

**RICAP'07 – Roma International Conference on
Astroparticle Physics**

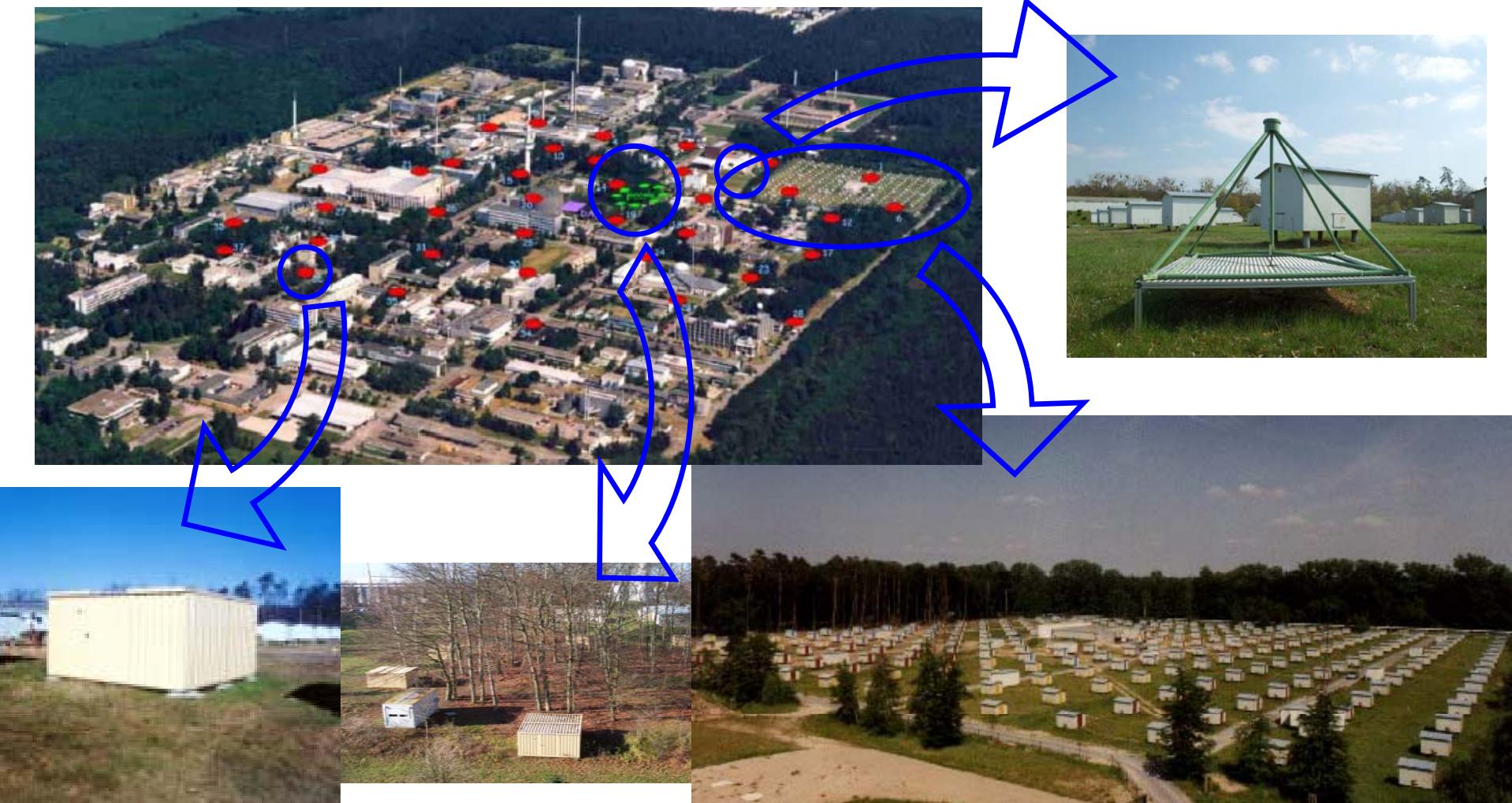
KASCADE-Grande: an overview and first results

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KASCADE-Grande

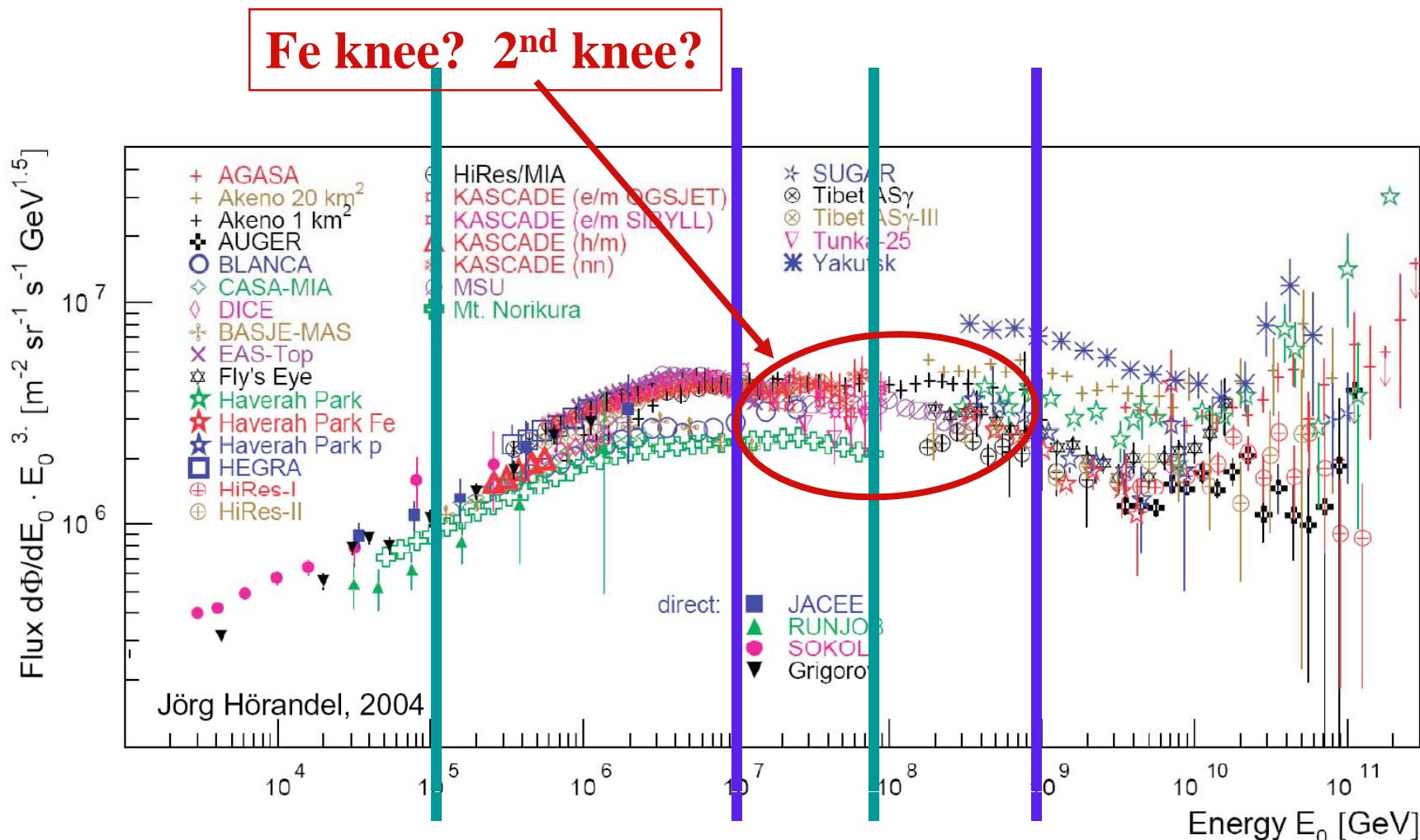
= KArlsruhe Shower Core and Array DEtector + Grande
and LOPES

Measurements of air showers in the energy range $E_0 = 100 \text{ TeV} - 1 \text{ EeV}$



KASCADE: $10^{14} - 8 \cdot 10^{16}$ eV

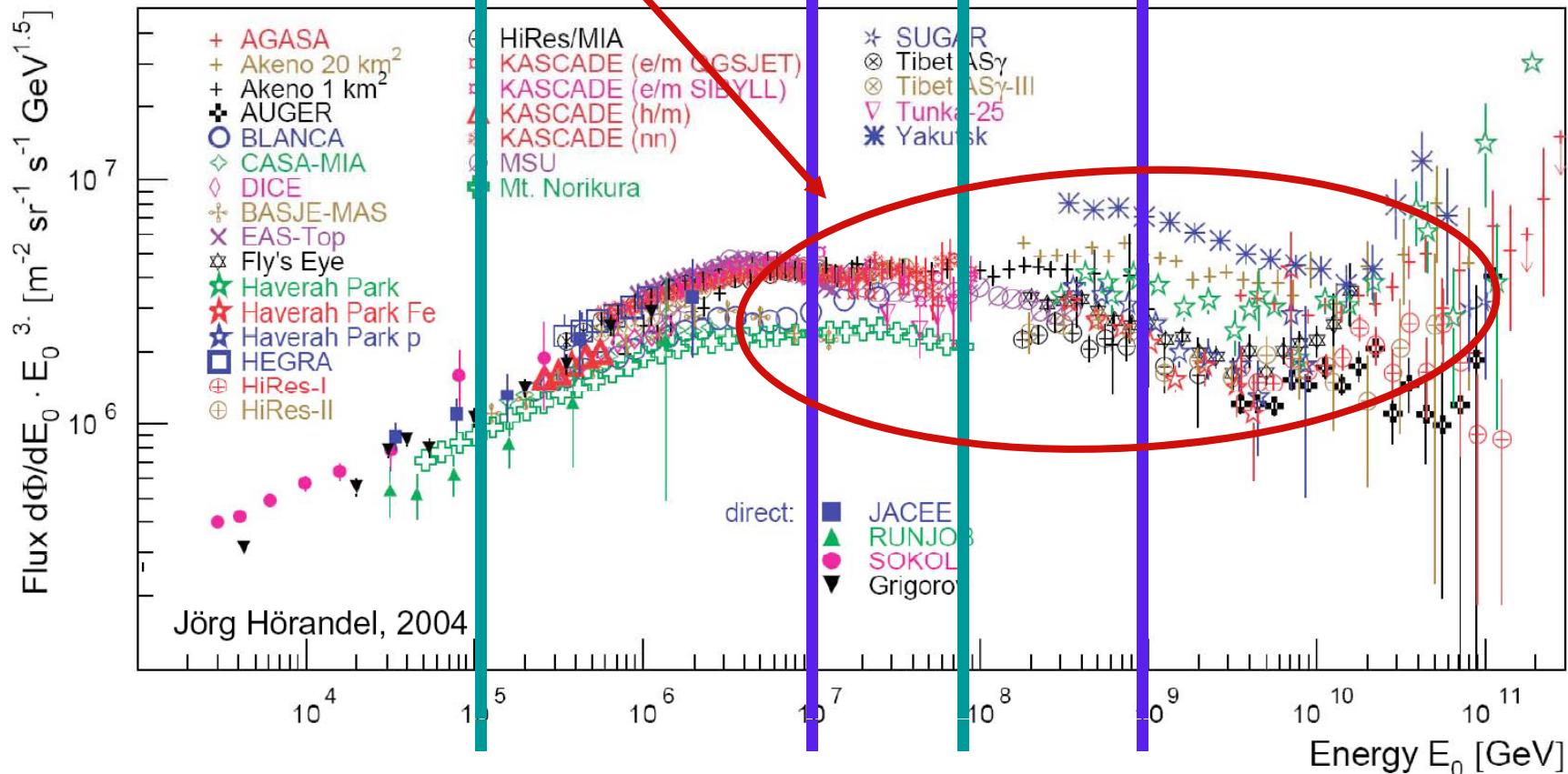
-Grande: $10^{16} - 10^{18}$ eV



KASCADE: $10^{14} - 8 \cdot 10^{16}$ eV

-Grande: $10^{16} - 10^{18}$ eV

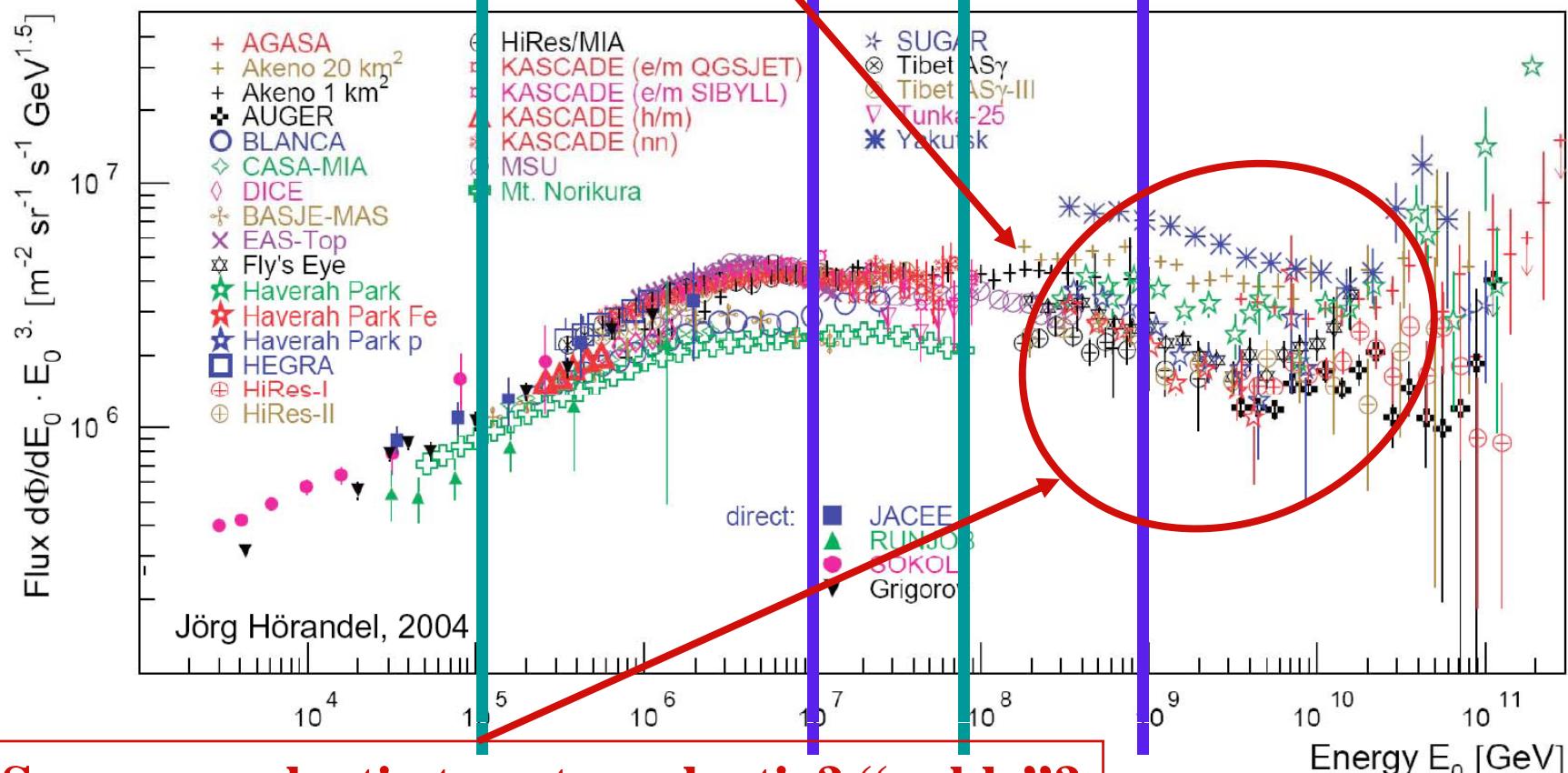
Hadronic interactions



KASCADE: 10^{14} - $8 \cdot 10^{16}$ eV

-Grande: 10^{16} - 10^{18} eV

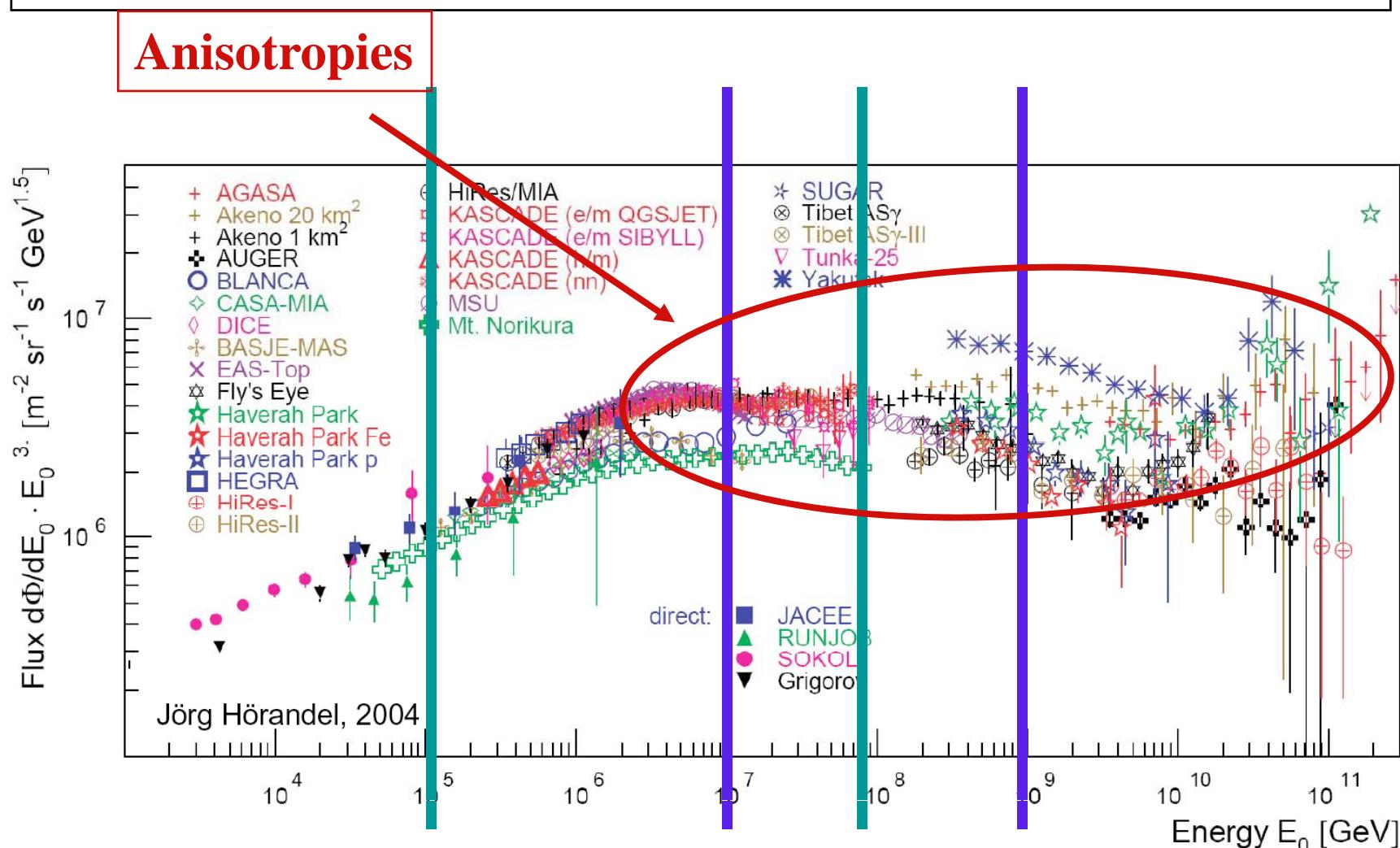
Technical: energy calibration



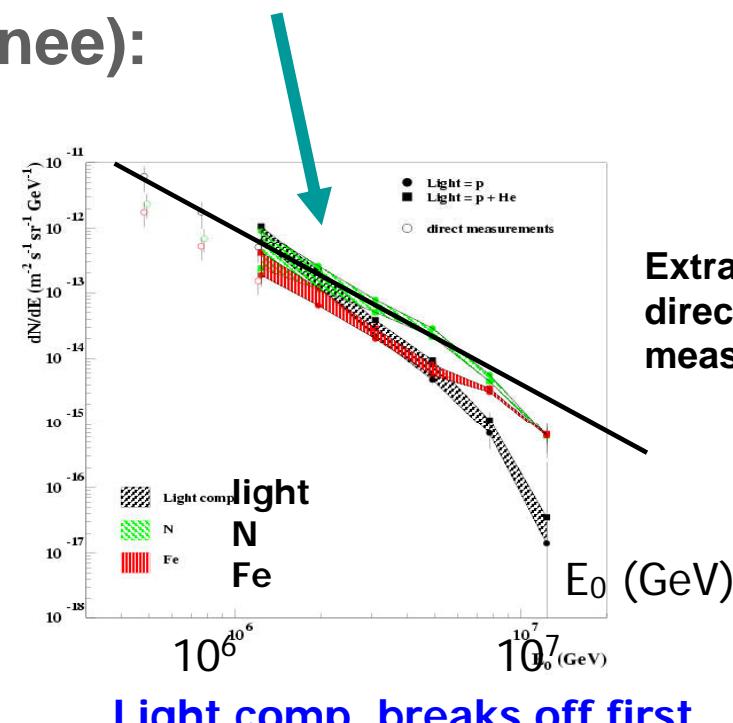
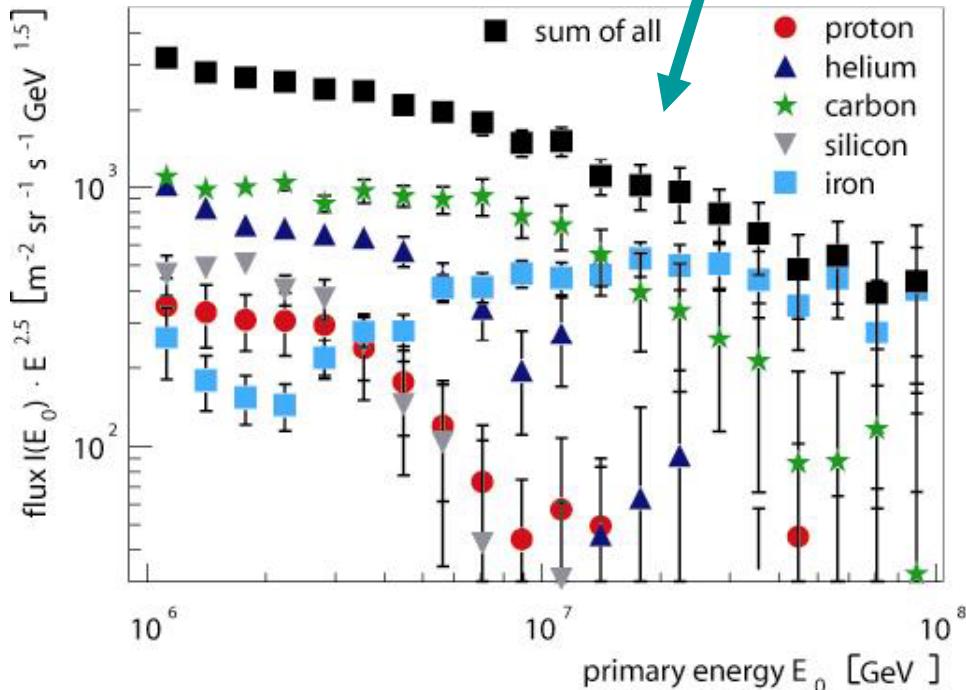
Sources: galactic to extragalactic? “ankle”?

KASCADE: $10^{14} - 8 \cdot 10^{16}$ eV

-Grande: $10^{16} - 10^{18}$ eV



Summary KASCADE and EAS-TOP Results (first knee):

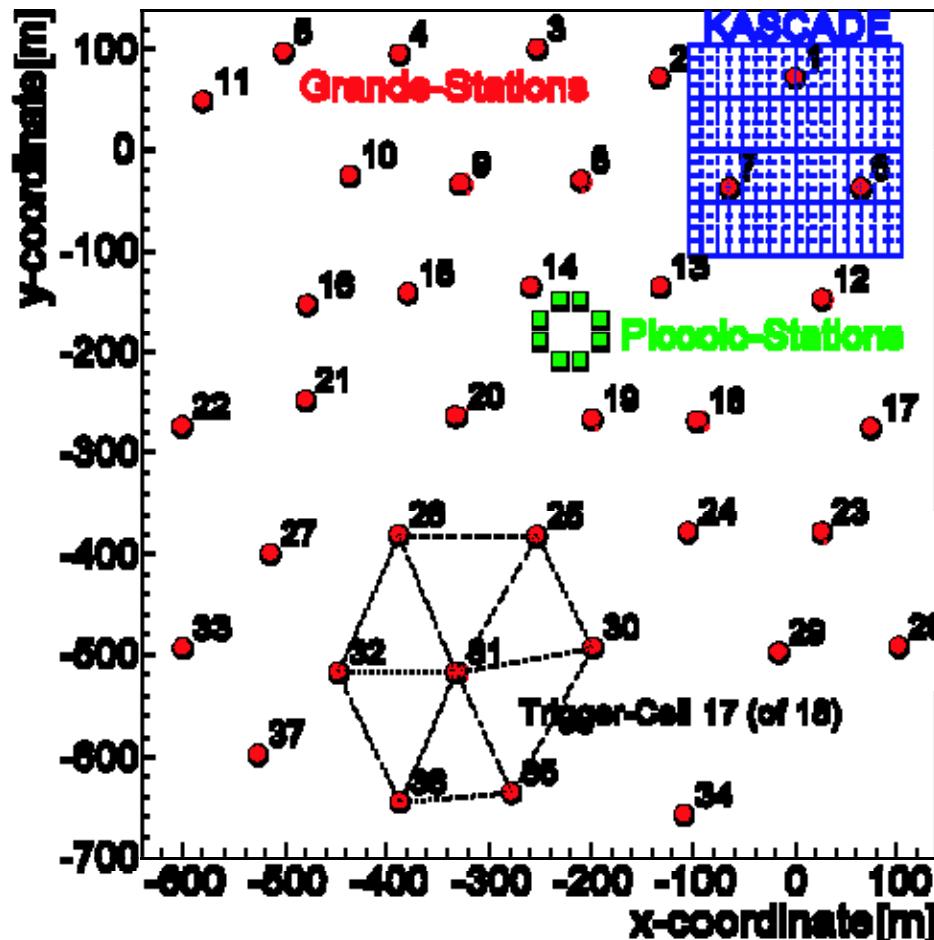


Extrapolation
direct
measurements

Light comp. breaks off first

-) Analysis: Correlation studies are required (→ multiparameter measurements needed)
(Analyzing mean values of data and simulations appears inadequate)
-) Knee is due to decrease in flux in light primaries! (model independent; most experiments)
-) How precise are the models ? (no new physics needed, compare proton spectrum)
-) Distinguishing between astrophysical models (Investigation of Anisotropy for different primaries)
-) Knee position dependence: $\propto Z$ or $\propto A$?

The experimental set-up



Detector	Detected EAS component	Sensitive area (m ²)
Grande	Charged particles	37x10
Piccolo	Charged particles	8x10
KASCADE array e/γ	Electrons, γ	490
KASCADE array μ	Muons ($E\mu^{\text{th}}=230 \text{ MeV}$)	622
MTD	Muons (Tracking) ($E\mu^{\text{th}}=800 \text{ MeV}$)	3x128
MWPCs/LSTs	Muons ($E\mu^{\text{th}}=2.4 \text{ GeV}$)	3x129
LOPES 30	Radio	
Trigger Plane	Muons ($E\mu^{\text{th}}=490 \text{ MeV}$)	208
Calorimeter	Hadrons	9x304

The strength of KASCADE-Grande is the multi observables information
Requirement: keep a good accuracy till 10^{18} eV

The KASCADE Array

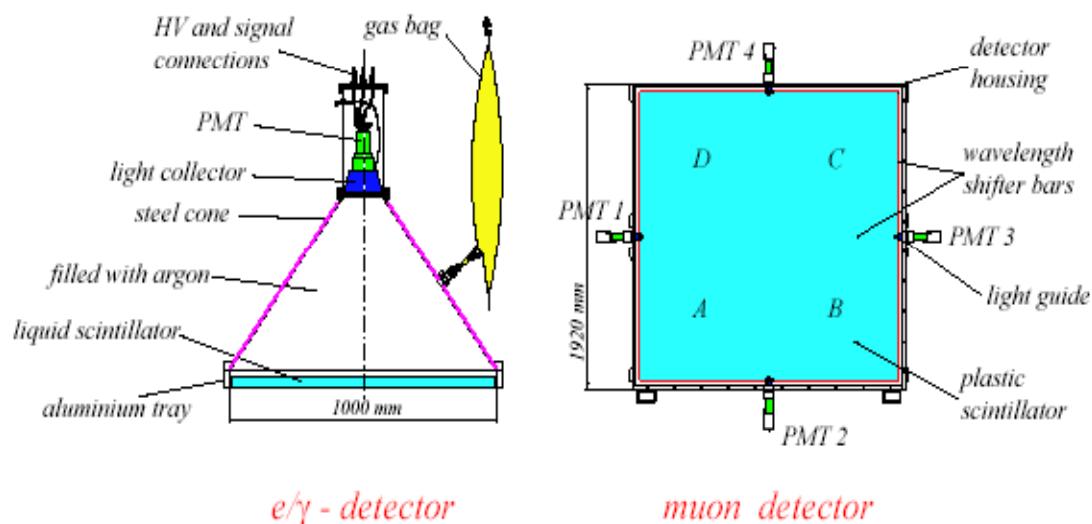


*e/γ - detector
(liquid scintillator)*

lead/iron absorber

*muon detector
(plastic scintillator)*

- 252 detectors
- 3.2 m² each
- 13 m distant
- 200 x 200 m²
- e/γ : liquid, 48 mm
- μ : plastic, 30 mm

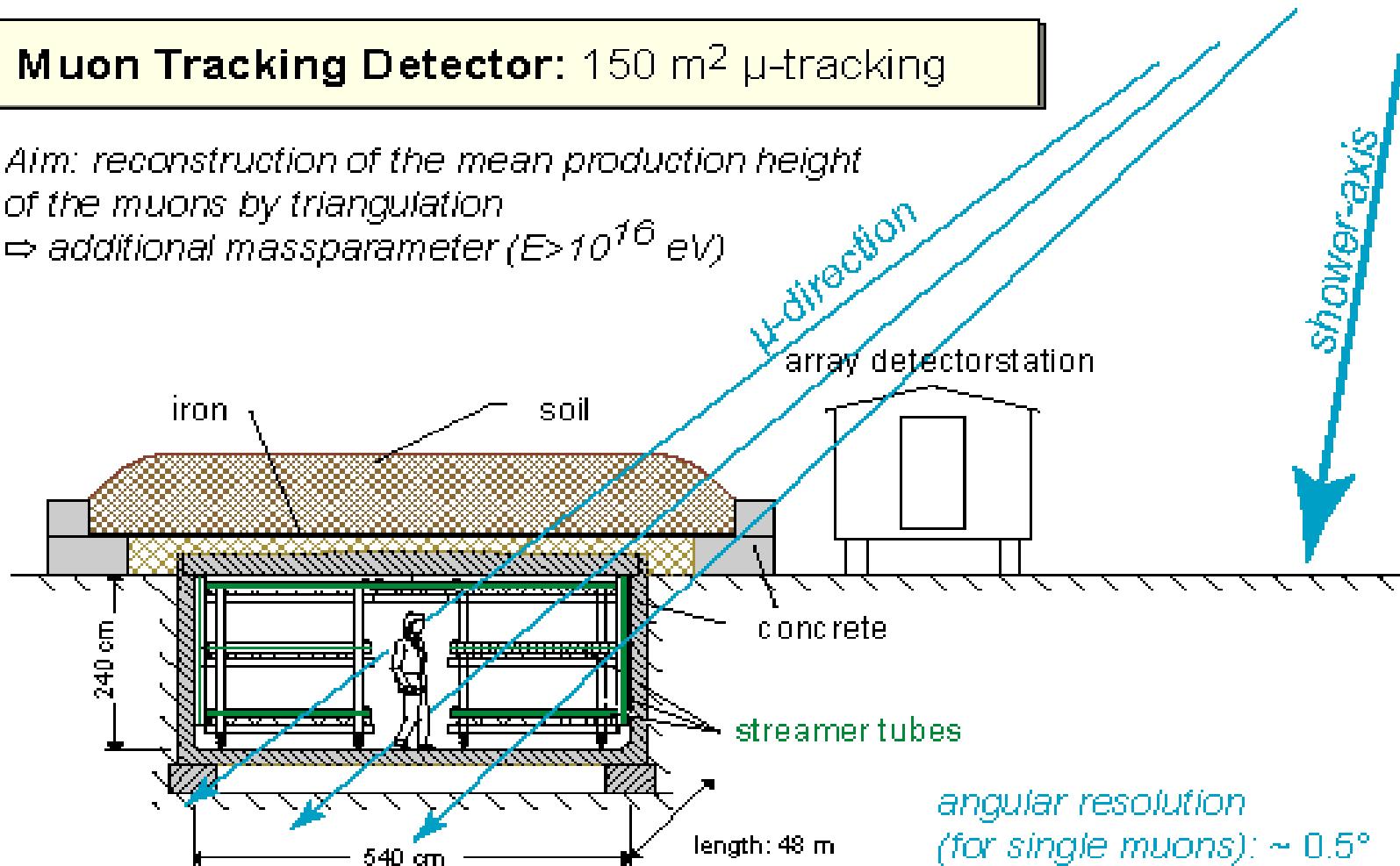


The muon tunnel

Muon Tracking Detector: 150 m^2 μ -tracking

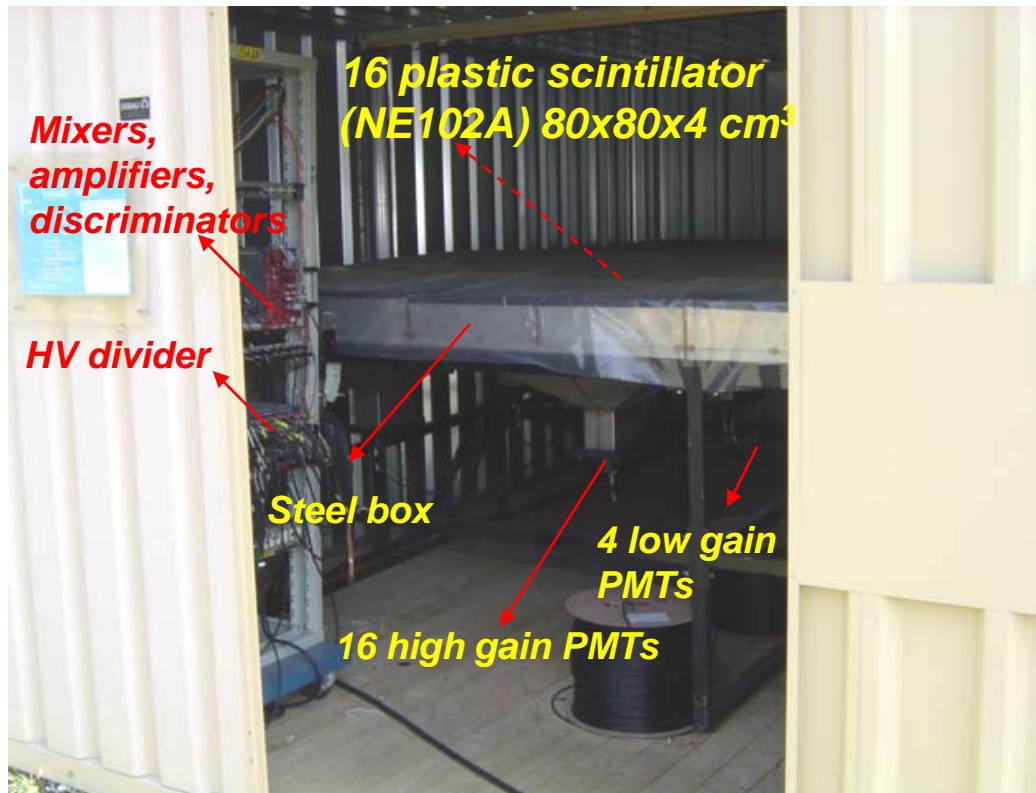
Aim: reconstruction of the mean production height
of the muons by triangulation

⇒ additional massparameter ($E > 10^{16} \text{ eV}$)



The GRANDE Array

The detector station



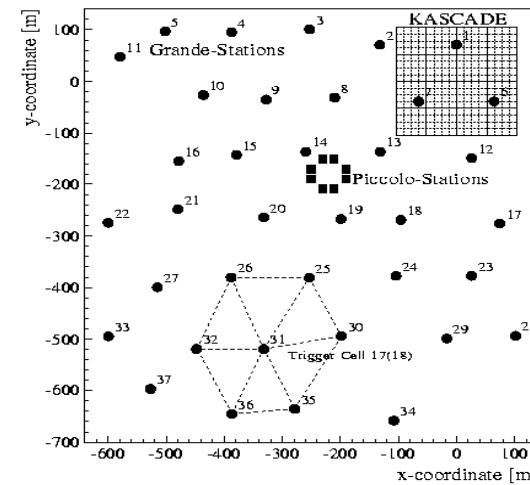
37 Stations of plastic scintillators 10 m²
140 m average distance
0.5 km² total surface



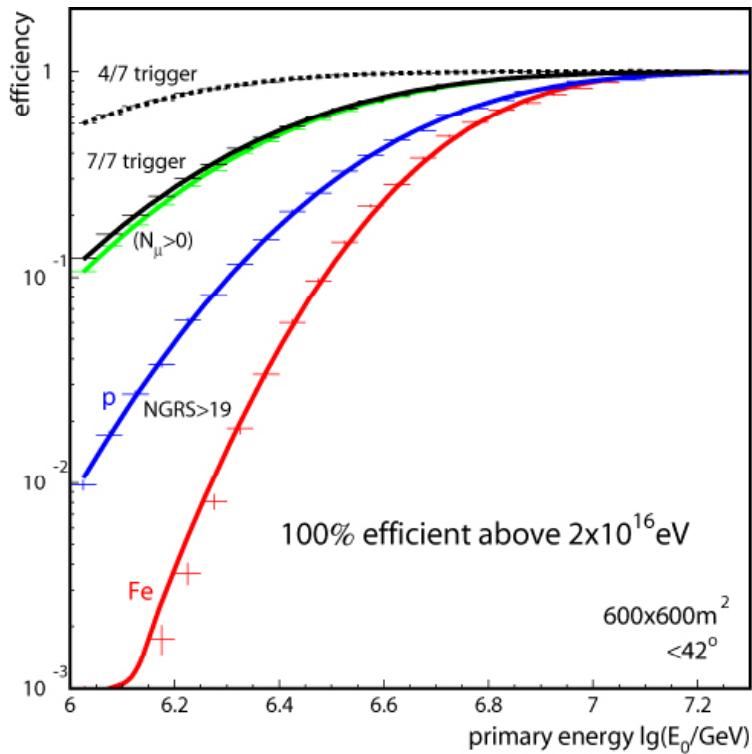
Peak sensing ADC:
Particle density range: 10^4 m.i.p./10m²
TDC: 0.8 ns resolution
Flash ADC...

KASCADE-Grande : Reconstruction steps

- 1) core position and angle-of-incidence
from Grande array data**
- 2a) shower size (charged particles)
from Grande array data**
- 2b) muon number
from KASCADE muon detectors**
- 3) electron number
from Grande by subtraction of muon content**
- 4) two dimensional size spectrum for the analysis**



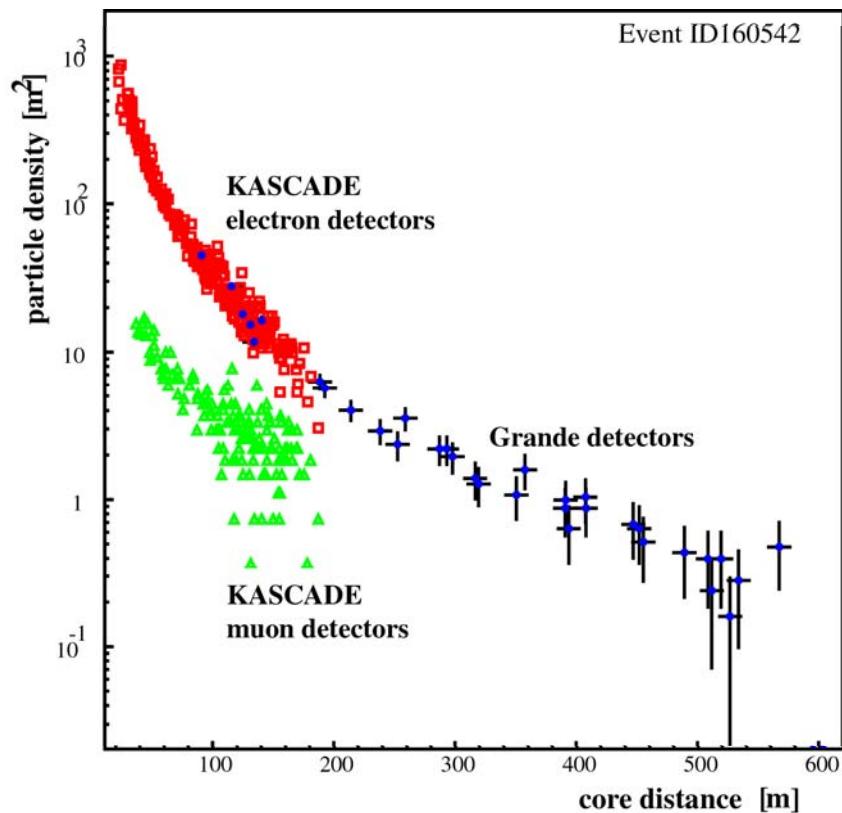
KASCADE-Grande : Efficiency



- Common events
(all detector components)
measured since December 2003
- Trigger: 7 of 7 stations at one
of 18 hexagons

KASCADE-Grande : Single event measurement

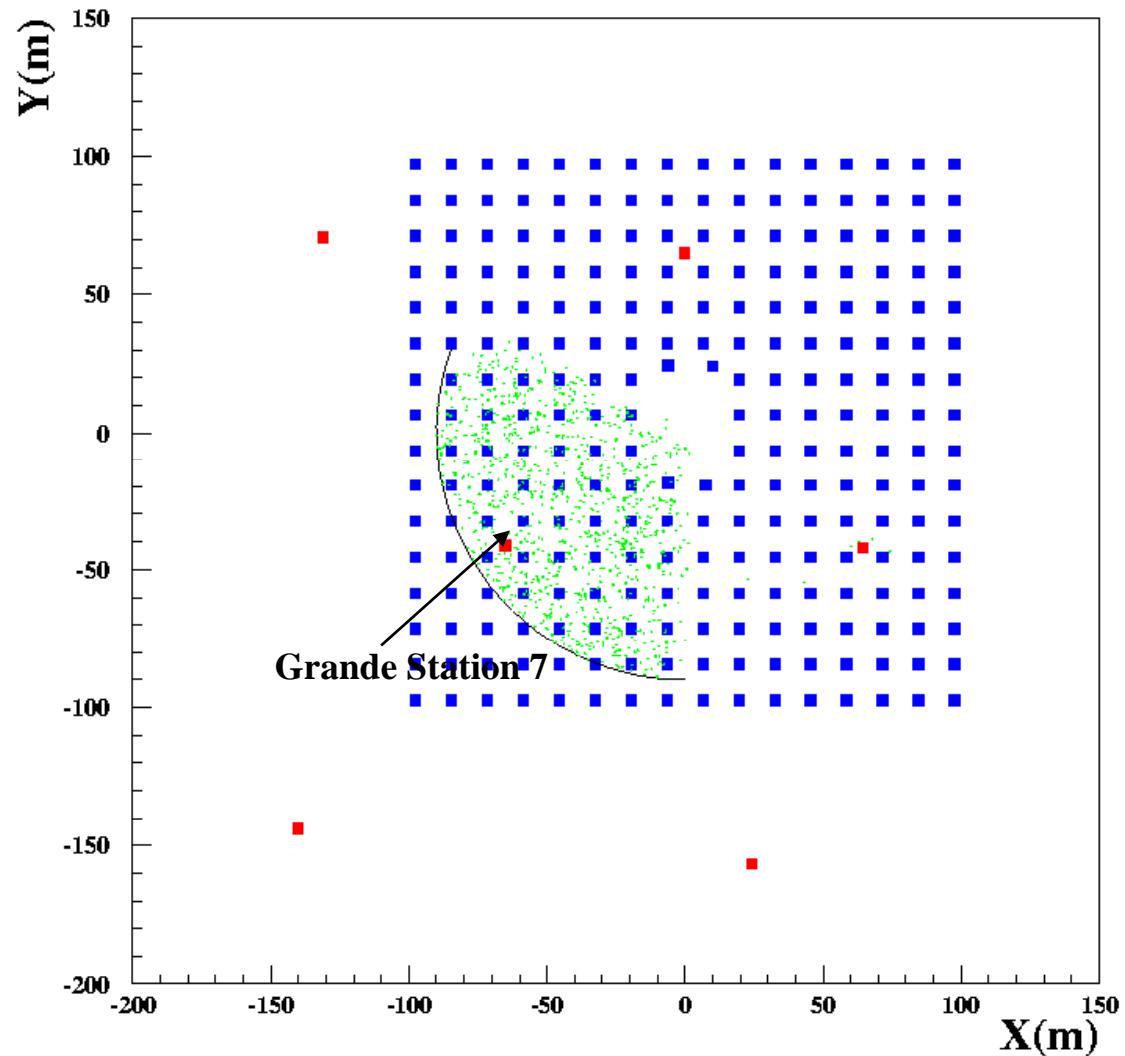
lateral distribution of a single event measured by KASCADE-Grande:
 $E_0 \approx 2 \cdot 10^{17}$ eV, $\Theta = 33^\circ$



Reconstruction : comparison Grande – KASCADE array

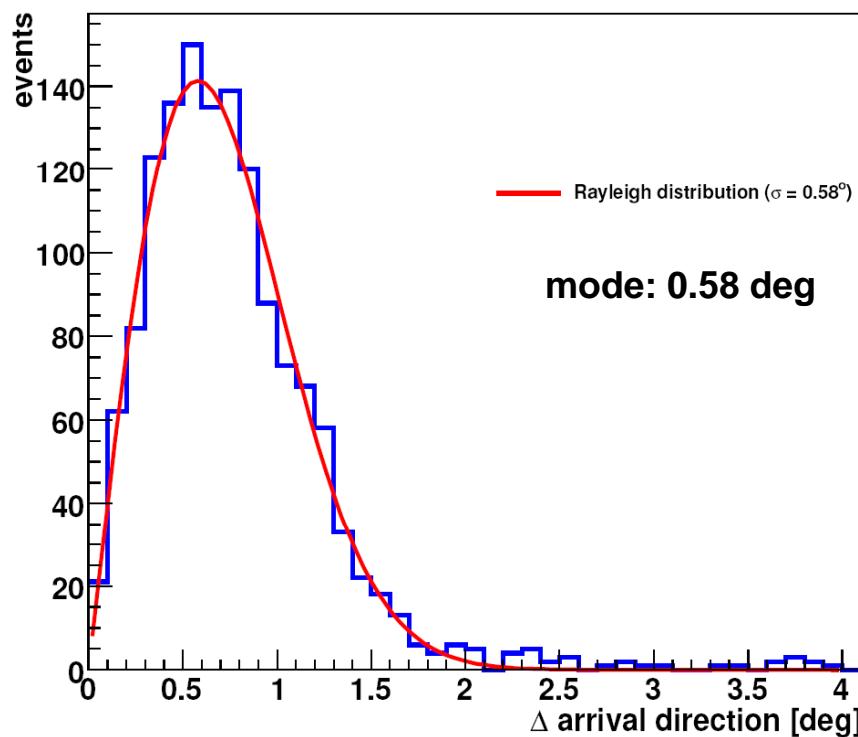
Event Selection:

- At least 10 out of 32 pmt inside a KASCADE cluster
- 7/7 coincidence in Grande
- ΔE_{\max} in Grande Station 7
- $0 \text{ m} < r_{\text{KA}} < 90 \text{ m}$
- $\log N e_{\text{KA}} > 6.0$
- $0.4 < s_{\text{Grande}} < 1.4$

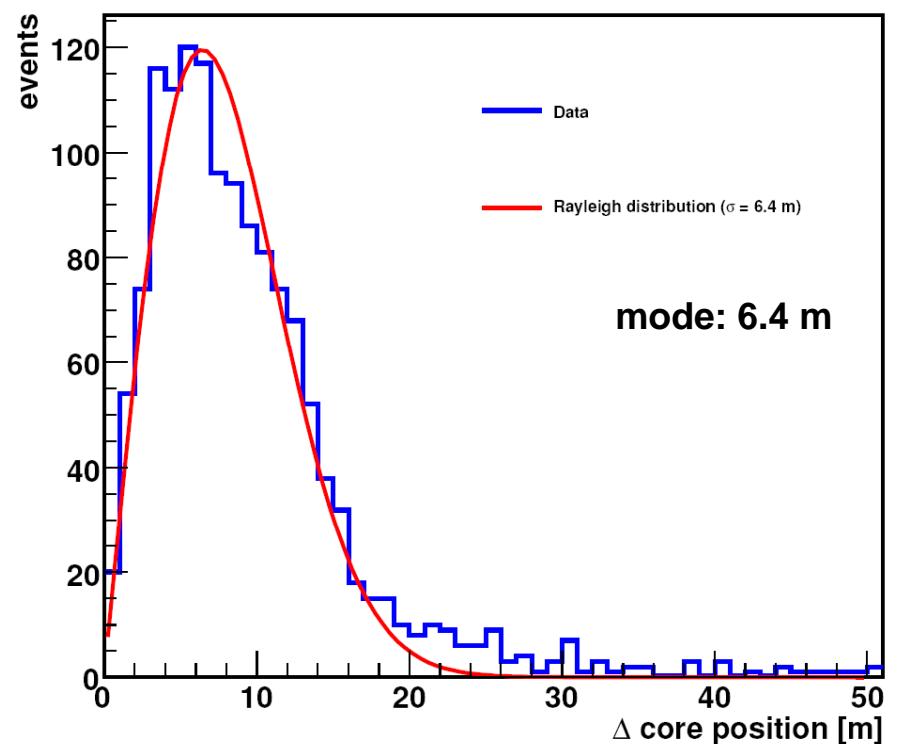


Reconstruction : comparison Grande – KASCADE array

Arrival direction



Core position



$\omega_k - \omega_g$ (deg)

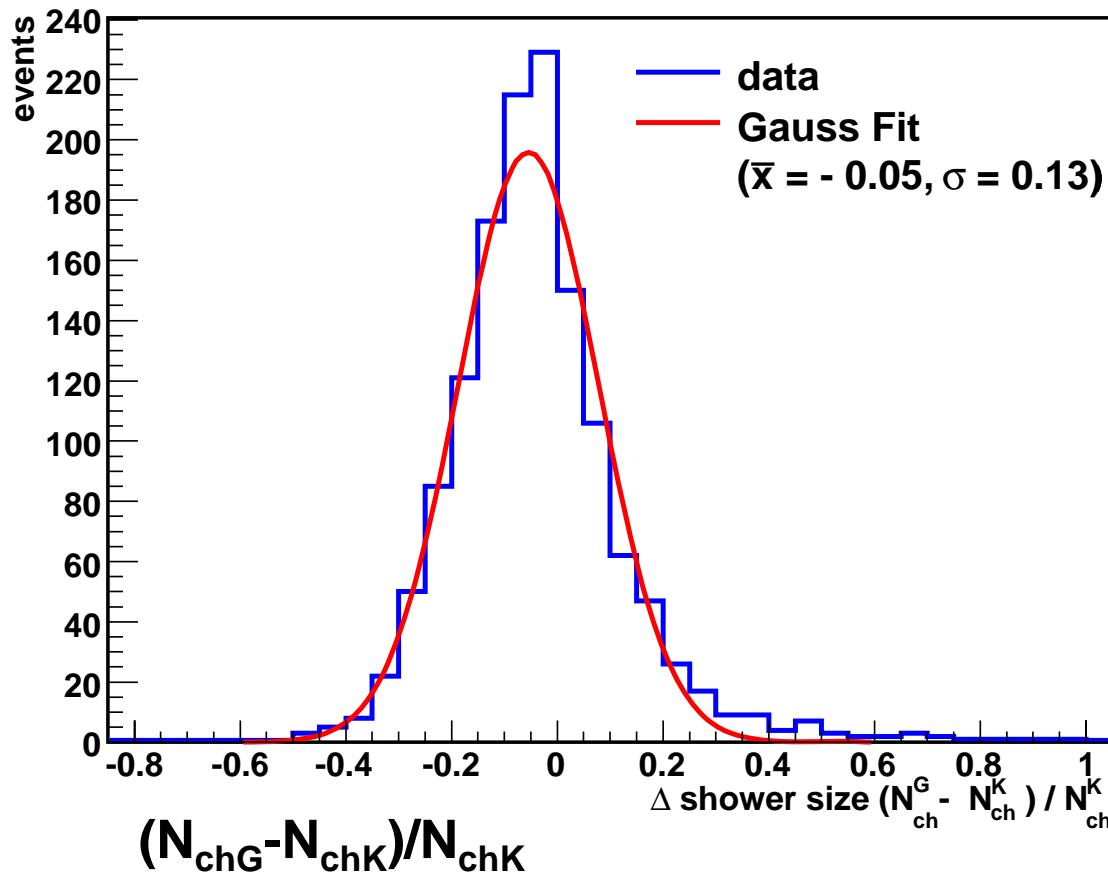
Δr (core position)

Federico Di Pierro et al. – KASCADE-Grande coll., ICRC (2007)

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Reconstruction : comparison Grande – KASCADE array

Size



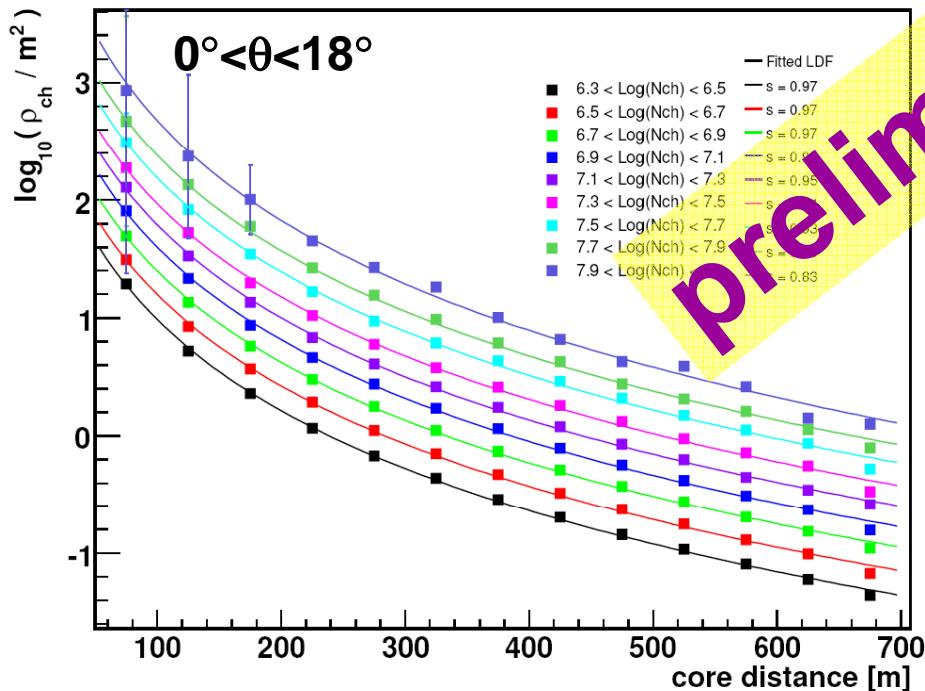
Cross-check between
EAS-TOP and KASCADE

Federico Di Pierro et al. – KASCADE-Grande coll., ICRC (2007)

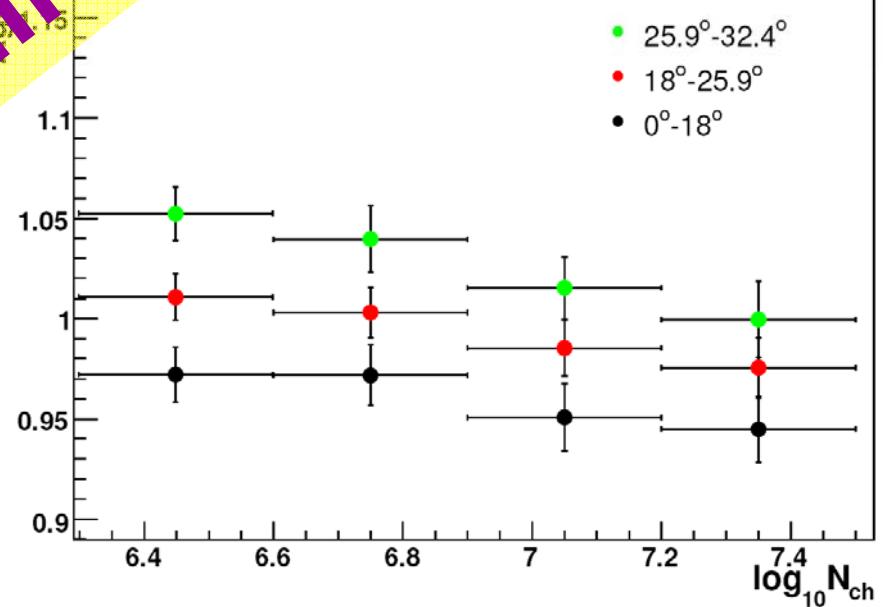
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KASCADE-Grande : lateral distributions

charged particle lateral distribution



s parameter



2-year data, core inside Grande array

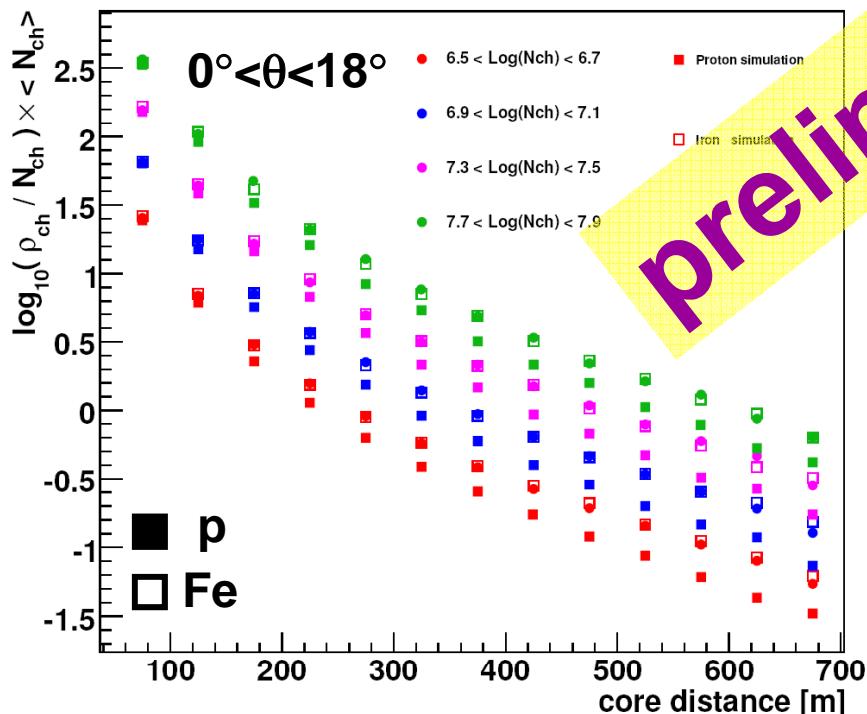
$$\rho_{\text{ch}} = N_{\text{ch}} \cdot C(s) \cdot (r/r_0)^{s-\alpha} \cdot (1+r/r_0)^{s-\beta}$$

$$\alpha=1.5, \beta=3.6, r_0=40\text{m}$$

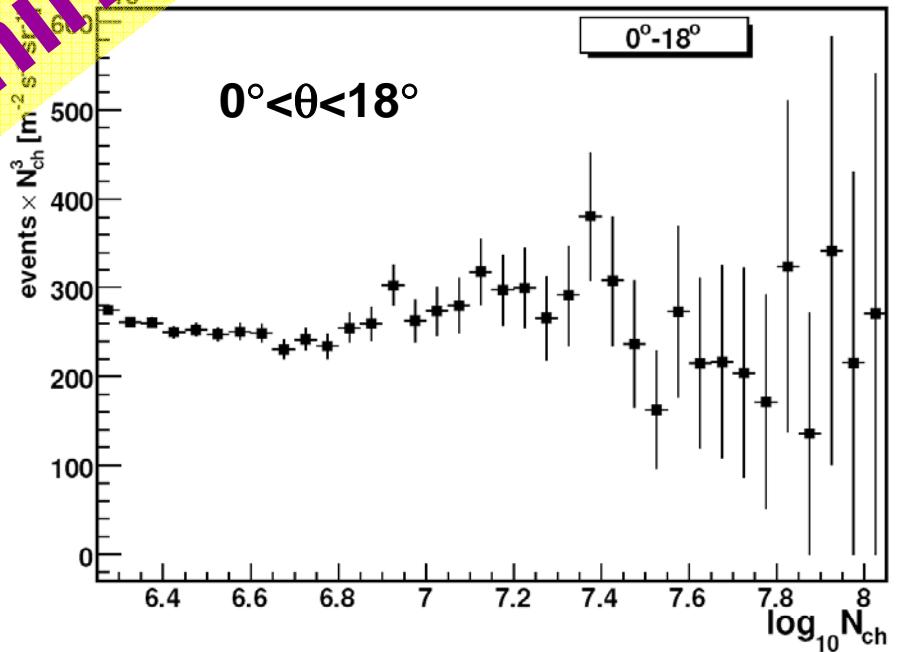
Federico Di Pierro et al. – KASCADE-Grande coll., ICRC (2007)

KASCADE-Grande : lateral distributions and size spectrum

charged particle lateral distribution



Size spectrum parameter



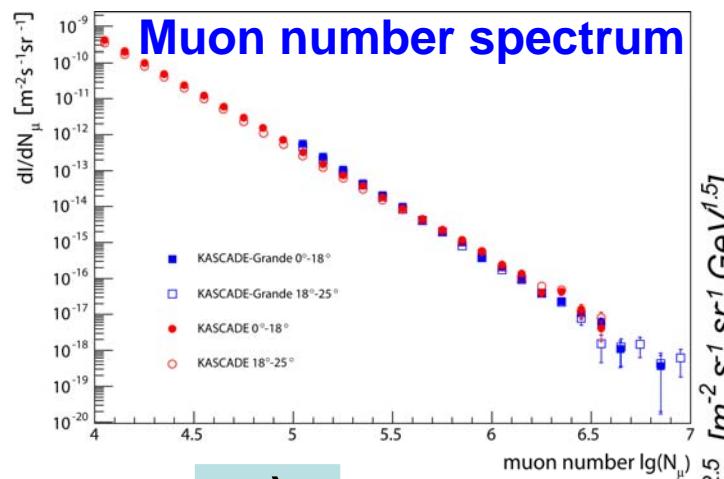
2-year data, core inside Grande array

reconstruction gives reasonable spectra
careful checks of systematic effects in work

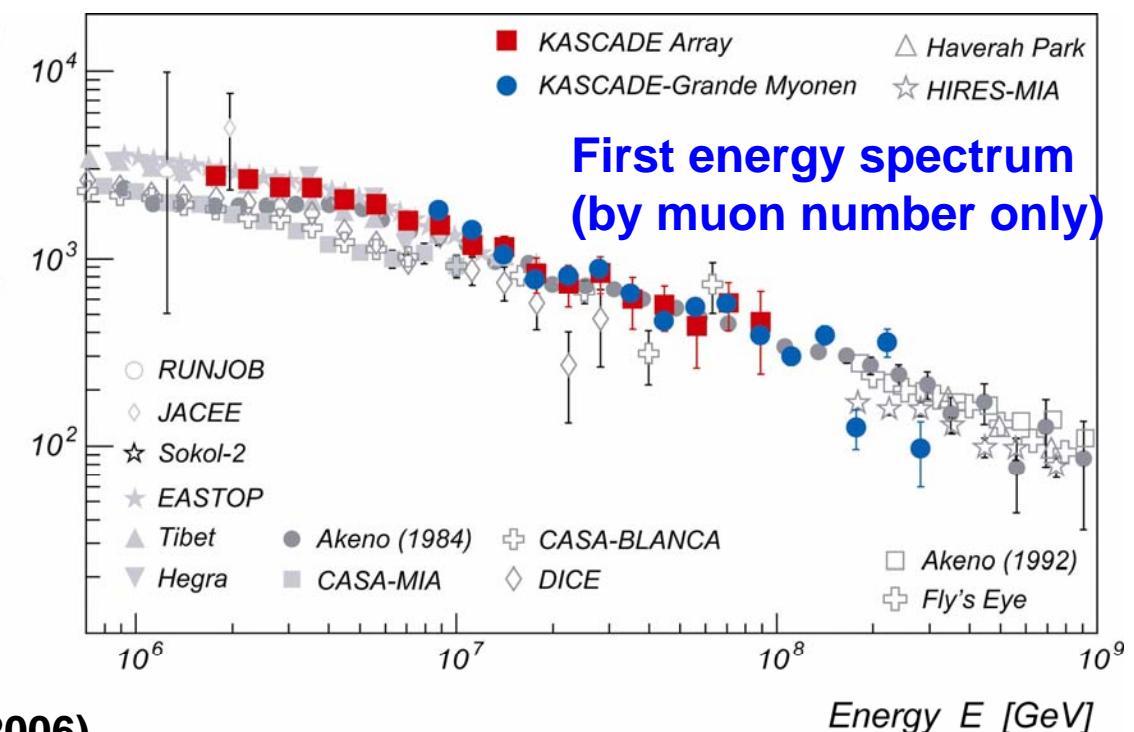
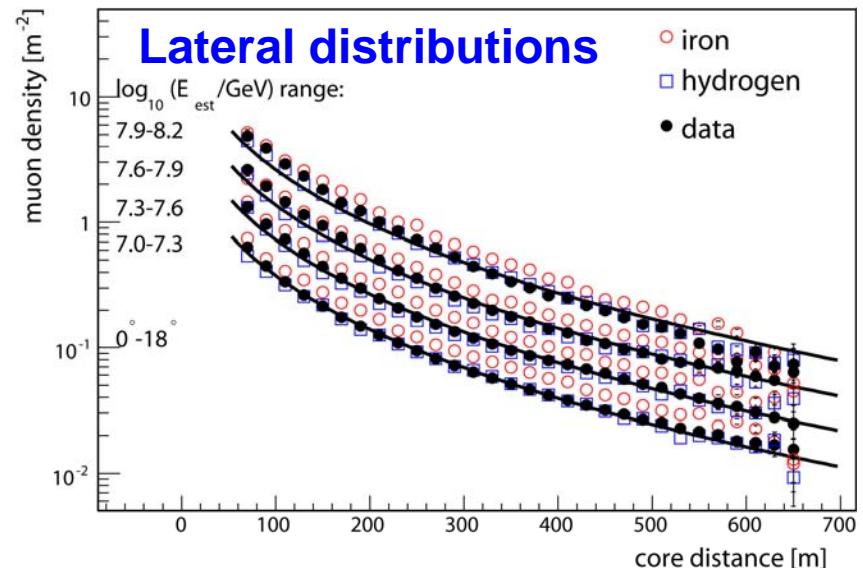
Federico Di Pierro et al. – KASCADE-Grande coll., ICRC (2007)

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Muon lateral distribution number spectrum and 1-dim unfolding



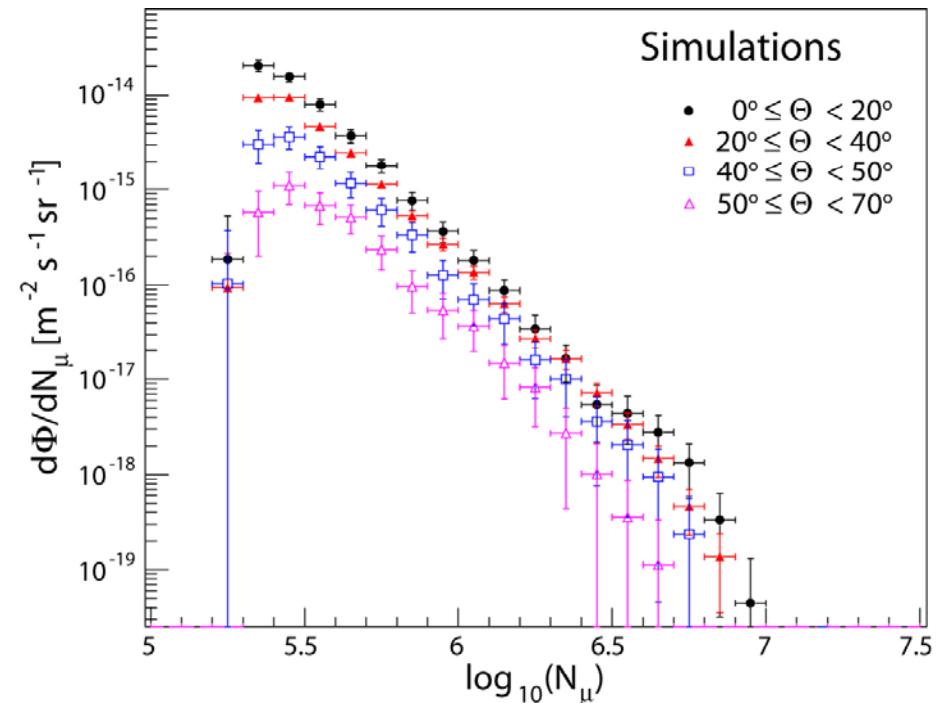
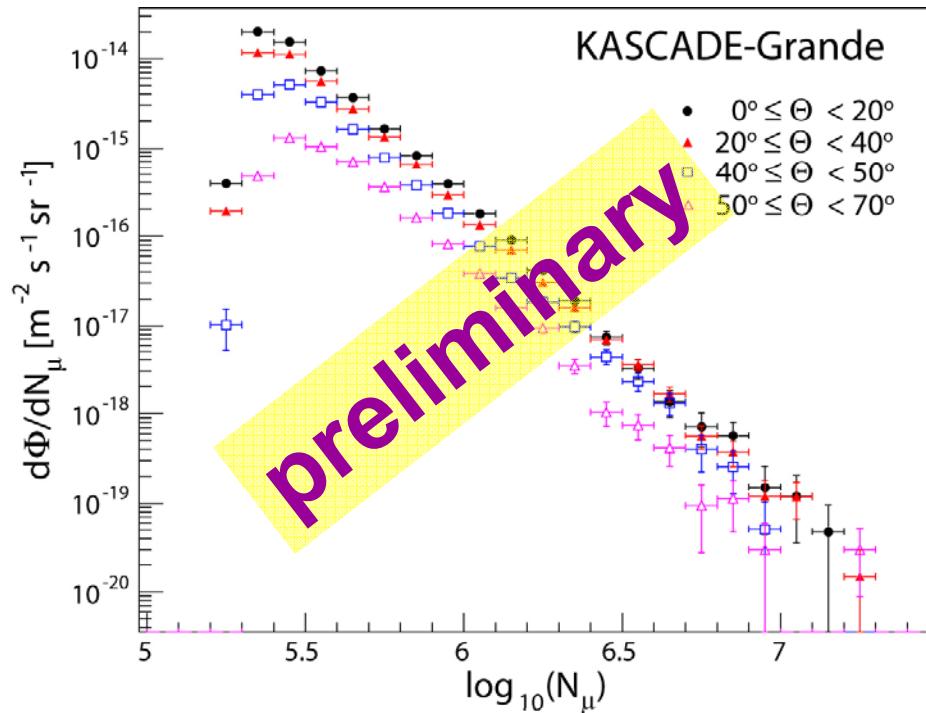
1-dim-
unfolding



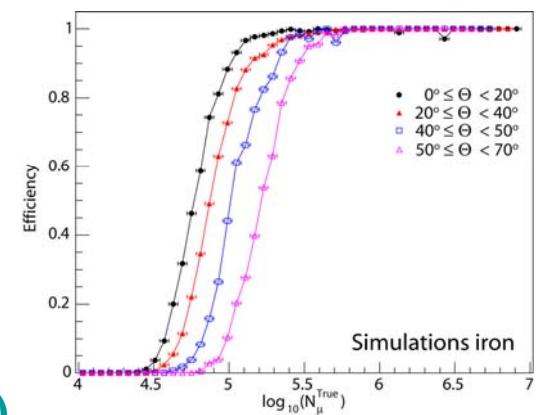
Jurriaan van Buren, PhD thesis (2006)

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KASCADE-Grande : first analyses muon reconstruction at inclined showers

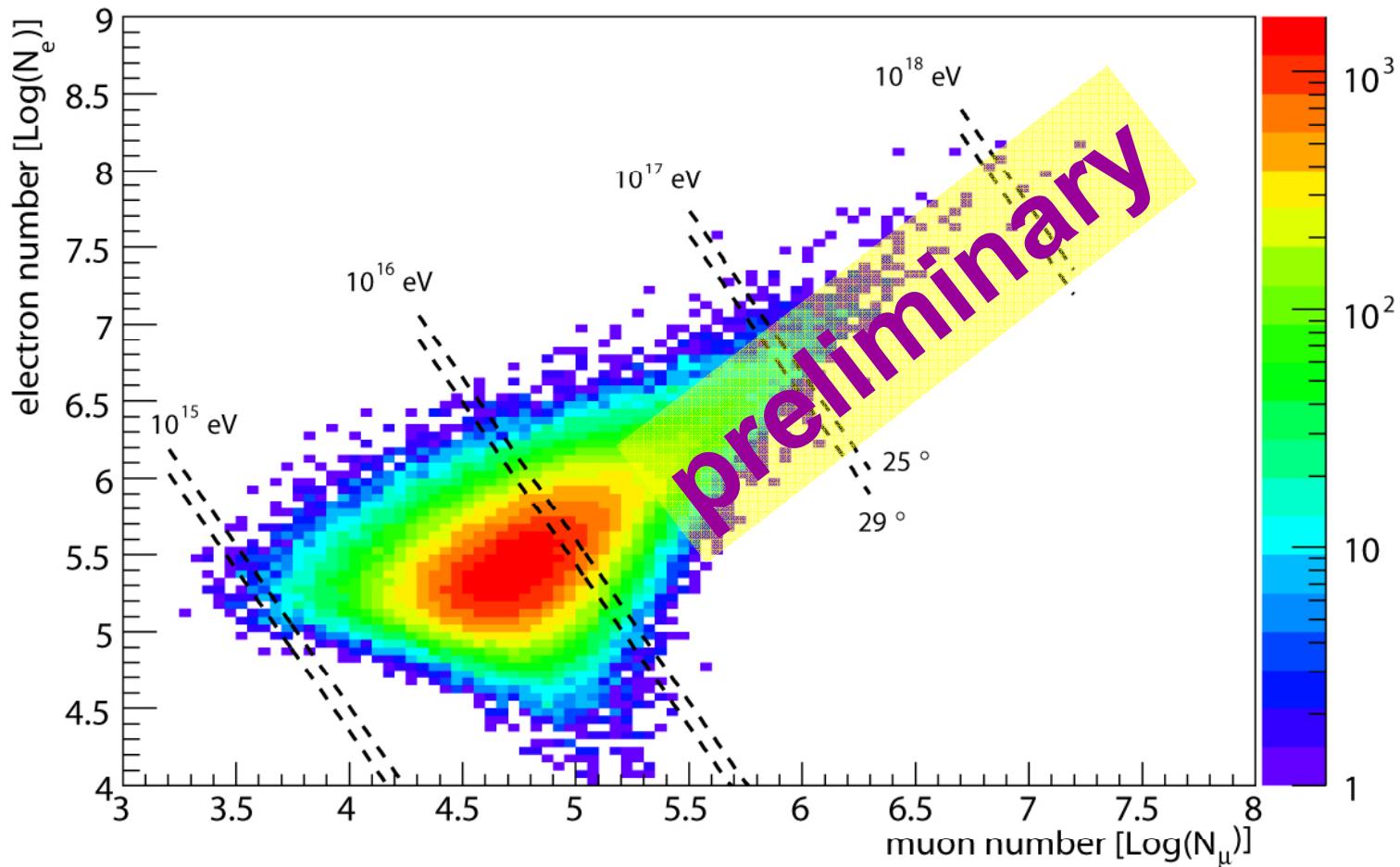


- muon number reconstruction possible up to 70°
- investigation of attenuation of muon component
- model tests
- Increasing of KASCADE-Grande statistics



Juan Carlos Arteaga et al. – KASCADE-Grande coll., ICRC (2007)

2-dimensional analysis (Ne-N μ)

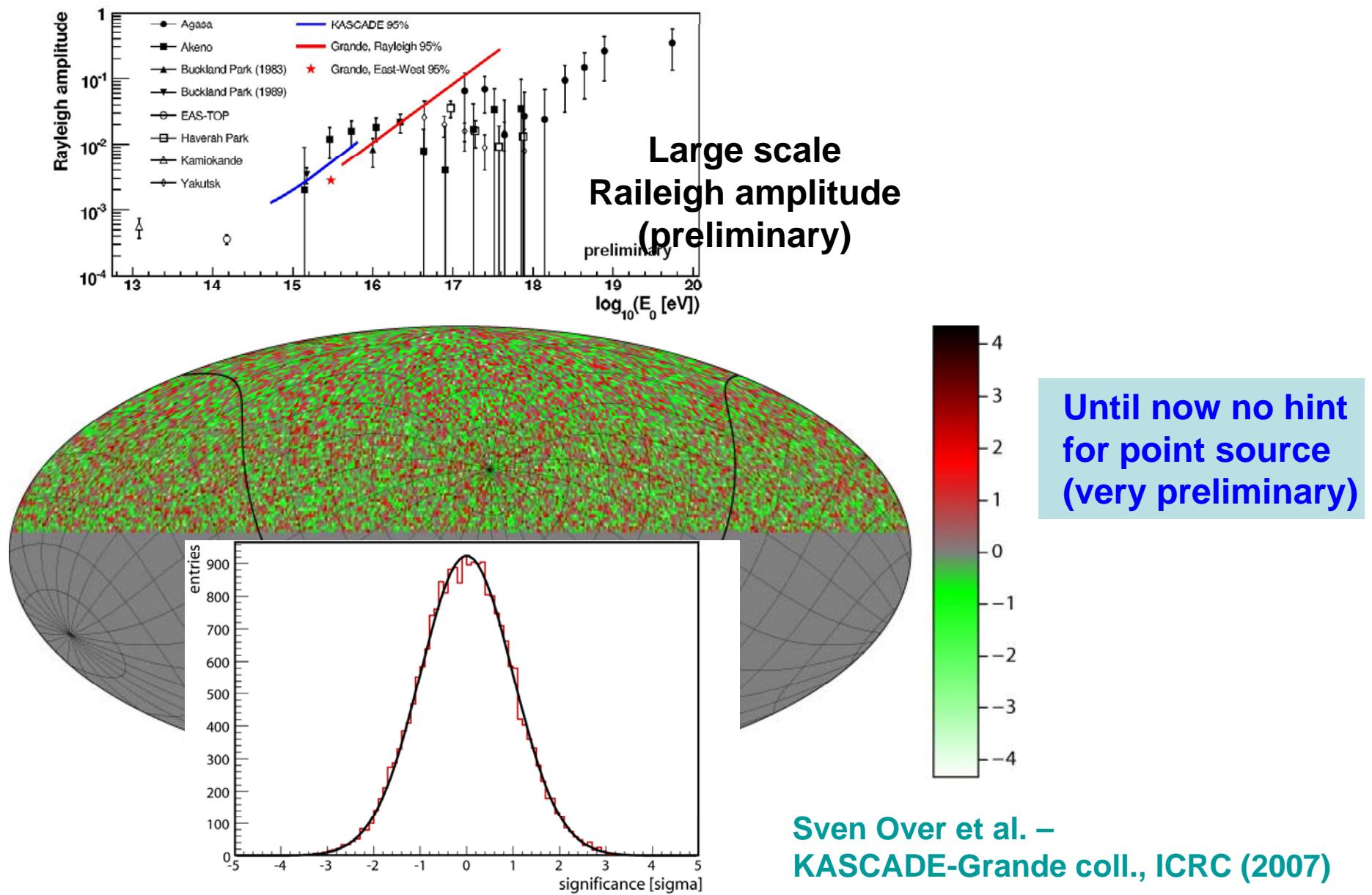


Unfolding of 2-dimensional shower size spectrum possible
→ energy & composition, but still improvements in systematics needed

Fabiana Cossavella et al. – KASCADE-Grande coll., ICRC (2007)

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Anisotropies: first analyses



Conclusions

- KASCADE-Grande is in continuous and stable data taking since 2004
- The detector has been studied and understood
- The performance of the detector looks promising
- First analysis show reasonable preliminary results
- Lot of physics to be explored...



THANK YOU for your attention!

KASCADE-Grande Collaboration

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