

Prototype string for a km³ Baikal neutrino telescope

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Collaboration

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- St.Petersburg State Marine University, Russia.
- Kurchatov Institute, Moscow, Russia.

Outline

1. Introduction
Objectives of prototype string installation
2. Prototype string design
3. Prototype string as a part of NT200+,
modernization of NT200+ data acquisition system
4. Preparative works (2007)
 - prototype of a FADC based system
 - new optical module
 - PM selection for the km3 prototype string

Summary

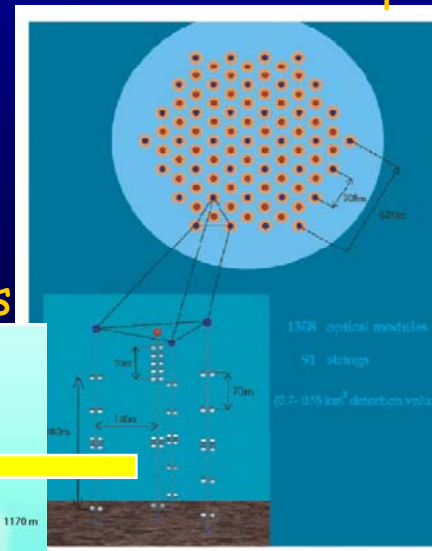
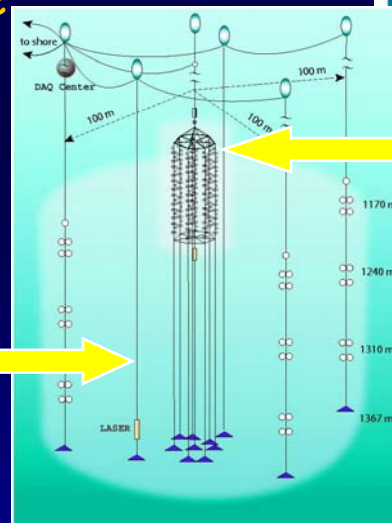
Introduction

Project of km3 Baiaki Neutrino Telescope

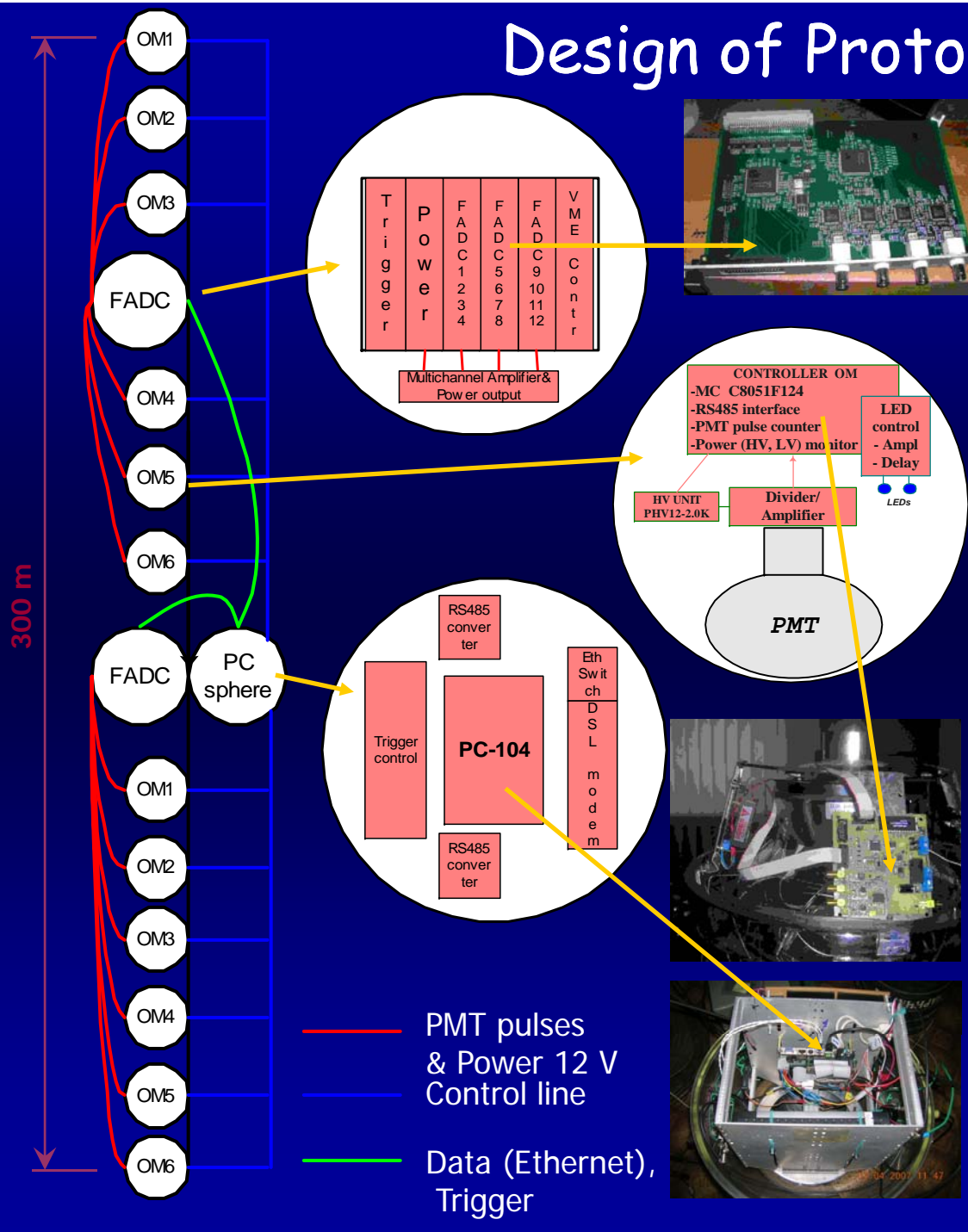
Installation of a "new technology" prototype string as a part of NT200+ (spring 2008)

- Investigations and in-situ tests of basic knots of future detector: optical modules, DAQ system, new cable communications.
- Studies of basic DAQ/Triggering approach for the km3-detector.
- Confrontation of classical TDC/ADC approach with FADC readout.

NT200+
Prototype string current status



Design of Prototype string



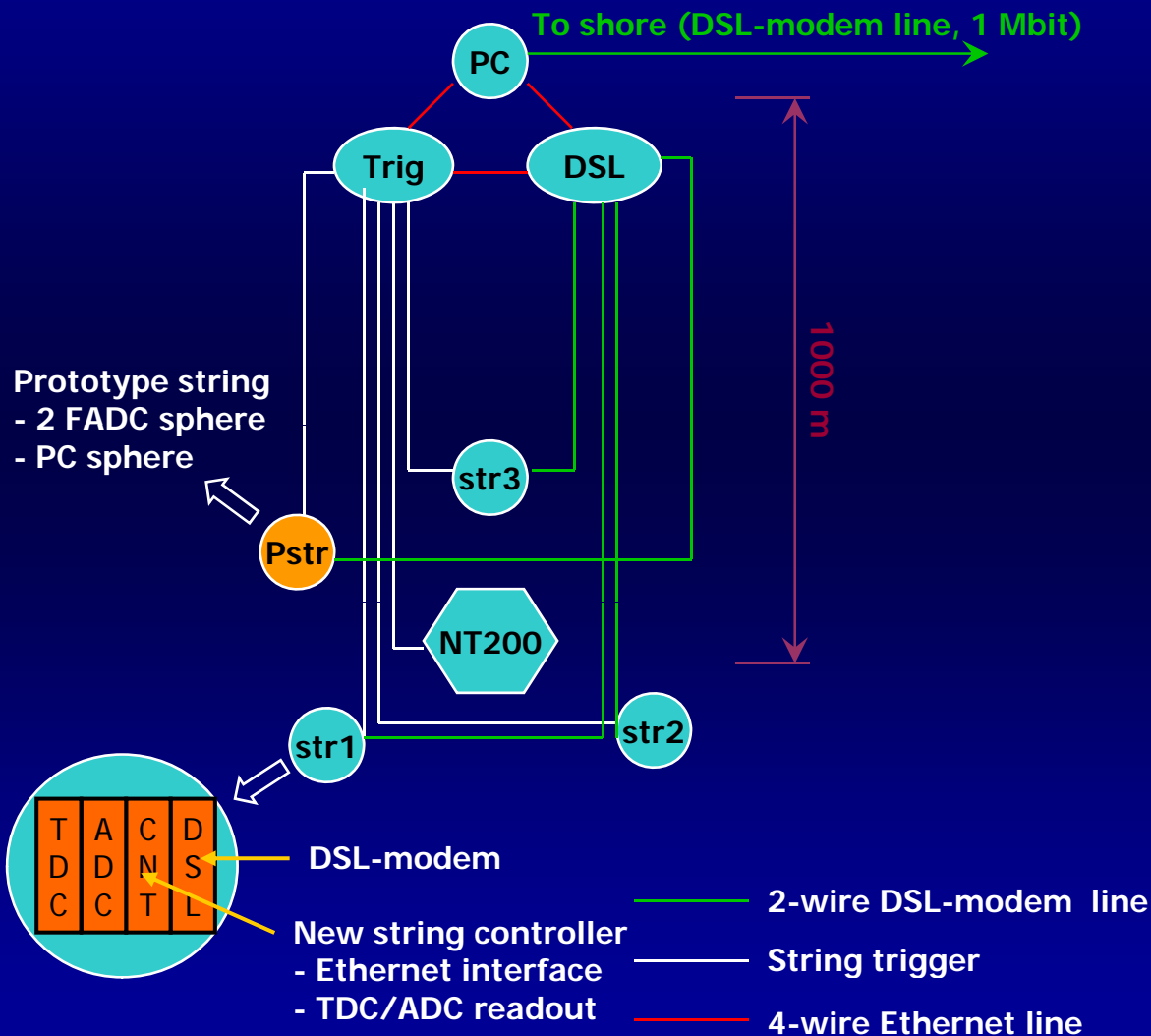
FADC unit is operating now in Tunka detector (astro-ph/0511229)

Basic features

- String lengths ~300 m
- String contains 12..16 OM
- Optical modules contains only PM and control electronics
- 12 bit 200 MHz FADC readout is designed as multi channel separate unit.
- Half-string FADC controllers with ethernet-interface connected to string PC unit
- String PC connected by string DSL-modem to central PC unit

Prototype string as a part of NT200+

Modernization of data acquisition system during expedition 2007



Basic goal of DAQ modernization is increase of the uw-data rate:

- Transmission FADC data;
- Trigger algorithm optimization

2005: data/control TCP/IP connection between the shore station and central uw-PCs.

2007: TCP/IP communication between uw-PCs and string controller.

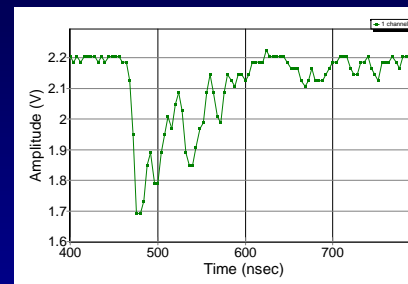
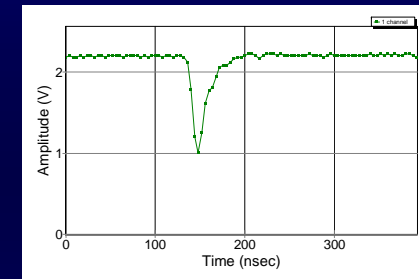
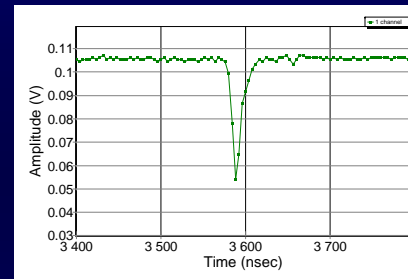
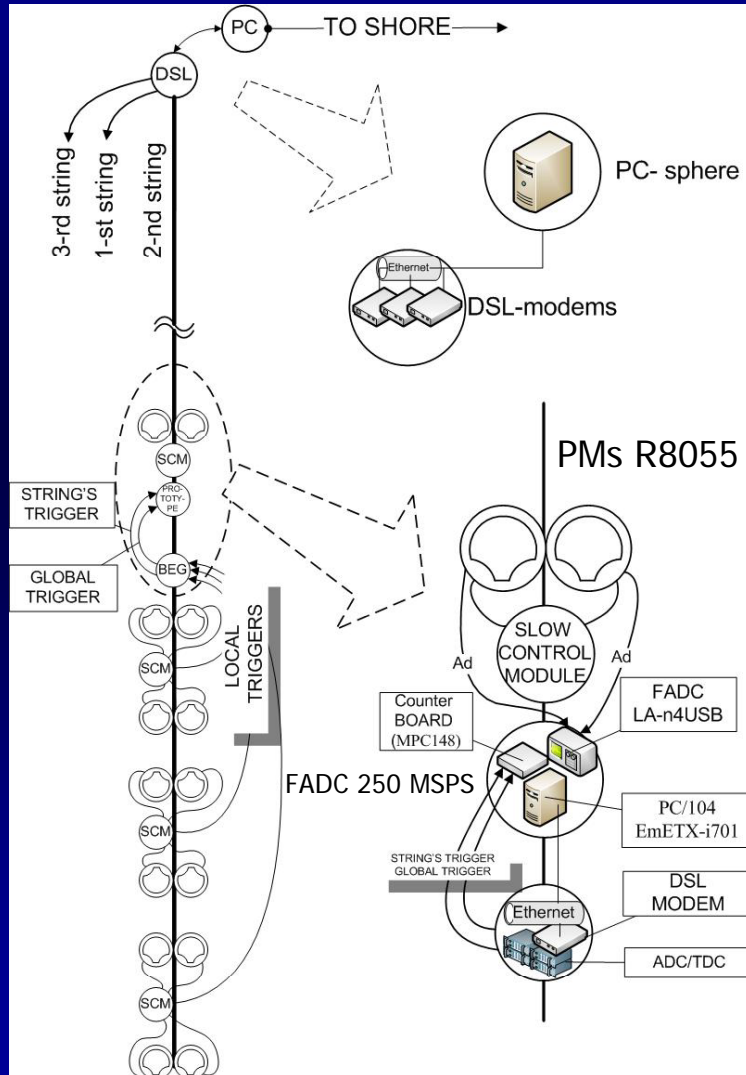
Time synchronization: measuring of each string trigger time with 2 ns precision).

Prototype of FADC based system

2-channel FADC prototype was installed during expedition 2007

Purposes:

- optimal sampling time window
- dynamic range
- obtainable pulse parameter precision
- algorithms for online data handling

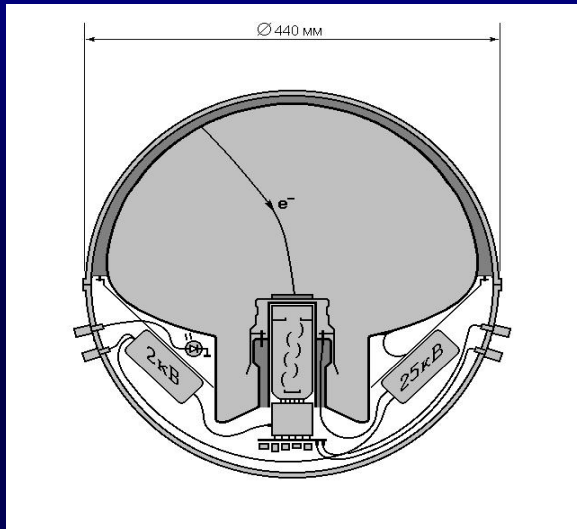


Examples of FADC pulses for different classes of events:

1. One p.e. noise hit
2. A muon trigger (multi-p.e.)
3. Backward illumination by a calibration laser

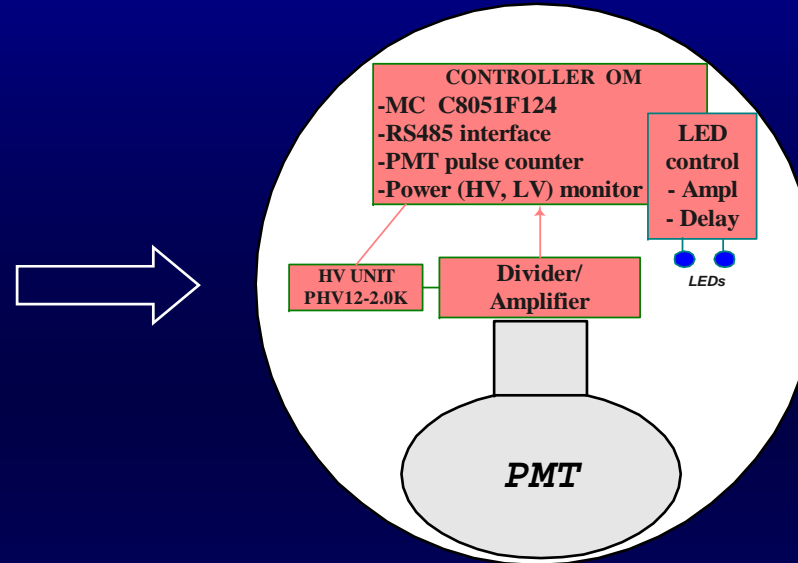
New optical module (OM)

OM NT200+



1. Hybrid photodetector
Quasar-370: ~80...100 ns FWHM
2. Outer control system
3. 25 kV Quasar power supply
4. 1-LED calibration

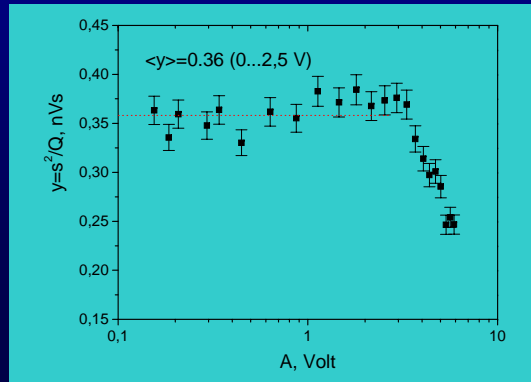
NEW OM designed
for km3 Baikal telescope



1. New PMT: R8055, XP1807, ... : ~20 ns FWHM
2. Control and monitoring system inside OM:
HV value, power supply, PMT noise rate,
temperature.
3. 2-LED calibration system for operation with
FADC: LED amplitude and pulse delay control:
dynamic range $>10^8$, delay range 0...1000 ns
(+/- 1 ns).

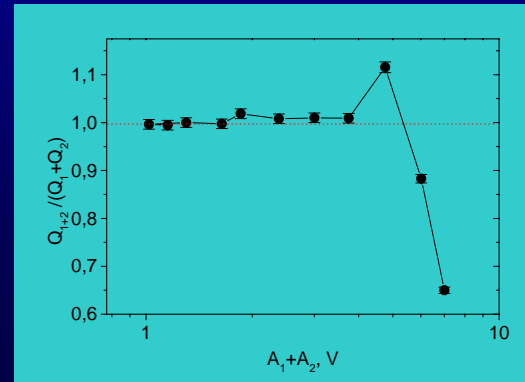
PM parameter monitoring with new OM control system

PM monitoring with one LED

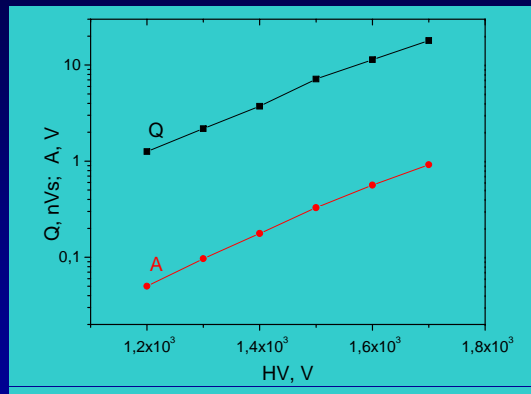
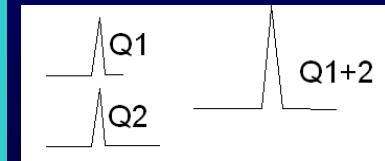


σ^2/Q dependence on pulse amplitude ($s^2/Q \sim \text{Gain_of_PM}$)

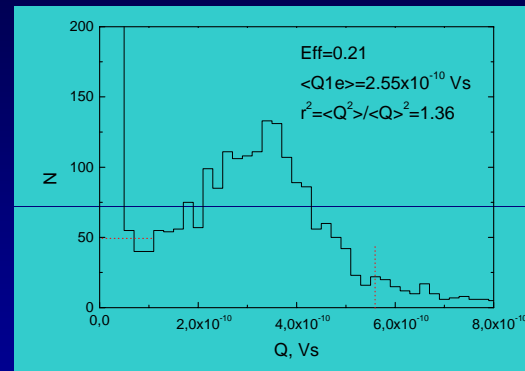
PM monitoring with two LED



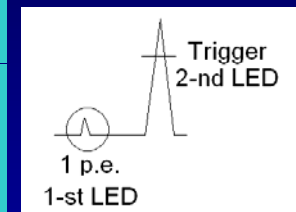
PM linearity test with 2 LEDs (1% precision)



Pulse area and amplitude dependence's on HV value



Single electron PM spectrum



PM selection for the km3 prototype string

Basic criteria of PM selection is its effective sensitivity to Cherenkov light: fraction of registered photons per photon flux unit.
(Photocathode area \times Quantum efficiency \times Collection efficiency)



Quasar-370

D \approx 14.6"

Quantum efficiency \approx 0.15

\approx ?

Hamamatsu R8055

D \approx 13"

Quantum efficiency \approx 0.20

\approx ?

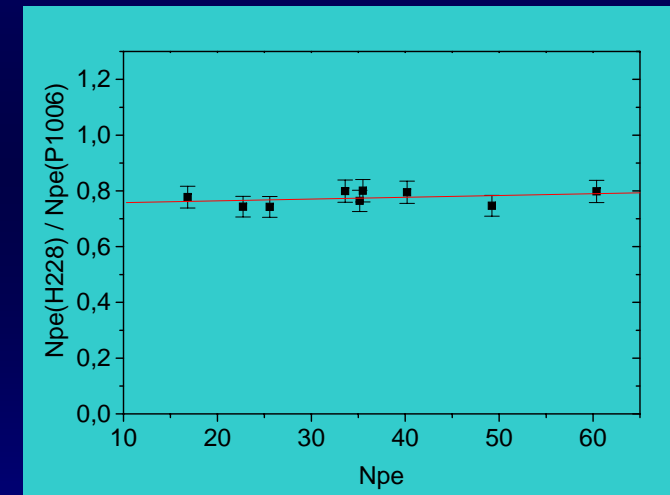
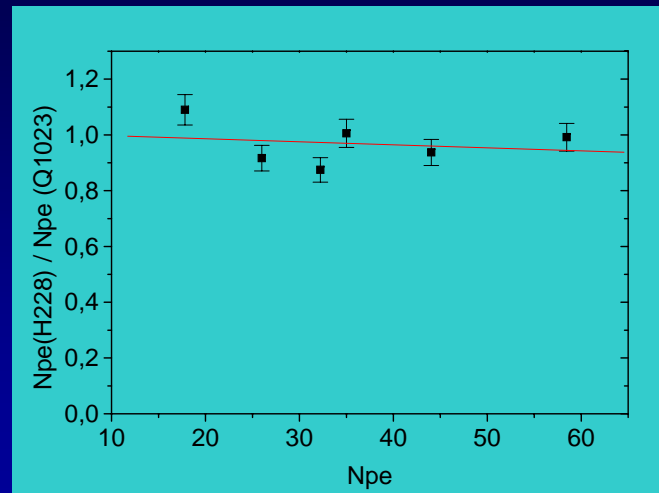
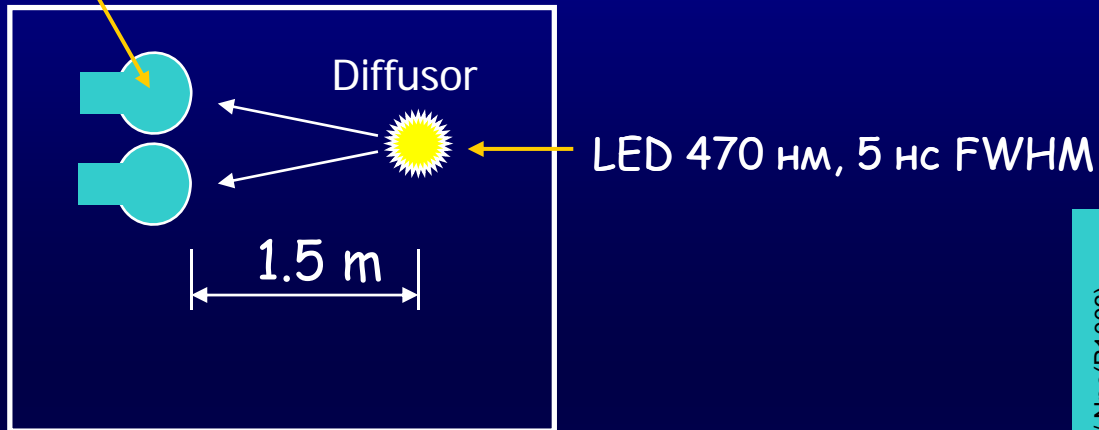
Photonis XP1807

D \approx 12"

Quantum efficiency \approx 0.24

PM selection : Laboratory tests

PMs R8055, XP1807, Quasar-370

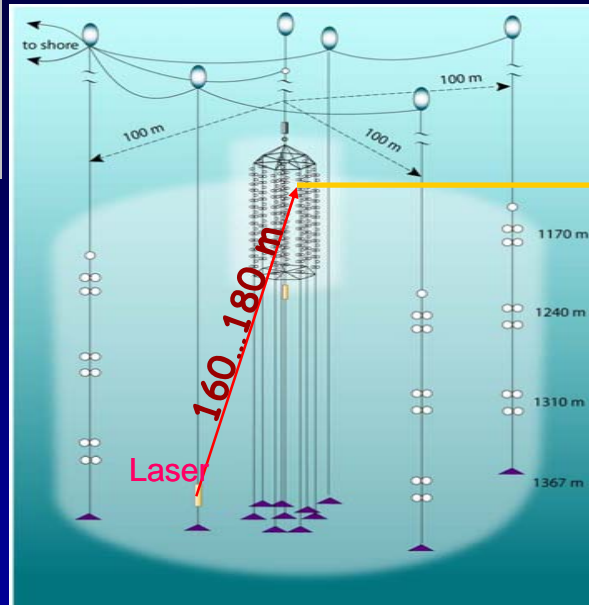


Examples of relative sensitivities measurements for Quasar-370, R8055, and XP1807

PM selection: Underwater tests

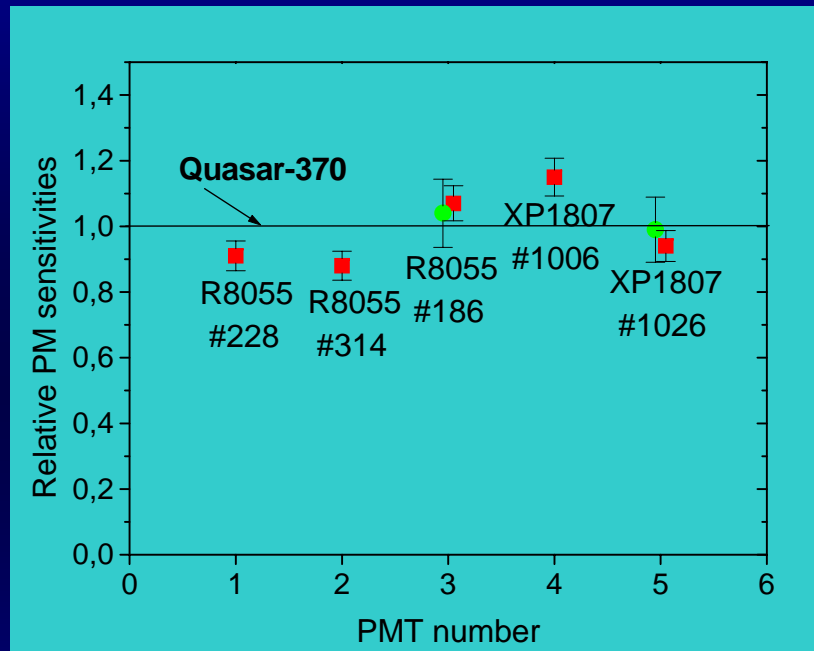
4 PM R8055 (Hamamatsu)
и 2 XP1807 (Photonis)
were installed to NT200+
detector (April 2007).

4 PM: central telescope
NT 200;
2 PM R8055: outer string,
FADC prototype.



- ● ← Quasar-370
- ● ← Quasar-370
- ● ← R8055
- ● ← XP1807
- ● ← Quasar-370
- ● ← Quasar-370

Ratio of effective sensitivities of large area PMs (preliminary results)



Smaller size (R8055, XP1807) tends to be compensated by larger photocathode sensitivities.

Ratio of effective sensitivity of large area PMs R8055/13" and XP1807/12" to Quasar-370/14.6". Laboratory (squares), in-situ (dots).

Summary

1. For the planned km³-detector in lake Baikal, R&D-activities have been started.
2. The existing NT200+ allows to verify all key elements and design principles of km³-detector.
3. Modernization of NT200+ DAQ system allowed to install a prototype FADC PM readout at spring 2007.
4. Six large area hemispherical PM (2 Photonis XP1807 and 4 Hamamatsu R8055) have been integrated into NT200+ to facilitate the optimal PM choice.
5. A full scale "new technology" prototype string will be installed in spring 2008 as a part NT200+.