



Searches for BSM physics using long-lived signatures with the CMS detector

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What are LLPs?

- Particles that have relatively long lifetime (> O(ns)), that travel significant distances before decaying (0.01 1 m)
- Naturally arise in many BSM scenarios like SUSY, dark matter models and hidden sector theories
- Main signature: displaced origin
 - may have other signature depending on the analysis



distance travelled

From H. Russell's Slides



Why are they long-lived?

- \checkmark Small couplings to SM \rightarrow decay channels are suppressed
- \checkmark Kinematics \rightarrow small phase-space volume, suppressing decay
- $\checkmark\,$ Decays via Heavy Mediator \rightarrow the decay width is suppressed by a factor of $(m/M)^p$

Trigger and Dataset



LLP Searches at CMS



From H. Russell's Slides



Neutrino oscillations \Rightarrow Neutrinos are massive

Not explained by SM

Possible solution: Extend SM to vMSM

 \succ Add right-handed sterile neutrinos v_R

 $HNL \sim (\nu_R + \nu_R^C)$

> Gets mass by Seesaw mechanism

$$\mathcal{L}_{ ext{mass}} = -rac{1}{2} egin{pmatrix} ar{
u}_L & ar{
u}_R^c \end{pmatrix} egin{pmatrix} 0 & m_D \ m_D & M \end{pmatrix} egin{pmatrix}
u_L^c \
u_R \end{pmatrix} + ext{h.c}$$

Outcomes:

- > Mass eigenstates N_1 , N_2 , N_3 (ordered in mass)
- \succ Lightest neutrino N₁ serves as DM candidate

 \succ N₂ and N₃ are LLPs



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EXO-22-019

SM

 $m_{\nu} =$

 $\nu \sim (\nu_L + \nu_L^C)$

$\ell_{P} = \mu \text{ or } e^{\pm}$ $\ell_{P} = \mu \text{ or } e^{\pm}$ $\psi_{\ell_{P}}$ $\psi_{\ell_{P}}$

- > Analysis performed using B-parking dataset
 - \checkmark Triggered by muon

В

- \checkmark 3 flavor channels explored in the analysis: $\mu\mu, \mu e, e\mu$
- > Strategy:
 - Reconstruct displaced vertex using l^{\pm} and π^{\pm} tracks
 - Bump hunt in $(l^{\pm}\pi^{\pm})$ mass spectrum
- > Event selection using parametric neural network (list of variables)
- > Background for training obtained by taking 1/1000 of events available in data



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- > Signal rate = F(m_N, V_N) (link) $|V_N|^2 \equiv |V_{eN}|^2 + |V_{\mu N}|^2 + |V_{\tau N}|^2$ $r_{\ell} \equiv |V_{\ell N}|^2 / |V_N|^2 \quad \ell = (e, \mu, \tau)$
- \succ The limits are calculated in terms of the mass $m_{\rm N}$ and mixing $V_{\rm N}$
- \succ The best limit is obtained for muon exclusive mixing (0, 1, 0) at 95% CL

| $(r_{ m e},r_{\mu},r_{	au})$ | Scenario | $ V_{\rm N} ^2$ | Mass~(GeV) |
|------------------------------|------------|----------------------|------------|
| (0,1,0) | Majorana | 2.0×10^{-5} | 1.95 |
| (0, 1/2, 1/2) | Majorana | 4.0×10^{-5} | 1.42 |
| (1/2, 1/2, 0) | Majorana | 3.3×10^{-5} | 2.15 |
| (1/3, 1/3, 1/3) | Majorana | 5.0×10^{-5} | 2.15 |
| (0,1,0) | Dirac-like | 3.2×10^{-5} | 1.68 |
| (0, 1/2, 1/2) | Dirac-like | 6.5×10^{-5} | 1.68 |
| (1/2, 1/2, 0) | Dirac-like | 5.7×10^{-5} | 1.68 |
| (1/3, 1/3, 1/3) | Dirac-like | 8.5×10^{-5} | 1.68 |

EXO-22-019

Heavy Stable Charged Particle Search

- > Model independent search for HSCPs
- \blacktriangleright Main signature \rightarrow Isolated high p_{T} track with Large energy loss per unit length (dE/dx)

> m_{HSCP} > 100 GeV ⇒ β = v/c significantly smaller than 1 ✓ Distinguishable from ultra-relativistic SM particles

- \succ Search performed using Run 2 data of 2017 and 2018 \rightarrow Total luminosity $101~{
 m fb^{-1}}$
- > Only barrel region of the tracker is used
 - 4 pixel detector layer (after upgrade in 2017) and 10 layers of strip detectors
 - Coverage: $|\eta| < 1.4$
 - Track resolution:
 - $p_T \! \rightarrow \! 2.8\%$ at 100 GeV
 - position \rightarrow 10(30) μm in transverse (longitudinal) IP





EXO-18-002





> For Pixel detector $F_{i}^{\text{Pixels}} = 1 - \prod_{j=1}^{n} P_{j}' \sum_{k=0}^{n-1} \frac{\left[-\ln(\prod_{j=1}^{n} P_{j}')\right]^{k}}{k!}$ $P_i \rightarrow \text{MIP}$ hit probability $n \rightarrow \text{number of pixel hits (excluding layer 1)}$ For Strip detector $G_{i}^{\text{Strips}} = \frac{3}{N} \left(\frac{1}{12N} + \sum_{i=1}^{N} \left| P_{j} \left(P_{j} - \frac{2j-1}{2N} \right)^{2} \right| \right)$ $P'_i \rightarrow MIP$ hit probability to produce charge $N \rightarrow$ number of hits equal or less than *i*th measured charge 5,0.9 CMSSimulation 5 0.9 CMS Simulation

0.6 0.5 0.5 0.4^E 0.4^E 0.3Ē 0.3E 10^{-2} 0.2 0.2E 0203040506070809 □Pixels -Pixels

Mass Method

> Ionization Estimator

$$I_{\rm h} = \left(\frac{1}{N}\sum_{j}^{N} (\mathrm{d}E/\mathrm{d}x_j)^{-2}\right)^{-1/2}$$

- ► Mass of HSCP $I_{\rm h} = K \frac{m^2}{p^2} + C$ $m \to \text{mass}$ $p \to \text{momentum}$
 - Constants K and C are determined from a sample of lowmomentum particles (link)
 - Equation is inverted to calculate the mass of the candidate particle from dE/dx measurement



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Heavy Stable Charged Particle Search

Results

 Particle Search
 EXO-18-002

 101 fb⁻¹ (13 TeV)
 101 fb⁻¹ (13 TeV)

 ber Limits
 95% CL Upper Limits

 it for g (inas)
 05. limit for pair prod. ₹ (mass)

 it for t (mass)
 05. limit for gMSB SPS7 ₹ (mass)



- Results interpreted based on many signal production scenarios
- Limits calculated for production cross-section vs. mass of HSCP at 95% CL









- Search for LLPs decaying to hadronic final state
- > Focusing on low mass LLP :

$$m_{LLP} = 10 - 60 \text{ GeV}$$

 $c\tau = 1 - 1000 \text{ mm}$



- \blacktriangleright Main signature \rightarrow displaced-jet vertex (DV) and tracks
- \blacktriangleright Benchmark Model \rightarrow Higgs portal to hidden sector
- Search performed using Run 3 data of 2022 (13.6 TeV)
 - \rightarrow Total luminosity 34.7 fb^-1
- > Dedicated displaced-jet triggers used to collect data



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EXO-23-013

LLP Search with Displaced Jets



The tagger scores are used as variables for the ABCD method to estimate the background in SR.



Results

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EX0-21-008





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р

22.8° producing shower in 20.7° the muon system 15.4° 14.0°

Search for LLPs



- > Main signature \rightarrow Cluster of hits in the muon system in the direction of missing transverse momentum (MET or $\vec{p}_{\rm T}^{\rm miss}$)
- \succ Benchmark Model \rightarrow Twin Higgs and Dark Shower Model
- Search performed using full Run 2 data (13.6 TeV)
 - \rightarrow Total luminosity $138~fb^{-1}$
- Events selected based on MET
 - ✓ MET > 120 GeV HLT
 - \checkmark MET $> 200~{\rm GeV}$ in offline selection





3 mutually exclusive search categories:

- Double Clusters: CSC-CSC, DT-DT and CSC-DT
- 1 DT Cluster
- 1 CSC Cluster



> Two discriminating variables defined for the analysis

- * $N_{hits} \rightarrow Number$ of hits in the cluster
- $\Delta \phi(\vec{\rho}_{\tau}^{\text{miss}}, \text{ cluster}) \rightarrow \text{Angular separation between MET and the cluster}$
- These are correlated for signal and uncorrelated for background. So, ABCD method employed to estimate background in SR

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Results



 \succ Results interpreted with signal models with varying mass of S

 \blacktriangleright Limits calculated for BR of H \rightarrow SS vs. $c\tau$ of S at 95% CL



 \blacktriangleright Results interpreted with signal models with varying mass of S

 \blacktriangleright Limits calculated for BR of H \rightarrow SS vs. $c\tau$ of S at 95% CL

Future Outlook

- ✤ New ML triggers
 - ★ Auto-encoder based AXOL1TL trigger for selecting anomalous events
 - CNN based calorimeter trigger CICADA
- * Inclusion of tracker information at L1 trigger \rightarrow improved selection efficiency for events with displaced vertices
- Proposed detector upgrades at CMS like Massive Timing Hodoscope for Ultra Stable neutral pArticles (MATHUSLA), a surface based dedicated detector for neutral LLPs

More LLP searches coming-up, STAY TUNED!

Thank you for your attention!



ABCD Method



CMS Detector





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List of Input Variables for pNN

- 1. Transverse momenta: $p_{\rm T}(\ell_{\rm B}), p_{\rm T}(\ell^{\pm}), p_{\rm T}(\pi^{\mp}).$
- 2. Invariant-masses: $m(\ell_{\rm B}\pi^{\mp}), m(\ell_{\rm B}\ell^{\pm}), m(\ell_{\rm B}\ell^{\pm}\pi^{\mp}).$
- 3. Track separation in the η - φ space (where φ is the azimuthal angle) $\Delta R \equiv \sqrt{(\Delta \eta)^2 + (\Delta \varphi)^2} \approx \Delta R(\ell_{\rm B}, \ell^{\pm}), \Delta R(\ell_{\rm B}, \pi^{\mp})$
- 4. Displaced vertex properties: $\cos \theta$, fit *p*-value.
- 5. Displacement-related quantities: $L_{xy}/\sigma_{L_{xy}}$ and $d_{xy}/\sigma_{d_{xy}}$ of the pion.
- 6. Track-related information: number of layers of the CMS silicon pixel and strip tracker traversed by the lepton(s) and pion from the DV.
- 7. Lepton isolation, defined in a cone of ΔR smaller than 0.3 around the lepton momentum vector

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pNN Validation



Invariant Mass distribution



Estimation of Signal Events

$$N_{\rm sig}\left(\ell_{\rm B}\ell, m_{\rm N}, c\tau_{\rm N}, \vec{r}\right) = \frac{\sigma_{\rm B}^{\rm eff}}{f_u} \mathcal{L}\sum_{\rm q} F_{\rm N}^{\rm q}\left(\ell_{\rm B}\ell, m_{\rm N}, c\tau_{\rm N}, \vec{r}\right)\epsilon_{\rm sig}^{\rm q}\left(\ell_{\rm B}\ell, m_{\rm N}, c\tau_{\rm N}\right)$$



$$\frac{1}{c\tau_{\rm N}} = |V_{\rm N}|^2 \left(r_{\rm e} \widetilde{\Gamma}_{\rm e} \left({\rm N} \right) + r_{\mu} \widetilde{\Gamma}_{\mu} \left({\rm N} \right) + r_{\tau} \widetilde{\Gamma}_{\tau} \left({\rm N} \right) \right)$$

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Control Region and Signal Region for Mass Method

control region A defined as $G_{i}^{\text{Strips}} < 0.018$ and $55 < p_{T} < 70 \text{ GeV}$ control region B defined as $G_{i}^{\text{Strips}} > 0.22$ and $55 < p_{T} < 70 \text{ GeV}$ control region C defined as $G_{i}^{\text{Strips}} < 0.018$ and $p_{T} > 70 \text{ GeV}$ signal region D is defined as $G_{i}^{\text{Strips}} > 0.22$ and $p_{T} > 70 \text{ GeV}$.

$$N_{bkg}(D) = \frac{N(B)N(C)}{N(A)}$$

Determining constants K and C

| Parameter | Da | ata | Simulation | | |
|-----------|-----------------|---------------|---------------|---------------|--|
| (MeV/cm) | 2017 | 2018 | 2017 | 2018 | |
| K | 2.54 ± 0.01 | 2.55 ± 0.01 | 2.50 ± 0.01 | 2.49 ± 0.01 | |
| С | 3.14 ± 0.01 | 3.14 ± 0.01 | 3.18 ± 0.01 | 3.18 ± 0.01 | |

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Mass Spectrum



Pass and Fail Region for Ionization Method



Exclusion Limits

| Model | Ionizatio | n method | Mass method | | |
|--------------------------------------|-----------------|------------|---------------|------------|--|
| | Exp. (TeV) | Obs. (TeV) | Exp. (TeV) | Obs. (TeV) | |
| ĝ | 2.06 ± 0.06 | 2.06 | 2.08 ± 0.02 | 2.08 | |
| ĩ | 1.43 ± 0.05 | 1.40 | 1.47 ± 0.02 | 1.47 | |
| GMSB SPS7 $\tilde{\tau}$ | 0.86 ± 0.07 | 0.85 | 0.87 ± 0.05 | 0.85 | |
| pair-prod. ${\widetilde 	au}_{ m R}$ | 0.53 ± 0.03 | 0.52 | 0.50 ± 0.07 | 0.51 | |
| pair-prod. $\tilde{\tau}_{\rm L}$ | 0.66 ± 0.04 | 0.64 | 0.67 ± 0.06 | 0.61 | |
| pair-prod. $\tilde{\tau}_{L/R}$ | 0.71 ± 0.04 | 0.69 | 0.75 ± 0.08 | 0.64 | |
| τ' ($Q = 1e$) from DY prod. | 1.05 ± 0.05 | 1.02 | 1.14 ± 0.03 | 1.14 | |
| τ' ($Q = 2e$) from DY prod. | 1.35 ± 0.05 | 1.32 | 1.41 ± 0.02 | 1.41 | |
| $Z'_\psi 	o 	au' 	au'$ | 3.99 ± 0.21 | 3.95 | 4.03 ± 0.01 | 4.03 | |
| $Z'_{ m SSM} ightarrow 	au' 	au'$ | 4.53 ± 0.23 | 4.38 | 4.56 ± 0.01 | 4.57 | |

Control Region and Signal Region

- Region A: events with $0.95 < g_{\text{displaced}} < 0.9985$, $0.95 < g_{\text{prompt-veto}} < 0.985$;
- Region B: events with $0.95 < g_{\text{displaced}} < 0.9985, 0.985 < g_{\text{prompt-veto}} < 1.0;$
- Region C: events with $0.9985 < g_{displaced} < 1.0, 0.95 < g_{prompt-veto} < 0.985$; and
- Region D, the signal region (SR): events with $0.9985 < g_{\text{displaced}} < 1.0, 0.985 < g_{\text{prompt-veto}} < 1.0.$

| Sdisplaced | Predicted background | Observation | Z-value | |
|-----------------|----------------------|-------------|---------|-------------------------------|
| (0.96, 0.97) | 68.39 ± 12.60 | 52 | -1.06 | |
| (0.97, 0.98) | 67.55 ± 9.46 | 77 | 0.80 | |
| (0.98, 0.99) | 76.18 ± 8.95 | 72 | -0.27 | |
| (0.99, 0.995) | 38.82 ± 5.08 | 45 | 0.84 | $g_{\rm prompt-veto} > 0.985$ |
| (0.995, 0.998) | 25.41 ± 3.87 | 26 | 0.22 | |
| (0.998, 0.9985) | 2.83 ± 1.17 | 5 | 1.25 | |
| (0.9985, 1.0) | 3.34 ± 1.28 | 3 | 0.19 | |

Exclusion Limit Plots



Exclusion Limit Plots





3 mutually exclusive search categories:

- Double Clusters: CSC-CSC, DT-DT and CSC-DT
- 1 DT Cluster
- 1 CSC Cluster

Control Region and Signal Region

Double Clusters

Bin A includes events with the CSC cluster with $N_{\rm hits}>100$ and the DT cluster with $N_{\rm hits}>80;$

Bin B includes events with the CSC cluster with $N_{\rm hits} > 100$ and the DT cluster with $N_{\rm hits} \leq 80$;

Bin C includes events with the CSC cluster with $N_{\rm hits} \leq 100$ and the DT cluster with $N_{\rm hits} \leq 80;$

Bin D includes events with the CSC cluster with $N_{\rm hits} \leq 100$ and the DT cluster with $N_{\rm hits} > 80.$

| Category | Validation region | $N_{\rm B}$ | N _C | $N_{\rm D}$ | $N_{\rm BD}$ | $\lambda_{\mathbf{A}}$ | $N_{\rm A}$ |
|----------|--|-------------|----------------|-------------|--------------|------------------------------|-------------|
| DT-DT | Inverted $\Delta \phi(\vec{p}_{\mathrm{T}}^{\mathrm{miss}}, \mathrm{cluster})$ Inverted N_{hits} | | 11 2 | | 1 1 | $0.02 + 0.05 \\ 0.12 + 0.27$ | 0 0 |
| CSC-CSC | Inverted $\Delta \phi(\vec{p}_{\mathrm{T}}^{\mathrm{miss}}, \mathrm{cluster})$ Inverted N_{hits} | | 8 4 | | 2 2 | $0.12 + 0.18 \\ 0.25 + 0.38$ | 0 0 |
| DT-CSC | Inverted $\Delta \phi(\vec{p}_{\mathrm{T}}^{\mathrm{miss}}, \mathrm{cluster})$ Inverted N_{hits} | 0 2 | 19 11 | 3 1 | | $0+0.3 \\ 0.18+0.23$ | 0 0 |

Single Clusters

Bin A includes events with $\Delta \phi(\vec{p}_{T}^{\text{miss}}, \text{cluster}) < 0.75 \text{ and } N_{\text{hits}} > 130;$ bin B includes events with $\Delta \phi(\vec{p}_{T}^{\text{miss}}, \text{cluster}) \ge 0.75$ and $N_{\text{hits}} > 130;$ bin C includes events with $\Delta \phi(\vec{p}_{T}^{\text{miss}}, \text{cluster}) \ge 0.75$ and $N_{\text{hits}} \le 130;$ bin D includes events with $\Delta \phi(\vec{p}_{T}^{\text{miss}}, \text{cluster}) < 0.75$ and $N_{\text{hits}} \le 130$.

| Validation region | $N_{\rm B}$ | $N_{\rm C}$ | N_{D} | $\lambda_{ m A}$ | $N_{\rm A}$ |
|--------------------|-------------|-------------|------------------|------------------|-------------|
| Out-of-time region | 8 | 442 | 121 | 2.2 + 0.8 | 3 |
| In-time region | 8 | 317 | 87 | 2.2 + 0.8 | 2 |

Double Cluster Results





Single Cluster Results

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