#### **Non-Standard Neutrino Interactions**

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### Introduction: NSI

Generic new physics affecting v oscillations can be parameterized as 4-fermion Non-Standard Interactions:

Production or detection of a  $v_{\beta}$  associated to a  $l_{\alpha}$ 

$$2\sqrt{2}G_{F}\varepsilon_{\alpha\beta}\left(\overline{\nu}_{\beta}\gamma^{\mu}P_{L}l_{\alpha}\right)\left(\bar{f}\gamma_{\mu}P_{L,R}f'\right)$$

So that 
$$|\nu_{\alpha}^{s}\rangle = |\nu_{\alpha}\rangle + \sum_{\beta=e,\mu,\tau} \varepsilon_{\alpha\beta} |\nu_{\beta}\rangle$$

 $\pi \to \mu + \nu_{\beta} \qquad n + \nu_{\beta} \to p + l_{\alpha}$ 

Y. Grossman hep-ph/9507344

### Direct bounds on prod/det NSI

$$2\sqrt{2}G_{F}\varepsilon^{ud}_{\alpha\beta}\left(\bar{l}_{\beta}\gamma^{\mu}P_{L}\nu_{\alpha}\right)\left(\bar{u}\gamma_{\mu}P_{L,R}d\right)$$

$$2\sqrt{2}G_{F}\varepsilon_{\alpha\beta}^{\mu e}\left(\overline{\mu}\gamma^{\mu}P_{L}\nu_{\beta}\right)\left(\overline{\nu}_{\alpha}\gamma_{\mu}P_{L}e\right)$$

	(0.041	0.025	0.041	(0.02)	5 0.03	0.03
$\left  \mathcal{E}^{ud} \right  <$	0.026	0.078	0.013	$\left  \varepsilon^{\mu e} \right  < 0.025$	5 0.03	0.03
	0.12	0.013	0.13	0.02	5 0.03	0.03

bounds order ~10<sup>-2</sup> from comparisons of measurements of G<sub>F</sub>:  $\mu$ ,  $\tau$ ,  $\pi$  decays, CKM universality, M<sub>W</sub> + M<sub>Z</sub>...

C. Biggio, M. Blennow and EFM 0907.0097

### Introduction: NSI

Non-Standard v scattering off matter can also be parameterized as 4-fermion Non-Standard Interactions:

$$2\sqrt{2}G_{F}\varepsilon_{\alpha\beta}^{m}\left(\overline{\nu}_{\beta}\gamma^{\mu}P_{L}\nu_{\alpha}\right)\left(\overline{f}\gamma_{\mu}P_{L,R}f\right)$$
  
so that  $\tilde{V}_{MSW} = a_{CC}\begin{pmatrix}1+\varepsilon_{ee}^{m}&\varepsilon_{e\mu}^{m}&\varepsilon_{e\tau}^{m}\\\varepsilon_{e\mu}^{m*}&\varepsilon_{\mu\mu}^{m}&\varepsilon_{\mu\tau}^{m}\\\varepsilon_{e\tau}^{m*}&\varepsilon_{\mu\tau}^{m*}&\varepsilon_{\tau\tau}^{m}\end{pmatrix}$ 

$$V_{\alpha} \rightarrow V_{\beta}$$
 in matter  $f = e, u, d$ 

If matter NSI are uncorrelated to production and detection direct bounds are mainly from  $\nu$  scattering off *e* and nuclei

$$2\sqrt{2}G_{F}\varepsilon_{\alpha\beta}^{m}\left(\overline{\nu}_{\beta}\gamma^{\mu}P_{L}\nu_{\alpha}\right)\left(\bar{f}\gamma_{\mu}P_{L,R}f\right)$$

 $\left| \varepsilon_m^e \right| < \begin{pmatrix} 0.14 & 0.1 & 0.44 \\ 0.1 & 0.03 & 0.1 \\ 0.44 & 0.1 & 0.5 \end{pmatrix} \left| \varepsilon_m^u \right| < \begin{pmatrix} 1 & 0.05 & 0.5 \\ 0.05 & 0.008 & 0.05 \\ 0.5 & 0.05 & 3 \end{pmatrix} \left| \varepsilon_m^d \right| < \begin{pmatrix} 0.6 & 0.05 & 0.5 \\ 0.05 & 0.015 & 0.05 \\ 0.5 & 0.05 & 6 \end{pmatrix}$ 

Rather weak bounds...

... can they be saturated avoiding additional constraints?

S. Davidson, C. Peña garay, N. Rius and A. Santamaria hep-ph/0302093 J. Barranco, O. G. Miranda, C. A. Moura and J. W. F. Valle hep-ph/0512195 J. Barranco, O. G. Miranda, C. A. Moura and J. W. F. Valle 0711.0698 C. Biggio, M. Blennow and EFM 0902.0607



Tension between MINOS nu and antinu data P. Vahle @ Neutrino 2010



Tension between MINOS nu and antinu data P. Vahle @ Neutrino 2010



R. Van de Water @ Neutrino 2010



R. Van de Water @ Neutrino 2010

Can be accommodated with production/detection NSI + sterile neutrinos  $\mathcal{E}_{e\mu} \sim 0.01$ 

E. Akhmedov and T. Schwetz 1007.4171

### Gauge invariance

However 
$$2\sqrt{2}G_F \varepsilon^m_{lphaeta} \left( \overline{v}_{eta} \gamma^\mu P_L v_{lpha} \right) \left( \overline{f} \gamma_\mu P_{L,R} f \right)$$

is related to 
$$2\sqrt{2}G_F \varepsilon^m_{lphaeta} (\bar{l}_{eta} \gamma^\mu P_L l_{lpha}) (\bar{f} \gamma_\mu P_{L,R} f)$$

#### by gauge invariance and very strong bounds exist

$$\mathcal{E}_{e\mu}^{m} < \sim 10^{-6}$$
  
 $\mathcal{E}_{e\tau}^{m} < \sim 10^{-4}$   
 $\mathcal{E}_{\mu\tau}^{m} < \sim 10^{-4}$ 

$$\mu \rightarrow e \gamma$$
  
$$\mu \rightarrow e \text{ in nuclei}$$
  
$$\tau \text{ decays}$$

S. Bergmann et al. hep-ph/0004049 Z. Berezhiani and A. Rossi hep-ph/0111147 S. Antusch, M. Blennow, EFM and T. Ota, 1005.0756

We search for gauge invariant SM extensions satisfying:

- Matter NSI are generated at tree level
- 4-charged fermion ops not generated at the same level
- No cancellations between diagrams with different messenger particles to avoid constraints
- The Higgs Mechanism is responsible for EWSB

S. Antusch, J. Baumann and EFM 0807.1003 B. Gavela, D. Hernández, T. Ota and W. Winter 0809.3451

At d=6 only one direct possibility: charged scalar singlet



M. Bilenky and A. Santamaria hep-ph/9310302

Since  $\lambda_{\alpha\beta} = -\lambda_{\beta\alpha}$  only  $\varepsilon_{\mu\mu}$ ,  $\varepsilon_{\mu\tau}$  and  $\varepsilon_{\tau\tau} \neq 0$ Very constrained:

$$\begin{array}{lll} |\varepsilon_{\mu\mu}^{m,e_{\rm L}}| &< 8.2 \cdot 10^{-4} & \mu \rightarrow e \gamma \\ |\varepsilon_{\tau\tau}^{m,e_{\rm L}}| &< 8.4 \cdot 10^{-3} & \mu \, {\rm decays} \\ |\varepsilon_{\mu\tau}^{m,e_{\rm L}}| &< 1.9 \cdot 10^{-3} & {\rm CKM \ unitarity} \end{array}$$

F. Cuypers and S. Davidson hep-ph/9310302 S. Antusch, J. Baumann and EFM 0807.1003

#### At d=6 indirect way: fermion singlets







$$L = i \overline{\nu}_{\alpha} \partial K_{\alpha\beta} \nu_{\beta} + \overline{\nu}_{\alpha} M_{\alpha\beta} \nu_{\beta} - \frac{g}{\sqrt{2}} \left( W_{\mu}^{+} \overline{l}_{\alpha} \gamma^{\mu} P_{L} \nu_{\alpha} + h.c. \right) - \frac{g}{\cos \theta_{W}} \left( Z_{\mu} \overline{\nu}_{\alpha} \gamma^{\mu} P_{L} \nu_{\alpha} + h.c. \right) + \dots$$

$$L = i \overline{v}_{\alpha} \partial K_{\alpha\beta} v_{\beta} + \overline{v}_{\alpha} M_{\alpha\beta} v_{\beta} - \frac{g}{\sqrt{2}} \left( W_{\mu}^{+} \overline{l}_{\alpha} \gamma^{\mu} P_{L} v_{\alpha} + h.c. \right) - \frac{g}{\cos \theta_{W}} \left( Z_{\mu} \overline{v}_{\alpha} \gamma^{\mu} P_{L} v_{\alpha} + h.c. \right) + \dots$$
  
Diagonal mass and canonical kinetic terms  
$$L = i \overline{v}_{i} \partial v_{i} + \overline{v}_{i} m_{ii} v_{i} - \frac{g}{\sqrt{2}} \left( W_{\mu}^{+} \overline{l}_{\alpha} \gamma^{\mu} P_{L} N_{\alpha i} v_{i} + h.c. \right) - \frac{g}{\cos \theta_{W}} \left( Z_{\mu} \overline{v}_{i} \gamma^{\mu} P_{L} (N^{\dagger} N)_{ij} v_{j} + h.c. \right) + \dots$$

$$L = i \overline{v}_{\alpha} \partial K_{\alpha\beta} v_{\beta} + \overline{v}_{\alpha} M_{\alpha\beta} v_{\beta} - \frac{g}{\sqrt{2}} (W_{\mu}^{+} \overline{l}_{\alpha} \gamma^{\mu} P_{L} v_{\alpha} + h.c.) - \frac{g}{\cos \theta_{W}} (Z_{\mu} \overline{v}_{\alpha} \gamma^{\mu} P_{L} v_{\alpha} + h.c.) + \dots$$
  
Diagonal mass and canonical kinetic terms  
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$$V_{\alpha} = N_{\alpha i} v_{i} \qquad N \text{ is not unitary}$$

$$L = i \overline{\nu}_{\alpha} \partial K_{\alpha\beta} \nu_{\beta} + \overline{\nu}_{\alpha} M_{\alpha\beta} \nu_{\beta} - \frac{g}{\sqrt{2}} (W_{\mu}^{+} \overline{l}_{\alpha} \gamma^{\mu} P_{L} \nu_{\alpha} + h.c.) - \frac{g}{\cos \theta_{W}} (Z_{\mu} \overline{\nu}_{\alpha} \gamma^{\mu} P_{L} \nu_{\alpha} + h.c.) + \dots$$
Diagonal mass and canonical kinetic terms
$$L = i \overline{\nu}_{i} \partial \nu_{i} + \overline{\nu}_{i} m_{ii} \nu_{i} - \frac{g}{\sqrt{2}} (W_{\mu}^{+} \overline{l}_{\alpha} \gamma^{\mu} P_{L} N_{\alpha i} \nu_{i} + h.c.) - \frac{g}{\cos \theta_{W}} (Z_{\mu} \overline{\nu}_{i} \gamma^{\mu} P_{L} (N^{\dagger} N)_{ij} \nu_{j} + h.c.) + \dots$$
unchanged
$$\nu_{\alpha} = N_{\alpha i} \nu_{i}$$
 $N \text{ is not unitary}$ 

### $(NN^{\dagger})$ from decays

• W decays

 $\sum_{i=1}^{V_i} \sum_{i=1}^{V_i} \sum_{j=1}^{V_i} \sum_{j=1}^{V_i}$ 

Invisible Z

Universality tests

Rare leptons decays

Ιαα  $\sum (NN^{\dagger})$ Info on  $(NN^{\dagger})_{\alpha\alpha}$ NN Ιαα Info on  $(NN^{\dagger})_{\alpha\beta}$ W

After integrating out W and Z neutrino NSI induced

 $v_i$ 

 $l_{\beta}$ 

 $l_{\alpha}$ 

### $(NN^{\dagger})$ from decays

$$\left|\varepsilon\right| = 2\left|\delta - NN^{\dagger}\right| < \begin{pmatrix} 2.0 \cdot 10^{-3} & 5.9 \cdot 10^{-5} & 1.6 \cdot 10^{-3} \\ 5.9 \cdot 10^{-5} & 8.2 \cdot 10^{-4} & 1.0 \cdot 10^{-3} \\ 1.6 \cdot 10^{-3} & 1.0 \cdot 10^{-3} & 2.6 \cdot 10^{-3} \end{pmatrix}$$
 Experimentally

E. Nardi, E. Roulet and D. Tommasini hep-ph/9503228 D. Tommasini, G. Barenboim, J. Bernabeu and C. Jarlskog hep-ph/9503228 S. Antusch, C. Biggio, EFM, B. Gavela and J. López Pavón hep-ph/0607020 S. Antusch, J. Baumann and EFM 0807.1003

## Non-Unitarity at a vFactory



Golden channel at vFactory is sensitive to  $\varepsilon_{re}$ 

 $v_{\mu}$  disappearance channel linearly sensitive to  $\varepsilon_{\tau\mu}$  through matter effects Near  $\tau$  detectors can improve the bounds on  $\varepsilon_{\tau e}$  and  $\varepsilon_{\tau\mu}$ Combination of near and far detectors sensitive to the new CP phases

> S. Antusch, M. Blennow, EFM and J. López-pavón 0903.3986 See also EFM, B. Gavela, J. López Pavón and O. Yasuda hep-ph/0703098; S. Goswami and T. Ota 0802.1434; G. Altarelli and D. Meloni 0809.1041,....

Without avoiding 4-charged fermion ops at the same level there is no direct way at d=6 to induce quark NSI

Constraints from gauge invariance at d=6 on flavour changing NSI very strong:

Operator	$(\mathcal{C}^1_{LQ})_{\alpha}^{\tau}$	$(\mathcal{C}_{LQ}^{\Im})_{\alpha}{}^{\tau}$	$(\mathcal{C}_{ED})_{\alpha}^{\ \tau}$	$(\mathcal{C}_{EU})_{\alpha}^{\ \tau}$	$(\mathcal{C}_{ED}^{\dagger})_{\alpha}^{\tau}$	$(\mathcal{C}_{EU}^{\dagger})_{\alpha}^{\tau}$
$\alpha = \mu$	$2.1 \cdot 10^{-4}$	$1.7 \cdot 10^{-4}$	$7.2 \cdot 10^{-4}$	$7.2 \cdot 10^{-4}$	$6.2 \cdot 10^{-4}$	$6.2 \cdot 10^{-4}$
$\alpha = e$	$2.4 \cdot 10^{-4}$	$1.6 \cdot 10^{-4}$	$6.9 \cdot 10^{-4}$	$7.1 \cdot 10^{-4}$	$6.0 \cdot 10^{-4}$	$6.1 \cdot 10^{-4}$

But pseudoscalar operators have chiral enhancement in pion decay  $\rightarrow$  production NSI ~ 0.01 in pion decay

S. Antusch, M. Blennow, EFM and T. Ota, 1005.0756

## Chiral "enhancement" of NSI

Production NSI from meson decay can have be enhanced with respect to SM for pseudoscalar ops.

$$2\sqrt{2}G_{F}\left(\overline{\nu}_{\alpha}\gamma^{\mu}P_{L}l_{\alpha}\right)\left(\overline{d}\gamma_{\mu}P_{L,R}u\right) \qquad 2\sqrt{2}G_{F}\varepsilon_{\alpha\beta}\left(\overline{\nu}_{\beta}P_{L}l_{\alpha}\right)\left(\overline{d}P_{L,R}u\right)$$

 $\pi$  or K decay requires chirality in SM

$$\left\langle 0 \middle| \overline{u} \gamma_{\mu} \gamma_{5} d \middle| \pi^{-} \right\rangle = i f_{\pi} p_{\mu} \qquad \left\langle 0 \middle| \overline{u} \gamma_{5} d \middle| \pi^{-} \right\rangle = -i f_{\pi} \frac{m_{\pi}^{2}}{m_{\mu} + m_{d}}$$
$$\frac{m_{\pi}^{2}}{m_{\mu} (m_{\mu} + m_{d})} \approx 20$$

0

## MINSIS

Main Injector Non Standard Interactions Search

OPERA-like detector at MINOS near detector site to search for  $v_{\tau}$  appearance in the NuMI beam at 1 km baseline

Could reach sensitivities ~10<sup>-6</sup>  $\rightarrow$  test production/detection NSI mixing the  $\mu$  and  $\tau$  sectors down to 10<sup>-3</sup>

- Models leading "naturally" to NSI imply:
  - O(10<sup>-3</sup>) bounds on the NSI
  - Relations between matter and production/detection NSI
- Probing O(10<sup>-3</sup>) NSI at future facilities very challenging but not impossible, near detectors good probes
- Saturating the mild model-independent bounds on matter NSI is hard: chiral enhancement on π decay or flavour conserving NSI

### **Sterile Neutrinos**

- Similar to NSI if new ∆m<sup>2</sup> is large oscillations will happen at short baselines:
  - Appearance seraches limited by background
  - Disappearance searches limited by systematics
- If E is small, oscillation length can be few m oscillations inside the detector!
- Expectation better than for NSI, favoured region for O(10<sup>-2</sup>)- O(10<sup>-3</sup>) probabilities