Dielectric-top membranes in plane-wave backgrounds

Georgios Linardopoulos

Asia Pacific Center for Theoretical Physics (APCTP) Interfaces and defects in strongly coupled matter research group





1st INPP Demokritos-APCTP meeting National Center for Scientific Research Demokritos, 04 October 2024

based on my work with M. Axenides and M. Floratos, PLB **773** (2017) 265 [arxiv:1707.02878], PRD **97** (2018) 126019 [arxiv:1712.06544], PRD **104** (2021) 106002 [arxiv:2109.01088] (with D. Katsinis), as well as work in progress

Table of Contents

Introduction



Spherical dielectric tops in plane-wave backgrounds







Section 1

Introduction

・ロト ・日 ・ ・ 目 ・ ・ 目 ・ つ へ ()
3 / 53

Plane-fronted gravitational waves with parallel rays (pp-waves)

• Plane-fronted (gravitational) waves with parallel rays (or pp-waves) are solutions of the 4-dimensional (vacuum) Einstein equations. In Brinkmann coordinates,

$$ds^{2} = 2dudv + H(u, x, y)du^{2} + dx^{2} + dy^{2}, \quad \nabla^{2}H(u, x, y) = 0.$$

Brinkmann (1925)

Plane-fronted gravitational waves with parallel rays (pp-waves)

• Plane-fronted (gravitational) waves with parallel rays (or pp-waves) are solutions of the 4-dimensional (vacuum) Einstein equations. In Brinkmann coordinates,

$$ds^{2} = 2dudv + H(u, x, y)du^{2} + dx^{2} + dy^{2}, \quad \nabla^{2}H(u, x, y) = 0.$$

Brinkmann (1925)

• Equivalently, pp-waves can be defined as spacetimes that admit a covariantly constant null Killing vector:

$$\nabla_m k_n = 0, \qquad k^n k_n = 0.$$

Ehlers-Kundt (1962)

Plane-fronted means that pp-waves can be completely covered by 2d wave fronts orthogonal to the wave vector k. The wave fronts are planes which propagate parallel to each other in the direction of k = constant ("parallel rays").

Plane-fronted gravitational waves with parallel rays (pp-waves)

• Plane-fronted (gravitational) waves with parallel rays (or pp-waves) are solutions of the 4-dimensional (vacuum) Einstein equations. In Brinkmann coordinates,

$$ds^2 = 2dudv + H(u, x, y)du^2 + dx^2 + dy^2, \qquad \nabla^2 H(u, x, y) = 0.$$

Brinkmann (1925)

• Equivalently, pp-waves can be defined as spacetimes that admit a covariantly constant null Killing vector:

$$\nabla_m k_n = 0, \qquad k^n k_n = 0.$$

Ehlers-Kundt (1962)

Plane-fronted means that pp-waves can be completely covered by 2d wave fronts orthogonal to the wave vector k. The wave fronts are planes which propagate parallel to each other in the direction of k = constant ("parallel rays").

• By choosing H(u, x, y), Brinkmann metric also solves Einstein-Maxwell theory... Plane waves are special pp-waves:

$$H(u, x, y) = a(u)(x^{2} - y^{2}) + 2b(u)xy + c(u)(x^{2} + y^{2}), \quad \text{(in vacuum, } c(u) = 0),$$

gravitational analogs of plane electromagnetic waves... providing the field very far from finite gravity sources...

• Most general metric of a d + 1 dimensional spacetime with a covariantly constant null Killing vector k:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + 2A_{j}(x^{+}, x^{i})dx^{+}dx^{j} + g_{jk}(x^{+}, x^{i})dx^{j}dx^{k}, \quad x^{\pm} \equiv \frac{1}{\sqrt{2}}\left(x^{0} \pm x^{d}\right),$$

where $i, j = 1, 2, ..., d - 1 \& F(u, x^i), A_j(u, x^i), g_{jk}(u, x^i)$ are determined from the sugra equations of motion...

• Most general metric of a d + 1 dimensional spacetime with a covariantly constant null Killing vector k:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + 2A_{j}(x^{+}, x^{i})dx^{+}dx^{j} + g_{jk}(x^{+}, x^{i})dx^{j}dx^{k}, \quad x^{\pm} \equiv \frac{1}{\sqrt{2}}\left(x^{0} \pm x^{d}\right),$$

where $i, j = 1, 2, ..., d - 1 \& F(u, x^i), A_j(u, x^i), g_{jk}(u, x^i)$ are determined from the sugra equations of motion...

• For $A_j = 0$, $g_{jk} = \delta_{jk}$, we retrieve the d + 1 dimensional Brinkmann metric:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + dx^{i}dx^{i}.$$

In this form, pp-waves are α' -exact solutions of supergravity & string theory...

• Most general metric of a d + 1 dimensional spacetime with a covariantly constant null Killing vector k:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + 2A_{j}(x^{+}, x^{i})dx^{+}dx^{j} + g_{jk}(x^{+}, x^{i})dx^{j}dx^{k}, \quad x^{\pm} \equiv \frac{1}{\sqrt{2}}\left(x^{0} \pm x^{d}\right),$$

where $i, j = 1, 2, \dots d - 1$ & $F(u, x^i), A_j(u, x^i), g_{jk}(u, x^i)$ are determined from the sugra equations of motion...

• For $A_j = 0$, $g_{jk} = \delta_{jk}$, we retrieve the d + 1 dimensional Brinkmann metric:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + dx^{i}dx^{i}.$$

イロト イロト イヨト イヨト ヨー わへの

5 / 53

In this form, pp-waves are $\alpha'\text{-exact}$ solutions of supergravity & string theory...

• Plane-waves are pp-waves for which $F(x^+, x^i) = f_{ij}(x^+)x^i x^j$, $A_j = 0$ and $g_{jk} = \delta_{jk}$: $ds^2 = -2dx^+ dx^- - f_{ij}(x^+)x^i x^j dx^+ dx^+ + dx^i dx^i.$

• Most general metric of a d + 1 dimensional spacetime with a covariantly constant null Killing vector k:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + 2A_{j}(x^{+}, x^{i})dx^{+}dx^{j} + g_{jk}(x^{+}, x^{i})dx^{j}dx^{k}, \quad x^{\pm} \equiv \frac{1}{\sqrt{2}}\left(x^{0} \pm x^{d}\right),$$

where $i, j = 1, 2, \dots d - 1$ & $F(u, x^i), A_j(u, x^i), g_{jk}(u, x^i)$ are determined from the sugra equations of motion...

• For $A_j = 0$, $g_{jk} = \delta_{jk}$, we retrieve the d + 1 dimensional Brinkmann metric:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + dx^{i}dx^{i}.$$

In this form, pp-waves are $\alpha'\text{-exact}$ solutions of supergravity & string theory...

- Plane-waves are pp-waves for which $F(x^+, x^i) = f_{ij}(x^+)x^ix^j$, $A_j = 0$ and $g_{jk} = \delta_{jk}$: $ds^2 = -2dx^+dx^- - f_{ij}(x^+)x^ix^jdx^+dx^+ + dx^idx^i.$
- Homogeneous plane-waves have $f_{ij}(x^+) = \mu_{ij}^2$, constant:

$$ds^{2} = -2dx^{+}dx^{-} - \mu_{ij}^{2}x^{i}x^{j}dx^{+}dx^{+} + dx^{i}dx^{i}.$$

4 ロ ト 4 部 ト 4 差 ト 4 差 ト 差 の Q ()
5 / 53

• Most general metric of a d + 1 dimensional spacetime with a covariantly constant null Killing vector k:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + 2A_{j}(x^{+}, x^{i})dx^{+}dx^{j} + g_{jk}(x^{+}, x^{i})dx^{j}dx^{k}, \quad x^{\pm} \equiv \frac{1}{\sqrt{2}}\left(x^{0} \pm x^{d}\right),$$

where $i, j = 1, 2, \dots d - 1$ & $F(u, x^i), A_j(u, x^i), g_{jk}(u, x^i)$ are determined from the sugra equations of motion...

• For $A_j = 0$, $g_{jk} = \delta_{jk}$, we retrieve the d + 1 dimensional Brinkmann metric:

$$ds^{2} = -2dx^{+}dx^{-} - F(x^{+}, x^{i})dx^{+}dx^{+} + dx^{i}dx^{i}.$$

In this form, pp-waves are $\alpha'\text{-exact}$ solutions of supergravity & string theory...

- Plane-waves are pp-waves for which $F(x^+, x^i) = f_{ij}(x^+)x^i x^j$, $A_j = 0$ and $g_{jk} = \delta_{jk}$: $ds^2 = -2dx^+ dx^- - f_{ij}(x^+)x^i x^j dx^+ dx^+ + dx^i dx^i.$
- Homogeneous plane-waves have $f_{ij}(x^+) = \mu_{ij}^2$, constant:

$$ds^2 = -2dx^+dx^- - \mu_{ij}^2 x^i x^j dx^+ dx^+ + dx^i dx^i.$$

• Homogeneous and isotropic plane-waves have $\mu_{ij} = \mu$:

$$ds^{2} = -2dx^{+}dx^{-} - \mu^{2}x^{i}x^{i}dx^{+}dx^{+} + dx^{i}dx^{i}.$$

5 / 53

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

• Penrose limit: d + 1 dimensional Brinkmann metric descends from any metric by blowing up spacetime around null geodesics ("zooming in" to them)... New solutions of Einstein's equations can be constructed from known ones...

Penrose (1976) & Güven (2000)

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

Penrose limit: d + 1 dimensional Brinkmann metric descends from any metric by blowing up spacetime around null
geodesics ("zooming in" to them)... New solutions of Einstein's equations can be constructed from known ones...

Penrose (1976) & Güven (2000)

• The d + 1 dimensional Brinkmann metric can be written in a form which resembles linearized gravity:

$$g_{mn} = \eta_{mn} + h_{mn}, \qquad h_{mn} \equiv -F(x^+, x^i)k_mk_n,$$

and solves Einstein's equations even when h_{mn} is not small... As a consequence, many properties of flat Minkowski spaces can be uplifted to pp-wave backgrounds with only minor modifications...

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

Penrose limit: d + 1 dimensional Brinkmann metric descends from any metric by blowing up spacetime around null geodesics ("zooming in" to them)... New solutions of Einstein's equations can be constructed from known ones...

Penrose (1976) & Güven (2000)

• The d + 1 dimensional Brinkmann metric can be written in a form which resembles linearized gravity:

$$g_{mn} = \eta_{mn} + h_{mn}, \qquad h_{mn} \equiv -F(x^+, x^i)k_mk_n,$$

and solves Einstein's equations even when h_{mn} is not small... As a consequence, many properties of flat Minkowski spaces can be uplifted to pp-wave backgrounds with only minor modifications...

• Pp-wave spacetimes are not globally hyperbolic (as opposed to Minkowski spaces)... no global Cauchy hypersurface...

Penrose (1965)

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

Penrose limit: d + 1 dimensional Brinkmann metric descends from any metric by blowing up spacetime around null geodesics ("zooming in" to them)... New solutions of Einstein's equations can be constructed from known ones...

Penrose (1976) & Güven (2000)

• The d + 1 dimensional Brinkmann metric can be written in a form which resembles linearized gravity:

$$g_{mn} = \eta_{mn} + h_{mn}, \qquad h_{mn} \equiv -F(x^+, x^i)k_mk_n,$$

and solves Einstein's equations even when h_{mn} is not small... As a consequence, many properties of flat Minkowski spaces can be uplifted to pp-wave backgrounds with only minor modifications...

• Pp-wave spacetimes are not globally hyperbolic (as opposed to Minkowski spaces)... no global Cauchy hypersurface...

Penrose (1965)

Pp and plane wave spacetimes cannot contain black holes & event horizons (Hubeny-Rangamani, 2002)...

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

Penrose limit: d + 1 dimensional Brinkmann metric descends from any metric by blowing up spacetime around null geodesics ("zooming in" to them)... New solutions of Einstein's equations can be constructed from known ones...

Penrose (1976) & Güven (2000)

• The d + 1 dimensional Brinkmann metric can be written in a form which resembles linearized gravity:

$$g_{mn} = \eta_{mn} + h_{mn}, \qquad h_{mn} \equiv -F(x^+, x^i)k_mk_n,$$

and solves Einstein's equations even when h_{mn} is not small... As a consequence, many properties of flat Minkowski spaces can be uplifted to pp-wave backgrounds with only minor modifications...

• Pp-wave spacetimes are not globally hyperbolic (as opposed to Minkowski spaces)... no global Cauchy hypersurface...

Penrose (1965)

• Pp and plane wave spacetimes cannot contain black holes & event horizons (Hubeny-Rangamani, 2002)... although the opposite is always true: any singular spacetime has a (singularity-free) plane-wave limit...

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

Penrose limit: d + 1 dimensional Brinkmann metric descends from any metric by blowing up spacetime around null geodesics ("zooming in" to them)... New solutions of Einstein's equations can be constructed from known ones...

Penrose (1976) & Güven (2000)

• The d + 1 dimensional Brinkmann metric can be written in a form which resembles linearized gravity:

$$g_{mn} = \eta_{mn} + h_{mn}, \qquad h_{mn} \equiv -F(x^+, x')k_mk_n,$$

and solves Einstein's equations even when h_{mn} is not small... As a consequence, many properties of flat Minkowski spaces can be uplifted to pp-wave backgrounds with only minor modifications...

Pp-wave spacetimes are not globally hyperbolic (as opposed to Minkowski spaces)... no global Cauchy hypersurface...

Penrose (1965)

- Pp and plane wave spacetimes cannot contain black holes & event horizons (Hubeny-Rangamani, 2002)... although the opposite is always true: any singular spacetime has a (singularity-free) plane-wave limit...
- Vanishing scalar invariant (VSI): spacetimes which admit covariantly constant null Killing vectors k have all their scalar invariants (constructed from the Riemann tensor and its covariant derivatives) vanish...

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

Penrose limit: d + 1 dimensional Brinkmann metric descends from any metric by blowing up spacetime around null geodesics ("zooming in" to them)... New solutions of Einstein's equations can be constructed from known ones...

Penrose (1976) & Güven (2000)

• The d + 1 dimensional Brinkmann metric can be written in a form which resembles linearized gravity:

$$g_{mn} = \eta_{mn} + h_{mn}, \qquad h_{mn} \equiv -F(x^+, x^i)k_mk_n,$$

and solves Einstein's equations even when h_{mn} is not small... As a consequence, many properties of flat Minkowski spaces can be uplifted to pp-wave backgrounds with only minor modifications...

Pp-wave spacetimes are not globally hyperbolic (as opposed to Minkowski spaces)... no global Cauchy hypersurface...

Penrose (1965)

- Pp and plane wave spacetimes cannot contain black holes & event horizons (Hubeny-Rangamani, 2002)... although the opposite is always true: any singular spacetime has a (singularity-free) plane-wave limit...
- Vanishing scalar invariant (VSI): spacetimes which admit covariantly constant null Killing vectors k have all their scalar invariants (constructed from the Riemann tensor and its covariant derivatives) vanish...
- Brinkmann spacetimes are α' -exact solutions of supergravity/string theory (with or without flux terms)...

Amati-Klimčík (1988), Horowitz-Steif (1990)

イロト イロト イヨト イヨト ヨー わへの

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

• Penrose-Güven limits preserve susy... maximally susy backgrounds of 11d/IIB sugra $AdS_{4/5/7} \times S^{7/5/4}$, give rise to two maximally susy homogeneous plane-wave solutions in 10 & 11d...

Figueroa-O'Farrill & Papadopoulos (2003)

These backgrounds are known as Hpp-waves (Cahen-Wallach plane-waves with homogeneous fluxes)... along with the 3 AdS solutions & flat space in 10 & 11d, these are the 7 maximally susy backgrounds in 10 & 11d...

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

• Penrose-Güven limits preserve susy... maximally susy backgrounds of 11d/IIB sugra $AdS_{4/5/7} \times S^{7/5/4}$, give rise to two maximally susy homogeneous plane-wave solutions in 10 & 11d...

Figueroa-O'Farrill & Papadopoulos (2003)

These backgrounds are known as Hpp-waves (Cahen-Wallach plane-waves with homogeneous fluxes)... along with the 3 AdS solutions & flat space in 10 & 11d, these are the 7 maximally susy backgrounds in 10 & 11d...

In 11d, the maximally susy homogeneous plane-wave background is part of the Kowalski-Glikman (KG) solution:

$$ds^{2} = -2dx^{+}dx^{-} - \left[\frac{\mu^{2}}{9}\sum_{i=1}^{3}x_{i}x_{i} + \frac{\mu^{2}}{36}\sum_{j=1}^{6}y_{j}y_{j}\right]dx^{+}dx^{+} + \sum_{i=1}^{3}dx_{i}dx_{i} + \sum_{j=1}^{6}dy_{j}dy_{j}, \qquad F_{123+} = \mu.$$

Kowalski-Glikman (1984)

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

 Penrose-Güven limits preserve susy... maximally susy backgrounds of 11d/IIB sugra AdS_{4/5/7} × S^{7/5/4}, give rise to two maximally susy homogeneous plane-wave solutions in 10 & 11d...

Figueroa-O'Farrill & Papadopoulos (2003)

These backgrounds are known as Hpp-waves (Cahen-Wallach plane-waves with homogeneous fluxes)... along with the 3 AdS solutions & flat space in 10 & 11d, these are the 7 maximally susy backgrounds in 10 & 11d...

In 11d, the maximally susy homogeneous plane-wave background is part of the Kowalski-Glikman (KG) solution:

$$ds^{2} = -2dx^{+}dx^{-} - \left[\frac{\mu^{2}}{9}\sum_{i=1}^{3}x_{i}x_{i} + \frac{\mu^{2}}{36}\sum_{j=1}^{6}y_{j}y_{j}\right]dx^{+}dx^{+} + \sum_{i=1}^{3}dx_{i}dx_{i} + \sum_{j=1}^{6}dy_{j}dy_{j}, \qquad F_{123+} = \mu.$$

Kowalski-Glikman (1984)

• IIB superstring σ model exactly solvable & quantizable on the 10-dimensional maximally susy background...

Metsaev (2001), Metsaev-Tseytlin (2002)

Pp and plane-wave spacetimes stand out thanks to a set of remarkable properties... Here's an outline of them:

• Penrose-Güven limits preserve susy... maximally susy backgrounds of 11d/IIB sugra $AdS_{4/5/7} \times S^{7/5/4}$, give rise to two maximally susy homogeneous plane-wave solutions in 10 & 11d...

Figueroa-O'Farrill & Papadopoulos (2003)

These backgrounds are known as Hpp-waves (Cahen-Wallach plane-waves with homogeneous fluxes)... along with the 3 AdS solutions & flat space in 10 & 11d, these are the 7 maximally susy backgrounds in 10 & 11d...

In 11d, the maximally susy homogeneous plane-wave background is part of the Kowalski-Glikman (KG) solution:

$$ds^{2} = -2dx^{+}dx^{-} - \left[\frac{\mu^{2}}{9}\sum_{i=1}^{3}x_{i}x_{i} + \frac{\mu^{2}}{36}\sum_{j=1}^{6}y_{j}y_{j}\right]dx^{+}dx^{+} + \sum_{i=1}^{3}dx_{i}dx_{i} + \sum_{j=1}^{6}dy_{j}dy_{j}, \qquad F_{123+} = \mu.$$

Kowalski-Glikman (1984)

IIB superstring σ model exactly solvable & quantizable on the 10-dimensional maximally susy background...

Metsaev (2001), Metsaev-Tseytlin (2002)

• BMN sector of AdS₅/CFT₄: Penrose limit of IIB string theory on AdS₅ × S⁵ \leftrightarrow BMN limit of \mathcal{N} = 4 SYM...

Berenstein-Maldacena-Nastase (2002)

M-theory on a plane wave

• The matrix model of Berenstein, Maldacena and Nastase (BMN),

$$H = H_0 + \frac{R}{2} \cdot \operatorname{Tr}\left[\sum_{i=1}^3 \frac{m^2}{9} \mathbf{X}_i^2 + \sum_{j=4}^9 \frac{m^2}{36} \mathbf{X}_j^2 + \sum_{i,j,k=1}^3 \frac{2m}{3} i\epsilon_{ijk} \mathbf{X}_i \mathbf{X}_j \mathbf{X}_k - \frac{m}{2} i \Psi^T \gamma_{123} \Psi\right],$$

Berenstein-Maldacena-Nastase (2002)

describes M-theory on the 11d maximally supersymmetric KG background (homogeneous plane-wave background)...

M-theory on a plane wave

• The matrix model of Berenstein, Maldacena and Nastase (BMN),

$$H = H_0 + \frac{R}{2} \cdot \operatorname{Tr}\left[\sum_{i=1}^3 \frac{m^2}{9} \mathbf{X}_i^2 + \sum_{j=4}^9 \frac{m^2}{36} \mathbf{X}_j^2 + \sum_{i,j,k=1}^3 \frac{2m}{3} i\epsilon_{ijk} \mathbf{X}_i \mathbf{X}_j \mathbf{X}_k - \frac{m}{2} i \Psi^T \gamma_{123} \Psi\right],$$

Berenstein-Maldacena-Nastase (2002)

describes M-theory on the 11d maximally supersymmetric KG background (homogeneous plane-wave background)....

• H_0 is the Hamiltonian of the BFSS matrix model which describes M-theory in flat ($\mu = 0$) space...

$$H_0 = \frac{R}{2} \cdot \operatorname{Tr} \left[\dot{\mathbf{X}}^2 - \frac{1}{2} \left[\mathbf{X}_A, \mathbf{X}_B \right]^2 - \Psi^T \gamma_A [\mathbf{X}_A, \Psi] \right], \quad A, B = 1, \dots, 9,$$

Banks-Fischler-Shenker-Susskind (1996)

where the vectors X_A and 16d Majorana spinor Ψ are $N \times N$ Hermitian matrices... γ_A are the 9d (16 × 16) Euclidean Dirac matrices, R is the DLCQ compactification radius, and $m \equiv \mu/R$...

M-theory on a plane wave

• The matrix model of Berenstein, Maldacena and Nastase (BMN),

$$H = H_0 + \frac{R}{2} \cdot \operatorname{Tr}\left[\sum_{i=1}^3 \frac{m^2}{9} \mathbf{X}_i^2 + \sum_{j=4}^9 \frac{m^2}{36} \mathbf{X}_j^2 + \sum_{i,j,k=1}^3 \frac{2m}{3} i\epsilon_{ijk} \mathbf{X}_i \mathbf{X}_j \mathbf{X}_k - \frac{m}{2} i \Psi^T \gamma_{123} \Psi\right],$$

Berenstein-Maldacena-Nastase (2002)

describes M-theory on the 11d maximally supersymmetric KG background (homogeneous plane-wave background)...

• H_0 is the Hamiltonian of the BFSS matrix model which describes M-theory in flat ($\mu = 0$) space...

$$H_{0} = \frac{R}{2} \cdot \operatorname{Tr}\left[\dot{\mathbf{X}}^{2} - \frac{1}{2}\left[\mathbf{X}_{A}, \mathbf{X}_{B}\right]^{2} - \Psi^{T} \gamma_{A}[\mathbf{X}_{A}, \Psi]\right], \quad A, B = 1, \dots, 9,$$

Banks-Fischler-Shenker-Susskind (1996)

where the vectors X_A and 16d Majorana spinor Ψ are $N \times N$ Hermitian matrices... γ_A are the 9d (16 × 16) Euclidean Dirac matrices, R is the DLCQ compactification radius, and $m \equiv \mu/R$...

The BMN matrix model constitutes a deformation of the BMN matrix model by mass terms and a Myers term...

• The mass terms of the BMN matrix model lift the flat directions of BFSS, making the supermembrane spectrum discrete... On the other hand, the Myers term allows for static fuzzy sphere solutions:

 $X_i = r \cdot J_i, \quad i = 1, 2, 3$ & $X_j = 0, \quad j = 4, \dots, 9,$

where the matrices J_i furnish a N-dimensional representation of $\mathfrak{su}(2)$. The radii,

$$r=0,$$
 $r=\frac{\mu}{3},$ $r=\frac{\mu}{6},$

• The mass terms of the BMN matrix model lift the flat directions of BFSS, making the supermembrane spectrum discrete... On the other hand, the Myers term allows for static fuzzy sphere solutions:

 $X_i = r \cdot J_i, \quad i = 1, 2, 3$ & $X_j = 0, \quad j = 4, \dots, 9,$

where the matrices J_i furnish a N-dimensional representation of $\mathfrak{su}(2)$. The radii,

$$r=0,$$
 $r=\frac{\mu}{3},$ $r=\frac{\mu}{6},$

correspond to the max susy vacuum $X_i = 0$, a 1/2-BPS solution and, an unstable, non-susy, positive-energy solution...

• The BMN matrix model describes the discrete light-cone quantization (DLCQ) of M-theory on the KG background...

• The mass terms of the BMN matrix model lift the flat directions of BFSS, making the supermembrane spectrum discrete... On the other hand, the Myers term allows for static fuzzy sphere solutions:

 $X_i = r \cdot J_i, \quad i = 1, 2, 3$ & $X_j = 0, \quad j = 4, \dots, 9,$

where the matrices J_i furnish a N-dimensional representation of $\mathfrak{su}(2)$. The radii,

$$r=0,$$
 $r=\frac{\mu}{3},$ $r=\frac{\mu}{6},$

- The BMN matrix model describes the discrete light-cone quantization (DLCQ) of M-theory on the KG background...
- As shown by Dasgupta, Sheikh-Jabbari, Van Raamsdonk (2002), the BMN matrix model can be derived by regularizing the light-cone (super)membrane in the 11d maximally susy KG background...

• The mass terms of the BMN matrix model lift the flat directions of BFSS, making the supermembrane spectrum discrete... On the other hand, the Myers term allows for static fuzzy sphere solutions:

 $X_i = r \cdot J_i, \quad i = 1, 2, 3$ & $X_j = 0, \quad j = 4, \dots, 9,$

where the matrices J_i furnish a N-dimensional representation of $\mathfrak{su}(2)$. The radii,

$$r=0,$$
 $r=\frac{\mu}{3},$ $r=\frac{\mu}{6},$

- The BMN matrix model describes the discrete light-cone quantization (DLCQ) of M-theory on the KG background...
- As shown by Dasgupta, Sheikh-Jabbari, Van Raamsdonk (2002), the BMN matrix model can be derived by regularizing the light-cone (super)membrane in the 11d maximally susy KG background...
- Equivalently, the light-cone (super)membrane on the maximally susy plane-wave background can be seen as the continuum (matrix dimensionality $N \rightarrow \infty$) limit of the BMN matrix model...

• The mass terms of the BMN matrix model lift the flat directions of BFSS, making the supermembrane spectrum discrete... On the other hand, the Myers term allows for static fuzzy sphere solutions:

 $X_i = r \cdot J_i, \quad i = 1, 2, 3$ & $X_j = 0, \quad j = 4, \dots, 9,$

where the matrices J_i furnish a N-dimensional representation of $\mathfrak{su}(2)$. The radii,

$$r = 0, \qquad r = \frac{\mu}{3}, \qquad r = \frac{\mu}{6},$$

- The BMN matrix model describes the discrete light-cone quantization (DLCQ) of M-theory on the KG background...
- As shown by Dasgupta, Sheikh-Jabbari, Van Raamsdonk (2002), the BMN matrix model can be derived by regularizing the light-cone (super)membrane in the 11d maximally susy KG background...
- Equivalently, the light-cone (super)membrane on the maximally susy plane-wave background can be seen as the continuum (matrix dimensionality $N \rightarrow \infty$) limit of the BMN matrix model...
- In the following we are going to work exclusively with the classical bosonic membrane on the 11-dimensional maximally susy plane-wave background...

• The mass terms of the BMN matrix model lift the flat directions of BFSS, making the supermembrane spectrum discrete... On the other hand, the Myers term allows for static fuzzy sphere solutions:

 $X_i = r \cdot J_i, \quad i = 1, 2, 3$ & $X_j = 0, \quad j = 4, \dots, 9,$

where the matrices J_i furnish a N-dimensional representation of $\mathfrak{su}(2)$. The radii,

$$r=0,$$
 $r=\frac{\mu}{3},$ $r=\frac{\mu}{6},$

- The BMN matrix model describes the discrete light-cone quantization (DLCQ) of M-theory on the KG background...
- As shown by Dasgupta, Sheikh-Jabbari, Van Raamsdonk (2002), the BMN matrix model can be derived by regularizing the light-cone (super)membrane in the 11d maximally susy KG background...
- Equivalently, the light-cone (super)membrane on the maximally susy plane-wave background can be seen as the continuum (matrix dimensionality N → ∞) limit of the BMN matrix model...
- In the following we are going to work exclusively with the classical bosonic membrane on the 11-dimensional maximally susy plane-wave background...
- We are going to construct solutions of spinning membranes, based on some tools and techniques that were introduced for flat space...

• The mass terms of the BMN matrix model lift the flat directions of BFSS, making the supermembrane spectrum discrete... On the other hand, the Myers term allows for static fuzzy sphere solutions:

 $X_i = r \cdot J_i, \quad i = 1, 2, 3$ & $X_j = 0, \quad j = 4, \dots, 9,$

where the matrices J_i furnish a N-dimensional representation of $\mathfrak{su}(2)$. The radii,

$$r=0,$$
 $r=\frac{\mu}{3},$ $r=\frac{\mu}{6},$

correspond to the max susy vacuum $X_i = 0$, a 1/2-BPS solution and, an unstable, non-susy, positive-energy solution...

- The BMN matrix model describes the discrete light-cone quantization (DLCQ) of M-theory on the KG background...
- As shown by Dasgupta, Sheikh-Jabbari, Van Raamsdonk (2002), the BMN matrix model can be derived by regularizing the light-cone (super)membrane in the 11d maximally susy KG background...
- Equivalently, the light-cone (super)membrane on the maximally susy plane-wave background can be seen as the continuum (matrix dimensionality N → ∞) limit of the BMN matrix model...
- In the following we are going to work exclusively with the classical bosonic membrane on the 11-dimensional maximally susy plane-wave background...
- We are going to construct solutions of spinning membranes, based on some tools and techniques that were introduced for flat space...
- Let us first briefly review the corresponding membrane action...

イロト イロト イヨト イヨト ヨー わへの

Subsection 3

Membranes in the light-cone gauge

Bosonic membrane in a curved background

• Bosonic membranes in curved backgrounds are described by the Dirac-Nambu-Goto (DNG) action:

$$S_{\text{DNG}} = -T \int d\tau d^2 \sigma \left\{ \sqrt{-h} + \dot{X}^m \partial_1 X^n \partial_2 X^r A_{rnm} \left(X \right) \right\}, \qquad T \equiv \frac{1}{(2\pi)^2 \ell_{11}^3},$$

where (m, n, r, s = 0, ..., 10),

$$h_{ij} \equiv G_{mn}\partial_i X^m \partial_j X^n$$
 (induced metric) $h \equiv \det h_{ij}$ & $F_{mnrs} = 4\partial_{[m}A_{nrs]}$ (field strength)

and Anrs is the (antisymmetric) 3-form field of 11-dimensional supergravity...

The light-cone gauge

• In the light-cone gauge, we write:

$$X^{\pm} = rac{1}{\sqrt{2}} \left(X^0 \pm X^{10}
ight)$$
 & $X^+ = au.$

Goldstone-Hoppe (1982)

• The light-cone Hamiltonian is then written as follows ($G_{--} = G_{a-} = 0$):

$$H = T \int d^2 \sigma \left\{ \frac{1}{2} \frac{G_{+-}}{P_{-} - C_{-}} \left[\left(P_a - C_a - \frac{P_{-} - C_{-}}{G_{+-}} G_{a+} \right)^2 + \frac{1}{2} G_{ab} G_{cd} \{ X^a, X^c \} \{ X^b, X^d \} \right] - \frac{1}{2} \frac{P_{-} - C_{-}}{G_{+-}} G_{++} - C_{+} + \frac{1}{P_{-} - C_{-}} \left[C_{-} C_{+-} - \{ X^a, X^b \} P_a C_{+-b} \right] \right\},$$

de Wit-Peeters-Plefka (1998)

where (a, b, c, d = 1, ..., 9),

$$C_{\pm} \equiv C_{\pm ab} - \partial_1 X^a \partial_2 X^b, \quad C_{+-} \equiv -C_{+-a} \{X^-, X^a\}, \quad C_a \equiv -\left(C_{-ab} \{X^b, X^-\} + C_{abc} \partial_1 X^b \partial_2 X^c\right).$$
Poisson brackets

The Poisson bracket is defined as:

$$\{f,g\} \equiv \frac{\epsilon_{rs}}{\sqrt{w(\sigma)}} \partial_r f \, \partial_s g = \frac{1}{\sqrt{w(\sigma)}} \left(\partial_1 f \, \partial_2 g - \partial_2 f \, \partial_1 g\right),$$

where $d^{2}\sigma = \sqrt{w(\sigma)} \ d\sigma_{1} \ d\sigma_{2}$. In a flat space-sheet, $w(\sigma) = 1$.

Section 2

Spherical Euler-top membranes in flat backgrounds

M. Axenides, E. Floratos, L. Perivolaropoulos Metastability of spherical membranes in supermembrane and matrix theory JHEP 11 (2000) 020 [arXiv:hep-th/0007198] M. Axenides, E. Floratos Euler-top dynamics of Nambu-Goto p-branes JHEP **03** (2007) 093 [arXiv:hep-th/0608017]

Light-cone gauge in flat space

In a flat background

 $G_{+-} = -1, \quad G_{ab} = \delta_{ab}, \qquad G_{++} = G_{--} = G_{a\pm} = 0, \qquad C_{\pm} = C_{+-} = C_a = 0,$

therefore the light-cone Hamiltonian becomes $(P_{-} = -1)$:

$$H = \frac{T}{2} \int d^2 \sigma \left[P^2 + \frac{1}{2} \{ X^i, X^j \}^2 \right].$$

The corresponding equations of motion and the Gauss law constraint become:

$$\ddot{X}^{i} = \{\{X^{i}, X^{j}\}, X^{j}\}$$
 & $\sum_{i=1}^{9}\{\dot{X}^{i}, X^{i}\} = 0.$

Euler-top membranes in flat space

• Consider the ansatz:

$$X^{i}=R^{ij}\left(au
ight)X_{0}^{j}\left(\sigma
ight),\qquad R\equiv\exp\left(\Omega\, au
ight),\qquad \Omega^{T}=-\Omega$$

If we define the angular momentum and moment of inertia matrices of the membrane as

$$\mathbf{I}^{ij} = \mathcal{T} \int d^2 \sigma \, X^i X^j \qquad \& \qquad \mathbf{L}^{ij} = \mathcal{T} \int d^2 \sigma \left(\dot{X}^i X^j - \dot{X}^j X^i \right),$$

we can prove that the energy of the membrane is given by

$$E = -\frac{3}{4} \cdot \frac{\text{Tr} \left[\Omega \cdot L\right]^2}{2 \text{ Tr} \left[\Omega^2 \cdot I\right]}$$

Axenides-Floratos (2006)

which is the generalization of the familiar from point-particle mechanics Euler-top Hamiltonian:

$$E = rac{\ell_x^2}{2I_x} + rac{\ell_y^2}{2I_y} + rac{\ell_z^2}{2I_z}$$

< □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □

• Consider the following spherical configuration:

$$\begin{split} X_i &\equiv x_i \left(\tau \right) \cdot e_1, & i = 1, 2, \dots, q_1 \\ Y_j &\equiv X_{q_1+j} = y_j \left(\tau \right) \cdot e_2, & j = 1, 2, \dots, q_2, & \& \quad q_1 + q_2 + q_3 = 9 \\ Z_j &\equiv X_{q_2+k} = z_k \left(\tau \right) \cdot e_3, & k = 1, 2, \dots, q_3, \end{split}$$

Collins-Tucker (1976)

that breaks the manifest $\mathfrak{so}(9)$ symmetry of the action to $\mathfrak{so}(q_1) \times \mathfrak{so}(q_2) \times \mathfrak{so}(q_3)$. We have defined:

$$egin{aligned} &(e_1,e_2,e_3)=(\cos\phi\sin heta,\sin\phi\sin heta,\cos heta), &\phi\in[0,2\pi), & heta\in[0,\pi]\ &\{e_i,e_j\}=\epsilon_{ijk}\,e_k, &\int e_i\,e_j\,d^2\sigma=rac{4\pi}{3}\,\delta_{ij} \end{aligned}$$

and the membrane area element is given by:

$$d^2\sigma = d\sigma_1 d\sigma_2 = \sin\theta d\phi d\theta$$
 & $\sqrt{w(\theta)} = \sin\theta.$

<ロト < 団ト < 巨ト < 巨ト < 巨ト 三 のQで 17 / 53

• Here's the energy of the bubble:

$$E = \frac{2\pi T}{3} \left[\dot{\mathbf{x}}^2 + \dot{\mathbf{y}}^2 + \dot{\mathbf{z}}^2 + \mathbf{x}^2 \mathbf{y}^2 + \mathbf{y}^2 \mathbf{z}^2 + \mathbf{z}^2 \mathbf{x}^2 \right].$$

• The corresponding equations of motion are:

$$\ddot{x}_i + (\mathbf{y}^2 + \mathbf{z}^2) x_i = 0, \qquad \ddot{y}_j + (\mathbf{z}^2 + \mathbf{x}^2) y_j = 0, \qquad \ddot{z}_k + (\mathbf{y}^2 + \mathbf{x}^2) z_k = 0,$$

while the Gauss law constraint

$$\sum_{i=1}^{q_1} \{ \dot{x}^i, x^i \} + \sum_{j=1}^{q_2} \{ \dot{y}^j, y^j \} + \sum_{k=1}^{q_3} \{ \dot{z}^k, z^k \} = 0,$$

is automatically satisfied by this ansatz.

• Let us switch to the notation:

$$r_{x}^{2} \equiv \mathbf{x}^{2} = \sum_{i=1}^{q_{1}} x_{i}x_{i}, \qquad r_{y}^{2} \equiv \mathbf{y}^{2} = \sum_{j=1}^{q_{2}} y_{j}y_{j}, \qquad r_{z}^{2} \equiv \mathbf{z}^{2} = \sum_{k=1}^{q_{3}} z_{k}z_{k}$$

$$(\ell_{x})_{ij} \equiv \dot{x}_{i}x_{j} - x_{i}\dot{x}_{j}\Big|_{\mathfrak{so}(q_{1})}, \qquad (\ell_{y})_{ij} \equiv \dot{y}_{i}y_{j} - y_{i}\dot{y}_{j}\Big|_{\mathfrak{so}(q_{2})}, \qquad (\ell_{z})_{ij} \equiv \dot{z}_{i}z_{j} - z_{i}\dot{z}_{j}\Big|_{\mathfrak{so}(q_{3})} \qquad \text{conserved}$$

$$\dot{\mathbf{x}}^{2} \equiv \sum_{i=1}^{q_{1}} \dot{x}_{i}\dot{x}_{i} = \dot{r}_{x}^{2} + \frac{\ell_{x}^{2}}{r_{x}^{2}}, \qquad \dot{\mathbf{y}}^{2} \equiv \sum_{j=1}^{q_{2}} \dot{y}_{j}\dot{y}_{j} = \dot{r}_{y}^{2} + \frac{\ell_{y}^{2}}{r_{y}^{2}}, \qquad \dot{\mathbf{z}}^{2} \equiv \sum_{k=1}^{q_{3}} \dot{z}_{k}\dot{z}_{k} = \dot{r}_{z}^{2} + \frac{\ell_{z}^{2}}{r_{z}^{2}},$$

which allows to write the energy of the membrane as follows:

$$E = \frac{2\pi T}{3} \left(E_{\rm kin} + V_{\rm eff} \right), \qquad E_{\rm kin} \equiv \dot{r}_x^2 + \dot{r}_y^2 + \dot{r}_z^2 \quad \& \quad V_{\rm eff} \equiv \frac{\ell_x^2}{r_x^2} + \frac{\ell_y^2}{r_y^2} + \frac{\ell_z^2}{r_z^2} + r_x^2 r_y^2 + r_y^2 r_z^2 + r_z^2 r_x^2,$$

where

$$\ell_x^2 = \frac{1}{2} (\ell_x)_{ij} (\ell_x)_{ij}, \qquad \ell_y^2 = \frac{1}{2} (\ell_y)_{ij} (\ell_y)_{ij}, \qquad \ell_z^2 = \frac{1}{2} (\ell_z)_{ij} (\ell_z)_{ij}.$$

Spherical Euler tops

• As shown by Axenides-Floratos (2006), the radii $r_x = x_0^2$, $r_y = y_0^2$, $r_z = z_0^2$ of the Euler top solutions

$$\mathbf{x}(\tau) = e^{\Omega_x \tau} \cdot \mathbf{x}_0, \qquad \mathbf{y}(\tau) = e^{\Omega_y \tau} \cdot \mathbf{y}_0, \qquad \mathbf{z}(\tau) = e^{\Omega_z \tau} \cdot \mathbf{z}_0,$$

can be determined for all antisymmetric matrices Ω_x , Ω_y , Ω_z in terms of the corresponding angular momenta ℓ_x , ℓ_y , ℓ_z , by minimizing the effective potential:

$$V_{\text{eff}} \equiv \frac{\ell_x^2}{r_x^2} + \frac{\ell_y^2}{r_y^2} + \frac{\ell_z^2}{r_z^2} + r_x^2 r_y^2 + r_y^2 r_z^2 + r_z^2 r_x^2 ,$$

i.e. by solving

$$\frac{dV_{\text{eff}}}{dr_{x}} = -\frac{2\ell_{x}^{2}}{r_{x}^{3}} + 2r_{x}\left(r_{y}^{2} + r_{z}^{2}\right) = \frac{dV_{\text{eff}}}{dr_{y}} = -\frac{2\ell_{y}^{2}}{r_{y}^{3}} + 2r_{y}\left(r_{z}^{2} + r_{x}^{2}\right) = \frac{dV_{\text{eff}}}{dr_{z}} = -\frac{2\ell_{z}^{2}}{r_{z}^{3}} + 2r_{z}\left(r_{x}^{2} + r_{y}^{2}\right) = 0.$$

• Equivalently we can plug the above ansatz into the equations of motion in order to determine the relation between the radii r_x , r_y , r_z and the components of the matrices Ω_x , Ω_y , Ω_z .

Symmetric & axially symmetric Euler spheres

• For a single radius $r = r_x = r_y = r_z$, $\ell = \ell_x = \ell_y = \ell_z$ the effective potential becomes:

$$V_{
m eff}\equiv rac{3\ell}{r^2}+3r^4,$$

finding

$$r_{(\min)} = rac{\ell^{1/3}}{2^{1/6}}, \qquad V_{eff(\min)} = rac{9\ell^{4/3}}{4^{1/3}}.$$

• The axially symmetric (two-radii) $r_{\alpha} = r_x = r_y$, $\ell_{\alpha} = \ell_x = \ell_y$ effective potential is:

$$V_{\rm eff} \equiv \frac{2\ell_\alpha^2}{r_\alpha^2} + \frac{\ell_z^2}{r_z^2} + r_\alpha^4 + 2r_\alpha^2 r_z^2$$

with

$$r_{\alpha(\min)}^{2} = \frac{2\ell_{\alpha}^{4/3}}{\left(\ell_{z} + \sqrt{\ell_{z}^{2} + 8\ell_{\alpha}^{2}}\right)^{2/3}}, \quad r_{z(\min)}^{2} = \frac{\ell_{z}}{2\ell_{\alpha}^{2/3}} \left(\ell_{z} + \sqrt{\ell_{z}^{2} + 8\ell^{2}}\right)^{1/3}$$

$$V_{\text{eff}(\min)} = \frac{6\ell_{\alpha}^{2/3}}{\left(\ell_{z} + \sqrt{\ell_{z}^{2} + 8\ell_{\alpha}^{2}}\right)^{4/3}} \left[\ell_{z} \left(\ell_{z} + \sqrt{\ell_{z}^{2} + 8\ell_{\alpha}^{2}}\right) + 2\ell_{\alpha}^{2}\right].$$

21 / 53

Section 3

Spherical dielectric tops in plane-wave backgrounds

M. Axenides, E. Floratos, D. Katsinis, GL M-theory as a dynamical system generator [arXiv:2007.07028]

M. Axenides, E. Floratos, GL to appear

Light-cone gauge in the plane-wave background

• In the maximally supersymmetric plane background,

$$\begin{aligned} G_{+-} &= -1, \qquad G_{ab} = \delta_{ab}, \qquad G_{++} = -\frac{\mu^2}{9} \sum_{i=1}^3 x^i x^i - \frac{\mu^2}{36} \sum_{j=1}^6 y^j y^j, \qquad G_{--} = G_{a\pm} = 0 \\ C_{-} &= C_{+-} = C_a = 0, \qquad C_{+} = \frac{\mu}{3} \epsilon_{ijk} \partial_1 x^i \partial_2 x^j x^k, \end{aligned}$$

the light-cone Hamiltonian becomes (for $P_{-} = -1$):

$$H = \frac{T}{2} \int d^2 \sigma \left[\pi_i^2 + \frac{1}{2} \left\{ x^i, x^j \right\}^2 + \frac{1}{2} \left\{ y^i, y^j \right\}^2 + \left\{ x^i, y^j \right\}^2 + \frac{\mu^2 x^2}{9} + \frac{\mu^2 y^2}{36} - \frac{\mu}{3} \epsilon_{ijk} \left\{ x^i, x^j \right\} x^k \right].$$

Light-cone gauge in the plane-wave background

• In the maximally supersymmetric plane background,

$$\begin{aligned} G_{+-} &= -1, \qquad G_{ab} = \delta_{ab}, \qquad G_{++} = -\frac{\mu^2}{9} \sum_{i=1}^3 x^i x^i - \frac{\mu^2}{36} \sum_{j=1}^6 y^j y^j, \qquad G_{--} = G_{a\pm} = 0 \\ C_{-} &= C_{+-} = C_a = 0, \qquad C_{+} = \frac{\mu}{3} \epsilon_{ijk} \partial_1 x^i \partial_2 x^j x^k, \end{aligned}$$

which can also be expressed as a sum of squares:

$$H = \frac{T}{2} \int d^2 \sigma \left[\pi^2 + \left(\frac{\mu}{3} x_i - \frac{1}{2} \epsilon_{ijk} \{ x_j, x_k \} \right)^2 + \frac{1}{2} \{ y_i, y_j \}^2 + \frac{\mu^2}{36} y_j y_j + \{ x_i, y_j \}^2 \right].$$

Light-cone gauge in the plane-wave background

• In the maximally supersymmetric plane background,

$$\begin{aligned} G_{+-} &= -1, \qquad G_{ab} = \delta_{ab}, \qquad G_{++} = -\frac{\mu^2}{9} \sum_{i=1}^3 x^i x^i - \frac{\mu^2}{36} \sum_{j=1}^6 y^j y^j, \qquad G_{--} = G_{a\pm} = 0 \\ C_{-} &= C_{+-} = C_a = 0, \qquad C_{+} = \frac{\mu}{3} \epsilon_{ijk} \partial_1 x^i \partial_2 x^j x^k, \end{aligned}$$

which can also be expressed as a sum of squares:

$$H = \frac{T}{2} \int d^2 \sigma \left[\pi^2 + \left(\frac{\mu}{3} x_i - \frac{1}{2} \epsilon_{ijk} \{ x_j, x_k \} \right)^2 + \frac{1}{2} \{ y_i, y_j \}^2 + \frac{\mu^2}{36} y_j y_j + \{ x_i, y_j \}^2 \right].$$

• The corresponding equations of motion and the Gauss law constraint read:

$$\ddot{x}_{i} = \left\{ \left\{ x_{i}, x_{j} \right\}, x_{j} \right\} + \left\{ \left\{ x_{i}, y_{j} \right\}, y_{j} \right\} - \frac{\mu^{2}}{9} x_{i} + \frac{\mu}{2} \epsilon_{ijk} \left\{ x_{j}, x_{k} \right\}, \qquad \sum_{i=1}^{3} \left\{ \dot{x}^{i}, x^{i} \right\} + \sum_{j=1}^{6} \left\{ \dot{y}^{j}, y^{j} \right\} = 0$$

$$\ddot{y}_{i} = \left\{ \left\{ y_{i}, y_{j} \right\}, y_{j} \right\} + \left\{ \left\{ y_{i}, x_{j} \right\}, x_{j} \right\} - \frac{\mu^{2}}{36} y_{i}.$$

23 / 53

Here's the generalization of the flat spherical ansatz to the maximally supersymmetric plane-wave background:

$$x_i \equiv x_{1i} = \tilde{x}_{1i}(\tau) e_1(\sigma), \quad i = 1, \dots, q_1, \qquad y_i \equiv y_{1i} = \tilde{y}_{1i}(\tau) e_1(\sigma), \quad i = 1, \dots, s_1$$

$$x_{q_{1}+j} \equiv x_{2j} = \tilde{x}_{2j}(\tau) e_{2}(\sigma), \quad j = 1, \dots, q_{2}, \qquad \qquad y_{s_{1}+j} \equiv y_{2j} = \tilde{y}_{2j}(\tau) e_{2}(\sigma), \quad j = 1, \dots, s_{2}$$

$$x_{q_{1}+q_{2}+k} \equiv x_{3k} = \tilde{x}_{3k} (\tau) e_{3} (\sigma), \quad k = 1, \dots, q_{3}, \qquad y_{s_{1}+s_{2}+k} \equiv y_{3k} = \tilde{y}_{3k} (\tau) e_{3} (\sigma), \quad k = 1, \dots, s_{3k} = 0, \dots$$

where

$$q_1 + q_2 + q_3 = 3$$
 & $s_1 + s_2 + s_3 = 6$,

and again,

$$\begin{aligned} (\mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3) &= (\cos\phi\sin\theta, \sin\phi\sin\theta, \cos\theta), \qquad \phi \in [0, 2\pi), \quad \theta \in [0, \pi] \\ \{e_i, e_j\} &= \epsilon_{ijk} e_k, \qquad \int e_i e_j d^2\sigma = \frac{4\pi}{3} \delta_{ij} \\ d^2\sigma &= d\sigma_1 d\sigma_2 = \sin\theta d\phi d\theta \qquad \& \qquad \sqrt{w(\theta)} = \sin\theta. \end{aligned}$$

Now switch to the notation:

$$r_{xx}^{2} \equiv \tilde{x}_{1}^{2} = \sum_{i=1}^{q_{1}} \tilde{x}_{1i} \tilde{x}_{1i}, \qquad r_{xy}^{2} \equiv \tilde{x}_{2}^{2} = \sum_{i=1}^{q_{2}} \tilde{x}_{2i} \tilde{x}_{2i}, \qquad r_{xz}^{2} \equiv \tilde{x}_{3}^{2} = \sum_{i=1}^{q_{3}} \tilde{x}_{3i} \tilde{x}_{3i}$$

$$(\ell_{xx})_{ij} \equiv \dot{\tilde{x}}_{1i}\tilde{x}_{1j} - \tilde{x}_{1i}\dot{\tilde{x}}_{1j}\Big|_{\mathfrak{so}(q_1)}, \quad (\ell_{xy})_{ij} \equiv \dot{\tilde{x}}_{2i}\tilde{x}_{2j} - \tilde{x}_{2i}\dot{\tilde{x}}_{2j}\Big|_{\mathfrak{so}(q_2)}, \quad (\ell_{xz})_{ij} \equiv \dot{\tilde{x}}_{3i}\tilde{x}_{3j} - \tilde{x}_{3i}\dot{\tilde{x}}_{3j}\Big|_{\mathfrak{so}(q_3)} \quad \underline{\text{conserved}}$$

$$\dot{\tilde{x}}_{1}^{2} \equiv \sum_{i=1}^{q_{1}} \dot{\tilde{x}}_{1i} \dot{\tilde{x}}_{1i} = \dot{r}_{xx}^{2} + \frac{\ell_{xx}^{2}}{r_{xx}^{2}}, \qquad \dot{\tilde{x}}_{2}^{2} \equiv \sum_{i=1}^{q_{2}} \dot{\tilde{x}}_{2i} \dot{\tilde{x}}_{2i} = \dot{r}_{xy}^{2} + \frac{\ell_{xy}^{2}}{r_{xy}^{2}}, \qquad \dot{\tilde{x}}_{3}^{2} \equiv \sum_{i=1}^{q_{3}} \dot{\tilde{x}}_{3i} \dot{\tilde{x}}_{3i} = \dot{r}_{xz}^{2} + \frac{\ell_{xz}^{2}}{r_{xz}^{2}},$$

and similarly for the six coordinates y:

$$\begin{split} r_{yx}^{2} &\equiv \tilde{y}_{1}^{2} = \sum_{j=1}^{s_{1}} \tilde{y}_{1j} \tilde{y}_{1j}, \qquad r_{yy}^{2} \equiv \tilde{y}_{2}^{2} = \sum_{j=1}^{s_{2}} \tilde{y}_{2j} \tilde{y}_{2j}, \qquad r_{yz}^{2} \equiv \tilde{y}_{3}^{2} = \sum_{j=1}^{s_{3}} \tilde{y}_{3j} \tilde{y}_{3j} \\ (\ell_{yx})_{ij} &\equiv \dot{\tilde{y}}_{1i} \tilde{y}_{1j} - \tilde{y}_{1i} \dot{\tilde{y}}_{1j} \Big|_{\mathfrak{so}(s_{1})}, \quad (\ell_{yy})_{ij} \equiv \dot{\tilde{y}}_{2i} \tilde{y}_{2j} - \tilde{y}_{2i} \dot{\tilde{y}}_{2j} \Big|_{\mathfrak{so}(s_{2})}, \quad (\ell_{yz})_{ij} \equiv \dot{\tilde{y}}_{3i} \tilde{y}_{3j} - \tilde{y}_{3i} \dot{\tilde{y}}_{3j} \Big|_{\mathfrak{so}(s_{3})} \quad \text{conserved} \\ \dot{\tilde{y}}_{1}^{2} &\equiv \sum_{j=1}^{s_{1}} \dot{\tilde{y}}_{1j} \dot{\tilde{y}}_{1j} = \dot{r}_{yx}^{2} + \frac{\ell_{yx}^{2}}{r_{yx}^{2}}, \qquad \dot{\tilde{y}}_{2}^{2} \equiv \sum_{j=2}^{s_{2}} \dot{\tilde{y}}_{2j} \dot{\tilde{y}}_{2j} = \dot{r}_{yy}^{2} + \frac{\ell_{yy}^{2}}{r_{yy}^{2}}, \qquad \dot{\tilde{y}}_{3}^{2} \equiv \sum_{j=1}^{s_{3}} \dot{\tilde{y}}_{3j} \dot{\tilde{y}}_{3j} = \dot{r}_{yz}^{2} + \frac{\ell_{yz}^{2}}{r_{yz}^{2}}. \end{split}$$

• Here's the resulting effective potential:

$$\begin{split} V_{\text{eff}} &= \frac{2\pi T}{3} \left[\frac{\ell_{xx}^2}{r_{xx}^2} + \frac{\ell_{xy}^2}{r_{xy}^2} + \frac{\ell_{xz}^2}{r_{xz}^2} + \frac{\ell_{yx}^2}{r_{yx}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yz}^2}{r_{yz}^2} + r_{xx}^2 r_{xy}^2 + r_{xy}^2 r_{xz}^2 + r_{xz}^2 r_{xx}^2 + r_{yy}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz$$

• Here's the resulting effective potential:

$$\begin{aligned} V_{\text{eff}} &= \frac{2\pi T}{3} \left[\frac{\ell_{xx}^2}{r_{xx}^2} + \frac{\ell_{xy}^2}{r_{xy}^2} + \frac{\ell_{xz}^2}{r_{xz}^2} + \frac{\ell_{yx}^2}{r_{yx}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yz}^2}{r_{yz}^2} + r_{xx}^2 r_{xy}^2 + r_{xy}^2 r_{xz}^2 + r_{xz}^2 r_{xx}^2 + r_{yx}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{$$

- The effective potential is made up of four basic types of terms: (1) kinetic/angular momentum terms (repulsive),
 - (2) quartic interaction terms (attractive), (3) mass terms (attractive), and (4) cubic Myers terms (repulsive).

• Here's the resulting effective potential:

$$\begin{aligned} V_{\text{eff}} &= \frac{2\pi T}{3} \left[\frac{\ell_{xx}^2}{r_{xx}^2} + \frac{\ell_{xy}^2}{r_{xy}^2} + \frac{\ell_{xz}^2}{r_{xz}^2} + \frac{\ell_{yx}^2}{r_{yx}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yz}^2}{r_{yz}^2} + r_{xx}^2 r_{xy}^2 + r_{xy}^2 r_{xz}^2 + r_{xz}^2 r_{xx}^2 + r_{yx}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{$$

The effective potential is made up of four basic types of terms: • (1) kinetic/angular momentum terms (repulsive),
 • (2) quartic interaction terms (attractive), • (3) mass terms (attractive), and • (4) cubic Myers terms (repulsive).

The last two types of terms (i.e. mass terms and Myers terms) are μ -dependent and therefore absent from the flat space case ($\mu \rightarrow 0$), which was studied in Axenides-Floratos (2006)...

• Here's the resulting effective potential:

$$\begin{split} V_{\text{eff}} &= \frac{2\pi T}{3} \left[\frac{\ell_{xx}^2}{r_{xx}^2} + \frac{\ell_{xy}^2}{r_{xy}^2} + \frac{\ell_{yx}^2}{r_{xz}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yz}^2}{r_{yz}^2} + r_{xx}^2 r_{xy}^2 + r_{xy}^2 r_{xz}^2 + r_{xz}^2 r_{xy}^2 + r_{yy}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2$$

The effective potential is made up of four basic types of terms: • (1) kinetic/angular momentum terms (repulsive),
 • (2) quartic interaction terms (attractive), • (3) mass terms (attractive), and • (4) cubic Myers terms (repulsive).

The last two types of terms (i.e. mass terms and Myers terms) are μ -dependent and therefore absent from the flat space case ($\mu \rightarrow 0$), which was studied in Axenides-Floratos (2006)...

In either case ($\mu = 0$ or $\mu \neq 0$), it is the equilibration between attraction and repulsion which determines the minima of the effective potential...

• Here's the resulting effective potential:

$$\begin{aligned} V_{\text{eff}} &= \frac{2\pi T}{3} \left[\frac{\ell_{xx}^2}{r_{xx}^2} + \frac{\ell_{xy}^2}{r_{xy}^2} + \frac{\ell_{xz}^2}{r_{xz}^2} + \frac{\ell_{yx}^2}{r_{yx}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yz}^2}{r_{yz}^2} + r_{xx}^2 r_{xy}^2 + r_{xy}^2 r_{xz}^2 + r_{xz}^2 r_{xx}^2 + r_{yx}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{$$

The effective potential is made up of four basic types of terms: • (1) kinetic/angular momentum terms (repulsive),
 • (2) quartic interaction terms (attractive), • (3) mass terms (attractive), and • (4) cubic Myers terms (repulsive).

The last two types of terms (i.e. mass terms and Myers terms) are μ -dependent and therefore absent from the flat space case ($\mu \rightarrow 0$), which was studied in Axenides-Floratos (2006)...

In either case ($\mu = 0$ or $\mu \neq 0$), it is the equilibration between attraction and repulsion which determines the minima of the effective potential... Yet another interesting aspect of these systems is the existence of closed periodic orbits which do not correspond to critical points...

• Here's the resulting effective potential:

$$\begin{split} V_{\text{eff}} &= \frac{2\pi T}{3} \left[\frac{\ell_{xx}^2}{r_{xx}^2} + \frac{\ell_{xy}^2}{r_{xy}^2} + \frac{\ell_{yx}^2}{r_{xz}^2} + \frac{\ell_{yy}^2}{r_{yx}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yy}^2}{r_{yz}^2} + r_{xx}^2 r_{xy}^2 + r_{xy}^2 r_{xz}^2 + r_{xz}^2 r_{xx}^2 + r_{yy}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 r_{yz}^2 + r_{yz}^2 r_{yz}^2$$

• As in flat spacetime, minimization of the effective potential leads to (dielectric) top solutions of the form:

$$\begin{split} \tilde{\mathbf{x}}_{1}\left(\tau\right) &= e^{\Omega_{xx}\tau} \cdot \tilde{\mathbf{x}}_{10}, \qquad \tilde{\mathbf{x}}_{2}\left(\tau\right) = e^{\Omega_{xy}\tau} \cdot \tilde{\mathbf{x}}_{20}, \qquad \tilde{\mathbf{x}}_{3}\left(\tau\right) = e^{\Omega_{xz}\tau} \cdot \tilde{\mathbf{x}}_{30} \\ \tilde{\mathbf{y}}_{1}\left(\tau\right) &= e^{\Omega_{yx}\tau} \cdot \tilde{\mathbf{y}}_{10}, \qquad \tilde{\mathbf{y}}_{2}\left(\tau\right) = e^{\Omega_{yy}\tau} \cdot \tilde{\mathbf{y}}_{20}, \qquad \tilde{\mathbf{y}}_{3}\left(\tau\right) = e^{\Omega_{yz}\tau} \cdot \tilde{\mathbf{y}}_{30}. \end{split}$$

• We can identify 3 cases, based on the ways we can distribute the 3 spatial coordinates x_i into 3 groups:

Case I: $x_1, x_2, x_3 \sim e_1$ Case II: $x_1, x_2 \sim e_1$ & $x_3 \sim e_3$ Case III: $x_1 \sim e_1, x_2 \sim e_2, x_3 \sim e_3$.

• In each case, we obtain a set of different effective potentials and (dielectric or not) membrane tops.

・ロト・西ト・ヨト・ヨト・ 日・ つへぐ

Case I:
$$q_1 = 3$$
, $q_2 = q_3 = 0$

• For $q_1 = 3$, $q_2 = q_3 = 0$ the spherical ansatz for the x-coordinates takes the form:

$$(x_1, x_2, x_3) = (\tilde{x}_1(\tau) e_1, \ \tilde{x}_2(\tau) e_1, \ \tilde{x}_3(\tau) e_1) \qquad \& \qquad r_x^2 \equiv \sum_{i=1}^3 \tilde{x}_i(\tau) \tilde{x}_i(\tau) (\ell_x)_{ij} \equiv \dot{\tilde{x}}_i(\tau) \tilde{x}_j(\tau) - \tilde{x}_i(\tau) \dot{\tilde{x}}_j(\tau).$$

The effective potential becomes:

$$\begin{split} V_{\text{eff}} &= \frac{2\pi T}{3} \left[\frac{\ell_x^2}{r_x^2} + \frac{\ell_{yx}^2}{r_{yx}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yz}^2}{r_{yz}^2} + r_{yx}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yx}^2 + r_x^2 \left(r_{yy}^2 + r_{yz}^2 \right) + \frac{\mu^2 r_x^2}{9} + \frac{\mu^2}{36} \left(r_{yx}^2 + r_{yy}^2 + r_{yz}^2 \right) \right]. \end{split}$$

- Completely symmetric (single-radius) configuration: $r = r_x = r_{yx} = r_{yy} = r_{yz}$, $\ell = \ell_x = \ell_{yx} = \ell_{yy} = \ell_{yz}$.
- There are 5 different axially symmetric (2-radii) configurations.
- There are 4 configurations with 3 different radii.

Case II: $q_1 = 2$, $q_2 = 0$, $q_3 = 1$

• For $q_1 = 2$, $q_2 = 0$, $q_3 = 1$ our ansatz is written:

$$(x_1, x_2, x_3) = (\tilde{x}_{11}(\tau) e_1, \ \tilde{x}_{12}(\tau) e_1, \ r_{xz}(\tau) e_3) \qquad \& \qquad r_{xx}^2 \equiv \tilde{x}_{11}^2(\tau) + \tilde{x}_{12}^2(\tau) \\ (\ell_{xx})_{ij} \equiv \dot{x}_{1i}(\tau) \tilde{x}_{1j}(\tau) - \tilde{x}_{1i}(\tau) \dot{x}_{1j}(\tau).$$

• The effective potential is given by:

$$V_{\text{eff}} = \frac{2\pi T}{3} \left[\frac{\ell_{xx}^2}{r_{xx}^2} + \frac{\ell_{yx}^2}{r_{yx}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yz}^2}{r_{yz}^2} + r_{xx}^2 r_{xz}^2 + r_{yx}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yx}^2 + r_{xx}^2 \left(r_{yy}^2 + r_{yz}^2 \right) + r_{xx}^2 \left(r_{yy}^2 + r_{yx}^2 \right) + \frac{\mu^2}{9} \left(r_{xx}^2 + r_{xz}^2 \right) + \frac{\mu^2}{36} \left(r_{yx}^2 + r_{yy}^2 + r_{yz}^2 \right) \right].$$

- Completely symmetric configuration: $r = r_{xx} = r_{yx} = r_{yy} = r_{yz}$, $\ell = \ell_{xx} = \ell_{yx} = \ell_{yy} = \ell_{yz}$.
- There are 13 different axially symmetric (2-radii) configurations.
- There exist 21 three-radii configurations.

Example 1

• Take for example a type II configuration with all the SO(6) variables y_i set to zero:

$$x_1 = x(\tau) \cdot e_1, \quad x_2 = y(\tau) \cdot e_1, \quad x_3 = z(\tau) \cdot e_2 \qquad \& \qquad y_i = 0, \quad i = 1, \dots, 6,$$

The corresponding effective potential reads,

$$V_{\text{eff}} = \frac{2\pi T}{3} \left[\frac{\ell^2}{x^2 + y^2} + \left(x^2 + y^2\right) z^2 + \frac{\mu^2}{9} \left(x^2 + y^2 + z^2\right) \right],$$

where we have set $\ell_{x1} = \ell$. The corresponding minimization condition $\nabla V_{eff} = 0$ leads to

$$x z^{2} + \frac{\mu^{2} x}{9} - \frac{x \ell^{2}}{\left(x^{2} + y^{2}\right)^{2}} = y z^{2} + \frac{\mu^{2} y}{9} - \frac{y \ell^{2}}{\left(x^{2} + y^{2}\right)^{2}} = z \left(x^{2} + y^{2}\right) + \frac{\mu^{2} z}{9} = 0,$$

which has the following solution

$$x^2 + y^2 = \frac{3\ell}{\mu}$$
 & $z = 0.$

To agree with the form of the above ansatz we can choose, for instance,

$$x(\tau) = \sqrt{\frac{3\ell}{\mu}} \cos \frac{\mu \tau}{3}, \qquad y(\tau) = \sqrt{\frac{3\ell}{\mu}} \sin \frac{\mu \tau}{3}, \qquad z(\tau) = 0.$$

29 / 53

Example 1

• Alternatively we could have directly inserted the ansatz into the light-cone equations of motion,

$$\ddot{x} \cdot e_1 = -x \, z^2 \cdot e_1 - \frac{\mu^2 x}{9} \cdot e_1 + \mu \, y \, z \cdot e_3$$
$$\ddot{y} \cdot e_1 = -y \, z^2 \cdot e_1 - \frac{\mu^2 y}{9} \cdot e_1 + \mu \, x \, z \cdot e_3$$
$$\ddot{z} \cdot e_2 = -z \left(x^2 + y^2\right) \cdot e_2 - \frac{\mu^2 z}{9} \cdot e_2,$$

from which it can be seen that any solution of the type

$$\tilde{\mathbf{x}}_{1}\left(\tau\right)=e^{\Omega_{xx}\tau}\cdot\tilde{\mathbf{x}}_{10},\qquad\tilde{\mathbf{x}}_{2}\left(\tau\right)=e^{\Omega_{xy}\tau}\cdot\tilde{\mathbf{x}}_{20},\qquad\tilde{\mathbf{x}}_{3}\left(\tau\right)=e^{\Omega_{xz}\tau}\cdot\tilde{\mathbf{x}}_{30},$$

will satisfy

$$x^2 + y^2 = \frac{3\ell}{\mu}$$
 & $z = 0.$

↓ □ ▶ ↓ ⑦ ▶ ↓ 差 ▶ ↓ 差 ♪ ♪ ② へ (?)
29 / 53

Example 2

• Another interesting type II solution is the following:

$$x_1 = x(\tau) \cdot e_1, \quad x_2 = y(\tau) \cdot e_2, \quad x_3 = 0$$
 & $y_i = 0, \quad i = 1, \dots, 6,$

where again all the SO(6) variables y_i and the SO(3) coordinate x_2 are zero... The effective potential becomes,

$$V_{\text{eff}} = rac{2\pi T}{3} \Bigg[x^2 y^2 + rac{\mu^2}{9} \left(x^2 + y^2
ight) \Bigg],$$

so that there is only one trivial critical point at x = y = 0, which is obtained by minimizing the effective potential:

$$x y^{2} + \frac{\mu^{2} x}{9} = y x^{2} + \frac{\mu^{2} y}{9} = 0$$

Potentials of the above form (which are in fact generalizations of the YM potential $x^2y^2/2$) have a very interesting and rich set of (stable) periodic orbits... See e.g. Contopoulos-Harsoula (2023)...

Case III: $q_1 = q_2 = q_3 = 1$

• For $q_1 = q_2 = q_3 = 1$ the spherical ansatz becomes:

$$(x_1, x_2, x_3) = (r_{xx}(\tau) e_1, r_{xy}(\tau) e_2, r_{xz}(\tau) e_3).$$

In this case the effective potential reads:

$$\begin{split} V_{\text{eff}} &= \frac{2\pi T}{3} \left[\frac{\ell_{yx}^2}{r_{yy}^2} + \frac{\ell_{yy}^2}{r_{yy}^2} + \frac{\ell_{yz}^2}{r_{yz}^2} + r_{xx}^2 r_{xy}^2 + r_{xy}^2 r_{xz}^2 + r_{xz}^2 r_{xx}^2 + r_{yz}^2 r_{yy}^2 + r_{yy}^2 r_{yz}^2 + r_{yz}^2 r_{yx}^2 + \\ &+ r_{xx}^2 \left(r_{yy}^2 + r_{yz}^2 \right) + r_{xy}^2 \left(r_{yz}^2 + r_{yx}^2 \right) + r_{xz}^2 \left(r_{yx}^2 + r_{yy}^2 \right) + \frac{\mu^2}{9} \left(r_{xx}^2 + r_{xy}^2 + r_{xz}^2 \right) + \\ &+ \frac{\mu^2}{36} \left(r_{yx}^2 + r_{yy}^2 + r_{yz}^2 \right) - 2\mu r_{xx} r_{xy} r_{xz} \right]. \end{split}$$

- Completely symmetric configuration: $r = r_{xx} = r_{xy} = r_{yz} = r_{yy} = r_{yz}$, $\ell = \ell_{yx} = \ell_{yy} = \ell_{yz}$.
- There are 9 different axially symmetric (2-radii) configurations.
- There are 10 configurations with 3 different radii.

An example

• Consider the following (out the 9 in total) axially symmetric configuration of case III:

$$r_x = r_{xx} = r_{xy} = r_{xz}$$
 & $r_y = r_{yx} = r_{yy} = r_{yz}$ & $\ell_y = \ell_{yx} = \ell_{yy} = \ell_{yz}$

with effective potential:

$$V_{\text{eff}} = 2\pi T \left[\frac{\ell_y^2}{r_y^2} + r_x^4 + 2r_x^2 r_y^2 + \frac{\mu^2 r_x^2}{9} + \frac{\mu^2 r_y^2}{36} - \frac{2\mu}{3} r_x^3 \right].$$

• The minimization conditions read:

$$\frac{dV_{\text{eff}}}{dr_{r_x}} = r_x \left(r_x^2 - \frac{\mu}{2} r_x + r_y^2 + \frac{\mu^2}{18} \right) = \frac{dV_{\text{eff}}}{dr_{r_y}} = r_y^6 + \left(r_x^2 + \frac{\mu^2}{72} \right) r_y^4 - \frac{\ell_y^2}{2} = 0.$$

We obtain the following selection rule:

$$r_y \leq rac{\mu}{12}$$
 & $r_x \geq rac{144^2 \ell_y^2}{\mu^5} + rac{\mu}{12}$

e.g. in the marginal case $r_x = \mu/4$, $r_y = \mu/12$, $\ell_y = \mu^3/144\sqrt{6}$, it's $V_{\rm eff(min)} = 7\pi T \mu^4/1296$.

• We also find the static solutions $r_y = 0$, $r_x = \mu/3$ (BPS) and $r_x = \mu/6$ (for which $V_{\text{eff(min)}} = \pi T \mu^4/648$)...

Section 4

Static dielectric membranes in SO(3)

M. Axenides, E. Floratos, GL M2-brane dynamics in the classical limit of the BMN matrix model PLB **773** (2017) 265 [arxiv:1707.02878]

M. Axenides, E. Floratos, GL to appear

M. Axenides, E. Floratos, GL Multipole stability of spinning M2-branes in the classical limit of the BMN matrix model PRD **97** (2018) 126019 [arxiv:1712.06544]

The SO(3) solution

Setting all SO(6) coordinates of type III configurations to zero and switching to dimensionless time $t \equiv \mu \tau$,

$$x_i = \mu u_i e_i, \quad i = 1, 2, 3$$
 & $y_i = \mu v_i = 0, \quad i = 1, \dots 6,$

the membrane equations of motion become:

$$\ddot{u}_1 + \left(u_2^2 + u_3^2 + \frac{1}{9}\right)u_1 = u_2u_3, \qquad \ddot{u}_2 + \left(u_1^2 + u_3^2 + \frac{1}{9}\right)u_2 = u_1u_3, \qquad \ddot{u}_3 + \left(u_1^2 + u_2^2 + \frac{1}{9}\right)u_3 = u_1u_2, \\ \ddot{v}_i = 0, \quad i = 1, \dots, 6.$$

The dynamics is fully specified in terms of the Hamiltonian...

$$H = \frac{4\pi T \mu^4}{3} \cdot \mathcal{H}, \quad \mathcal{H} \equiv \frac{1}{2} \left[p_1^2 + p_2^2 + p_3^2 + u_1^2 u_2^2 + u_2^2 u_3^2 + u_1^2 u_3^2 + \frac{1}{9} \left(u_1^2 + u_2^2 + u_3^2 \right) - 2u_1 u_2 u_3 \right],$$

and Hamilton's equations of motion:

$$p_i = \dot{u}_i, \qquad \dot{p}_i = -\frac{\partial \mathcal{H}}{\partial u_i},$$

which evidently imply the above Lagrangian equations of motion...

The SO(3) solution

Setting all SO(6) coordinates of type III configurations to zero and switching to dimensionless time $t \equiv \mu \tau$,

$$x_i = \mu u_i e_i, \quad i = 1, 2, 3$$
 & $y_i = \mu v_i = 0, \quad i = 1, \dots 6,$

the membrane equations of motion become:

$$\ddot{u}_1 + \left(u_2^2 + u_3^2 + \frac{1}{9}\right)u_1 = u_2u_3, \qquad \ddot{u}_2 + \left(u_1^2 + u_3^2 + \frac{1}{9}\right)u_2 = u_1u_3, \qquad \ddot{u}_3 + \left(u_1^2 + u_2^2 + \frac{1}{9}\right)u_3 = u_1u_2, \\ \ddot{v}_i = 0, \quad i = 1, \dots, 6.$$

The effective potential energy of the static membrane is given by

$$V_{\text{eff}} = \frac{2\pi T \mu^4}{3} \left[\left(u_1^2 u_2^2 + u_2^2 u_3^2 + u_1^2 u_3^2 \right) + \frac{1}{9} \left(u_1^2 + u_2^2 + u_3^2 \right) - 2u_1 u_2 u_3 \right].$$

The SO(3) solution

Setting all SO(6) coordinates of type III configurations to zero and switching to dimensionless time $t \equiv \mu \tau$,

$$x_i = \mu u_i e_i, \quad i = 1, 2, 3$$
 & $y_i = \mu v_i = 0, \quad i = 1, \dots 6,$

the membrane equations of motion become:

$$\ddot{u}_1 + \left(u_2^2 + u_3^2 + \frac{1}{9}\right)u_1 = u_2u_3, \qquad \ddot{u}_2 + \left(u_1^2 + u_3^2 + \frac{1}{9}\right)u_2 = u_1u_3, \qquad \ddot{u}_3 + \left(u_1^2 + u_2^2 + \frac{1}{9}\right)u_3 = u_1u_2, \\ \ddot{v}_i = 0, \quad i = 1, \dots, 6.$$

The effective potential energy of the static membrane is given by

$$V_{\text{eff}} = \frac{2\pi T \mu^4}{3} \Bigg[\left(u_1^2 u_2^2 + u_2^2 u_3^2 + u_1^2 u_3^2 \right) + \frac{1}{9} \left(u_1^2 + u_2^2 + u_3^2 \right) - 2u_1 u_2 u_3 \Bigg].$$

This potential turns out to be a special case of the so-called generalized 3-dimensional Hénon-Heiles potential,

$$V_{\rm HH} = \frac{1}{2} \left(u_1^2 + u_2^2 + u_3^2 \right) + K_3 \, u_1 u_2 u_3 + K_0 \left(u_1^2 + u_2^2 + u_3^2 \right)^2 + K_4 \left(u_1^4 + u_2^4 + u_3^4 \right) \quad \text{(Efstathiou-Sadovskii, 2004)}.$$

For $K_3 = -9$, $K_0 = -K_4 = 9/4$, V_{HH} obviously reduces to the above effective potential V_{eff} .

The extrema of the potential solve the equilibrium conditions:

$$\begin{split} \partial_i V_{\text{eff}} &= 0 \Rightarrow \left(u_2^2 + u_3^2 + \frac{1}{9} \right) u_1 = u_2 u_3 \\ & \left(u_3^2 + u_1^2 + \frac{1}{9} \right) u_2 = u_3 u_1 \\ & \left(u_1^2 + u_2^2 + \frac{1}{9} \right) u_3 = u_1 u_2. \end{split}$$

Here are the corresponding roots:

$$\mathbf{u}_0 = 0, \qquad \mathbf{u}_{1/6} = rac{1}{6} \cdot (\pm 1, \pm 1, \pm 1), \qquad \mathbf{u}_{1/3} = rac{1}{3} \cdot (\pm 1, \pm 1, \pm 1),$$

The extrema of the potential solve the equilibrium conditions:

$$\partial_i V_{\text{eff}} = 0 \Rightarrow \left(u_2^2 + u_3^2 + \frac{1}{9} \right) u_1 = u_2 u_3$$
$$\left(u_3^2 + u_1^2 + \frac{1}{9} \right) u_2 = u_3 u_1$$
$$\left(u_1^2 + u_2^2 + \frac{1}{9} \right) u_3 = u_1 u_2.$$

Here are the corresponding roots:

$$u_0=0, \qquad u_{1/6}=\frac{1}{6}\cdot (\pm 1,\pm 1,\pm 1)\,, \qquad u_{1/3}=\frac{1}{3}\cdot (\pm 1,\pm 1,\pm 1)\,,$$

• The extrema are nine in total because the product of their components must be non-negative...

The extrema of the potential solve the equilibrium conditions:

$$\partial_i V_{\text{eff}} = 0 \Rightarrow \left(u_2^2 + u_3^2 + \frac{1}{9} \right) u_1 = u_2 u_3$$
$$\left(u_3^2 + u_1^2 + \frac{1}{9} \right) u_2 = u_3 u_1$$
$$\left(u_1^2 + u_2^2 + \frac{1}{9} \right) u_3 = u_1 u_2.$$

Here are the corresponding roots:

$$u_0=0, \qquad u_{1/6}=\frac{1}{6}\cdot (\pm 1,\pm 1,\pm 1)\,, \qquad u_{1/3}=\frac{1}{3}\cdot (\pm 1,\pm 1,\pm 1)\,,$$

- The extrema are nine in total because the product of their components must be non-negative...
- V_{eff} has the symmetry of a tetrahedron T_d that is generated by the 4 critical points $\mathbf{u}_{1/3}$ and $\mathbf{u}_{1/6}$...

The extrema of the potential solve the equilibrium conditions:

$$\begin{split} \partial_i V_{\text{eff}} &= 0 \Rightarrow \left(u_2^2 + u_3^2 + \frac{1}{9} \right) u_1 = u_2 u_3 \\ & \left(u_3^2 + u_1^2 + \frac{1}{9} \right) u_2 = u_3 u_1 \\ & \left(u_1^2 + u_2^2 + \frac{1}{9} \right) u_3 = u_1 u_2. \end{split}$$

Here are the corresponding roots:

$$u_0=0, \qquad u_{1/6}=\frac{1}{6}\cdot (\pm 1,\pm 1,\pm 1)\,, \qquad u_{1/3}=\frac{1}{3}\cdot (\pm 1,\pm 1,\pm 1)\,,$$

- The extrema are nine in total because the product of their components must be non-negative...
- V_{eff} has the symmetry of a tetrahedron T_d that is generated by the 4 critical points $\mathbf{u}_{1/3}$ and $\mathbf{u}_{1/6}$...
- u_0 (point-like membrane) and $u_{1/3}$ (Myers dielectric sphere) are global minima, while $u_{1/6}$ is a saddle point...
$$V_{\rm eff}\left(0\right)=V_{\rm eff}\left(\frac{1}{3}\right)=0, \qquad V_{\rm eff}\left(\frac{1}{6}\right)=\frac{2\pi T \mu^4}{6^4}.$$

• The value of the effective potential at the extremal points is

$$V_{\mathrm{eff}}\left(0
ight)=V_{\mathrm{eff}}\left(rac{1}{3}
ight)=0,\qquad V_{\mathrm{eff}}\left(rac{1}{6}
ight)=rac{2\pi T\mu^4}{6^4}.$$

• Hessian matrix is positive-definite for u_0 and $u_{1/3}$ (global minima) and indefinite for $u_{1/6}$ (saddle point)...

$$V_{
m eff}(0) = V_{
m eff}\left(rac{1}{3}
ight) = 0, \qquad V_{
m eff}\left(rac{1}{6}
ight) = rac{2\pi T \mu^4}{6^4}.$$

- Hessian matrix is positive-definite for u_0 and $u_{1/3}$ (global minima) and indefinite for $u_{1/6}$ (saddle point)...
- This result will be confirmed below by leading order (LO) radial and angular/mutlipole perturbations... next-to-leading
 order (NLO) perturbations will be studied right after...

$$V_{\mathrm{eff}}\left(0
ight)=V_{\mathrm{eff}}\left(rac{1}{3}
ight)=0,\qquad V_{\mathrm{eff}}\left(rac{1}{6}
ight)=rac{2\pi T\mu^4}{6^4}.$$

- Hessian matrix is positive-definite for \mathbf{u}_0 and $\mathbf{u}_{1/3}$ (global minima) and indefinite for $\mathbf{u}_{1/6}$ (saddle point)...
- This result will be confirmed below by leading order (LO) radial and angular/mutlipole perturbations... next-to-leading
 order (NLO) perturbations will be studied right after...
- When the *u_i* in are not all equal, the equations of motion are so complicated that exact time-dependent solutions can only be found numerically...

$$V_{\mathrm{eff}}\left(0
ight)=V_{\mathrm{eff}}\left(rac{1}{3}
ight)=0,\qquad V_{\mathrm{eff}}\left(rac{1}{6}
ight)=rac{2\pi T\mu^4}{6^4}.$$

- Hessian matrix is positive-definite for \mathbf{u}_0 and $\mathbf{u}_{1/3}$ (global minima) and indefinite for $\mathbf{u}_{1/6}$ (saddle point)...
- This result will be confirmed below by leading order (LO) radial and angular/mutlipole perturbations... next-to-leading
 order (NLO) perturbations will be studied right after...
- When the *u_i* in are not all equal, the equations of motion are so complicated that exact time-dependent solutions can only be found numerically...
- When all the SO(3) membrane coordinates u_i are equal, an analytic solution is possible...

Subsection 2

Spherically symmetric membrane



The spherically symmetric membrane

• The ansatz for the fully symmetric membrane in SO(3) reads:

 $u \equiv u_1 = u_2 = u_3.$

• The membrane Hamiltonian is that of a double-well oscillator:

$$H = 2\pi T \mu^4 \left[p^2 + u^2 \left(u - \frac{1}{3} \right)^2 \right].$$

• Here are the corresponding equations of motion:

$$\dot{u}=p,$$
 $\dot{p}=-u\left[2u^2+rac{1}{9}-u
ight].$

Define E ≡ E/2πTµ⁴, E_c ≡ 6⁻⁴. There are three kinds of orbits:
 (1) oscillations of small energies (E < E_c) around either of the two stable global minima (u₀ = 0, 1/3) • (2) oscillations of larger energies (E > E_c) around the local maximum (u₀ = 1/6) • (3) two homoclinic orbits through the unstable equilibrium point at u₀ = 1/6 with energy equal to the potential height (E = E_c).



(日)(同)(日)(日)(日)(日)

The spherically symmetric membrane

• The orbits may be computed from the initial conditions:

$$\dot{u}_0\left(0
ight)=0, \qquad u_0\left(0
ight)=rac{1}{6}\pm\sqrt{rac{1}{6^2}+\sqrt{\mathcal{E}}},$$

where the \pm signs correspond to the right/left side of the double-well potential.

Integrating the energy integral we find the solution:

$$u_0(t) = \frac{1}{6} \pm \sqrt{\frac{1}{6^2} + \sqrt{\mathcal{E}}} \cdot cn\left[\sqrt{2\sqrt{\mathcal{E}}} \cdot t \left| \frac{1}{2} \left(1 + \frac{1}{36\sqrt{\mathcal{E}}}\right) \right],$$

where only the plus sign should be kept for $\mathcal{E} \geq \mathcal{E}_c$.

• For the critical energy $\mathcal{E} = \mathcal{E}_c$ the homoclinic orbit is obtained:

$$u_0(t) = rac{1}{6} \pm rac{1}{3\sqrt{2}} \cdot \operatorname{sech}\left(rac{t}{3\sqrt{2}}
ight)$$



(日)(同)(日)(日)(日)(日)

The spherically symmetric membrane

• The period as a function of the energy is given in terms of the complete elliptic integral of the first kind:

$$\mathcal{T}\left(\mathcal{E}
ight)=2\sqrt{rac{2}{\sqrt{\mathcal{E}}}}\cdot\mathsf{K}\left(rac{1}{2}\left(1+rac{1}{36\sqrt{\mathcal{E}}}
ight)
ight),$$

it becomes infinite for the homoclinic orbit $\mathcal{E} = \mathcal{E}_c$. For more, see e.g. Brizard-Westland (2017).



Subsection 3

Leading order stability

・ロト・日本・モン・モン・モー りへぐ

40 / 53

Leading order stability analysis

The full type III equations of motion around each of the SO(3) extremal points read:

$$\begin{split} \ddot{u}_{1} + \left(u_{2}^{2} + u_{3}^{2} + \frac{r_{y2}^{2}}{\mu^{2}} + \frac{r_{y3}^{2}}{\mu^{2}} + \frac{1}{9}\right)u_{1} &= u_{2}u_{3}, \quad \ddot{v}_{i} + \left(\frac{r_{y2}^{2}}{\mu^{2}} + \frac{r_{y3}^{2}}{\mu^{2}} + u_{2}^{2} + u_{3}^{2} + \frac{1}{36}\right)v_{i} &= 0\\ \ddot{u}_{2} + \left(u_{3}^{2} + u_{1}^{2} + \frac{r_{y3}^{2}}{\mu^{2}} + \frac{r_{y1}^{2}}{\mu^{2}} + \frac{1}{9}\right)u_{2} &= u_{3}u_{1}, \quad \ddot{v}_{j} + \left(\frac{r_{y3}^{2}}{\mu^{2}} + \frac{r_{y1}^{2}}{\mu^{2}} + u_{3}^{2} + u_{1}^{2} + \frac{1}{36}\right)v_{j} &= 0\\ \ddot{u}_{3} + \left(u_{1}^{2} + u_{2}^{2} + \frac{r_{y1}^{2}}{\mu^{2}} + \frac{r_{y2}^{2}}{\mu^{2}} + \frac{1}{9}\right)u_{3} &= u_{1}u_{2}, \quad \ddot{v}_{k} + \left(\frac{r_{y1}^{2}}{\mu^{2}} + \frac{r_{y2}^{2}}{\mu^{2}} + v_{1}^{2} + v_{2}^{2} + \frac{1}{36}\right)v_{k} &= 0, \end{split}$$

where we have set $t \equiv \mu \tau$ and

$$x_i = \mu u_i e_i, \quad i = 1, 2, 3 \qquad \& \qquad y_i = \mu v_i e_1, \quad i = 1, \dots, s_1$$
$$y_j = \mu v_j e_2, \quad j = s_1 + 1, \dots, s_1 + s_2$$
$$y_k = \mu v_k e_3, \quad k = s_1 + s_2 + 1, \dots, s_1 + s_2 + s_3$$

<ロト < 団ト < 巨ト < 巨ト < 巨ト 三 のQで 41 / 53

Radial stability analysis

• By radially perturbing each of the 9 critical points as:

$$u_i = u_i^0 + \delta u_i(t), \quad i = 1, 2, 3, \qquad \& \qquad v_j = \delta v_j(t), \quad j = 1, \dots, 6,$$

we obtain the following system of fluctuation equations

$$\delta \ddot{\mathbf{u}} = - \begin{bmatrix} 2u_0^2 + \frac{1}{9} & 2u_1^0 u_2^0 - u_3^0 & 2u_1^0 u_3^0 - u_2^0 \\ 2u_2^0 u_1^0 - u_3^0 & 2u_0^2 + \frac{1}{9} & 2u_2^0 u_3^0 - u_1^0 \\ 2u_3^0 u_1^0 - u_2^0 & 2u_3^0 u_2^0 - u_1^0 & 2u_0^2 + \frac{1}{9} \end{bmatrix} \cdot \delta \mathbf{u} \quad \& \quad \delta \ddot{\mathbf{v}} = -\left(2u_0^2 + \frac{1}{36}\right) \cdot \delta \mathbf{v},$$

where we have defined

$$u_0^2 \equiv (u_1^0)^2 = (u_2^0)^2 = (u_3^0)^2$$

for the common value of the square of each extremum's components. Then we plug the particular solution

$$\begin{bmatrix} \delta \mathbf{u} \\ \delta \mathbf{v} \end{bmatrix} = e^{\lambda t} \boldsymbol{\xi},$$

we solve the resulting eigenvalue/eigenvector problem...

Radial stability analysis

• By radially perturbing each of the 9 critical points as:

$$u_i = u_i^0 + \delta u_i(t), \quad i = 1, 2, 3, \qquad \& \qquad v_j = \delta v_j(t), \quad j = 1, \dots, 6.$$

This way we confirm the conclusion we derived above from the corresponding Hessian matrix, i.e. that \mathbf{u}_0 and $\mathbf{u}_{1/3}$ are global minima (positive-definite Hessian) and $\mathbf{u}_{1/6}$ is a saddle point (indefinite Hessian):

extremum	location	eigenvalues $r=\lambda^2$ (degeneracy)	stability
u ₀	0	$-\frac{1}{9}$ (3), $-\frac{1}{36}$ (6)	center (stable)
u _{1/6}	$\left(\pm\tfrac{1}{6},\pm\tfrac{1}{6},\pm\tfrac{1}{6}\right)$	$\frac{1}{18}$ (1), $-\frac{5}{18}$ (2), $-\frac{1}{12}$ (6)	saddle point
u _{1/3}	$\left(\pm\tfrac{1}{3},\pm\tfrac{1}{3},\pm\tfrac{1}{3}\right)$	$-rac{1}{9}(1),-rac{4}{9}(2),-rac{1}{4}(6)$	center (stable)

Axenides-Floratos-GL (2017a)

where the negative eigenvalues $r = \lambda^2 < 0$ correspond to stable directions and the positive eigenvalues $r = \lambda^2 > 0$ lead to stable/unstable directions (depending on the sign of the real eigenvalue λ)...

Angular stability analysis

• We may also perform more general (angular/multipole) perturbations of the following form:

$$x_{i}(t) = x_{i}^{0} + \delta x_{i}(t), \qquad i = 1, 2, 3$$

where δx_i is expanded in spherical harmonics $Y_{jm}(\theta, \phi)$:

$$x_i(t) = \mu u_i(t) e_i, \qquad x_i^0 = \mu u_i^0 e_i, \qquad \delta x_i(t) = \mu \cdot \sum_{j=1}^{\infty} \sum_{m=-j}^{J} \eta_i^{jm}(t) Y_{jm}(\theta, \phi).$$

• For the critical points u_0 , $u_{1/6}$, $u_{1/3}$ we find the eigenvalues (Axenides-Floratos-GL, 2017b):

$$\begin{split} \mathbf{u}_{0} : \quad \lambda_{P}^{2} &= \lambda_{\pm}^{2} = -\frac{1}{9}, \qquad \lambda_{\theta}^{2} = -\frac{1}{36} \\ \mathbf{u}_{1/6} : \lambda_{P}^{2} &= 0, \qquad \lambda_{\pm}^{2} = -\frac{1}{36} \left(j + 1 \right) \left(j + 4 \right), \qquad \lambda_{-}^{2} = -\frac{j \left(j - 3 \right)}{36}, \qquad \lambda_{\theta}^{2} = -\frac{1}{36} \left(j^{2} + j + 1 \right) \\ \mathbf{u}_{1/3} : \lambda_{P}^{2} &= 0, \qquad \lambda_{\pm}^{2} = -\frac{1}{36} \left(j + 1 \right)^{2}, \qquad \lambda_{-}^{2} = -\frac{j^{2}}{9}, \qquad \lambda_{\theta}^{2} = -\frac{1}{36} \left(2j + 1 \right)^{2}, \end{split}$$

with multiplicities $d_P = 2j + 1$, $d_+ = 2j + 3$, $d_- = 2j - 1$ and $d_\theta = 6(2j + 1)$, respectively.

Angular stability analysis

- The critical point **u**₀ (point-like membrane) is obviously stable.
- $\mathbf{u}_{1/3}$ has a zero mode of degeneracy $2d_P$ while all its other eigenvalues are stable for j = 1, 2, ...
- u_{1/6} has one 2d_P-degenerate zero mode for every j and a 10-fold degenerate zero mode for j = 3. It is unstable for j = 1 (2-fold degenerate) and j = 2 (6-fold degenerate).
- The above results were first obtained by (Dasgupta, Sheikh-Jabbari, Van Raamsdonk (2002)) from the BMN matrix model point-of-view.
- In the flat-space limit ($\mu \rightarrow 0$), we recover the results of (Axenides-Floratos-Perivolaropoulos, 2000, 2001).

Section 5

The $SO(3) \times SO(6)$ symmetric membrane

M. Axenides, E. Floratos, GL M2-brane dynamics in the classical limit of the BMN matrix model PLB **773** (2017) 265 [arxiv:1707.02878]

M. Axenides, E. Floratos, GL to appear

M. Axenides, E. Floratos, GL Multipole stability of spinning M2-branes in the classical limit of the BMN matrix model PRD **97** (2018) 126019 [arxiv:1712.06544]

The $SO(3) \times SO(6)$ sector

• Similar analysis can be carried out in the $SO(3) \times SO(6)$ sector where the equations of motion become:

$$\begin{split} \ddot{u}_{1} + \left(u_{2}^{2} + u_{3}^{2} + \frac{r_{y2}^{2}}{\mu^{2}} + \frac{r_{y3}^{2}}{\mu^{2}} + \frac{1}{9}\right) u_{1} &= u_{2}u_{3}, \quad \ddot{v}_{i} + \left(\frac{r_{y2}^{2}}{\mu^{2}} + \frac{r_{y3}^{2}}{\mu^{2}} + u_{2}^{2} + u_{3}^{2} + \frac{1}{36}\right) v_{i} &= 0, \quad i, j, k = 1, 2, 3, \\ \ddot{u}_{2} + \left(u_{3}^{2} + u_{1}^{2} + \frac{r_{y3}^{2}}{\mu^{2}} + \frac{r_{y1}^{2}}{\mu^{2}} + \frac{1}{9}\right) u_{2} &= u_{3}u_{1}, \quad \ddot{v}_{j} + \left(\frac{r_{y3}^{2}}{\mu^{2}} + \frac{r_{y1}^{2}}{\mu^{2}} + u_{3}^{2} + u_{1}^{2} + \frac{1}{36}\right) v_{j} &= 0 \\ \ddot{u}_{3} + \left(u_{1}^{2} + u_{2}^{2} + \frac{r_{y1}^{2}}{\mu^{2}} + \frac{r_{y2}^{2}}{\mu^{2}} + \frac{1}{9}\right) u_{3} &= u_{1}u_{2}, \quad \ddot{v}_{k} + \left(\frac{r_{y1}^{2}}{\mu^{2}} + \frac{r_{y2}^{2}}{\mu^{2}} + v_{1}^{2} + v_{2}^{2} + \frac{1}{36}\right) v_{k} &= 0. \end{split}$$

• A solution of these equations of motion is

$$u_i^0 = u_0, \qquad v_j^0(t) = v_0 \cos\left(\omega t + \varphi_j\right), \qquad w_j^0(t) \equiv v_{j+3}^0(t) = v_0 \sin\left(\omega t + \varphi_k\right),$$

where (u_0, v_0) are the critical points of the corresponding (axially symmetric) potential

$$V \equiv \frac{V_{\rm eff}}{2\pi T \mu^4} = u^4 + 2u^2 v^2 + v^4 + \frac{u^2}{9} + \frac{v^2}{36} - \frac{2u^3}{3} + \frac{\ell^2}{v^2}, \qquad \ell \mu^3 \equiv \ell_1 = \ell_2 = \ell_3.$$

• It can be proven that the critical points (u_0, v_0) always lie within the interval:

$$\frac{1}{6} \le u_0 \le \frac{1}{3} \qquad \& \qquad 0 \le v_0 \le \frac{1}{12}.$$

46 / 53

Radial stability analysis

Setting,

$$u_{i} = u_{i}^{0} + \delta u_{i}(t), \qquad v_{i} = v_{i}^{0}(t) + \delta v_{i}'(t), \qquad w_{i} = w_{i}^{0}(t) + \delta w_{i}'(t),$$

we find the radial eigenvalues



Angular stability analysis

• Going further, we again set out to perform angular/multipole perturbations of the form:

$$x_i = x_i^0 + \delta x_i, \quad i = 1, 2, 3$$
 & $y_k = y_k^0 + \delta y_k, \quad k = 1, \dots, 6$

where the δx_i , δy_k are expanded in spherical harmonics $Y_{jm}(\theta, \phi)$ as:

$$\delta x_{i} = \mu \cdot \sum_{j,m} \eta_{i}^{jm}(\tau) Y_{jm}(\theta,\phi) \qquad \delta y_{k} = \mu \cdot \sum_{j,m} \epsilon_{k}^{jm}(\tau) Y_{jm}(\theta,\phi) \qquad \delta y_{l} = \mu \cdot \sum_{j,m} \zeta_{l}^{jm}(\tau) Y_{jm}(\theta,\phi),$$

and
$$i = 1, 2, 3$$
, $k = 1, 3, 5$ and $l = 2, 4, 6$.

• One of the eigenvalues always vanishes, two others are given by the following analytic expression

$$\lambda_{P}^{2} = \frac{1}{2} \left(j^{2} + j + 2\right) u_{0} - \frac{1}{18} \left(1 + j \left(j + 1\right) \pm 3 \sqrt{144 \left(j^{2} + j - 2\right) u_{0}^{3} - 12 \left(j^{2} + j - 14\right) u_{0}^{2} - 24 u_{0} + 1}\right),$$

while 6 more eigenvalues λ_{\pm} are also known in closed forms but are too complicated to be included here.

• The corresponding multiplicities of the eigenvalues are $d_P = 2j + 1$, $d_+ = 2j + 3$, $d_- = 2j - 1$.

Lowest-lying modes

• For j = 1 four eigenvalues vanish, while two others coincide with those found from radial perturbations:

$$\begin{split} \lambda_P^2 &= 4u_0 + \frac{1}{3}, \qquad \lambda_+^2 = \frac{5u_0}{2} - \frac{1}{9} \pm \sqrt{\frac{1}{9^2} - \frac{u_0}{9} - \frac{5u_0^2}{12} + 4u_0^3} \\ \lambda_-^2 &= \frac{5u_0}{2} - \frac{5}{18} \pm \sqrt{\frac{5^2}{18^2} - \frac{35u_0}{18} + \frac{163u_0^2}{12} - 20u_0^3} \end{split}$$

• For j = 2 there's one zero eigenvalue while $\lambda_P > 0$. We can also plot the eigenvalues of λ_{\pm} :



Multipole stability

- The nonzero j = 1 eigenvalues are all positive/stable in the interval $\frac{1}{6} \le u_0 \le \frac{1}{3}$, $0 \le v_0 \le \frac{1}{12}$, except $\lambda_{-(-)}^2$ which is positive/stable only for $u_{\text{crit}} < u_0 < 1/3$, where $u_{\text{crit}} \equiv \frac{1}{60} \left(11 + \sqrt{21}\right)$.
- For j = 2, the λ_P , λ_+ and one of the λ_- eigenvalues are positive/stable. The remaining λ_- eigenvalue is negative/unstable in the interval $\frac{1}{6} \leq u_0 \leq 0.207245 < u_{crit}$.

eigenvalues	j = 1	<i>j</i> = 2	$j \ge 3$	degeneracy
λ_P^2	0,0,+	0,+,+	0,+,+	$d_P = 2j + 1$
λ_+^2	0,+,+	+, +, +	+, +, +	$d_+=2j+3$
λ_{-}^{2}	$\begin{array}{c} 0,+,\{0,\pm\} \\ \left(\begin{array}{c} \text{positive for} \\ u_0 > u_{\text{crit}} \end{array} \right) \end{array}$	$\left(\begin{array}{c} +,+,\{0,\pm\}\\ \left(\begin{array}{c} \text{positive for}\\ u_0>0.207245 \end{array}\right)\right)$	+, +, +	d=2j-1

• Here's a summary of the angular/multipole spectrum:

Instability cascade

• By examining higher orders in perturbation theory beyond the linear level (in the interval $1/6 \le u_0 \le u_{crit}$) we expect to obtain a cascade of instabilities that originates from the j = 1, 2 sectors and propagates towards the higher multipoles...

$$\begin{aligned} x_i &= \sum_{n=0}^{\infty} \varepsilon^n \delta x_i^n = x_i^0 + \sum_{n=1}^{\infty} \varepsilon^n \delta x_i^n, \quad i = 1, 2, 3 \\ y_i &= \sum_{n=0}^{\infty} \varepsilon^n \delta y_i^n = y_i^0 + \sum_{n=1}^{\infty} \varepsilon^n \delta y_i^n, \quad i = 1, \dots, 6. \end{aligned}$$

• This is due to the fact that the various (constant *j*) multipoles at a given order in perturbation theory couple to all the *j*'s of the previous orders through an effective forcing term that arises in the corresponding fluctuation equation...

$$\delta x_i^n = \mu \cdot \sum_{j,m} \eta_i^{njm}(\tau) Y_{jm}(\theta, \phi), \qquad \eta_i^{njm}(0) = 0, \qquad i = 1, 2, 3$$
$$\delta y_i^n = \mu \cdot \sum_{j,m} \theta_i^{njm}(\tau) Y_{jm}(\theta, \phi), \qquad \theta_i^{njm}(0) = 0, \qquad i = 1, \dots, 6.$$

• E.g. the lowest order instabilities (j = 1, 2) couple to all the modes (having different j's) of the first order...

イロト イロト イヨト イヨト ヨー わへの

Section 6

Conclusions

・ロト・日本・モート・モート モークへの
52 / 53

• The spherically symmetric membrane in SO(3) is integrable (Axenides-Floratos-GL, 2017a).

Conclusions

- The spherically symmetric membrane in SO(3) is integrable (Axenides-Floratos-GL, 2017a).
- Radial & angular perturbation analysis for the elliptic SO(3) membrane was carried out in (Axenides-Floratos-GL, 2017a, 2017b). Found instabilities...

Conclusions

- The spherically symmetric membrane in SO(3) is integrable (Axenides-Floratos-GL, 2017a).
- Radial & angular perturbation analysis for the elliptic SO(3) membrane was carried out in (Axenides-Floratos-GL, 2017a, 2017b). Found instabilities...
- Radial & angular perturbation analysis in the SO(3) × SO(6) case in (Axenides-Floratos-GL, 2017a, 2017b). Studied instabilities...

Conclusions

- The spherically symmetric membrane in SO(3) is integrable (Axenides-Floratos-GL, 2017a).
- Radial & angular perturbation analysis for the elliptic SO(3) membrane was carried out in (Axenides-Floratos-GL, 2017a, 2017b). Found instabilities...
- Radial & angular perturbation analysis in the SO(3) × SO(6) case in (Axenides-Floratos-GL, 2017a, 2017b). Studied instabilities...
- Analysis of higher orders in perturbation theory... instability cascade... (Axenides-Floratos-Katsinis-GL, 2020, 2021).

Conclusions

- The spherically symmetric membrane in SO(3) is integrable (Axenides-Floratos-GL, 2017a).
- Radial & angular perturbation analysis for the elliptic SO(3) membrane was carried out in (Axenides-Floratos-GL, 2017a, 2017b). Found instabilities...
- Radial & angular perturbation analysis in the SO(3) × SO(6) case in (Axenides-Floratos-GL, 2017a, 2017b). Studied instabilities...
- Analysis of higher orders in perturbation theory... instability cascade... (Axenides-Floratos-Katsinis-GL, 2020, 2021).

Ευχαριστώ!