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1 Terrestrial Magnetic Field Effects on Large Photomultipliers

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6 Abstract

7 The effects of the Earth's magnetic field on the performance of large PMTs for a cubic-kilometer-scale neutrino telescope has
8 been studied. Measurements were performed for three Hamamatsu PMTs: two 8" R5912 types; one with a standard and the
9 other with a super bialkali photocathode, and a 10" R7081 type with a standard bialkali photocathode. The main
10 characteristics of the PMTs, such as detection efficiency, transit time, transit time spread, gain, peak-to-valley ratio, charge
11 resolution and fractions of spurious pulses were measured while varying the PMT orientations with respect to the Earth's
12 magnetic field. The measurements were performed both with and without a mu-metal cage magnetic shielding. For the 8"
13 PMTs the impact of the magnetic field was found to be smaller than for the 10" PMT. The magnetic shielding strongly
14 reduced the orientation-dependent variations measured for the 10" PMT and even improved the performance. Although less
15 pronounced, improvements were also measured for the 8" PMTs.

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19 1. Introduction

20 The performance of a large area photomultiplier
21 tube (PMT) is subject to significant variation due to
22 magnetic fields, in particular of the long trajectories
23 of electrons from the photocathode to the anode [1].
24 The main effect is de-focalization of the
25 photoelectrons arriving at the first dynode, which
26 affects timing properties, such Transit Time (TT) and

27 Transit Time Spread (TTS), and even the energy of
28 photoelectrons hitting the first dynode. This has an
29 influence on detection efficiency, gain and peak to
30 valley ratio of the PMT [2]. A secondary effect is the
31 deviation of electron trajectory in the amplification
32 chain, in particular between first and second dynodes,
33 can also contribute to the decrease of the gain and to
34 the degradation of the charge spectrum. With this in
35 mind, the influence of the Earth's magnetic field on
36 large area PMT candidates for a cubic-kilometer-
37 scale neutrino telescope was measured within the

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1 framework of the KM3NeT design study [3], in order
 2 to evaluate variations in PMT performance and to
 3 decide whether the use of magnetic shielding is
 4 necessary in the design of an optical module
 5 containing a single large area PMT. In this study,
 6 three large PMTs produced by Hamamatsu were
 7 measured. Two were R5912 types, with an 8"
 8 photocathode, and 10 stages. One of these (8" STD)
 9 had a standard bialkali photocathode ($QE \approx 25\%$ @
 10 400nm), while the other (8" HQE) had a super-
 11 bialkali photocathode ($QE \approx 32\%$ @ 400nm). The
 12 third PMT was a R7081 type, with a 10" standard
 13 bialkali photocathode (10" STD) and the same
 14 dynode structure as the R5912 [4].

15 2. Experimental procedure and setup

16 PMT responses to an injected light were
 17 measured while varying the orientation and
 18 inclination of the PMT relative to the Earth's
 19 magnetic field. First, the performance of "naked"
 20 PMTs without magnetic shielding was measured. To
 21 this purpose, a light-tight dark box (1x0.5x0.5m) was
 22 constructed that can be rotated horizontally and of
 23 which the inclination can be changed. A laser source
 24 (Picoquant PDL 800-B) attenuated to the condition
 25 of single photo-electrons was used, with a head of
 26 410nm wavelength which emitted light pulses of 50
 27 ps FWHM. The laser was pulsed at a frequency of 10
 28 KHz using an external generator. A second fixed
 29 PMT was used as monitor of the light source
 30 stability. An optical diffuser (Thorlabs, D1-C50 [5]),
 31 provided homogeneous illumination over the
 32 photocathode.

33 The measured values of the Earth's magnetic
 34 field in the area selected for the box were around 40
 35 micro-Tesla. The magnetic shield used was a wire
 36 cage, made of 1 mm diameter wire of mu-metal [6],
 37 composed of a hemispherical part and a second flat
 38 part with a central hole for the neck of the PMT. The
 39 shadowing effect on the photocathode was calculated
 40 to be less than 4%. The magnetic reduction factor
 41 inside the volume of the cage was measured with an
 42 average value of 4. Three PMT inclinations were
 43 studied: vertically downwards (Tilt=0°), horizontal
 44 (Tilt=90°) and 50 deg downwards (Tilt=50°). For
 45 each inclination, the PMT under test was rotated 360°

46 in the horizontal plane in 30° steps. All PMTs were
 47 powered using an ISEG PMT active base (type
 48 PHQ7081-i-2m), and set at the same gain condition
 49 of $1.5 \cdot 10^7$, for supply voltages of around 1650V.

50 3. Measurements and results

51 For each PMT position, the detection efficiency,
 52 gain, peak to valley (P/V) ratio, charge resolution, TT
 53 and TTS were measured simultaneously. The fraction
 54 of spurious pulses was also measured.

55 Tables 1-6 show the measured values. For all sets
 56 of measured parameters, the average values are
 57 given, together with the percentage of the variation,
 58 calculated as the percentage of the difference
 59 between maximum and minimum value, divided by
 60 the average value.

61 3.1 Detection Efficiency

62 The ratio between the number of detected pulses
 63 and the number of pulses emitted by the laser source
 64 defines the detection efficiency. Figure 1 shows the
 65 detection efficiency for the three PMTs vertically
 66 inclined (tilt=0°) as a function of orientation, with
 67 and without the mu-metal cage. Table 1 summarizes
 68 the measurements at the three different inclinations.
 69 Values were normalized to the maximum over all
 70 measurements.
 71

Efficiency		naked			shielded		
Tilt		8" STD	8" HQE	10" STD	8" STD	8" HQE	10" STD
0°	Ave	0.80	0.91	0.89	0.80	0.93	0.97
	Var%	13.08	6.63	22.51	3.27	1.39	6.27
50°	Ave	0.73	0.87	0.76	0.77	0.93	0.97
	Var%	22.15	15.68	48.82	3.89	4.28	5.75
90°	Ave	0.70	0.82	0.66	0.76	0.93	0.95
	Var%	2.86	3.67	15.27	1.59	2.37	1.79

73 Table 1. Detection efficiency measurements

74
 75 In the "naked" 8" PMTs the impact of the
 76 magnetic field was smaller than that measured in the
 77 "naked" 10" PMT. The use of the magnetic shield
 78 reduces considerably the variations for the 10" PMT.
 79 The increased Quantum Efficiency (QE) in the HQE

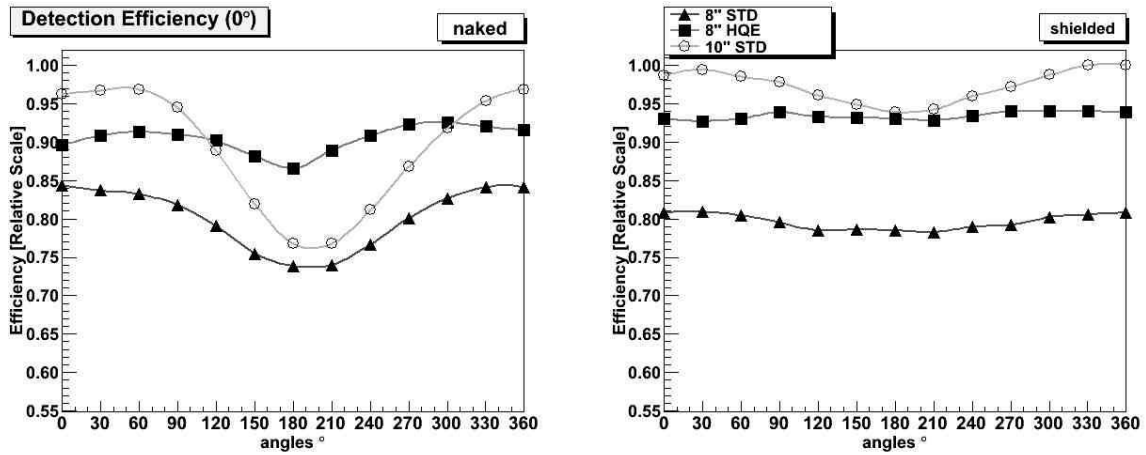


Fig. 1. Detection efficiency for PMTs vertically inclined (Tilt = 0°). On the left: PMTs naked. On the right: PMTs shielded

1 8'' PMT seems to compensate the smaller detection
2 area with respect to the 10'' PMT.

3

4 3.2 Charge Properties

5 The single photo-electron charge spectrum was
6 acquired for each PMT using a calibrated charge-
7 amplitude converter (mod. 7422 SILENA). Table 2
8 shows the results for each set of gain measurements.

9

Gain [1xE7]		naked			shielded		
Tilt		8'' STD	8'' HQE	10'' STD	8'' STD	8'' HQE	10'' STD
0°	Ave	1.51	1.63	1.50	1.59	1.63	1.56
	Var%	9.33	10.46	17.32	3.40	4.49	6.85
50°	Ave	1.45	1.64	1.35	1.56	1.67	1.53
	Var%	10.18	5.00	31.80	3.97	4.50	6.88
90°	Ave	1.42	1.60	1.35	1.54	1.64	1.52
	Var%	8.04	8.48	30.68	3.30	3.72	5.79

10 Table 2. Gain measurements

11

12 The variation in gain in the three orientations was
13 less than 10% for both "naked" 8'' PMTs, and
14 considerable (up to 30%) in the case of the "naked"
15 10'' PMT. The magnetic shield reduces variations in
16 both 8'' PMTs, with larger effect in the case of the
17 10'' PMT. Considering the P/V ratio (Table 3),
18 considerable variations for all the PMTs were
19 measured without the shield. Significant reductions
20 of these variations were seen with the magnetic

21 shield, with a small improvement in the average
22 values.

23

P/V ratio		naked			shielded		
Tilt		8'' STD	8'' HQE	10'' STD	8'' STD	8'' HQE	10'' STD
0°	Ave	2.19	2.15	1.61	2.41	2.22	1.90
	Var%	32.34	23.93	37.27	11.86	10.28	15.39
50°	Ave	1.92	1.93	1.39	2.30	2.22	1.83
	Var%	34.01	27.94	53.10	11.60	9.64	14.95
90°	Ave	1.73	1.77	1.21	2.24	2.12	1.76
	Var%	15.50	9.97	17.16	6.34	7.18	10.95

23 Table 3. Peak to Valley ratio measurements

24

25 With regard to the charge resolution
26 measurements (Table 4), the large effects due to the
27 magnetic field measured for the "naked" 10'' PMT
28 were largely reduced through use of the magnetic
29 shield.

Charge Res. %		naked			shielded		
Tilt		8'' STD	8'' HQE	10'' STD	8'' STD	8'' HQE	10'' STD
0°	Ave	49.31	55.59	64.18	46.02	55.66	57.90
	Var%	17.18	14.57	47.35	10.28	9.13	23.56
50°	Ave	51.66	55.78	87.38	46.63	54.70	60.49
	Var%	17.98	13.36	73.81	9.80	9.98	22.62
90°	Ave	53.24	57.51	97.73	46.98	55.89	62.70
	Var%	16.23	15.11	68.13	6.34	10.73	13.05

30 Table 4. Charge resolution (sigma) measurements

3.3 Time Properties

The results for TT on “naked” PMTs (Table 5), although not showing significant variations due to magnetic field, were slightly improved through the use of the mu-metal cage. Considering the TTS, calculated as FWHM, large variations for all “naked” PMTs were measured (Table 6). Strong reduction of these variations was seen with the magnetic shielding, but it was not accompanied by significant improvement of average values.

TT [ns]		naked			shielded		
Tilt		8" STD	8" HQE	10" STD	8" STD	8" HQE	10" STD
0°	Ave	103.0	101.6	110.0	103.0	101.6	109.9
	Var%	0.49	0.27	0.52	0.24	0.05	0.16
50°	Ave	103.3	101.9	110.1	103.2	101.5	110.1
	Var%	0.56	0.25	0.37	0.23	0.13	0.26
90°	Ave	103.5	101.8	110.3	103.2	101.7	110.1
	Var%	0.21	0.34	0.38	0.11	0.03	0.25

Table 5. Transit Time measurements

TIS [ns]		naked			shielded		
Tilt		8" STD	8" HQE	10" STD	8" STD	8" HQE	10" STD
0°	Ave	2.40	2.27	3.34	2.29	2.17	3.13
	Var%	22.13	11.88	15.28	7.85	5.52	1.60
50°	Ave	2.58	2.35	3.25	2.44	2.18	3.14
	Var%	16.69	12.74	11.07	8.20	4.14	4.14
90°	Ave	2.66	2.42	3.24	2.47	2.24	3.15
	Var%	10.52	7.43	10.18	5.67	1.79	5.72

Table 6. Transit Time Spread measurements (FWHM).

3.4 Fraction of Spurious Pulses

Spurious pulses are noise pulses, time-correlated with the PMT main response, which can be categorized into four different groups according to their causes and arrival times [7]: pre-pulses, delayed pulses, type 1 and type 2 after pulses. The percentage of spurious pulses with respect to the number of true pulses was measured for each of these groups. No significant magnetic field effects on the fraction of pre-pulses were measured. Considerable variation in delayed pulse fraction was measured only for the “naked” 10” PMT which was significantly reduced

by the mu-metal cage. In the case of type 1 and type 2 after pulses, no significant variations due to magnetic field were measured. Moreover, the standard bialkali 8” and 10” PMTs had similar fractions of type 1 and 2 after pulses, while the super bialkali photocathode 8” PMT had a larger fraction of type 1 and type 2 after pulses [8].

4. Summary

The influence of the Earth’s magnetic field on performance of three large photocathode area (8” and 10”) Hamamatsu PMTs was measured with and without magnetic shielding. Results confirmed that the performance of large area PMTs is significantly affected by orientation with respect to the Earth’s magnetic field. For the 8” PMTs the impact of the magnetic field was found to be smaller than in the 10” PMT. The magnetic shield significantly reduced the rotation and orientation-dependent performance variations in the 10” PMT and improved its performance. Less improvements were also seen in the case of 8” PMTs. The increased QE in the super bialkali 8” PMT almost compensates its smaller detection surface compared to the 10” PMT. No significant magnetic effects were measured on Transit Time and on the fraction of spurious pulses.

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References

- [1] E. Calvo et al., Nucl. Instr. Meth A 621, 2010, pp. 222-230
- [2] A. Tripathi UCLA - 3rd Beane Conference, 17-21 June 2002, Beane, France.
- [3] KM3NeT Technical Design Report ISBN 978-90-6488-033-9 km3net.org/KM3NeT-TDR.pdf
- [4] Hamamatsu official web-site, www.hamamatsu.com
- [5] www.thorlabs.com

- 1 [6] Institute for Theoretical and Experimental Physics, Moscow,
- 2 <http://www.itep.ru>
- 3 [7] S.Aiello et al., Nucl. Instr. Meth. A 605, (2009) 293-300
- 4 [8] N.Akchurin et al., Nucl. Instr. Meth. A 574 (2007) 121-126