

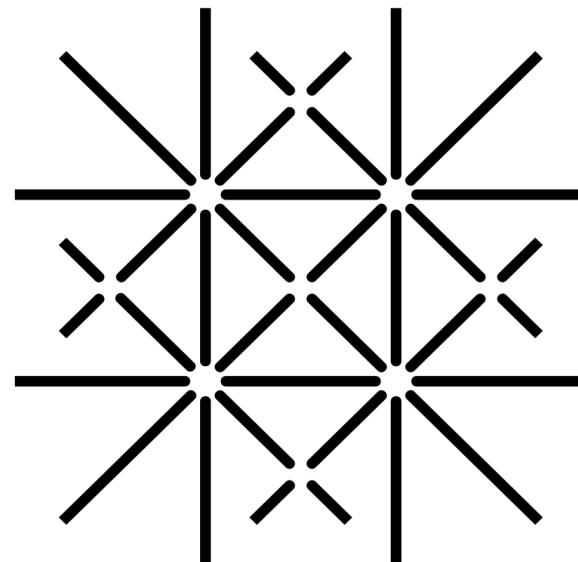
ALP phenomenology

(in concrete flavor models)

Alessandro Valenti

University of Basel

Based on [2407.02998](#) in collaboration with Admir Greljo, Aleks Smolkovic



Universität
Basel

Implications of LHCb measurements and future prospects

CERN

October 25th, 2024

Outline

1. What is an ALP?
2. Concrete model: Froggatt-Nielsen ALP
3. Phenomenology
4. Conclusion

1. What is an ALP?

ALP (axion-like particle):

spin-0 particle with an approximate shift symmetry $a \rightarrow a + \text{const}$

- Ubiquitous: remnant of spontaneously broken global symmetries
- Possible solution to Strong CP problem (QCD axion)
- Can account for the dark matter in our universe (misalignment)
- Can drive inflation
- Generic prediction of String Theory models ($N \gg 1$)
- ...

1. What is an ALP?

SM+ a EFT:

$$\mathcal{L}_{\text{SM}+a} = \mathcal{L}_{\text{SM}} + \frac{1}{2}(\partial a)^2$$

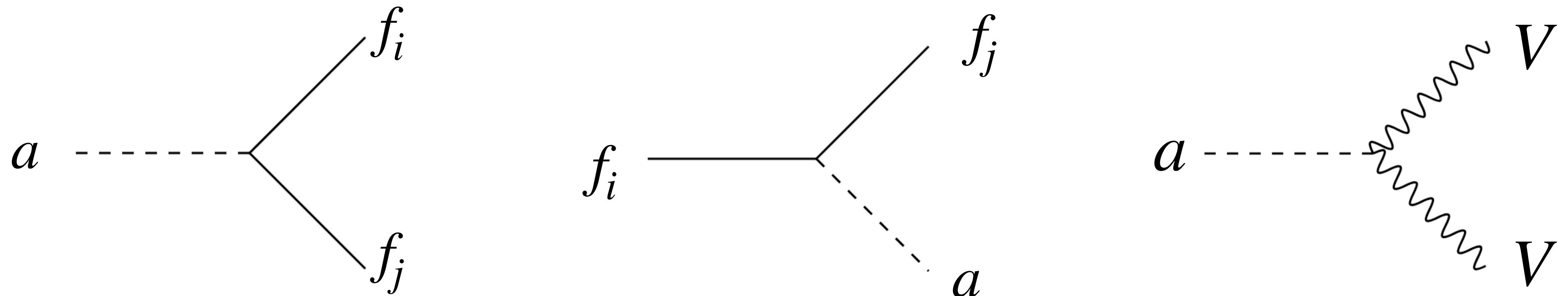
1. What is an ALP?

SM+ a EFT:

$$\mathcal{L}_{\text{SM}+a} = \mathcal{L}_{\text{SM}} + \frac{1}{2}(\partial a)^2 - \frac{1}{2}m_a^2 a^2$$

$-\frac{\partial_\mu a}{f_a} \sum_f c_{V(A),ij}^f \bar{f}_i \gamma^\mu (\gamma^5) f_j + \frac{a}{f_a} \frac{c_F}{32\pi^2} F \tilde{F} + \dots$

soft shift
symmetry breaking



2. Froggatt-Nielsen ALP

Very general. Focus on concrete popular flavor scenario.

Froggatt-Nielsen

- New discrete/continuous global symmetry+ SSB \longrightarrow **ALP** *

$f_a(\nu_\Phi)$ = scale of
new physics!

* If gauged no ALP \rightarrow see heavy vector bosons pheno
ALP phenomenology

2. Froggatt-Nielsen ALP

Very general. Focus on concrete popular flavor scenario.

Froggatt-Nielsen

- New discrete/continuous global symmetry+ SSB $\longrightarrow \text{ALP}^*$
- Predictive: c_{ij} fixed by SM flavor parameters $f_a(\nu_\Phi)$ = scale of new physics!

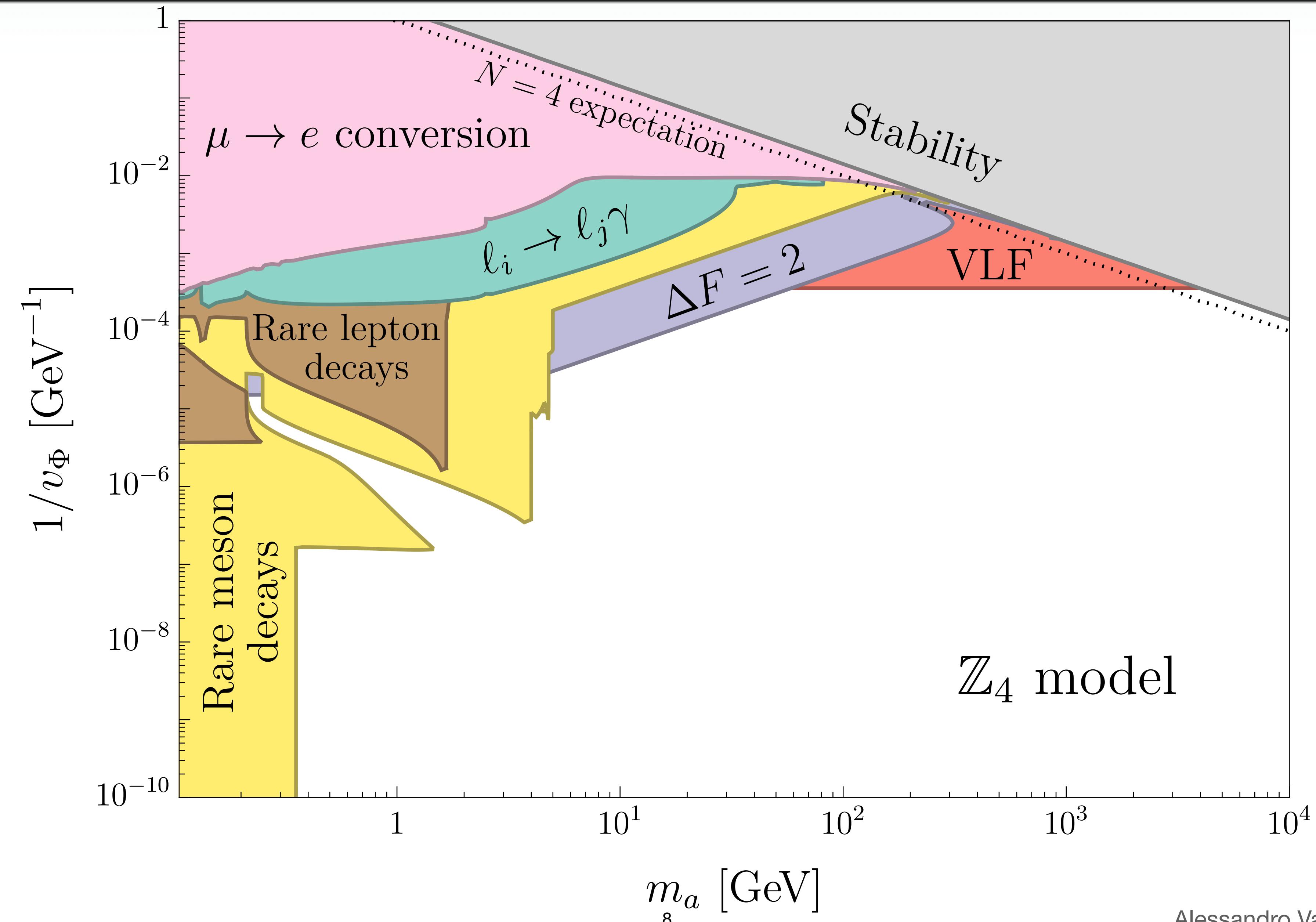
Example: Z_4 model

$$\mathcal{L} \supset -\frac{\partial_\mu a}{f_a} \sum_{f=q,u,d,l,e} c_{V(A),ij}^f \bar{f}_i \gamma^\mu (\gamma^5) f_j$$

$$c^d \sim \begin{pmatrix} 1 & \frac{m_d}{m_s} & \frac{m_d}{m_b} \\ \frac{m_d}{m_s} & 1 & \frac{m_s}{m_b} \\ \frac{m_d}{m_d} & \frac{m_s}{m_b} & \frac{m_s^2}{m_b^2} \end{pmatrix}, \quad c^u, c^e \sim c^d$$
$$\left(\frac{m_d}{m_s} \sim \lambda_c, \frac{m_s}{m_b} \sim \lambda_c^2, \frac{m_d}{m_b} \sim \lambda_c^3 \right)$$

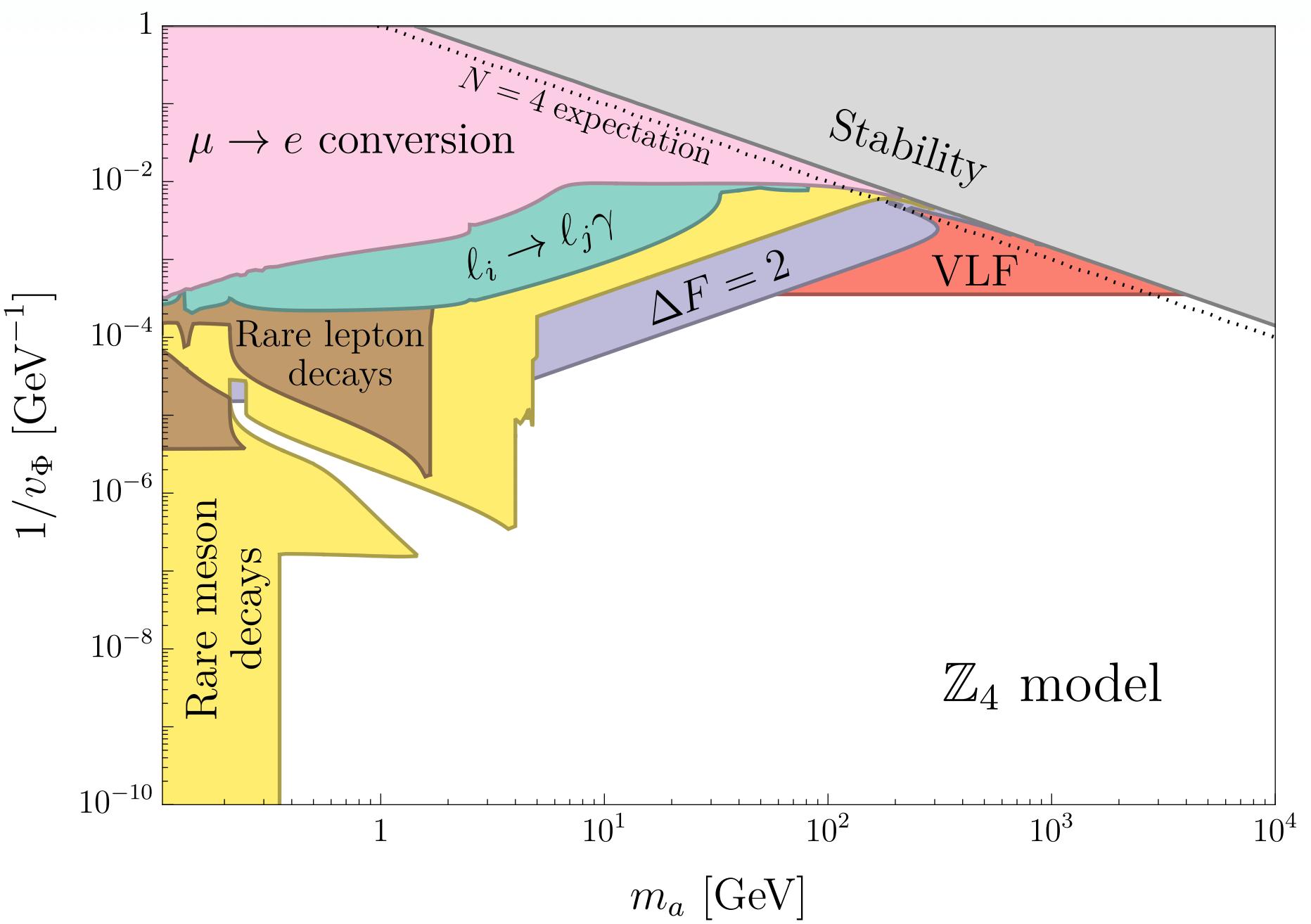
* If gauged no ALP \rightarrow see heavy vector bosons pheno
ALP phenomenology

3. ALP pheno



3. ALP pheno

MESON DECAYS



3. ALP pheno

MESON DECAYS

Invisible

$$\text{Br}(K^+ \rightarrow \pi^+ + a) < 9.5 \times 10^{-11}$$

@ E787+E949 2007

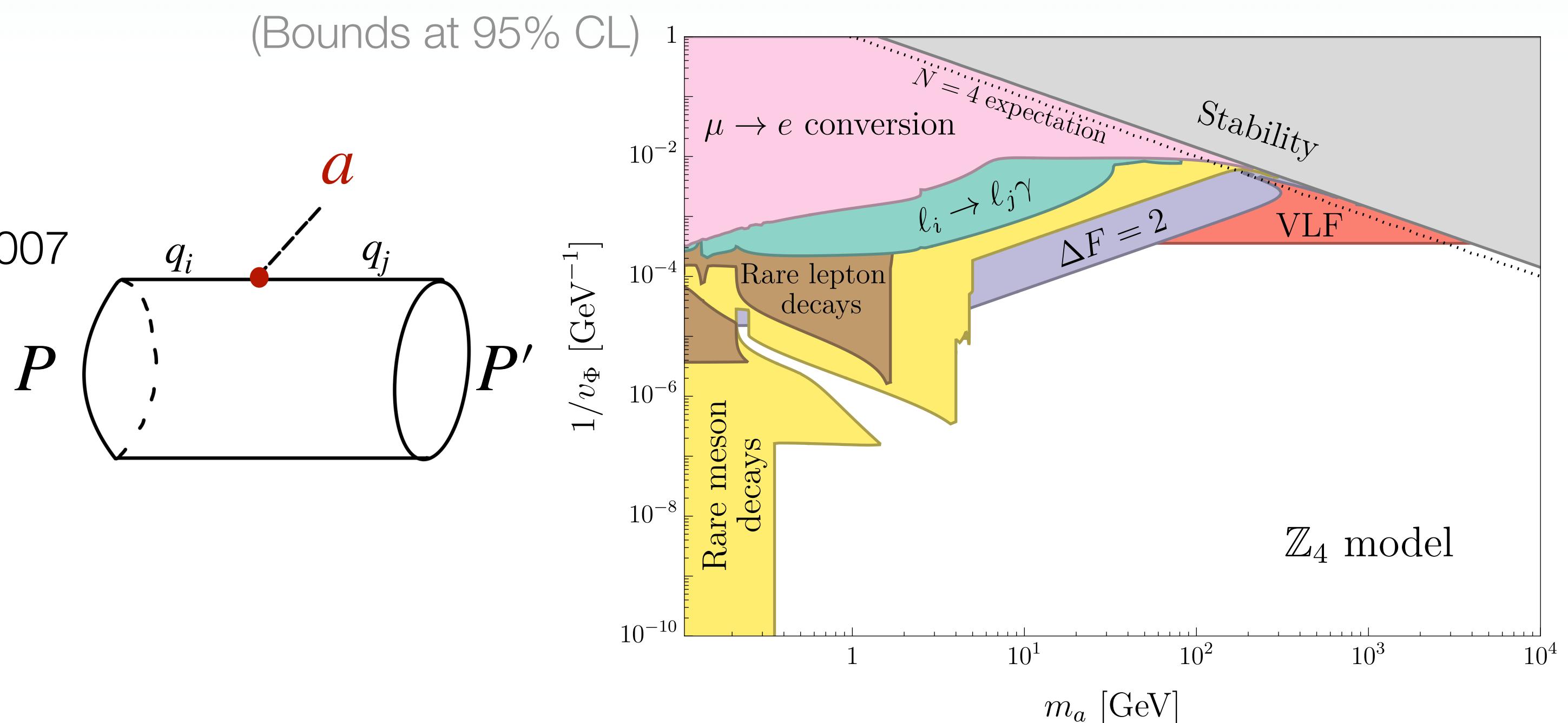
$$\text{Br}(B_0 \rightarrow K_0^* + a) < 1.0 \times 10^{-4}$$

@ BaBar 2013

$$\text{Br}(B^+ \rightarrow K^+ \bar{\nu}\nu) = (2.3 \pm 0.7) \times 10^{-5}$$

@ Belle II 2023

⋮



3. ALP pheno

MESON DECAYS

Invisible

$$\text{Br}(K^+ \rightarrow \pi^+ + a) < 9.5 \times 10^{-11}$$

@ E787+E949 2007

$$\text{Br}(B_0 \rightarrow K_0^* + a) < 1.0 \times 10^{-4}$$

@ BaBar 2013

$$\text{Br}(B^+ \rightarrow K^+ \bar{\nu}\nu) = (2.3 \pm 0.7) \times 10^{-5}$$

@ Belle II 2023

⋮

Visible - purely leptonic

$$\text{Br}(K_L \rightarrow \mu\mu) = (6.8 \pm 0.1) \times 10^{-11}$$

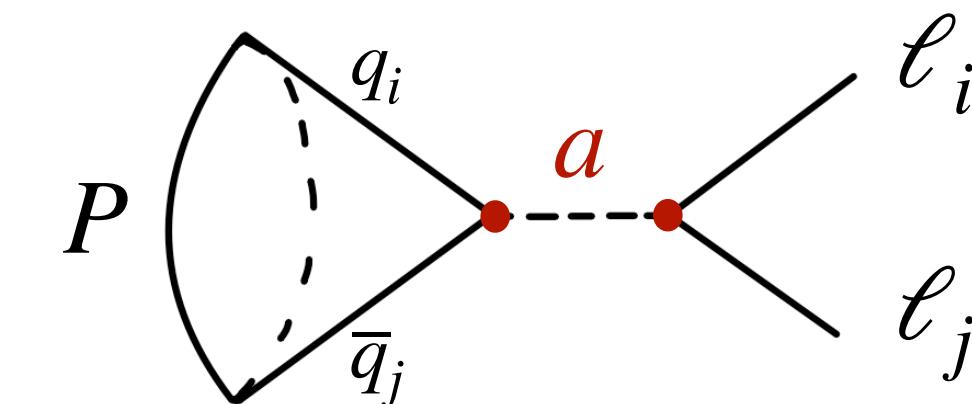
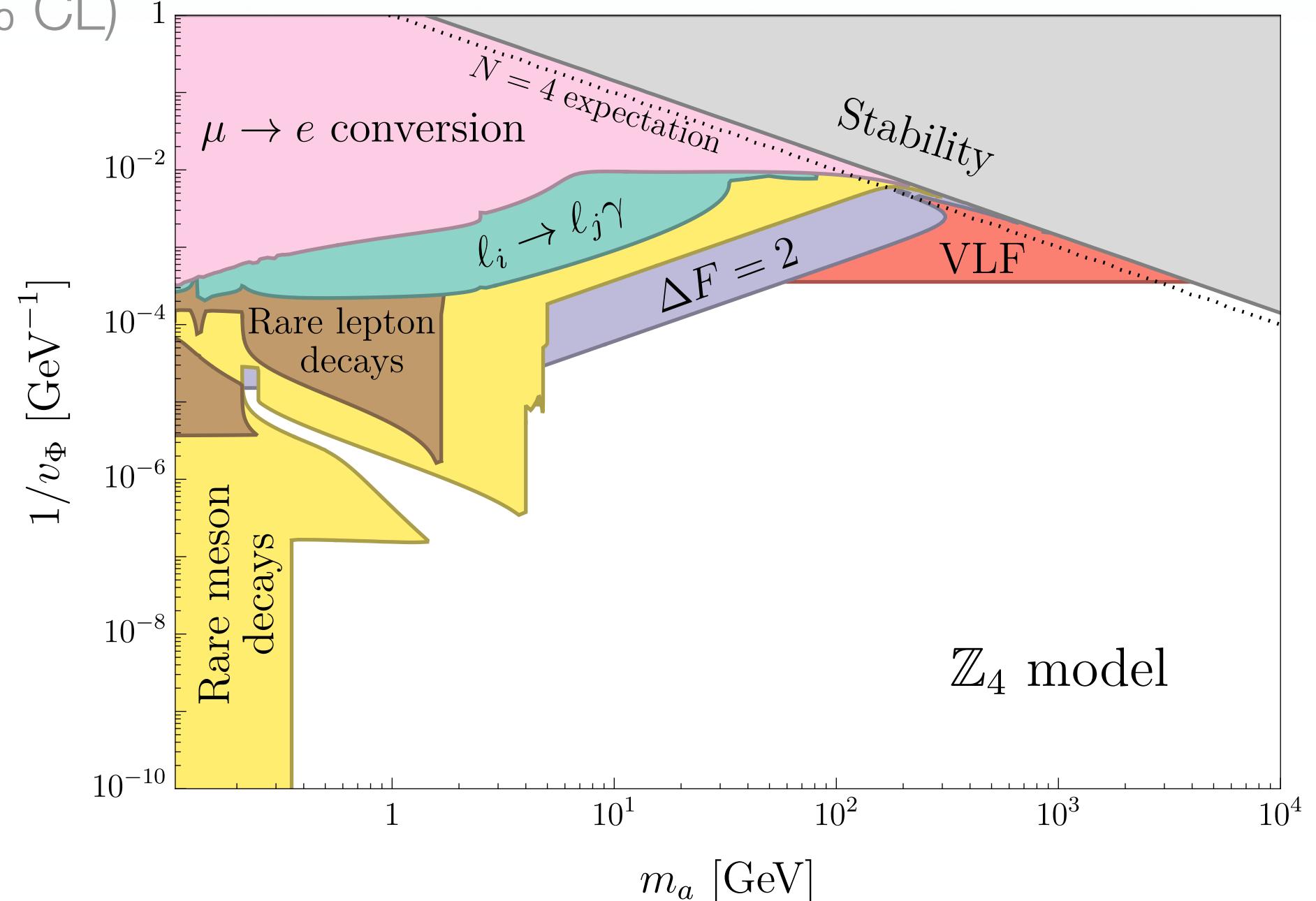
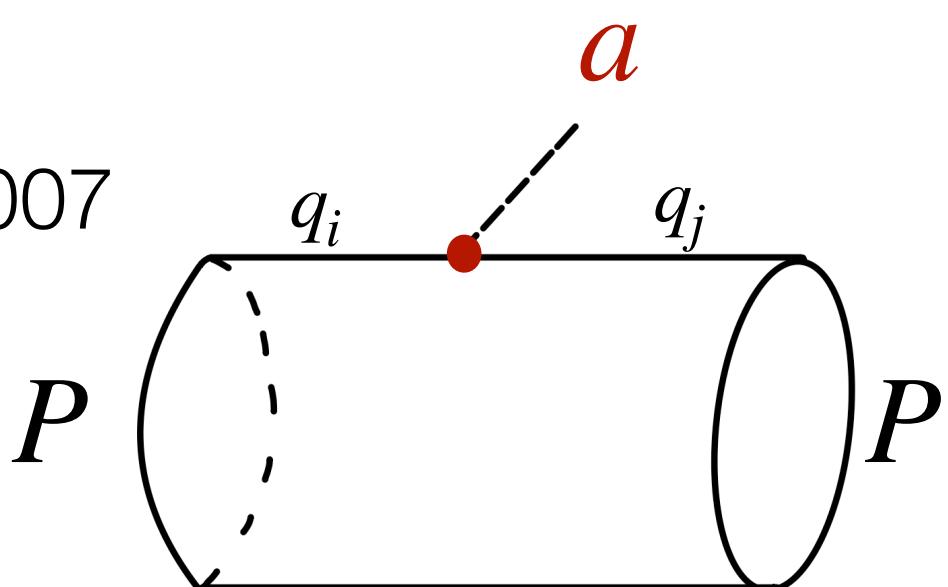
@ E871 2000

$$\text{Br}(K_L \rightarrow \mu e) < 6.3 \times 10^{-12}$$

@ E871 1998

⋮

(Bounds at 95% CL)



3. ALP pheno

MESON DECAYS

Invisible

$$\text{Br}(K^+ \rightarrow \pi^+ + a) < 9.5 \times 10^{-11}$$

@ E787+E949 2007

$$\text{Br}(B_0 \rightarrow K_0^* + a) < 1.0 \times 10^{-4}$$

@ BaBar 2013

$$\text{Br}(B^+ \rightarrow K^+ \bar{\nu}\nu) = (2.3 \pm 0.7) \times 10^{-5}$$

@ Belle II 2023

⋮

Visible - purely leptonic

$$\text{Br}(K_L \rightarrow \mu\mu) = (6.8 \pm 0.1) \times 10^{-11}$$

@ E871 2000

$$\text{Br}(K_L \rightarrow \mu e) < 6.3 \times 10^{-12}$$

@ E871 1998

⋮

Visible - prompt

$$\text{Br}(B^+ \rightarrow K^+ \tau\mu) < 7.7 \times 10^{-6}$$

@ Belle 2022, ...

$$\text{Br}(B_0 \rightarrow K_0^* \tau\mu) < 9.8 \times 10^{-6}$$

@ LHCb 2022

$$R_K, R_{K^*}$$

⋮

Visible - displaced

\longrightarrow D decays?

$$\text{Br}(B_0 \rightarrow K_0^* \mu\mu) \lesssim 10^{-10} f(m_a, \tau_a)$$

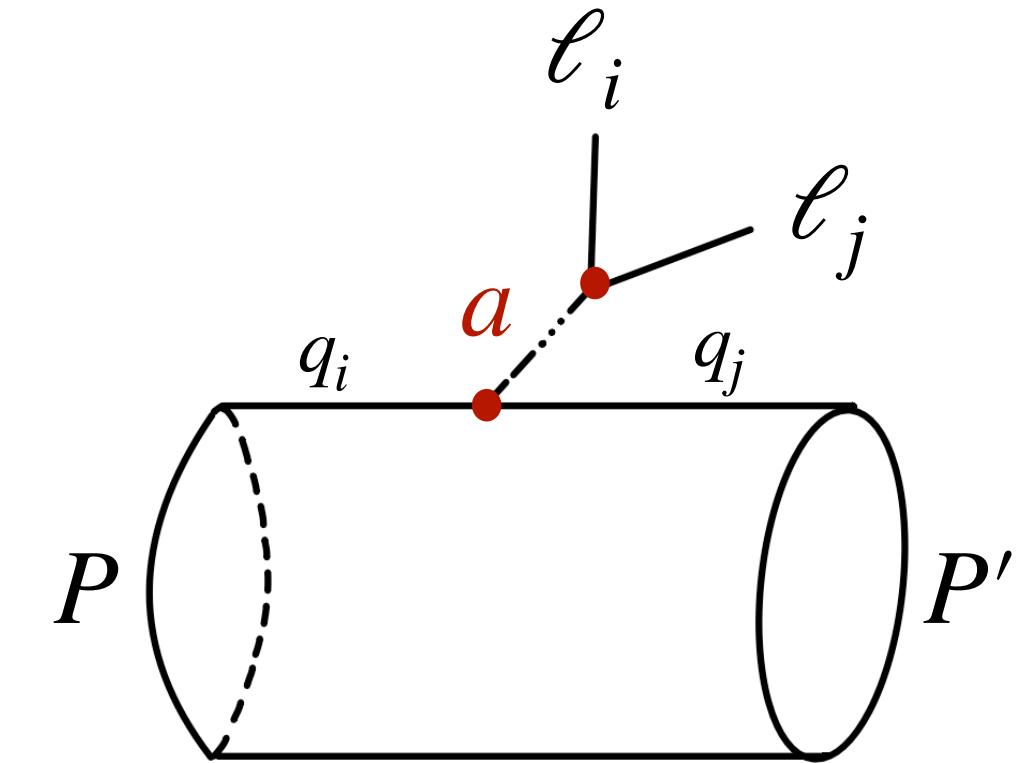
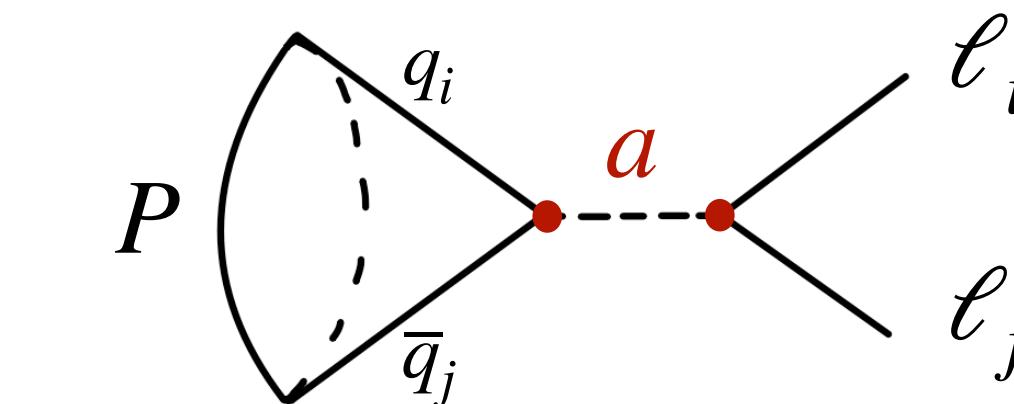
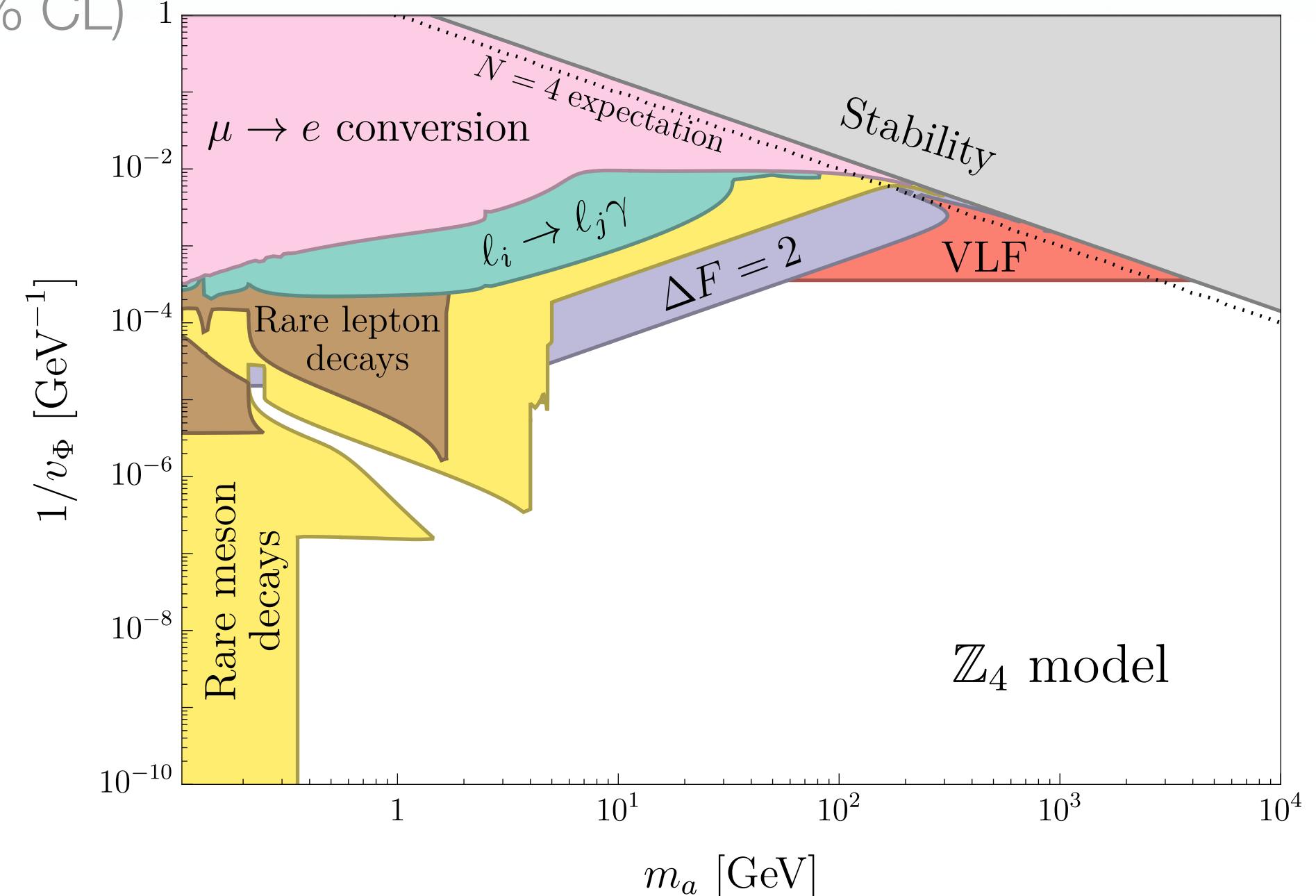
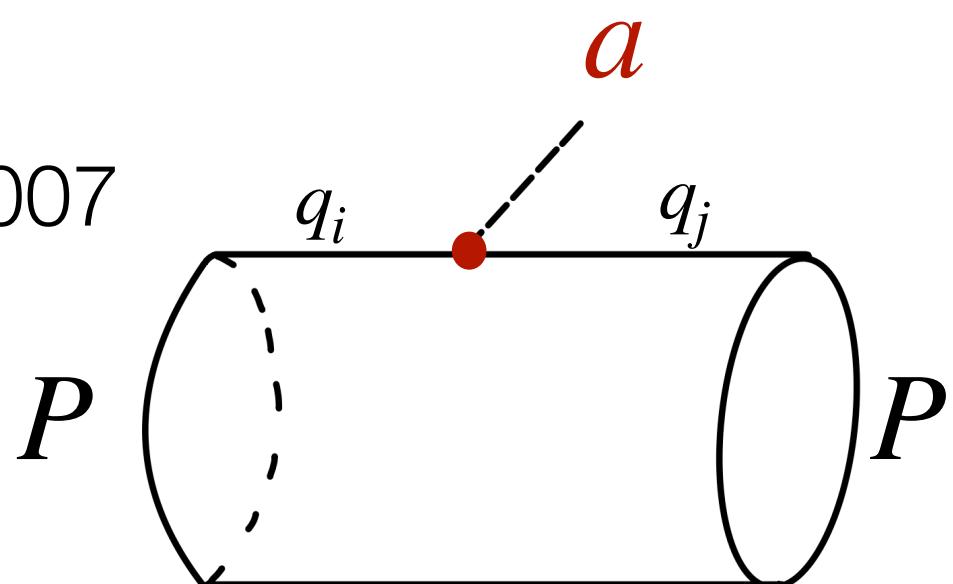
@ LHCb 2015

$$\text{Br}(B^+ \rightarrow K^+ \mu\mu) \lesssim 10^{-10} f(m_a, \tau_a)$$

@ LHCb 2016

⋮
ALP phenomenology

(Bounds at 95% CL)



3. ALP pheno

LEPTON DECAYS

Invisible

$\text{Br}(\tau \rightarrow \mu + a) < 10^{-4} - 10^{-3}$ @ Belle II 2022

$\text{Br}(\tau \rightarrow e + a) < 10^{-4} - 10^{-3}$ @ Belle II 2022

$\text{Br}(\mu \rightarrow e + a) \lesssim 10^{-5}$ @ TRIUMF 1986, 2014

\vdots (TWIST)

Visible - prompt

$\text{Br}(\tau \rightarrow K\mu) < 3 \times 10^{-8}$ @ Belle 2010, BaBar 2009, ...

$\text{Br}(\tau \rightarrow 3\mu) < 2.5 \times 10^{-8}$ @ Belle II 2024, LHCb 2015, ...

$\text{Br}(\tau \rightarrow \mu\mu e) < 2.2 \times 10^{-8}$ @ Belle 2010, BaBar 2010, ...

\vdots

Visible - displaced

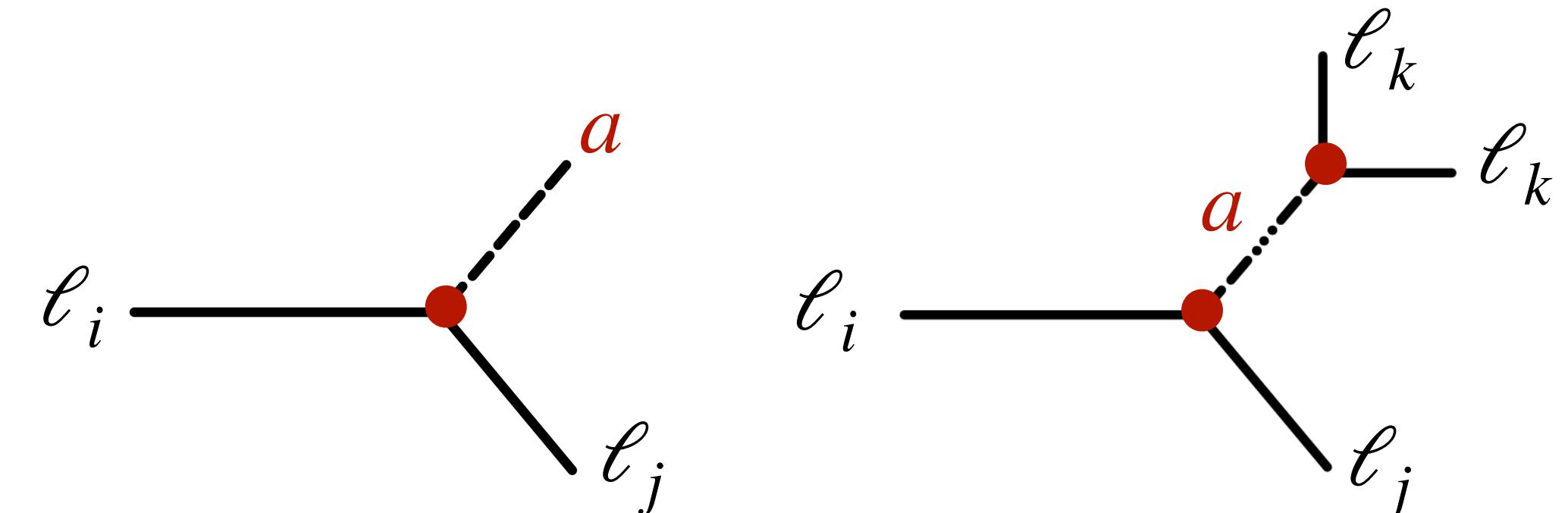
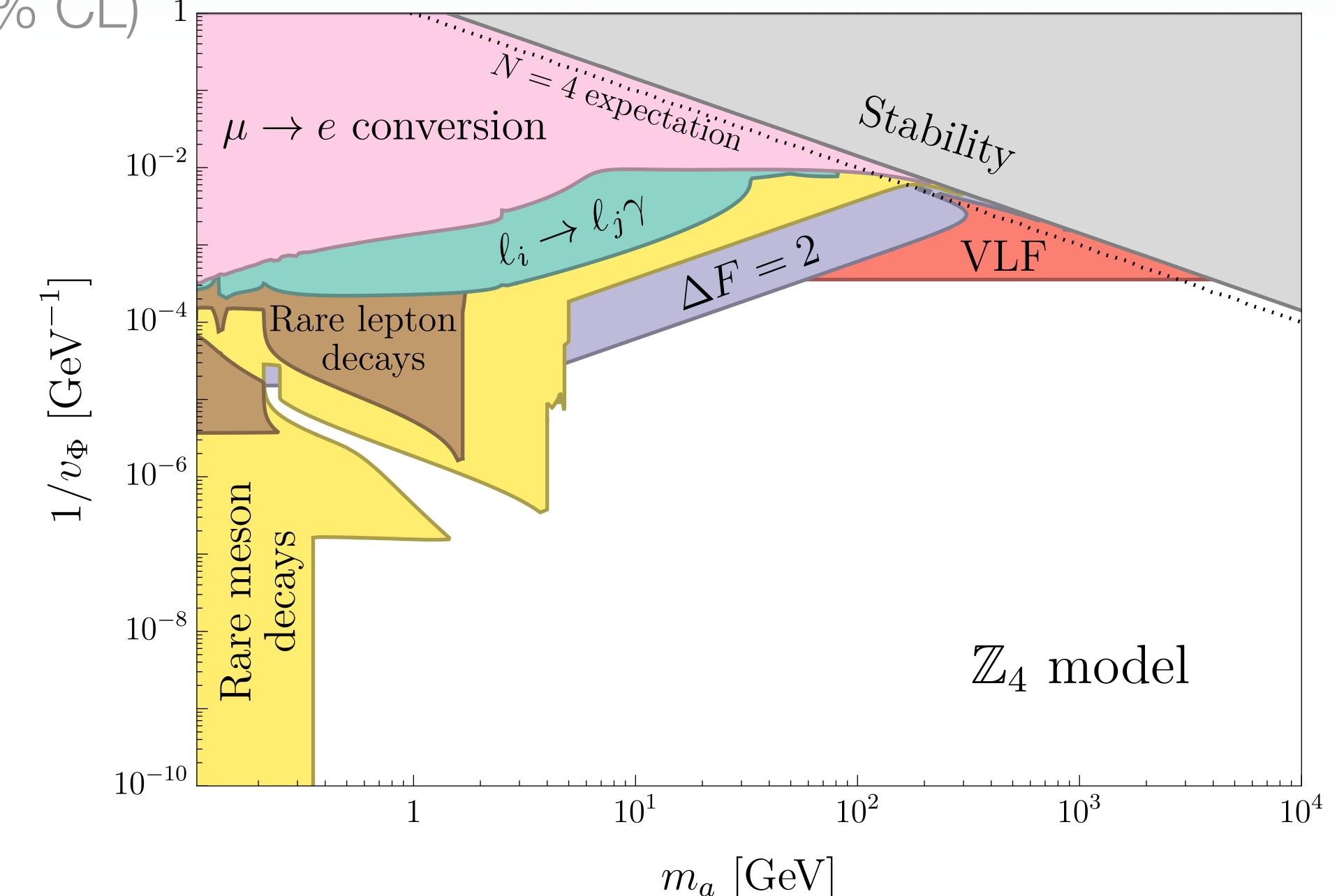
$\text{Br}(\mu \rightarrow 3e) \lesssim 10^{-12} f(m_a, c\tau_a)$ @ SINDRUM 1986

$\text{Br}(\mu \rightarrow e\gamma\gamma) \lesssim 10^{-11} f(m_a, c\tau_a)$ @ CrystalBox 1988, MEG 2020

**Need searches for displaced
 τ decays! ($\tau \rightarrow 3\mu, \mu\mu e, \dots$)**

Heeck,
Rodejohann (2017)

(Bounds at 95% CL)



3. ALP pheno

MESONS MIXING

$P_0 - \bar{P}_0$ oscillations with $P = K, B, B_s, D$

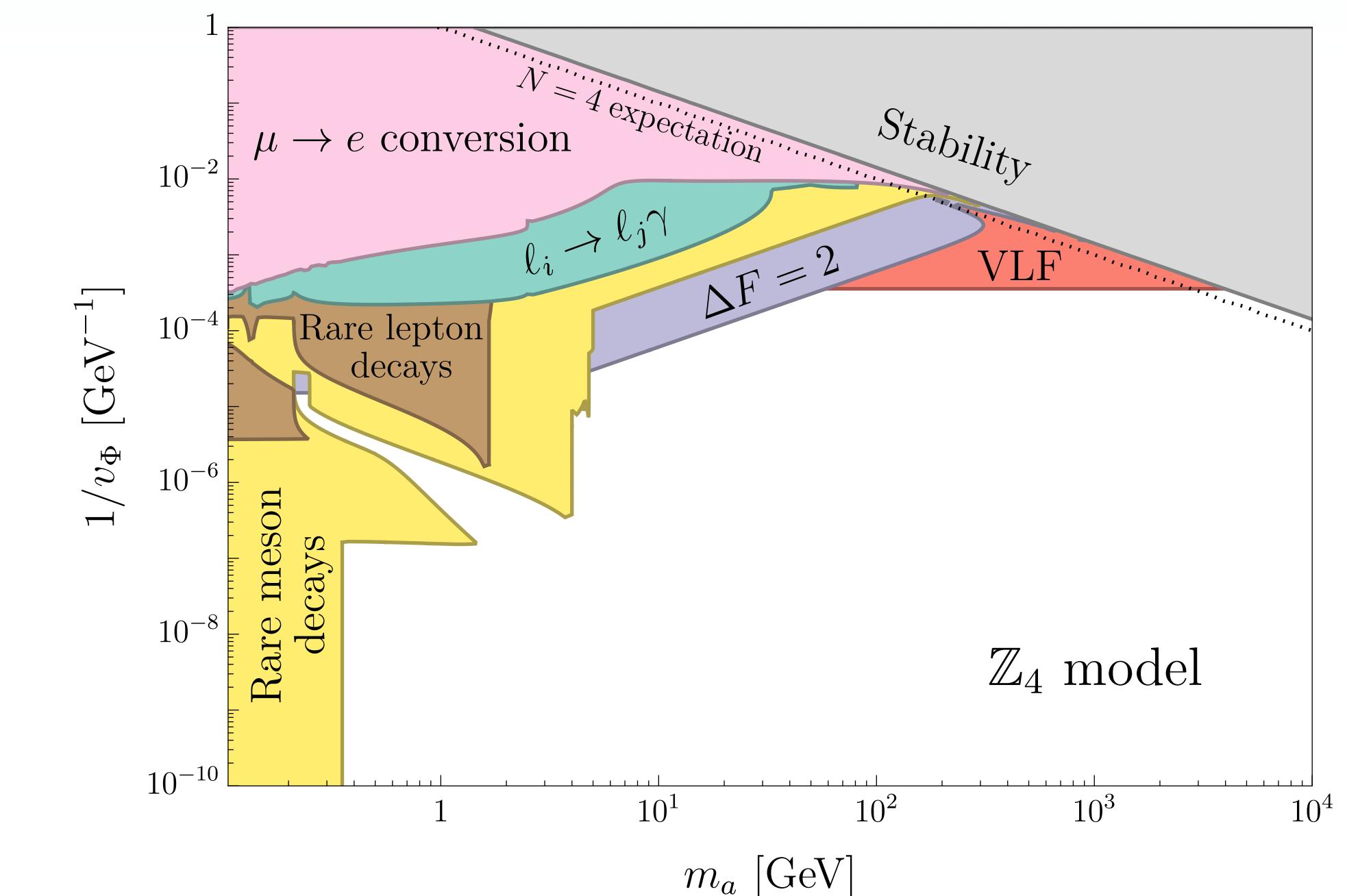
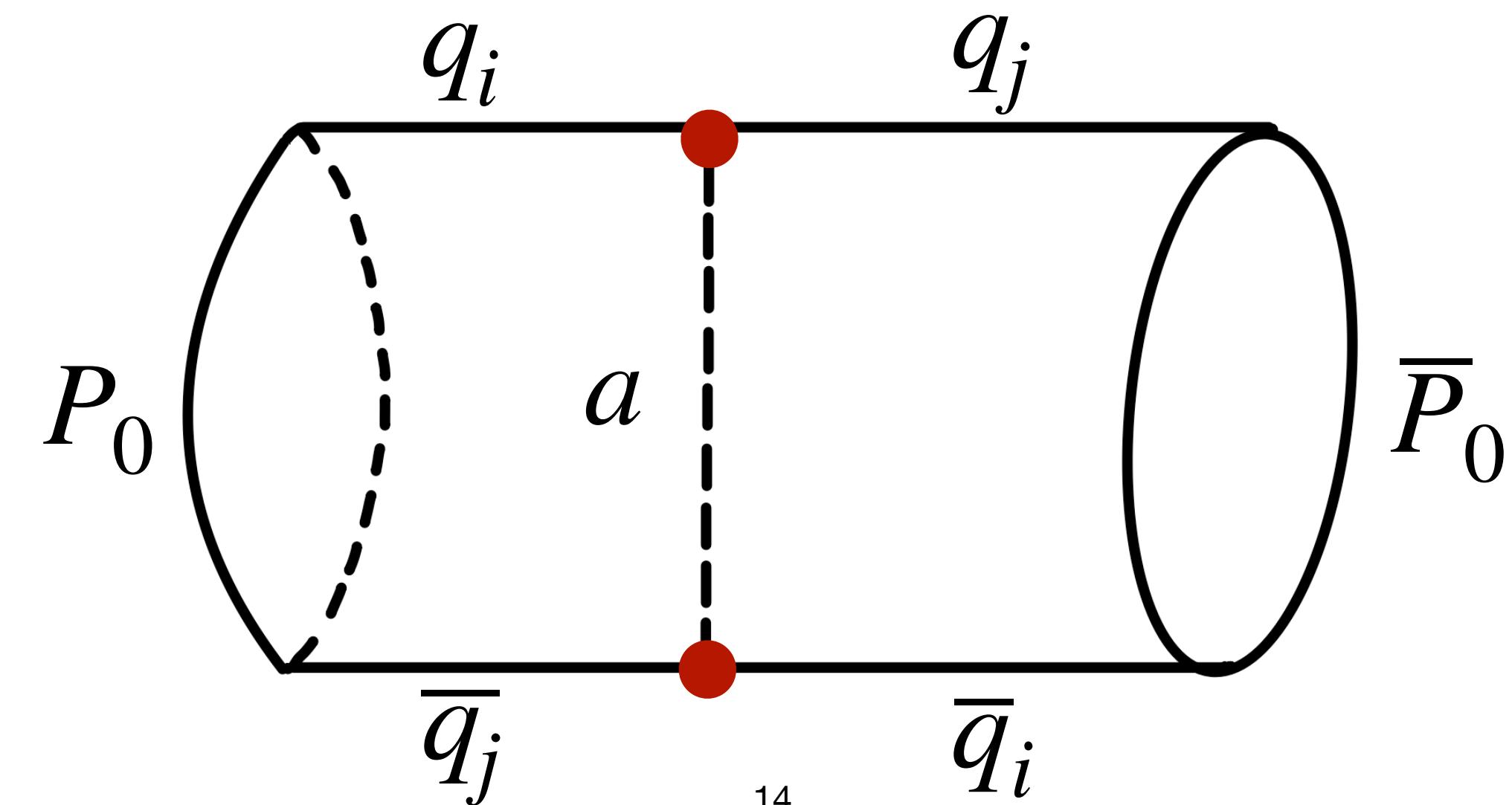
$$\epsilon_K = (2.23 \pm 0.01) \times 10^{-3}$$

$$\Delta M_s = 17.765 \pm 0.006 \text{ ps}^{-1}$$

$$\Delta M_d = 0.506 \pm 0.002 \text{ ps}^{-1}$$

⋮

@ PDG world average 2024
(LHCb dominating $\Delta M_d, \Delta M_s$)



3. ALP pheno

Lepton dipoles

$\text{Br}(\tau \rightarrow \mu\gamma) < 5.7 \times 10^{-8}$ @ BaBar 2009, Belle 2008, ...

$\text{Br}(\tau \rightarrow e\gamma) < 4.3 \times 10^{-8}$ @ BaBar 2009, Belle 2008, ...

$\text{Br}(\mu \rightarrow e\gamma) < 5.5 \times 10^{-13}$ @ MEG 2016

$\mu \rightarrow e$ in nuclei

$\text{Br}(\mu \rightarrow e) < 9 \times 10^{-13}$ @ SINDRUM II 2006 (Au)

Lepton EDMs

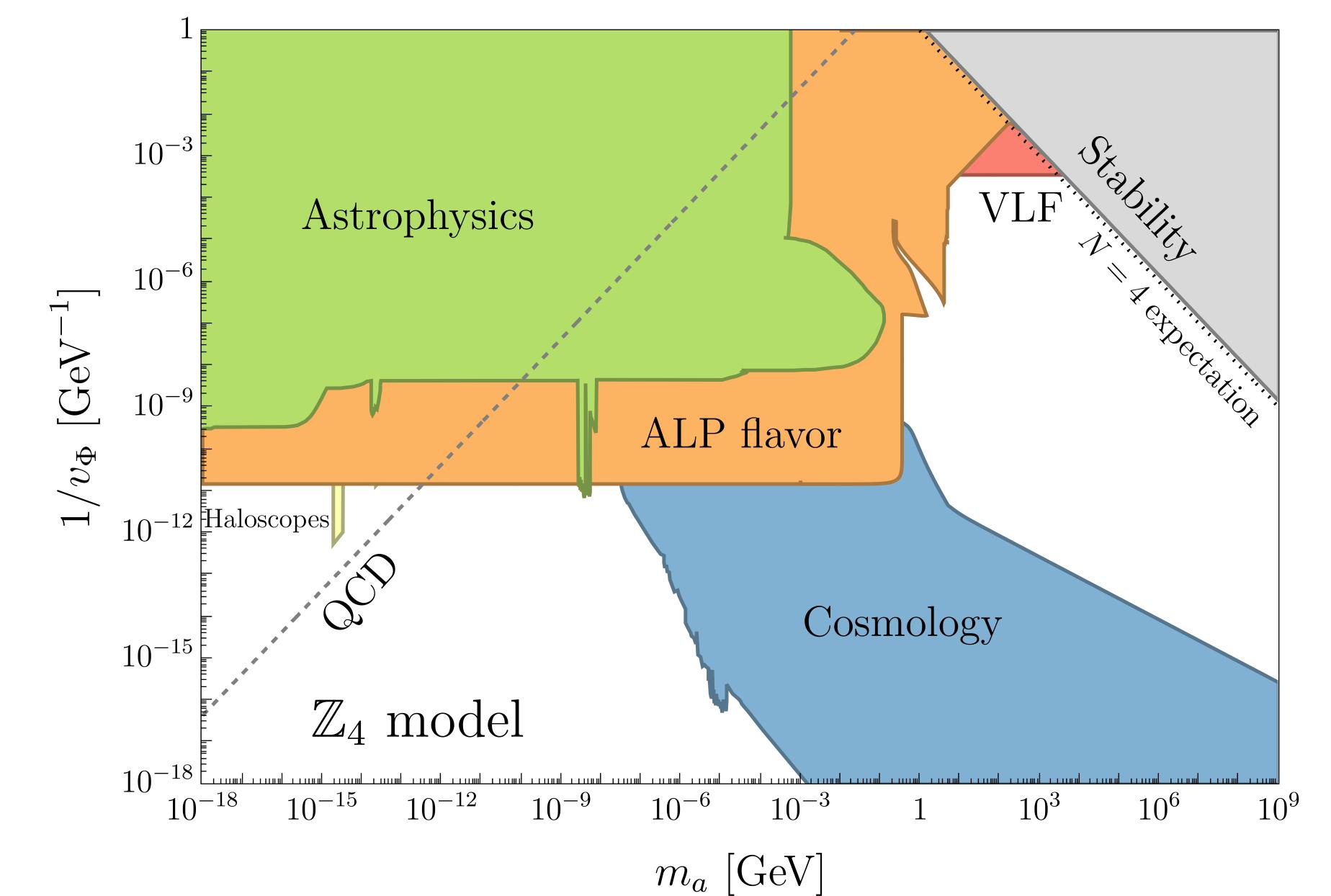
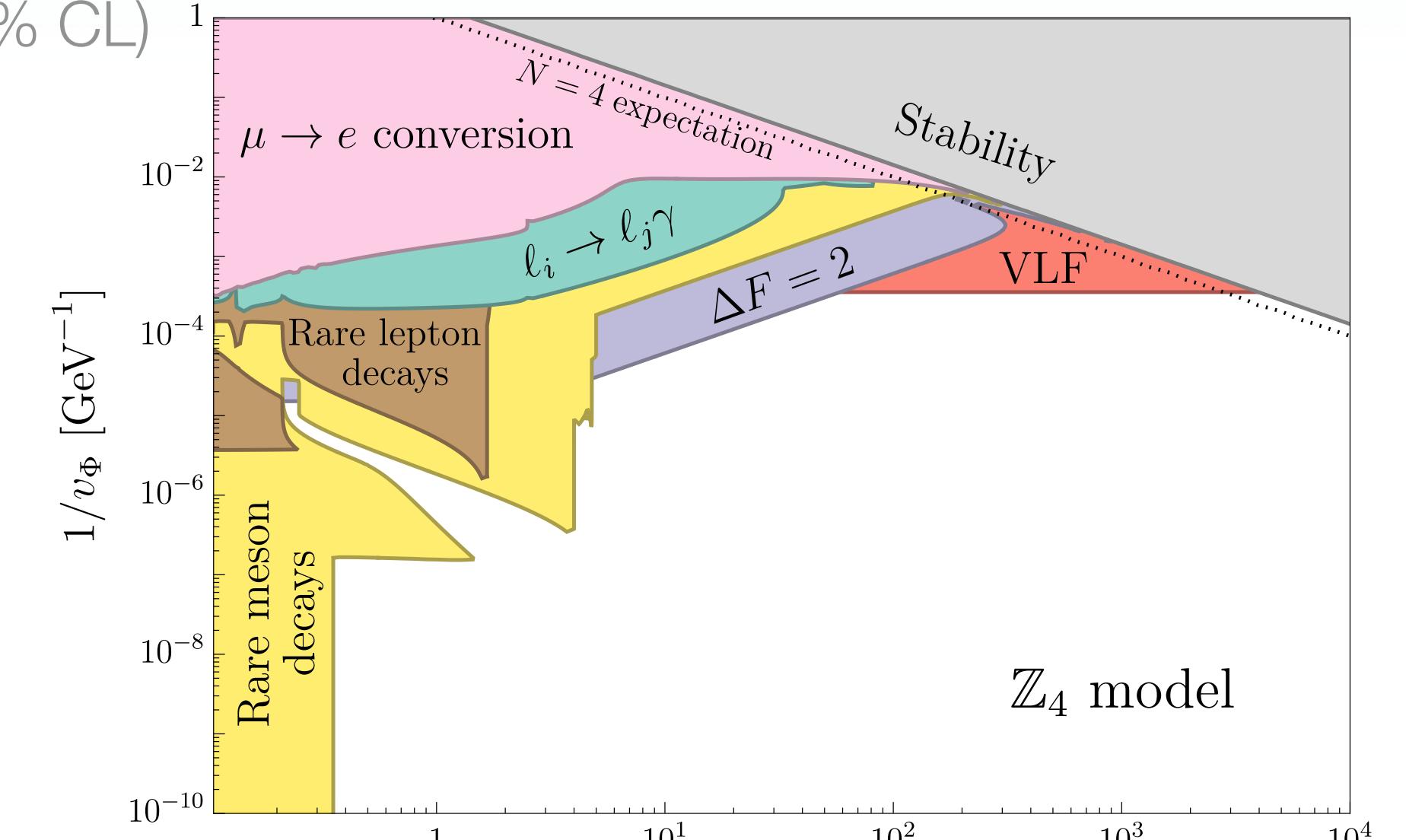
$|d_e|/e < 5.4 \times 10^{-30}$ cm @ Cornell et al. 2022

Neutron EDM

$|d_n|/e < 2.6 \times 10^{-26}$ cm @ Abel et al. 2020

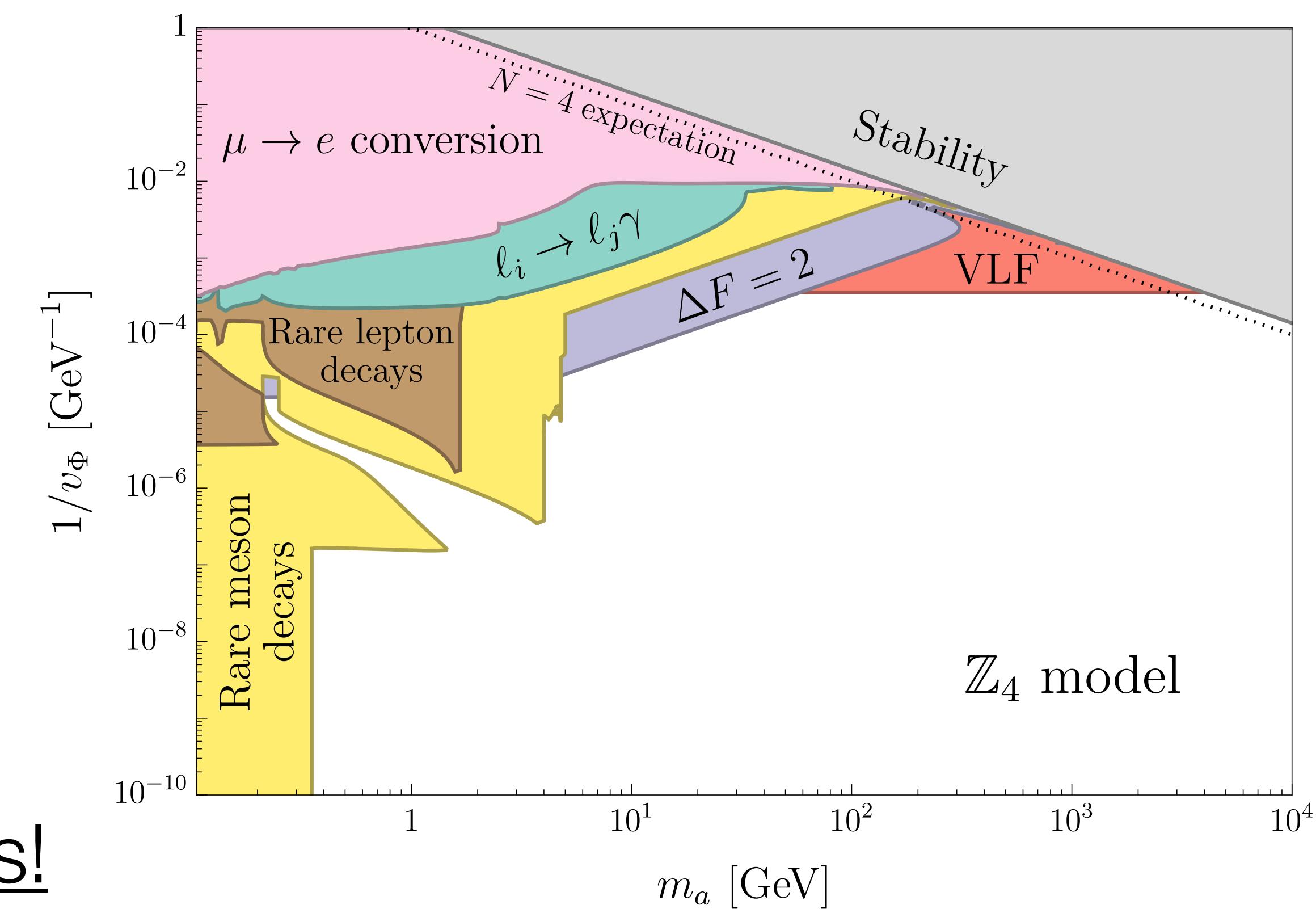
Astrophysics, Cosmology, Haloscopes, ...

(Bounds at 95% CL)



4. Conclusion

- ALPs: very popular,
theoretically motivated particles
- Concrete models (FN) able to provide
guidelines and correlations in observables
- LHCb + other experiments delivered
key observations for
constraining/discovering ALPs
- Would be nice:
displaced τ , D (and more) decays searches!



Thank you for your attention!

Some observables (not exhaustive)

Process	Exp. limit	Ref.	Process	Exp. limit	Ref.	Process	Exp. limit	Ref.
$\pi^0 \rightarrow e\mu$	4.8×10^{-10}	[30]	$\tau \rightarrow e\pi^0$	1.1×10^{-7}	[30]	$\tau \rightarrow \mu\pi^0$	1.5×10^{-7}	[30]
$\eta \rightarrow e\mu$	8.1×10^{-6}	[30]	$\tau \rightarrow e\eta$	1.2×10^{-7}	[30]	$\tau \rightarrow \mu\eta$	8.7×10^{-8}	[30]
$\eta' \rightarrow e\mu$	6.3×10^{-4}	[30]	$\tau \rightarrow e\eta'$	2.1×10^{-7}	[30]	$\tau \rightarrow \mu\eta'$	1.8×10^{-7}	[30]
			$\tau \rightarrow e\rho^0$	2.4×10^{-8}	[30]	$\tau \rightarrow \mu\rho^0$	1.6×10^{-8}	[30]
$K_L \rightarrow e\mu$	6.3×10^{-12}	[30]	$\tau \rightarrow eK_S$	3.5×10^{-8}	[30]	$\tau \rightarrow \mu K_S$	3.1×10^{-8}	[30]
$K^+ \rightarrow \pi^+ \mu^- e^+$	1.1×10^{-10}	[30]	$\tau \rightarrow eK^{*0}$	4.3×10^{-8}	[30]	$\tau \rightarrow \mu K^{*0}$	7.9×10^{-8}	[30]
$K_L \rightarrow \pi^0 e\mu$	1.0×10^{-10}	[30]						
$D^0 \rightarrow e\mu$	1.7×10^{-8}	[30]	$\tau \rightarrow eD^0$	–				
$D^+ \rightarrow \pi^+ e\mu$	4.5×10^{-7}	[38]						
$D_s \rightarrow K^+ e\mu$	1.5×10^{-6}	[38]						
$B^0 \rightarrow e\mu$	1.3×10^{-9}	[30]	$B^0 \rightarrow e\tau$	2.1×10^{-5}	[30]	$B^0 \rightarrow \mu\tau$	1.4×10^{-5}	[30]
$B^+ \rightarrow \pi^+ e\mu$	1.2×10^{-7}	[39]	$B^+ \rightarrow \pi^+ e\tau$	1.0×10^{-4}	[30]	$B^+ \rightarrow \pi^+ \mu\tau$	9.7×10^{-5}	[30]
$B^0 \rightarrow \rho^0 e\mu$	–		$B^0 \rightarrow \rho^0 e\tau$	–		$B^0 \rightarrow \rho^0 \mu\tau$	–	
$B_s \rightarrow K^0 e\mu$	–		$B_s \rightarrow K^0 e\tau$	–		$B_s \rightarrow K^0 \mu\tau$	–	
$B_s \rightarrow e\mu$	7.2×10^{-9}	[30]	$B_s \rightarrow e\tau$	1.9×10^{-3}	[10]	$B_s \rightarrow \mu\tau$	4.2×10^{-5}	[30]
$B^+ \rightarrow K^+ e\mu$	1.8×10^{-8}	[9]	$B^+ \rightarrow K^+ e\tau$	4.1×10^{-5}	[10]	$B^+ \rightarrow K^+ \mu\tau$	4.1×10^{-5}	[10]
$B^0 \rightarrow K^* e\mu$	1.2×10^{-8}	[9]	$B^0 \rightarrow K^* e\tau$	–		$B^0 \rightarrow K^* \mu\tau$	2.2×10^{-5}	[9]
$B_s \rightarrow \phi e\mu$	2.0×10^{-8}	[9]	$B_s \rightarrow \phi e\tau$	–		$B_s \rightarrow \phi \mu\tau$	–	
$\phi \rightarrow e\mu$	2.7×10^{-6}	[30]	$\tau \rightarrow e\phi$	5.5×10^{-8}	[30]	$\tau \rightarrow \mu\phi$	1.1×10^{-7}	[30]
$J/\psi \rightarrow e\mu$	6.1×10^{-9}	[40]	$J/\psi \rightarrow e\tau$	1.0×10^{-7}	[40]	$J/\psi \rightarrow \mu\tau$	2.7×10^{-6}	[30]
$\Upsilon \rightarrow e\mu$	5.2×10^{-7}	[30]	$\Upsilon \rightarrow e\tau$	3.6×10^{-6}	[30]	$\Upsilon \rightarrow \mu\tau$	3.6×10^{-6}	[30]

Plakias, Sumensari (2023)

Decay mode	Exp. limit	Future prospects	Ref.
$\mu \rightarrow e\gamma$	4.2×10^{-13}	$\approx 6 \times 10^{-14}$	[43]
$\mu \rightarrow 3e$	1.0×10^{-12}	$\approx 10^{-16}$	[43]
$\mu \rightarrow e\gamma\gamma$	7.2×10^{-11}	–	[43]
$\mu \rightarrow e + \text{inv}$	$\approx 10^{-5}$	–	[44]
$\mu^- \text{Ti} \rightarrow e^- \text{Ti}$	4.3×10^{-12}	–	[45]
$\mu^- \text{Au} \rightarrow e^- \text{Au}$	7×10^{-13}	$\approx 10^{-17}$	[42]
$\mu^- \text{Al} \rightarrow e^- \text{Al}$	–	$\approx 10^{-17}$	[17, 19]
$\tau \rightarrow e\gamma$	3.3×10^{-8}	$\approx 3 \times 10^{-9}$	[43]
$\tau \rightarrow 3e$	2.7×10^{-8}	$\approx 5 \times 10^{-10}$	[43]
$\tau \rightarrow e\mu^+ \mu^-$	1.7×10^{-8}	$\approx 6 \times 10^{-10}$	[43]
$\tau \rightarrow e + \text{inv}$	$\approx 5 \times 10^{-3}$	–	[46]
$\tau \rightarrow \mu\gamma$	4.4×10^{-8}	$\approx 10^{-9}$	[43]
$\tau \rightarrow 3\mu$	2.1×10^{-8}	$\approx 4 \times 10^{-10}$	[43]
$\tau \rightarrow \mu e^+ e^-$	1.8×10^{-8}	$\approx 3 \times 10^{-10}$	[43]
$\tau \rightarrow \mu + \text{inv}$	$\approx 5 \times 10^{-3}$	–	[46]

Cornella, Paradisi, Sumensari (2019)

flavio observables

flav-io.github.io/docs/observables

Backup

Z_8 model

$$c^d \sim \begin{pmatrix} 1 & \frac{m_d}{m_s} & \frac{m_d}{m_b} \\ \frac{m_d}{m_s} & 1 & \frac{m_s}{m_b} \\ \frac{m_d}{m_b} & \frac{m_s}{m_b} & 1 \end{pmatrix}$$

$$c^e \sim c^d$$

$$c^u \sim \begin{pmatrix} 1 & \sqrt{\frac{m_u}{m_c}} & \sqrt{\frac{m_u}{m_t}} \\ \sqrt{\frac{m_u}{m_c}} & 1 & \sqrt{\frac{m_c}{m_t}} \\ \sqrt{\frac{m_u}{m_t}} & \sqrt{\frac{m_c}{m_t}} & \frac{m_c}{m_t} \end{pmatrix}$$

Better masses fit

Up physics more important! (D mesons)

ALP phenomenology

