)$$



A (tiny)-bit of General relativity

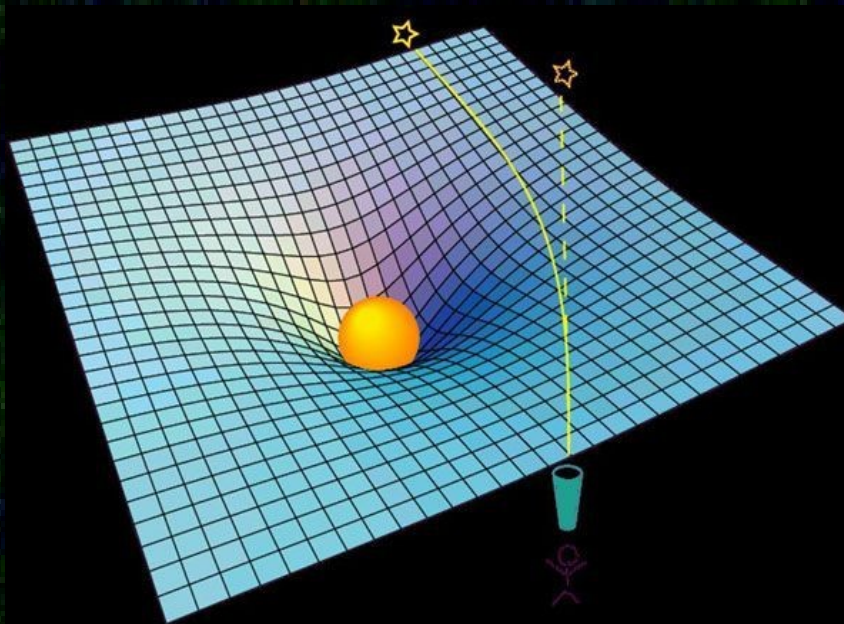
Equivalence Principle - A. Einstein

- No difference could be found between **inertial mass** (in acceleration) and **gravitational mass** (in gravity forces)
 - ⇒ Implies that acceleration of a body in a gravitational field is independent of the nature of the body
 - Tested extensively in vacuum tower
- Thus there is no way to distinguish between
 - free-fall movement in gravitational field
 - accelerated movement in absence of field
 - ⇒ Gravity can be understood as a property of space and not of the falling body



General Relativity vs Newtonian

- Newtonian Gravity: Universe is flat and immutable, trajectories are curved due to a force (non-inertial movement)
- General relativity: Gravity is a geometric property of space, not a force. Trajectories are always inertial (geodesics) in a curved space
- Major conclusion: massless particles (light) are also affected, confirmed by measure of deflection of stars (Eddington, 1919)



Evolving Universe – Tensor Algebra

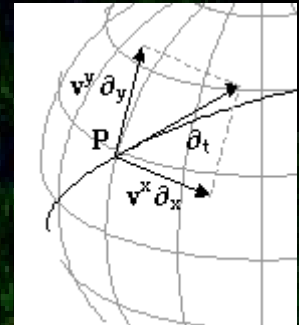
□ We consider a space time, in which we have a base of vectors $\{\vec{e}_\mu\}$

□ The **metric** is defined by the cross-product of vectors:

$$g_{\mu\nu} = \vec{e}_\mu \cdot \vec{e}_\nu$$

□ Any vector can be decomposed on the base: $\vec{x} = x^\mu \vec{e}_\mu$

Covariant coordinates



□ Several **bases** can describe the same Universe, transformation given by

$$d x^\mu = \frac{\partial x^\mu}{\partial y^\nu} d y^\nu = \Lambda^\mu_\nu d y^\nu, \quad \vec{e}_\mu = \Lambda_\mu^\nu \vec{f}_\nu$$

□ **Tensors** are objects of higher rank (2, 3, ...) which transform in a similar manner

$$T^{\mu\nu} = \Lambda^\mu_\alpha \Lambda^\nu_\beta T'^{\alpha\beta}$$

Norm & Invariants

- Scalars are invariant by change of coordinate, for instance:

$$A = U^\mu \cdot V_\mu = g^{\mu\nu} U_\mu V_\nu$$

- The elementary distance, defining the metric, can be expressed as:

$$ds^2 = dx^\mu \cdot dx_\mu = g^{\mu\nu} dx_\mu dx_\nu$$

Units where $c = 1$!

And is invariant by coordinate changes (such as the scalar product)

Tensor Algebra is the recipe to ensure that equations are Lorentz invariant, i.e. that equivalence principle is satisfied.

Curved Universe

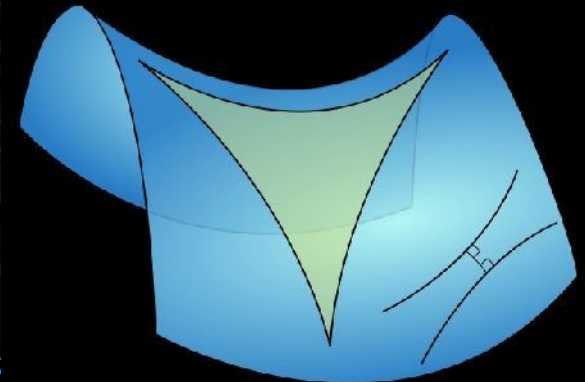
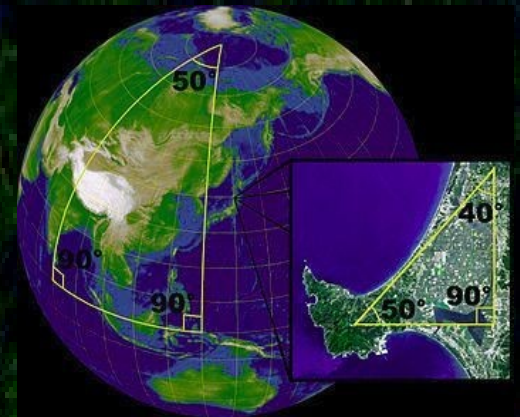
- In a flat Universe, the metric can be expressed in a diagonal form.
e.g. Minkowski space (flat space-time)

$$g_{\mu\nu} = \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix}$$

- This is not the case any more in curved Universe
- The “**curvature**” is a mathematical concept that is obtained from derivatives of the metric:

- Ricci tensor $R_{\mu\nu}$

- Scalar curvature $R = g^{\mu\nu} R_{\mu\nu}$

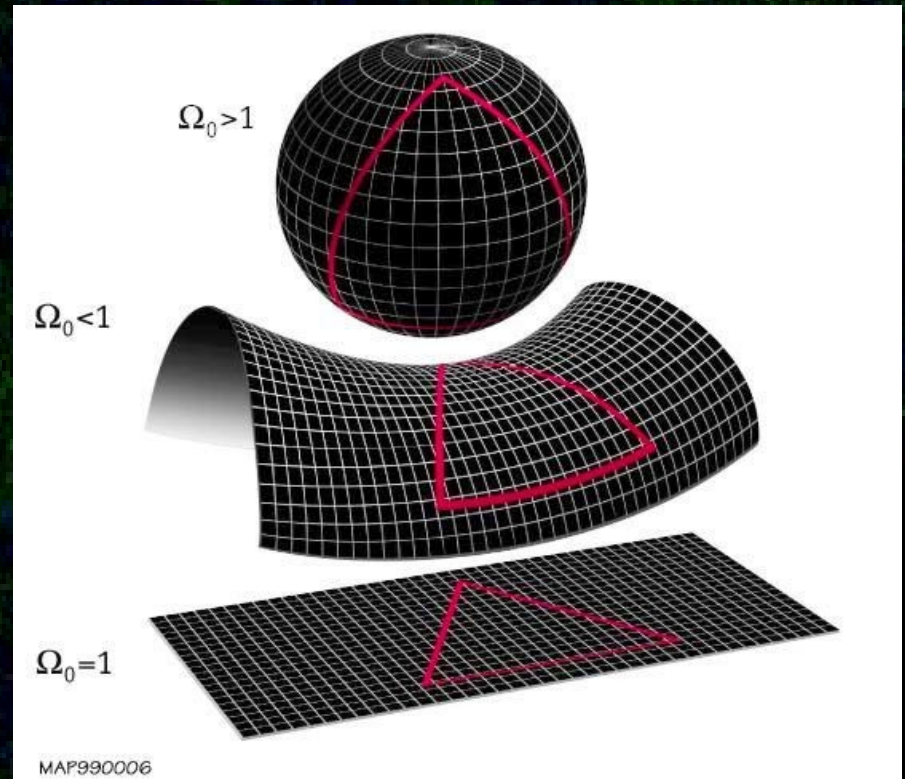


Uniform, Isotropic Universe

- A uniform, isotropic universe can be described by the Friedman-Lemaitre-Robertson-Walker metric

$$d s^2 \equiv d x^\mu d x_\mu = d t^2 - a^2(t) \left[\frac{d r^2}{1 - k r^2} + r^2 d \theta^2 + r^2 \sin^2 \theta d \phi^2 \right]$$

- $a(t)$ is a “**scale factor**” giving the size of a bubble of Universe
The **grid** itself is expanding, not the content!
- $k = 1$: Spherical space
(Sum of angles $> \pi$)
- $k = -1$: Hyperbolic space
(Sum of angles $< \pi$)
- $k = 0$: Euclidean space
(Sum of angles $= \pi$)



Einstein Equation – I

- General idea: find the minimum covariant formalism compatible with Newton gravity
- Start for the Poisson equation for gravitational potential

$$\text{Field} \quad \nabla^2 \Phi_p = -4 \pi \rho_g \quad \text{Matter Content}$$

- Construct a Lorentz-invariant (Covariant) version

$$\text{Covariant Derivative} \quad \left(\frac{\partial^2}{\partial t^2} - \nabla^2 \right) A^\mu = 4 \pi j^\mu \quad \text{Matter Quadri-current (Density is NOT Lorentz invariant)}$$

Energy Momentum Tensor

- Need a covariant (Lorentz invariant) formulation of energy conservation
- In special relativity Energy & Momentum are coupled

$$\frac{\partial \rho}{\partial t} + \text{div}(\rho \vec{v}) = 0 \quad \longrightarrow \quad \nabla_{\mu} T^{\mu}_{\nu} = 0$$

- Energy momentum tensor for a perfect fluid (Lorentz Invariant)

$$T_{\mu\nu} = n(\tilde{x}) \frac{p_{\mu} p_{\nu}}{E} = \rho u_{\mu} u_{\nu} + P(g_{\mu\nu} + u_{\mu} u_{\nu})$$

u_{μ} is the four velocity

- In the rest frame of fluid, $u^{\mu} = (1, 0, 0, 0)$ and thus:

$$T_{\mu\nu} = \begin{pmatrix} \rho(t) & & & \\ & -P(t) & & \\ & & -P(t) & \\ & & & -P(t) \end{pmatrix}$$

Energy Momentum Tensor

$$T_{\mu\nu} = \begin{array}{c} \text{Energy} \\ \text{Density} \end{array} \begin{array}{c} \text{Energy} \\ \text{Flux} \end{array} \begin{array}{c} T_{00} \\ T_{01} \quad T_{02} \quad T_{03} \\ T_{10} \\ T_{11} \quad T_{12} \quad T_{13} \\ T_{20} \\ T_{21} \quad T_{22} \quad T_{23} \\ T_{30} \\ T_{31} \quad T_{32} \quad T_{33} \end{array} \begin{array}{c} \text{Momentum} \\ \text{Density} \end{array} \begin{array}{c} \text{Momentum} \\ \text{Flux} \end{array} \begin{array}{c} \text{Viscosity} \\ \text{Pressure} \end{array}$$

The diagram illustrates the Energy-Momentum Tensor $T_{\mu\nu}$ as a 4x4 matrix. The components are organized into four quadrants:

- Top-Left (Energy Density):** A single component T_{00} .
- Top-Right (Energy Flux):** Three components T_{01} , T_{02} , and T_{03} .
- Bottom-Left (Momentum Density):** Three components T_{10} , T_{20} , and T_{30} .
- Bottom-Right (Momentum Flux):** A 3x3 sub-matrix of components T_{ij} for $i, j \in \{1, 2, 3\}$. This sub-matrix is shaded light green and has a diagonal line through it, representing the stress tensor.

Labels on the right side of the matrix indicate the physical quantities associated with the rows and columns of the bottom-right sub-matrix:

- Viscosity:** Associated with the diagonal elements T_{11} , T_{22} , and T_{33} .
- Pressure:** Associated with the off-diagonal elements T_{12} , T_{13} , T_{21} , T_{23} , T_{31} , and T_{32} .

Einstein Equation – II

- Minimum Covariant Equation

$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = 8\pi G T_{\mu\nu}$$

Curvature of Universe

Energy Content

- Energy Content:

$$T_{\mu\nu} = \sum_{\text{species}} (\rho u_{\mu} u_{\nu} + P (g_{\mu\nu} + u_{\mu} u_{\nu}))$$

- One can add a **Cosmological Constant** to force a static universe (Compensates for matter), no classical equivalent

$$G_{\mu\nu} + \Lambda g_{\mu\nu} = 8\pi G T_{\mu\nu}$$

General relativity in Friedman-Lemaitre-Robertson-Walker metric

- Einstein Equation (Isotropic Uniform Universe)

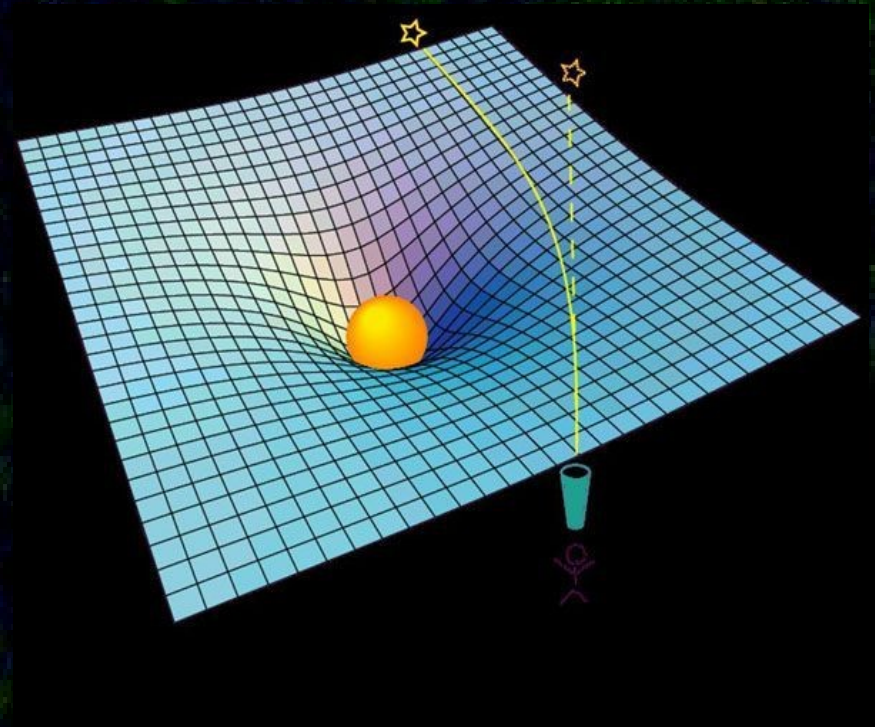
$$H^2 = \left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3} \sum_i \rho_i - \frac{k}{a^2}$$

- Acceleration

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} \sum_i (\rho_i + 3 p_i)$$

- So we still need:

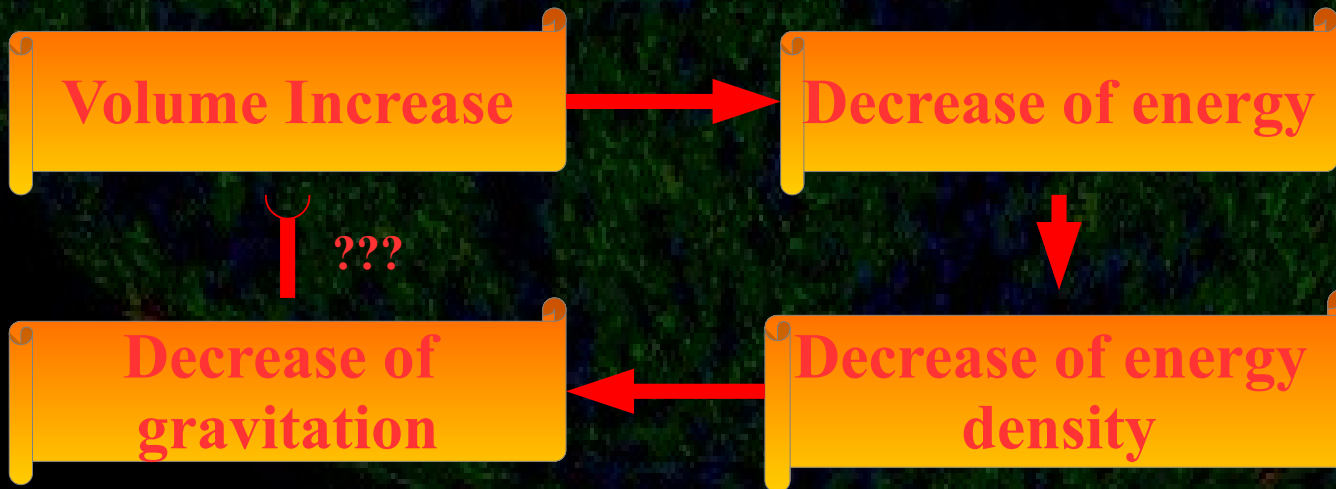
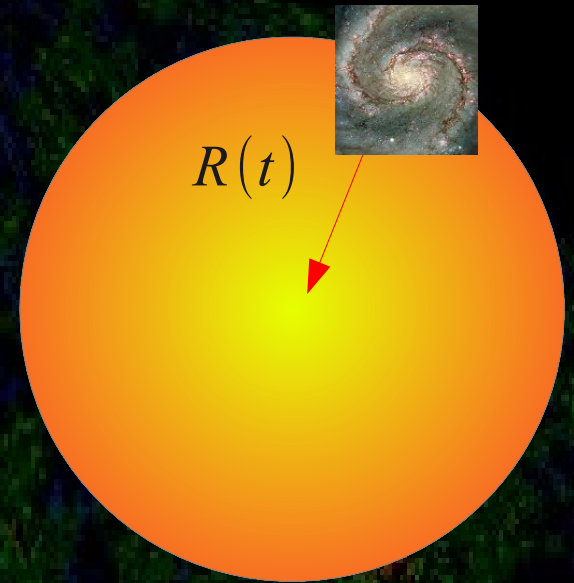
- Relation between pressure & density (equation of state)
- Corresponding evolution of density with time



Why pressure?

- Gravitation depends on energy content
- But what if the size of the Universe changes?
- Thermodynamics never lies and says:

$$dE = \delta W = -p dV$$



Thermodynamics – Evolution of density

□ Work of pressure:

$$dE = \delta W = -p dV$$

□ Expression of energy:

$$E = \rho V$$

$$\frac{dE}{dt} = \rho \frac{dV}{dt} + V \frac{d\rho}{dt} = -p \frac{dV}{dt}$$

□ Evolution of density:

$$\frac{d\rho}{dt} = -(p + \rho) \frac{1}{V} \frac{dV}{dt} = -3 \frac{\dot{a}}{a} (p + \rho)$$

□ In particular, constant density:

$$\frac{d\rho}{dt} = 0 \Leftrightarrow p = -\rho$$

Negative pressure?

□ Using equation of state:

$$P = w\rho \Rightarrow \rho(t) = \rho_0 \left(\frac{a}{a_0} \right)^{-3(1+w)}$$

Equation of state – Matter (cold)

- Normal matter:
 - Energy Density

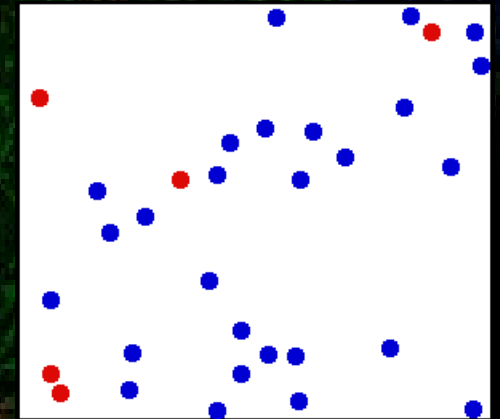
$$\frac{E}{V} = \rho_m \left(c^2 + \frac{1}{2} v^2 \right) \approx \rho_m c^2$$

- Pressure is related to kinetic energy (internal energy)

$$P = \frac{n R T}{V} = \frac{2}{3} \frac{\langle E_c \rangle}{V} \approx \frac{2}{3} \frac{\langle v^2 \rangle}{c^2} \times \frac{E}{V} \ll \frac{E}{V}$$

- For normal matter kinetic energy is negligible compared to mass energy

$$P = 0 = w \rho \quad \text{with} \quad w = 0$$



Equation of state – Radiation

□ Radiation:

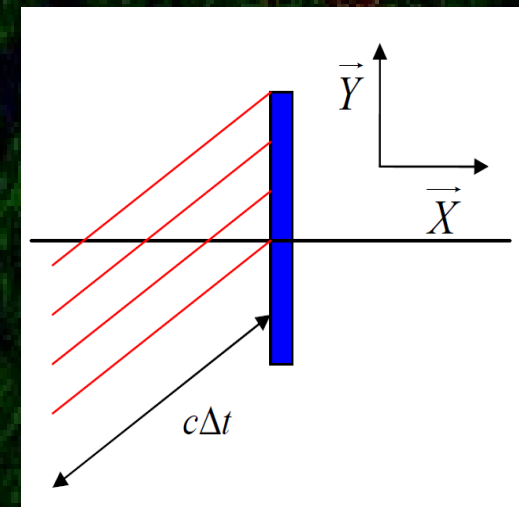
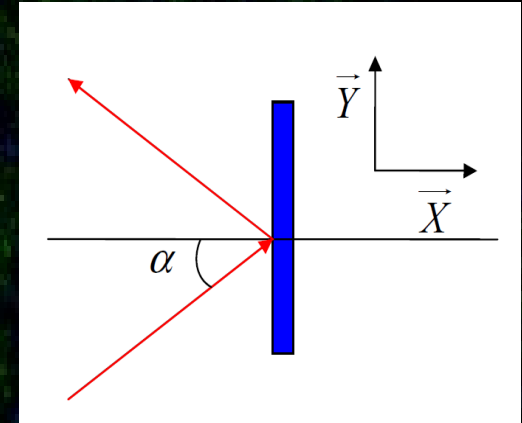
□ Energy Density

$$\frac{E}{V} = \frac{N}{V} \times pc$$

□ Simple calculation (reflection of photons with momentum transfer) shows

$$P = \frac{N}{V} \times pc \int \cos^2 \alpha d \cos \alpha$$
$$= \frac{1}{3} \frac{E}{V}$$

$$P = w \rho \quad \text{with} \quad w = \frac{1}{3}$$



Equation of state – Cosmological constant

- Cosmological constant is characterized by constant density

$$\rho = \text{constant}$$

- Thus

$$\frac{d\rho}{dt} = -3 \frac{\dot{a}}{a} (p + \rho) = 0$$

- This implies

$$P = -\rho = w \rho \quad \text{with} \quad w = -1$$

- Strange fluid with **negative pressure!**
⇒ Volume increase lead to **energy increase**

Cosmological Constant

- ❑ Introduced by Einstein to allow for a **static** Universe (counteracting the mass)
- ❑ **Positive** energy density, independent of size, implying **negative pressure**
- ❑ Kind of “vacuum energy”
- ❑ But in 1929 Edwin Hubble showed that the Universe is in expansion

Much later, when I was discussing cosmological problems with Einstein, he remarked that the introduction of the cosmological term was the biggest blunder of his life.

-- George Gamow, *My World Line*, 1970

General relativity in Friedman-Lemaitre-Robertson-Walker metric

- Einstein Equation (Isotropic Uniform Universe)

$$H^2 = \left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3} \sum_i \rho_i - \frac{k}{a^2}$$

- Acceleration

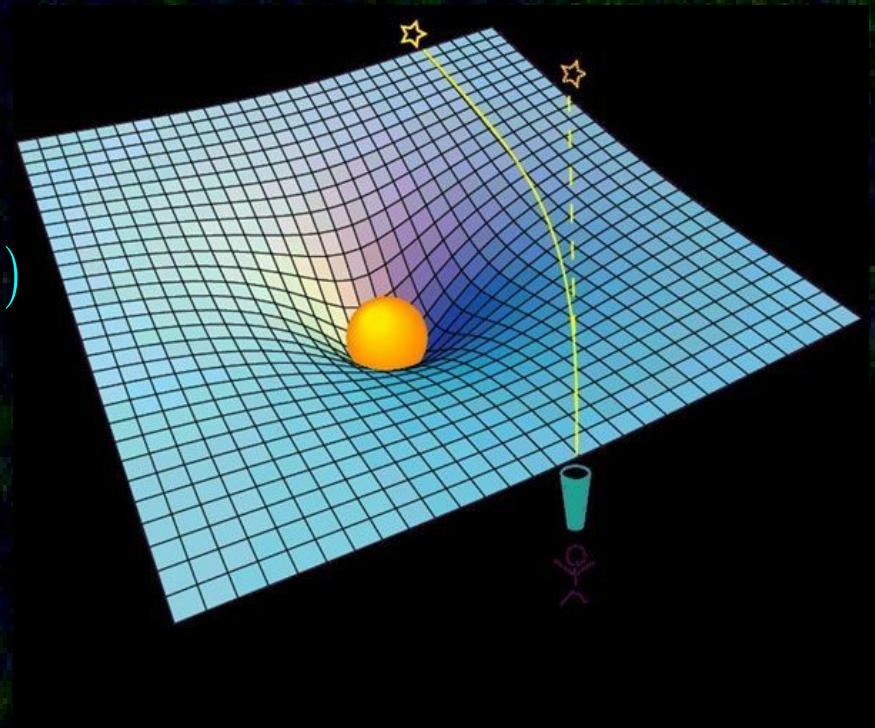
$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} \sum_i (\rho_i + 3p_i)$$

- Evolution of density

$$\frac{d\rho}{dt} = -(\rho + p) \frac{1}{V} \frac{dV}{dt} = -3 \frac{\dot{a}}{a} (\rho + p)$$

- Equation of state

$$P = w\rho \quad \Rightarrow \quad \rho(t) = \rho_0 \left(\frac{a}{a_0}\right)^{-3(1+w)}$$



Matter, radiation, ...

Content	State Equation	w	Dilution Law	Evolution
Cold Matter	$p \approx 0$	0	$\rho \propto a(t)^{-3}$	$a(t) \propto t^{2/3}$
Hot Radiation	$p = \frac{\rho}{3}$	1/3	$\rho \propto a(t)^{-4}$	$a(t) \propto t^{1/2}$
Curvature			$\left(\frac{\dot{a}}{a}\right)^2 = -\frac{k}{a^2}$	$a(t) \propto t$
Cosmological constant	$p = -\rho$	-1	$\rho = C_{ste} = \frac{\Lambda}{8\pi G_N}$	$a(t) \propto e^{H \times t}$
Generic	$p = w\rho$	w	$\rho \propto a(t)^{-3(1+w)}$	$a(t) \propto t^{1/3(1+w)}$

Evolution of the Universe

$$\left(\frac{H}{H_0}\right)^2 = \Omega_m^0 \left(\frac{a_0}{a}\right)^3 + \Omega_r^0 \left(\frac{a_0}{a}\right)^4 + \Omega_\Lambda + (1 - \Omega_{tot}^0) \left(\frac{a_0}{a}\right)^2$$

- ❑ (Cold) Matter: $\Omega_m^0 \left(a_0/a\right)^3$
- ❑ (Hot) Radiation: $\Omega_r^0 \left(a_0/a\right)^4$ Dominates in the early Universe
- ❑ Curvature: $(1 - \Omega_{tot}^0) \left(a_0/a\right)^2$
- ❑ Cosmological Constant: Ω_Λ
- ❑ Were we used dimensionless densities: $\Omega_i^0 = \frac{\rho_i^0}{\rho_c} = \frac{8\pi G}{3H_0^2} \rho_i^0, \quad \Omega_k^0 = \frac{-k}{a_0^2 H_0^2}$

Deceleration parameter

- Deceleration parameter

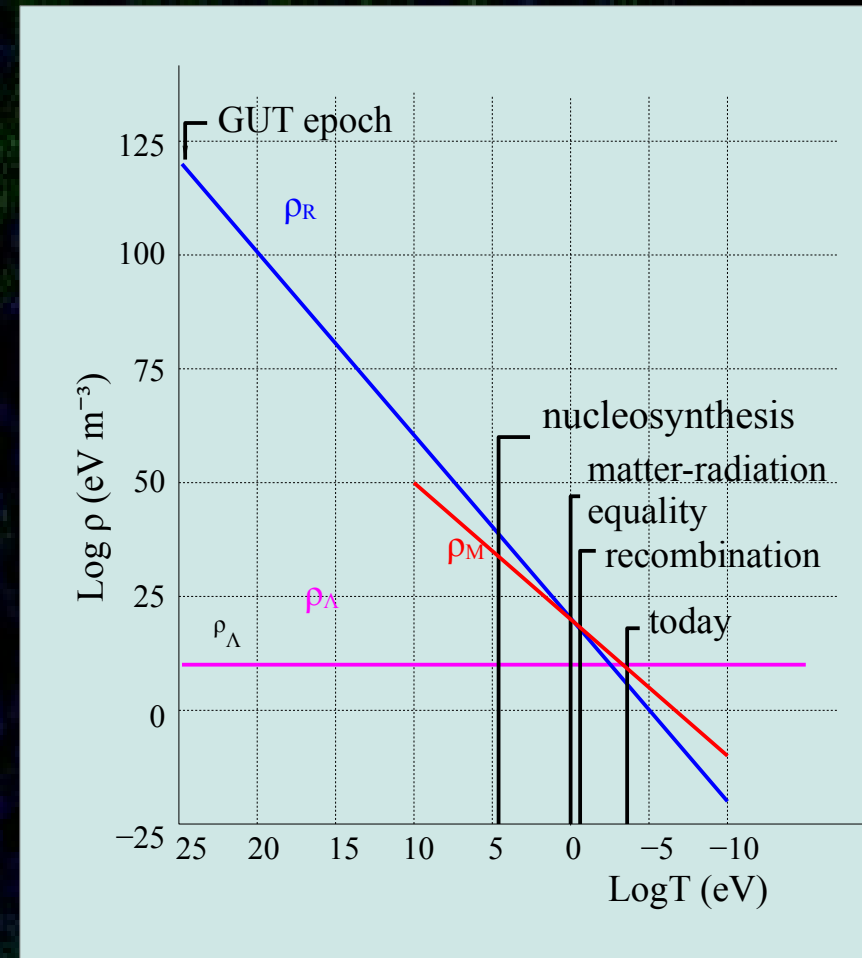
$$q = -\frac{1}{H^2} \left[\frac{\ddot{a}}{a} \right] = \frac{\Omega_m}{2} + \Omega_r - \Omega_\Lambda$$

- Matter and radiation **decelerates** expansion
- Cosmological constants **accelerates** expansion
- Curvature is **neutral**
- Null deceleration (**constant expansion rate** universe) if

$$\Omega_m + 2\Omega_r = 2\Omega_\Lambda$$

Epochs and Fate

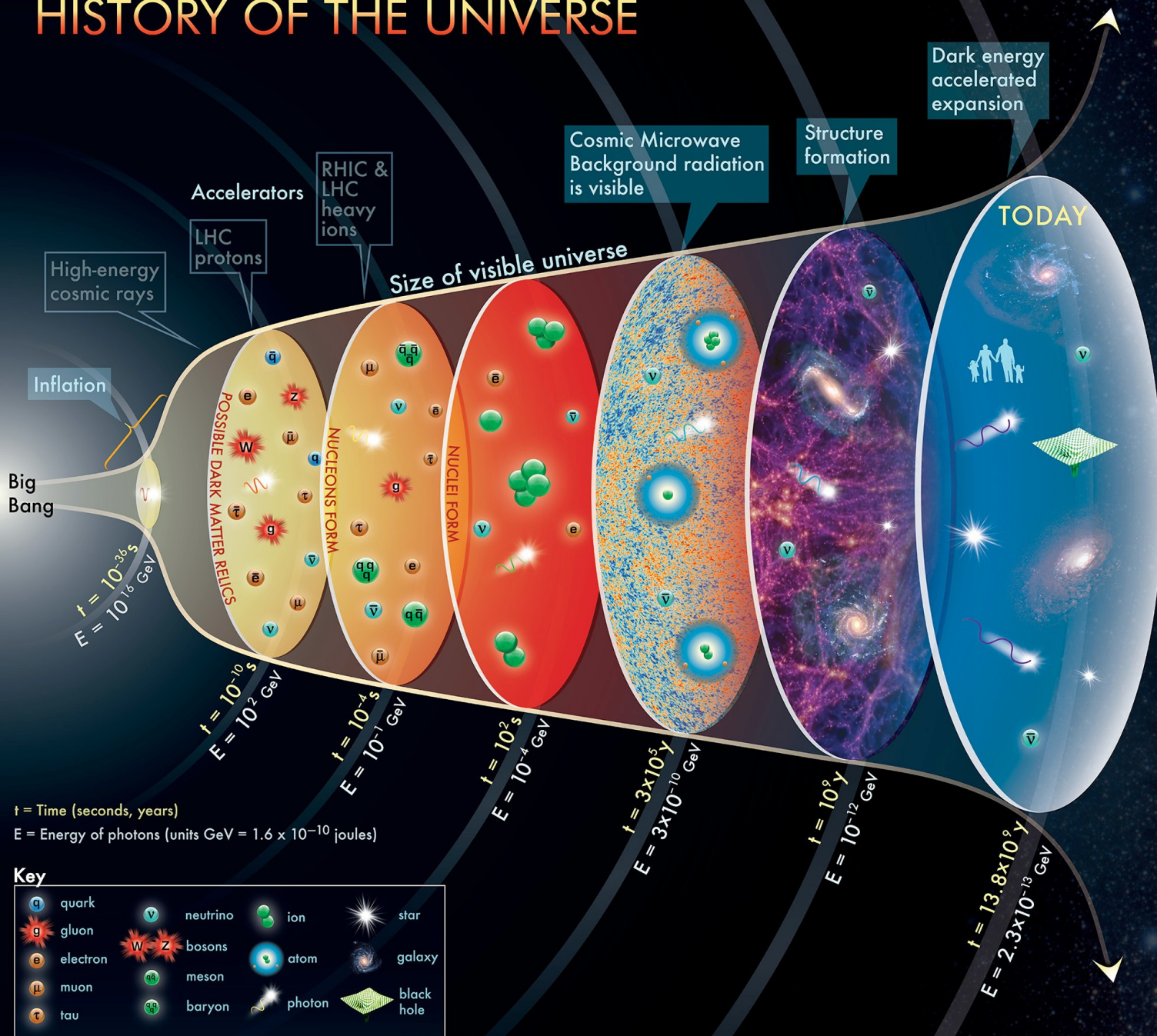
- Universe starts by a **radiation** dominated era
- After some times, **matters** dominates over the radiation and expansion slows down
- If $\Omega > 1$ and $\Omega_\Lambda \sim 0$, the Universe re-collapses and radiation dominates again
- If $\Omega < 1$ and $\Omega_\Lambda \sim 0$, the Universe ends in free expansion governed by curvature
- If $\Omega < 1$ and $\Omega_\Lambda > 0$, the Universe ends in accelerated exponential expansion governed by cosmological constant



Take-home message

- ❑ Evolution of Universe is given by
 - ❑ **General Relativity** (conceptual frame)
 - ❑ **Curvature** of the Universe (different Universes are possible, specificity of our Universe)
 - ❑ Energy content of the Universe, evolving with time
 - ❑ Matter
 - ❑ Radiation
 - ❑ Cosmological Constant – Dark Energy
- ❑ After 5 min break, Experimental status

HISTORY OF THE UNIVERSE



t = Time (seconds, years)
 E = Energy of photons (units GeV = 1.6×10^{-10} joules)

Key

quark	neutrino	ion	star
gluon	bosons	atom	galaxy
electron	meson	photon	black hole
muon	baryon		
tau			

The concept for the above figure originated in a 1986 paper by Michael Turner.



5 Min Break