# **Higgs Combination**

## The 3 analyses

**Z(II)H(XX):** neural to categorize in H flavour decay modes; fit on recoil distribution **Z(vv)H(XX):** neural to categorize in H flavour decay modes; 2D fit on visible and missing mass **Z(qq)H(qq):** multi-jet environment – categorization in flavours, 2D fit on recoil and dijet system

Final state	Z(II)H(jj) [%]	Z(vv)H(jj) [%]	Z(jj)H(jj) [%]	Comb. [%]	Previous results from Annecy, using 7.2 /ab Combination = quadrature sum	
$H \rightarrow bb$	0.68	0.3	0.25	0.18		
$H \rightarrow cc$	4.11	2.17	2.92	1.60		
$H \rightarrow gg$	2.28	0.92	2.00	0.78		
$H \rightarrow ss$	342	114	363	103		

## What we need for the mid-term

We fit directly  $\sigma(ZH)^*BR(H\rightarrow jj)$  (assume Z branching ratios perfectly measured)

The minimal deliverables for the mid-term report are:

- Combination of the 3 analyses at 240
- Combination of the 3 analyses at 365

Knowing ZH couplings, we can directly fit the BR and/or kappa's

- Combination of 240+365 GeV
- In first stage, combine assuming "absolute" ZH coupling known from Z(II)H
- Afterwards, we can attempt for a simultaneously with the Z(II)H analysis to extract the kappas:
  - At 240 GeV: *σ*(ZH) ~ 0.7%, at 365 GeV: *σ*(ZH) ~ 1.1%
  - Need to take into account possible correlations (some discussion needed)

#### It seems important to monitor the improvements due to better algorithms

- CDR (and MTR) precisions based on 2018 ILC efficiencies see table below
  - Let's see where we can do better and update the table accordingly

**Table 1.** From Ref. [4]: Relative uncertainty (in %) on  $\sigma_{\text{ZH}} \times \mathcal{B}(\text{H} \to X\overline{X})$  and  $\sigma_{\nu_e \bar{\nu}_e \text{H}} \times \mathcal{B}(\text{H} \to X\overline{X})$ , as expected from the FCC-ee data at 240 and 365 GeV.

$\sqrt{s}$	240	GeV	$365{ m GeV}$	
Integrated luminosity	5 a	$b^{-1}$	1.5	$\mathrm{ab}^{-1}$
Channel	ZH	$\nu_{\mathrm{e}} \bar{\nu}_{\mathrm{e}}$ H	ZH	$\nu_{\rm e}\bar{\nu}_{\rm e}~{\rm H}$
${\rm H} \rightarrow {\rm any}$	$\pm 0.5$		$\pm 0.9$	
$H \rightarrow b\bar{b}$	$\pm 0.3$	$\pm 3.1$	$\pm 0.5$	$\pm 0.9$
$H \to c \bar{c}$	$\pm 2.2$		$\pm 6.5$	$\pm 10$
$\mathrm{H} \to \mathrm{gg}$	$\pm 1.9$		$\pm 3.5$	$\pm 4.5$
$\rm H \rightarrow W^+W^-$	$\pm 1.2$		$\pm 2.6$	$\pm 3.0$
$\mathrm{H} \to \mathrm{ZZ}$	$\pm 4.4$		$\pm 12$	$\pm 10$
$H \rightarrow \tau^+ \tau^-$	$\pm 0.9$		$\pm 1.8$	$\pm 8$
${ m H}  ightarrow \gamma \gamma$	$\pm 9.0$		$\pm 18$	$\pm 22$
$H \rightarrow \mu^+ \mu^-$	$\pm 19$		$\pm 40$	
$\mathrm{H} \rightarrow \mathrm{invisible}$	< 0.3		< 0.6	

More lumi

**Table 3.** From Ref. [4]: Relative uncertainty (in %) on  $\sigma_{\text{ZH}} \times \mathcal{B}(\text{H} \to X\overline{X})$  and  $\sigma_{\nu_e \bar{\nu}_e \text{H}} \times \mathcal{B}(\text{H} \to X\overline{X})$ , as expected from the FCC-ee data at 240 and 365 GeV.

$\sqrt{s}$	$240{ m GeV}$		$365{ m GeV}$	
Integrated luminosity	$10.8  {\rm ab}^{-1}$		$3.0\mathrm{ab}^{-1}$	
Channel	ZH	$ u_{\mathrm{e}} \bar{\nu}_{\mathrm{e}} \mathrm{H}$	ZH	$ u_{ m e} ar{ u}_{ m e}  { m H}$
${\rm H} \rightarrow {\rm any}$	$\pm 0.36$		$\pm 0.6$	
${ m H}  ightarrow { m b} {ar b}$	$\pm 0.20$	$\pm 2.1$	$\pm 0.35$	$\pm 0.6$
$H \to c \bar c$	$\pm 1.5$	?	$\pm 4.4$	$\pm 7.1$
$\mathrm{H} \to \mathrm{gg}$	$\pm 1.3$	?	$\pm 2.5$	$\pm 3.2$
$\rm H \rightarrow W^+W^-$	$\pm 0.8$	?	$\pm 1.8$	$\pm 2.1$
$\mathrm{H} \to \mathrm{ZZ}$	$\pm 3.0$	?	$\pm 8.5$	$\pm 7.1$
$H \rightarrow \tau^+ \tau^-$	$\pm 0.6$	?	$\pm 1.3$	$\pm 5.7$
$H \rightarrow \gamma \gamma$	$\pm 6.1$	?	$\pm 13$	$\pm 16$
$\mathbf{H} \to \mathbf{Z} \gamma$	??	??	??	??
$H \rightarrow \mu^+ \mu^-$	$\pm 13$	?	$\pm 28$	
$\mathrm{H} \rightarrow \mathrm{invisible}$	< 0.2	?	< 0.4	

#### Chap. 1.3:

https://link.springer.com/content/pdf/10.1140/epjst/e2019-900045-4.pdf

## **Combination strategy**

Definition of signal processes:

- POIs and names
- $H \rightarrow$ (tautau, WW, ZZ) as auxiliary POIs

Definition of backgrounds

- Procs: WW, ZZ, Zgamma, rares?
- Constrained or unconstrained

### Separation of Z(vv) and vvH

- Necessary for 365 GeV, but also at 240 (for the width measurement)
- Split on missing mass at RECO/GEN level

## Fitting tools

All the fitting tools basically provide an implementation of the Poisson likelihood fit

- Combine (RooFit based)
- CombineTF (Tensorflow based)
- Custom RooFit

- $\rightarrow$  Combine and CombineTF are identical for binned fits
- $\rightarrow$  CombineTF does not support unbinned/analytic fits
- $\rightarrow$  Analytic fits should be identical to binned fits

## **Binning strategy**

Need to synchronize on the binning strategy:

- Important, especially for sensitivity of the rare process  $H \rightarrow ss$
- Extensive studies done with the hadronic analysis

Converged on:

- Start from uniform 1 GeV bins in recoil/mjj/etc. templates
- Rebin such that
  - At least one expected event in each bin (i.e. both signal and backgrounds)
  - Put 1e-4 event content in empty bins

## MC statistical uncertainties

Monte Carlo statistical uncertainties important to consider

Combine and CombineTF have the Barlow-Beeston light algo implemented

- Used in all CMS analyses
- Effectively assigns an equivalent stat. uncertainty of all processes combined -
- Should include the signal processes as well -

Usage:

- Combine: in the cards using: \* autoMCStats 0 1
- **CombineTF: arumgent line:** --binByBinStat -

Obviously, need to enable TH1::Sumw2()

## Systematic uncertainties

Currently only uncertainties on the background (normalization)

## Repository/versioning/lumi

Will make a repository so we can track the versionings of the cards

Need to converge on the lumis (final numbers to be used)