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¹Terrestrial Magnetic Field Effects on Large Photomultipliers

6 **Abstract**

The effects of the Earth's magnetic field on the performance of large PMTs for a cubic-kilometer-scale neutrino telescope has been studied. Measurements were performed for three Hamamatsu PMTs: two 8" R5912 types; one with a standard and the 9 other with a super bialkali photocathode, and a 10" R7081 type with a standard bialkali photocathode. The main
10 characteristics of the PMTs, such as detection efficiency, transit time, transit time spread, gain, peak-t characteristics of the PMTs, such as detection efficiency, transit time, transit time spread, gain, peak-to-valley ratio, charge resolution and fractions of spurious pulses were measured while varying the PMT orientations with respect to the Earth's magnetic field. The measurements were performed both with and without a mu-metal cage magnetic shielding. For the 8" PMTs the impact of the magnetic field was found to be smaller than for the 10" PMT. The magnetic shielding strongly reduced the orientation-dependent variations measured for the 10" PMT and even improved the performance. Although less pronounced, improvements were also measured for the 8" PMTs.

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17 PACS: 85.60. Ha;
18 Keywords: Large

18 Keywords: Large area photomultiplier; magnetic shielding; Earth's magnetic field

19 **1. Introduction**

26 affects timing properties, such Transit Time (TT) and 20 The performance of a large area photomultiplier 31 21 tube (PMT) is subject to significant variation due to 32 22 magnetic fields, in particular of the long trajectories 33 23 of electrons from the photocathode to the anode $[1]$. 34 24 The main effect is de-focalization of the 35 25 photoelectrons arriving at the first dynode, which 36

27 Transit Time Spread (TTS), and even the energy of 28 photoelectrons hitting the first dynode. This has an 29 influence on detection efficiency, gain and peak to 30 valley ratio of the PMT [2]. A secondary effect is the deviation of electron trajectory in the amplification chain, in particular between first and second dynodes, can also contribute to the decrease of the gain and to the degradation of the charge spectrum. With this in mind, the influence of the Earth's magnetic field on large area PMT candidates for a cubic-kilometerscale neutrino telescope was measured within the

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1 framework of the KM3NeT design study [3], in order 2 to evaluate variations in PMT performance and to 3 decide whether the use of magnetic shielding is 4 necessary in the design of an optical module 5 containing a single large area PMT. In this study, 6 three large PMTs produced by Hamamatsu were 7 measured. Two were R5912 types, with an 8" 8 photocathode, and 10 stages. One of these (8" STD) 9 had a standard bialkali photocathode (QE≈25% @ 10 400nm), while the other (8" HQE) had a super-

11 bialkali photocathode (QE≈ 32% @ 400nm). The

12 third PMT was a R7081 type, with a 10" standard

13 bialkali photocathode (10" STD) and the same

14 dynode structure as the R5912 [4].

15 **2. Experimental procedure and setup**

16 PMT responses to an injected light were 17 measured while varying the orientation and 18 inclination of the PMT relative to the Earth's 19 magnetic field. First, the performance of "naked" 20 PMTs without magnetic shielding was measured. To 21 this purpose, a light-tight dark box (1x0.5x0.5m) was 22 constructed that can be rotated horizontally and of 23 which the inclination can be changed. A laser source 24 (Picoquant PDL 800-B) attenuated to the condition 25 of single photo-electrons was used, with a head of 26 410nm wavelength which emitted light pulses of 50 27 ps FWHM. The laser was pulsed at a frequency of 10 28 KHz using an external generator. A second fixed 29 PMT was used as monitor of the light source 30 stability. An optical diffuser (Thorlabs, D1-C50 [5]), 31 provided homogeneous illumination over the 32 photocathode. 61 72

33 The measured values of the Earth's magnetic 34 field in the area selected for the box were around 40 35 micro-Tesla. The magnetic shield used was a wire 36 cage, made of 1 mm diameter wire of mu-metal [6], 37 composed of a hemispherical part and a second flat 38 part with a central hole for the neck of the PMT. The 39 shadowing effect on the photocathode was calculated 40 to be less than 4%. The magnetic reduction factor 41 inside the volume of the cage was measured with an 42 average value of 4. Three PMT inclinations were 43 studied: vertically downwards (Tilt=0°), horizontal 44 (Tilt=90°) and 50 deg downwards (Tilt=50°). For 45 each inclination, the PMT under test was rotated 360° 74

in the horizontal plane in 30° steps. All PMTs were 47 powered using an ISEG PMT active base (type PHQ7081-i-2m), and set at the same gain condition 49 of 1.5 10^7 , for supply voltages of around 1650V.

50 **3. Measurements and results**

51 For each PMT position, the detection efficiency, 52 gain, peak to valley (P/V) ratio, charge resolution, TT 53 and TTS were measured simultaneously. The fraction 54 of spurious pulses was also measured.

55 Tables 1-6 show the measured values. For all sets 56 of measured parameters, the average values are 57 given, together with the percentage of the variation, 58 calculated as the percentage of the difference 59 between maximum and minimum value, divided by 60 the average value.

62 *3.1 Detection Efficiency*

63 The ratio between the number of detected pulses 64 and the number of pulses emitted by the laser source defines the detection efficiency. Figure 1 shows the 66 detection efficiency for the three PMTs vertically inclined (tilt= 0°) as a function of orientation, with and without the mu-metal cage. Table 1 summarizes the measurements at the three different inclinations. Values were normalized to the maximum over all 71 measurements.

73 Table 1. Detection efficiency measurements

75 In the "naked" 8" PMTs the impact of the 76 magnetic field was smaller than that measured in the 77 "naked" 10" PMT. The use of the magnetic shield 78 reduces considerably the variations for the 10" PMT. 79 The increased Quantum Efficiency (QE) in the HQE

Fig. 1. Detection efficiency for PMTs vertically inclined (Tilt = 0°). On the left: PMTs naked. On the right: PMTs shielded^o

2 area with respect to the 10" PMT.

4 *3.2 Charge Properties*

5 The single photo-electron charge spectrum was 6 acquired for each PMT using a calibrated charge-7 amplitude converter (mod. 7422 SILENA). Table 2 8 shows the results for each set of gain measurements. 9

10 Table 2. Gain measurements

11

3

12 The variation in gain in the three orientations was 13 less than 10% for both "naked" 8" PMTs, and 14 considerable (up to 30%) in the case of the "naked" 15 10" PMT. The magnetic shield reduces variations in 16 both 8" PMTs, with larger effect in the case of the 17 10" PMT. Considering the P/V ratio (Table 3), 18 considerable variations for all the PMTs were 19 measured without the shield. Significant reductions 20 of these variations were seen with the magnetic

1 8" PMT seems to compensate the smaller detection 21 shield, with a small improvement in the average 22 values

vaiues.									
P/V ratio		naked			shielded				
Tilt		8" STD	$\mathbf{R}^{\prime\prime}$ HQE	10" STD	\mathbf{R}^{11} STD	8" HQE	10" STD		
0°	Ave	2.19	2.15	1.61	2.41	2.22	1.90		
	$\bf Var$ %	32.34	23.93	37.27	11.86	10.28	15.39		
50°	Ave	1.92	1.93	1.39	2.30	2.22	1.83		
	$Var_{\%}$	34.01	27.94	53.10	11.60	9.64	14.95		
90°	Ave	1.73	1.77	1.21	2.24	2.12	1.76		
	$Var_{\%}$	15.50	9.97	17.16	6.34	7.18	10.95		

23 Table 3. Peak to Valley ratio measurements

25 With regard to the charge resolution 26 measurements (Table 4), the large effects due to the 27 magnetic field measured for the "naked" 10" PMT 28 were largely reduced through use of the magnetic 29 shield.

Charge Res. $\frac{6}{10}$		naked			shielded		
Tilt		\mathbf{R}^{11} STD	\mathbf{R}^{11} HOE	10" STD	\mathbf{R}^{11} STD	$\mathbf{R}^{\prime\prime}$ HOE	10" STD
0°	Ave	49.31	55.59	64.18	46.02	55.66	57.90
	Var%	17.18	14.57	47.35	10.28	9.13	23.56
50°	Ave	51.66	55.78	87.38	46.63	54.70	60.49
	Var%	17.98	13.36	73.81	9.80	9.98	22.62
90°	Ave	53.24	57.51	97.73	46.98	55.89	62.70
	Var%	16.23	15.11	68.13	6.34	10.73	13.05

³⁰ Table 4. Charge resolution (sigma) measurements

1 *3.3 Time Properties*

2 The results for TT on "naked" PMTs (Table 5), 29 3 although not showing significant variations due to 4 magnetic field, were slightly improved through the 5 use of the mu-metal cage. Considering the TTS, 6 calculated as FWHM, large variations for all "naked" 7 PMTs were measured (Table 6). Strong reduction of 8 these variations was seen with the magnetic

- 9 shielding, but it was not accompanied by significant
- 10 improvement of average values.

11

12 Table 5. Transit Time measurements 13

14 Table 6. Transit Time Spread measurements (FWHM). 15

16 *3.4 Fraction of Spurious Pulses*

17 Spurious pulses are noise pulses, time-correlated 58 18 with the PMT main response, which can be 19 categorized into four different groups according to 20 their causes and arrival times [7]: pre-pulses , delayed 21 pulses, type 1 and type 2 after pulses. The percentage 22 of spurious pulses with respect to the number of true 23 pulses was measured for each of these groups. No 24 significant magnetic field effects on the fraction of 63 25 pre-pulses were measured. Considerable variation in 26 delayed pulse fraction was measured only for the 27 "naked" 10" PMT which was significantly reduced

28 by the mu-metal cage. In the case of type 1 and type 29 2 after pulses, no significant variations due to magnetic field were measured. Moreover, the standard bialkali 8" and 10" PMTs had similar fractions of type 1 and 2 after pulses, while the super bialkali photocathode 8" PMT had a larger fraction of type 1 and type 2 after pulses [8].

35 **4. Summary**

36 The influence of the Earth's magnetic field on 37 performance of three large photocathode area (8" and 38 10") Hamamatsu PMTs was measured with and 39 without magnetic shielding. Results confirmed that 40 the performance of large area PMTs is significantly 41 affected by orientation with respect to the Earth's 42 magnetic field. For the 8" PMTs the impact of the 43 magnetic field was found to be smaller than in the 44 10" PMT. The magnetic shield significantly reduced 45 the rotation and orientation-dependent performance 46 variations in the 10" PMT and improved its 47 performance. Less improvements were also seen in 48 the case of 8" PMTs. The increased QE in the super 49 bialkali 8" PMT almost compensates its smaller 50 detection surface compared to the 10" PMT. No 51 significant magnetic effects were measured on 52 Transit Time and on the fraction of spurious pulses.

53 **Acknowledgment**

54 The KM3NeT project is supported under EU FP6 55 Contract no. 011937 and FP7 Grant agreement no. 56 212525. The author thanks O. Kalekin, P. Keller and 57 P. Vernin for their presence and technical support at the start of this work.

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