

Exploring neutrino masses and mixing in R-Parity Violating supersymmetric models

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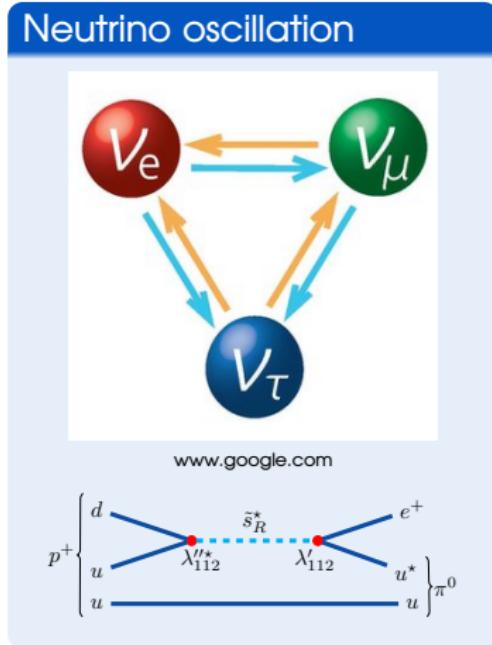


Outline

- Introduction of Model
- Bilinear RPV SUSY model
 - Neutrino mass generation
 - Observables and constraints
 - Parameters
 - Analysis details
 - Results and discussion
- Trilinear RPV model
 - Parameters
 - Results

Existence of neutrino mass

- Neutrino oscillation → one of the most robust indications towards the existence of physics BSM
- Within Standard Model (SM) framework neutrino is massless → no right handed neutrino
- No neutrino mass from RPC MSSM → leads to RPV MSSM
- R-parity, $R_p = (-1)^{(3B-2L+S)}$



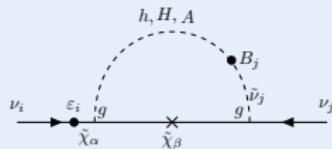
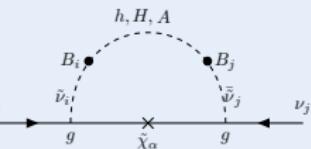
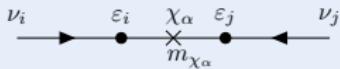
- $W_{\cancel{R_p}} = \varepsilon_i L_i H_u + \frac{1}{2} \lambda_{ijk} L_i L_j \bar{e}_k + \lambda'_{ijk} L_i Q_j \bar{d}_k + \frac{1}{2} \lambda''_{ijk} \bar{u}_i \bar{d}_j \bar{d}_k$
- Two separate analyses - lepton number violating Bilinear RPV model and Trilinear RPV model

Bilinear Model definition

Bilinear R-Parity violating Superpotential

$$W_{Bp} = \varepsilon_i L_i H_u; \quad \mathcal{L}_{Bp} = [\varepsilon_i (\tilde{H}_u^0 \nu_{iL} - \tilde{H}_u^+ l_{iL})]; \quad \mathcal{L}_{soft} = B_i \tilde{L}_i H_u$$

Tree and loop level diagrams



Tree level

BB loop

ϵB loop

- Only one neutrino becomes massive at tree level → The highest neutrino mass eigenstate
- BB loop is the dominant one
- Second mass becomes heavy mainly from BB loop
- Lowest mass eigenstate will become heavy from ϵB loop

Observables

Neutrino observables

- Two mass square splitting values (Δm_{21}^2 and Δm_{31}^2)
- Three mixing angles ($\theta_{12}, \theta_{13}, \theta_{23}$)

Source: JHEP 02 (2021) 071.

Constraints from Higgs

- Higgs Mass: We have considered ± 3 GeV as theoretical uncertainty around 125 GeV. [Phys. Rev. Lett. 114 191803 \(2015\)](#)
- Higgs coupling strength data from LHC at $\sqrt{s} = 13$ TeV \rightarrow Higgs coupling to Z, W, b, t, μ , τ , and γ particle. [CMS-PAS-HIG-19-005, 2020](#)

Constraints from flavor physics

- rare b -hadron decays as $\mathcal{B}(B \rightarrow X_s + \gamma)$ and $\mathcal{B}(B_s^0 \rightarrow \mu^+ + \mu^-)$ [Eur. Phys. J. C 81 226 \(2021\)](#) and [Phys. Rev. Lett. 128 041801 \(2022\)](#)
- Total 15 observables

Parameters

- Considered minimal set of parameters

List of fixed parameters

$$M_1 = 300 \text{ GeV}$$

$$M_{\tilde{q}} = 3 \text{ TeV}$$

$$M_2 = 1.2 \text{ TeV}$$

$$M_{\tilde{l}} = 2 \text{ TeV}$$

$$M_3 = 3 \text{ TeV}$$

$$A_t = -3.5 \text{ TeV}$$

$$M_A = 3 \text{ TeV}$$

Range of input parameters for scanning

From literature study we came up with some exhaustive range of each parameter such as

$$\mu: 1 \text{ to } 3 \text{ TeV}$$

$$\tan \beta: 1 \text{ to } 60$$

$$\varepsilon_i (i = 1, 2, 3): -1.0 \text{ to } 1.0 \text{ GeV}$$

$$B_i (i = 1, 2, 3): 0.1 \text{ GeV to } 10 \text{ TeV}$$

$$v_i (i = 1, 2, 3): 10^{-8} \text{ to } 0.1 \text{ GeV}$$

- So we have total 11 free parameters

Analysis details

- For scanning we use Markov Chain Monte Carlo (MCMC) based likelihood analysis → emcee ([Publications of the Astronomical Society of the Pacific, 125 306 \(2013\)](#))

- We find the maximum likelihood function $L \propto \exp(-\mathcal{L})$

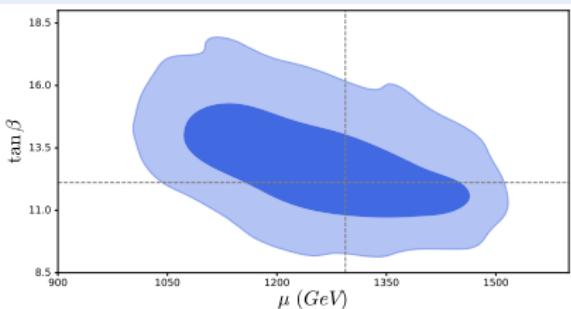
- Log likelihood $\mathcal{L} = \frac{\chi^2}{2} = \frac{1}{2} \sum_{i=1}^{n_{\text{obs}}} \left[\frac{\Gamma_i^{\text{obs}} - \Gamma_i^{\text{th}}}{\sigma_i} \right]^2$

- Maximum likelihood means we find the minimum χ^2
- Degrees of freedom(D.O.F) = 15 independent observables - 11 free parameters = 4
- We use a flat prior on all the parameters
- We use 500 walkers and 400 steps for each walker. Total sample generated = $500 \times 400 \times n_{\text{core}} = 200000 \times n_{\text{core}}$

Results - Normal Hierarchy ($\nu_3 > \nu_2 > \nu_1$)

- We have got $\chi^2_{min} = 3.46$
- $\sum m_{\nu_i} = 0.059 \text{ eV} \rightarrow \text{satisfies} \rightarrow \sum m_{\nu_i} < 0.12 \text{ eV}$

$\mu - \tan \beta$ contour

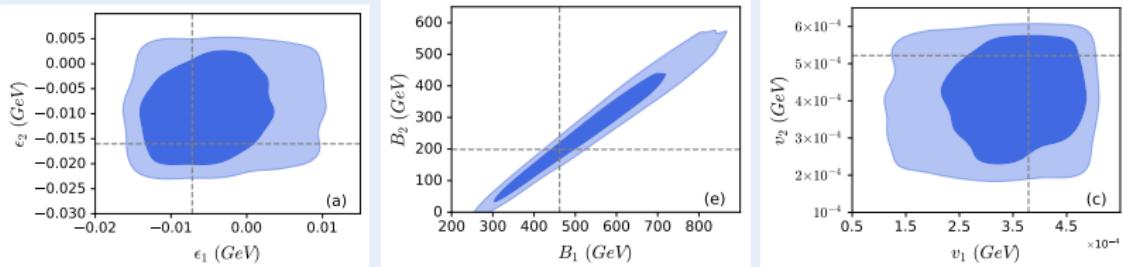


- Tree level \rightarrow only third neutrino
- BB loop \rightarrow second neutrino
- ϵB loop \rightarrow first neutrino
- From theory $\tan \beta$ should not be large or very low
- It also depends on the choice of M_A and A_t parameters
- Most stringent limit comes from neutrino oscillation data

- $[m_\nu]_{ij}^{(\varepsilon\varepsilon)} \propto \frac{1}{\mu \tan^2 \beta}$
- $[m_\nu]_{ij}^{(BB)} \propto \tan^2 \beta$
- $[m_\nu]_{ij}^{(\varepsilon B)} \propto \tan \beta$

Results - Normal Hierarchy ($\nu_3 > \nu_2 > \nu_1$)

Contour plots



- $m_{\text{highest}} \propto (\epsilon_1^2 + \epsilon_2^2 + \epsilon_3^2) \sin^2 \xi$, ξ represents alignment between ϵ_i and ν_i ([JHEP 02 \(2024\) 004](#))
- Heaviest one is τ flavored $\rightarrow \epsilon_3$ and ν_3 should be largest one
- Also it has next to zero admixture of electron neutrino $\rightarrow \epsilon_1$ and ν_1 should be lowest one
- $m_2 \propto B_i B_j$ and $m_1 \propto \epsilon_i B_j + \epsilon_j B_i$
- Second one has comparable admixture of all three neutrino flavors \rightarrow nice correlations among B_i parameters

Results - Normal Hierarchy ($\nu_3 > \nu_2 > \nu_1$)

- Loop contributions are already suppressed and $\tan \beta$ is already restricted by tree level mass
- For these contributions to neutrino masses to be significant, the B_i parameters have to be much larger compared to ϵ_i parameters
- ϵB loop contribution is further suppressed due to their dependence on ϵ_i
- As a result, B_1 is expected to be relatively larger than B_2 since the lightest state is dominantly electron neutrino-like
- B_3 will have larger value compared to others

Best-fit point

$\epsilon_1 = -0.0072$	$v_1 = 0.00038$	$B_1 = 461$	$\mu = 1293$
$\epsilon_2 = -0.0160$	$v_2 = 0.00052$	$B_2 = 198$	$\tan \beta = 12$
$\epsilon_3 = -0.0279$	$v_3 = 0.00091$	$B_3 = 1760$	

All are in GeV unit except $\tan \beta$ (JHEP 02 (2024) 004)

Results - Inverted Hierarchy ($\nu_2 > \nu_1 > \nu_3$)

- $\chi^2_{min} = 3.38$ and $\sum m_{\nu_i} = 0.1$ eV \rightarrow satisfies $\sum m_{\nu_i} < 0.15$ eV
- Second one is heaviest \rightarrow an almost equal admixture of all three neutrino flavors
- ϵ_2 and ν_2 have largest values
- ν_1 is the second heaviest one and have larger values than NH scenario \rightarrow larger B_1 and B_2 values required
- As neutrino oscillation parameters are more constraint in IH scenario \rightarrow allowed parameter space is also more constraint

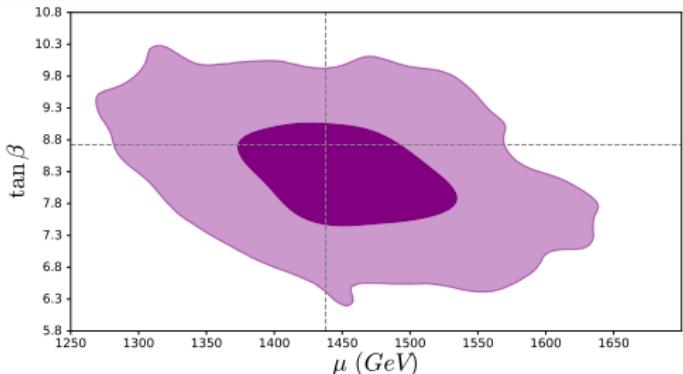
Best-fit point

$\epsilon_1 = -0.0216$	$\nu_1 = 0.00086$	$B_1 = 894$	$\mu = 1437$
$\epsilon_2 = -0.0833$	$\nu_2 = 0.00140$	$B_2 = 982$	$\tan \beta = 8$
$\epsilon_3 = -0.0499$	$\nu_3 = 0.00110$	$B_3 = 1609$	

All are in GeV unit except $\tan \beta$ (JHEP 02 (2024) 004)

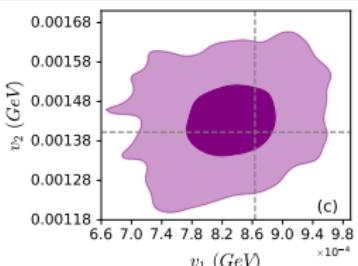
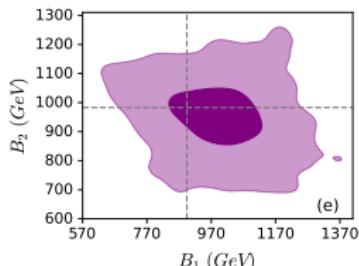
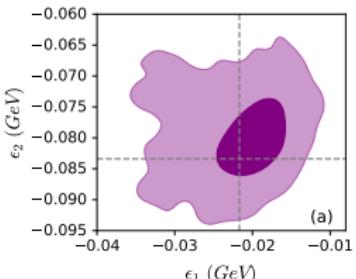
Results - Inverted Hierarchy ($\nu_2 > \nu_1 > \nu_3$)

$\mu - \tan \beta$ contour



- $[m_\nu]_{ij}^{(\varepsilon\varepsilon)} \propto \frac{1}{\mu \tan^2 \beta}$
- ν_3 has lowest mass
→ μ must have larger value than NH scenario

Contour plots



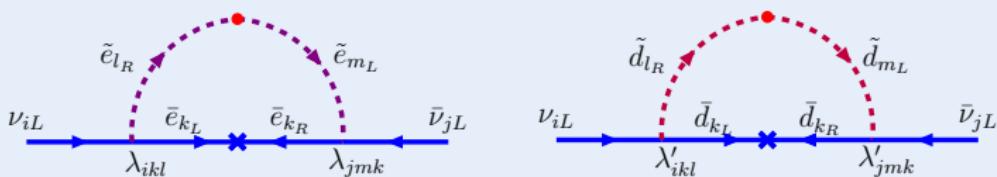
Trilinear Model

Superpotential

$$W_L = \frac{1}{2} \lambda_{ijk} L_i L_j E_k^c + \lambda'_{ijk} L_i Q_j D_k^c$$

λ_{ijk} is antisymmetric $\rightarrow 9 \lambda_{ijk} + 27 \lambda'_{ijk}$ parameters

Loop diagrams



- Due to fermion mass hierarchy, we consider only third generation couplings λ_{i33} and λ'_{i33}
- $M_\nu = \frac{1}{8\pi^2 \tilde{m}} [\lambda_{i33} \lambda_{j33} m_\tau^2 + 3\lambda'_{i33} \lambda'_{j33} m_b^2]$
- Leading contribution to heaviest neutrino
 $m_{\nu_3} = \frac{3m_b^2}{8\pi^2 \tilde{m}} \sum_i \lambda'^2_{i33}$

Parameters space

- we have 2 λ_{i33} ($i = 1, 2$) and 3 λ'_{i33} ($i = 1, 2, 3$) parameters
- We also consider μ and $\tan \beta$ as before
- Total 7 parameters
- We have added one other observable $B \rightarrow \tau\nu$
- we have 16 observables \rightarrow d.o.f = 9

Range of parameters

$$\begin{aligned} |\lambda_{i33}| (i = 1, 2) &: 0 - 0.001 \text{ GeV} \\ |\lambda'_{i33}| (i = 1, 2, 3) &: 0 - 0.001 \text{ GeV} \\ \mu &= 1000 - 3000 \text{ GeV} \\ \tan \beta &= 1 - 60 \end{aligned}$$

- only LLE coupling $\rightarrow \Delta m^2_{31}$ and θ_{12}
- only LQD coupling \rightarrow can satisfy all except Δm^2_{21}

Results - Normal hierarchy ($\nu_3 > \nu_2 > \nu_1$)

- The minimum χ^2 we obtained 4.14 for d.o.f 9
- It also satisfies the cosmological bound
- Here $m_{\nu_3} = \frac{3m_b^2}{8\pi^2\tilde{m}} \sum_i \lambda'_{i33}$
- λ'_{333} must have higher value than others
- Second and first neutrinos get masses mostly from λ_{i33} couplings $\rightarrow \lambda_{233}$ coupling must have larger value than λ_{133}

Best-fit point

$$\lambda_{133} = 1.71 \times 10^{-4}$$

$$\lambda_{233} = 2.52 \times 10^{-4}$$

$$\mu = 1996$$

$$\tan \beta = 6.68$$

$$\lambda'_{133} = -7.61 \times 10^{-5}$$

$$\lambda'_{233} = -7.65 \times 10^{-5}$$

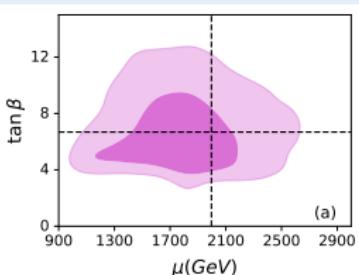
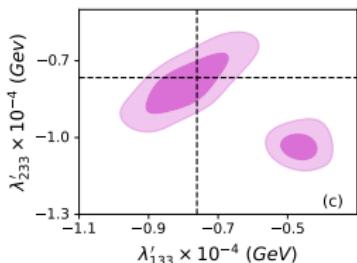
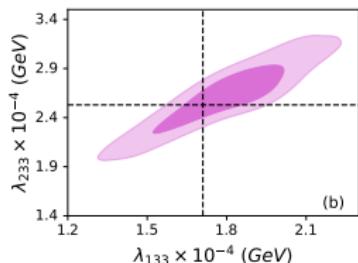
$$\lambda'_{333} = -1.34 \times 10^{-4}$$

All parameters are in GeV unit except $\tan \beta$

Results - Normal hierarchy

- With the λ_{i33} ($i = 1, 2$), the bino-type LSP ($\tilde{\chi}_1^0$) decay final states $\rightarrow \tau^\pm e^\mp \nu, \tau^\pm \tau^\mp \nu$, and $\tau^\pm \mu^\mp \nu$
- λ'_{i33} ($i = 1, 2, 3$) $\rightarrow l + b + j$ and $2b + \cancel{E}_T$
- At the best-fit point branching fraction corresponding to λ_{i33} and λ'_{i33} coupling $\sim 83\%$ and 17% respectively
- As the coupling values of λ_{i33} are larger than the values of λ'_{i33} , the branching ratio corresponding to the λ_{i33} coupling is also comparatively higher

contour plot



Results - Inverted hierarchy

- Minimum χ^2 obtained 4.56 for d.o.f 9
- Second one is the heaviest one $\rightarrow \lambda_{233}$ has the largest value than others
- Third neutrino gets mass from $\lambda'_{i33} \rightarrow \lambda'_{333}$ must have higher value and it is lower than NH scenario
- Lowest neutrino eigenstate has mass very close to second one and to get that higher mass we need contribution from both couplings $\rightarrow \lambda_{133}$ must have larger value as well as λ'_{133}
- the total branching ratio corresponding to λ_{i33} and λ'_{i33} couplings are 93% and 7% respectively due the larger values of *LLE* type RPV couplings than *LQD* type couplings.

Results - Inverted hierarchy

Best-fit point

$$\lambda_{133} = 2.21 \times 10^{-4}$$

$$\lambda_{233} = 4.88 \times 10^{-4}$$

$$\mu = 1676$$

$$\tan \beta = 8.38$$

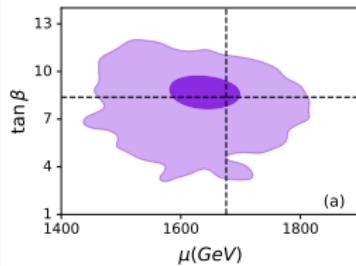
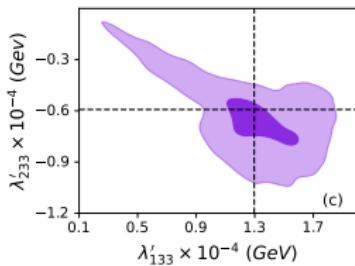
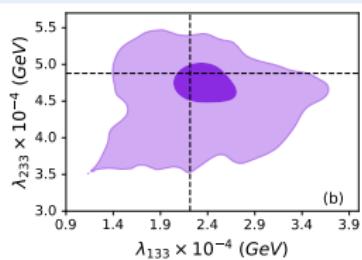
$$\lambda'_{133} = 1.29 \times 10^{-4}$$

$$\lambda'_{233} = -5.92 \times 10^{-5}$$

$$\lambda'_{333} = -1.25 \times 10^{-4}$$

All parameters are in GeV unit except $\tan \beta$

contour plot



The allowed parameter space for IH scenario is more constraint than NH scenario as BRPV model

Conclusion

- We have considered neutrino observables along with recent higgs data and flavor physics data
- We have done two separate analyses for Bilinear RPV and Trilinear RPV model
- To scan the parameter space we have used MCMC based likelihood analysis
- We obtained that the both the models can explain neutrino and other experimental data.
- We have also shown the allowed 1σ and 2σ region for each parameter space along with their correlation
- But the allowed parameter space is tightly constrained

THANK YOU

Collider constraints

Gluino search

- Limit is 2.0-2.5 TeV for various couplings $\rightarrow m_{\tilde{g}} = 3 \text{ TeV}$

Squark search

Limit is 0.8-1.9 TeV for different couplings $\rightarrow m_{\tilde{q}} = 3 \text{ TeV}$ (fixed)

Slepton search

- Limit is 0.86-1.2 TeV depending on couplings \rightarrow all the slepton masses fixed at 3 TeV

Chargino search

- We have considered a scenario with bino-type LSP and wino-type NLSP
- $m_{\tilde{\chi}_2^0}/m_{\tilde{\chi}_1^\pm}$ excluded upto 1.14 TeV for λ_{i33} coupling
- We consider $m_{\tilde{\chi}_2^0}/m_{\tilde{\chi}_1^\pm}$ masses fixed at 1.2 TeV and $m_{\tilde{\chi}_1^0} = 300 \text{ GeV}$

Results with only *LLE* coupling

- Mass matrix for this model is $M_\nu|_{\lambda} = \frac{1}{8\pi^2\tilde{m}} \lambda_{i33}\lambda_{j33} m_\tau^2$
- After diagonalization only third neutrino becomes heavy,
 $m_{\nu_3} = \frac{m_\tau^2}{8\pi^2\tilde{m}} \sum_{i=1,2} \lambda_{i33}^2$
- $\sin \theta_{12} = \frac{\lambda_{133}}{\sqrt{\lambda_{133}^2 + \lambda_{233}^2}}$
- This model can satisfy only Δm_{31}^2 and $\sin \theta_{12}$

Parameter	Value	Observable	Value	χ^2 contribution
λ_{133}	2.76×10^{-4}	Δm_{21}^2	4.15×10^{-13}	1162
λ_{233}	4.06×10^{-4}	Δm_{31}^2	2.56×10^{-3}	0.11
μ	2151	θ_{13}	7.67×10^{-25}	4305
$\tan \beta$	8.02	θ_{12}	34.27	0.001
		θ_{23}	90.0	2659

Results with only LQD couplings

- Mass matrix for this model $M_\nu|_{\lambda'} = \frac{3}{8\pi^2 \tilde{m}} \lambda'_{i33} \lambda'_{j33} m_b^2$
- After diagonalization only third neutrino becomes heavy which is already mentioned before
- $\sin \theta_{13} = \frac{\lambda'_{133}}{\sum_{i=1,2,3} \lambda'^2_{i33}}$ and $\sin \theta_{23} = \frac{\lambda'_{233}}{\sum_{i=2,3} \lambda'^2_{i33}}$
- It can satisfy all the observables except Δm_{21}^2 which is reflected in the result

Parameter	Value	Observable	Value	χ^2 contribution
λ'_{133}	-6.66×10^{-5}	Δm_{21}^2	6.55×10^{-11}	1162
λ'_{233}	-1.25×10^{-4}	Δm_{31}^2	2.61×10^{-3}	4.0
λ'_{333}	-1.21×10^{-4}	θ_{13}	8.74	2.60
μ	1867	θ_{12}	35.66	1.85
$\tan \beta$	7.73	θ_{23}	48.86	0.25

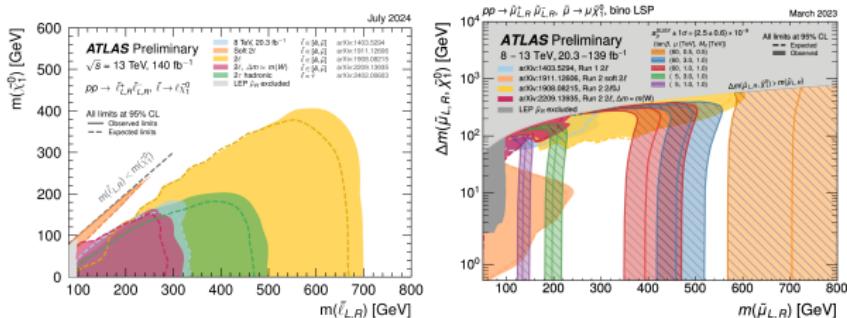
Trilinear mass matrix

$$\begin{aligned}
 M_{ij}^\nu|_\lambda &= \frac{1}{16\pi^2} \sum_{k,l,m} \lambda_{ikl} \lambda_{jmk} m_{e_k} \frac{(\tilde{m}_{LR}^{e^2})_{ml}}{m_{\tilde{e}_{Rl}}^2 - m_{\tilde{e}_{Lm}}^2} \ln\left(\frac{m_{\tilde{e}_{Rl}}^2}{m^2 \tilde{e}_{Lm}}\right) + (i \leftrightarrow j) \\
 M_{ij}^\nu|_{\lambda'} &= \frac{3}{16\pi^2} \sum_{k,l,m} \lambda'_{ikl} \lambda'_{jmk} m_{d_k} \frac{(\tilde{m}_{LR}^{d^2})_{ml}}{m_{\tilde{d}_{Rl}}^2 - m_{\tilde{d}_{Lm}}^2} \ln\left(\frac{m_{\tilde{d}_{Rl}}^2}{m^2 \tilde{d}_{Lm}}\right) + (i \leftrightarrow j)
 \end{aligned} \tag{1}$$

$$\begin{aligned}
 M_{ij}^\nu|_\lambda &\simeq \frac{1}{8\pi^2} \frac{A^e - \mu \tan \beta}{\bar{m}_{\tilde{e}}^2} \sum_{k,l} \lambda_{ikl} \lambda_{jkl} m_{e_k} m_{e_l} \\
 M_{ij}^\nu|_{\lambda'} &\simeq \frac{3}{8\pi^2} \frac{A^d - \mu \tan \beta}{\bar{m}_{\tilde{d}}^2} \sum_{k,l} \lambda'_{ikl} \lambda'_{jkl} m_{d_k} m_{d_l}
 \end{aligned} \tag{2}$$

Result - Anomalous muon magnetic moment

- $\Delta a_\mu = a_\mu^{\text{Exp}} - a_\mu^{\text{SM}} = (25.1 \pm 5.9) \times 10^{-10}$
- Sneutrino-chargino and slepton-neutralino loops contribution
- Lower the smuon masses keeping all other slepton masses decoupled



Input parameters		Output observables			
Parameters	BP-I	BP-II	Output	BP-I	BP-II
M_1 (GeV)	128	183	$m_{\tilde{\chi}_1^0}$ (GeV)	125	180
M_2 (GeV)	1200	1200	$m_{\tilde{\chi}_1^\pm}$ (GeV)	1198	1192
$m_{\tilde{\mu}_L}$ (GeV)	120	200	$m_{\tilde{\mu}_1}$ (GeV)	164	224
$m_{\tilde{\mu}_R}$ (GeV)	190	240	$m_{\tilde{\mu}_2}$ (GeV)	175	235
$\tan \beta$	13.75	11.94	$\Delta a_\mu (\times 10^{-10})$	25.41	13.52

Table: (JHEP 02 (2024) 004)