



# Outlook for Theoretical Precision of the Luminosity at Future Lepton Colliders<sup>a</sup>

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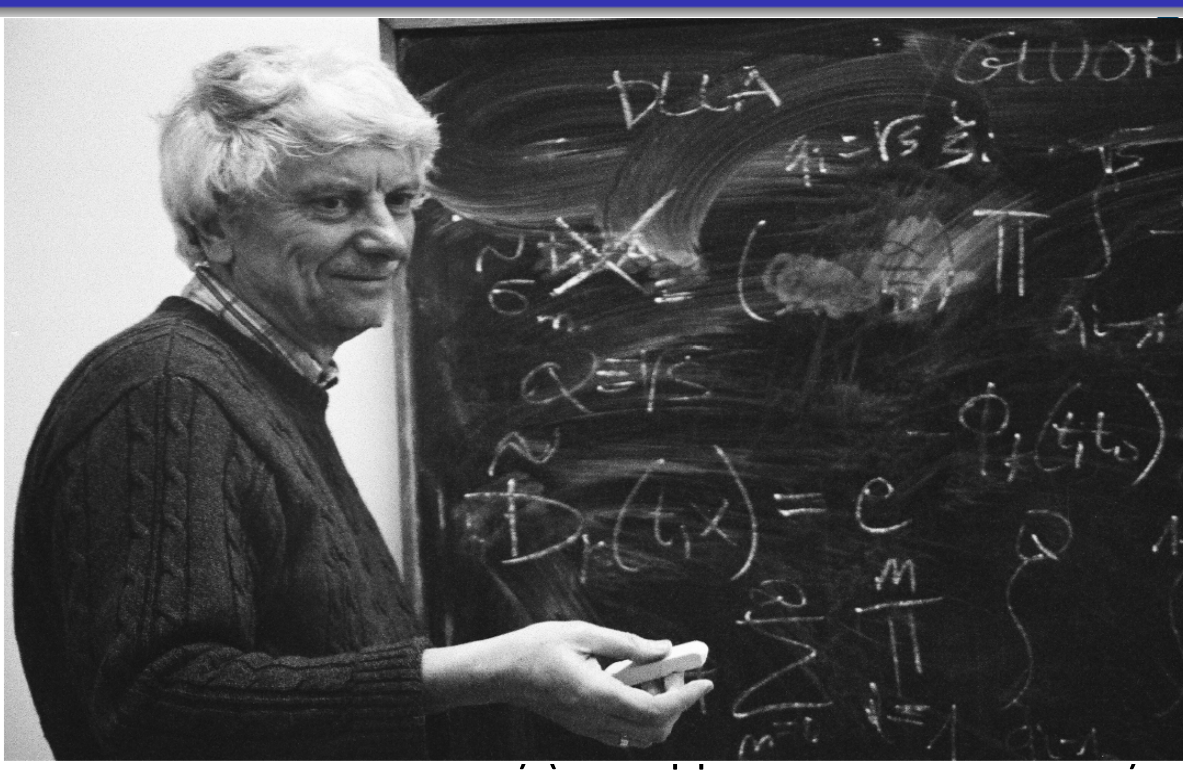
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## In Memoriam to Staszek Jadach



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## Introduction, Lumi Basics

Bhabha cross sect. depends on detector acceptance angles

$$\sigma_{Bh} \approx 4\pi\alpha^2 \left( \frac{1}{t_{min}} - \frac{1}{t_{max}} \right) = 4\pi\alpha^2 \left( \frac{t_{max} - t_{min}}{t^2} \right), \quad \bar{t} = \sqrt{t_{min} t_{max}}$$

$\bar{t}$  is the characteristic scale of the process  
 $\bar{t}/s$  is the suppression factor between s- and t-channel contributions

Machine	$\theta_{min}^+ \theta_{max}^-$ [mrad]	$\sqrt{s}$ [GeV]	$\bar{t}/s \approx \theta^2/4$	$\sqrt{t}$ [GeV]
LEP	28-50	$M_Z$	$3.5 \times 10^{-4}$	1.70
FCCee	64-86	$M_Z$	$13.7 \times 10^{-4}$	3.37
FCCee	64-86	240	$13.7 \times 10^{-4}$	8.9
FCCee	64-86	350	$13.7 \times 10^{-4}$	13.0
ILC	31-77	500	$6.0 \times 10^{-4}$	12.2
ILC	31-77	1000	$6.0 \times 10^{-4}$	24.4
CLIC	39-134	3000	$13.0 \times 10^{-4}$	108
CEPC	26-105	$M_Z$	$8.8 \times 10^{-4}$	2.38
CEPC	26-105	240	$8.8 \times 10^{-4}$	6.27

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## Current Situation, Related to LEP

### LEP update 2018(2019)

Type of correction / Error	1999	Update 2018
(a) Photonic $O(L_s \alpha^2)$	0.027% [5]	0.027%
(b) Photonic $O(L_s^2 \alpha^2)$	0.015% [6]	0.015%
(c) Vacuum polariz.	0.040% [7,8]	0.013% (0.011% (J))
(d) Light pairs	0.030% [10]	0.010% [18,19]
(e) s-channel Z-exchange	0.015% [11,12]	0.015%
(f) Up-down interference	0.0014% [27]	0.0014%
(f) Technical Precision	-	(0.027%)
Total	0.061% [13]	0.038% (0.037% (J))

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J <-> Janot and Jadach, Phys. Lett. B 803 (2019) 135319

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## Current Situation, Related to LEP

- Implied Upgrade of BHLUMI: 'still not urgent' in Review for US DOE OHEP Funding Program
- Recent Work by Banerjee et al., PLB 820 (2021) 136547, T. Engel et al., in arXiv:2203.12557, contacts

Type of correction / Error	Update 2018	FCCee forecast
(a) Photonic $O(L_s \alpha^2)$	0.027%	$0.6 \times 10^{-3}$
(b) Photonic $O(L_s^2 \alpha^2)$	0.015%	$0.1 \times 10^{-4}$
(c) Vacuum polariz.	0.014% [25]	$0.6 \times 10^{-4}$
(d) Light pairs	0.010% [18,19]	$0.5 \times 10^{-4}$
(e) Z and s-channel $\gamma$ exchange	0.009% [11]	$0.1 \times 10^{-4}$
(f) Up-down interference	0.000% [27]	$0.1 \times 10^{-4}$
(f) Technical Precision	0.027%	$0.1 \times 10^{-4}$
Total	0.097%	$1.0 \times 10^{-4}$

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## Current Situation, Related to LEP

- References on Virtual Correction to Bremsstrahlung at  $O(\alpha^2 L)$
- S. Jadach et al., Phys. Lett. B 377 (1996) 168.
  - S. Jadach et al., Phys. Lett. B 450 (1999) 262.
  - S. Jadach et al., Comput. Phys. Commun. 70 (1992) 305.
  - S. Jadach et al., Comput. Phys. Commun. 102 (1997) 229.
  - S. Jadach et al., Acta Phys. Pol. B30 (1999) 1745.

Reference to  $O(\alpha^2 L^2)$

- S. Jadach and B.F.L. Ward, Phys. Lett. B 389 (1996) 129.

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## $(\gamma_s + Z_s + \gamma_t + Z_t) \otimes \gamma$ EW interferences

- Included in BHLUMI:  $(\gamma_s + Z_s) \otimes \gamma$
- To be added:
  - complete Born - trivial
  - complete  $O(\alpha_{EW})$  - known, e.g. BHWIDE
- Error:  $O(\alpha_{EW}^2)$ 
  - estimated at FCCee( $M_Z$ ) based on analysis of S. Jadach et al. Phys. Lett. B790 (2019) 314 - from BHWIDE
  - estimated at other energies/angles based on analysis done with  $O(\alpha_{EW})$  DIZET/ZFITTER (by changing switch NPAR(2) from 2 to 3) M. Battaglia, S. Jadach, D. Bardin, eConf C010630 (2001) E3015, <http://www.slac.stanford.edu/econf/C010630/papers/E3015.PDF> for the energies of 800 GeV and 3 TeV.
  - Extrapolation from 800 to 350/240 GeV not done  $\Rightarrow$  error likely overestimated (factor of 2-3 ???)
  - Error at higher  $\bar{t}/M_Z^2$  almost entirely from  $\gamma_t \otimes Z_t$  interference
  - Amplitude-level exponentiation (KKMC-style) needed to account for leading  $O(\alpha_{EW}^2)$  corrs.

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## QED photonic up-down interference

- Missing in BHLUMI
  - size at  $O(\alpha)$ :  $0.07 \times \bar{t}_{ex}/s$  - easy to include,  $\bar{t}_{ex}/s$  depends only on angles
  - LEP - FCCee:  $t/s$  grows 4 times (LEP  $\rightarrow$  ILC: 2 times)
- Error: h.o.t. - suppressed by  $(\alpha/\pi) \ln(\bar{t}_{ex}/m_e^2)$  times safety factor of 2 ( $O(\alpha_{EW}^2)$  calculations exist) - almost negligible

## Vacuum polarisation

- Uncertainty due to vacuum polarisation:
  - $\delta_{\nu_{\mu\mu}}(\bar{t}) = 2\delta_{\nu_{\mu\mu}}(\bar{t})/\alpha_{\mu\mu}(\bar{t})$
  - $\delta_{\nu_{\mu\mu}}(\bar{t})$  from (based on R-ratio measured at low energies) F. Jegerlehner, CERN Yellow Reports: Monographs 3 (2020) 9-37
  - $\alpha_{\mu\mu}(\bar{t})$  from F. Jegerlehner, Nucl. Phys. Proc. Suppl. 162 (2006) 22-32
- By FCCee operation time factor of 2 improvement expected (F. Jegerlehner)

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## Light pairs

- Current state of the art: BHLUMI + external four-fermion code + virtual semi-analytical corrections
  - P. Janot and S. Jadach, Phys. Lett. B 803 (2020) 135319
- Included components:
  - $\mu\mu$ -pair,  $\nu\nu$ -pair,  $\tau\tau$ -pair, qq-pair with s-channel photonic emissions (FERMUS, KORALW)
  - result for LEP:  $4 \times 10^{-4} \pm 1 \times 10^{-4}$
- future prospects for external 4fermion code scenario
  - error components:
    - $4f + \gamma$  (25% of  $4f$ ) - s vs. t mismatch  $\sim 30\%$
    - $O(\alpha)$  4fermion calculations exist for selected final states
    - $4f + 2\gamma$ ,  $6f$
- future prospects for BHLUMI upgrade scenario
  - error components:
    - $4f + \gamma$  - absent - correct f-channel behavior (LL+soft),  $O(\alpha)$  4fermion likely not needed
    - $O(\alpha)$  4fermion likely not needed
    - $4f + 2\gamma$  - included via exponentiation + LL,  $6f$

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## Light pairs

- Extrapolation to other energies/angles
- use LEP result for  $ff$ :  $4 \times 10^{-4} \pm 1 \times 10^{-4}$  and scale with  $\ln^2(\bar{t}_{ex}/m_{ff}^2) / \ln^2(\bar{t}_{LEP}/m_{ff}^2)$  (pairs)
  - use LEP result for  $f\bar{f}$  terms:  $20\% \times 4 \times 10^{-4}$  (G. Montagna, M. Moretti, O. Nicrosini, A. Pallavicini, and F. Piccinini, Nucl. Phys. B547 (1999) 39-59), and scale with  $\ln(\bar{t}_{ex}/m_f^2) / \ln(\bar{t}_{LEP}/m_f^2)$  (photons)
  - $\tau$ -pair (negligible at LEP) estimated relative to muon-pair as  $\ln^2(\bar{t}_{ex}/m_\tau^2) / \ln^2(\bar{t}_{ex}/m_\mu^2)$
  - hadron-pair estimated relative to muon-pair as  $R_{had} \times \ln^2(\bar{t}_{ex}/(0.5\text{GeV})^2) / \ln^2(\bar{t}_{ex}/m_\mu^2)$

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## Lumi at FCCee - Forecast

Type of correction / Error	FCCee <sub>240</sub> [1]	FCCee <sub>350</sub> [2]	FCCee <sub>500</sub> [2]
(a) Photonic $O(L_s \alpha^2)$	$0.10 \times 10^{-4}$	$0.10 \times 10^{-4}$	$0.13 \times 10^{-4}$
(b) Photonic $O(L_s^2 \alpha^2)$	$0.06 \times 10^{-4}$	$0.26 \times 10^{-4}$	$0.27 \times 10^{-4}$
(c) Vacuum polariz.	$0.6 \times 10^{-4}$	$1.0 \times 10^{-4}$	$1.1 \times 10^{-4}$
(d) Light pairs	$0.5 \times 10^{-4}$	$0.4 \times 10^{-4}$	$0.4 \times 10^{-4}$
(e) Z and s-channel $\gamma$ exch.	$0.1 \times 10^{-4}$	$1.0 \times 10^{-4}$	$1.0 \times 10^{-4}$
(f) Up-down interference	$0.1 \times 10^{-4}$	$0.09 \times 10^{-4}$	$0.1 \times 10^{-4}$
Total	$1.0 \times 10^{-4}$	$1.5 \times 10^{-4}$	$1.6 \times 10^{-4}$

- Numbers: (\*) likely overestimated, (a) include safety factor 2. Technical error is not included  
Precision dominated by:  
- Vacuum polarisation (c) - seems irreducible  
- The EW  $O(\alpha^2)$  uncertainty (e): Numbers (\*) are likely overestimated (taken from 800 GeV estimate) - factor 2 too big?  
Number (-) possibly underestimated ( $0.3 \times 10^{-4}$  ?)  
Precision loss at higher energies reasonable  
factor of 2 loss w.r.t.  $M_Z$

[1] S. Jadach, W. Placzek, M. Skrzypek, B. F. L. Ward, S. A. Yost, Phys. Lett. B 790 (2019) 314  
[2] S. Jadach, W. Placzek, M. Skrzypek, B. F. L. Ward, Eur. Phys. J. C (2021) 81:1047

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## Lumi at FCCee<sub>MZ</sub> - Forecast study

Type of correction / Error	Forecast [1]	Redone
(a) Photonic $O(L_s \alpha^2)$	$0.10 \times 10^{-4}$	$0.10 \times 10^{-4}$
(b) Photonic $O(L_s^2 \alpha^2)$	$0.06 \times 10^{-4}$	$0.06 \times 10^{-4}$
(c) Vacuum polariz.	$0.6 \times 10^{-4}$	$0.6 \times 10^{-4}$
(d) Light pairs	$0.5 \times 10^{-4}$	$0.27 \times 10^{-4}$
(e) Z and s-channel $\gamma$ exch.	$0.1 \times 10^{-4}$	$0.1 \times 10^{-4}$
(f) Up-down interference	$0.1 \times 10^{-4}$	$0.08 \times 10^{-4}$
Total	$1.0 \times 10^{-4}$	$0.70 \times 10^{-4}$

- (d) light pairs are re-analysed w.r.t. [1] (safety factor 1.25 is removed;  $ff$ -non-leading contrib. less conservative;  $Z_{cut} \leq .5$  can help; hadr-pair uncertainty is set to few % as in [2])
  - (f) value not rounded up is used as compared to Ref. [1]
  - "Total" value not rounded up is used as compared to Ref. [1] (the above three entries corrected at 240 and 350 GeV as well)
  - (b) missing non-logarithmic  $O(\alpha^2 L^2)$  correction added for completeness
  - (e): size of  $O(\alpha^2)_{EW}$  corrs. to be revisited - available BHWIDE (conservative scaling  $0.3 \times 10^{-4}$ ) & DIZET (switches, at higher energy) CEEX amplitude level exponentiation instrumental (KKMC style) ?
- [1] S. Jadach, W. Placzek, M. Skrzypek, B. F. L. Ward, S. A. Yost, Phys. Lett. B 790 (2019) 314  
[2] ALEPH Collaboration, D. Buskulic et al., Z. Phys. C 66 (1995) 3-18
- Possible precision  $\sim 0.7 \times 10^{-4}$  within the reach ??

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## Lumi via gamma-gamma

- Another option for Lumi measurement:  $e^+e^- \rightarrow \gamma\gamma$  [1],[2]
- (KLOE, CLEO, BESIII)
- statistical error  $\sim 2 \times 10^{-5}$
  - not sensitive to hadronic corrections (vac. pol. contributes at NNLO)  $\leq 10^{-5}$
  - not sensitive to detector geometry (measurement at wide angles), low angle Bhabha needs  $\mu$ -meter alignment
  - huge background due to large angle Bhabhas ( $\sim 100$  times bigger than signal)
  - channel sensitive to new physics, must be well controlled
  - precision below  $10^{-4}$  requires full  $O(\alpha^2)$  QED and EW corrections

[1] Electroweak corrections to  $e^+e^- \rightarrow \gamma\gamma$  as a luminosity process at FCC-ee Carlo M. Carloni Calame, Mauro Chiesa, Guido Montagna, Oreste Nicrosini, Fulvio Piccinini, arXiv:1906.08056  
[2] Precision studies of quantum electrodynamics at future  $e^+e^-$  colliders J. Alcaraz Maestre, arXiv:2206.07564

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## Bhabha lumi: $0.1 \times 10^{-4}$ at Z-peak?

### Vacuum polarisation

Note: Lattice methods with Jegerlehner's results allow, in principle, (c)  $\rightarrow$  (c)/6

$$\Delta\alpha_{had}(\bar{t}) = \Delta\alpha_{had}(-Q_0^2)_{lat} + [\Delta\alpha_{had}(\bar{t}) - \Delta\alpha_{had}(-Q_0^2)]_{pQCD}^{Adler}$$

Lattice results are mainly limited now by statistics (?), so if enough computing resources are available, the  $0.1 \times 10^{-4}$  precision at  $\sim 1$  GeV<sup>2</sup> may be feasible.

The above is more optimistic than the 3.5 $\sigma$  tension with estimates based on exp. data of R-ratio reported in arXiv: 2203.08676, 2211.11401 [hep-lat] for  $\Delta\alpha_{had}^{(5)}(-Q^2)$ ,  $Q^2 = 3 \div 7$  GeV<sup>2</sup>.

The precision of lattice results given in the above papers is  $\Delta\alpha_{had}(-5\text{GeV}^2) = 0.00716 \pm 0.9 \times 10^{-4}$  - on par with R-ratio method.

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## Bhabha lumi: $0.1 \times 10^{-4}$ at Z-peak?

### EW corrections

In S. Jadach, W. Placzek, M. Skrzypek, B. F. L. Ward, and S. A. Yost, Phys. Lett. B 790 (2019) 314-321 we estimated

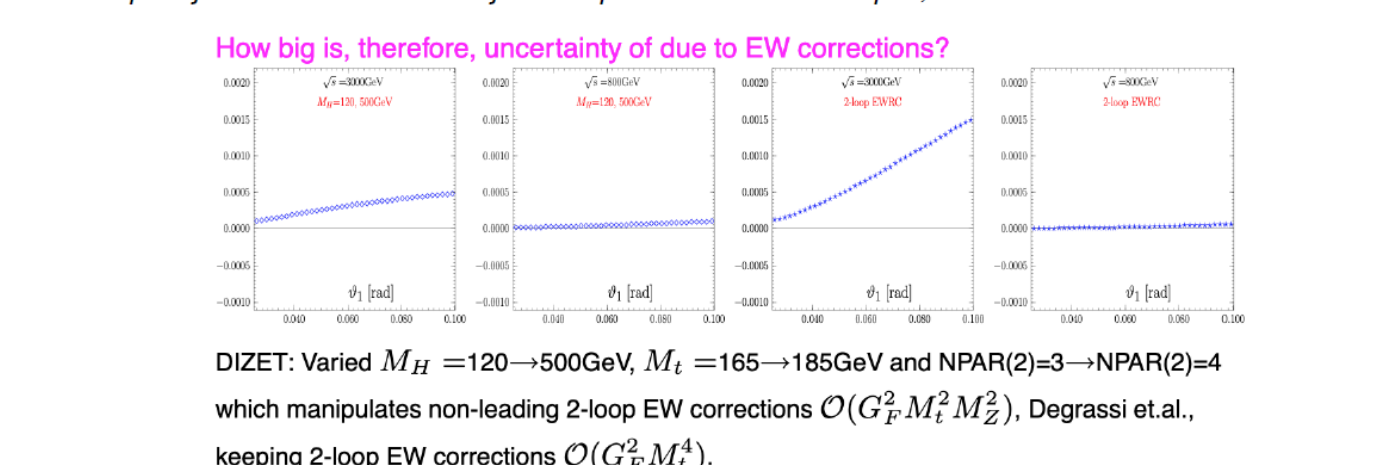
- $O(\alpha_{EW}^2)$  uncertainties in BHWIDE at Z-peak:
  - Conservatively estimated as  $\frac{1}{2} \ln \frac{1}{\alpha_{EW}^2} \times O(\alpha_{EW}^2)$  times safety factor of 2.
  - This gives  $0.7 \times 10^{-4}$  for QED part and  $0.3 \times 10^{-4}$  for EW part. Added linearly one obtains  $1 \times 10^{-4}$ .
  - But we are interested only in EW part!
- More aggressive estimate (no safety factor, added in quadratures) would give  $0.4 \times 10^{-4}$  for total and  $0.15 \times 10^{-4}$  for EW part.

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## Bhabha lumi: $0.1 \times 10^{-4}$ at Z-peak?

DIZET analysis of EW corrs. done above Z-peak. At the peak different graphs contribute ( $\gamma_t \otimes Z_t$  vs  $\gamma_t \otimes Z_t$ ), but rough idea could be valid? M. Battaglia, S. Jadach, and D. Bardin, eConf C010630 (2001) E3015  
S. Jadach, "MC tools for extracting luminosity spectra. What do we need?", <https://jadach.web.cern.ch/jadach/public/LumiL.Cslac.pdf>, 2002



At 800 GeV  $O(\alpha_{EW}^2)$  contributes below  $0.4 \times 10^{-4}$  and decreases with energy decrease ( $1.5 \times 10^{-4}$  at 3 TeV).  
Bottom line: leading  $\alpha_{EW}^2$  contribs may be needed

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## Bhabha lumi: $0.1 \times 10^{-4}$ at Z-peak?

### Fermion pairs

- One will probably need  $O(\alpha)$  corrections to four fermion final state.
- Calculations of Denner et al. (PLB 612(2005) 223) exist for charged current final states. Claimed physical precision (due to higher orders) at WW threshold is few  $\cdot 0.1\%$  of the 4f Born.
- The whole pair contribution to Bhabha is  $\sim 4 \times 10^{-4}$ . Assuming precision of 1% for NC final states we are well below  $0.1 \times 10^{-4}$  target, provided f-channel multiphotons are properly resummed. Note, that above  $\sim 500$  GeV Sudakov logs must be resummed.

Bottom line  
 $0.1 \times 10^{-4}$  precision a priori not excluded

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## Lumi forecast at ILC and CLIC

Type of correction / Error	ILC <sub>500</sub>	ILC <sub>1000</sub>	CLIC <sub>3000</sub>
(a) Photonic $O(L_s \alpha^2)$	$0.13 \times 10^{-4}$	$0.15 \times 10^{-4}$	$0.20 \times 10^{-4}$
(b) Photonic $O(L_s^2 \alpha^2)$	$0.27 \times 10^{-4}$	$0.37 \times 10^{-4}$	$0.63 \times 10^{-4}$
(c) Vacuum polariz.	$1.1 \times 10^{-4}$	$1.1 \times 10^{-4}$	$1.2 \times 10^{-4}$
(d) Light pairs	$0.4 \times 10^{-4}$	$0.5 \times 10^{-4}$	$0.7 \times 10^{-4}$
(e) Z and s-channel $\gamma$ exch.	$1.0 \times 10^{-4}$	$2.4 \times 10^{-4}$	$1.6 \times 10^{-4}$
(f) Up-down interference	$0.1 \times 10^{-4}$	$0.1 \times 10^{-4}$	$0.1 \times 10^{-4}$
Total	$1.6 \times 10^{-4}$	$2.7 \times 10^{-4}$	$1.6 \times 10^{-4}$

- Number (\*) is somewhat overestimated (taken from 800 GeV estimate)
- Precision at high energies totally due to the EW  $O(\alpha^2)$  hard process uncertainty (e).
  - EW interferences are dominated by  $\gamma_t \otimes Z_t$  (15% of  $\gamma_t \otimes \gamma_t$  at CLIC) and  $Z_t \otimes Z_t$  (2% of  $\gamma_t \otimes \gamma_t$  at CLIC)
  - useful for  $O(\alpha_{EW}^2)$  calculation ?
  - CEEX amplitude level exponentiation mandatory ?

At 3 TeV loss of precision is dramatic, dominant  $O(\alpha_{EW}^2)$  and CEEX are a must!

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## Lumi forecast at CEPC

Type of correction / Error	CEPC <sub>240</sub>	CEPC <sub>240</sub>
(a) Photonic $O(L_s \alpha$		