

# Inflationary Gravitational waves as test for low scale leptogenesis

Friday 13 December 2024 15:50 (17 minutes)

We study thermal and non-thermal resonant leptogenesis in a general setting where a heavy scalar  $\phi$  decays to right-handed neutrinos (RHNs) whose further out-of-equilibrium decay generates the required lepton asymmetry. Domination of the energy budget of the Universe by the  $\phi$  or the RHNs alters the evolution history of the primordial gravitational waves (PGW) of inflationary origin, which re-enter the horizon after inflation, modifying the spectral shape. The decays of  $\phi$  and RHNs release entropy into the early Universe while nearly degenerate RHNs facilitate low and intermediate-scale leptogenesis. We show that depending on the coupling  $y_R$  of  $\phi$  to radiation species, RHNs can achieve thermal abundance before decaying, which gives rise to thermal leptogenesis. A characteristic damping of the GW spectrum resulting in knee-like features would provide evidence for low-scale thermal and non-thermal leptogenesis. We explore the parameter space for the lightest right-handed neutrino mass  $M_1 \in [10^2, 10^{14}]$  GeV and washout parameter  $K$  that depends on the light-heavy neutrino Yukawa couplings  $\lambda$ , in the weak ( $K < 1$ ) and strong ( $K > 1$ ) washout regimes. The resulting novel features compatible with observed baryon asymmetry are detectable by future experiments like LISA and ET. By estimating signal-to-noise ratio (SNR) for upcoming GW experiments, we investigate the effect of the scalar mass  $M_\phi$  and reheating temperature  $T_\phi$ , which depends on the  $\phi - N$  Yukawa couplings  $y_N$ .

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**Session Classification:** Parallel Session