



Universität Hamburg

DER FORSCHUNG | DER LEHRE | DER BILDUNG

air pressure fluctuations

seismic noise

ocean waves

FAKULTÄT
FÜR MATHEMATIK, INFORMATIK
UND NATURWISSENSCHAFTEN

storage ring

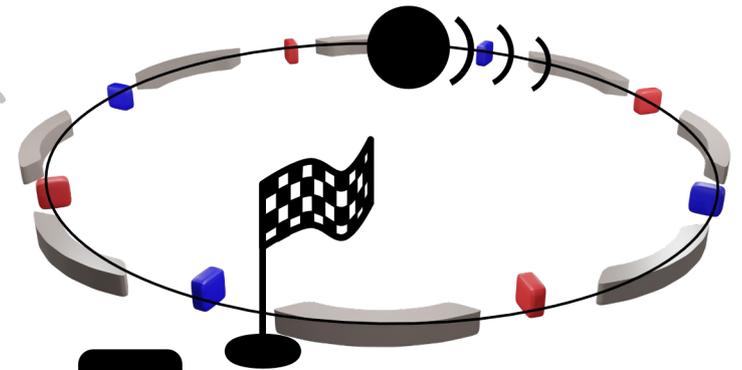
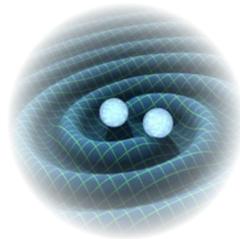
CONCEPT STUDY OF A STORAGE RING-BASED GRAVITATIONAL WAVE OBSERVATORY

Storage Rings & Gravitational Waves – mini-brainstorm

THORBEN SCHMIRANDER

10.02.2025

Idea



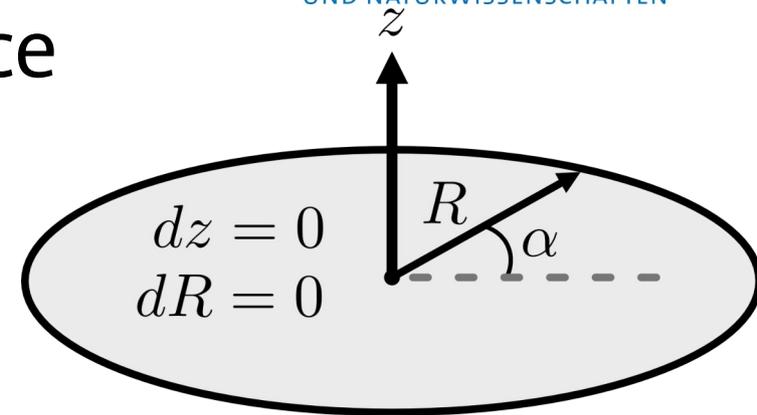
a **VERY** precise clock

- Variation of circulation time encodes GW signature
- Need: Particle on fixed circular trajectory in longitudinal free fall
- can circulate for minutes up to hours: mHz GW

Frame of reference

metric for ring with fixed radius (cyl. coords):

$$ds^2 = -c^2 dt^2 + (1 + h_{\theta\phi\psi}(t, \alpha)) R^2 d\alpha^2$$



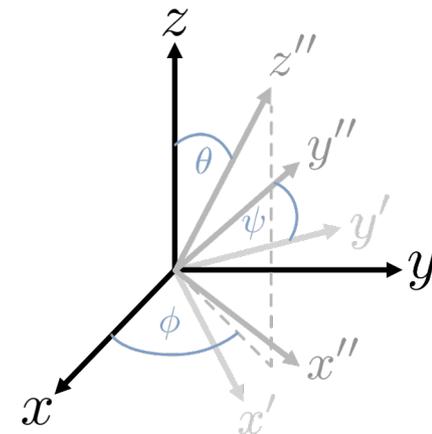
Effective GW strain includes Earth's rotation:

$$h_{\theta\phi\psi}(t, \alpha) = h_+(t)(f_s^+ \sin^2 \alpha + f_c^+ \cos^2 \alpha + \dots) + h_\times(t)(\dots)$$



Euler angles enter via e.g.:

$$f_s^+ = (\cos^2 \theta \cos^2 \phi - \sin^2 \phi) \cos 2\psi - (\cos \theta \sin 2\phi) \sin 2\psi$$



Particle in a Storage Ring

From geodesic equations:
$$\frac{d^2 l}{dt^2} = - \frac{1}{1+h_{\theta\phi\psi}(t)} \frac{dh_{\theta\phi\psi}(t)}{dt} v_0$$

longitudinal acceleration

For fixed radius:

$$a_{\parallel}(t) \approx -\dot{h}_{\theta\phi\psi}(t)v_0$$

$$\Delta T^{\text{fixed}} = \frac{\Delta l^{\text{fixed}}}{v_0}$$

$$= - \int_0^t (h_{\theta\phi\psi}(t') - h_{\theta\phi\psi}(0)) dt'$$

~ relativistic result

Main idea:

- GW strain as classical force
- Find transformation between ring with fixed radius and storage ring simulation
- Find reference to make noise terms comparable

Particle in a Storage Ring

For storage ring:

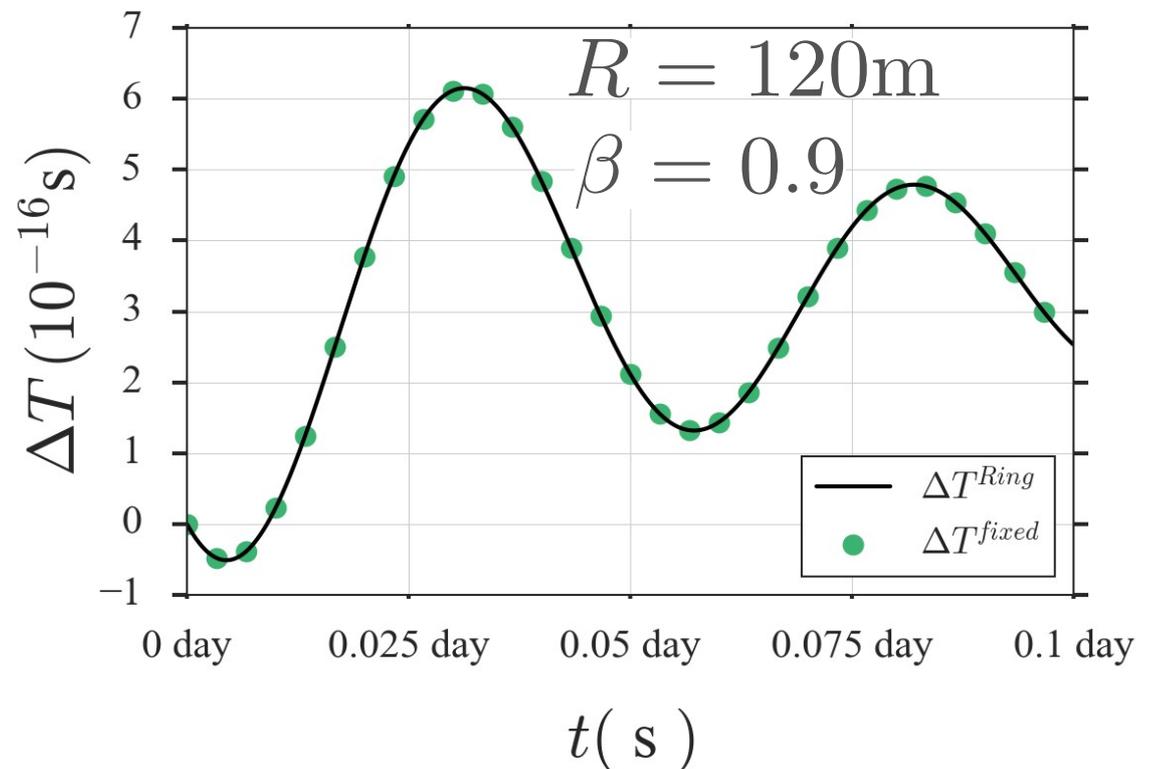
$$\Delta T^{\text{fixed}} = -\frac{1}{\eta} \frac{1}{\gamma^2} \left(1 - \frac{v_0^2}{2c^2}\right) \Delta T^{\text{Ring}} - \left(1 - \frac{v_0^2}{2c^2}\right) \overline{h_{\theta\phi\psi}(t_0)t}$$

“slip factor”

Inspiral @ $z=0.2$

$$m_{1/2} = 10^6 M_{\odot}$$

on order of 0.1 fs!!!



Particle tracking I

radial EOM:

$$\ddot{x} - \omega^2(\rho_0 + x) = -\frac{q}{p}v_s^2 B_y$$

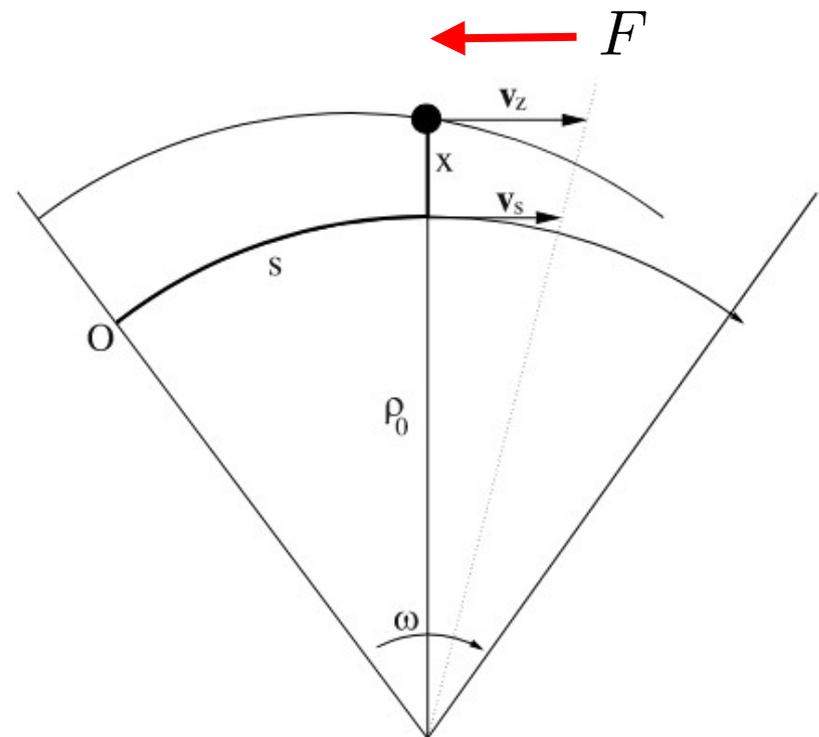
$$v_s = (\rho_0 + x)\omega + \frac{\langle F \rangle}{m\gamma^3\beta c} \cdot s$$

$$x'' + \frac{1-n}{\rho_0}x = \frac{\delta}{\rho_0} - \frac{2z\langle F \rangle}{\rho_0^2\omega m\gamma^3\beta c}$$

Green's function approach:

$$x(s) = x_0 c_x(s) + x'_0 s_x(s) + \delta d_{x1}(s) + d_{x2}(s)$$

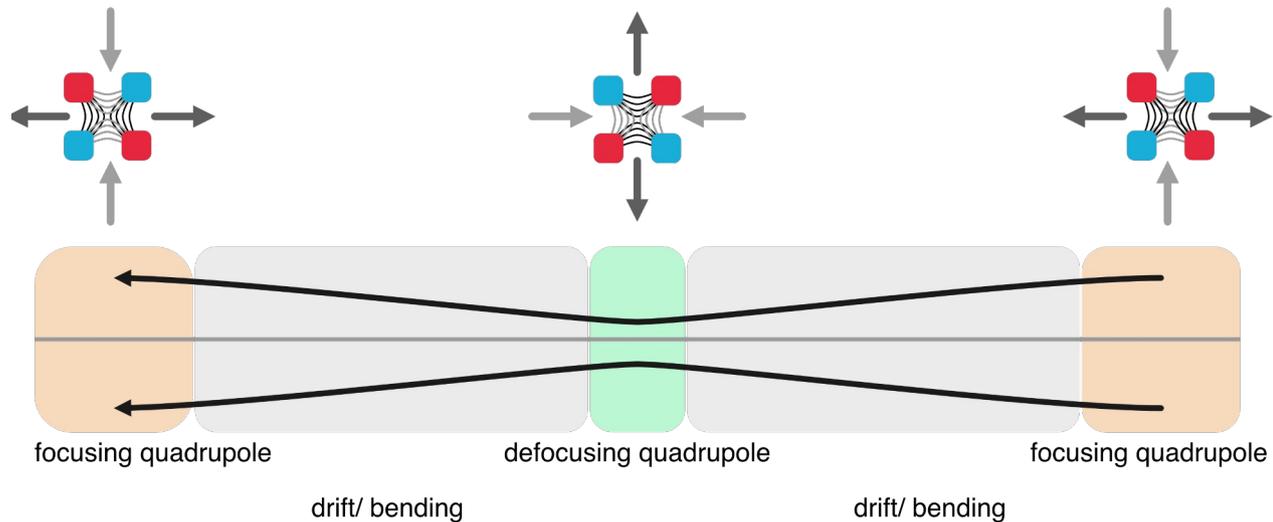
Include longitudinal force



First-order EOM particle tracking: $x_{\text{out}} = M_{\text{drift}} \cdot x_{\text{in}} + \Delta x$

Particle tracking II

$x = \text{function of } \left\{ \begin{array}{l} \text{radial} \\ \text{longitudinal} \\ \text{azimuthal} \end{array} \right\} \left\{ \begin{array}{l} \text{positions} \\ \text{momenta} \end{array} \right\} \text{ relative to a design particle}$



$$x_{\text{out}} = M_{\text{drift}} \cdot x_{\text{in}} + \Delta x_1$$

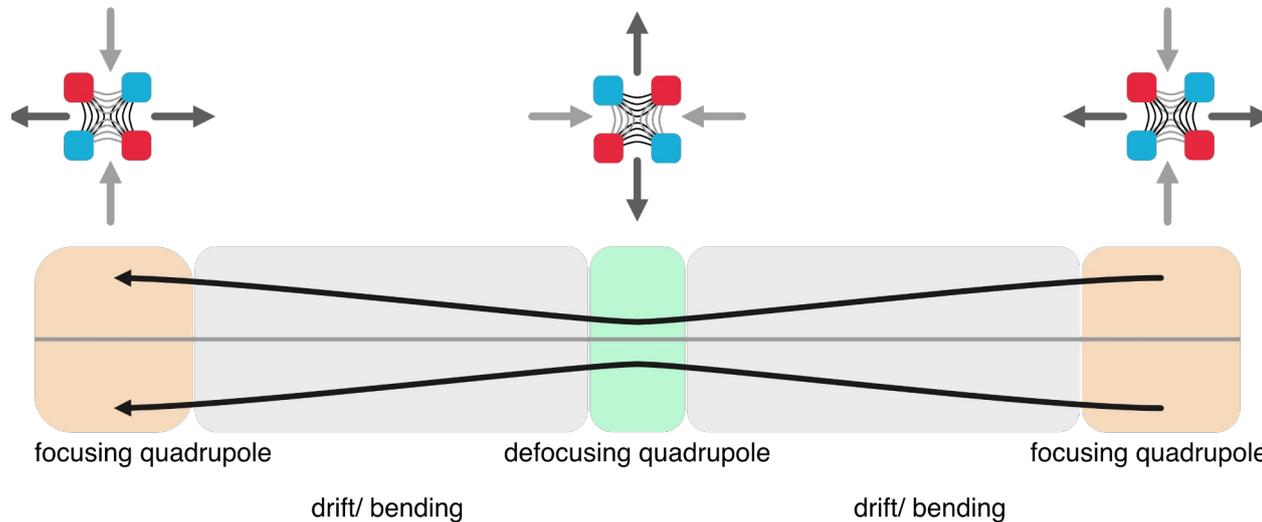
$$x_{\text{out}} = M_{\text{sector}} \cdot x_{\text{in}} + \Delta x_2$$

⋮

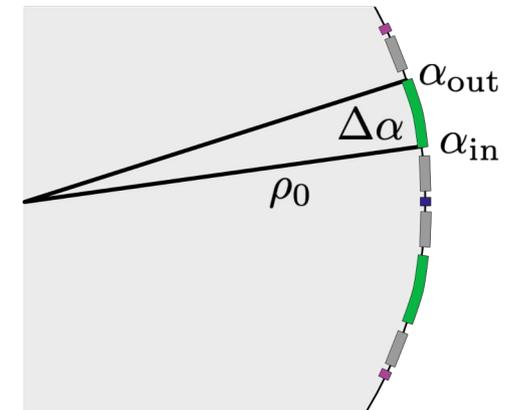
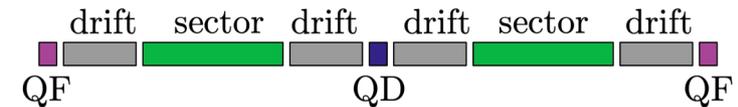
↑
GW force

Particle tracking II

$x = \text{function of } \left\{ \begin{array}{l} \text{radial} \\ \text{longitudinal} \\ \text{azimuthal} \end{array} \right\} \left\{ \begin{array}{l} \text{positions} \\ \text{momenta} \end{array} \right\}$ relative to a design particle



“FODO” cell



$$x_{out} = M_{drift} \cdot x_{in} + \Delta x_1$$

$$x_{out} = M_{sector} \cdot x_{in} + \Delta x_2$$

\vdots

↑
GW force

e.g. for U_2^+ :

@ $\beta = 0.32$

150 GeV

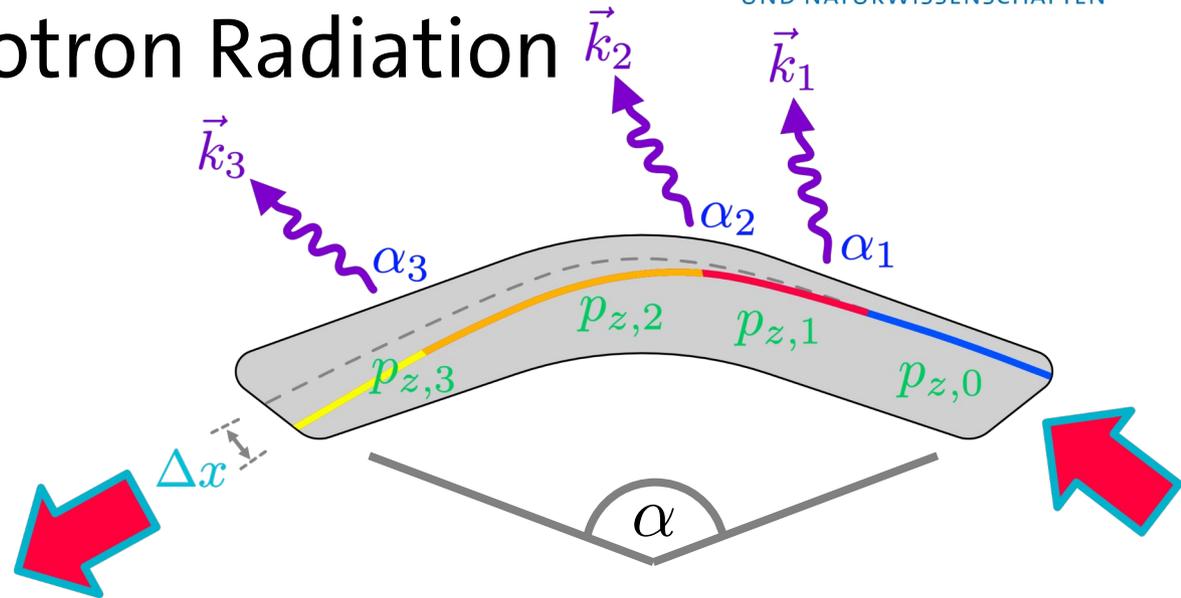
$$L_{FODO} = 133.5 \text{ m}$$

$$200 \cdot L_{FODO}$$

$$= 26.7 \text{ km}$$

$$= L_{LHC}$$

Synchrotron Radiation

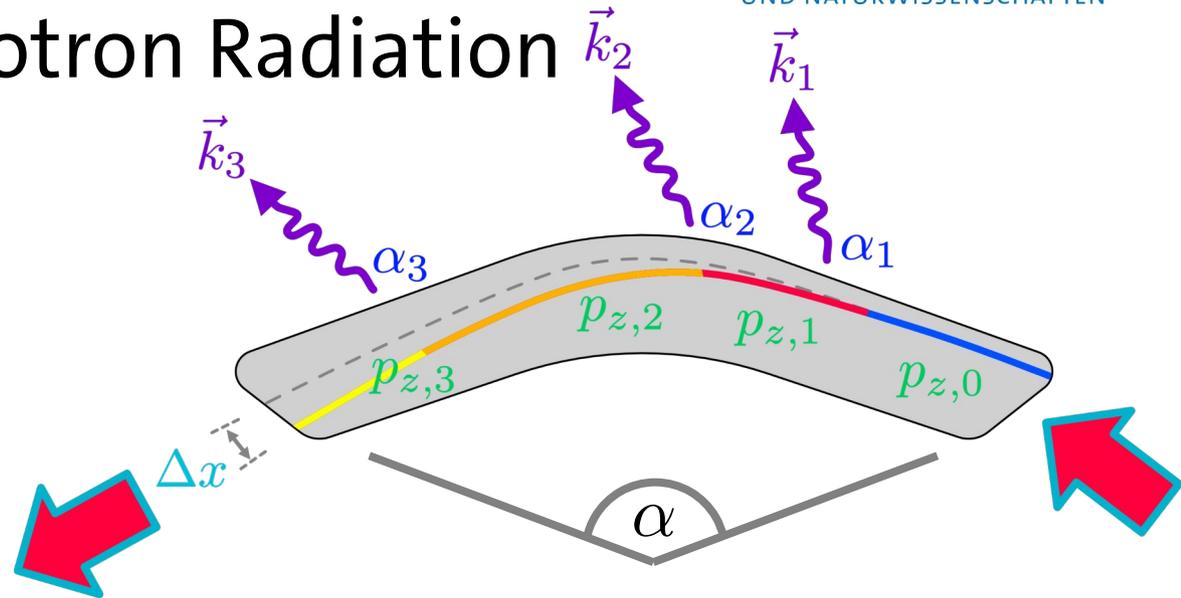


$$x_0(l) = M_{\text{sector}}(\alpha) \cdot x(0)$$

$$x_1(l) = M_{\text{sector}}(\alpha - \alpha_1) \left(M_{\text{sector}}(\alpha_1) \cdot x(0) + \Delta_\gamma(\vec{k}_1) \right)$$

$$\Delta x(\alpha_3, \dots, \vec{k}_3, \dots) = M_{\text{sector}}(\alpha - \alpha_3) \left(\dots + \Delta_\gamma(\vec{k}_3) \right) - x_0(l)$$

Synchrotron Radiation



$$x_0(l) = M_{\text{sector}}(\alpha) \cdot x(0)$$

$$x_1(l) = M_{\text{sector}}(\alpha - \alpha_1) \left(M_{\text{sector}}(\alpha_1) \cdot x(0) + \Delta_\gamma(\vec{k}_1) \right)$$

$$\Delta x(\alpha_3, \dots, \vec{k}_3, \dots) = M_{\text{sector}}(\alpha - \alpha_3) \left(\dots + \Delta_\gamma(\vec{k}_3) \right) - x_0(l)$$

- Is given by a “simple” shift: $x(l) = M_{\text{sector}} \cdot x(0) + \Delta x$
- Heavy ions!
- Need distribution for angles and momenta

Sampling of Synchrotron Photons

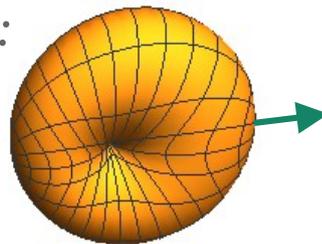
$$F(\theta, \phi) = \frac{(1 - \beta \cos \theta)^2 - (1 - \beta^2) \sin^2 \theta \cos^2 \phi}{(1 - \beta \cos \theta)^5}$$

$$\beta = 0.1$$

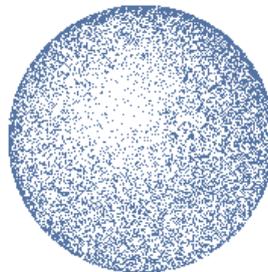
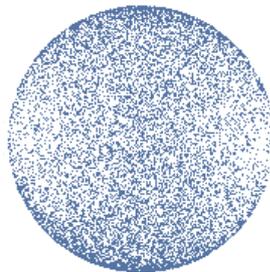
front:



side:



propagation
direction



1. $\{\theta_i, \phi_i\} \in [0, \pi] \times [0, 2\pi]$

2. stratification

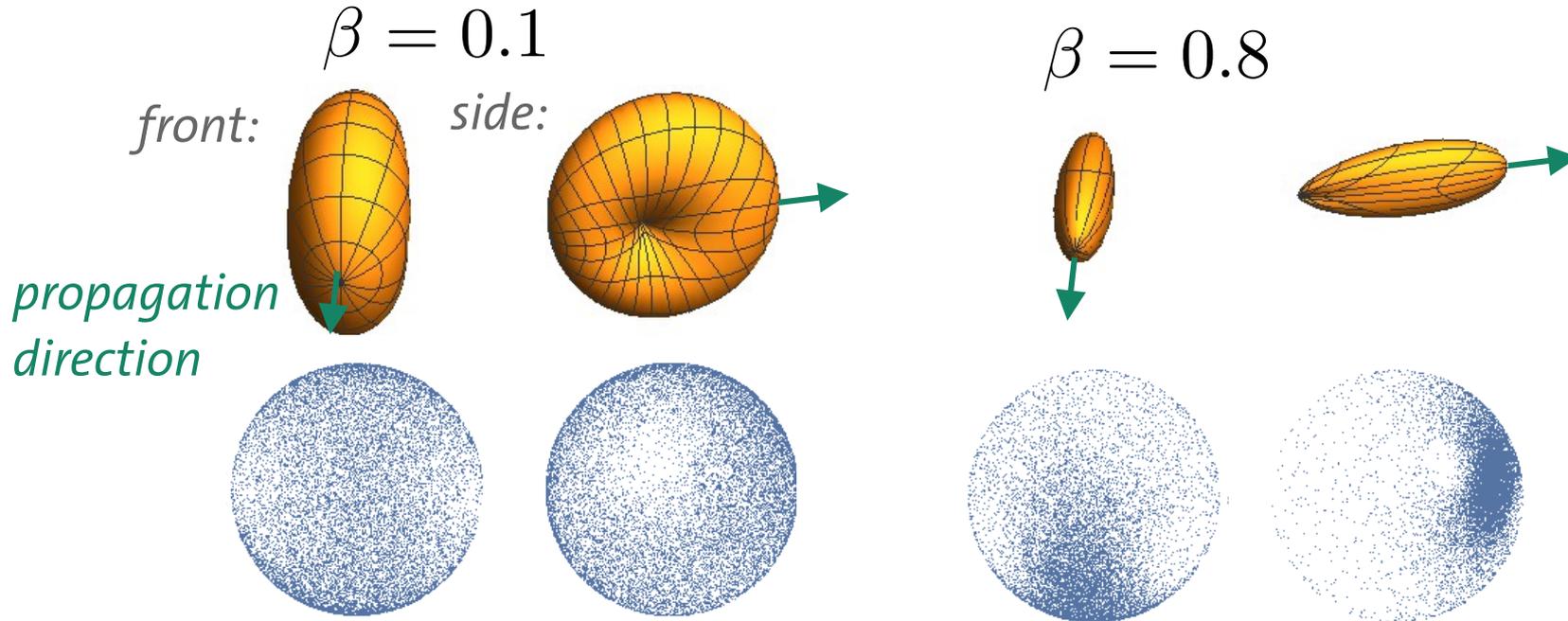
“rejection sampling”

3. $f_i \in [0, F(\theta, \phi)]$

See: Arvo et al. SIGGRAPH (2004)

Sampling of Synchrotron Photons

$$F(\theta, \phi) = \frac{(1 - \beta \cos \theta)^2 - (1 - \beta^2) \sin^2 \theta \cos^2 \phi}{(1 - \beta \cos \theta)^5}$$



1. $\{\theta_i, \phi_i\} \in [0, \pi] \times [0, 2\pi]$

2. stratification

“rejection sampling”

3. $f_i \in [0, F(\theta_i, \phi_i)]$

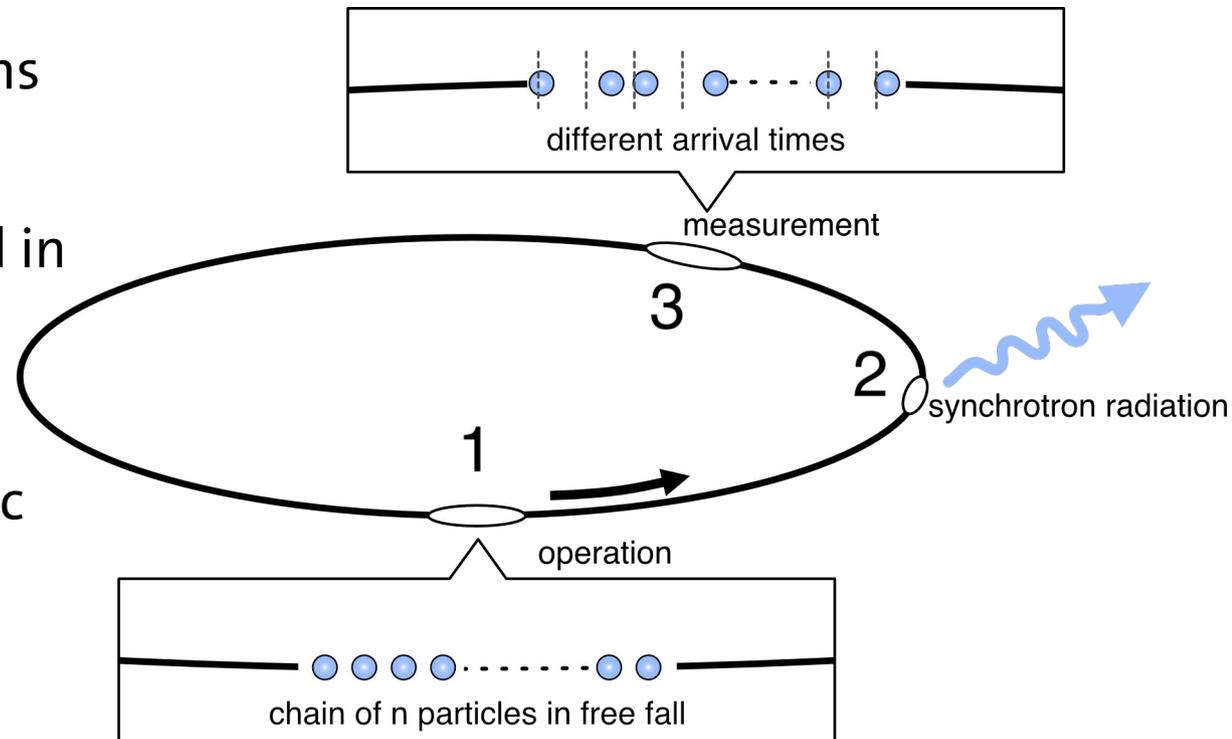
See: Arvo et al. SIGGRAPH (2004)

Concept Experiment Setup

1: operate under stable conditions for many rounds

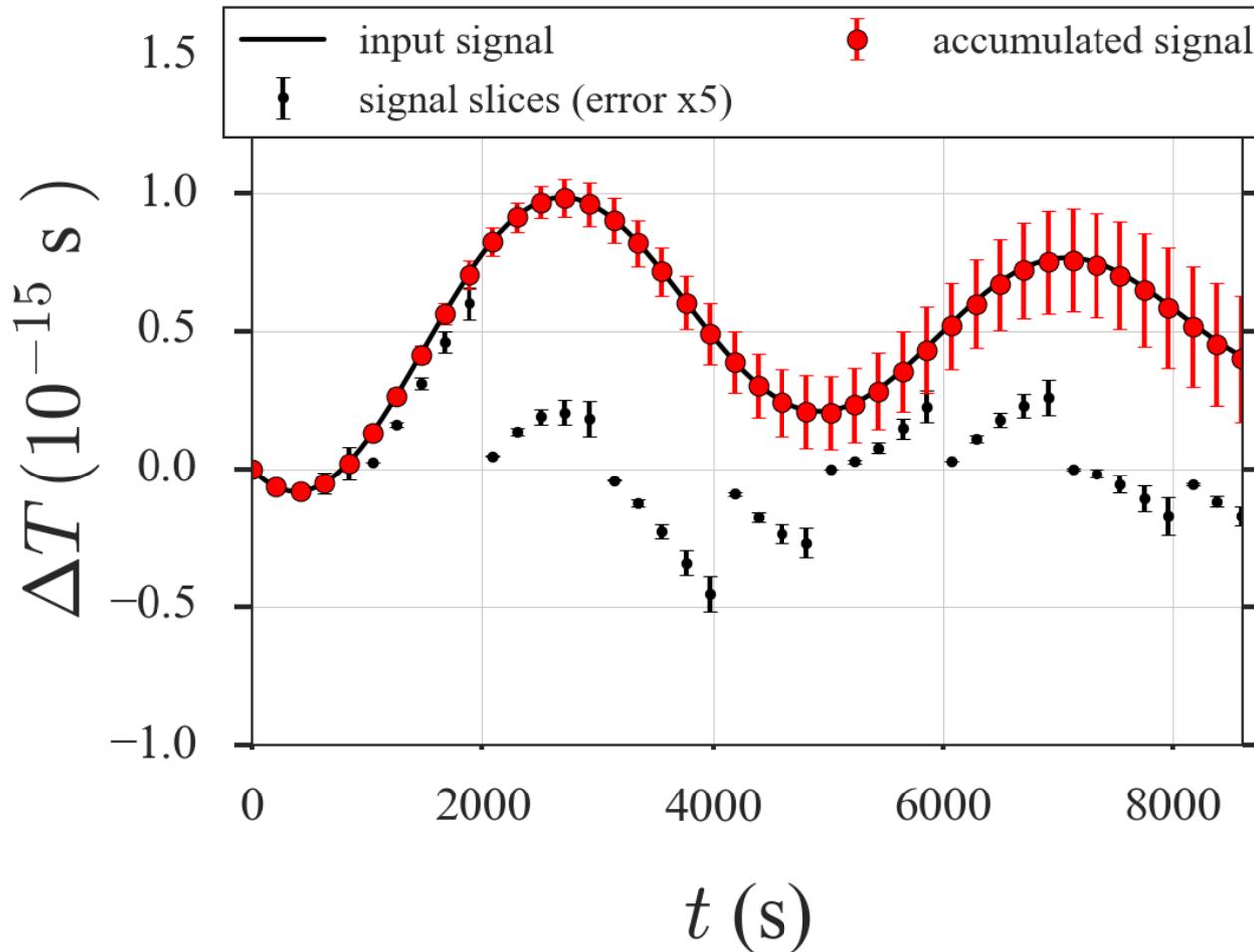
2: synchrotron radiation emitted in sector magnets

3: measure timing relative to external ticking (from e.g. atomic clock)



- all particle subject to **SAME** GW force + **INDIVIDUAL** SR emission
- expected positions (dashed lines) incl. **AVERAGE** SR emission

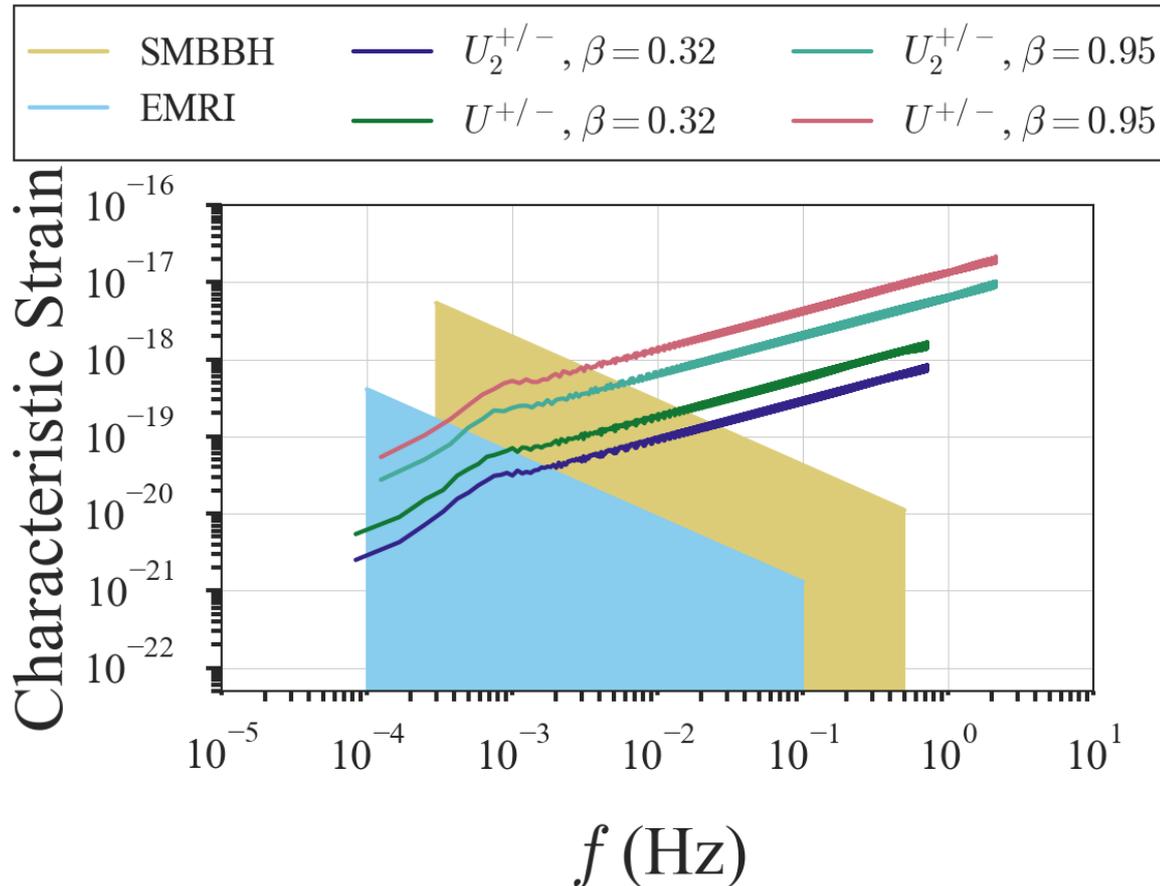
Results



- chain of 100 U_2^+ ions
- dumped after $t=10^3$ s
- $\beta = 0.32$
- $L = 26.7$ km

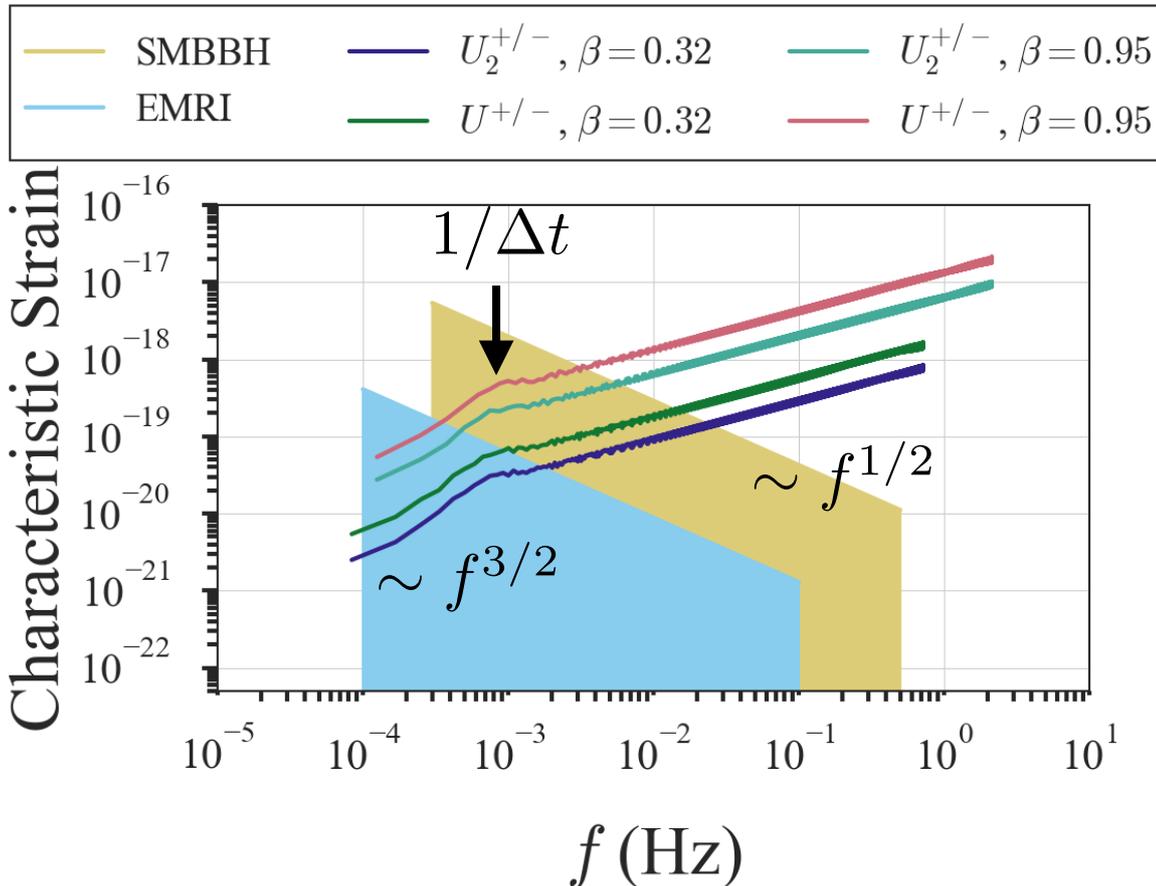
error of mean arrival time **ONLY** due to synch. rad. emission power fluctuations

Characteristic Noise Strain



ONLY due to synch. rad. emission power fluctuations , no other noise (read-out, etc.)

Characteristic Noise Strain



$$f < \Delta t$$

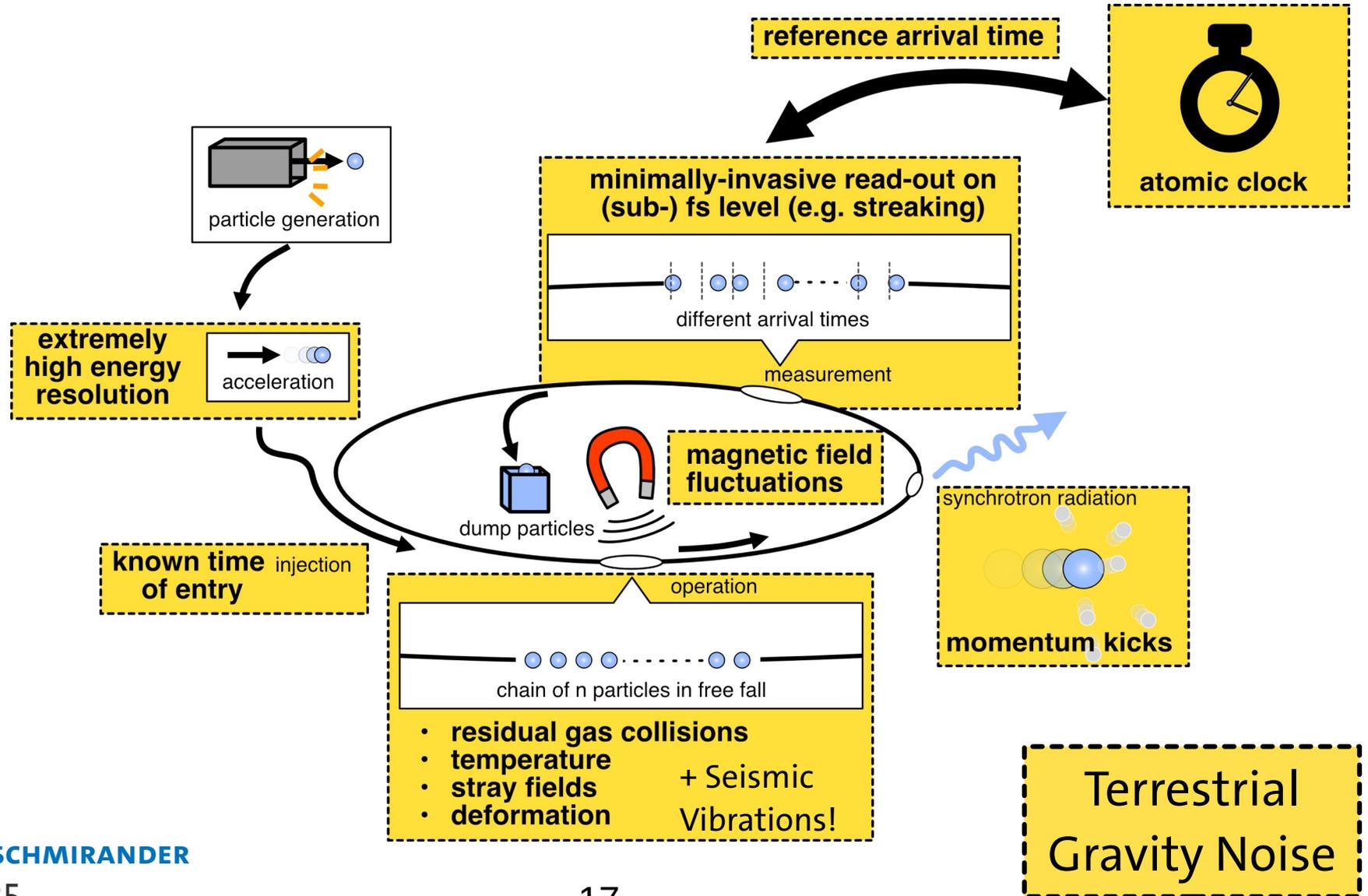
$$\sqrt{f S_n(f)} \sim f^{3/2}$$

$$f > \Delta t$$

$$\sqrt{f S_n(f)} \sim f^{1/2}$$

ONLY due to synch. rad. emission power fluctuations , no other noise (read-out, etc.)

Still open design specifics:





References

PHYSICAL REVIEW D **102**, 122006 (2020)

Detection of gravitational waves in circular particle accelerators

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Hamburger Sternwarte, University of Hamburg, Gojenbergsweg 112, 21029 Hamburg, Germany

(Received 10 July 2020; accepted 30 November 2020; published 22 December 2020)

Storage Rings and Gravitational Waves: Summary and Outlook

A. Berlin¹, M. Brüggem², O. Buchmueller³, P. Chen⁴, R. T. D'Agnolo⁵, R. Deng⁶,
J. R. Ellis^{7,*}, S. Ellis⁵, G. Franchetti⁸, A. Ivanov⁹, J. M. Jowett⁸,
A. P. Kobushkin¹⁰, S. Y. Lee¹¹, J. Liske², K. Oide¹², S. Rao², J. Wenninger¹³,
M. Wellenzohn⁹, M. Zanetti¹⁴, F. Zimmermann^{13,*}

^x *Editors*

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PHYSICAL REVIEW D **110**, 022007 (2024)

Detection of gravitational waves in circular particle accelerators II. Response analysis and parameter estimation using synthetic data

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PHYSICAL REVIEW D **110**, 082002 (2024)

Concept study of a storage ring-based gravitational wave observatory: Gravitational wave strain and synchrotron radiation noise

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Conclusion

- Particle in a storage ring: time delay encodes GW signature
- Include GW force & Synchrotron radiation (numerics)
- Concept of experimental setup
- Characteristic noise strain of radiation fluctuation
- open experiment setup questions/ noise sources

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air pressure fluctuations

seismic noise

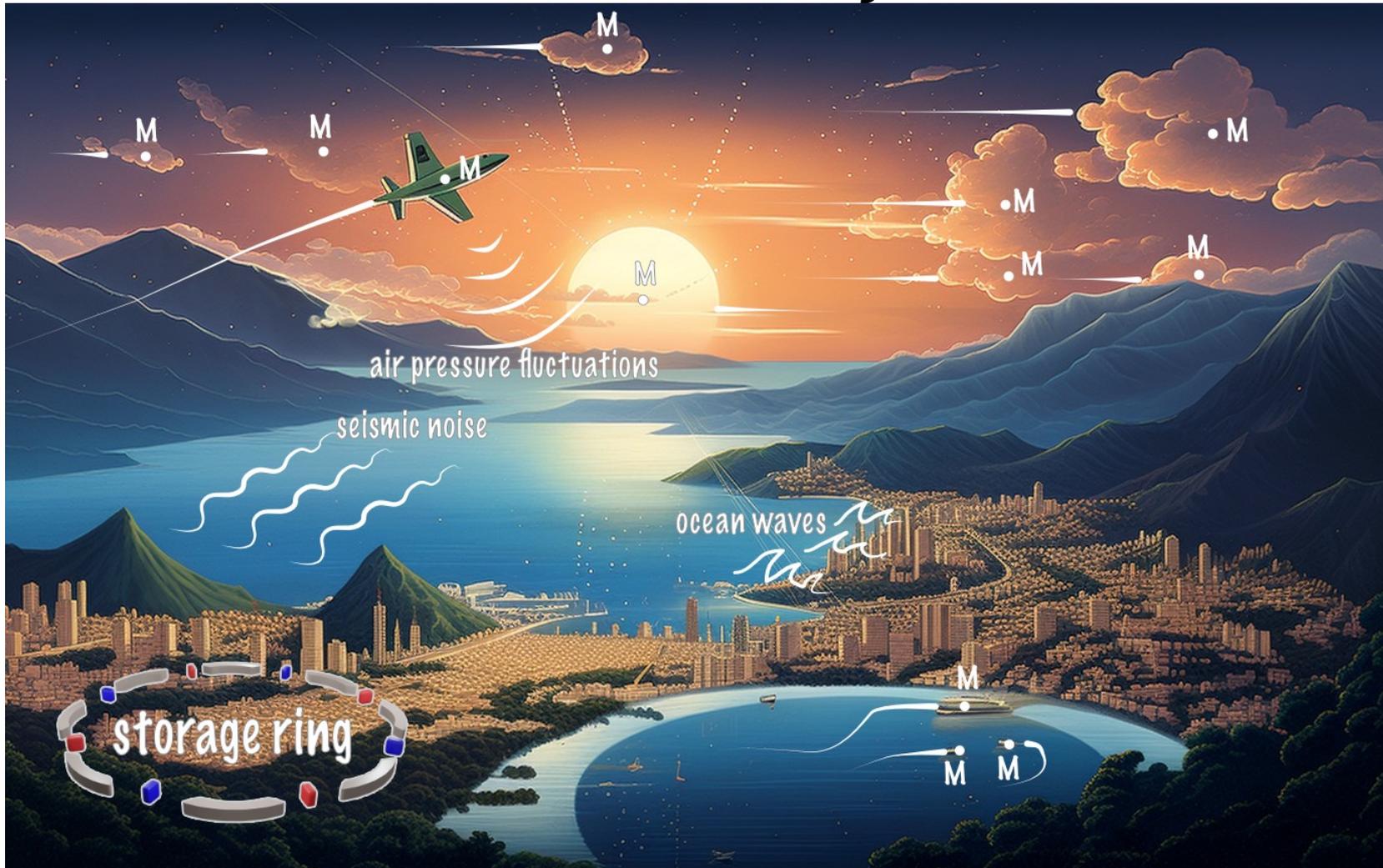
ocean waves

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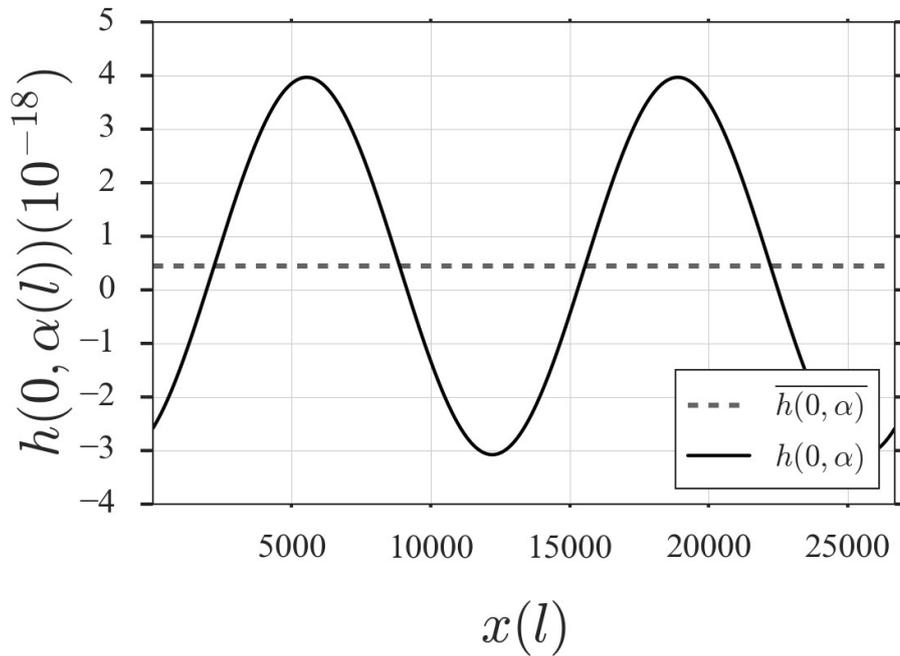


BACKUP

Terrestrial Gravity Noise



Signal Modeling



Chirp mass (red-shift corrected)

$$\mathcal{M} = \frac{(1+z)(m_1 m_2)^{3/5}}{(m_1 + m_2)^{1/5}}$$

Frequency $f = \frac{1}{1+z} \frac{\sqrt{G(m_1 + m_2)}}{\pi r^{3/2}}$

Luminosity distance

d_L inclination angle

$$h_+ = \frac{4}{d_L} \left(\frac{G\mathcal{M}}{c^2} \right)^{5/3} \left(\frac{\pi f}{c} \right)^{2/3} \frac{1 + \cos^2(i)}{2} \cos(2\pi f t + \delta_0)$$

$$h_\times = \frac{4}{d_L} \left(\frac{G\mathcal{M}}{c^2} \right)^{5/3} \left(\frac{\pi f}{c} \right)^{2/3} \cos(i) \sin(2\pi f t + \delta_0)$$

Sampling of Synchrotron Photons I

By parametrizing the surface of a sphere via $\vec{r}(\theta, \phi)$, the function

$$\sigma(\theta, \phi) = \left| \frac{d\vec{r}}{d\theta} \times \frac{d\vec{r}}{d\phi} \right|$$

is computed, used in the definition of two cumulative distribution functions

$$G(s) := \frac{\int_0^{2\pi} \int_0^s \sigma(\theta, \phi) d\theta d\phi}{\int_0^{2\pi} \int_0^\pi \sigma(\theta, \phi) d\theta d\phi},$$

$$H(\theta, t) := \frac{\int_0^t \sigma(\theta, \phi) d\phi}{\int_0^{2\pi} \sigma(\theta, \phi) d\phi}.$$

These functions are next inverted, such that

$$g(s_1) = G^{-1}(s_1),$$

$$h(s_1, s_2) = H^{-1}(s_1, s_2).$$

While this procedure is more general, in the particular case of a sphere it follows that $H^{-1}(s_1, s_2) = \text{Id}(s_2)$. The inverted functions can be used to map a distributions of angles $\{\theta, \phi\} \in [0, \pi] \times [0, 2\pi]$ onto itself by $(g(s_1), h(s_1, s_2)) : [0, \pi] \times [0, 2\pi] \rightarrow [0, \pi] \times [0, 2\pi]$, which turns a stratified sampling of cartesian space into a stratified sampling of the surface of a sphere.

Sampling of Synchrotron Photons II

classically

emitted power:

$$P_\gamma = \frac{2}{3} r_c m c^3 \frac{\beta^4 \gamma^4}{\rho^2}$$

$$u_c = \frac{3 \hbar c \gamma^3}{2 \rho}$$

rate of photons:

$$\dot{N} = \frac{15\sqrt{3}}{8} \frac{P_\gamma}{u_c}$$

$$\dot{N}_{\text{exp}} = \dot{N} \frac{2\pi\rho}{\beta c}$$

Universal Synchrotron spectrum:

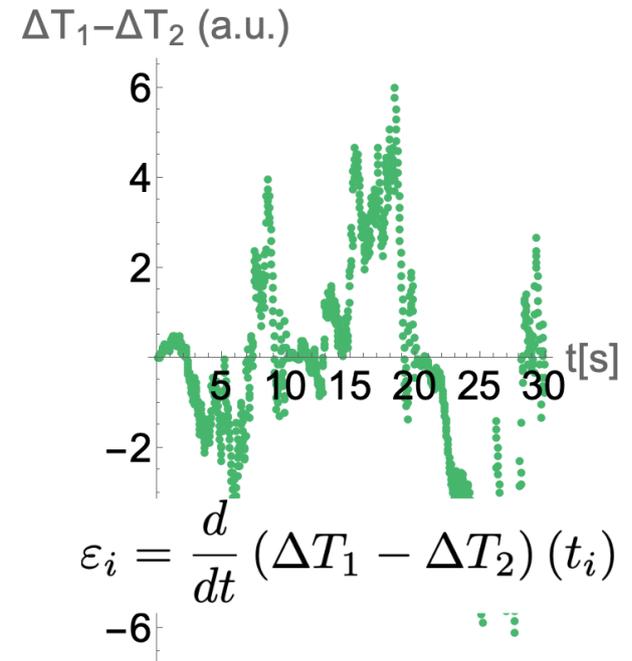
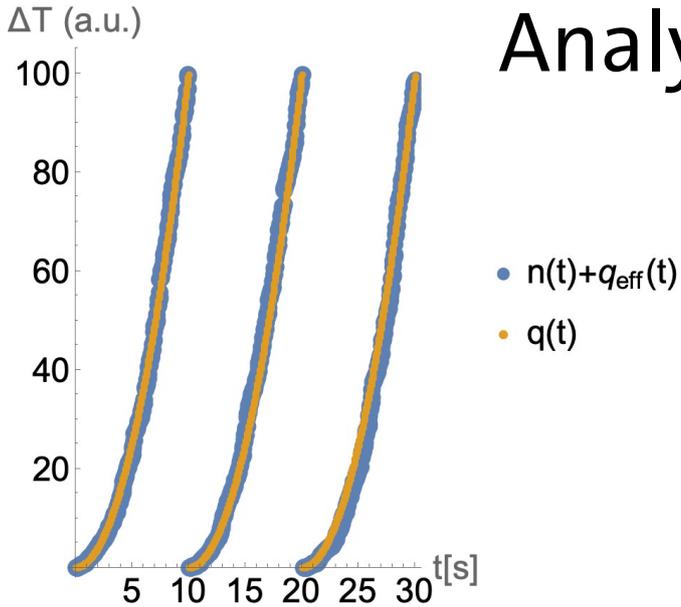
$$S(\xi) = \frac{9\sqrt{3}}{8\pi} \xi \int_\xi^\infty K_{5/3}(x) dx$$

$$n(\xi) = \frac{P_\gamma}{u_c^2} \frac{1}{\xi} S(\xi)$$

Probability distribution:

$$\int_0^\infty u_c \frac{n(\xi)}{\dot{N}_{\text{exp}}} d\xi = 1 \quad \xi \in [10^{-14}, 20]$$

Analytical Noise Models



$$\Delta T_1 = \frac{2}{\Delta t} \sum_n^N \frac{1}{k_n} \sum_j^{k_n} (t - (n-1)\Delta t - t_j) \times$$

$$\Theta(t - (n-1)\Delta t - t_j) \Theta(n\Delta t - t)$$

$\sim q_{\text{eff}}(t) + n_T(t)$ effectively “quadratic”

For many photons (k) the difference becomes zero, if emitted at average t_j !

$$\lim_{k \rightarrow \infty} (\tilde{q}_{\text{eff}}(f) + \tilde{n}_T(f)) \sim \tilde{q}(f)$$

$$\Delta T_2 = \frac{1}{\Delta t^2} \sum_n^N (t - (n-1)\Delta t)^2 \times$$

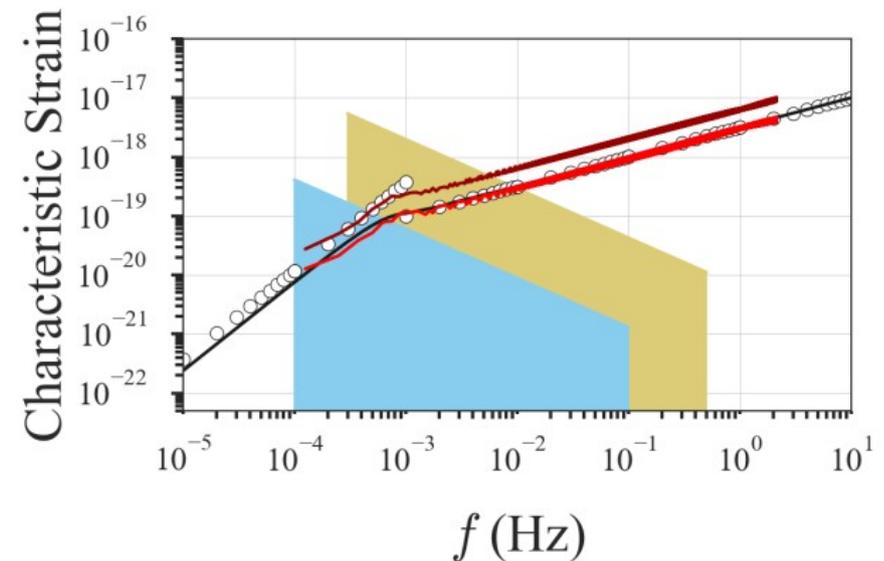
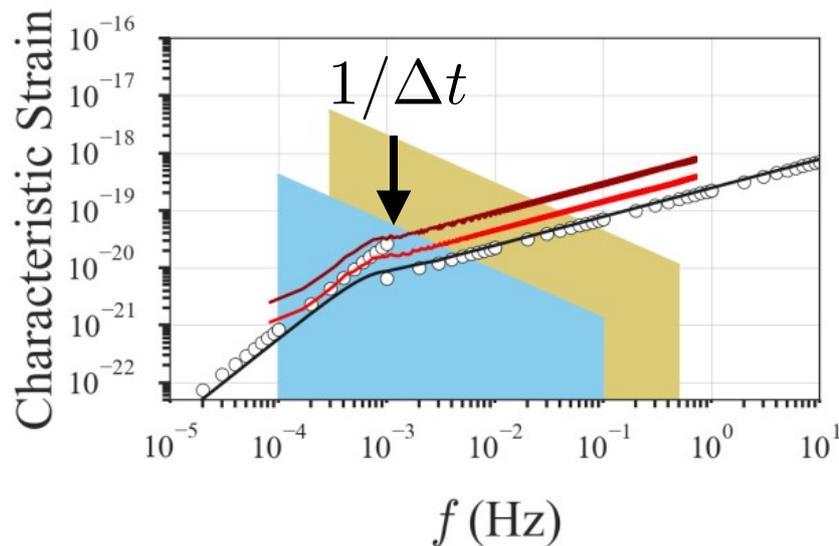
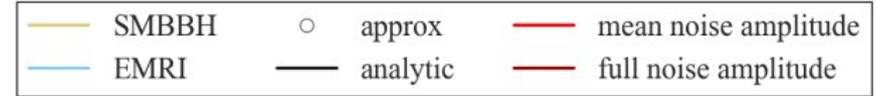
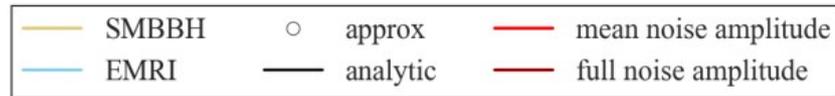
$$\Theta(t - (n-1)\Delta t) \Theta(n\Delta t - t)$$

$\sim q(t)$ “quadratic”

Analytics vs. Numerics

$\beta = 0.32$

$\beta = 0.95$



$$S_n(f) = 4\Delta t S_0^2 \left(\frac{\langle P_{\text{eff}} \rangle}{2p_0 c \gamma^2} \right)^2 \langle \varepsilon \rangle^2 \begin{cases} (2\pi f \Delta t)^2 / 4 & f < 1/\Delta t \\ 1 & f > 1/\Delta t \end{cases}$$

Geodesic Equations

$$\frac{d^2 t}{d\tau^2} + \frac{1}{2} \frac{dh_{\theta\phi\psi}(t)}{dt} \left(\frac{dl}{d\tau} \right)^2 = 0,$$

$$\frac{d^2 l}{d\tau^2} + \frac{1}{1 + h_{\theta\phi\psi}(t)} \frac{dh_{\theta\phi\psi}(t)}{dt} \left(\frac{dl}{d\tau} \right) \left(\frac{dt}{d\tau} \right) = 0,$$

$$- \left(\frac{dt}{d\tau} \right)^2 + (1 + h_{\theta\phi\psi}(t)) \left(\frac{dl}{d\tau} \right)^2 = g_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = -1.$$

$$h_{\theta\phi\psi}(t, \alpha) = h_+(t) \left(f_s^+ \sin^2 \alpha + f_c^+ \cos^2 \alpha + f_{sc}^+ \sin 2\alpha \right) \\ + h_\times(t) \left(f_s^\times \sin^2 \alpha + f_c^\times \cos^2 \alpha + f_{sc}^\times \sin 2\alpha \right)$$

$$f_s^+ = (\cos^2 \theta \cos^2 \phi - \sin^2 \phi) \cos 2\psi - (\cos \theta \sin 2\phi) \sin 2\psi,$$

$$f_c^+ = (\cos^2 \theta \sin^2 \phi - \cos^2 \phi) \cos 2\psi + (\cos \theta \sin 2\phi) \sin 2\psi,$$

$$f_{sc}^+ = \left(\frac{1}{2} (1 + \cos^2 \theta) \sin 2\phi \right) \cos 2\psi + (\cos \theta \cos 2\phi) \sin 2\psi,$$

$$f_s^\times = (\cos^2 \theta \cos^2 \phi - \sin^2 \phi) \sin 2\psi + (\cos \theta \sin 2\phi) \cos 2\psi,$$

$$f_c^\times = (\cos^2 \theta \sin^2 \phi - \cos^2 \phi) \sin 2\psi - (\cos \theta \sin 2\phi) \cos 2\psi,$$

$$f_{sc}^\times = \left(\frac{1}{2} (1 + \cos^2 \theta) \sin 2\phi \right) \sin 2\psi - (\cos \theta \cos 2\phi) \cos 2\psi.$$