# **Workshop on Advances,Innovations, and Prospects in High-Energy Nuclear Physics Deep Learning for nuclear EoS at extreme conditions**

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Oct. 20-24, WuHan



- The nuclear EoS describes the relationships between **pressure p, energy density e, temperature T, net baryon density ρ** and **chemical potential µ**. For instance, p(ρ, T),  $p(\mu, T)$  and  $e(\rho, T)$
- Crucial for understanding the **evolution of**<br> **early universe**, **supernova explosions**,<br> **neutron star stability**, heavy element<br>
synthesis, and heavy-jon collision **early universe**, **supernova explosions**, **neutron star stability**, **heavy element synthesis**, and **heavy-ion collision experiments**
- It also constrains two-body and three-body nuclear interactions as well as non perturbative Quantum Chromodynamics (QCD).

#### Nuclear EoS employed in astrophysical studies



F. Weber, IoP Publishing, Bristol(1999)

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T Kojo, PD Powell,YF Song,and G Baym, 2015 S. Rosswog et al., Astronomy and Astrophysics 341, 499 (1999).

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- $p(\mu, T)$  and  $e(\rho, T)$ <br>
 Crucial for understanding the **evolution of**<br> **early universe, supernova explosions, early universe**, **supernova explosions**, **neutron star stability**, **heavy element synthesis**, and **heavy-ion collision experiments**
- It also constrains two-body and three-body nuclear interactions as well as non perturbative Quantum Chromodynamics (QCD).

#### Nuclear EoS employed in HIC physics



LG Pang, H Petersen, XN Wang, PRC 2018

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# Nuclear EoS in relativistic hydrodynamics



#### **Name of CLVisc**:

1. CCNU-LBNL Viscous Hydro, CCNU = Central China Normal University 2. A 3+1D viscous hydro parallized on GPU using OpenCL

**Purpose:** Describe the non-equilibrium space-time evolution of hot QCD matter **Feature: 60 times faster** for hydrodynamic evolution, **100 times faster** for hadronization

> L.G. Pang, Q. Wang and X. N. Wang, PRC 86 (2012) 024911 L.G. Pang, B.W. Xiao, Y. Hatta, X.N.Wang, PRD 2015 L.G. Pang, H.Petersen, XN Wang, PRC97(2018)no.6,064918



# CLVisc for different EoS



**eta/s = 0 Lattice QCD EoS (smooth cross over)**

**eta/s = 0 First order phase transition**

**eta/s = 0.08 Lattice QCD EoS**

**eta/s = 0.08 First order phase transition eta/s: shear viscosity / entropy density**

Will the effect of EoS survive the dynamical evolution and exist in the final state hadrons?



# EoS for different phase transition types



baryon chemical potential  $\mu_B$ 



## Determine nuclear phase transitions



Nature Communications 2018, **LG. Pang**, K.Zhou, N.Su, H.Petersen, H. Stoecker, XN. Wang.



### Spinodal vs Maxwell 1<sup>st</sup> order phase transition



J. Steinheimer, L.G. Pang, K. Zhou, V. Koch and J. Randrup, JHEP 12 (2019) 122



## Looking for self similarity in momentum space



Self similarity, scaling invariance



Dynamical Edge Convolution Network

PLB 827(2022) 137001, Y.-G. Huang, L.-G. Pang, X.F. Luo and X.-N. Wang







Protons, Predicted labels

PLB 822 (2021) 136669, Y.J Wang, F.P. Li, Q.F. Li, H.L. L¨u, and K. Zhou



## Auto Encoder for order parameter

#### PHYSICAL REVIEW RESEARCH 2, 043202 (2020)

#### Nuclear liquid-gas phase transition with machine learning

Rui Wang  $\bullet$ , <sup>1,2,\*</sup> Yu-Gang Ma, <sup>1,2,†</sup> R. Wada,<sup>3</sup> Lie-Wen Chen  $\bullet$ ,<sup>4</sup> Wan-Bing He,<sup>1</sup> Huan-Ling Liu,<sup>2</sup> and Kai-Jia Sun<sup>3,5</sup>





## Jet eloss and medium response

Can Being Underwater Protect You From Bullets?



**66** If the bullet is shot from an angle of 30 Degrees, then being underwater in the range of  $3$ -5 feet (0.9-1.5 meters) can ensure safety from most guns.



Jet quenching in hot QGP





Nuclear EoS: 
$$
c_s^2 = \frac{dP}{d\epsilon} = \sin^2 \theta
$$

 $\frac{2}{s} = \frac{ar}{d\epsilon} = \sin^2 \theta$  Shear Viscosity: width of the shock wave



- Random production locations and propagating directions relative to collective flow
- ●Tilted by different path length and collective flow



L.M. Satarov, H. Stoecker, I.N. Mishustin, PLB 627 (2005) 64-70



# DL assisted jet tomography (gamma-jet)



Z Yang, YY He, W Chen, WY Ke, LG Pang, XN Wang, EPJC 83 (2023) 7, 652



## Training data: CoLBT(LBT + CLVisc)

$$
p\partial f(p) = -C(p) \quad (p \cdot u > p_{cut}^0)
$$

$$
\partial_{\mu} T^{\mu\nu}(x) = j^{\nu}(x)
$$

$$
j^{\nu} = \sum_{i} p_i^{\nu} \delta^{(4)}(x - x_i) \theta(p_{cut}^0 - p \cdot u)
$$

**LBT: YY He, T Luo, XN Wang, Y Zhu,** PRC 91 (2015) 054908, PRC 97 (2018) 1, 019902 -10

#### **CLVisc:**

**LG Pang, Q Wang, XN Wang, PRC 86 (2012) 024911**

**LG Pang, H Petersen, XN Wang,** PRC 97 (2018) 6, 064918

**XY Wu, GY Qin, LG Pang, XN Wang,**PRC 105 (2022) 3, 034909



**CoLBT**:

**W Chen, T Luo, SS Cao, LG Pang, XN Wang,** PLB 777 (2018) 86-90



# DL assisted jet tomography



Z Yang, YY He, W Chen, WY Ke, **LG Pang**, XN Wang, EPJC 83 (2023) 7, 652



## Enhance the Diffusion Wake signal



Z Yang, YY He, W Chen, WY Ke, LG Pang, XN Wang, EPJC 83 (2023) 7, 652 Z Yang, T Luo, W Chen, LG Pang, XN Wang, PRL 130 (2023) 5,052301



# Effective theory: DL For Quasi Particle Mass



### FuPeng Li, HL Lu, LG Pang, GY Qin, PLB 2023

$$
\ln Z(T) = \ln Z_g(T) + \ln Z_{u,d}(T) + \ln Z_s(T),
$$

Fermi-Dirac distributions,

$$
\ln Z_g(T) = -\frac{16V}{2\pi^2} \int_0^\infty p^2 dp
$$
  
\n
$$
\ln \left[ 1 - \exp\left( -\frac{1}{T} \sqrt{p^2 + m_g^2(T)} \right) \right], \quad (2)
$$
  
\n
$$
\ln Z_{q_i}(T) = +\frac{12V}{2\pi^2} \int_0^\infty p^2 dp
$$
  
\n
$$
\ln \left[ 1 + \exp\left( -\frac{1}{T} \sqrt{p^2 + m_{q_i}^2(T)} \right) \right], \quad (3)
$$



quarks,  $m_s(T, \theta_2)$  for strange quark and  $m_q(T, \theta_3)$  for gluons, where  $\theta_1$ ,  $\theta_2$  and  $\theta_3$  are the parameters in DNN shown in Fig. 1.

The resulting pressure and energy density are computed using the following statistical formulae,

$$
P(T) = T\left(\frac{\partial \ln Z(T)}{\partial V}\right)_T, \tag{5}
$$

$$
\epsilon(T) = \frac{T^2}{V} \left( \frac{\partial \ln Z(T)}{\partial T} \right)_V, \tag{6}
$$





FuPeng Li, HL Lu, LG Pang, GY Qin, PLB 2023



## Location of minimum eta/s



Thesis of Valeriya Mykhaylova, 2023



# Extend Quasi Parton Model to finite muB



#### **Model:**

Deep learning Quasi Parton Model Effective theory of strongly coupled QGP and nuclear matter at finite baryon density

**Training data: Lattice QCD + HRG** PRD 95, 054504 (2017) PRL118, 182301 (2017) PRD 90, 094503 (2014)

 $\frac{4}{2}$ 

 $0.08$ 

0.06

 $0.02$ 

 $0.00 -$ 

0.050

0.075

 $\approx 0.04$ 

 $0.1$ 

 $\frac{0.2}{T[\text{GeV}]}$ 

 $0.3$ 

0.100 0.125 0.150 0.175

T[GeV]

 $04$ 

 $-\bullet$  LQCD, and, HRG,  $\mu_B = 0$ 

 $\frac{0.2}{T[\text{GeV}]}$ 

0.050 0.075 0.100 0.125 0.150 0.175

T[GeV]

 $0.3$ 

 $0.4$ 

 $\rightarrow$  network,  $\mu_B = 0$ 

 $01$ 

15.0

 $12<sub>5</sub>$ 

 $10.0$  $\sqrt{7^3}$ 

 $0.15$ 

0.05

 $0.00$  –  $...$ 

 $\mathbb{R}^{\infty}$  0.10









●We explored 3 approaches to studying QCD EoS using deep learning

- ●For soft probes, DL serves as an EoS-meter
- ●For hard probes, DL assisted jet tomography aids in the investigation of QCD EoS through Mach cones
- ●DL and auto-diff are widely used to represent unknown functions to construct effective theories
- ●DL quasi parton model are extended to finite muB region