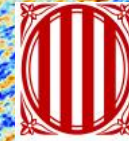


# Cosmic Birefringence: New Physics Explanations



Silvia Gasparotto (IFAE)

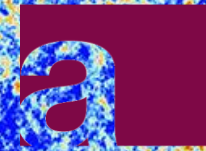
UNDARK Kick-Off 9/10/2024



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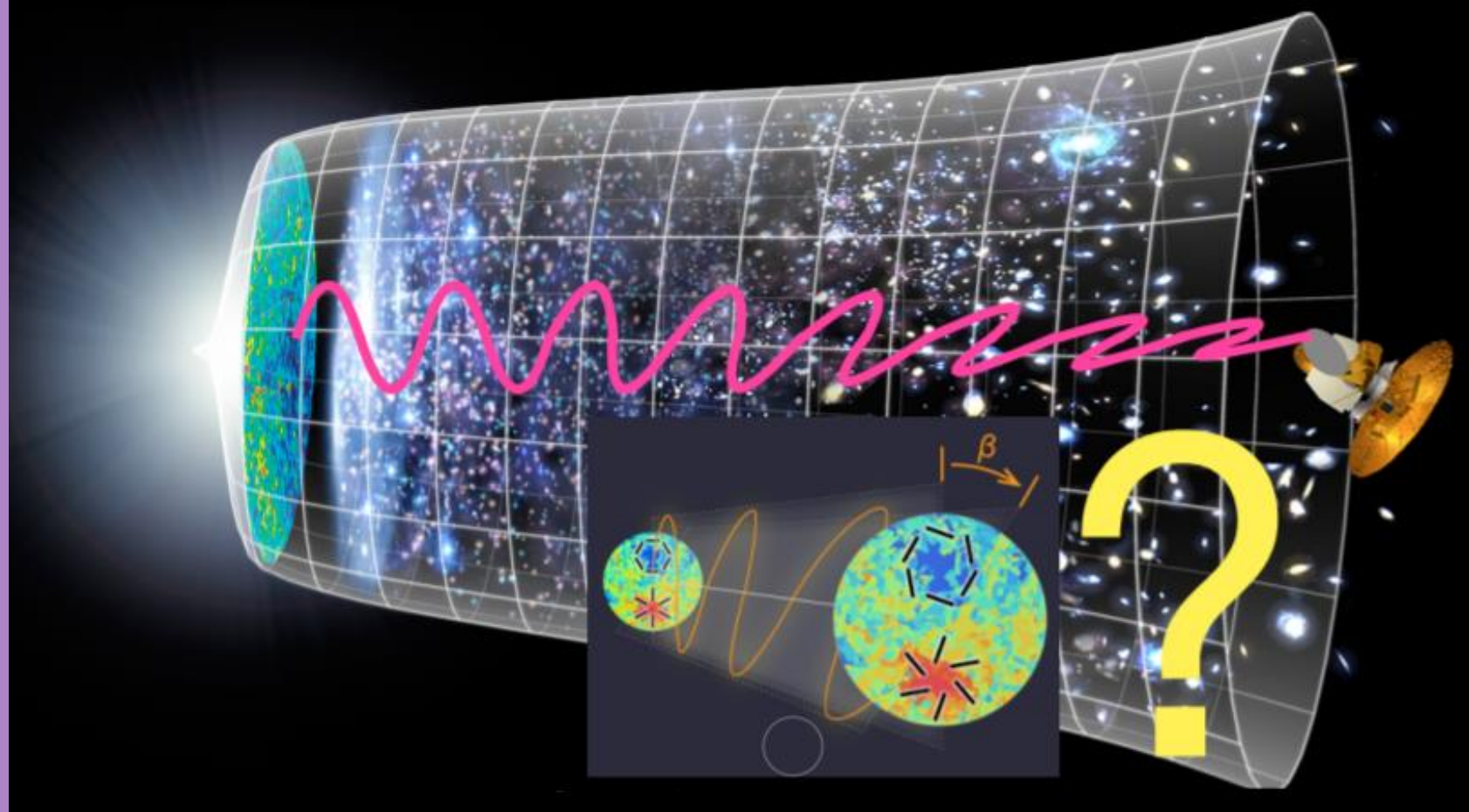


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Funded by  
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# Outlook



- Introduction: Axion Birefringence
- Signal from the Cosmic Microwave Background
- Possible explanations: different models
- Conclusion and future directions

# Axion Birefringence

Turner & Widrow (1988)

the effective Lagrangian for axion electrodynamics is

$$\mathcal{L} = -\frac{1}{2}\partial_\mu\theta\partial^\mu\theta - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \underbrace{g_a\theta F_{\mu\nu}\tilde{F}^{\mu\nu}}_{\text{Chern-Simons term}}, \quad (3.7)$$

$\tilde{F}^{\mu\nu} = \sum_{\alpha\beta} \frac{\epsilon^{\mu\nu\alpha\beta}}{2\sqrt{-g}} F_{\alpha\beta}$

where  $g_a$  is a coupling constant of the order  $\alpha$ , and the vacuum angle  $\theta = \phi_a / f_a$  ( $\phi_a =$  axion field). The equations

$$g_{a\gamma}\phi\vec{E}\cdot\vec{B} \quad \text{Parity odd}$$

Axion-photon interaction modifies the photon dispersion relation in a parity-violating way

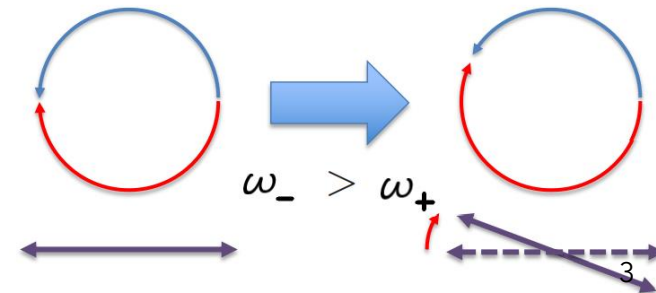
$$A''_{\pm}(\eta, k) + \underbrace{k^2}_{\omega_{\pm}^2} \left( 1 \mp \frac{g_{a\gamma}\phi'}{k} \right) A_{\pm}(\eta, k) = 0$$

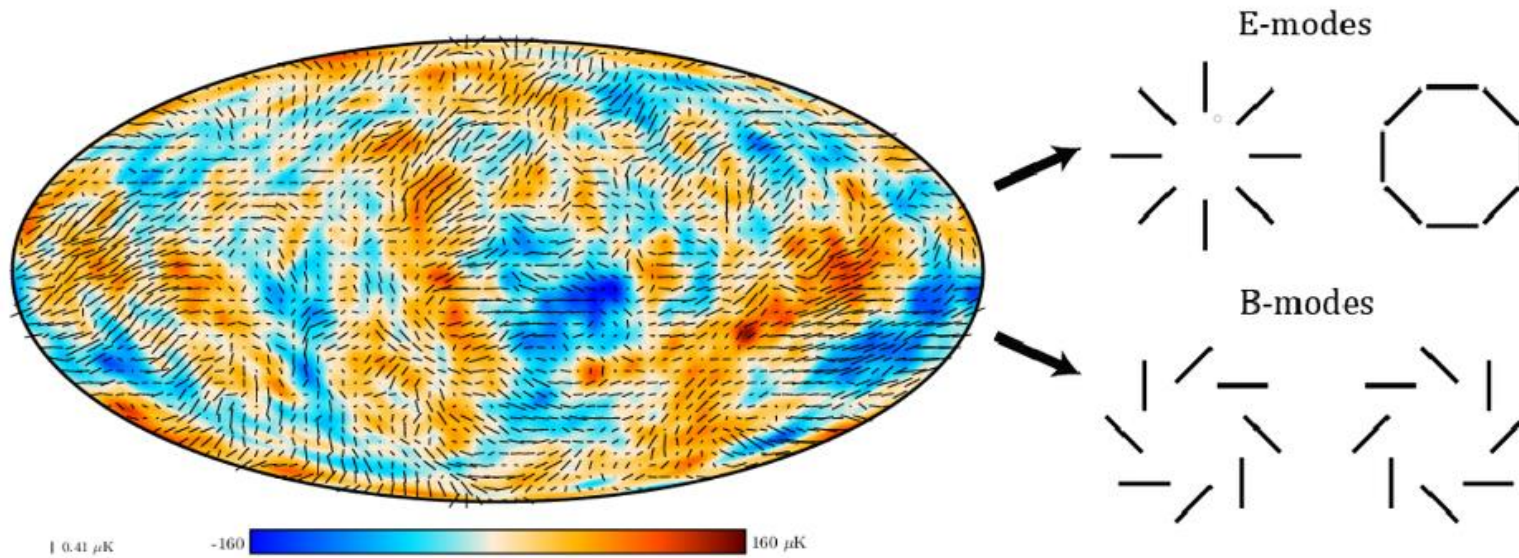
Left- and right handed photons propagate with a different speed

$$\omega_{\pm} \simeq k \mp \frac{g_{a\gamma}}{2} \phi'$$

- i. Frequency independent
- ii. Evolving field  $\phi' \neq 0$

Rotation of the photon linear polarization





CMB photons emitted 13.8 billion years ago

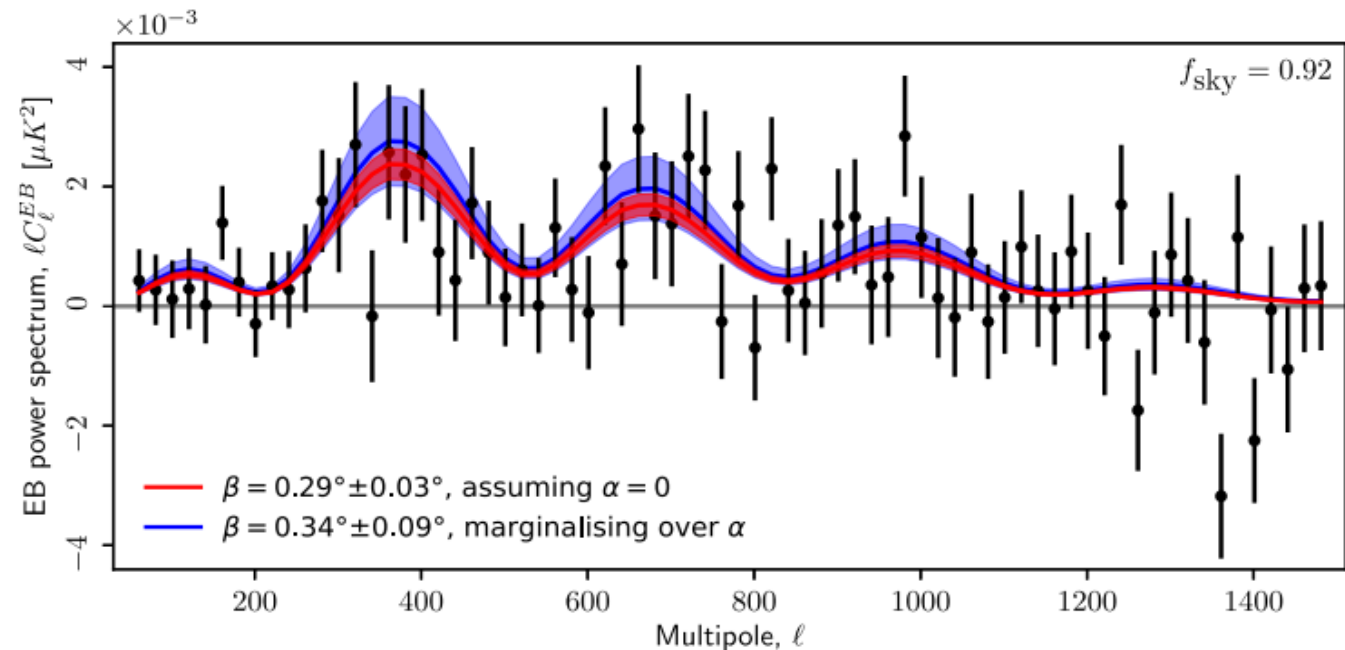
Rotation accumulates over distance

Birefringence of CMB is called  
COSMIC BIREFRINGENCE

Lue, Wang & Kamionkowski (1997); Feng et al. (2005,2006); Liu, Lee & Ng (2006)

# Cosmic Microwave Background as a perfect target

# Hint of parity violation



Uniform rotation of CMB polarization of an **angle  $\beta$**  generates an EB cross-correlation

$$C_1^{\text{EB,obs}} = \frac{1}{2} \sin(4\beta) (C_1^{\text{EE}} - C_1^{\text{BB}})$$

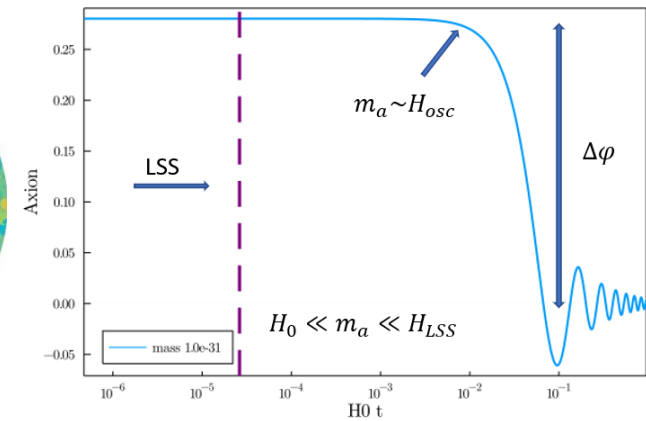
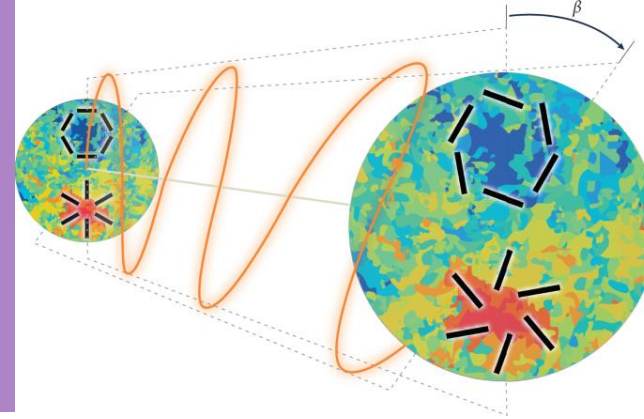
$\beta$  is degenerate with a miscalibration angle  $\rightarrow$  New method from Minami and Komatsu (2020) from reported  $\beta = 0.35^{\circ} \pm 0.14^{\circ}$  also Diego-Palazuelos et al. (2022), Eskilt (2022)

Current measure:

$$\beta = 0.342_{-0.091}^{+0.094} \text{ deg } (3.6\sigma) \quad \text{Frequency independent!}$$

Zero excluded at 99.987% C.L. from the joint analysis of Planck and WMAP data  
Eskilt et al (2023)

# Axion Explanation



$$\beta(\hat{n}) = \frac{1}{2} \int_{\eta_{em}}^{\eta_{obs}} d\eta (\omega_- - \omega_+) = \frac{g_{a\gamma}}{2} \int_{\eta_{em}}^{\eta_{obs}} d\eta \frac{d\phi}{d\eta} = \frac{g_{a\gamma}}{2} \underbrace{(\phi_{obs}(\hat{n}) - \phi_{em}(\hat{n}))}_{\text{Axion field displacement from CMB and today}}$$

Axion field displacement from CMB and today

The field evolution is model-dependent

$$\ddot{\phi} + 3H\dot{\phi} + V'(\phi) = 0$$

$$V(\phi) = \frac{f_a^2 m_a^2}{N^2} \left( 1 - \cos\left(\frac{\phi}{f_a N}\right) \right)$$

Different scenarios:

## Uniform axion-like background

- I. Pre-inflationary case ( $f_a \gg H_I$ )
- II. (Early-)Dark Energy or percentage of Dark matter
- III. Small fluctuations expected

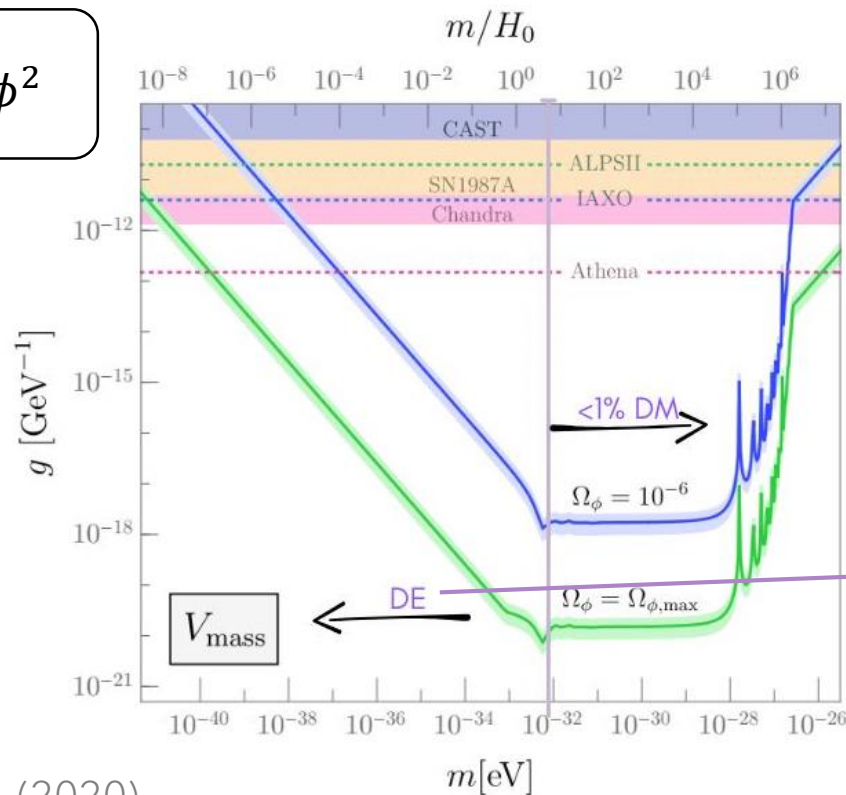
## Network of topological defects:

- I. Post-inflationary case ( $f_a \ll H_I$ ): cosmic strings and domain walls
- II. Large quantum diffusion ( $\sigma_{a_{in}} \gg f_a$ ): domain walls
- III. Large anisotropies, isotropic rotation on average

# Implication for single-field

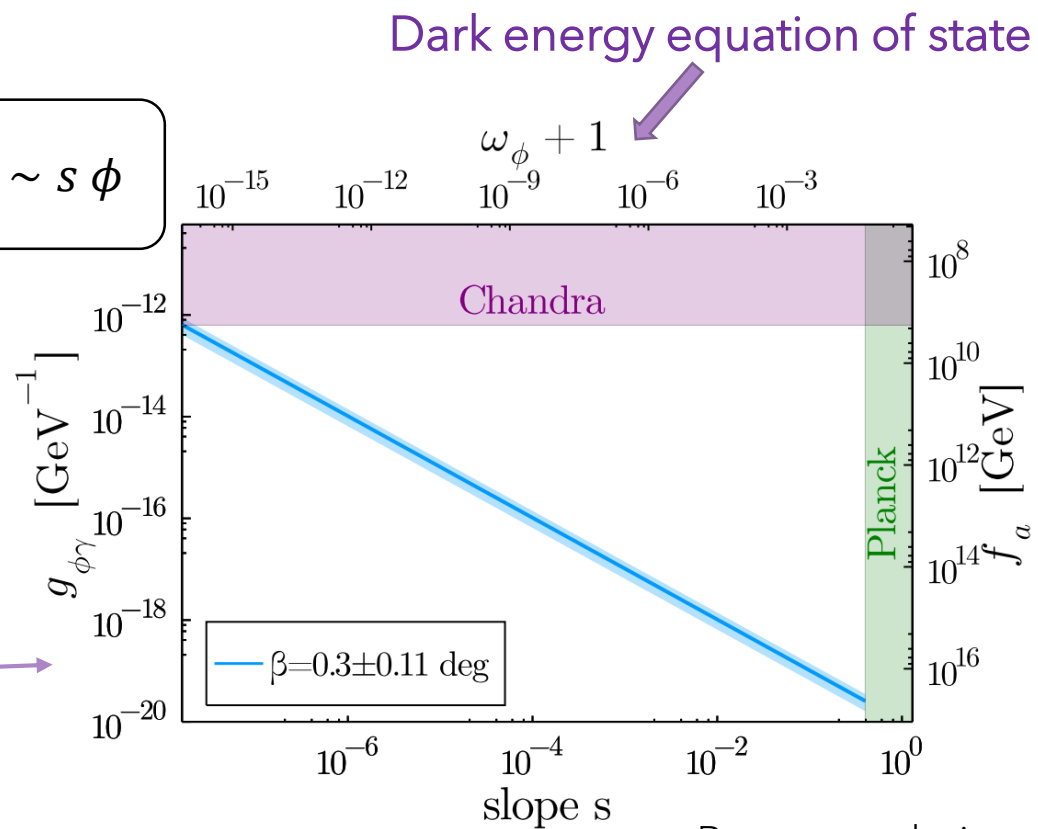
Parameter space explaining the CMB signal  $\beta \sim 0.3$  deg

$$V(\phi) = \frac{1}{2} m_a \phi^2$$



T. Fujita et al. (2020)

$$V(\phi) \sim s \phi$$



SG & I. Obata (2022)

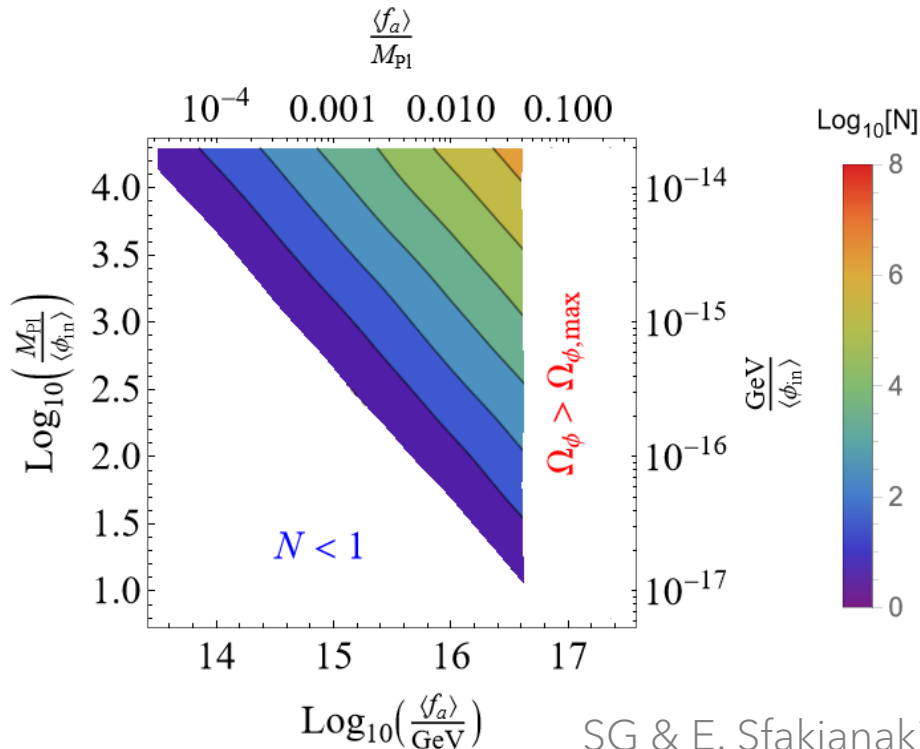
Recent evolution of the field

Axions within 15 orders of magnitude in mass could generate the signal even with a small energy density

→ AXIVERSE Arvanitaki et al (2009) ...

The total birefringence signal is given by the variance

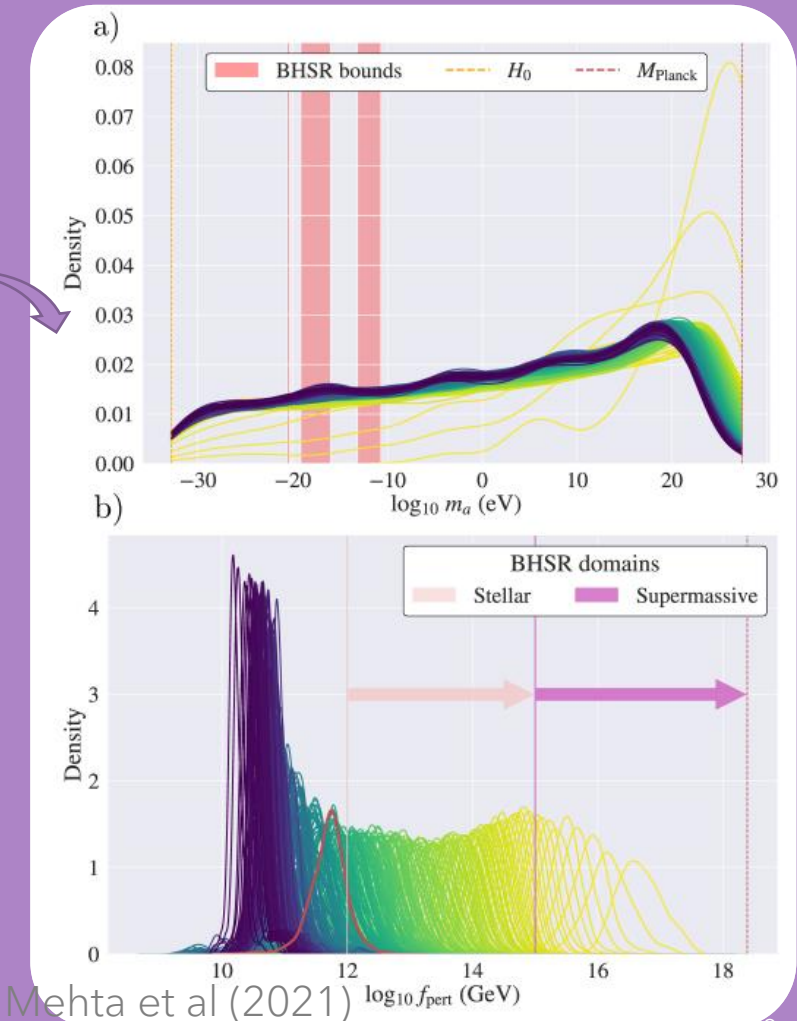
$$\beta = \sum_{i=0}^N \frac{\alpha_{em}}{2\pi f_{a,i}} \frac{\phi_{in,i}}{2} \Rightarrow \sqrt{\langle \beta^2 \rangle} = \frac{\alpha_{em}}{4\pi} \sqrt{\sum_{i=1}^N \left( \frac{\phi_{in,i}}{f_{a,i}} \right)^2} \propto \sqrt{N}$$



SG & E. Sfakianakis (2023)

The model depends on the probabilistic distribution of the model parameter  $(m_a, f_a)$

# Multi-fields and the Axiverse



Mehta et al (2021)



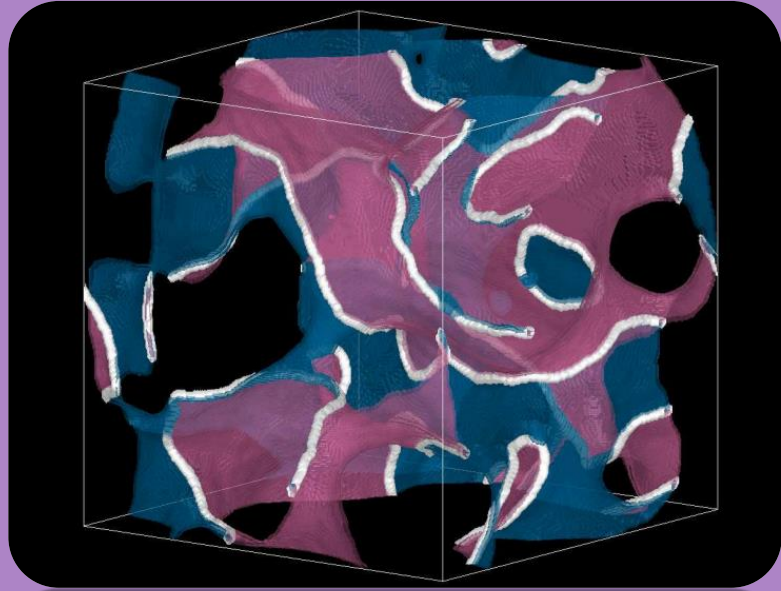
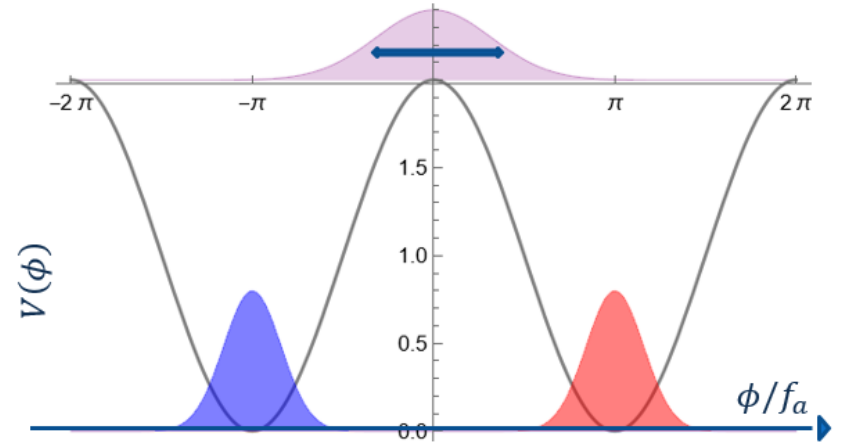
# Topological defects: Domain Wall



## Isotropic rotation

$$\langle \beta \rangle = \frac{\alpha_{em} c}{4\pi} (\theta_0 - \langle \theta_{LS}(\hat{n}) \rangle)$$

Symmetry breaking broken by the field local value

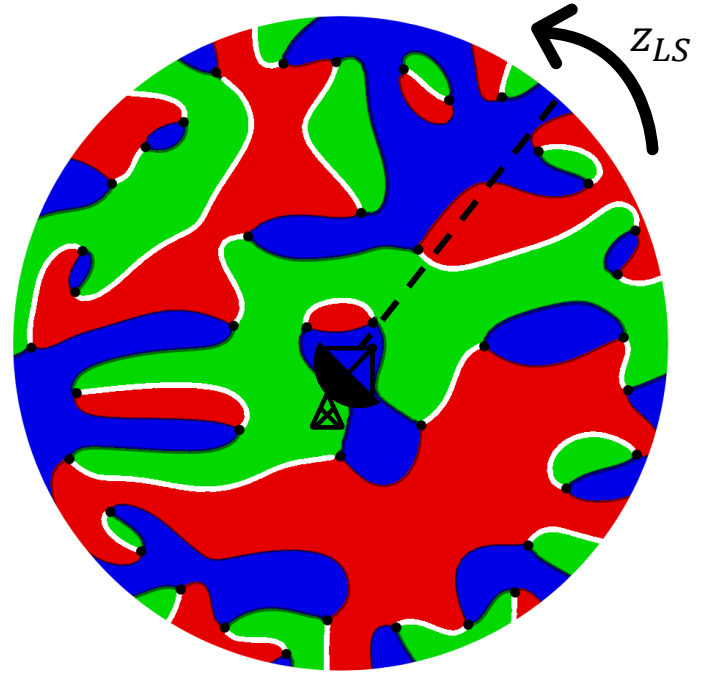


## Anisotropic rotation

Anisotropic power spectrum

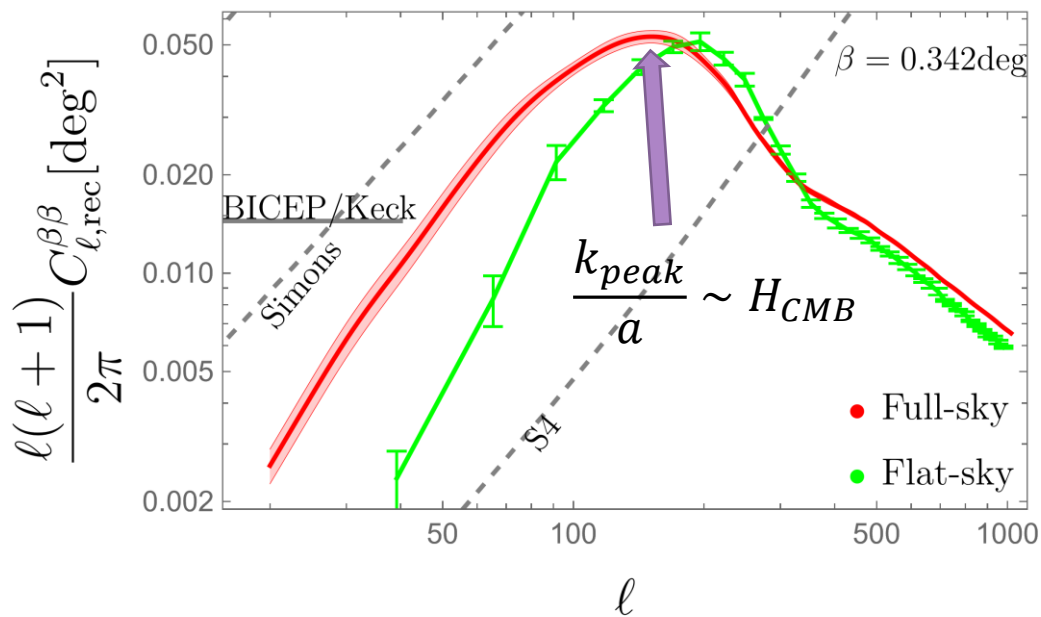
$$C_\ell^{\beta\beta}(\eta) = \frac{4}{\pi} \beta_{iso}^2 \int \frac{dk}{k} J_\ell^2(k\Delta\eta) P_\theta(k)$$

Depends on the field power spectrum at CMB



# Birefringence and gravitational waves

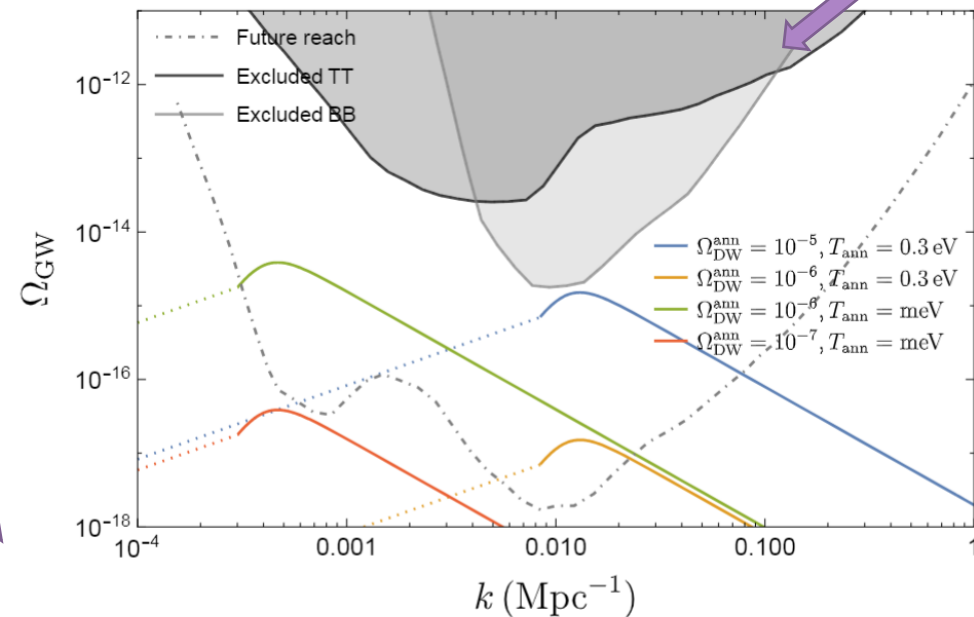
Anisotropic Birefringence:  
(still undetected)



Assuming “*scaling regime*” (1 domain wall per Hubble volume), the spectrum peaks at CMB horizon scale.  
Birefringence does not depend on the field energy density

Stochastic Gravitational-wave background:

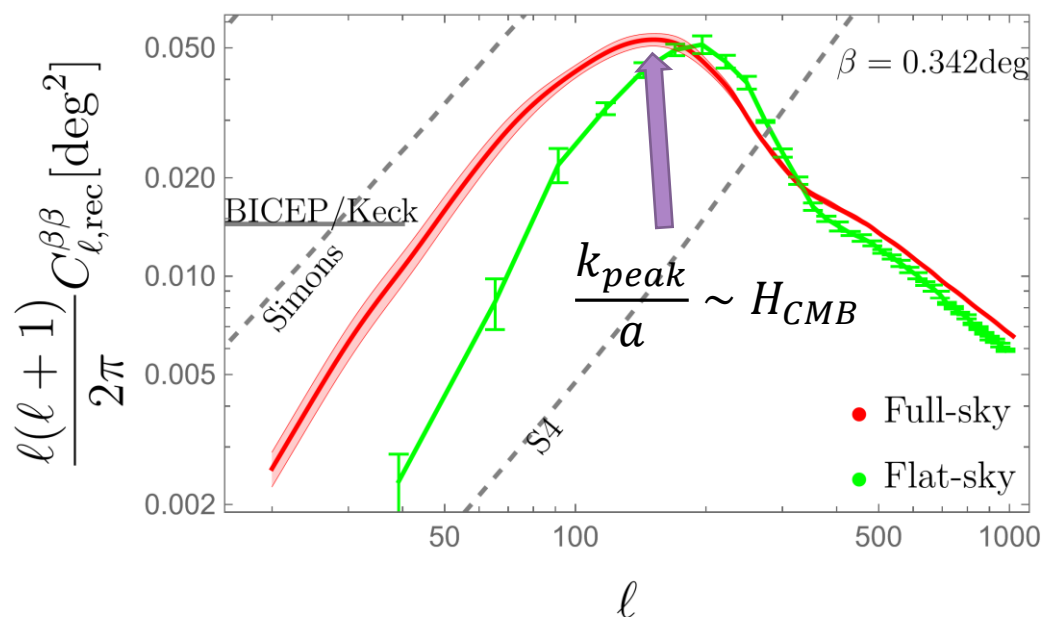
Current and future CMB constraint on the stochastic GW background from Namikawa et al (2019)



Bound on the **GW energy density** from domain wall network still present or annihilated after CMB

# Birefringence and gravitational waves

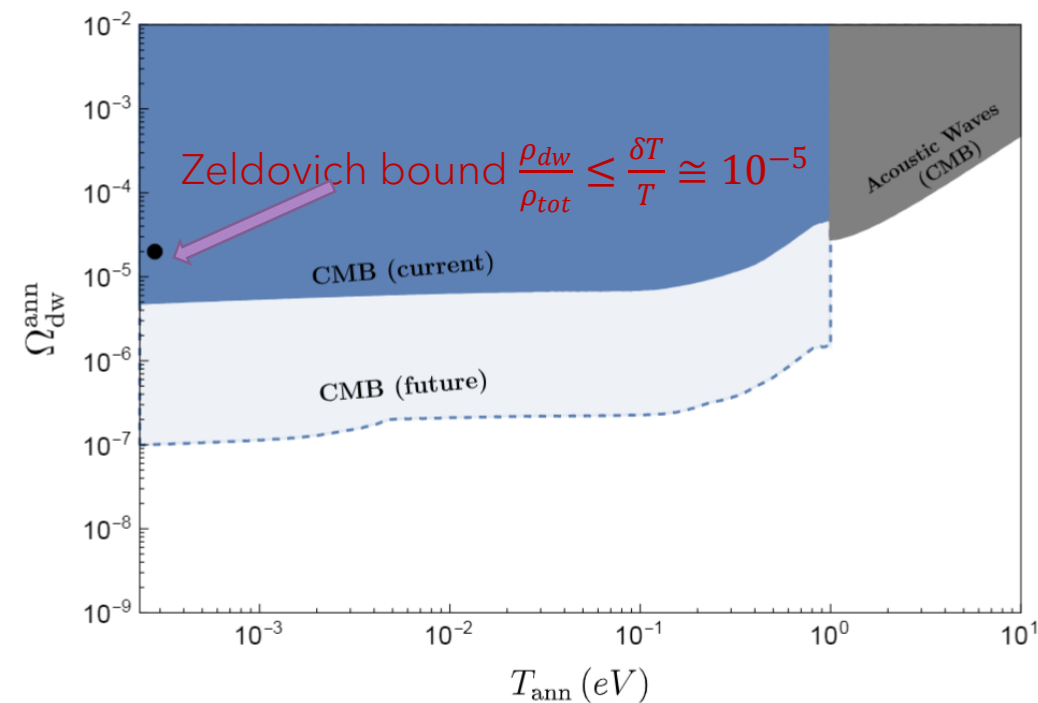
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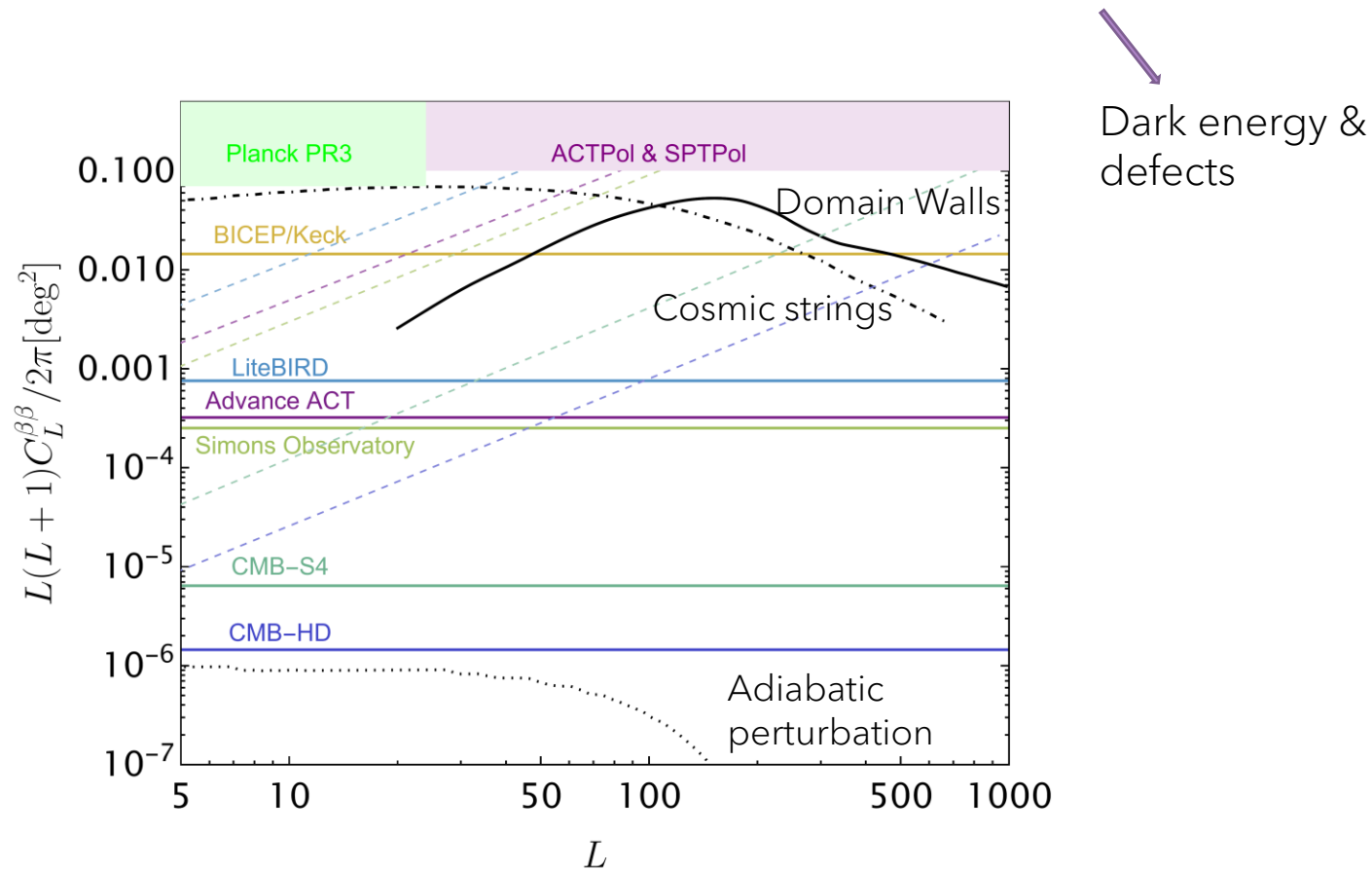
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Stochastic Gravitational-wave background:

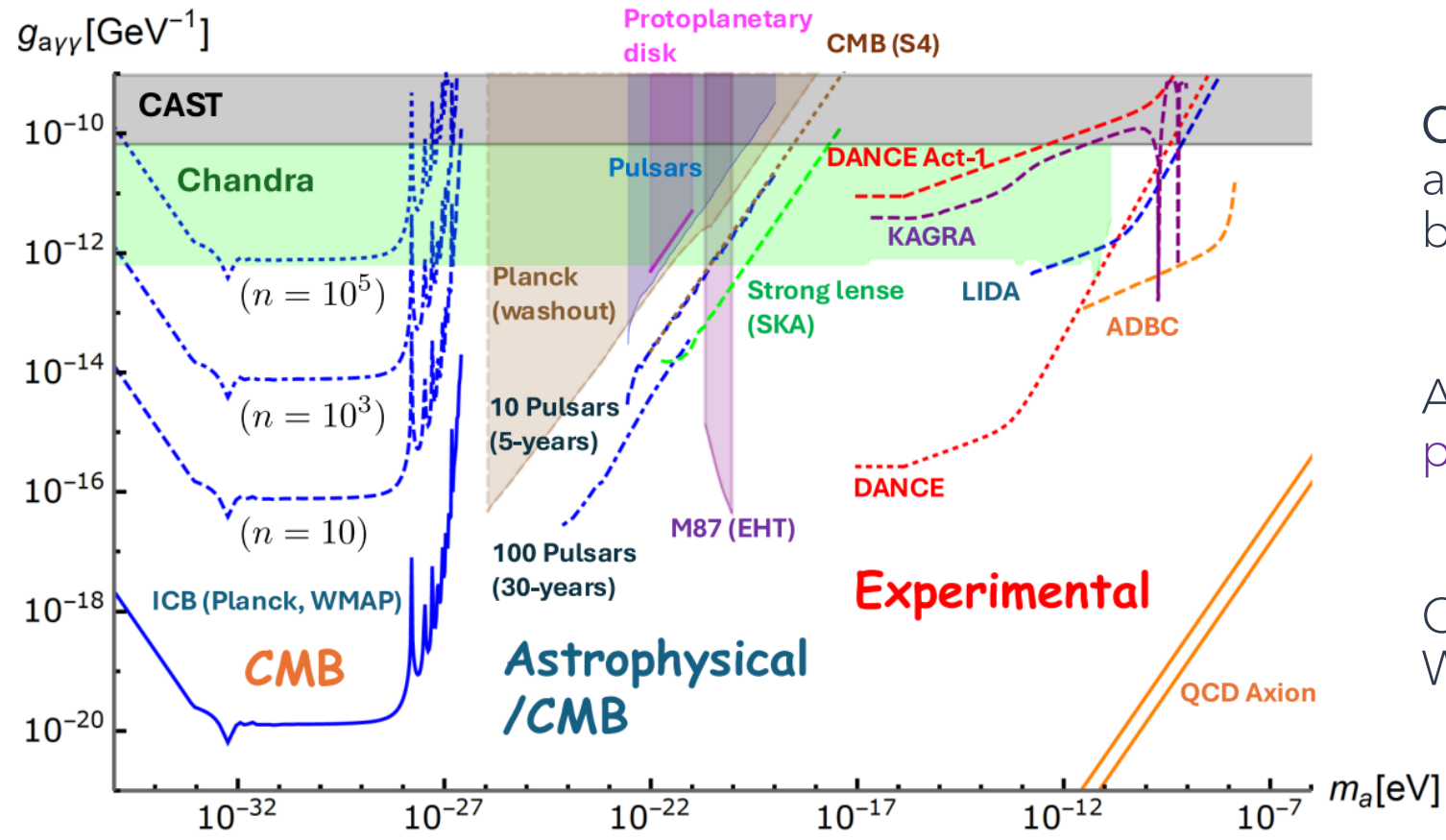


If the CMB signal is confirmed, how do we distinguish between models? The isotropic signal is very degenerate!

- I. Anisotropic counterpart
- II. Tomography to study its evolution (ask me)
- III. Looking for other polarized sources (from local universe)



# Current and future prospects



Overview of constraints and forecast from astrophysical and laboratory searches of birefringence

At higher axion masses the signal would be periodic with frequency  $\omega = m_a$

Contribution to White Paper of Cosmic WISPer COST action (in preparation)

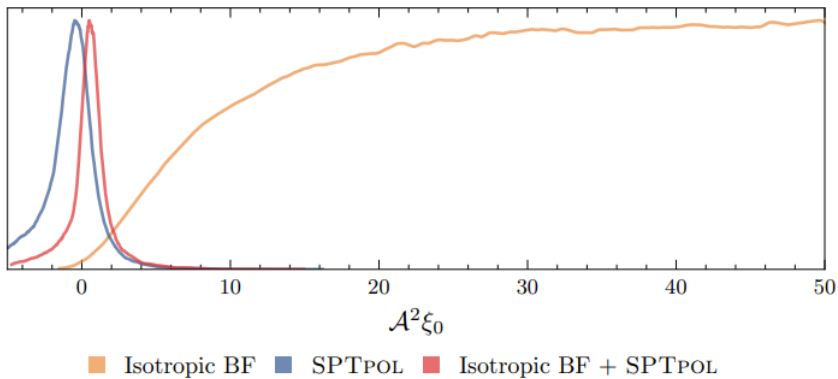
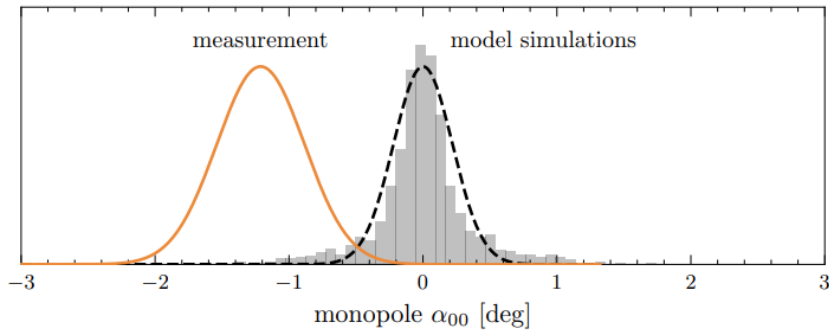
# Conclusions

Thank you!

Contact: [sgasparotto@ifae.es](mailto:sgasparotto@ifae.es)

# String-wall and String network case

It was considered by Agrawal et al 2020, Jain et al 2021 & 2022 → in the presence of strings the signal can be enhanced at small angles because after each loop-crossing birefringence gets a  $\Delta\theta = \pm\alpha_{em}c$  → it seemed that the non-detection of anisotropies is incompatible with the isotropic signal.



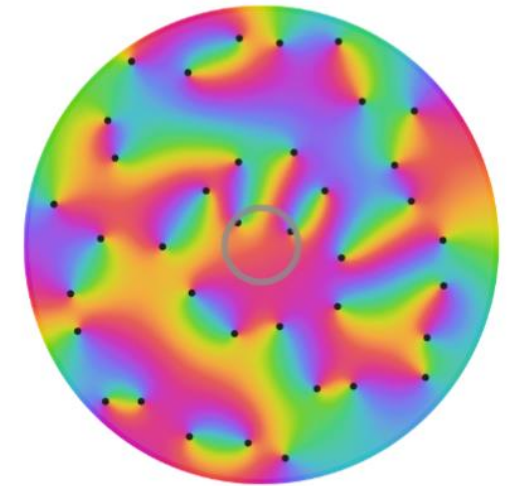
**BUT** the local gradient/value of the network which differs from the average value at recombination also contributes to the monopole which is not captured in the loop-crossing model

$$\langle\beta\rangle = \frac{0.21c}{2\pi} \left( \frac{\phi_{loc}}{f_a} - \langle\theta_s(\hat{n})\rangle \right)$$

*Environmental birefringence*



It seems that the isotropic birefringence naturally arises in every axionic network



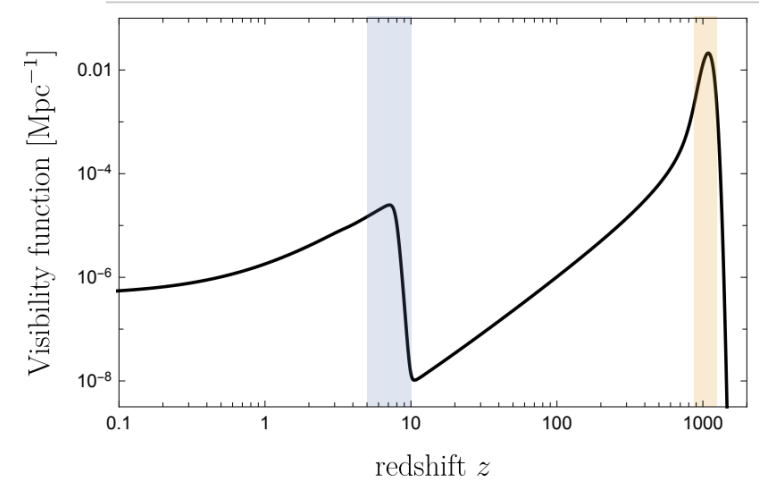
Jain et al 2022



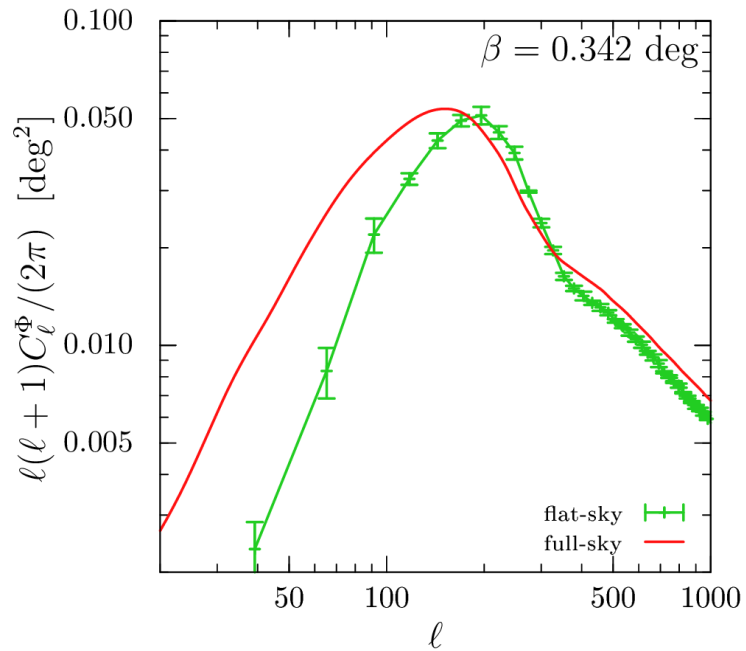
# DWs at recombination and reionization

Anisotropies in the scalar power spectrum translate into anisotropies in the cosmic birefringence,

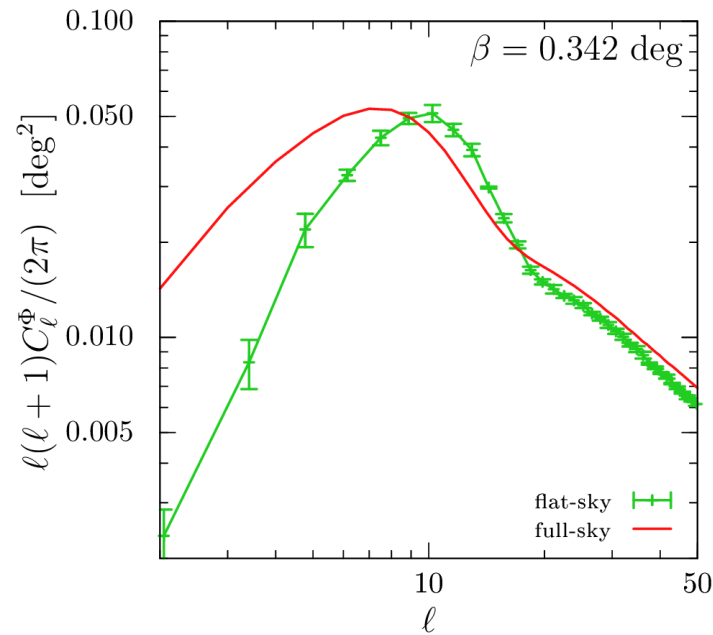
$$C_{\ell}^{\beta\beta}(\eta) = \frac{4}{\pi} \beta_{iso}^2 \int \frac{dk}{k} J_{\ell}^2(k\Delta\eta) P_{\theta}(k)$$



## Recombination



## Reionization



Contributions coming from the DW network at **recombination and reionization** which peak at different scales  
 → **birefringence tomography** can be used to distinguish different **formation/annihilation scenarios**