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# SRF for GW detection: MAGO

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**University of Geneva**

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# Conclusion This Morning

Identify regime of GW by hierarchy with respect to size of detector:

- Resonant regime:  $\omega_g \sim 1/L \gg c_s/L$  — use PDF and account for current & boundary changes

*e.g. axion cavity experiments* Berlin, Blas, D'Agnolo, SARE, Harnik, Kahn, Schutte-Engel (2021)

- High-frequency regime:  $\omega_g \gg 1/L \gg c_s/L$  — use TT gauge

*e.g. MADMAX* Domcke, Ellis, Kopp (2024)

- Low-frequency regime:  $1/L \gg \omega_g$  — use PDF and account for current & boundary changes

*e.g. Magneto-quasistatic experiments* Domcke, Garcia-Cely, Rodd (2022)

*e.g. Heterodyne experiments* Domcke, Garcia-Cely, Lee, Rodd (2023)  
Domcke, Ellis, Rodd (2024)  
Berlin, Blas, D'Agnolo, SARE, Harnik, Kahn, Schutte-Engel, Wentzel (2023)

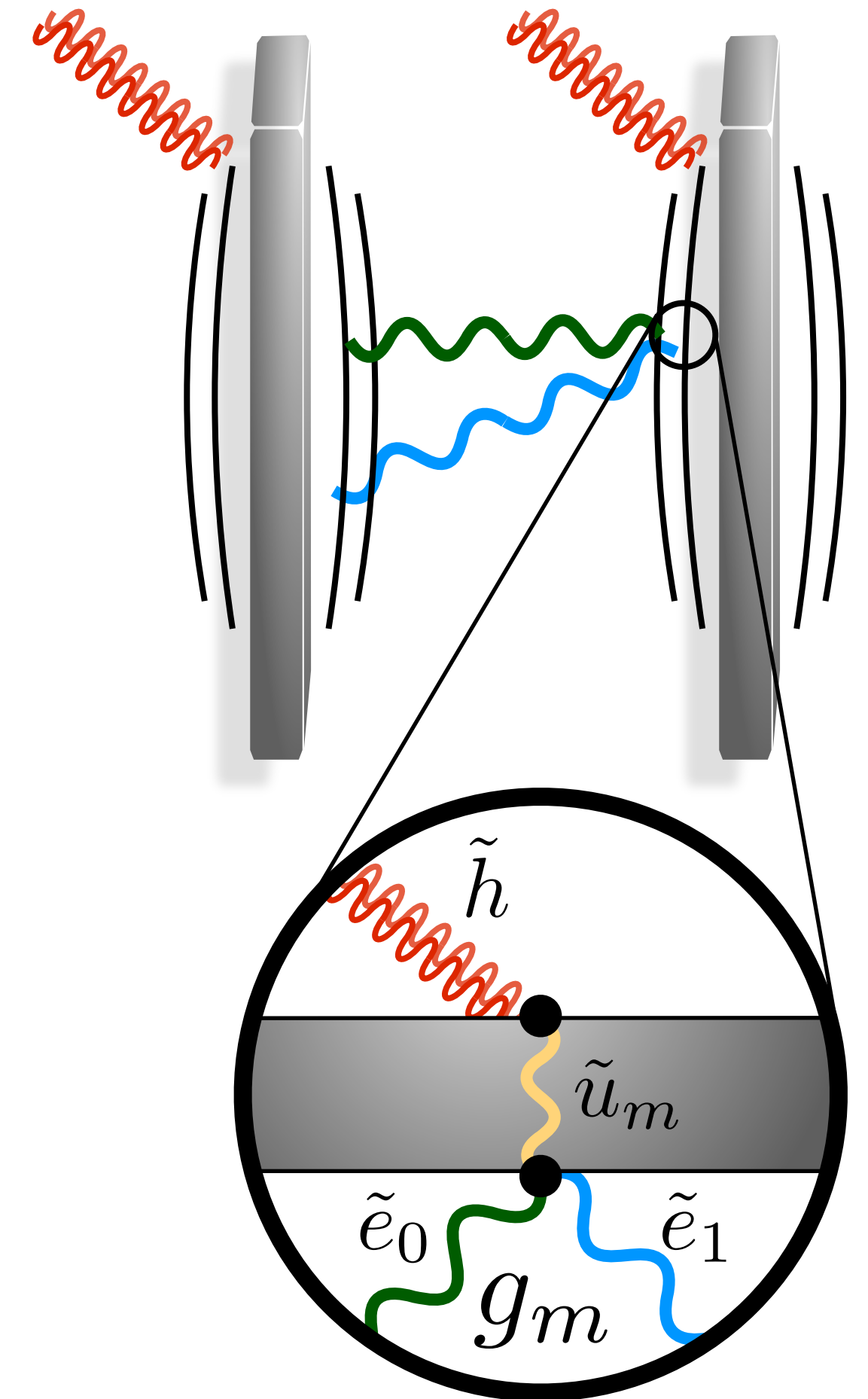
# Transfer function for mechanical transduction

$$\left( \omega_m^2 - \omega^2 + i \frac{\omega \omega_m}{Q_m} \right) \tilde{u}_m(\omega) \simeq -\frac{\omega_g^2 L}{2} \tilde{h}^{\text{TT}}(\omega)$$

$$\left( \omega_1^2 - \omega^2 + i \frac{\omega \omega_1}{Q} \right) \tilde{e}_1(\omega) \simeq \int d\omega' \tilde{e}_0(\omega - \omega') g_m \tilde{u}_m(\omega')$$

$$g_m \equiv -\frac{2\omega_1^2}{L}$$

$$\mathcal{T}_{\text{mech}}^2(\omega) = \frac{\omega_g^4 \omega_1^4}{\left( (\omega_1^2 - \omega^2)^2 + \frac{\omega^2 \omega_1^2}{Q^2} \right) \left( (\omega_m^2 - \omega^2)^2 + \frac{\omega_g^2 \omega_m^2}{Q_m^2} \right)}$$

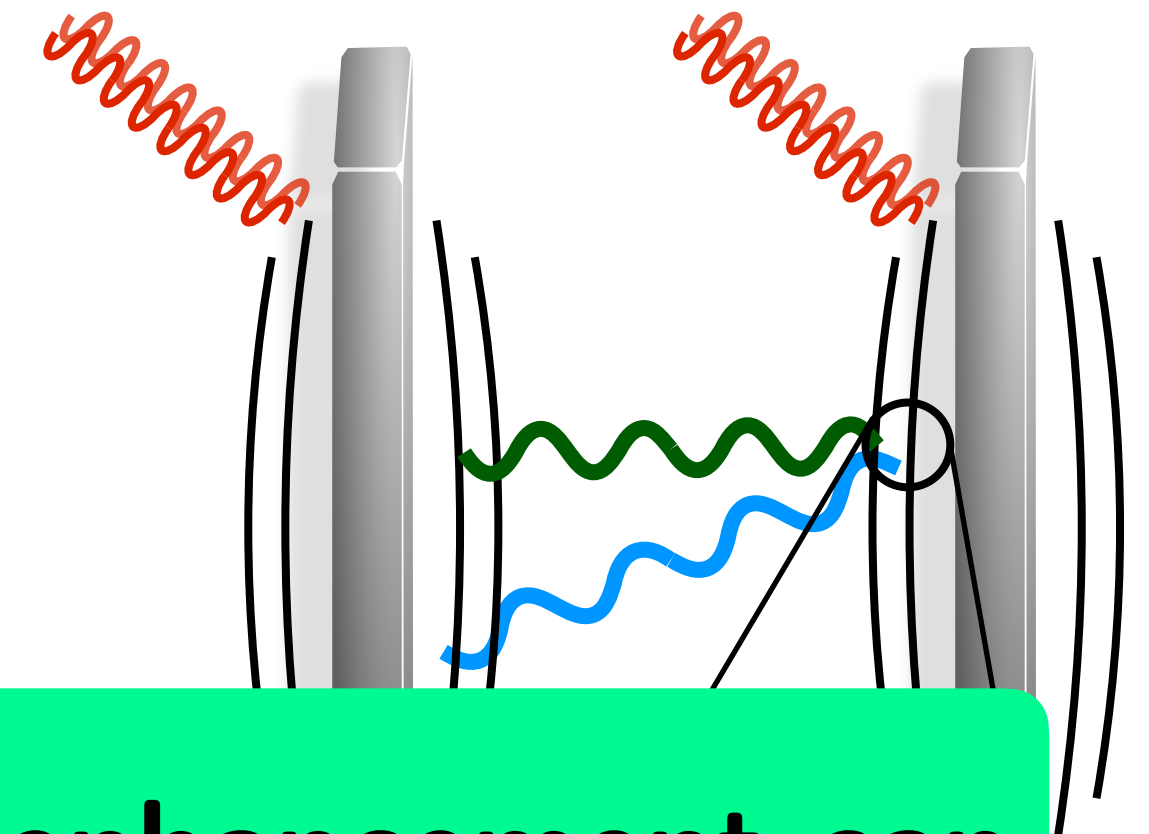


*D'Agnolo, SARE (gr-qc/2412.17897)*

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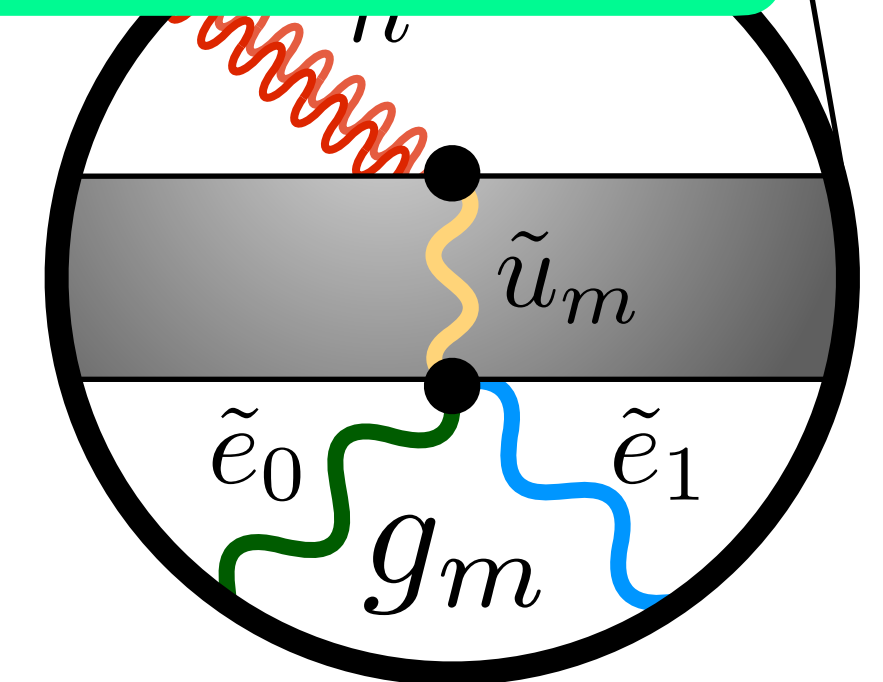
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Is there a compact physical system where EM resonant enhancement can be realised?

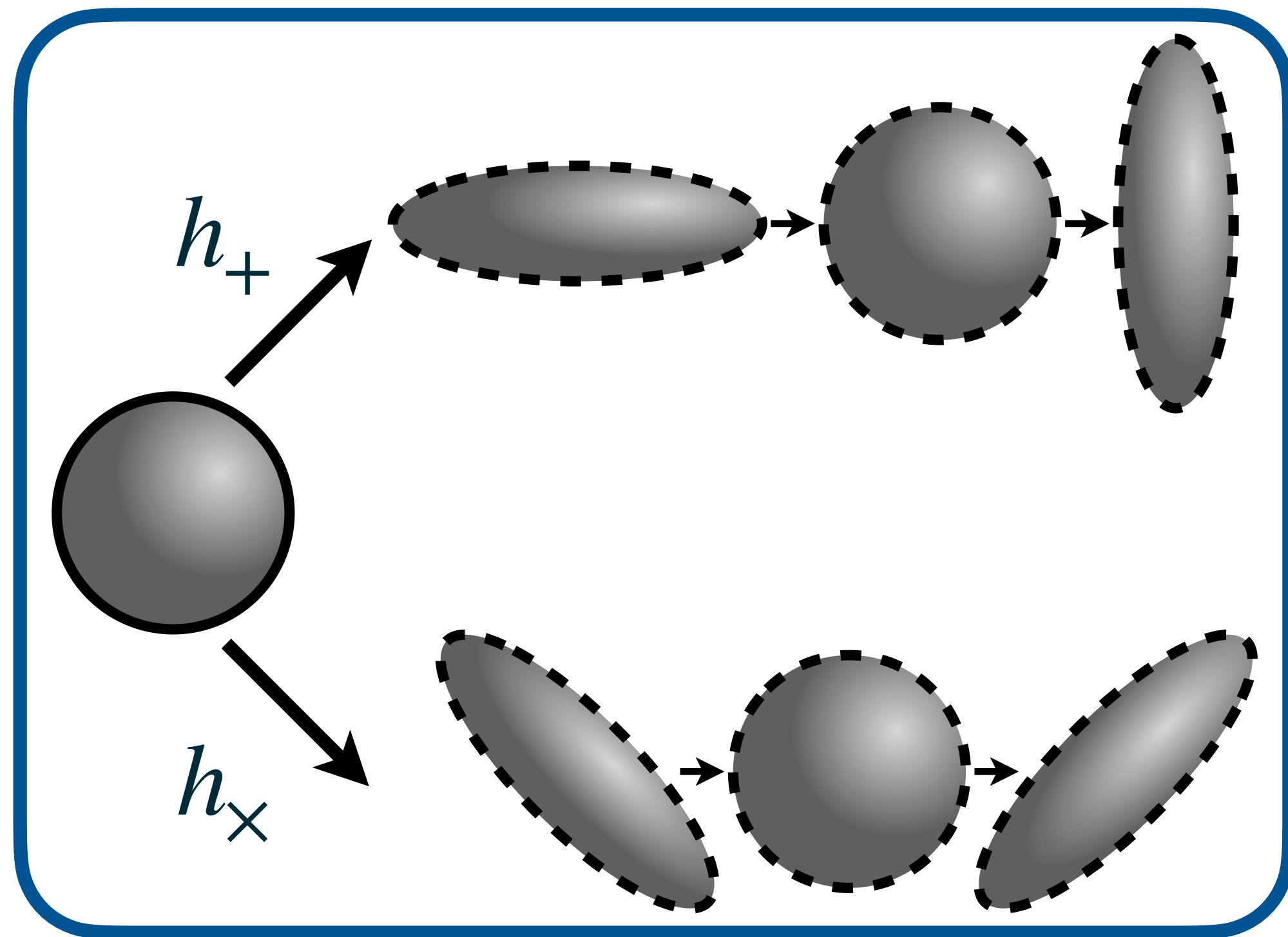
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*D'Agnolo, SARE (gr-qc/2412.17897)*



# Gravitational Wave and a Hollow Sphere

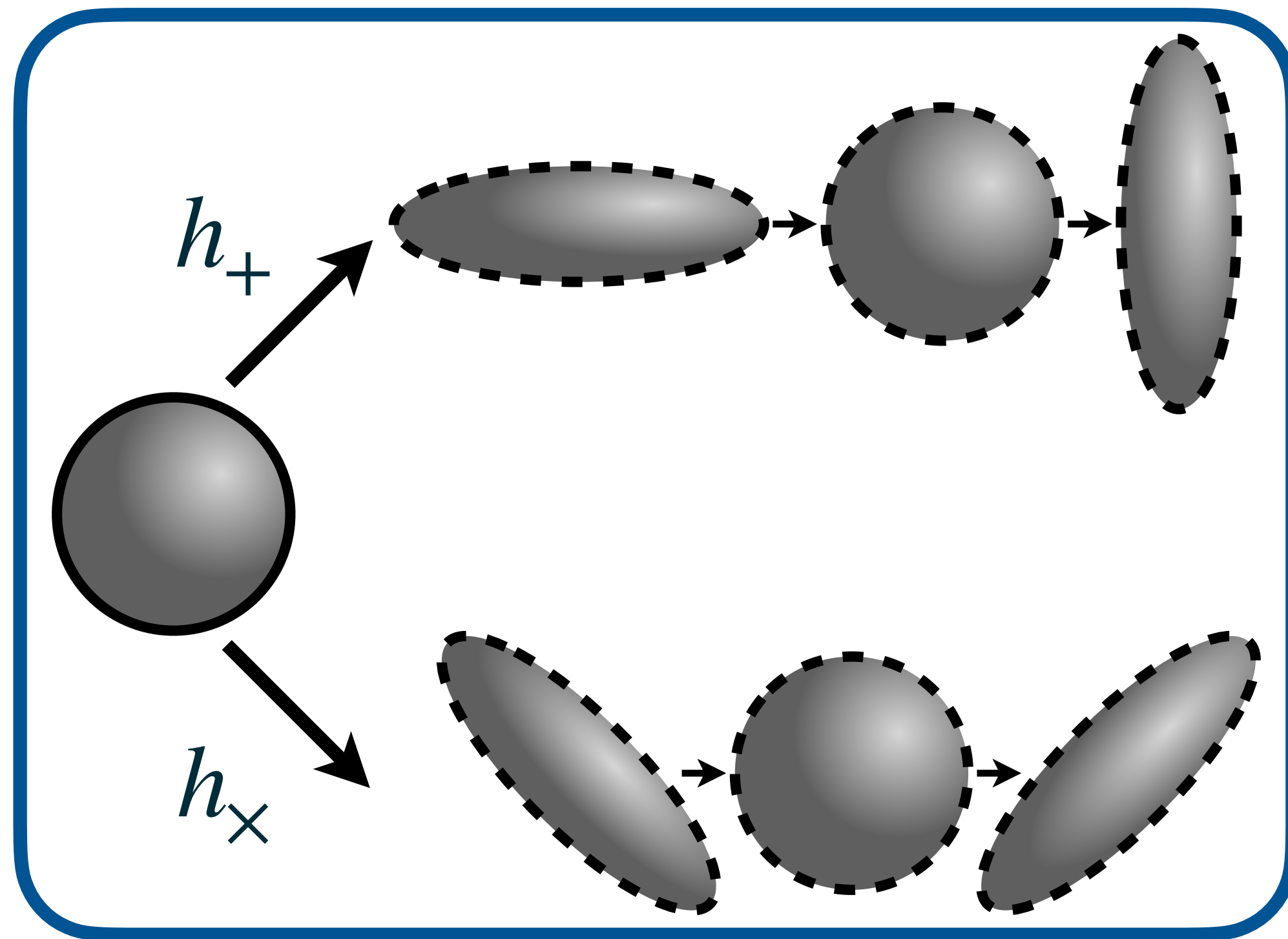


TT frame intuition

# Gravitational Wave and a Hollow Sphere

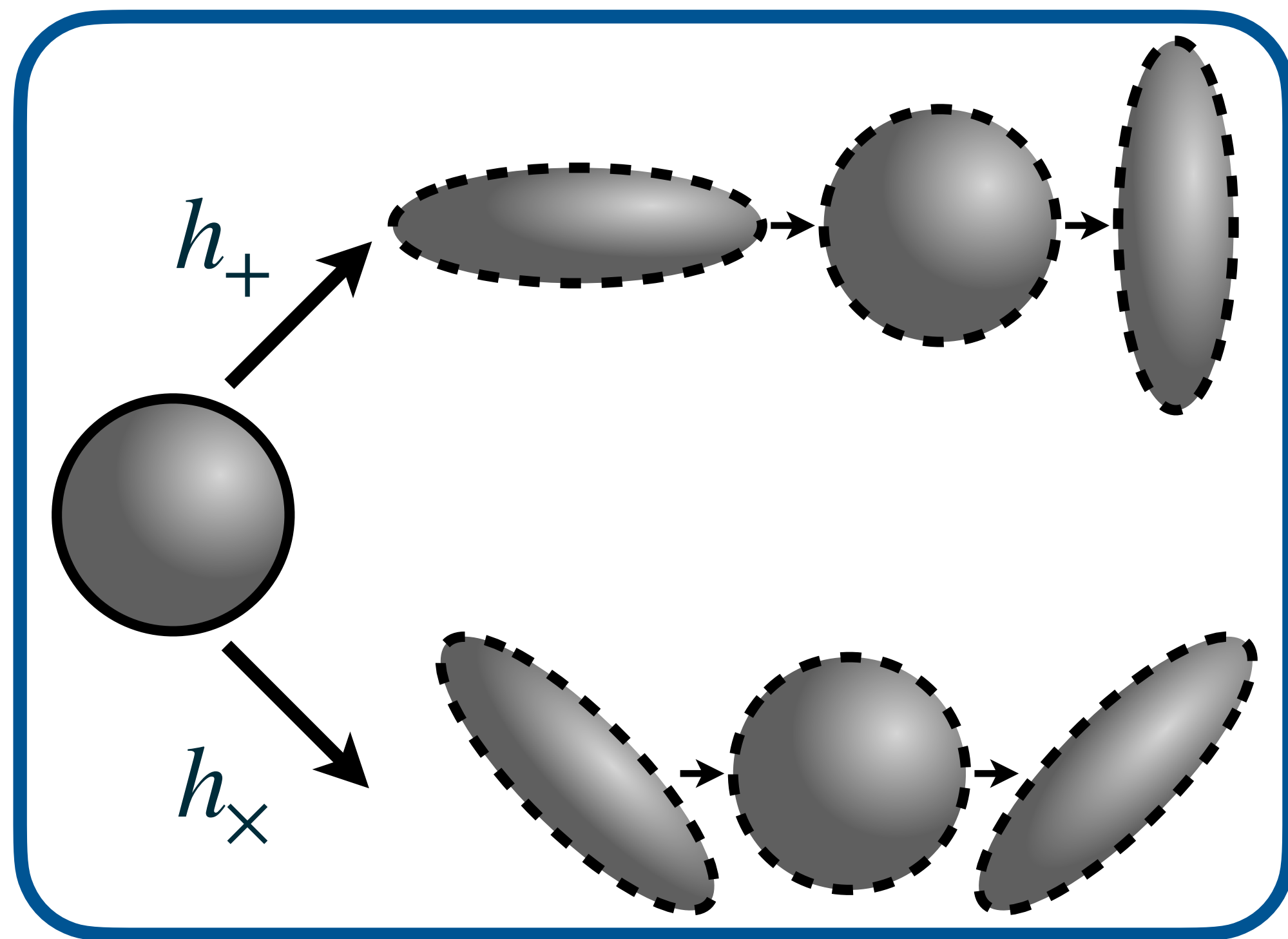
Mechanical modes of a sphere

$$\mathbf{U}_{lmn} = \nabla\phi_L + i\nabla \times \mathbf{L}\phi_{T_1} + i\mathbf{L}\phi_{T_2} .$$



TT frame intuition

# Gravitational Wave and a Hollow Sphere

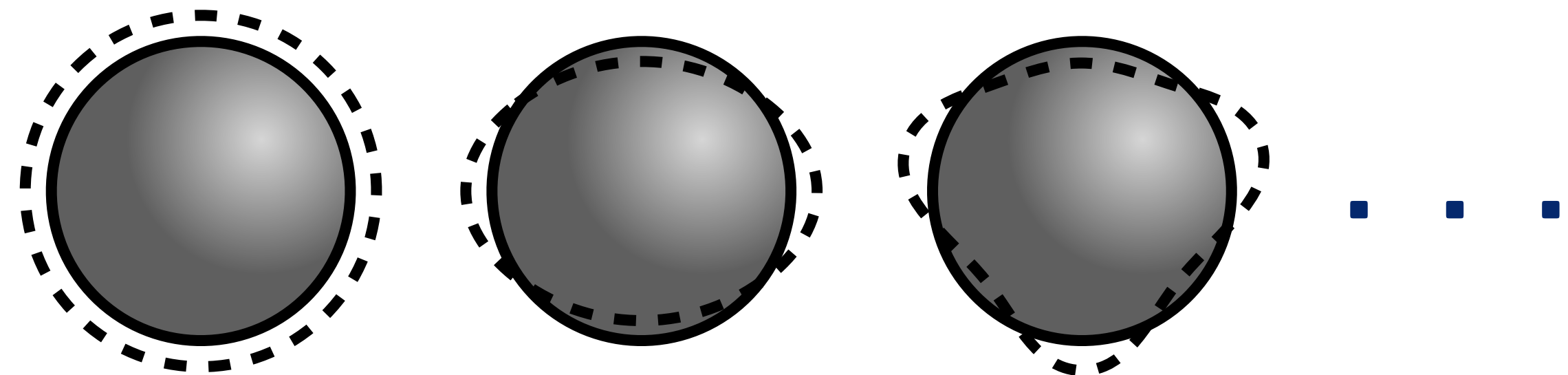


TT frame intuition

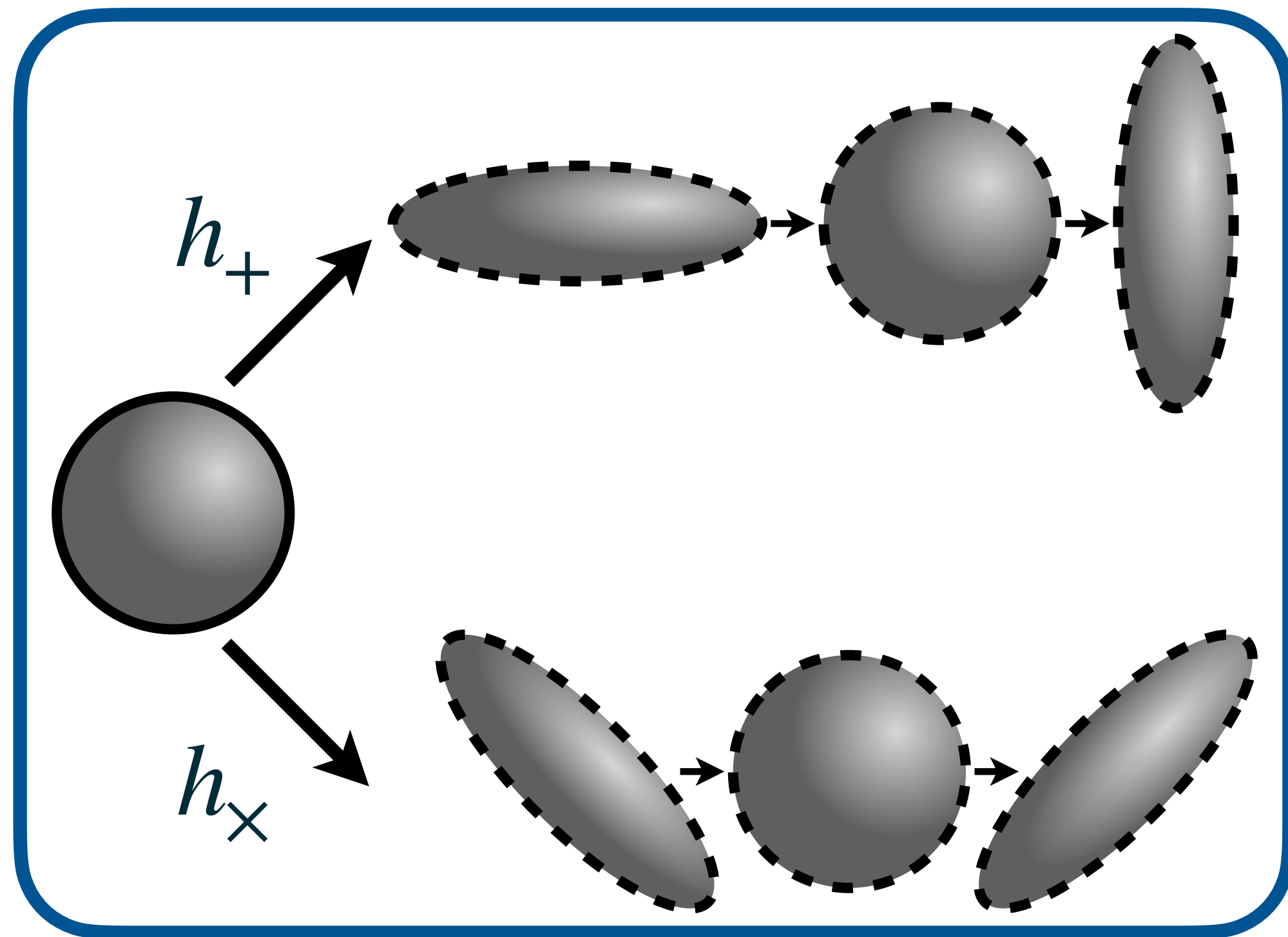
## Mechanical modes of a sphere

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Spheroidal



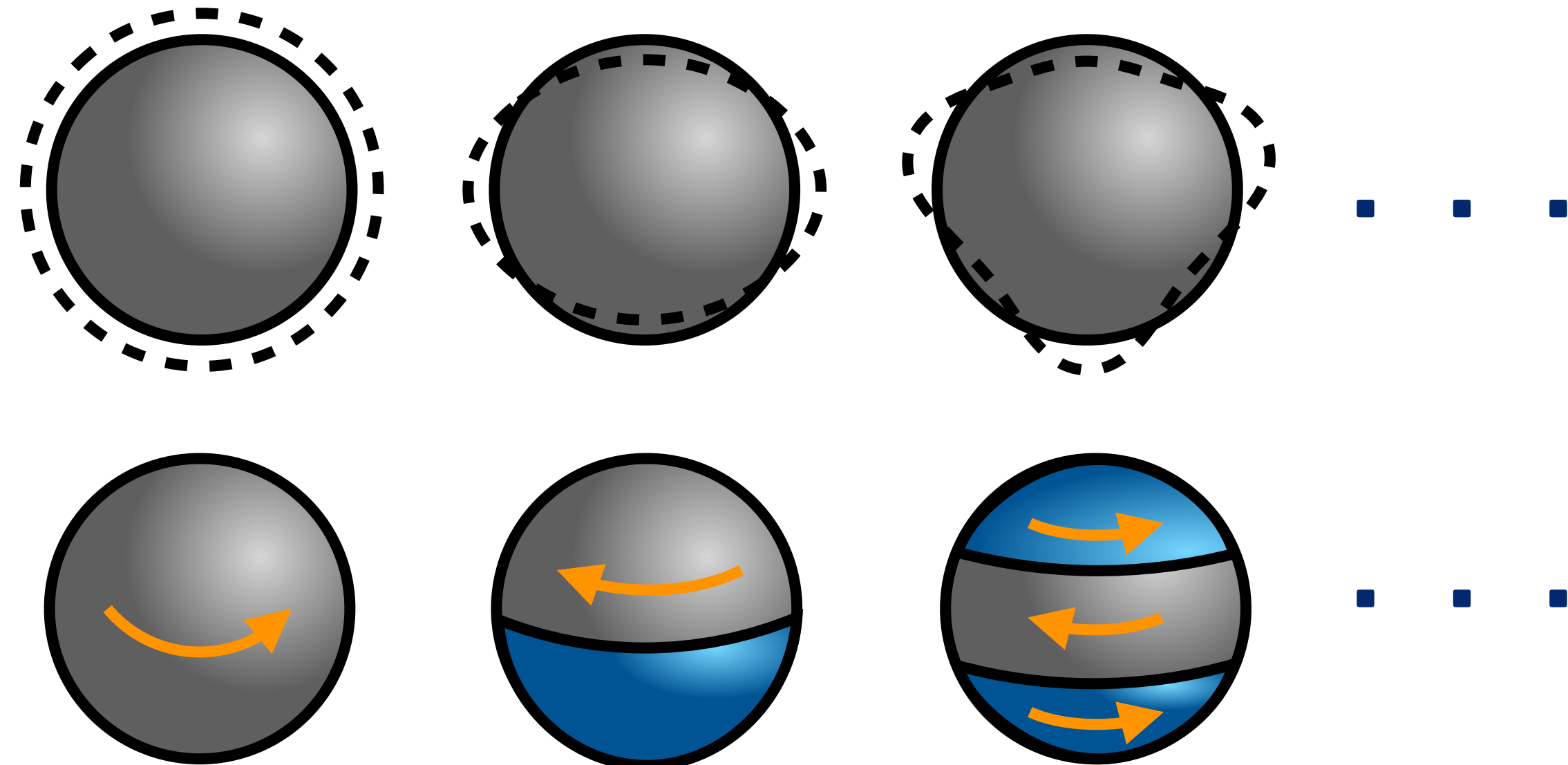
# Gravitational Wave and a Hollow Sphere



TT frame intuition

## Mechanical modes of a sphere

$$U_{lmn} = \underbrace{\nabla\phi_L + i\nabla \times \mathbf{L}\phi_{T_1}}_{\text{Spheroidal}} + \underbrace{i\mathbf{L}\phi_{T_2}}_{\text{Toroidal}}.$$



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# Gravitational Wave and a Hollow Sphere

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# Gravitational Wave and a Hollow Sphere

## Mechanical modes of a sphere

$$\mathbf{U}_{lmn} = \nabla\phi_L + i\nabla \times \mathbf{L}\phi_{T_1} + i\mathbf{L}\phi_{T_2} .$$

$$\mathbf{U}(\mathbf{x}, t) = u_p(t) \mathbf{U}_p(\mathbf{x})$$



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## Equation of motion

$$\ddot{u}_p + \frac{\omega_p}{Q_p} \dot{u}_p + \omega_p^2 u_p \simeq -\frac{1}{2} \omega_g^2 V_{\text{cav}}^{1/3} \eta_{\text{mech}}^g h_0 e^{i\omega_g t}$$

$$\eta_{\text{mech}}^g = \frac{\hat{h}_{ij}^{TT}}{V_{\text{cav}}^{1/3} V_{\text{shell}}} \int_{V_{\text{shell}}} d^3\mathbf{x} U_p^{*i} x^j$$

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$$\langle \mathbf{U}_p \rangle \sim h_0 V_{\text{cav}}^{1/3} \eta_{\text{mech}}^g \times \begin{cases} \frac{\omega_g^2}{\omega_g^2 - \omega_p^2} , & |\omega_g - \omega_p| \gg \omega_p / Q_p \\ Q_p , & |\omega_g - \omega_p| \ll \omega_p / Q_p \end{cases}$$

**Tiny displacement  $\ll$  nm**

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# MAGO 2.0: Mechanical and EM Signals

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**On the operation of a tunable electromagnetic detector for gravitational waves**

F Pegoraro<sup>†</sup>, E Picasso<sup>‡</sup> and L A Radicati<sup>‡§</sup>

<sup>†</sup>Scuola Normale Superiore, Pisa, Italy

<sup>‡</sup>CERN, Geneva, Switzerland

Received 6 December 1977, in final form 20 April 1978

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### Microwave Apparatus for Gravitational Waves Observation

R. Ballantini, A. Chincarini, S. Cuneo, G. Gemme<sup>\*</sup>, R. Parodi, A. Podestà, and R. Vaccarone  
*INFN and Università degli Studi di Genova, Genova, Italy*

Ph. Bernard, S. Calatroni, E. Chiaveri, and R. Losito  
*CERN, Geneva, Switzerland*

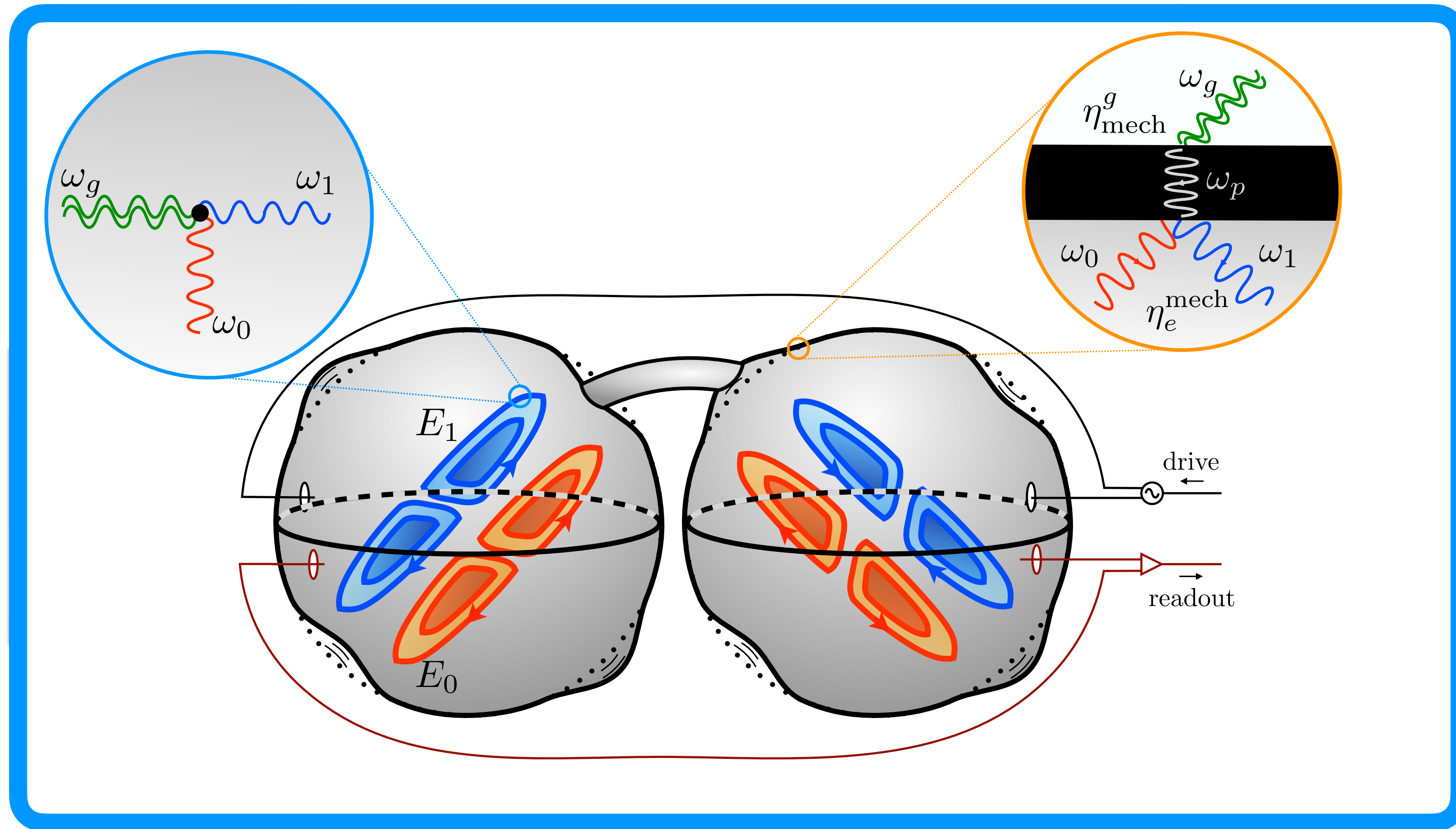
R.P. Croce, V. Galdi, V. Pierro, and I.M. Pinto  
*INFN, Napoli, and Università degli Studi del Sannio, Benevento, Italy*

E. Picasso  
*INFN and Scuola Normale Superiore, Pisa, Italy and  
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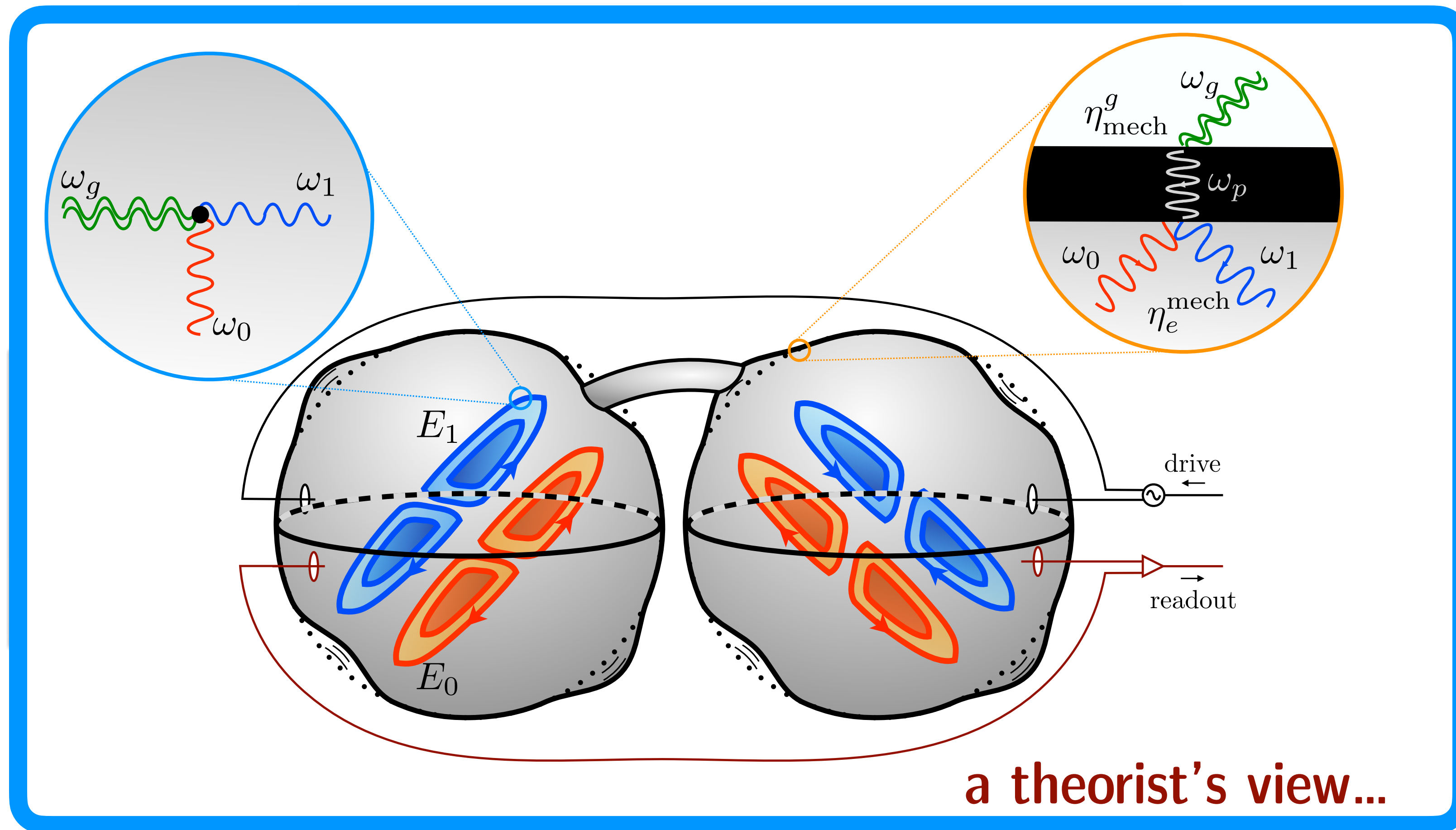


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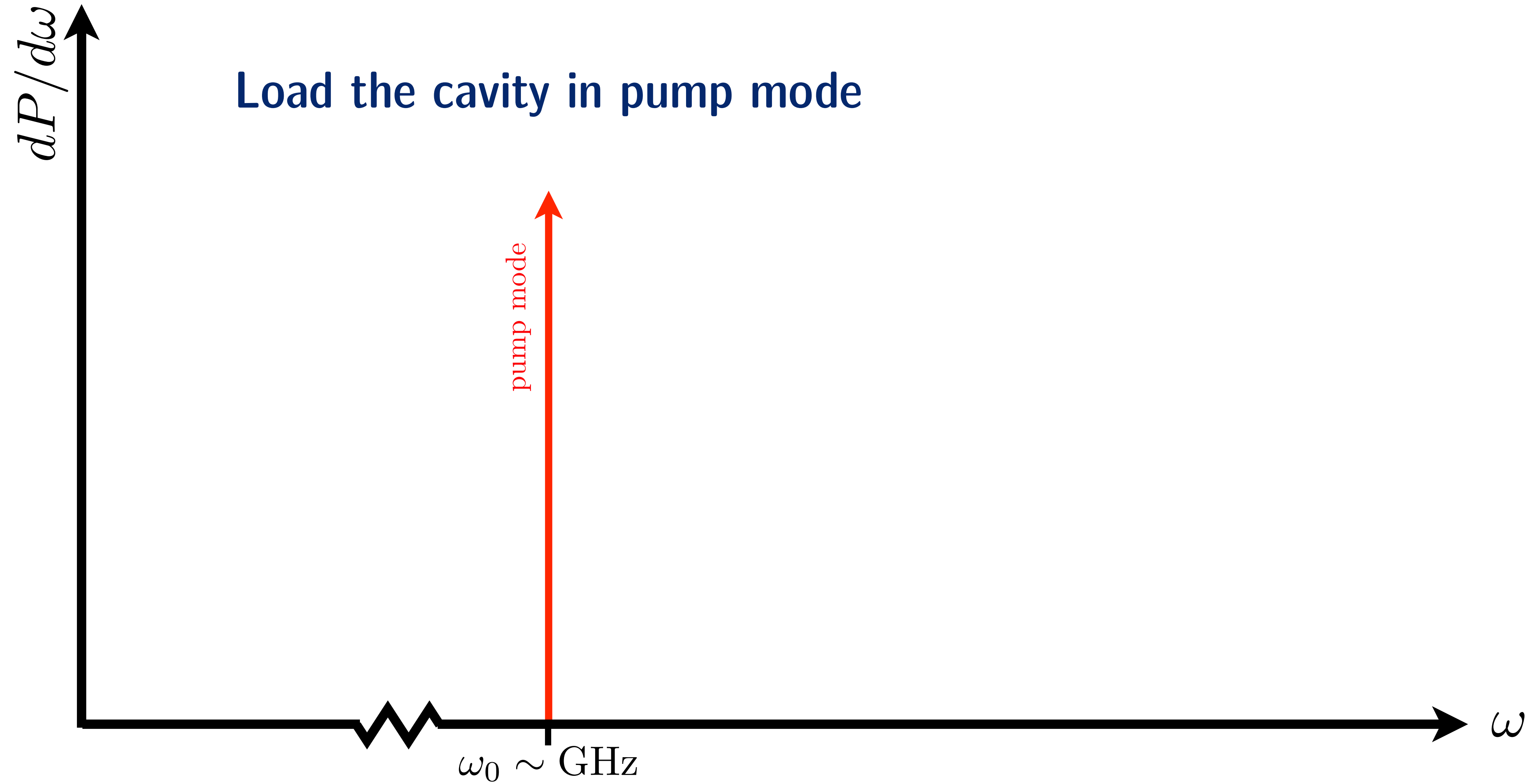




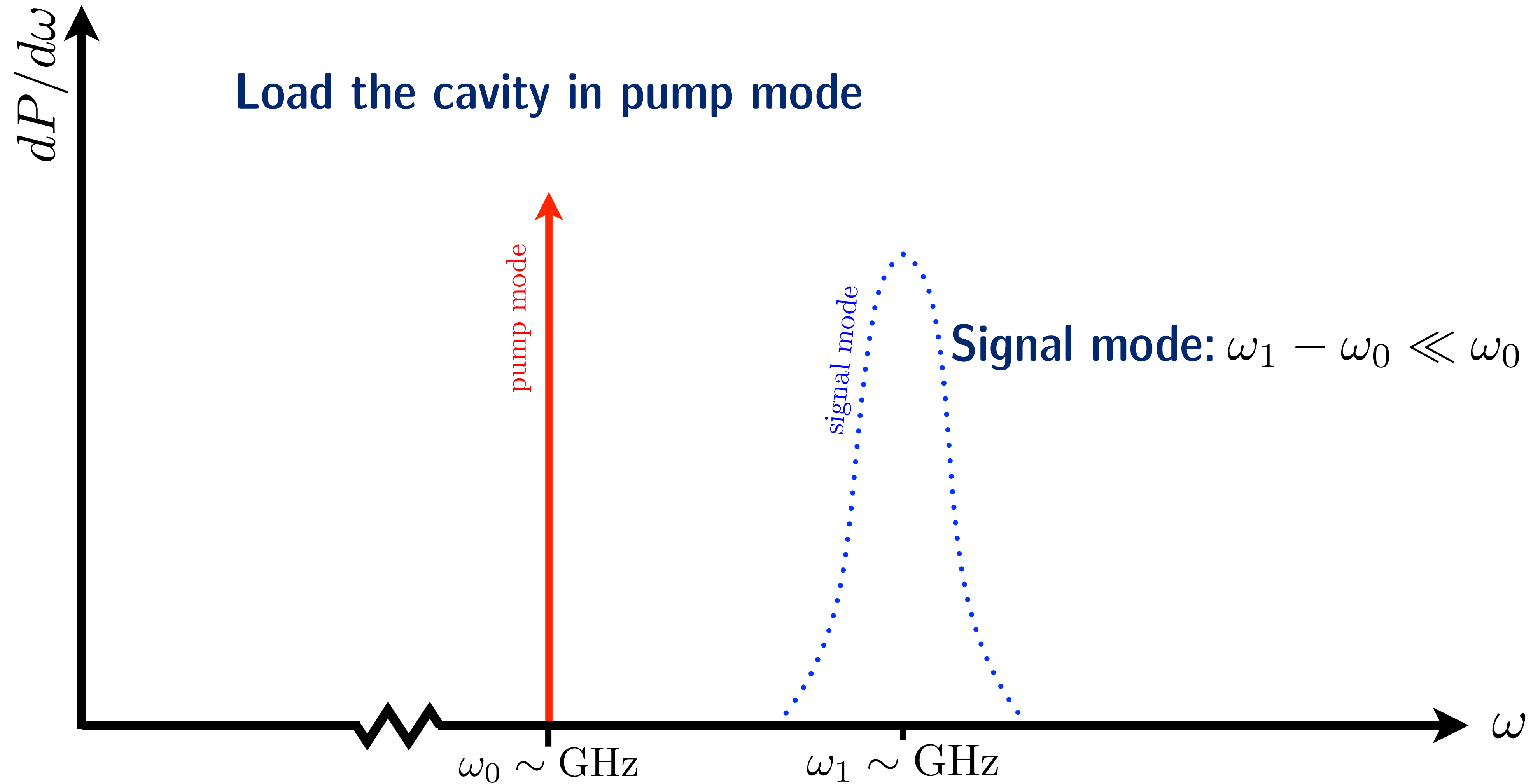
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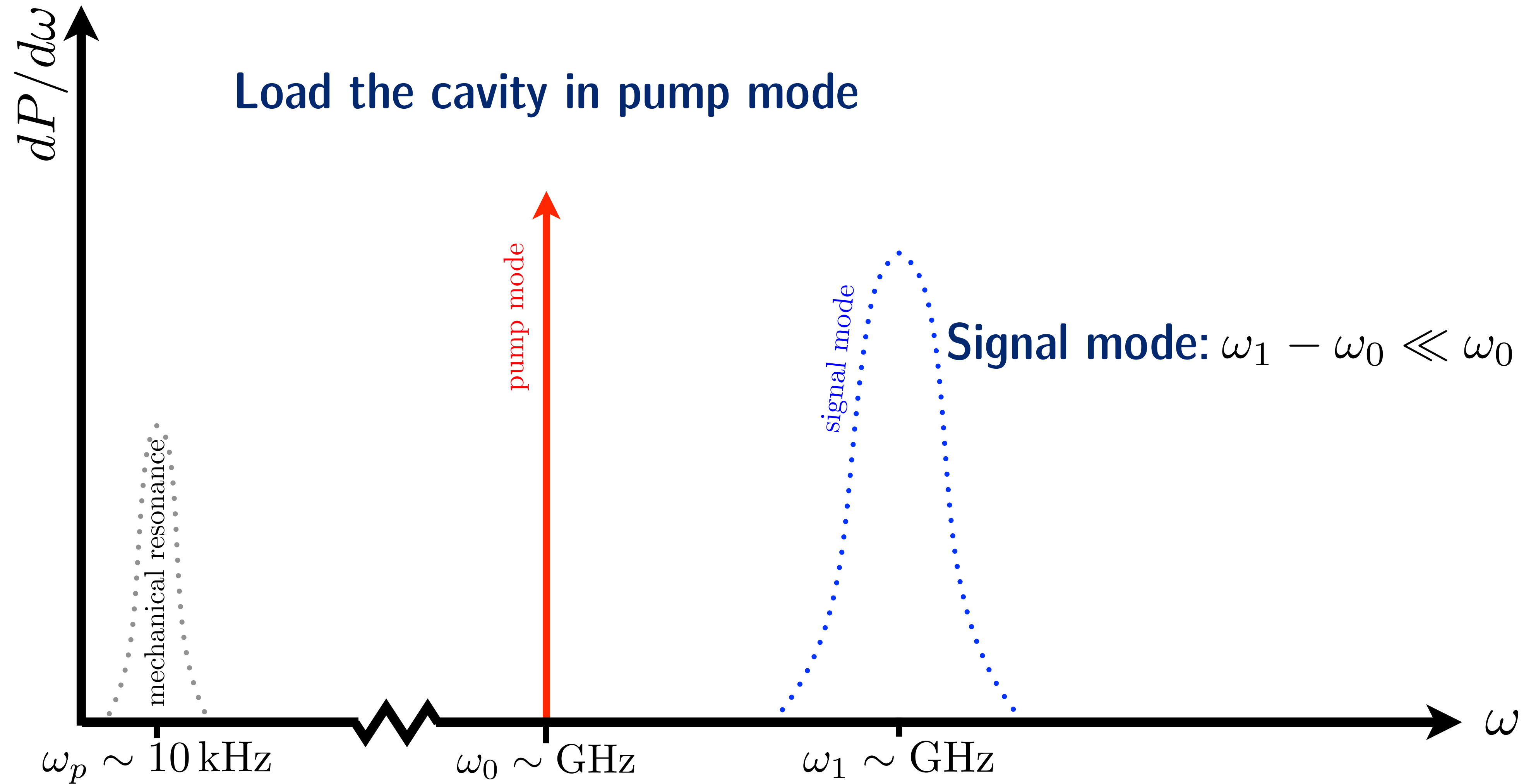
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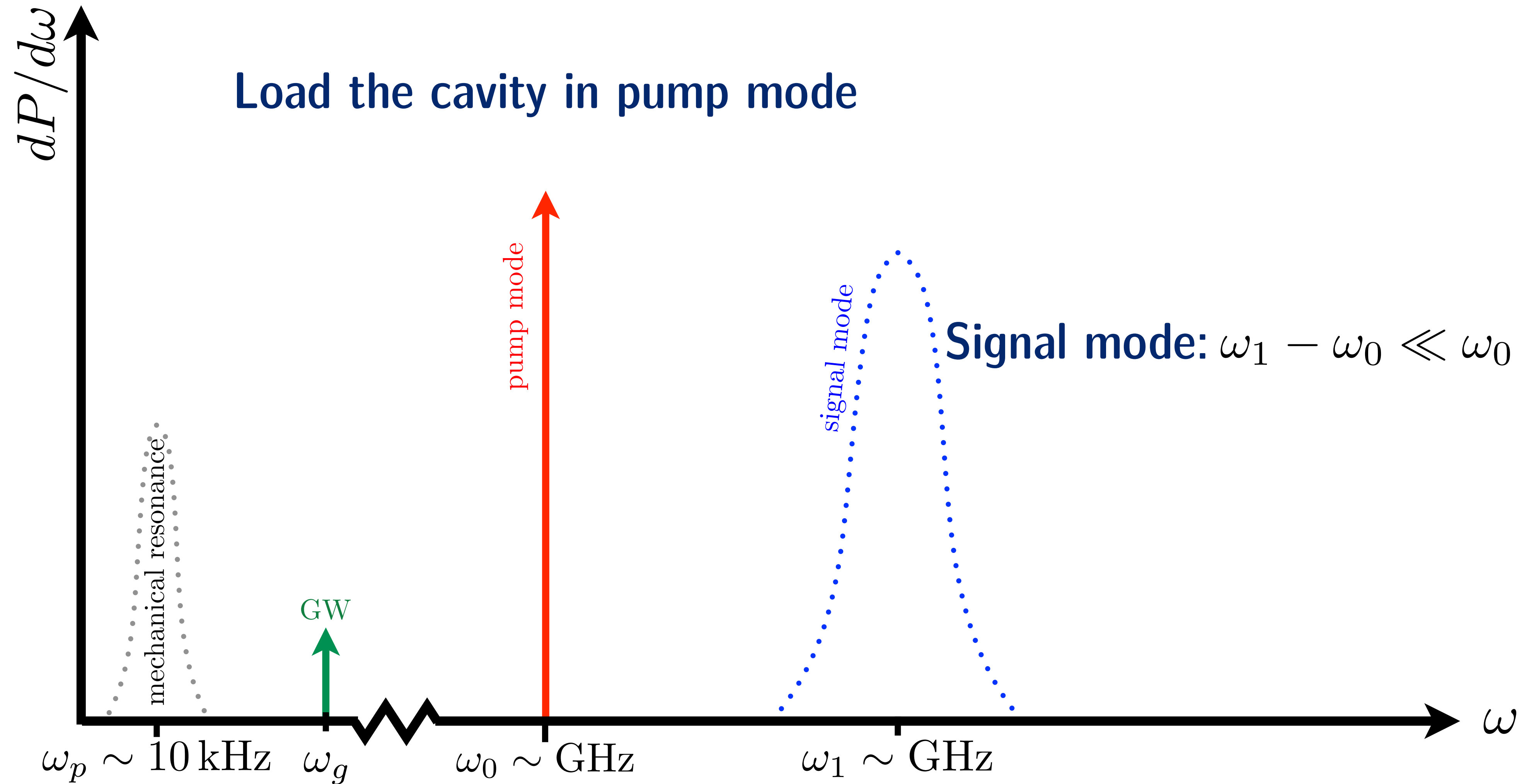
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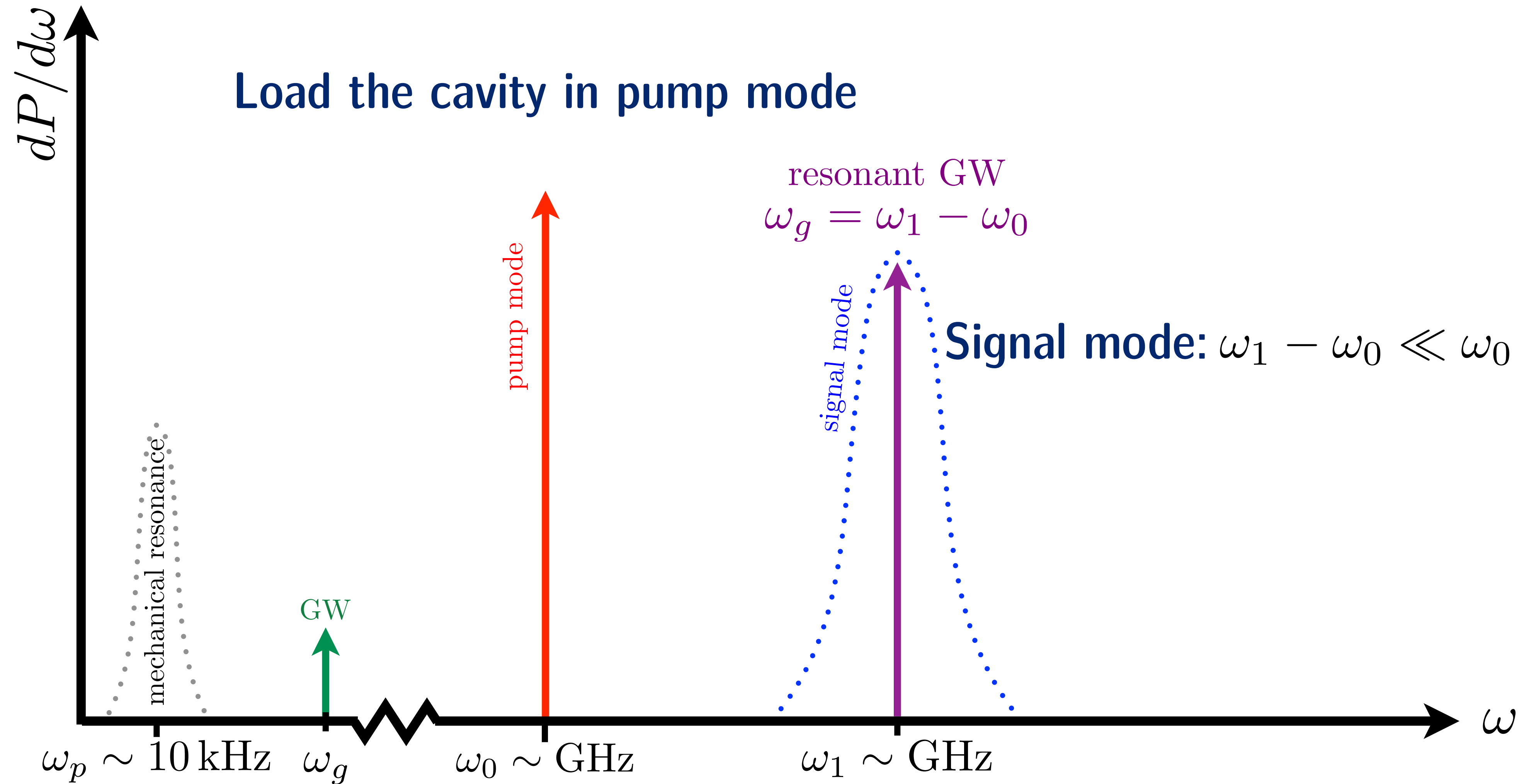
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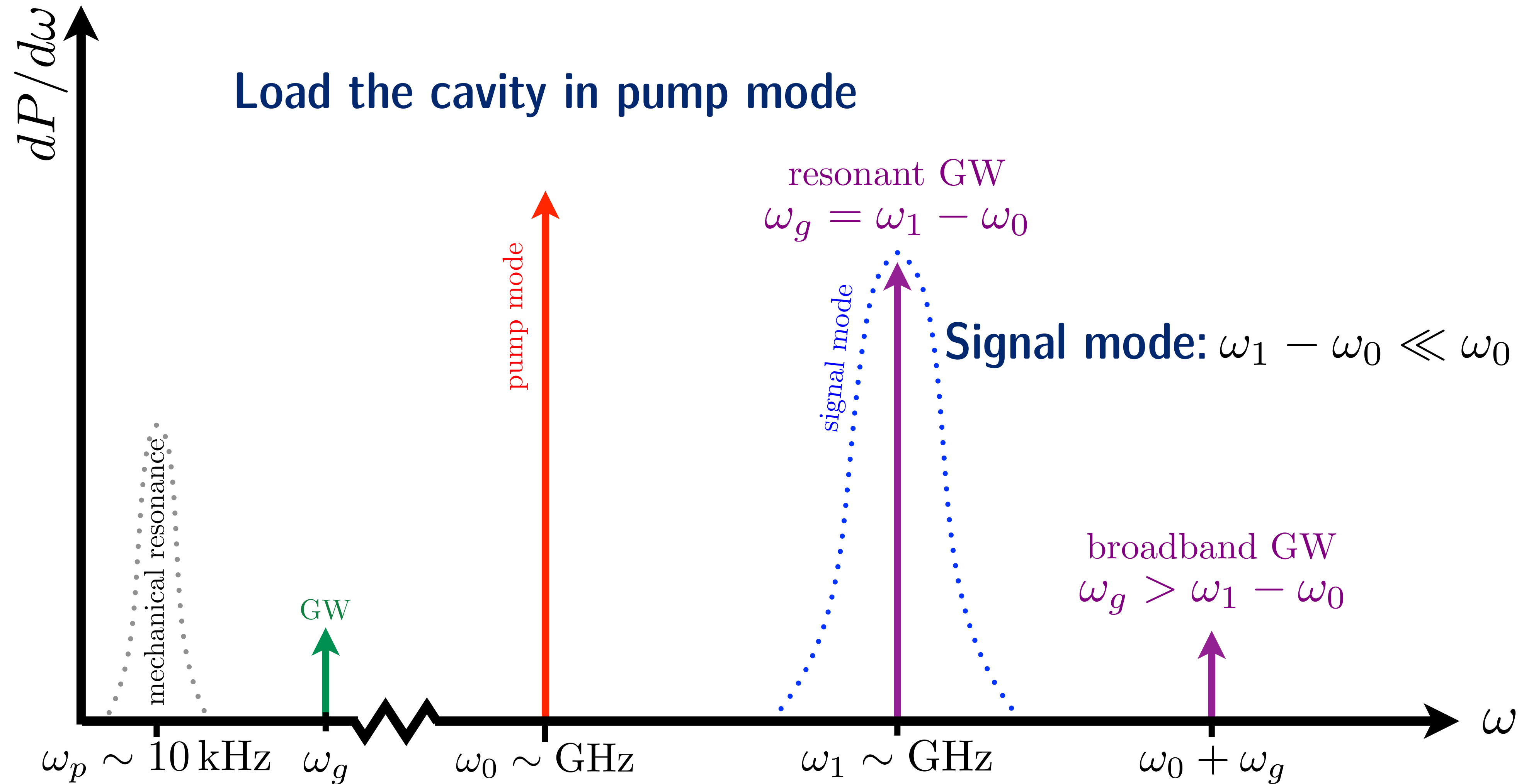


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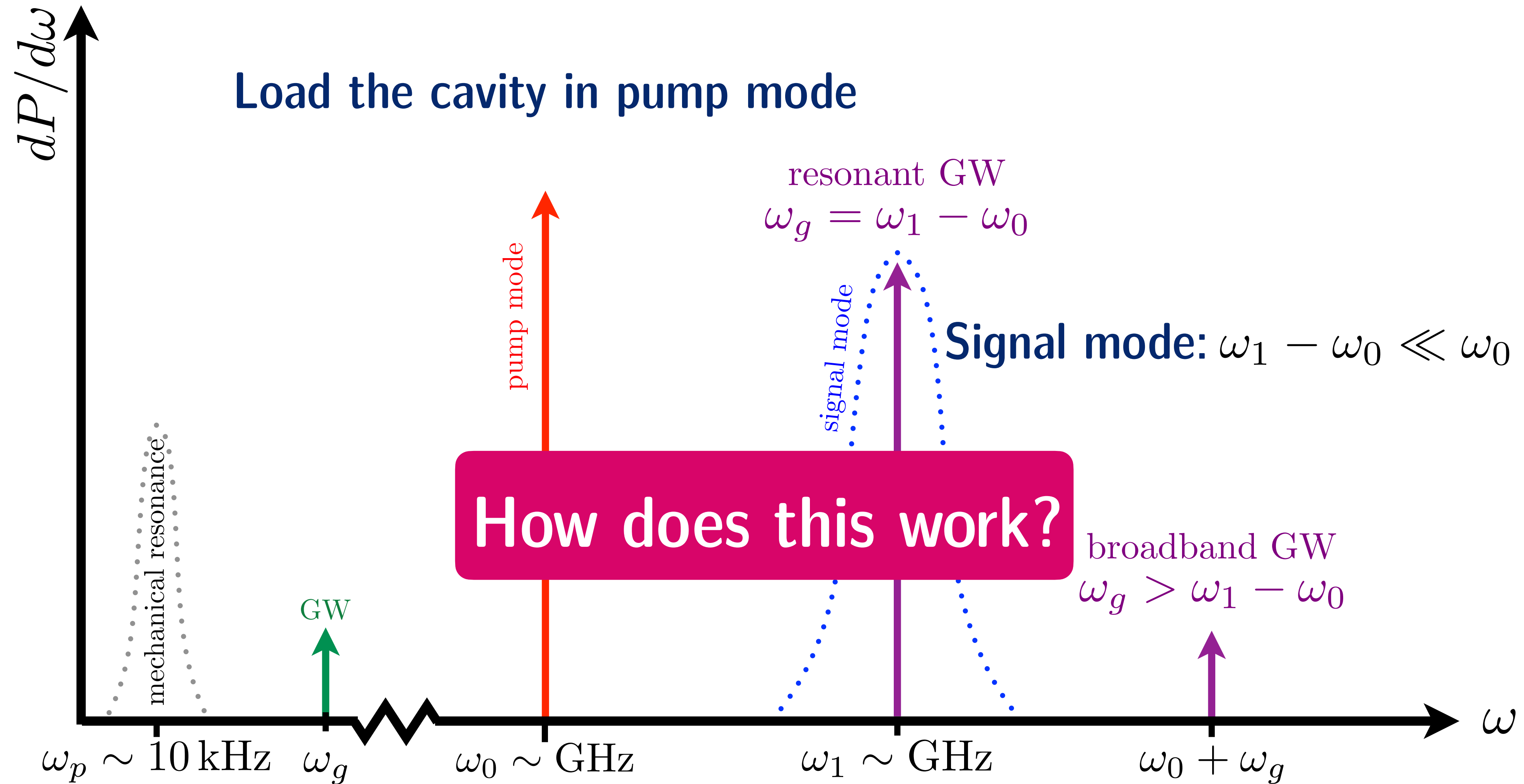




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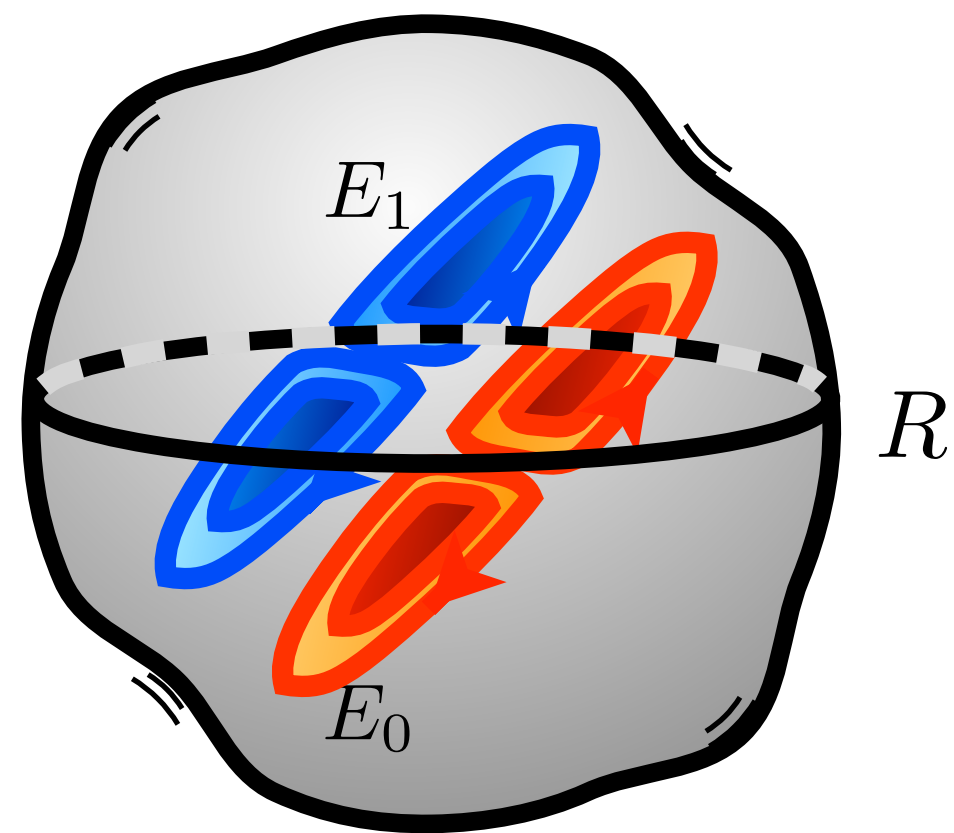


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# MAGO 2.0

$$\omega \sim \text{GHz}$$

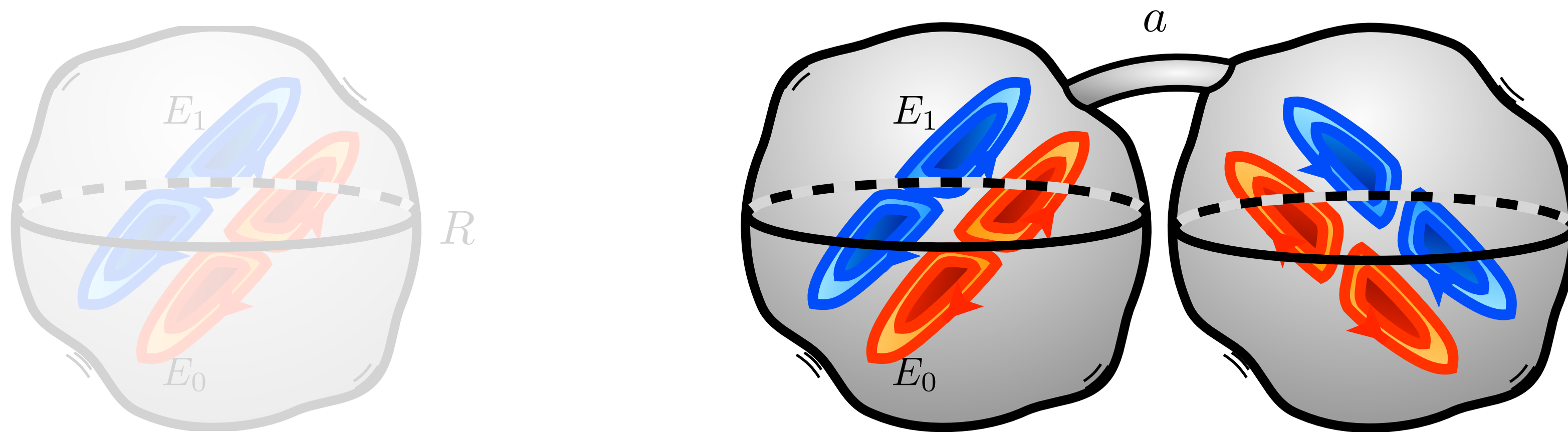
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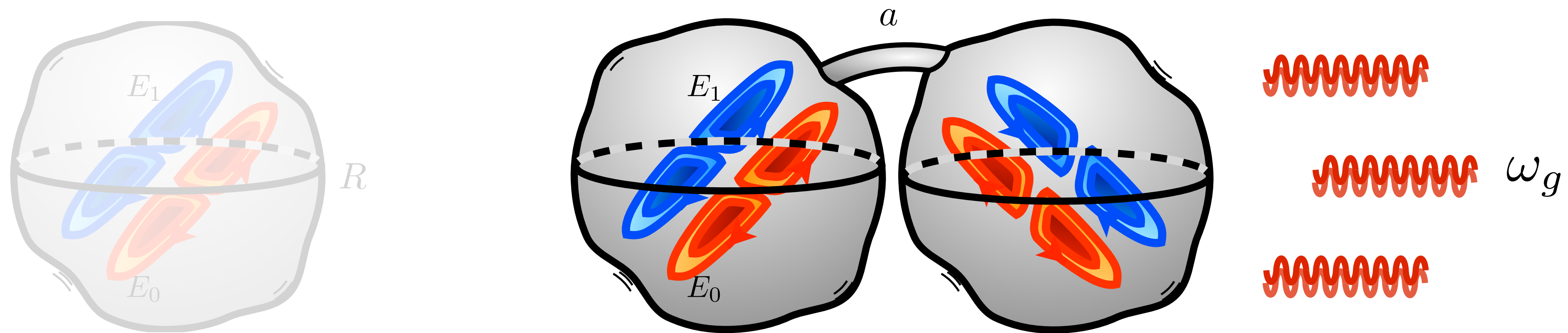
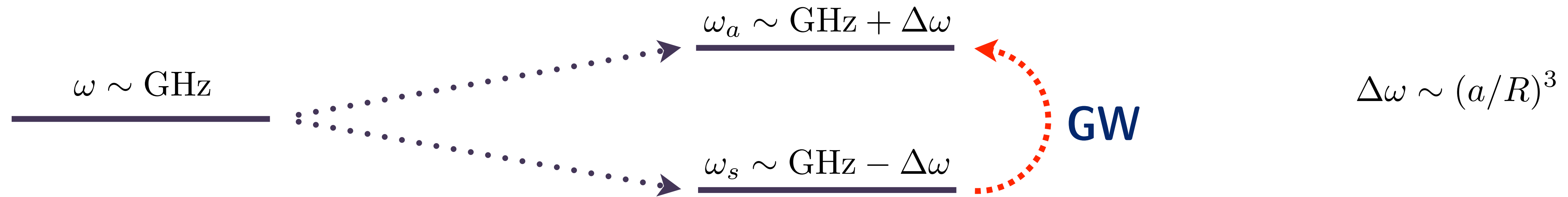
# MAGO 2.0



$$\Delta\omega \sim (a/R)^3$$



# MAGO 2.0



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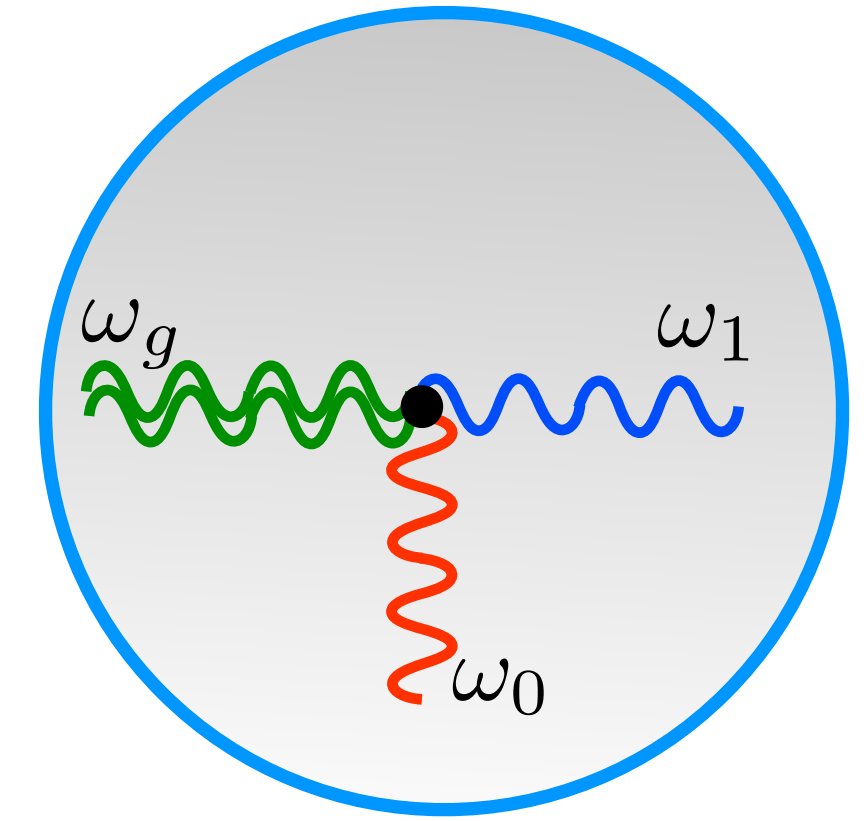
# EM and Mechanical signals



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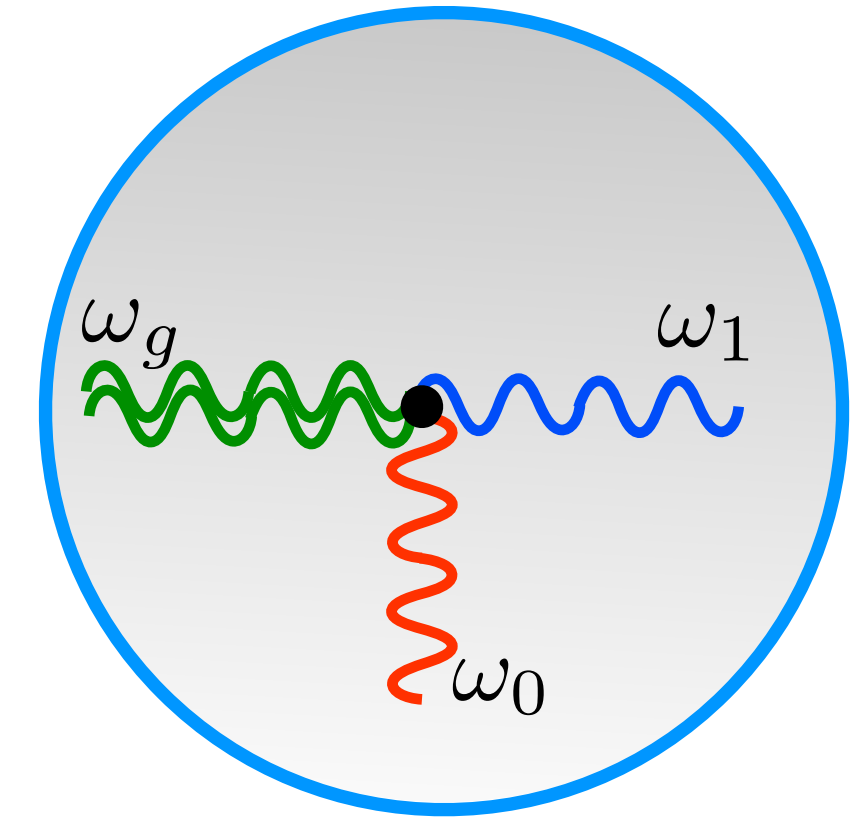
# EM and Mechanical signals

Parametrics of the EM signal:  $E_{\text{sig}}^{(\text{EM})} \sim Q_{\text{em}} (\omega_g L_{\text{cav}})^2 h^{\text{TT}} E_0$



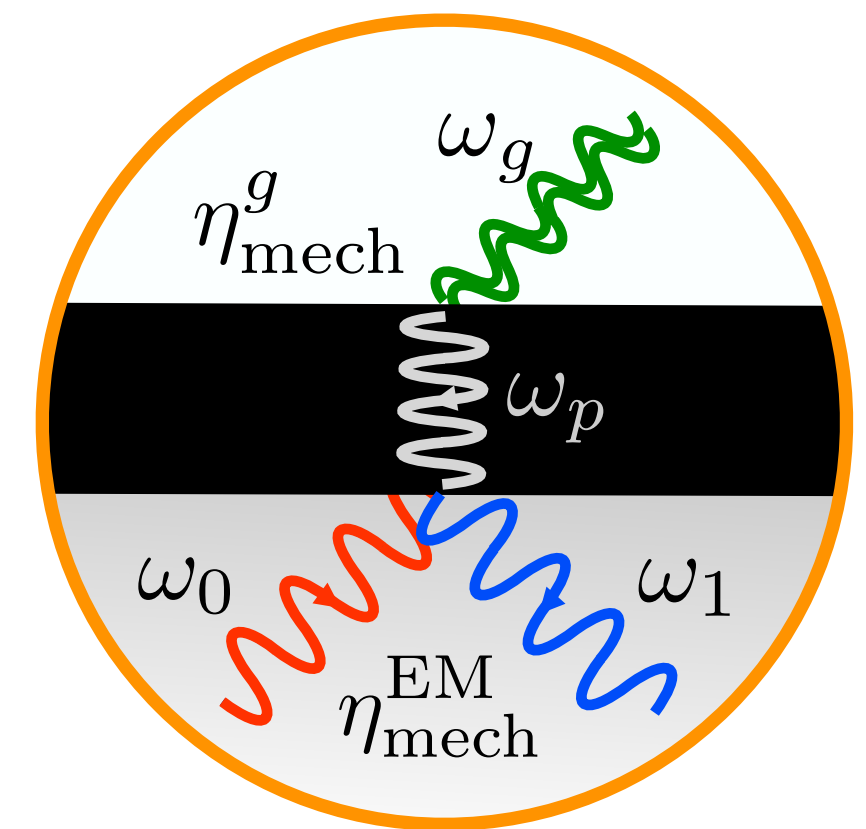
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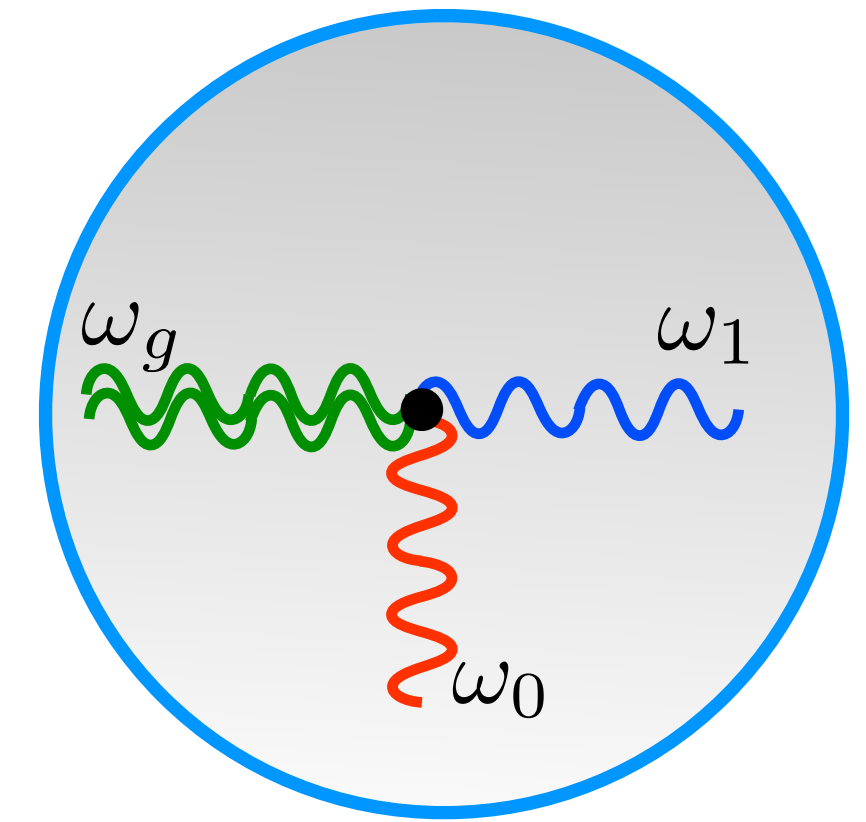
Mechanical signal:

$$E_{\text{sig}}^{(\text{mech})} \sim Q_{\text{em}} h^{\text{TT}} E_0 \min \left( 1, \frac{\omega_g L_{\text{cav}}}{c_s} \right)^2$$



# EM and Mechanical signals

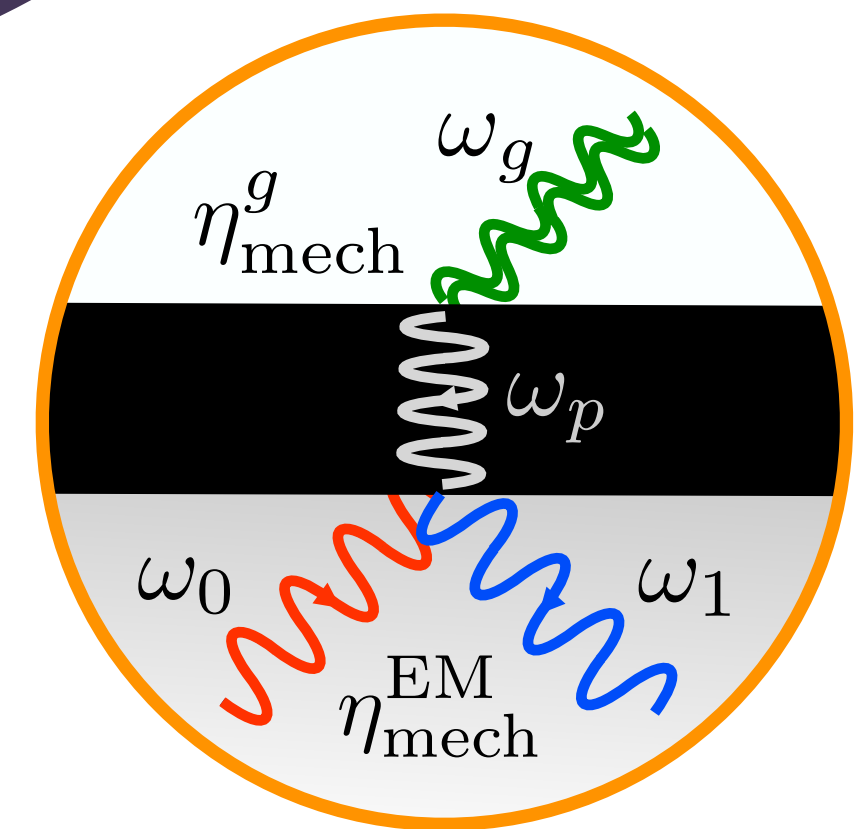
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*Enhanced by up to  $1/c_s^2 \gg 1$  (!)*

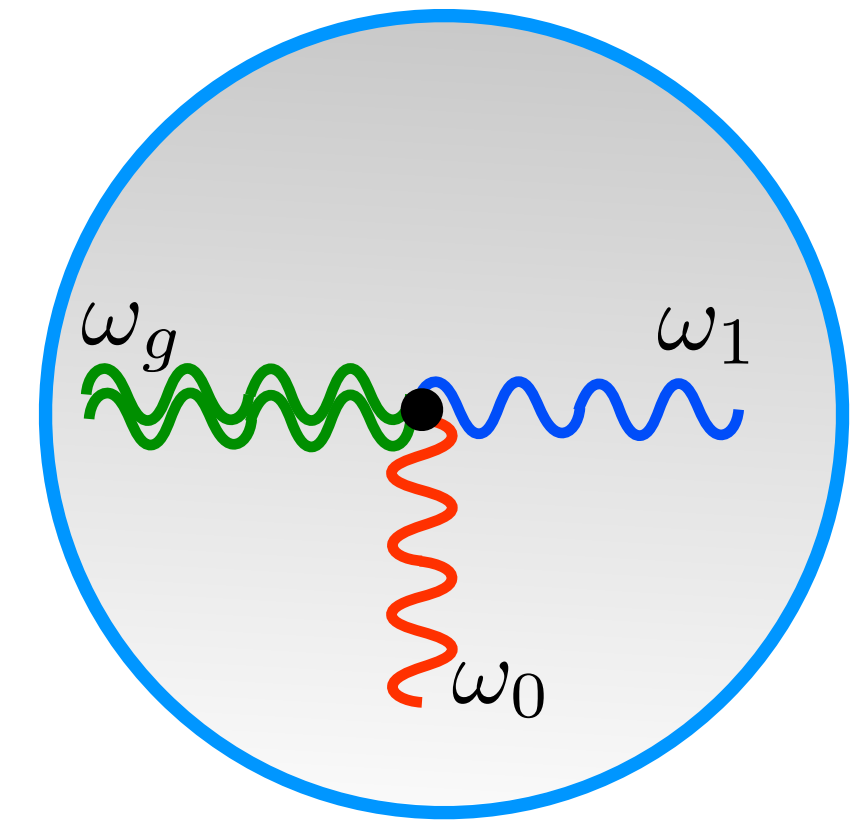
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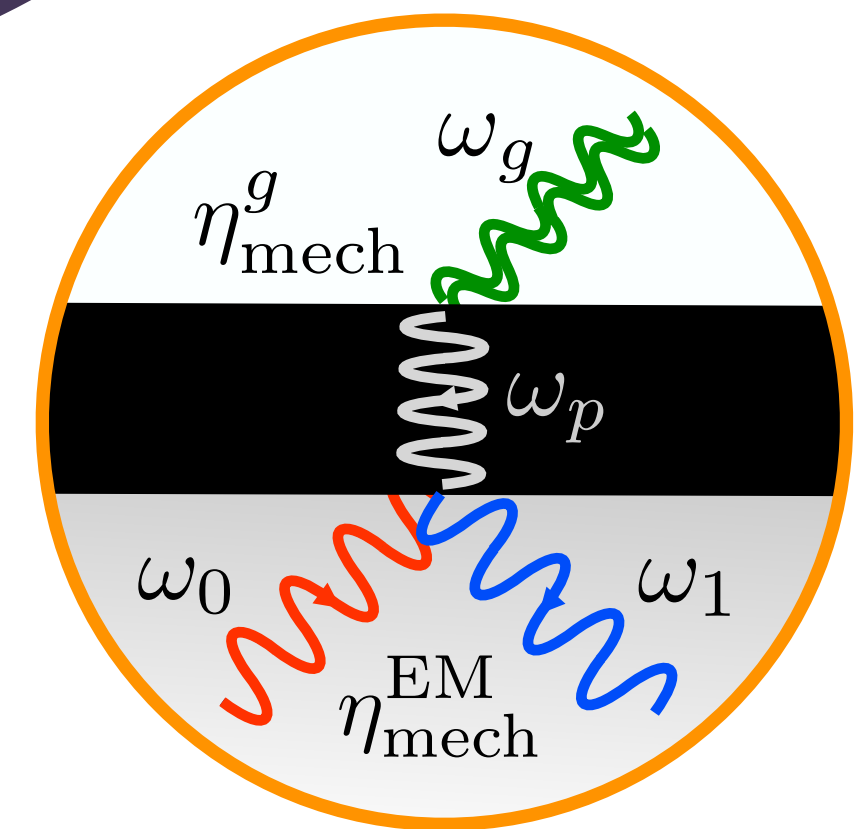


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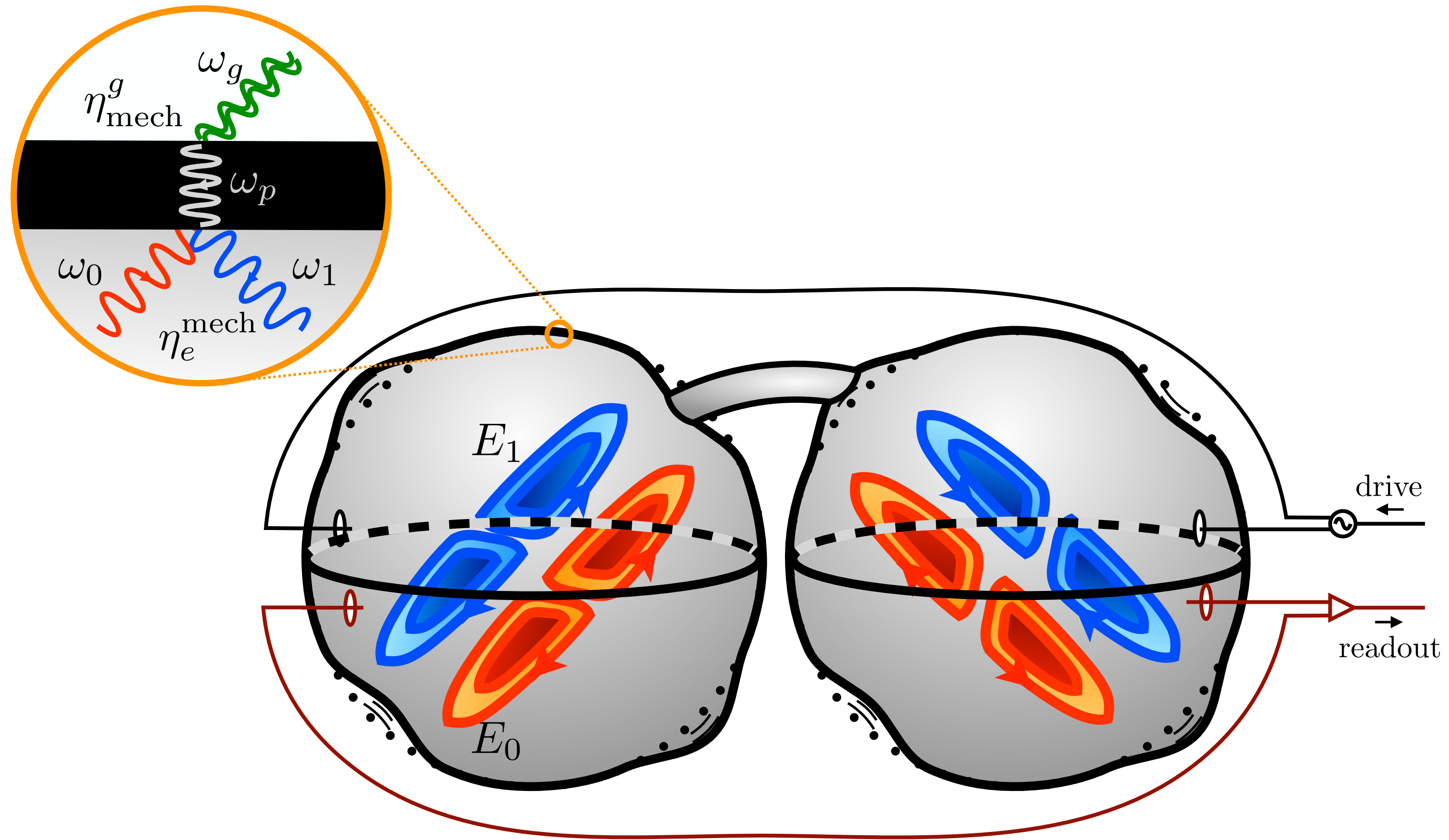
**Mechanical modes less “rigid” than EM modes**



# Gravitational Wave Signal: *Mechanical*

Mechanical modes of a hollow sphere:

$$\mathbf{q}_{mnl} = q_{0,mn} [f_{mn}(r)Y_{ml}(\theta, \varphi)\hat{r} - ig_{mn}(r)\hat{r} \times \mathbf{L}Y_{ml}(\theta, \varphi)]$$



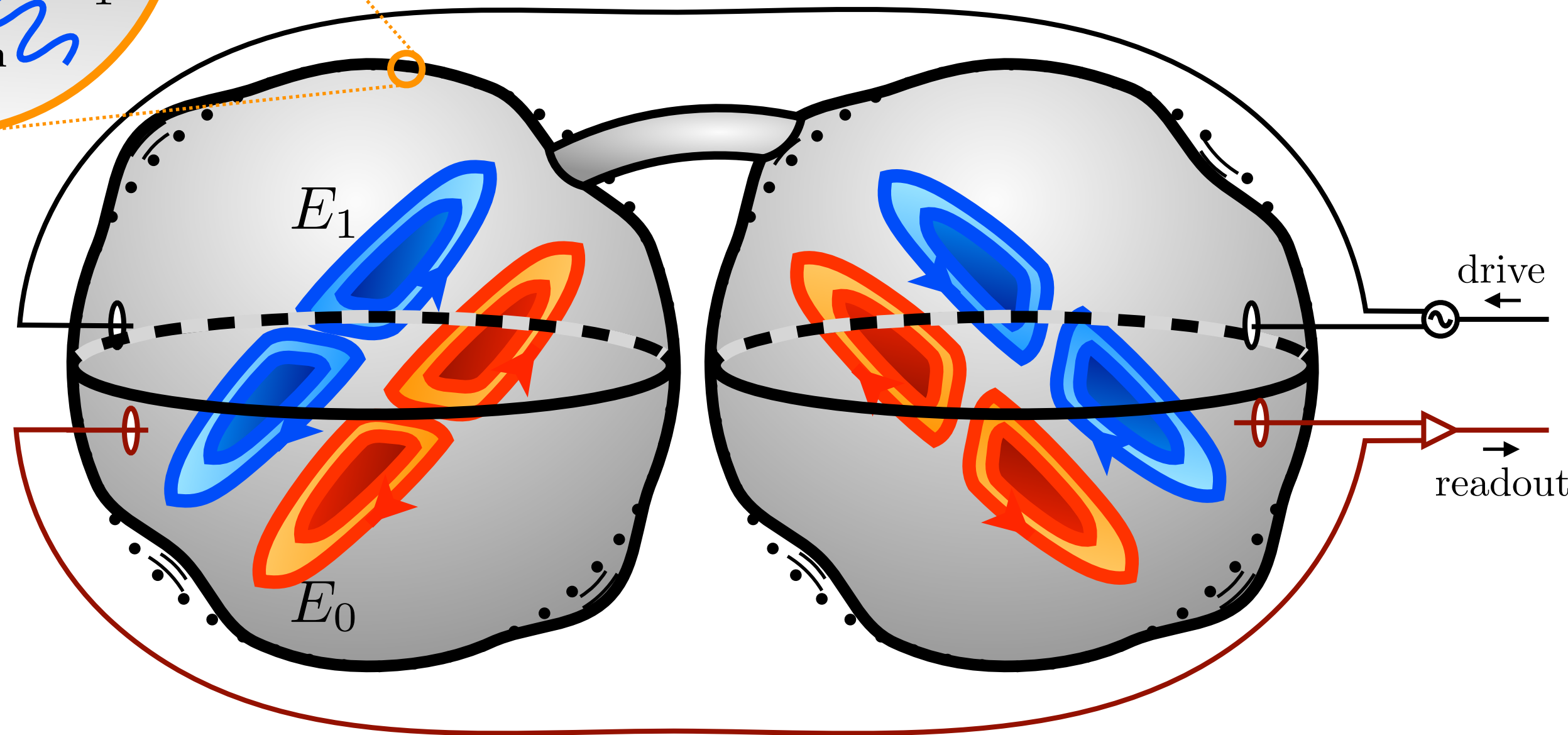
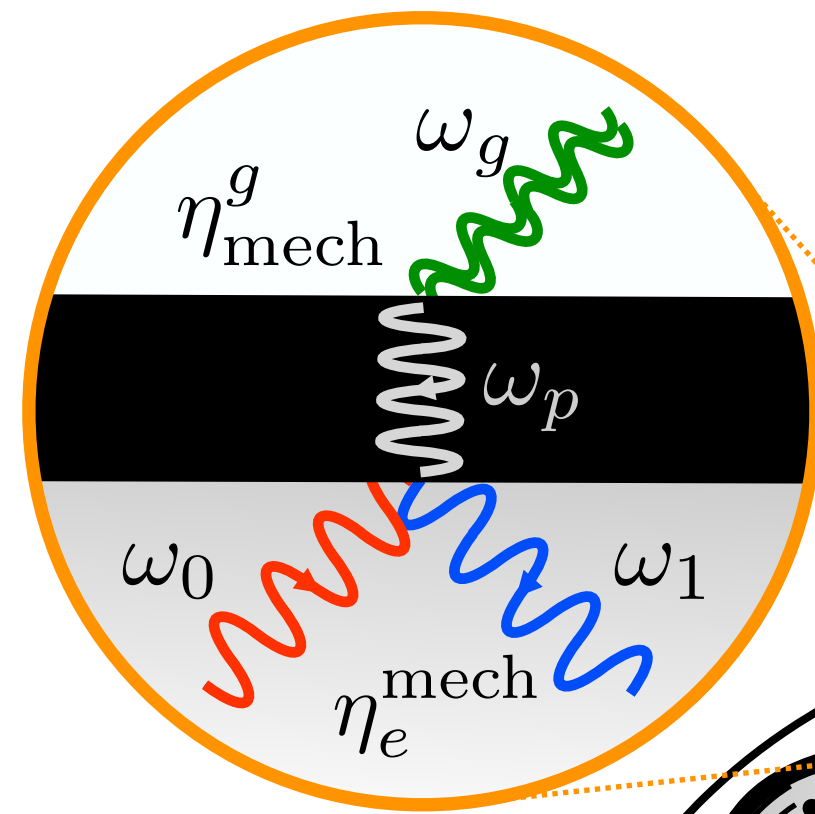
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GW coupling:

$$\ddot{q}_m(t) + \frac{\omega_m}{Q_m} \dot{q}_m(t) + \omega_m^2 q_m(t) = \frac{1}{2} \ddot{h}_{ij}^{\text{TT}} \int_V d^3\mathbf{x} x_j (q_{mnl})_i$$





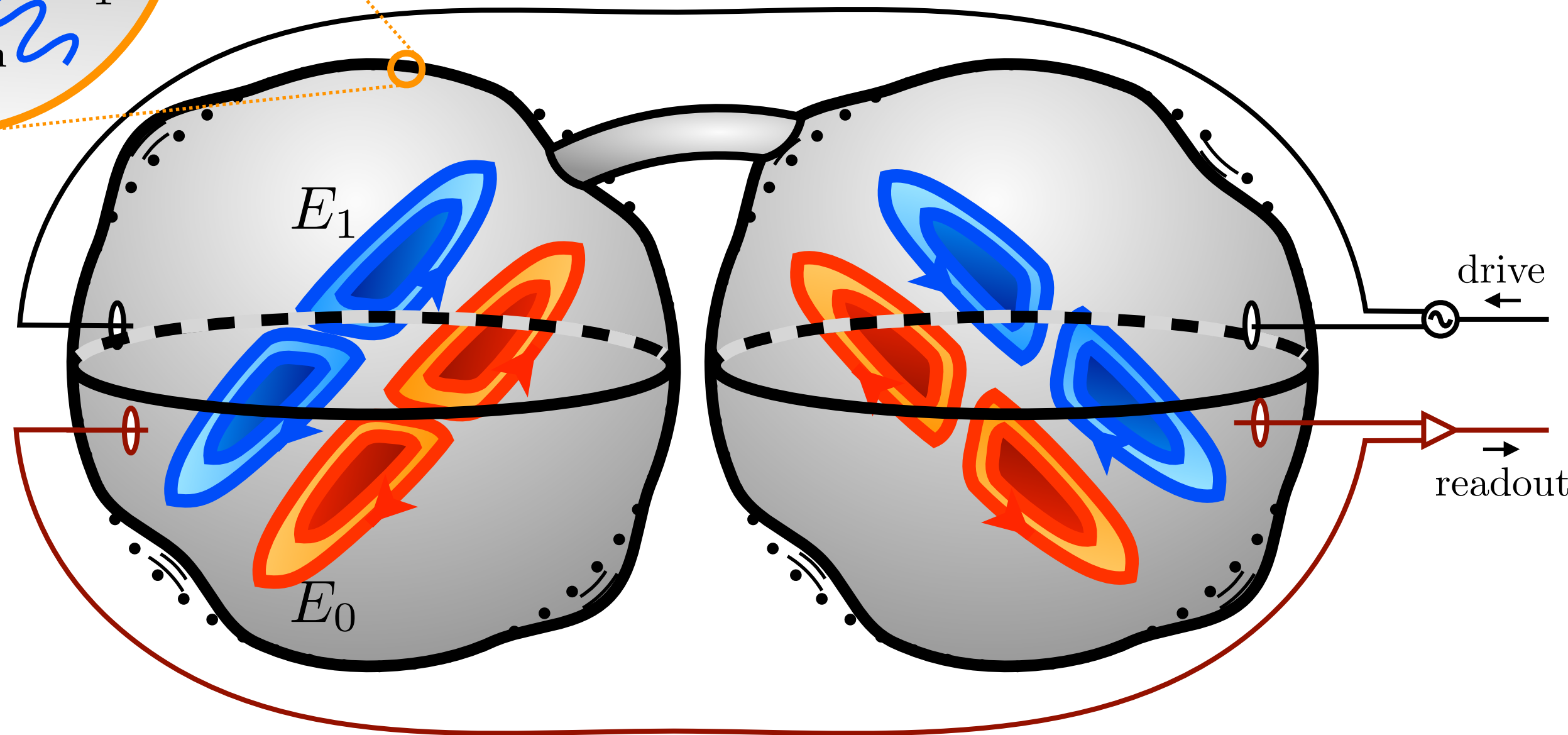
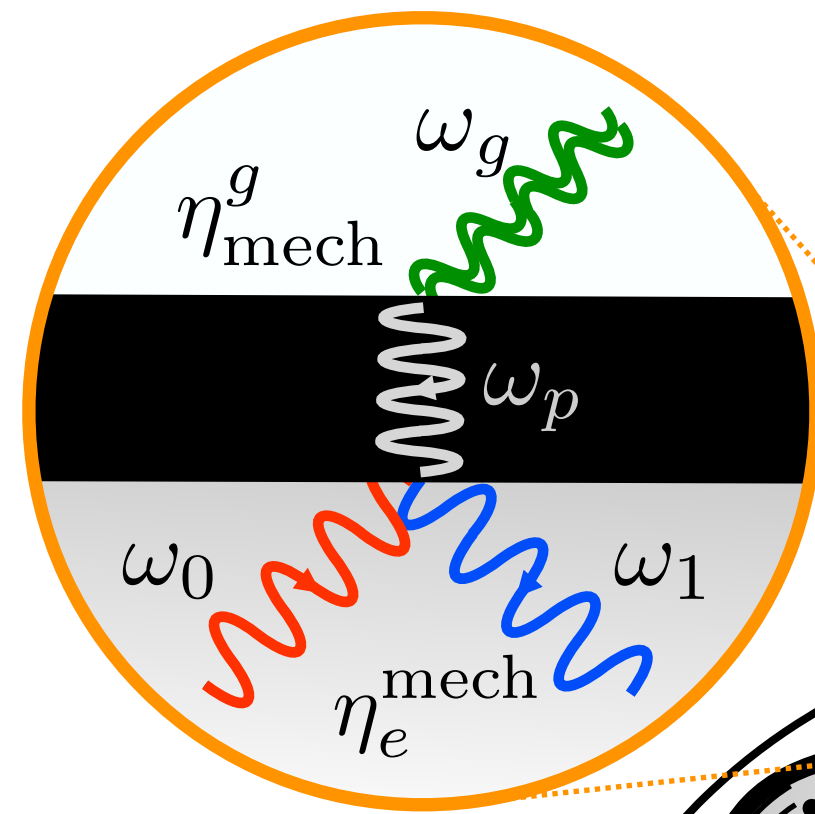
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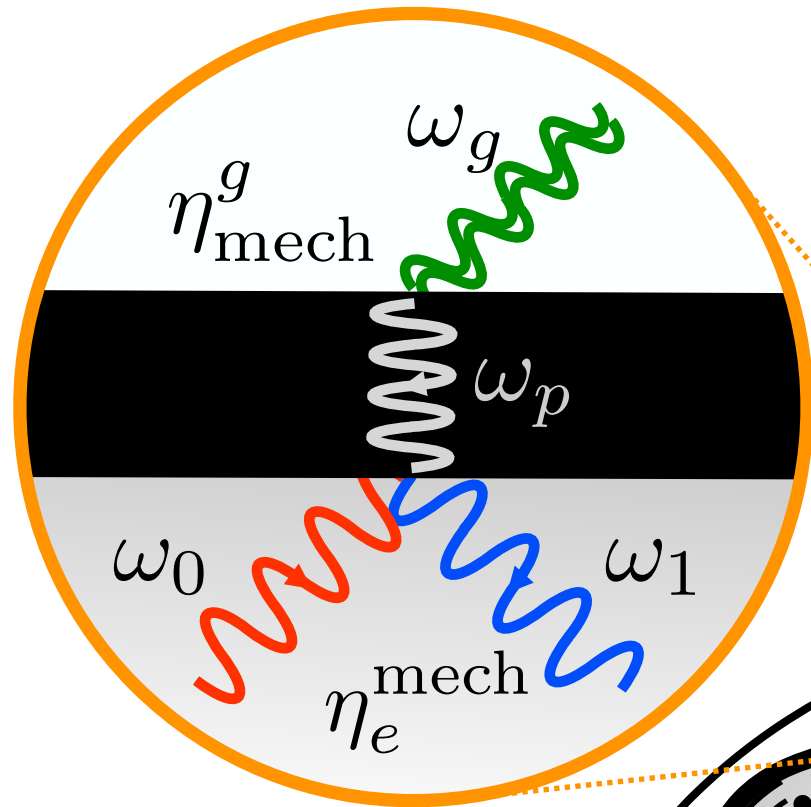
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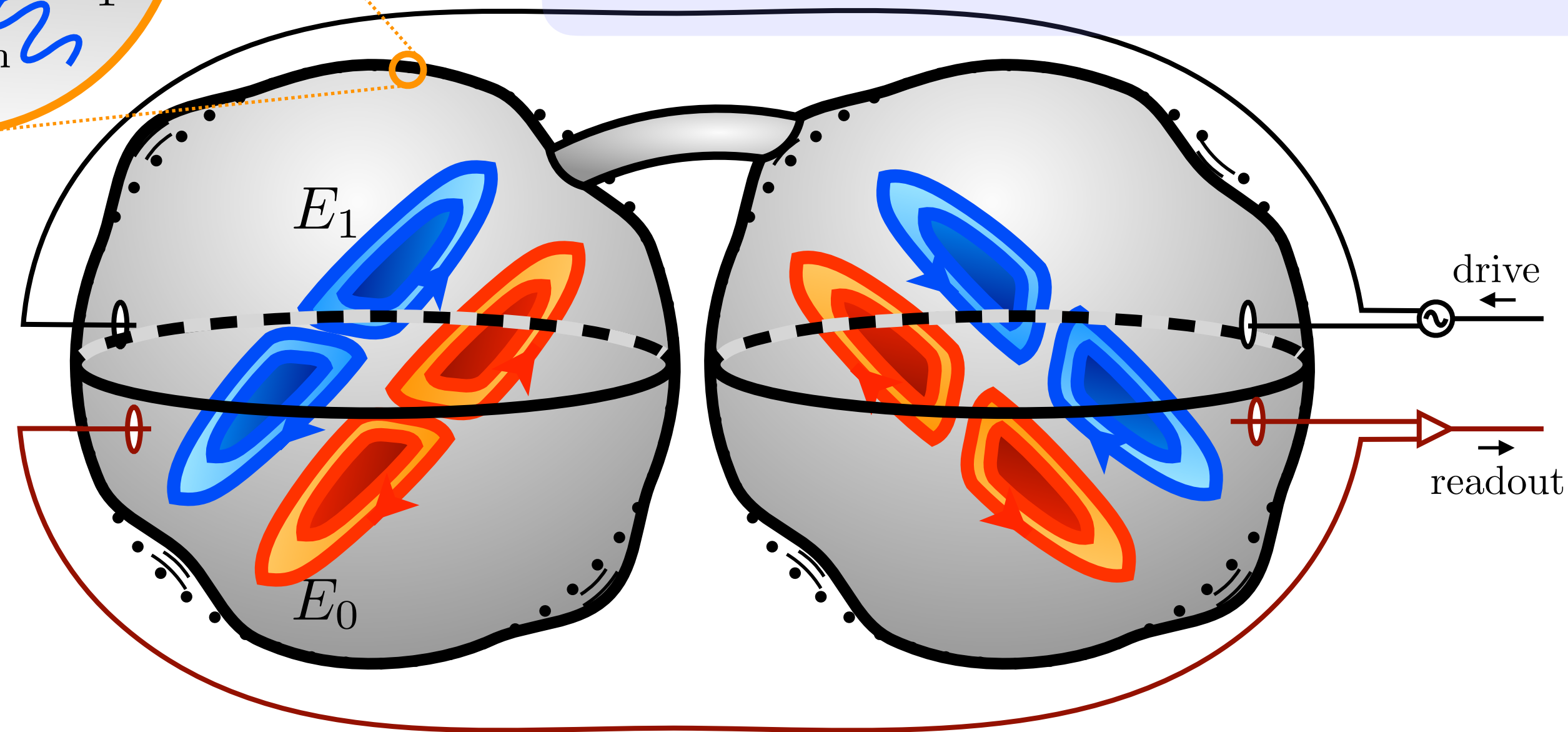


# Gravitational Wave Signal: *Mechanical*

Mechanical coupling to EM: 
$$\left( \partial_t^2 + \frac{\omega_1}{Q} \partial_t + \omega_1^2 \right) e_1 = -\omega_1^2 e_0 \frac{\int_{\Delta V} d^3x (\mathbf{E}_0 \cdot \mathbf{E}_1^* - (\omega_0/\omega_1) \mathbf{B}_0 \cdot \mathbf{B}_1^*)}{\int_{V_0} d^3x |\mathbf{E}_1|^2} + \mathcal{O}(\hbar^2)$$



$$\eta_{\text{mech}}^{\text{EM}} = \frac{V_{\text{cav}}^{1/3}}{\int d^3\mathbf{x} |\tilde{\mathbf{E}}|^2} \left[ \int_{S_{0L}} d\mathbf{A} \cdot \mathbf{U}_p \left( |\tilde{\mathbf{E}}|^2 - |\tilde{\mathbf{B}}|^2 \right) - \int_{S_{0R}} d\mathbf{A} \cdot \mathbf{U}_p \left( |\mathcal{R}\tilde{\mathbf{E}}|^2 - |\mathcal{R}\tilde{\mathbf{B}}|^2 \right) \right]$$

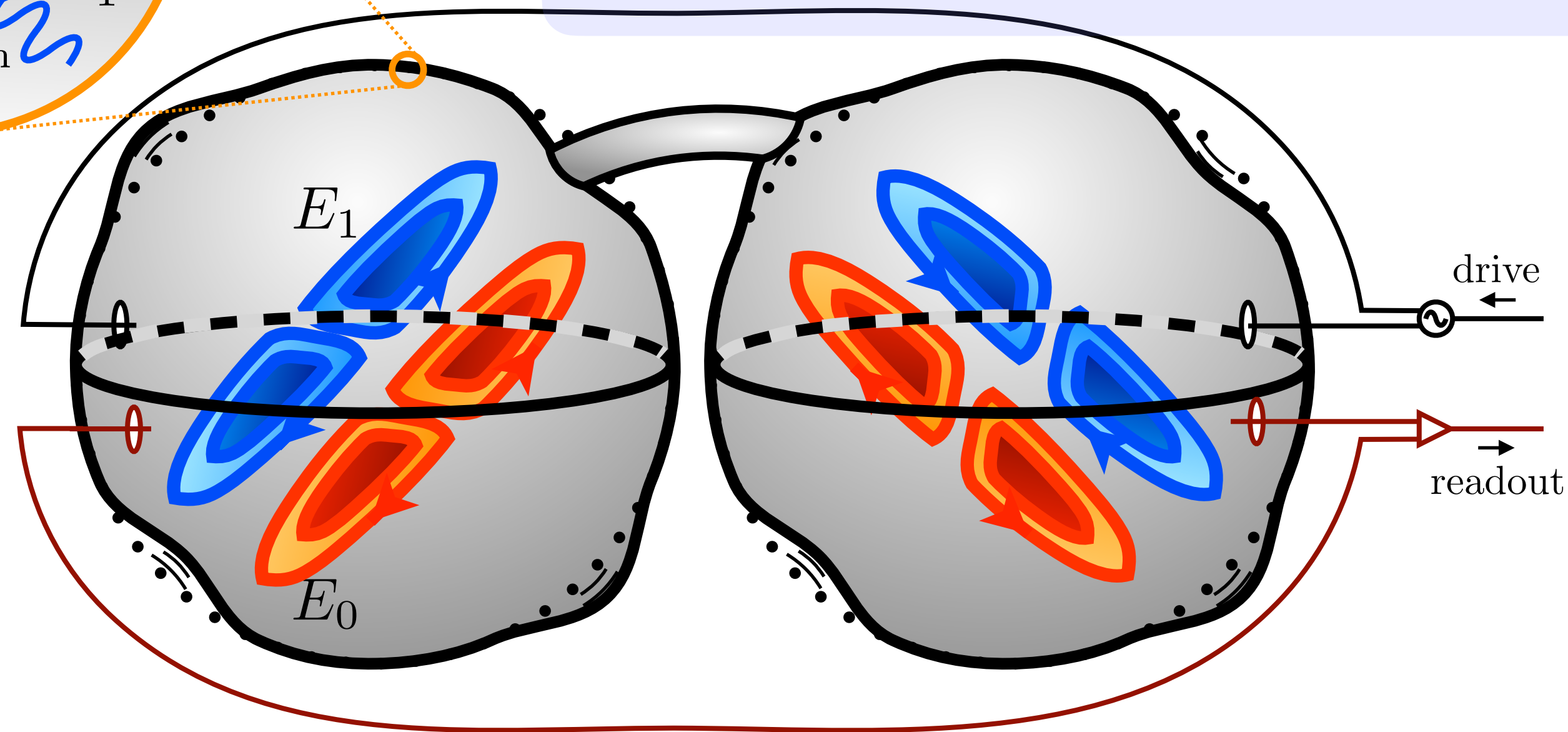
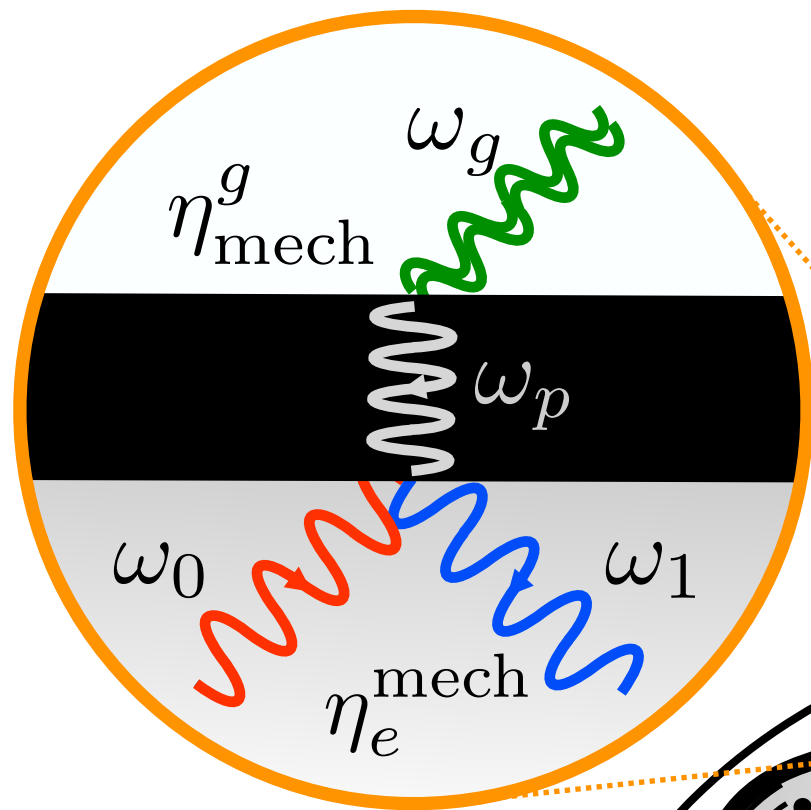


# Gravitational Wave Signal: *Mechanical*

Mechanical coupling to EM:  $\left(\partial_t^2 + \frac{\omega_1}{Q}\partial_t + \omega_1^2\right)e_1 = -\omega_1^2 e_0 \frac{\int_{\Delta V} d^3x (\mathbf{E}_0 \cdot \mathbf{E}_1^* - (\omega_0/\omega_1)\mathbf{B}_0 \cdot \mathbf{B}_1^*)}{\int_{V_0} d^3x |\mathbf{E}_1|^2} + \mathcal{O}(h^2)$

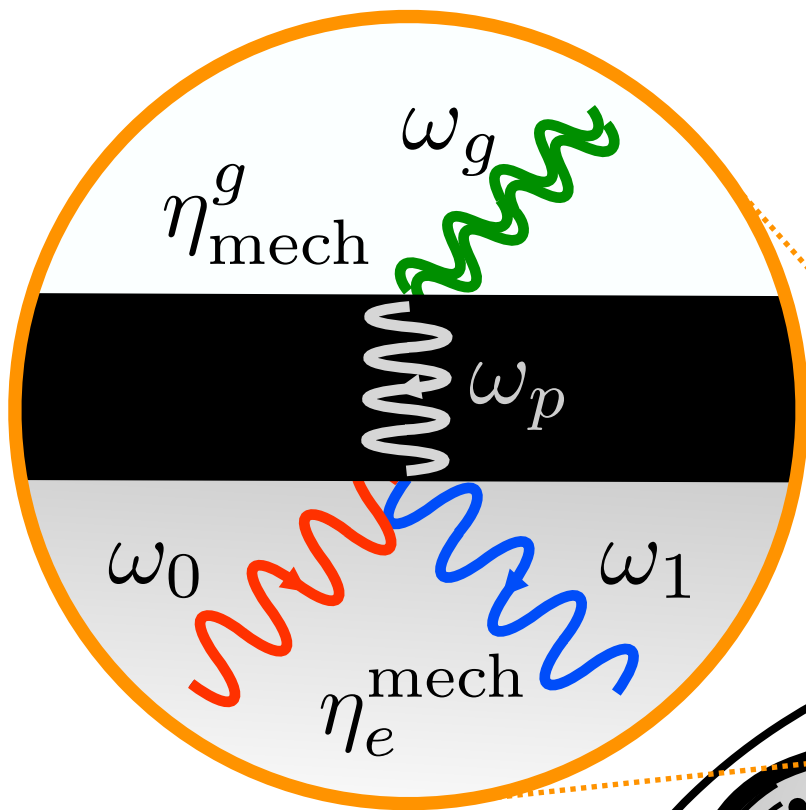
Integral over deformation to cavity volume

$$\eta_{\text{mech}}^{\text{EM}} = \frac{V_{\text{cav}}^{1/3}}{\int d^3\mathbf{x} |\tilde{\mathbf{E}}|^2} \left[ \int_{S_{0L}} d\mathbf{A} \cdot \mathbf{U}_p \left( |\tilde{\mathbf{E}}|^2 - |\tilde{\mathbf{B}}|^2 \right) - \int_{S_{0R}} d\mathbf{A} \cdot \mathbf{U}_p \left( |\mathcal{R}\tilde{\mathbf{E}}|^2 - |\mathcal{R}\tilde{\mathbf{B}}|^2 \right) \right]$$

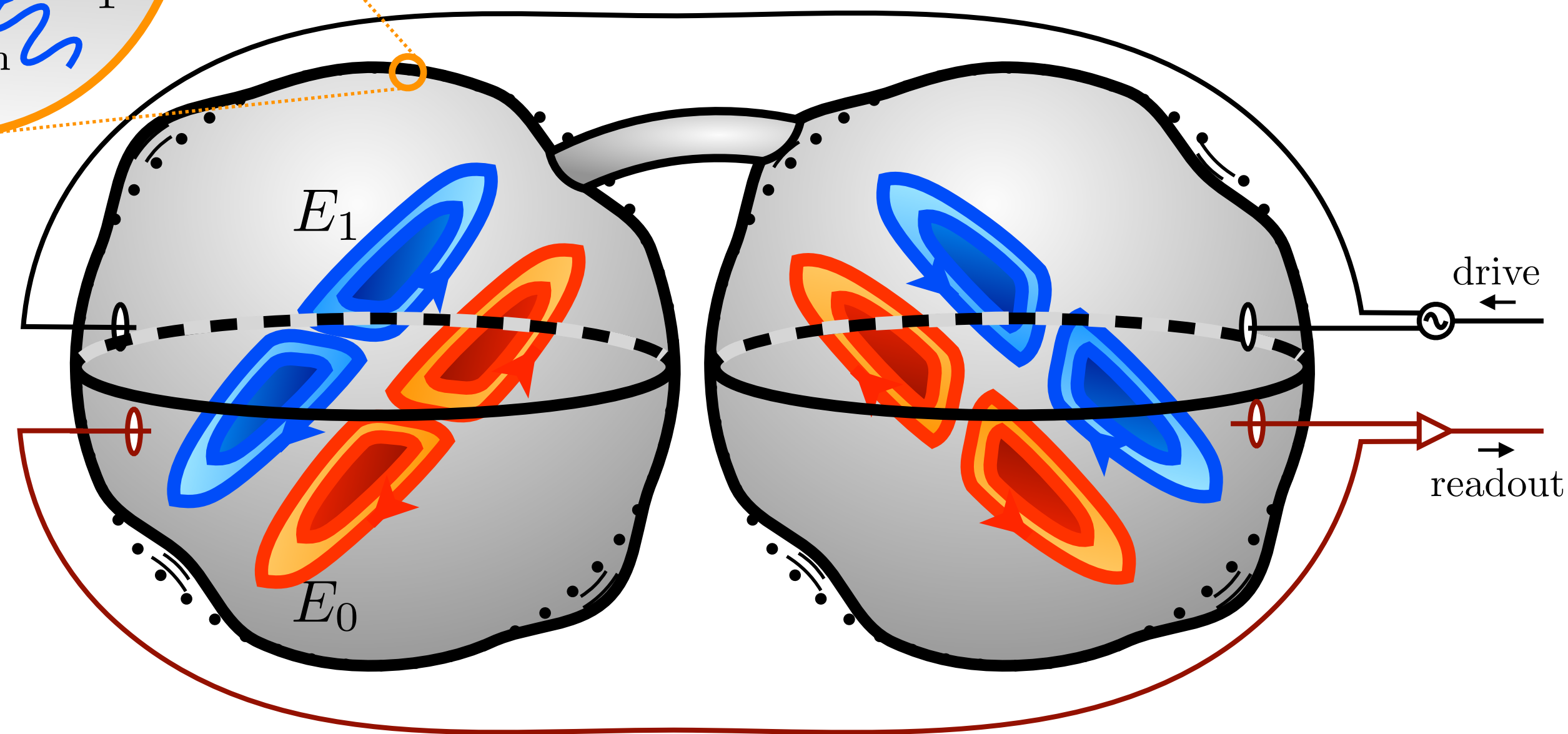


# Gravitational Wave Signal: *Mechanical*

Power in a heterodyne setup from monochromatic source



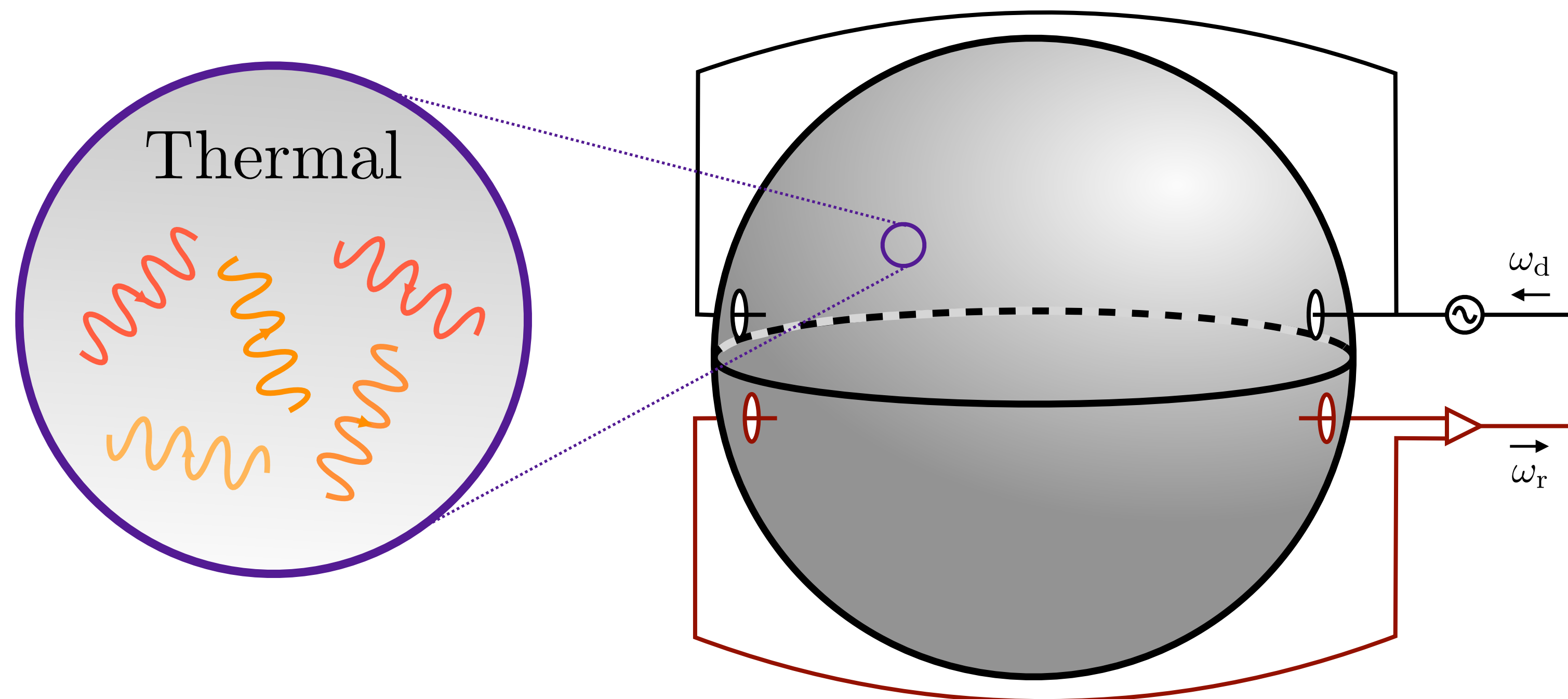
$$P_{\text{sig}} \simeq \frac{1}{16} Q_1 \omega_1 (|\eta_u^{\text{mech}}| |\eta_e^{\text{mech}}| h_0 E_0)^2 V_{\text{cav}} \times \begin{cases} Q_m^2 & \omega_g^2 - \omega_m^2 \ll \frac{\omega_m \omega_g}{Q_m} \\ \frac{\omega_g^4}{(\omega_g^2 - \omega_m^2)^2} & \omega_g^2 - \omega_m^2 \gg \frac{\omega_m \omega_g}{Q_m} \end{cases}$$



# Thermal Noise

Power spectral density:

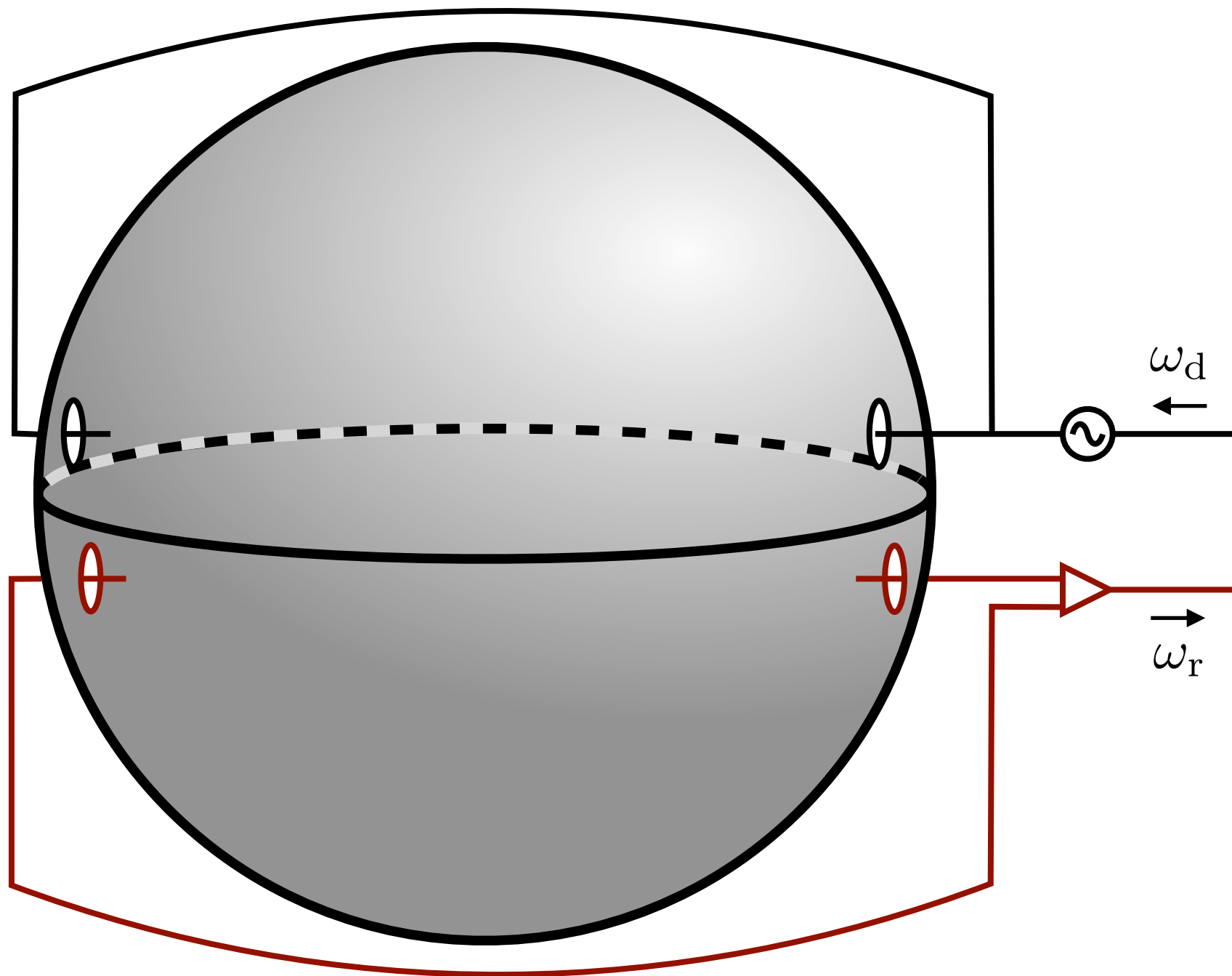
$$S_{\text{th}}(\omega) = \frac{Q_1}{Q_{\text{int}}} \frac{4\pi T (\omega \omega_1 / Q_1)^2}{(\omega^2 - \omega_1^2)^2 + (\omega \omega_1 / Q_1)^2}$$





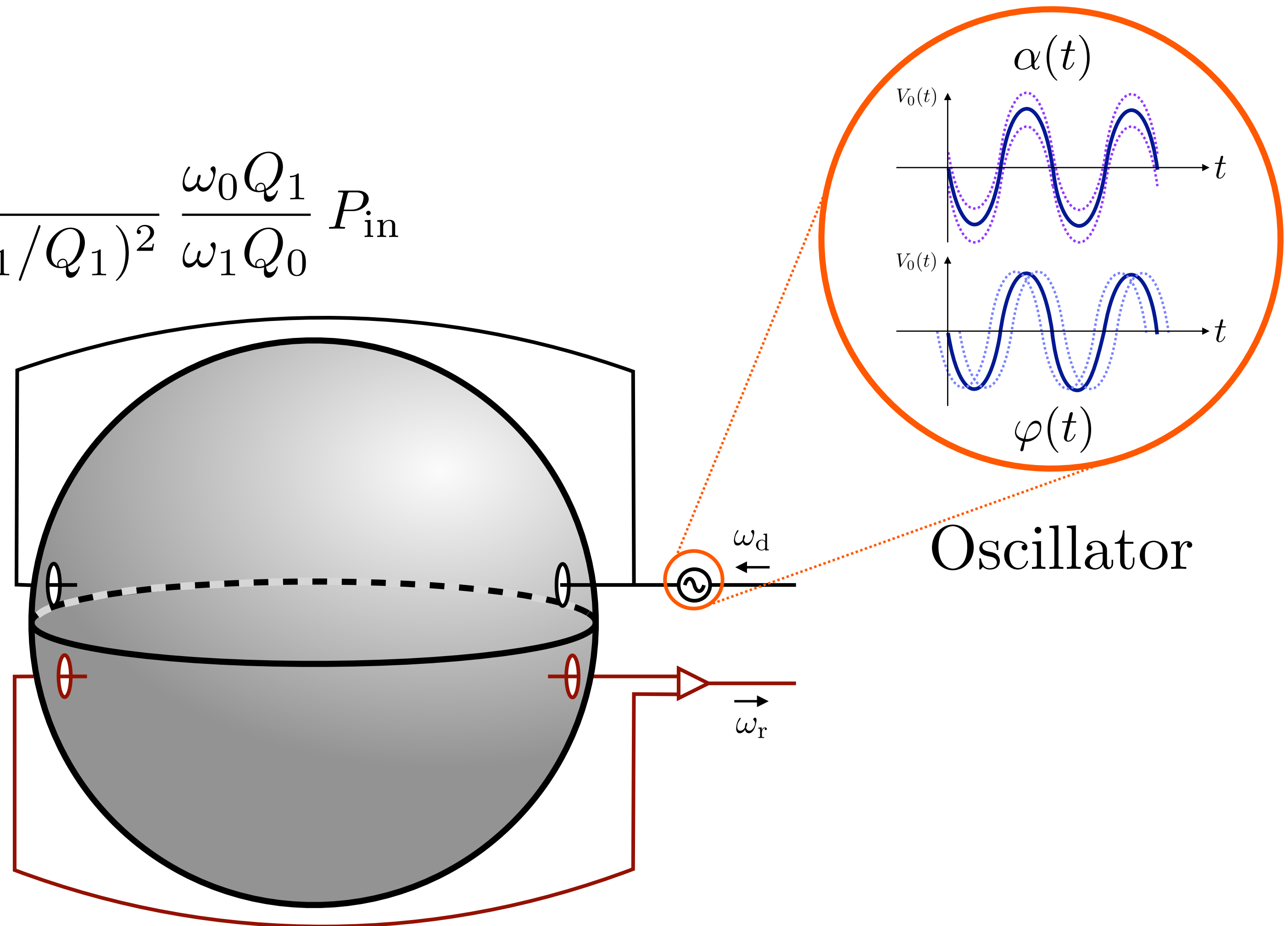
# Phase Noise

$$S_{\text{phase}}(\omega) \simeq \frac{1}{2} \epsilon_{1d}^2 S_{\varphi}(\omega - \omega_0) \\ \times \frac{(\omega \omega_1 / Q_1)^2}{(\omega^2 - \omega_1^2)^2 + (\omega \omega_1 / Q_1)^2} \frac{\omega_0 Q_1}{\omega_1 Q_0} P_{\text{in}}$$



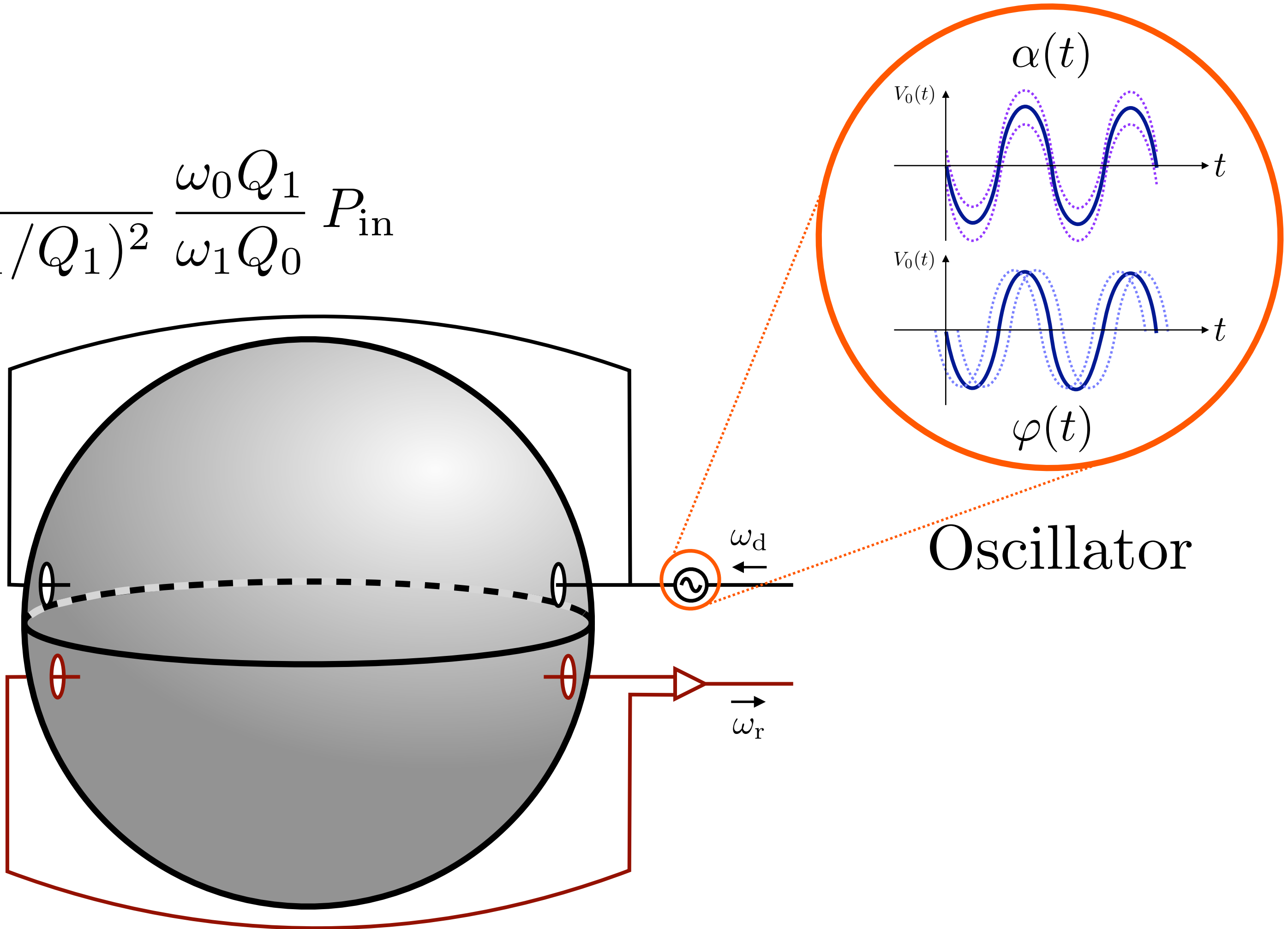
# Phase Noise

$$S_{\text{phase}}(\omega) \simeq \frac{1}{2} \epsilon_{1d}^2 S_{\varphi}(\omega - \omega_0) \times \frac{(\omega \omega_1 / Q_1)^2}{(\omega^2 - \omega_1^2)^2 + (\omega \omega_1 / Q_1)^2} \frac{\omega_0 Q_1}{\omega_1 Q_0} P_{\text{in}}$$



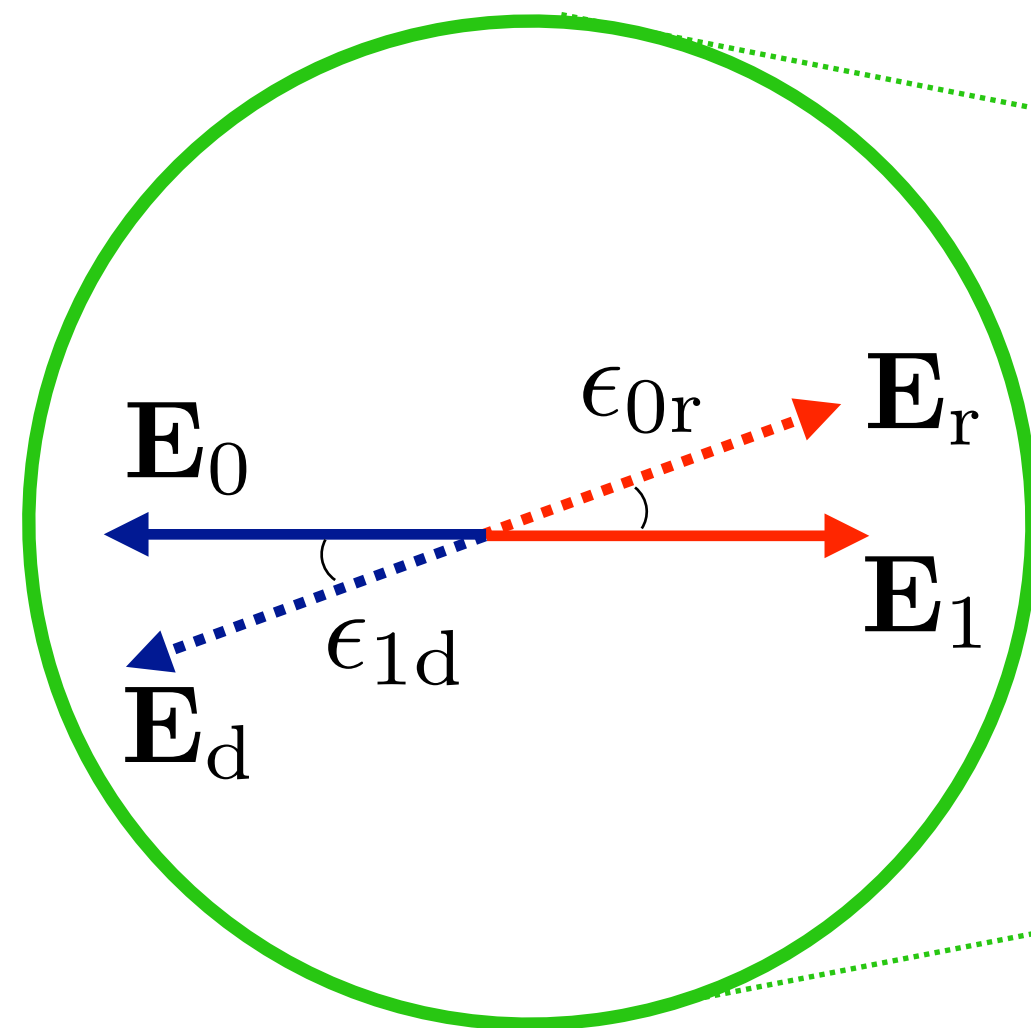
# Phase Noise

$$S_{\text{phase}}(\omega) \simeq \frac{1}{2} \epsilon_{1d}^2 S_{\varphi}(\omega - \omega_0) \times \frac{(\omega \omega_1 / Q_1)^2}{(\omega^2 - \omega_1^2)^2 + (\omega \omega_1 / Q_1)^2} \frac{\omega_0 Q_1}{\omega_1 Q_0} P_{\text{in}}$$

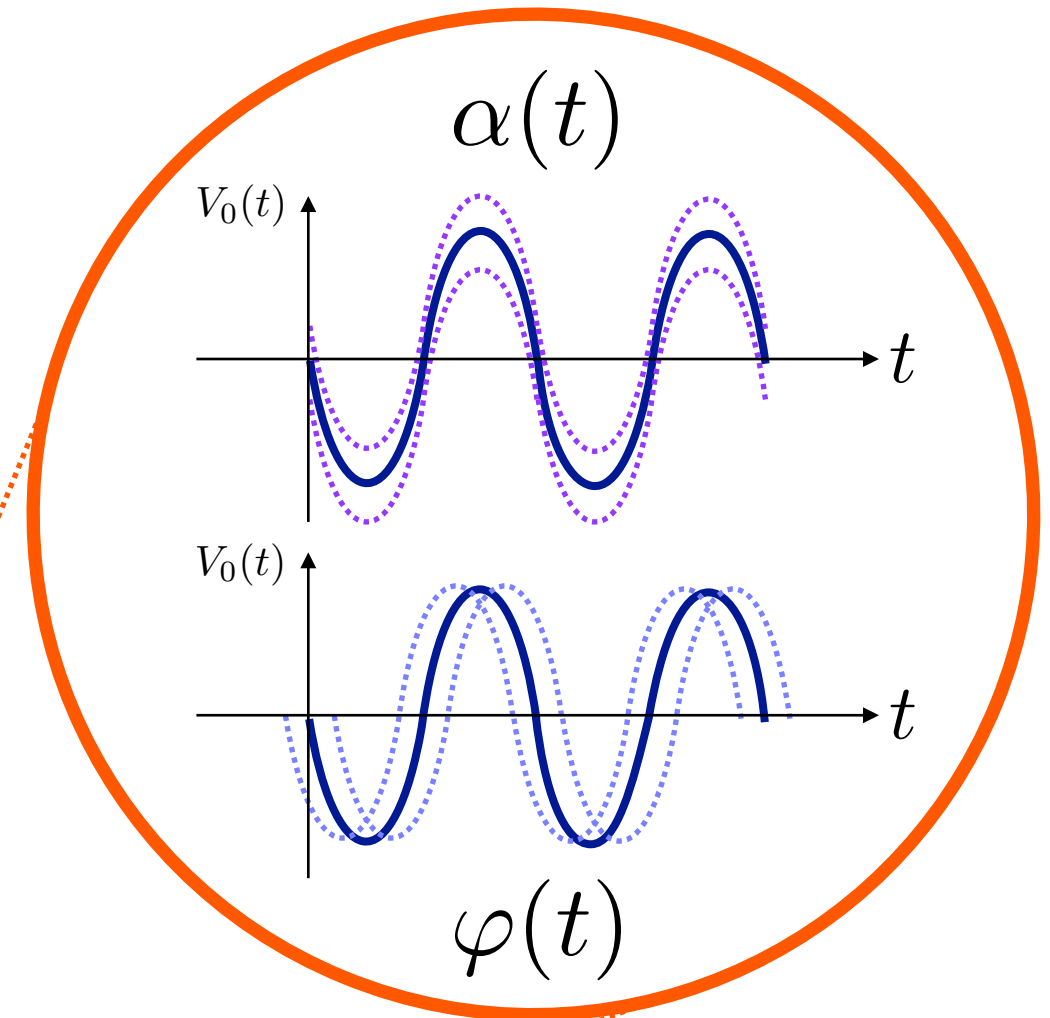
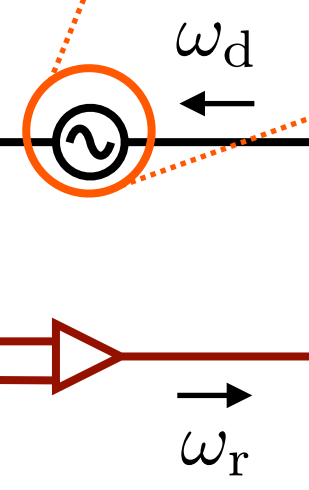
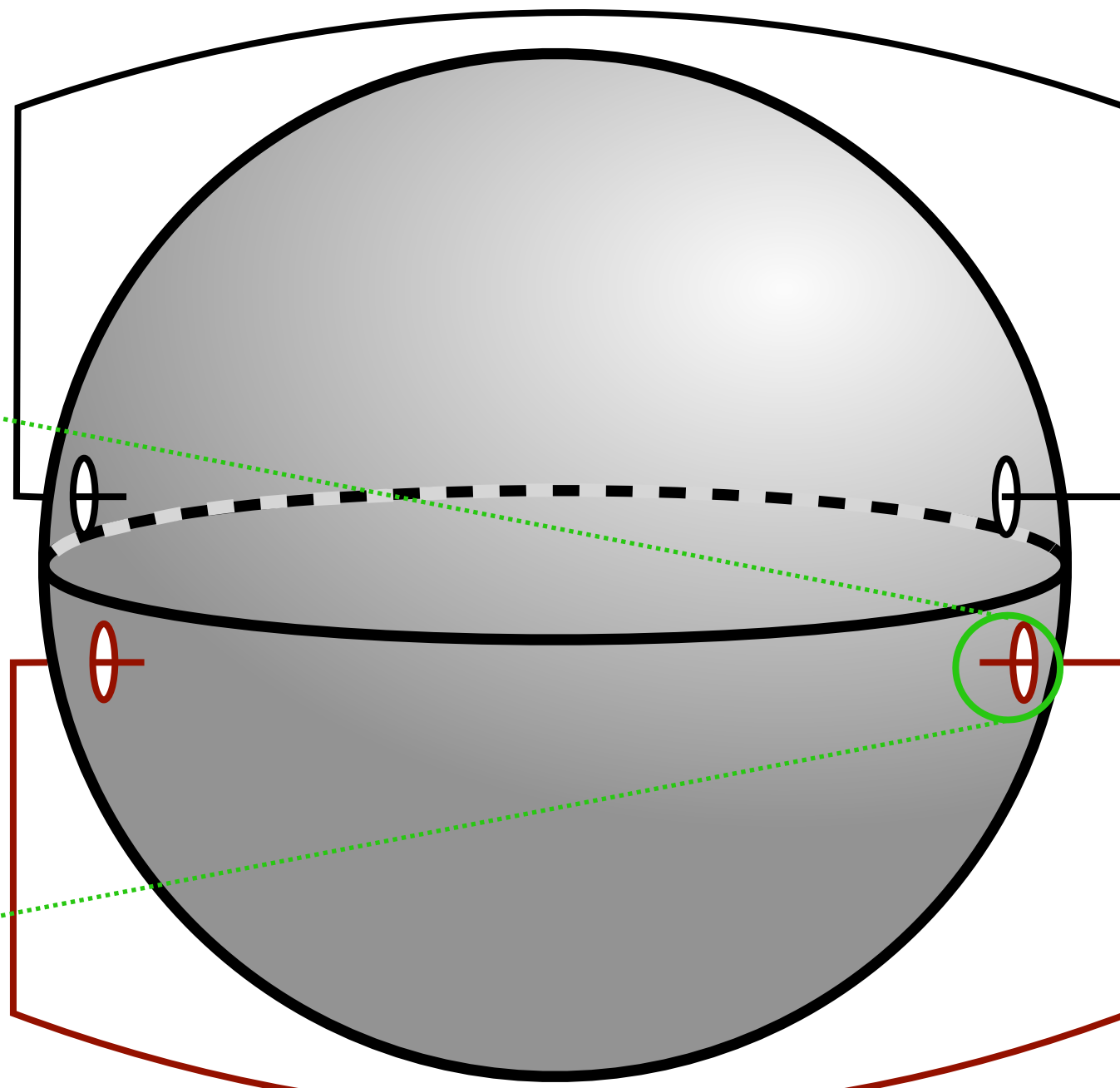


# Phase Noise

$$S_{\text{phase}}(\omega) \simeq \frac{1}{2} \epsilon_{1d}^2 S_{\varphi}(\omega - \omega_0) \times \frac{(\omega \omega_1 / Q_1)^2}{(\omega^2 - \omega_1^2)^2 + (\omega \omega_1 / Q_1)^2} \frac{\omega_0 Q_1}{\omega_1 Q_0} P_{\text{in}}$$



Loading  
&  
Readout

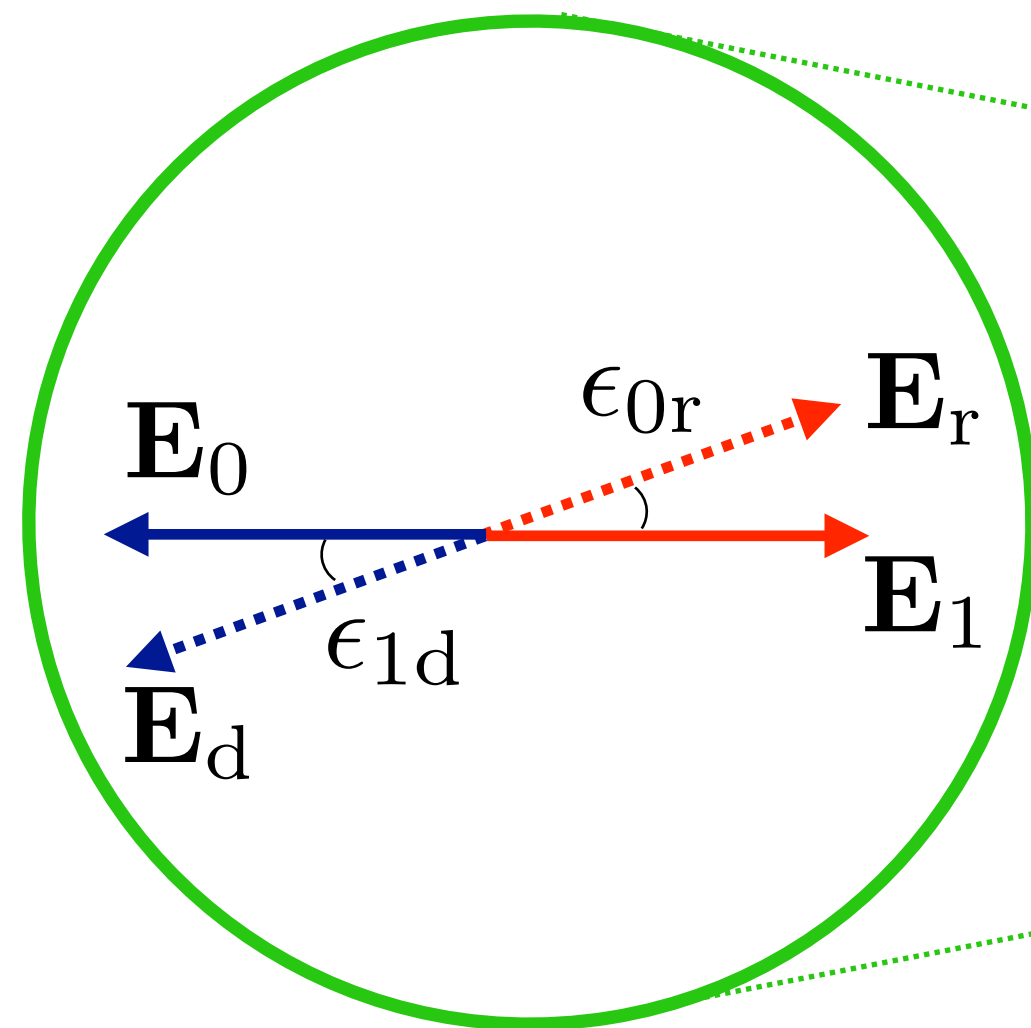


Oscillator

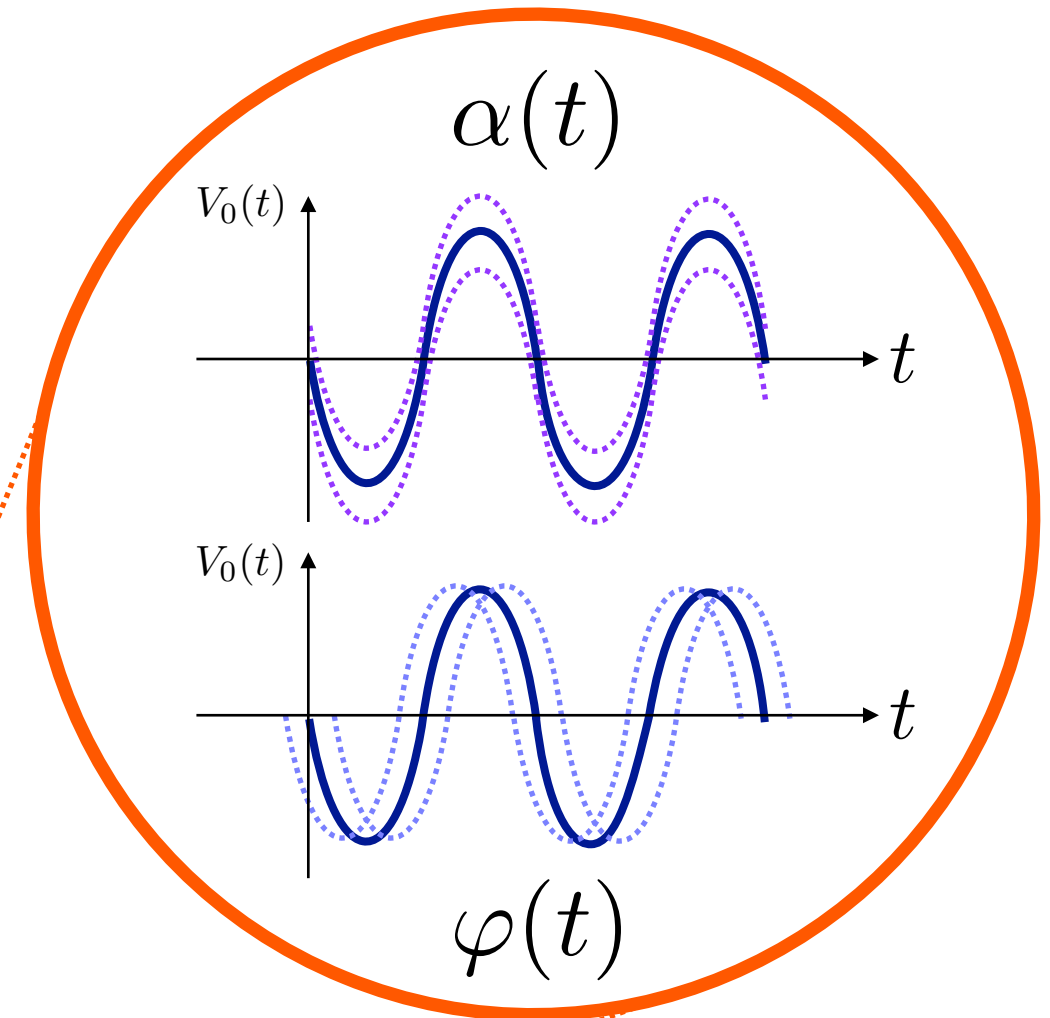
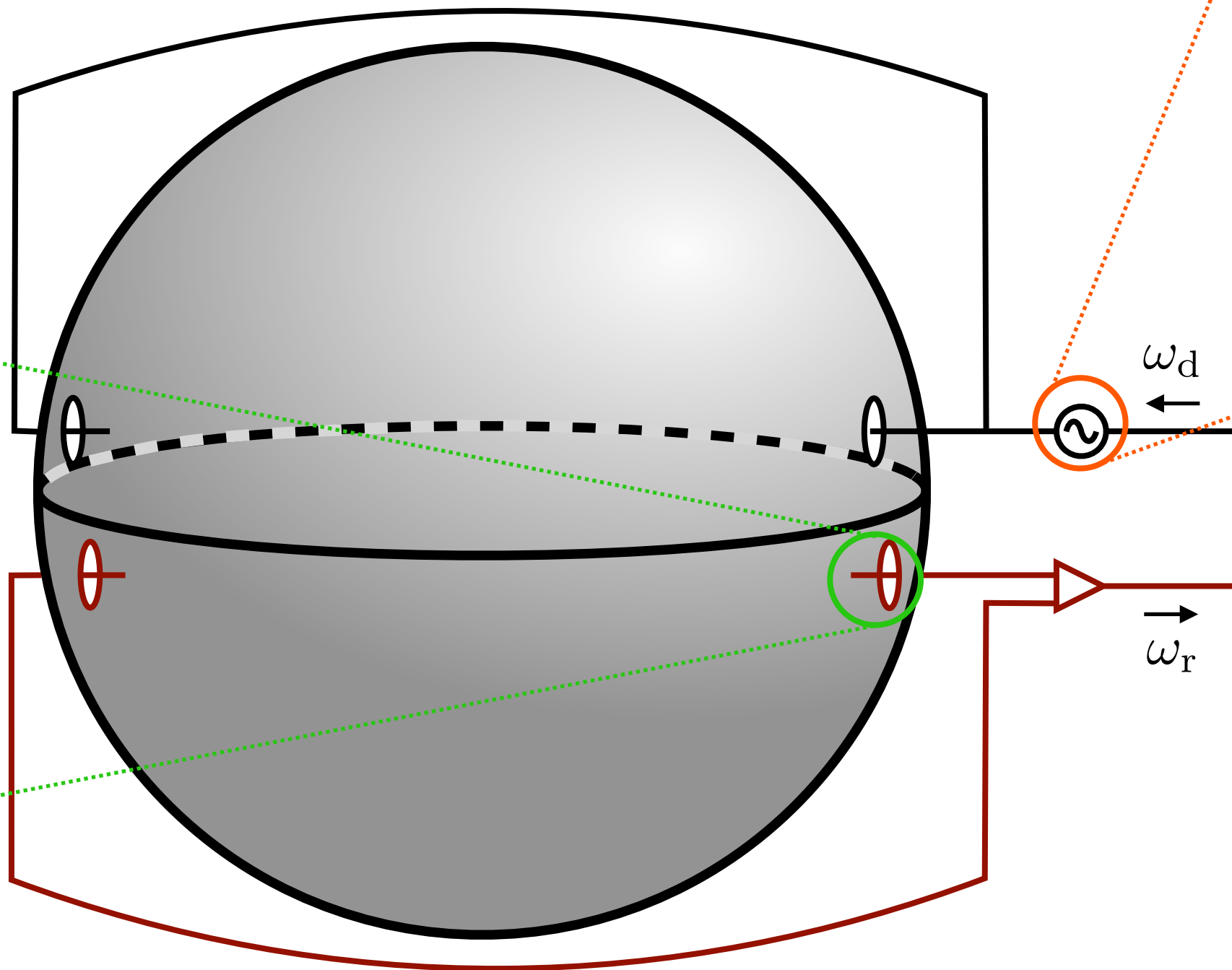


# Phase Noise

$$S_{\text{phase}}(\omega) \simeq \frac{1}{2} \epsilon_{1d}^2 S_{\varphi}(\omega - \omega_0) \times \frac{(\omega \omega_1 / Q_1)^2}{(\omega^2 - \omega_1^2)^2 + (\omega \omega_1 / Q_1)^2} \frac{\omega_0 Q_1}{\omega_1 Q_0} P_{\text{in}}$$



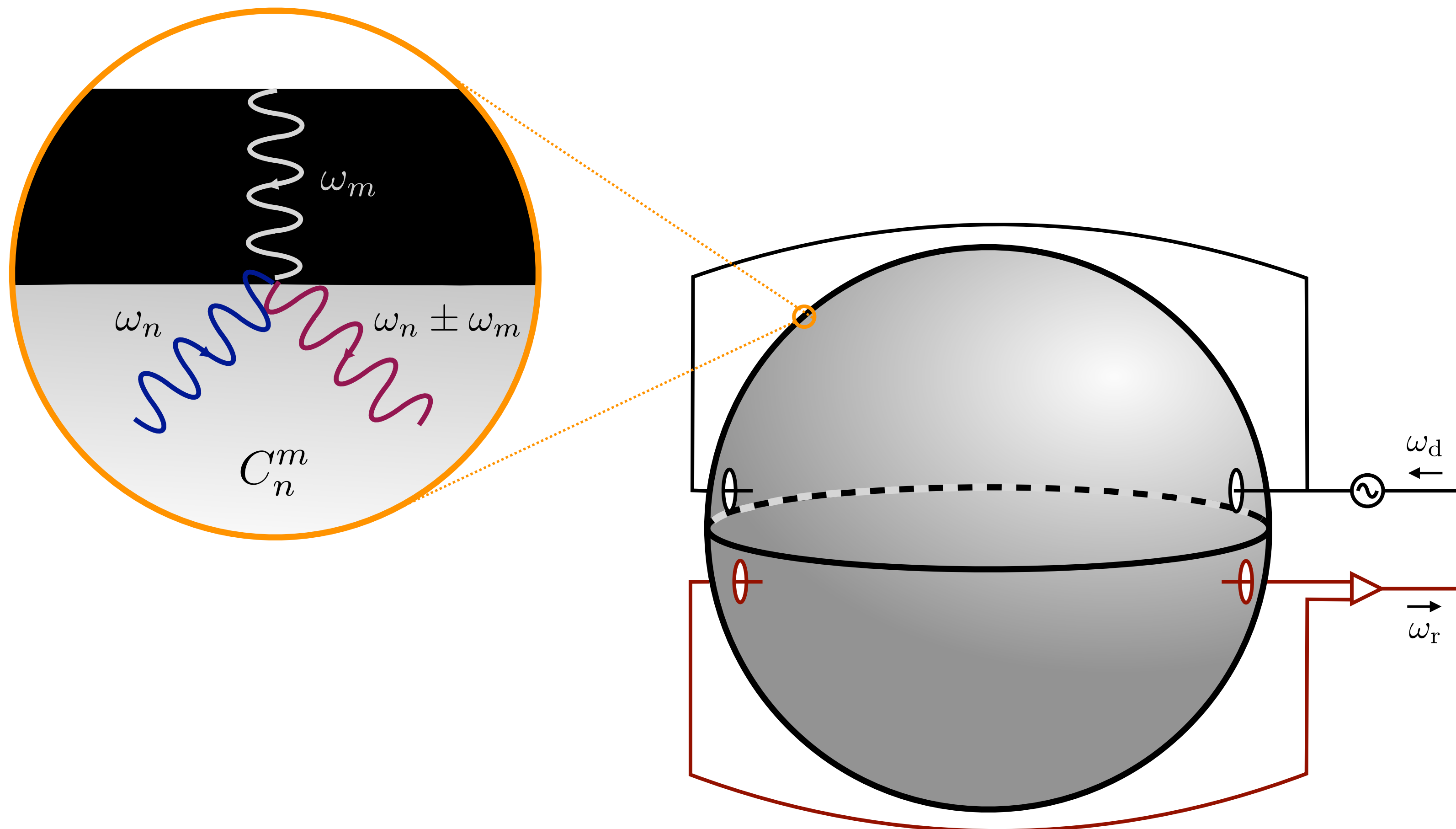
Loading  
&  
Readout



Oscillator

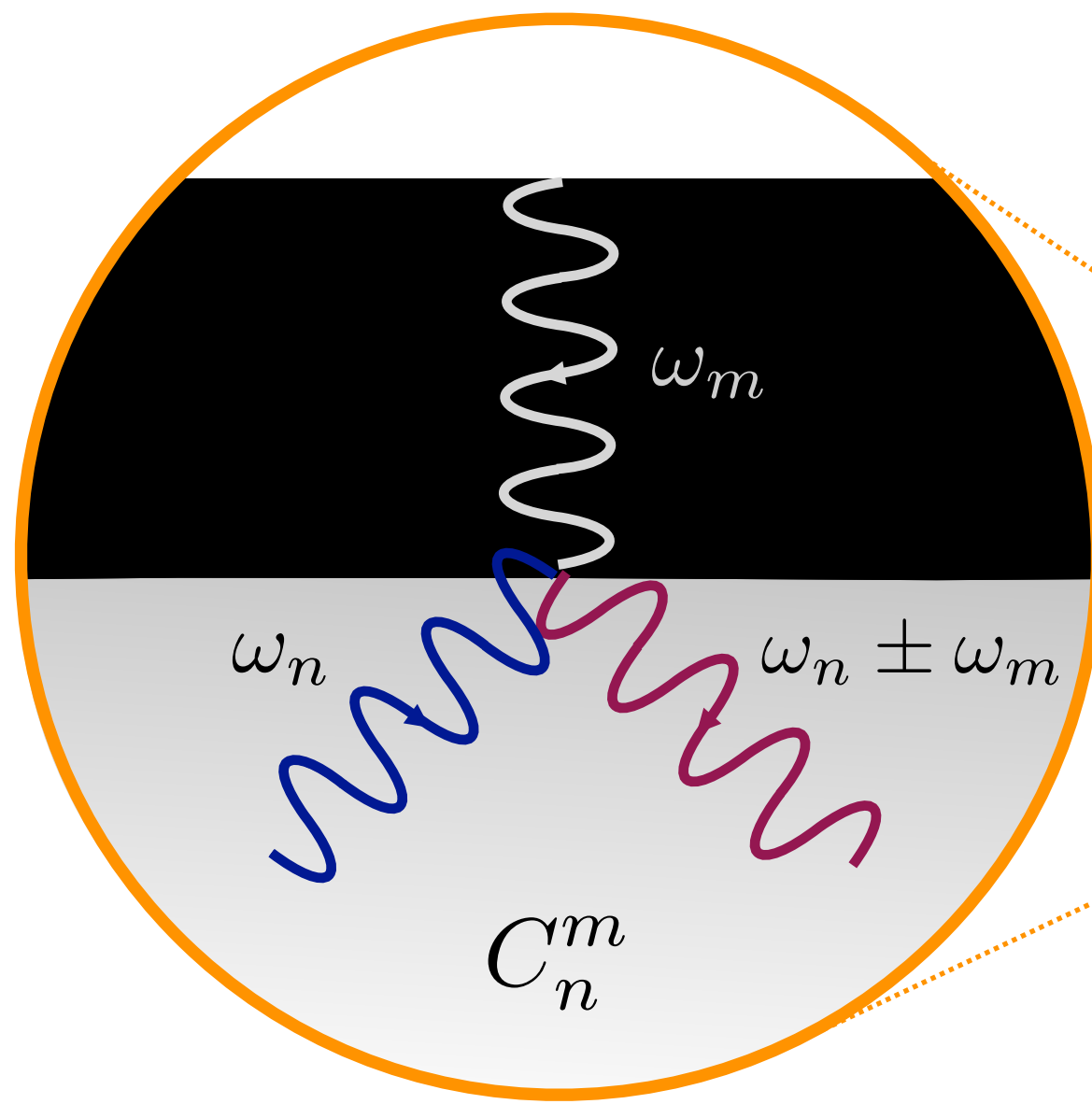
# Mechanical Noise

Vibrations



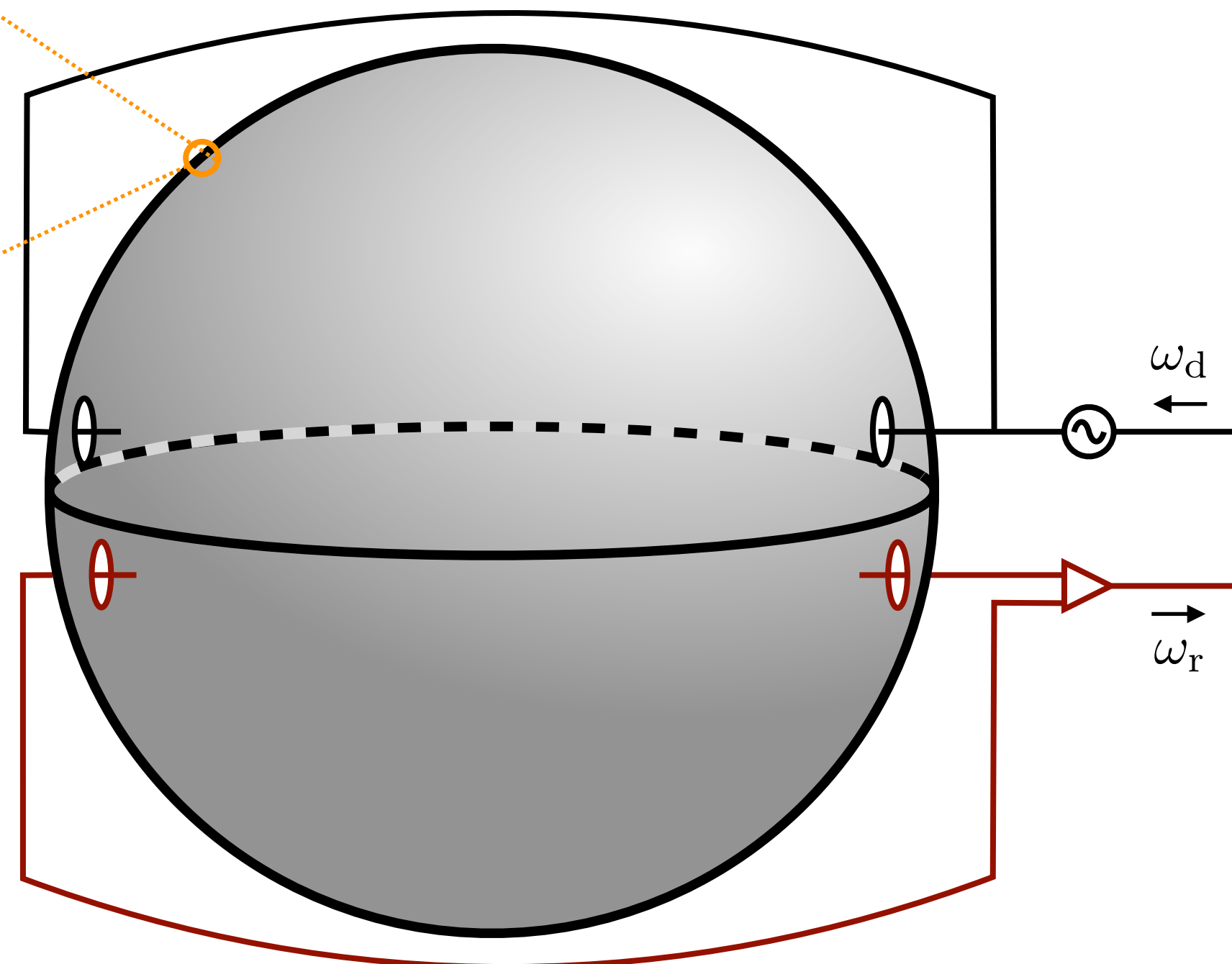
# Mechanical Noise

Vibrations



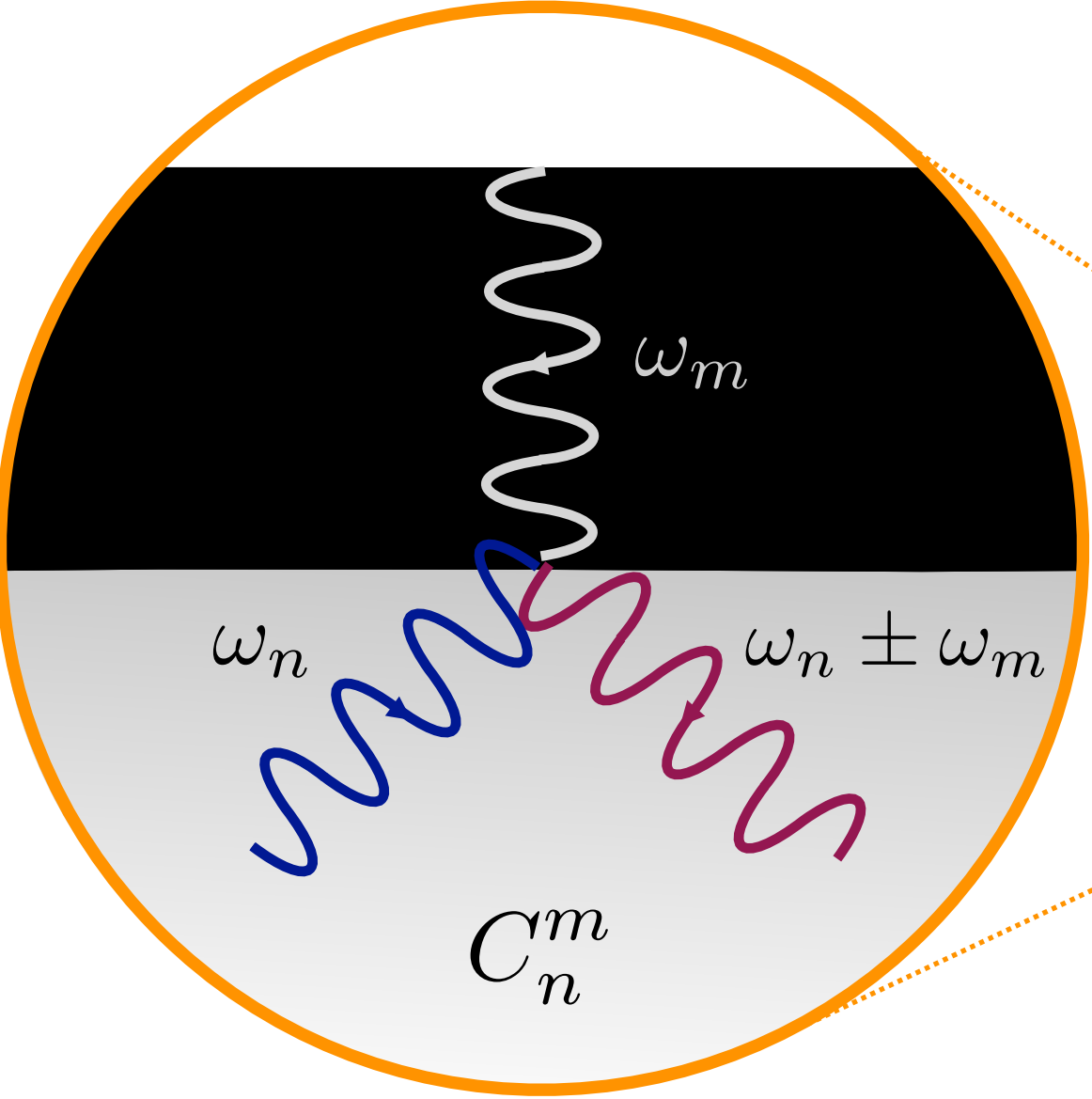
$$S_{\text{mech}}(\omega) \simeq \frac{\epsilon_{1d}^2}{4} \frac{\omega_0}{Q_0} P_{\text{in}} \times$$

$$\sum_{n=0,1} \frac{(S_{q_m}(\omega - \omega_0)/V^{2/3}) (\omega_n/Q_n) \omega_n^4 \omega^2}{[(\omega^2 - \omega_n^2)^2 + (\omega \omega_n/Q_n)^2] [(\omega_0^2 - \omega_n^2)^2 + (\omega_0 \omega_n/Q_n)^2]}$$



# Mechanical Noise

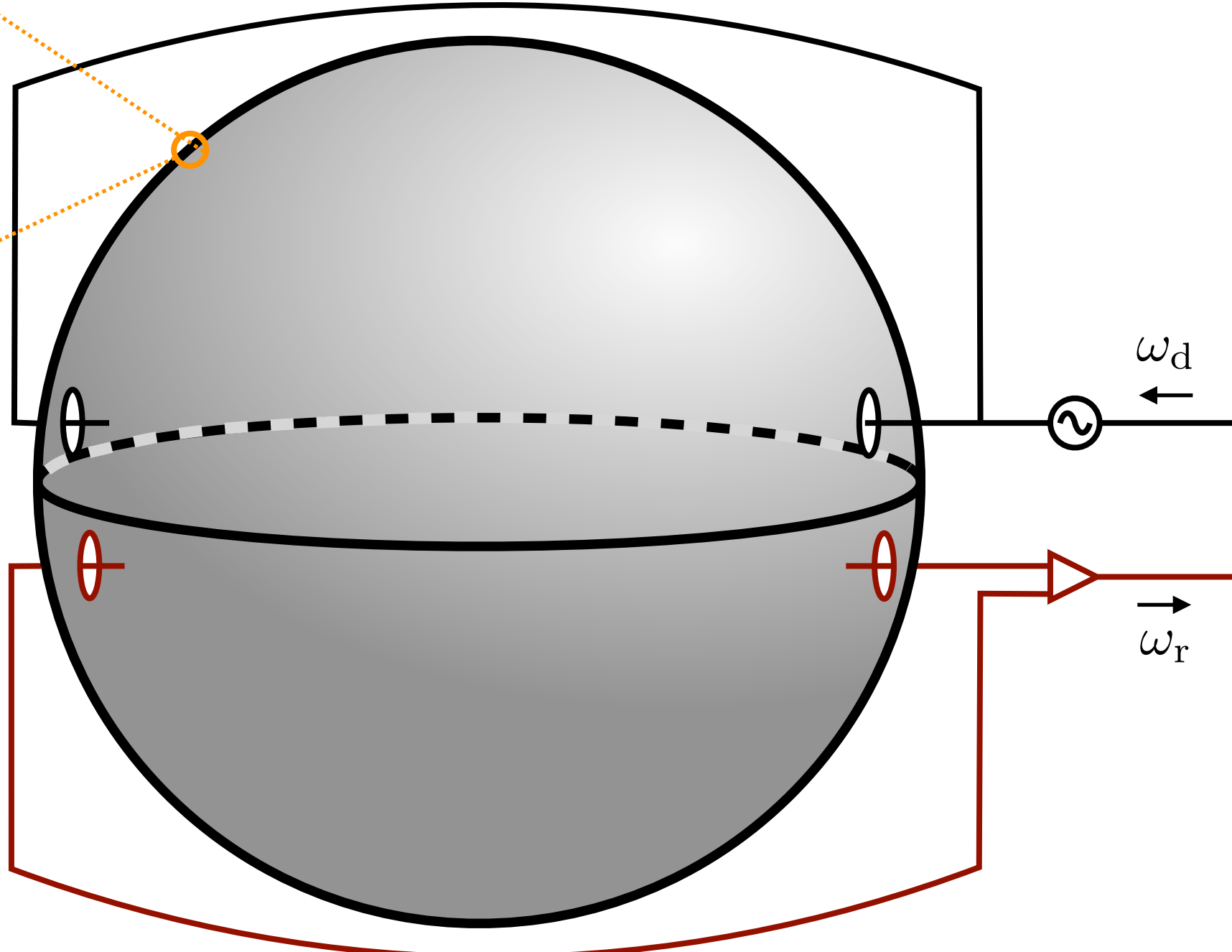
Vibrations



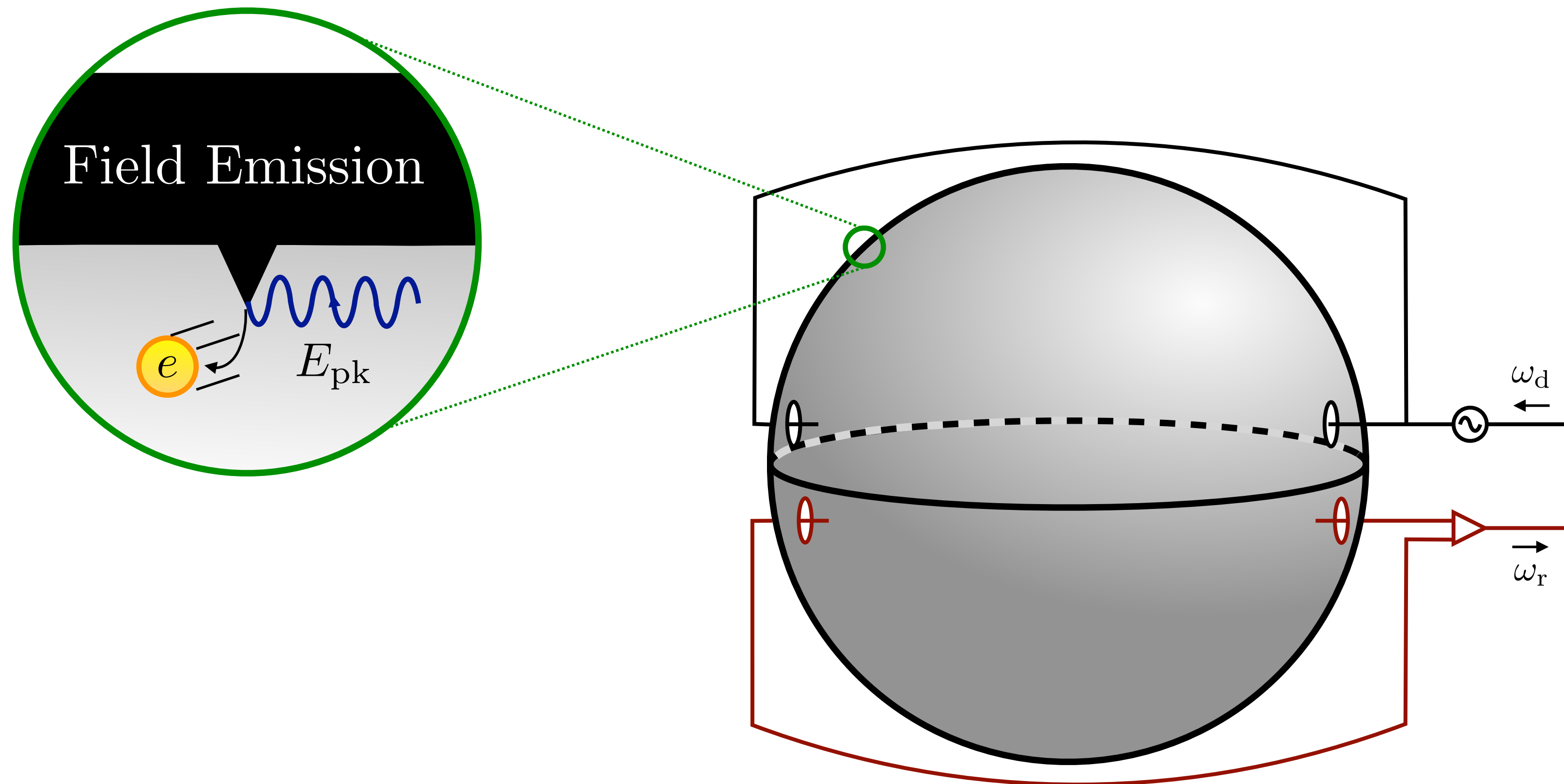
$$S_{\text{mech}}(\omega) \simeq \frac{\epsilon_{1d}^2}{4} \frac{\omega_0}{Q_0} P_{\text{in}} \times \sum_{n=0,1} \frac{(S_{q_m}(\omega - \omega_0)/V^{2/3}) (\omega_n/Q_n) \omega_n^4 \omega^2}{[(\omega^2 - \omega_n^2)^2 + (\omega \omega_n/Q_n)^2] [(\omega_0^2 - \omega_n^2)^2 + (\omega_0 \omega_n/Q_n)^2]}$$

Displacement PSD:

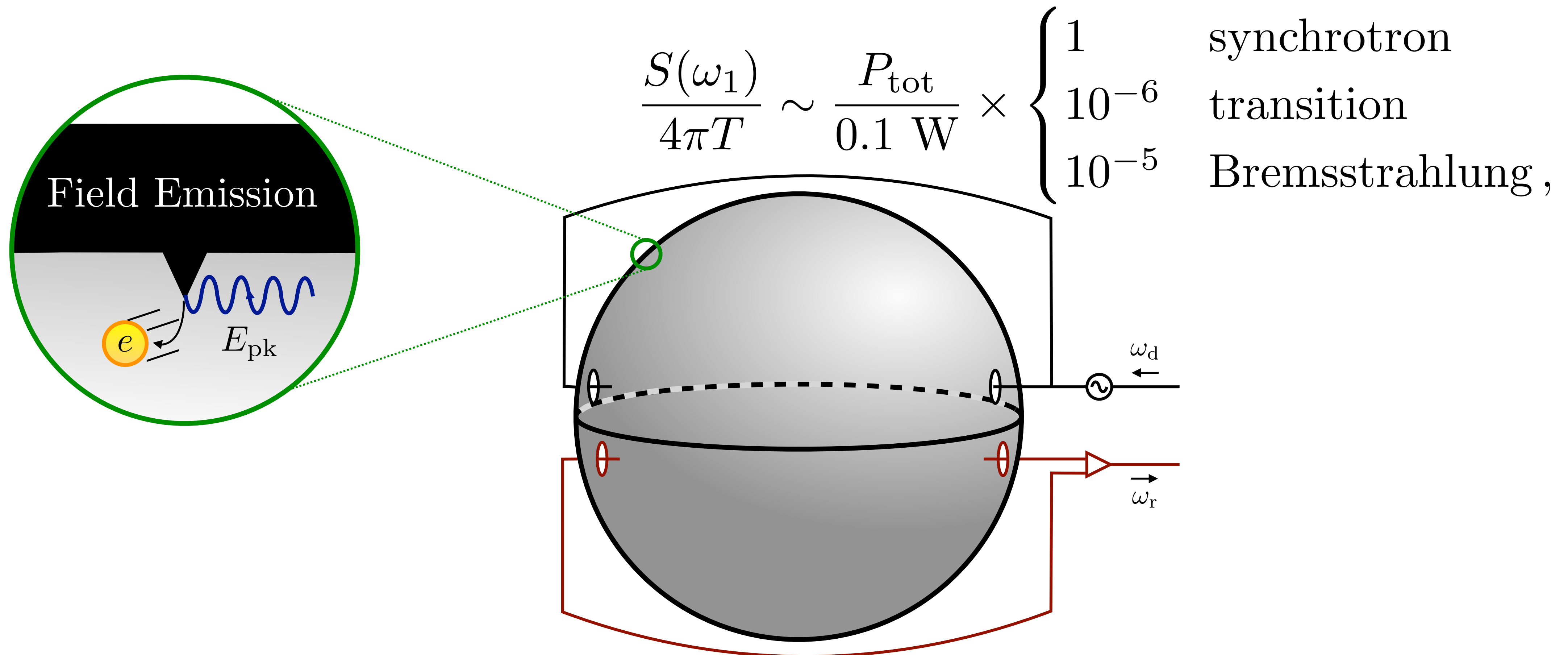
$$S_{q_m}(\omega) \simeq \frac{1}{M^2} \frac{S_{f_m}(\omega)}{(\omega^2 - \omega_m^2)^2 + (\omega_m \omega/Q_m)^2}$$



# Non-standard Noise Sources: Field Emission

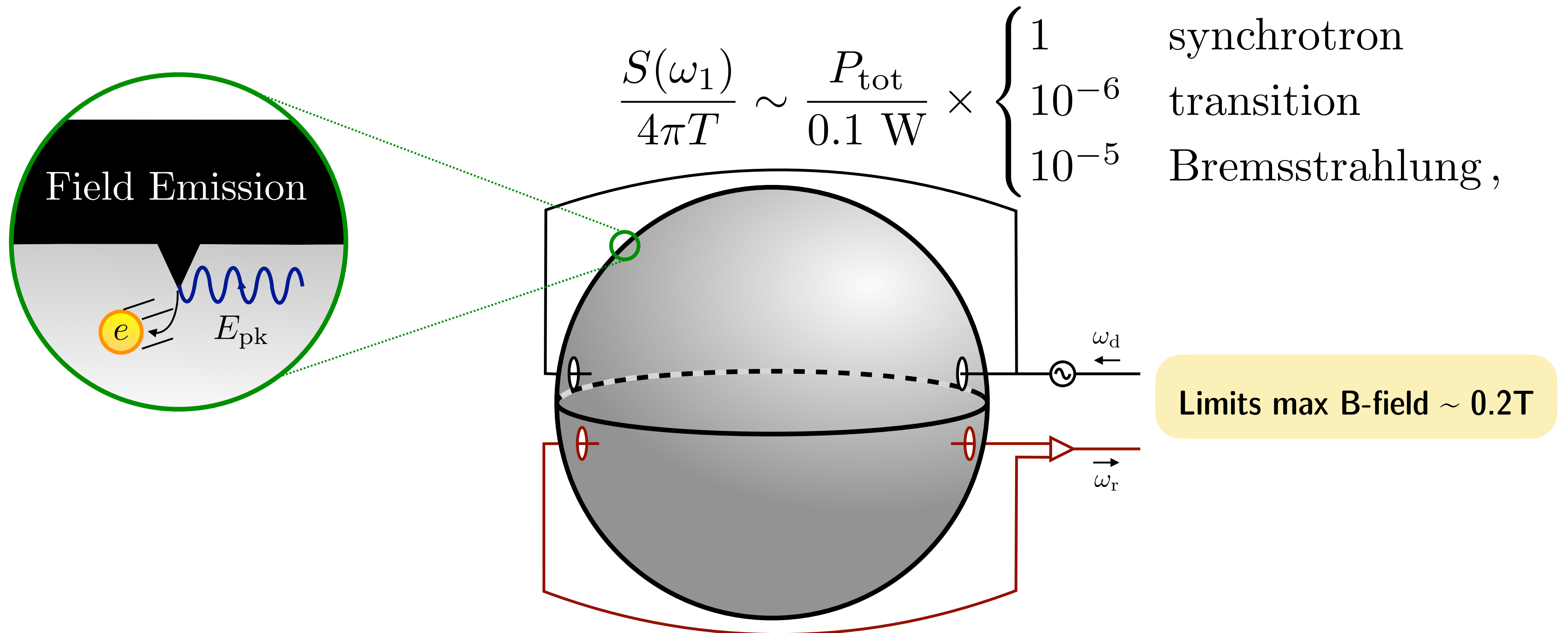


# Non-standard Noise Sources: Field Emission

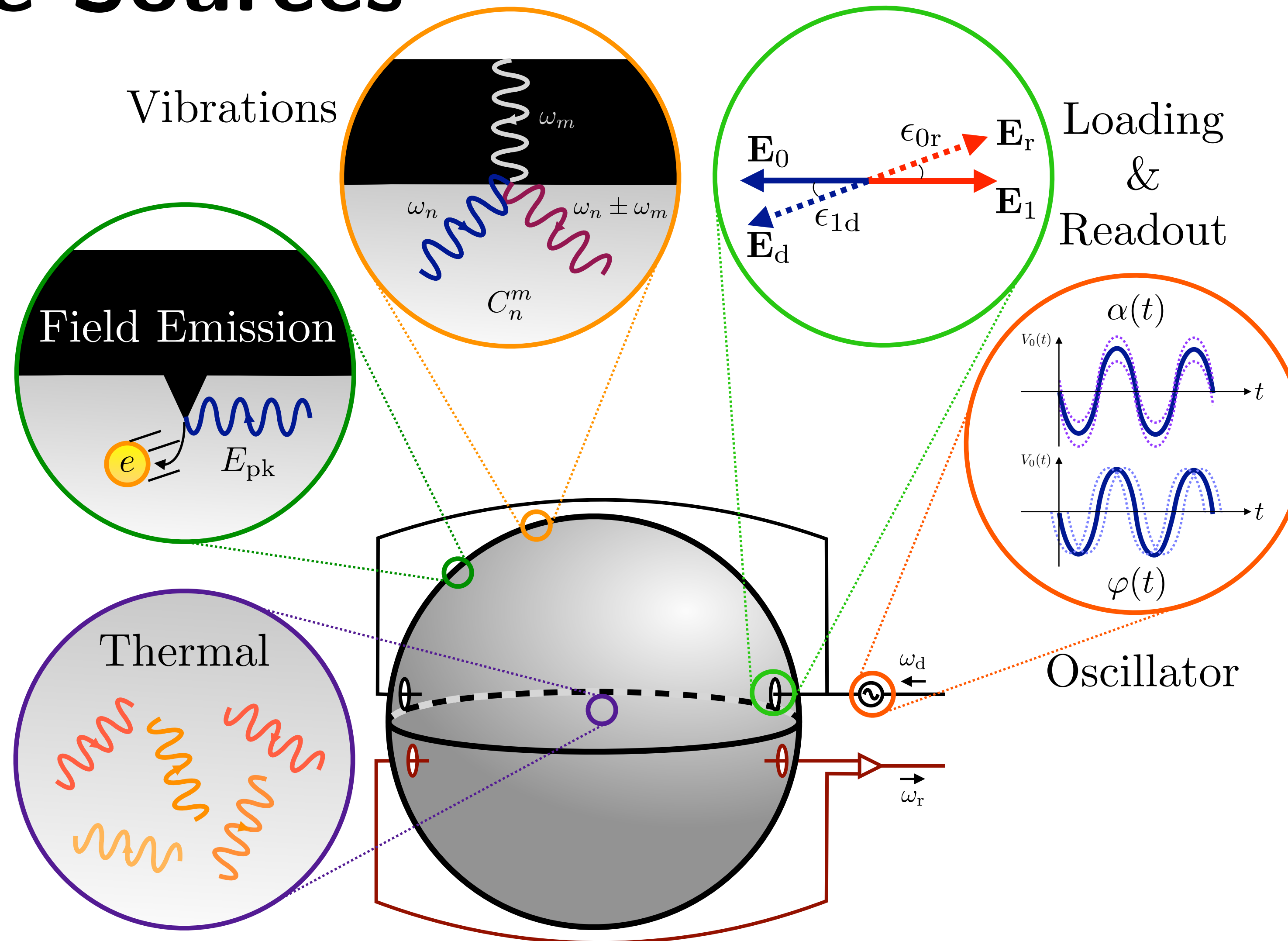




# Non-standard Noise Sources: Field Emission

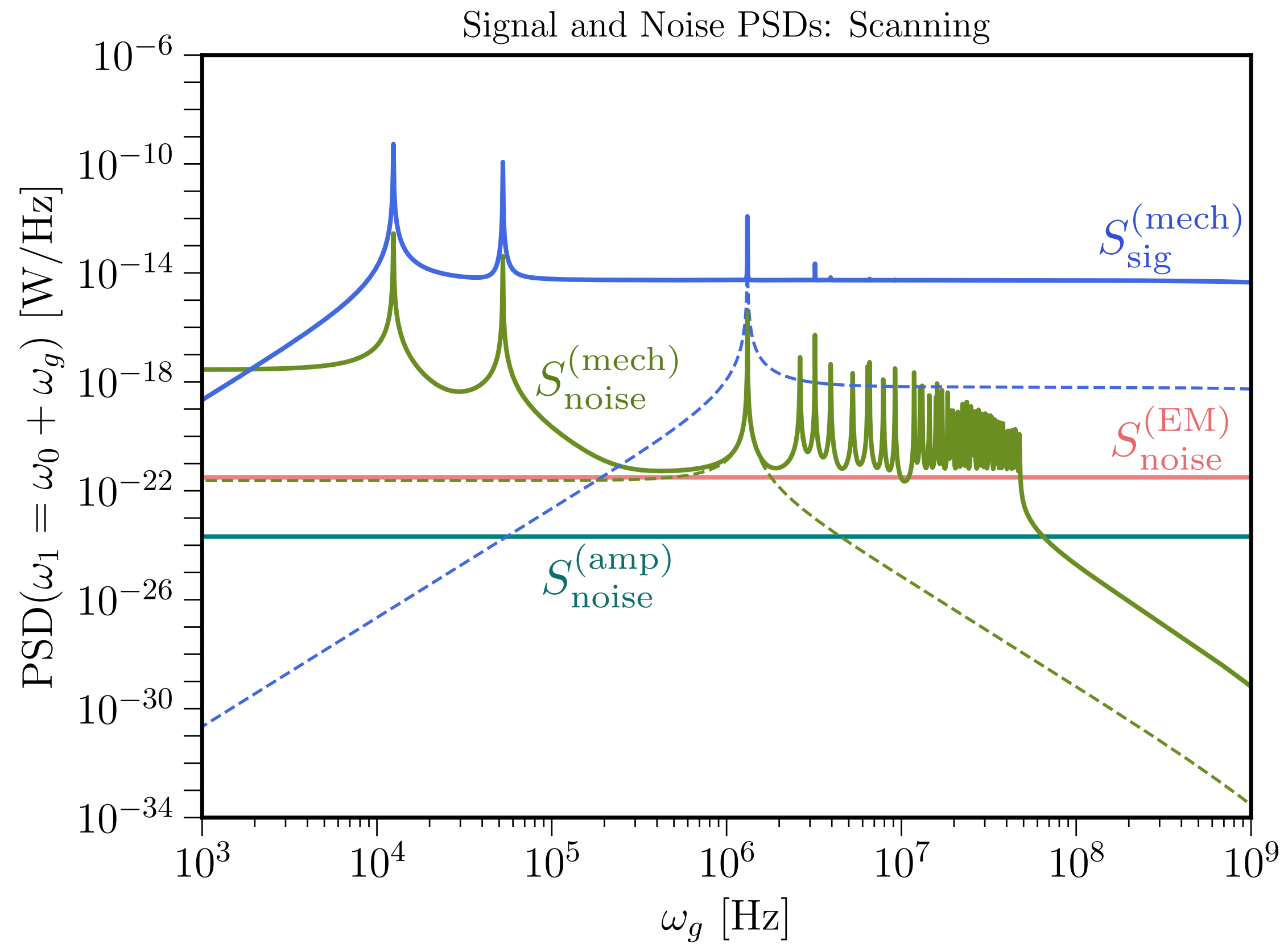


# All Noise Sources

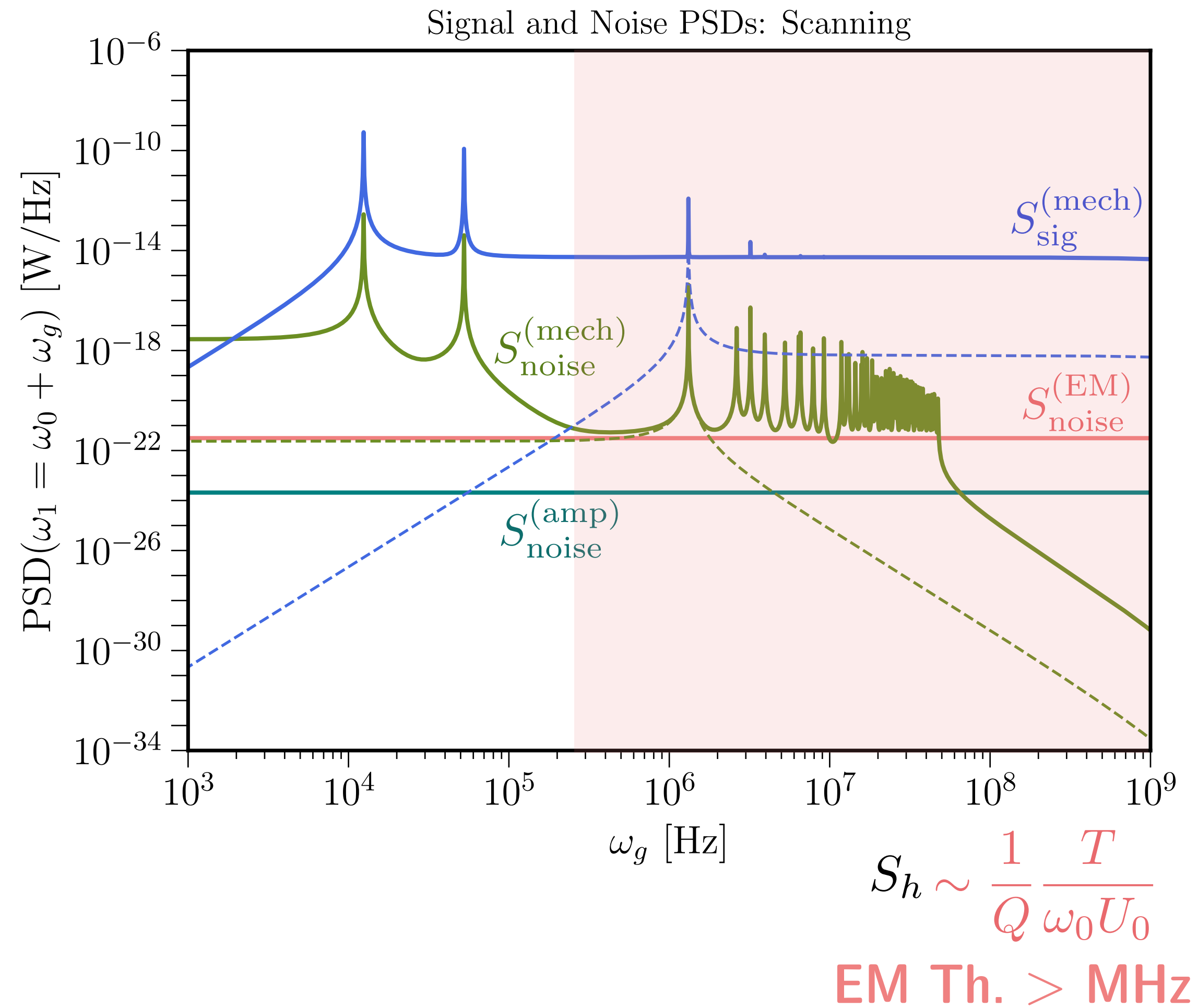




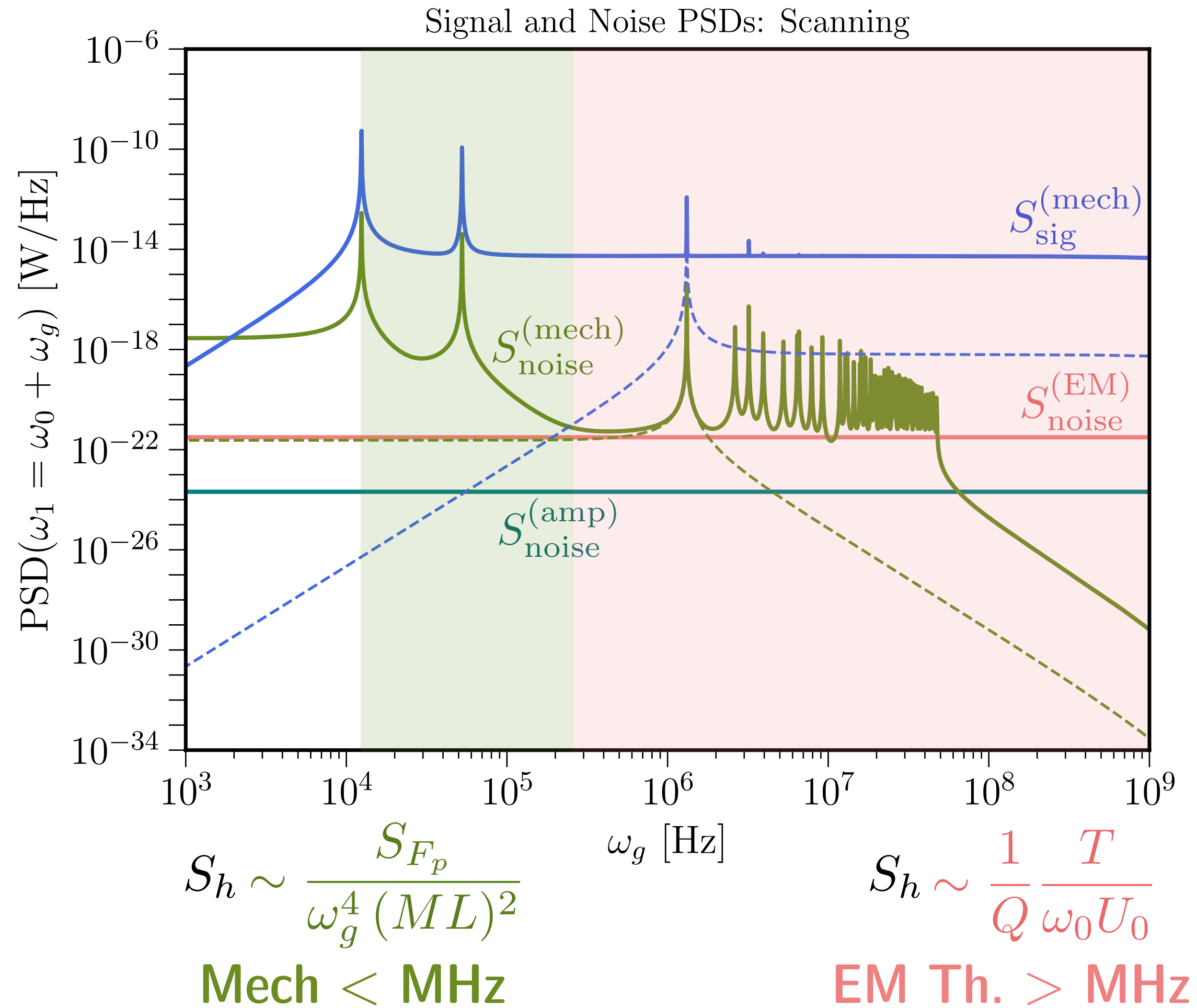
# Noise in MAGO 2.0



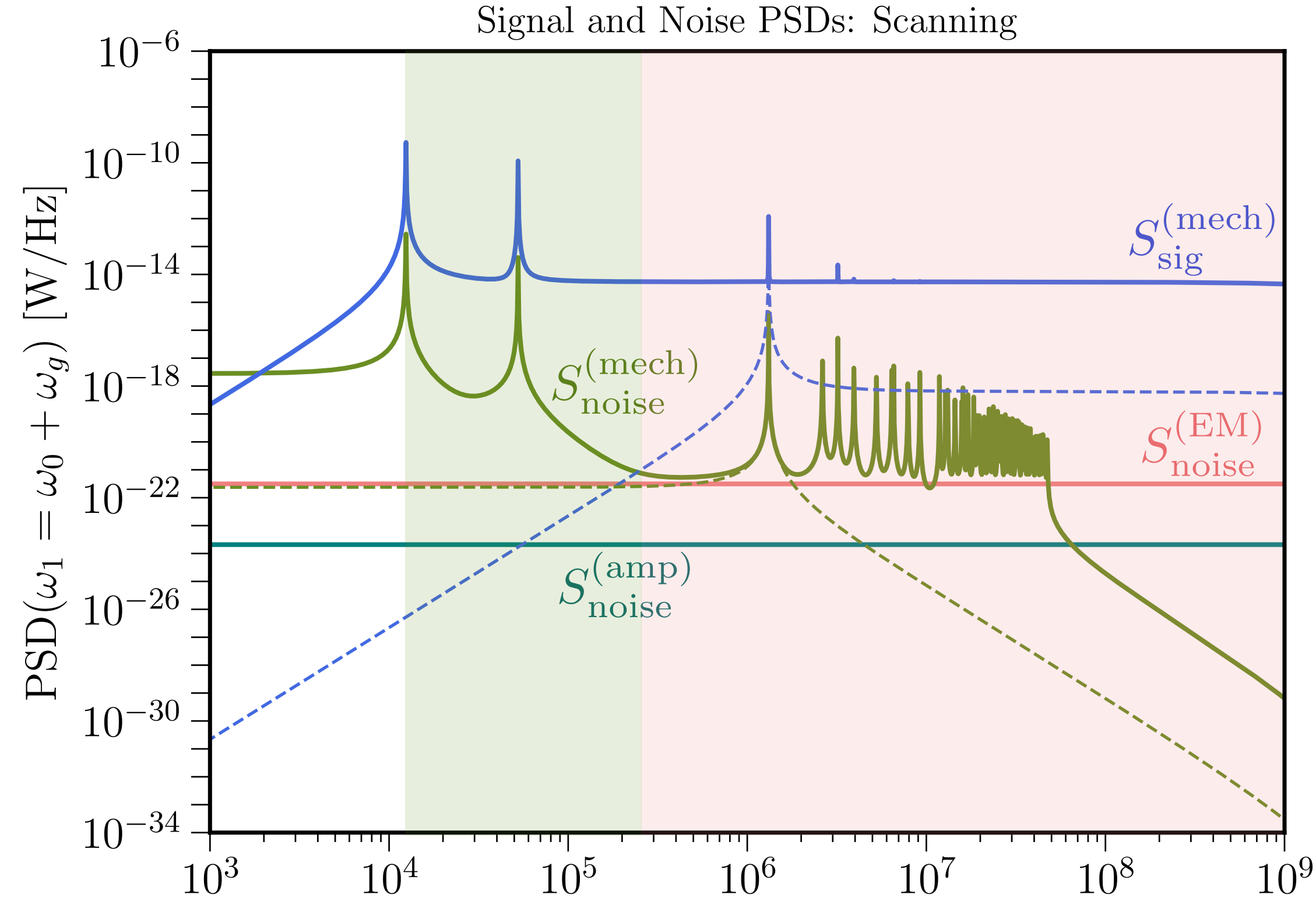
# Noise in MAGO 2.0



# Noise in MAGO 2.0



# Noise in MAGO 2.0

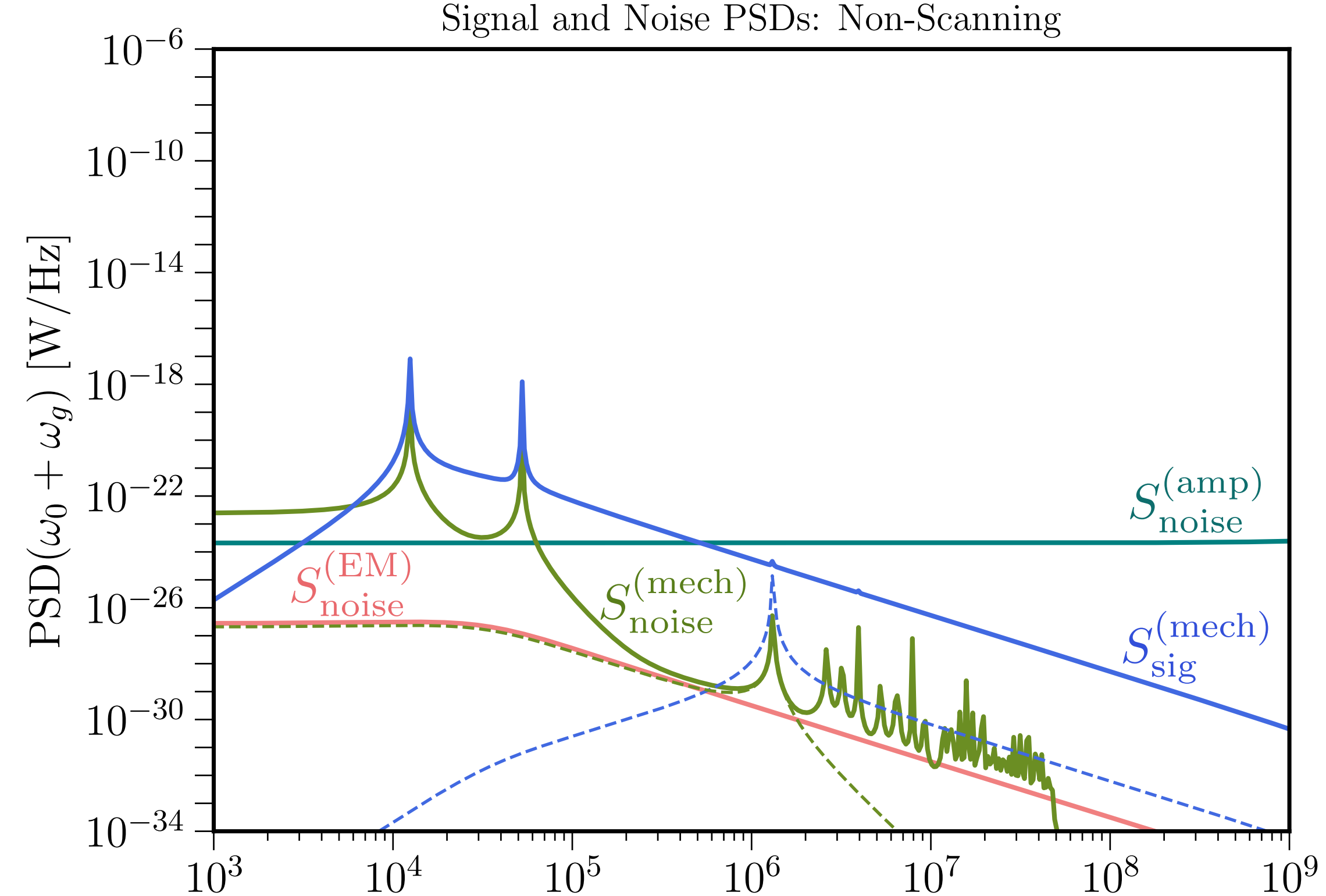


$$S_h \sim \frac{S_{F_p}}{\omega_g^4 (ML)^2}$$

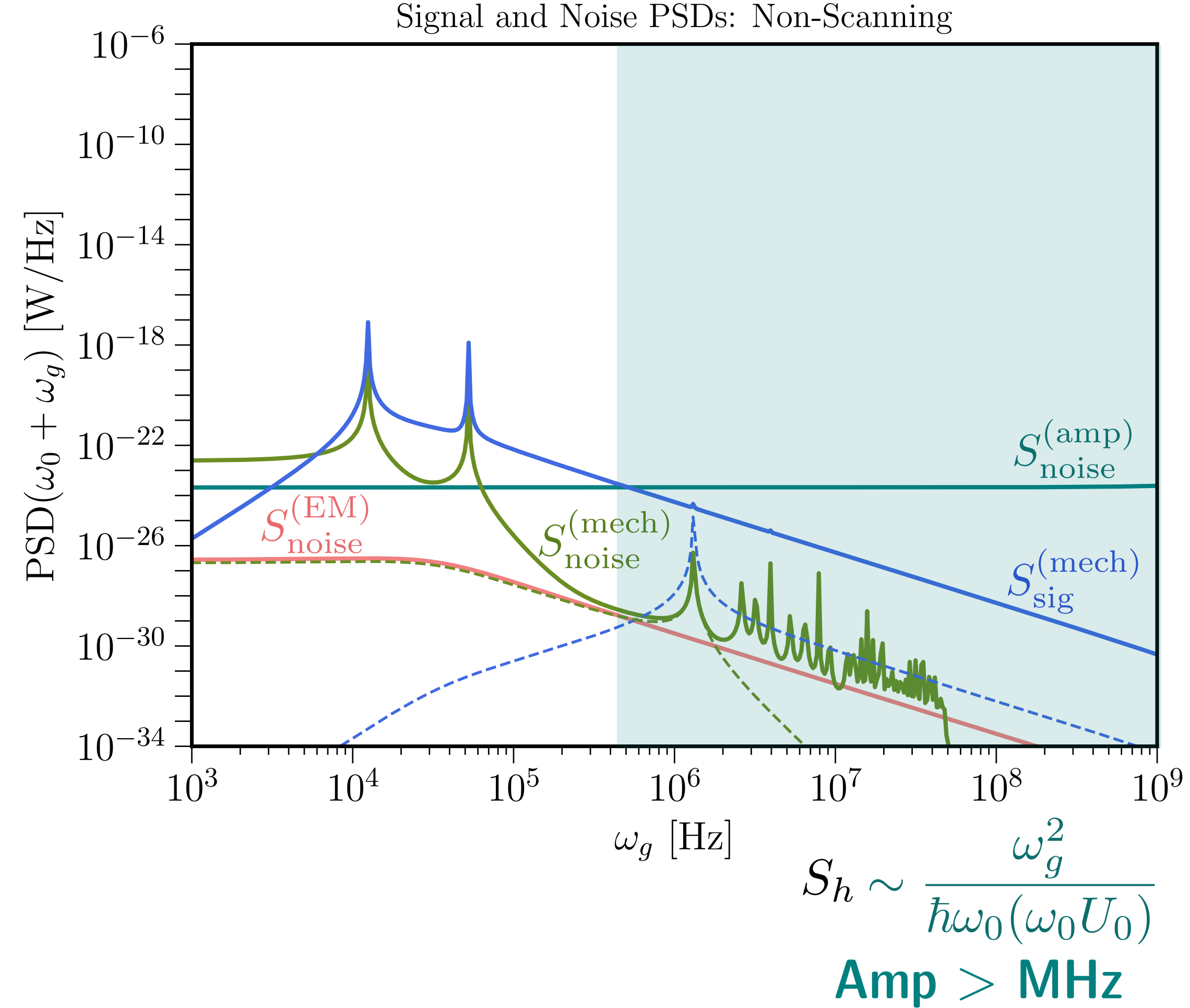
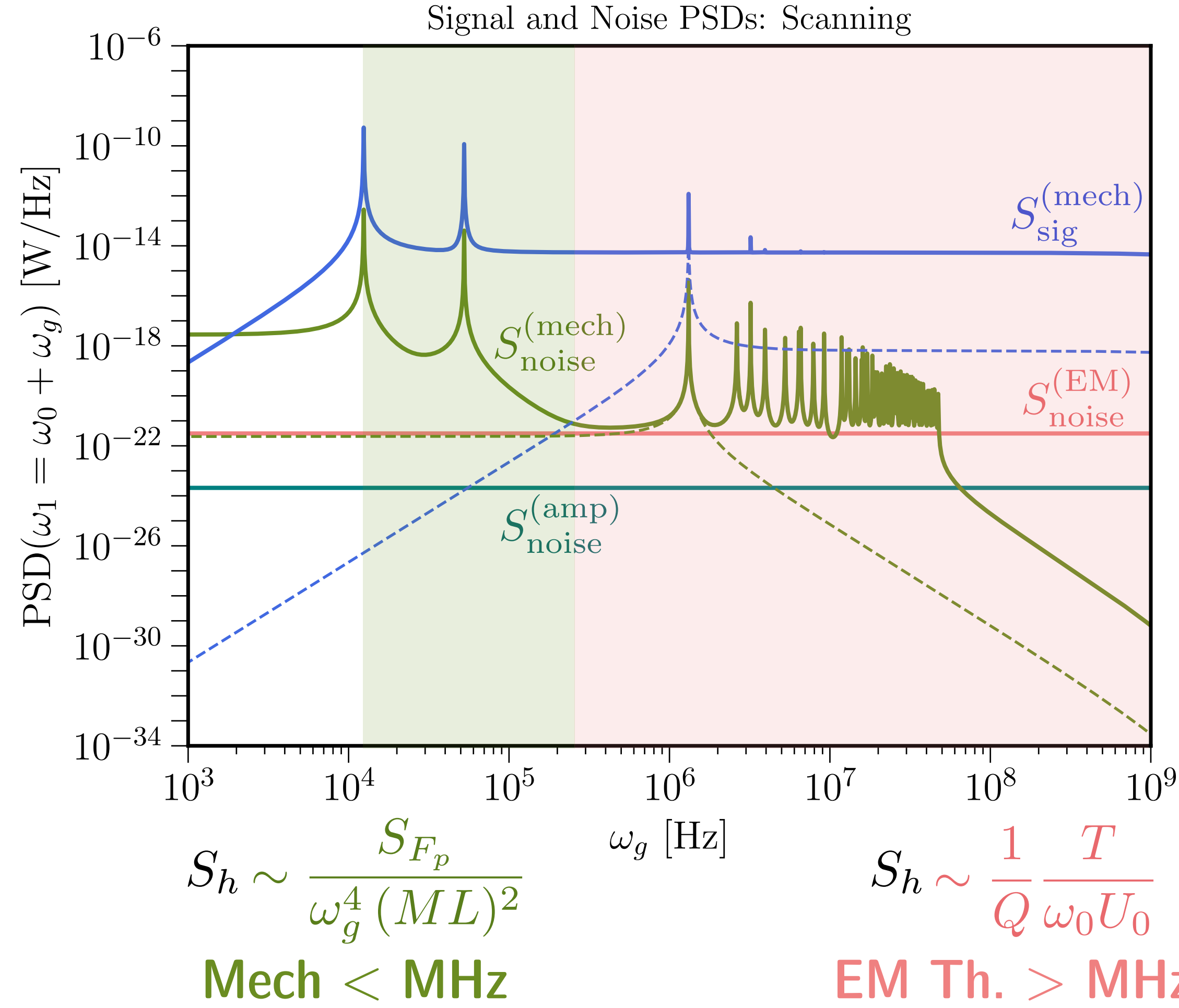
Mech < MHz

$$S_h \sim \frac{1}{Q} \frac{T}{\omega_0 U_0}$$

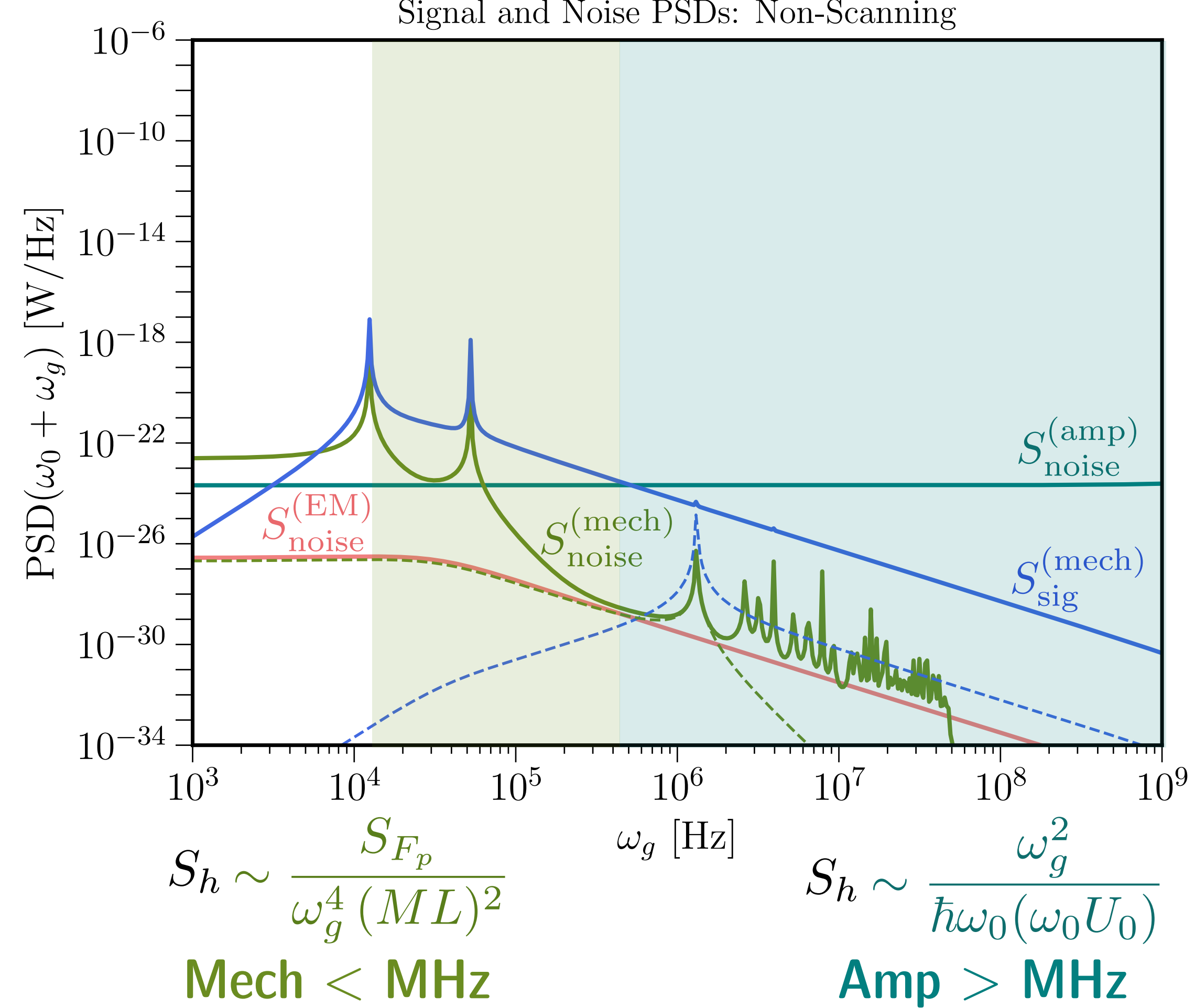
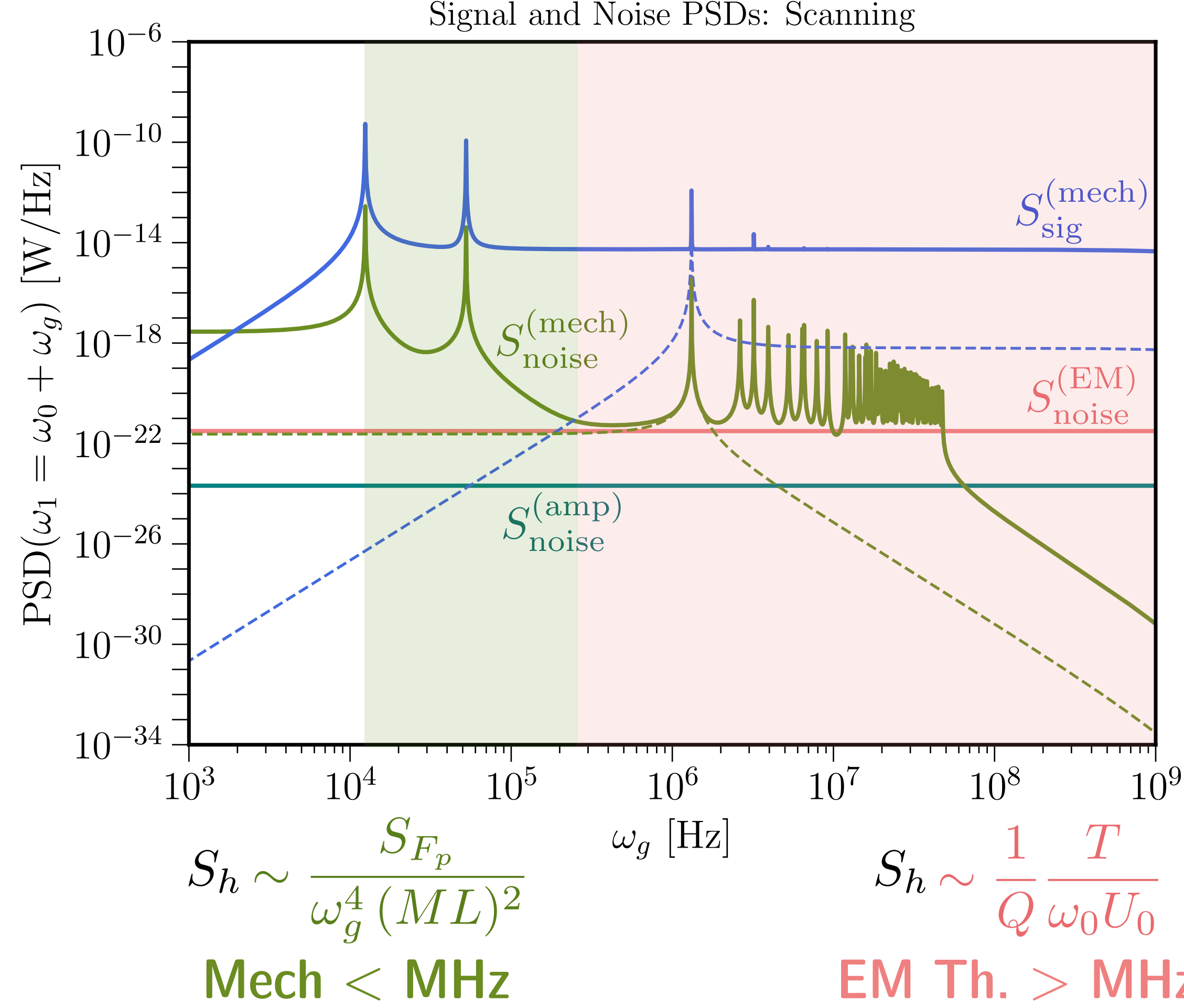
EM Th. > MHz



# Noise in MAGO 2.0

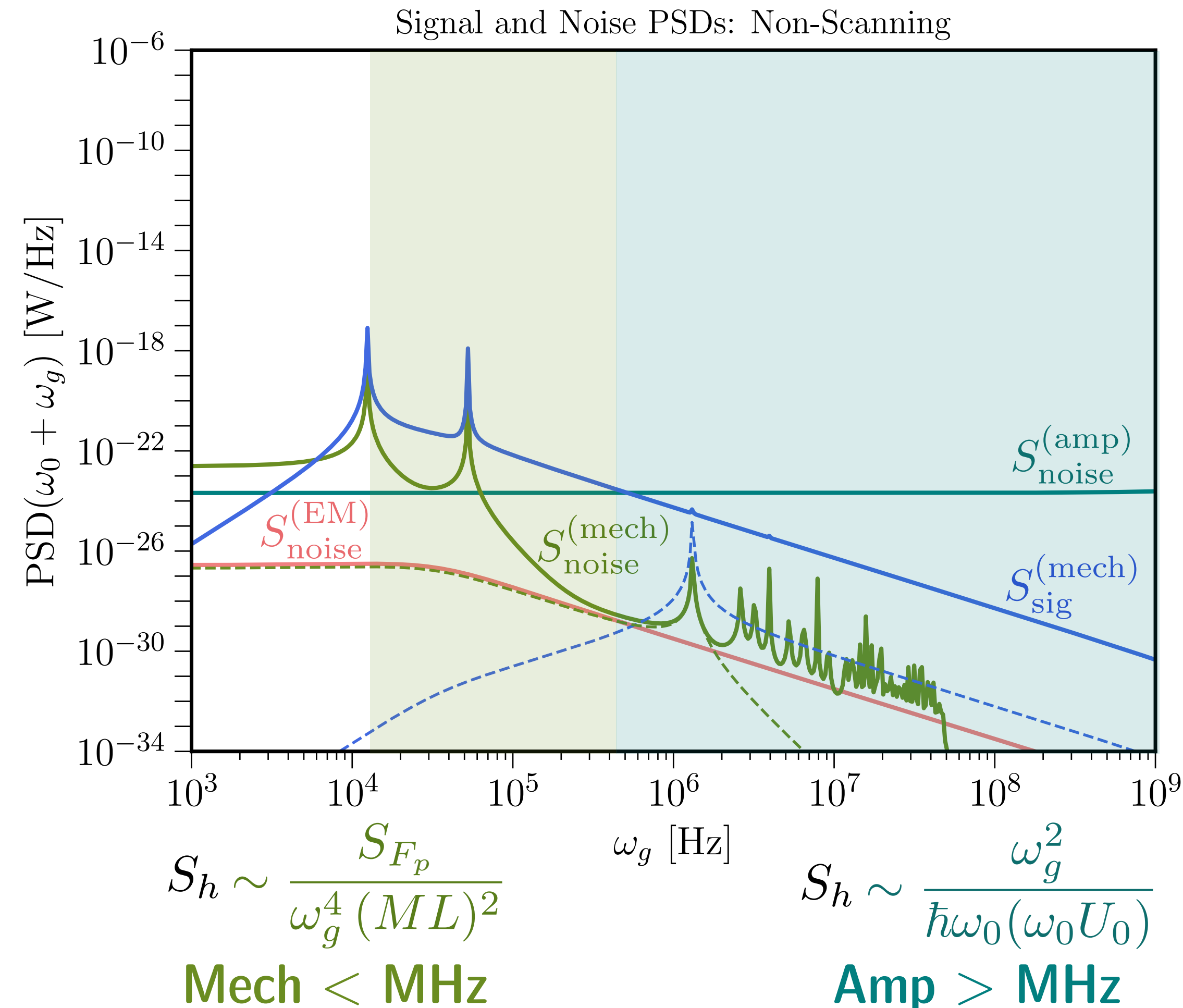
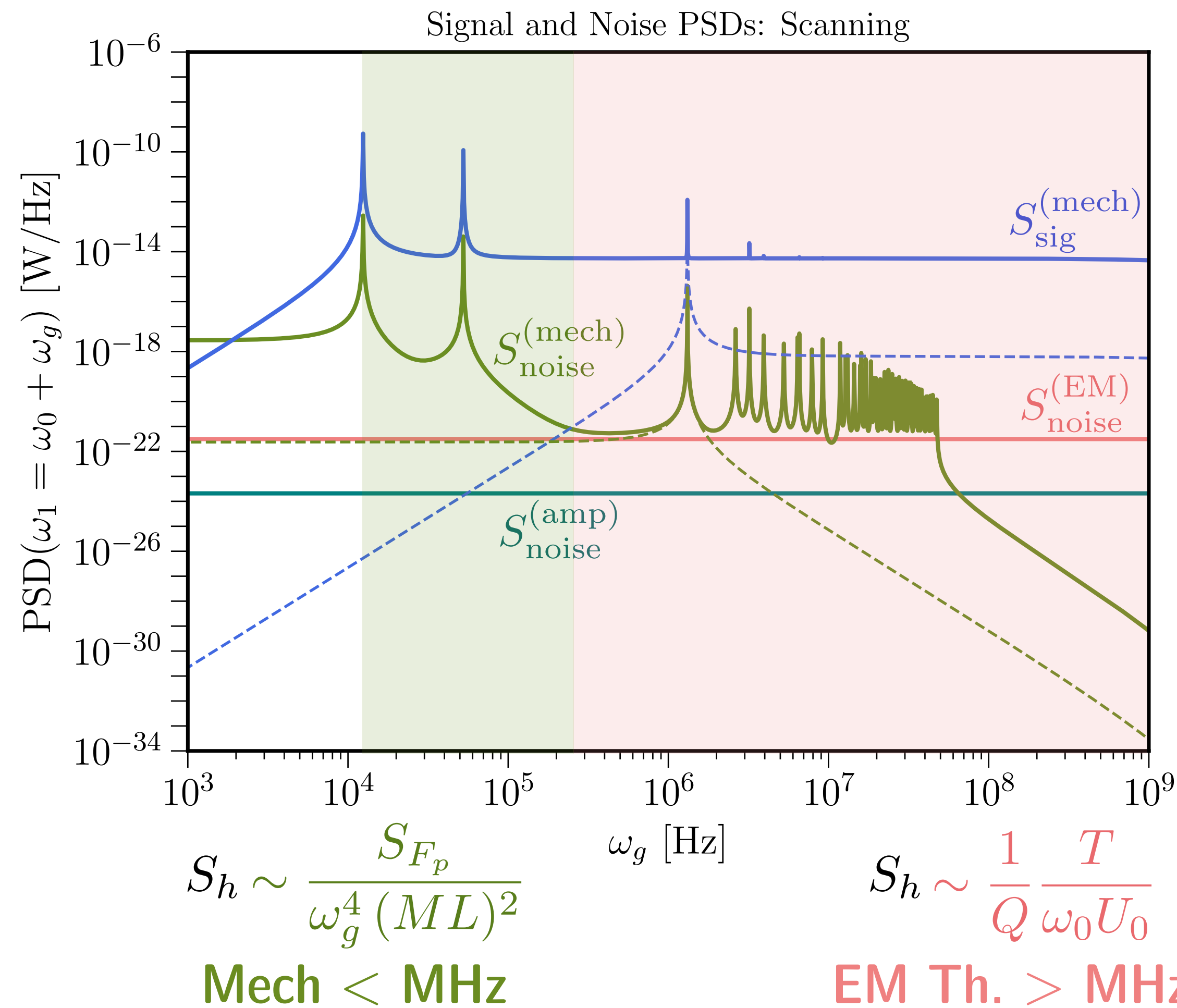


# Noise in MAGO 2.0





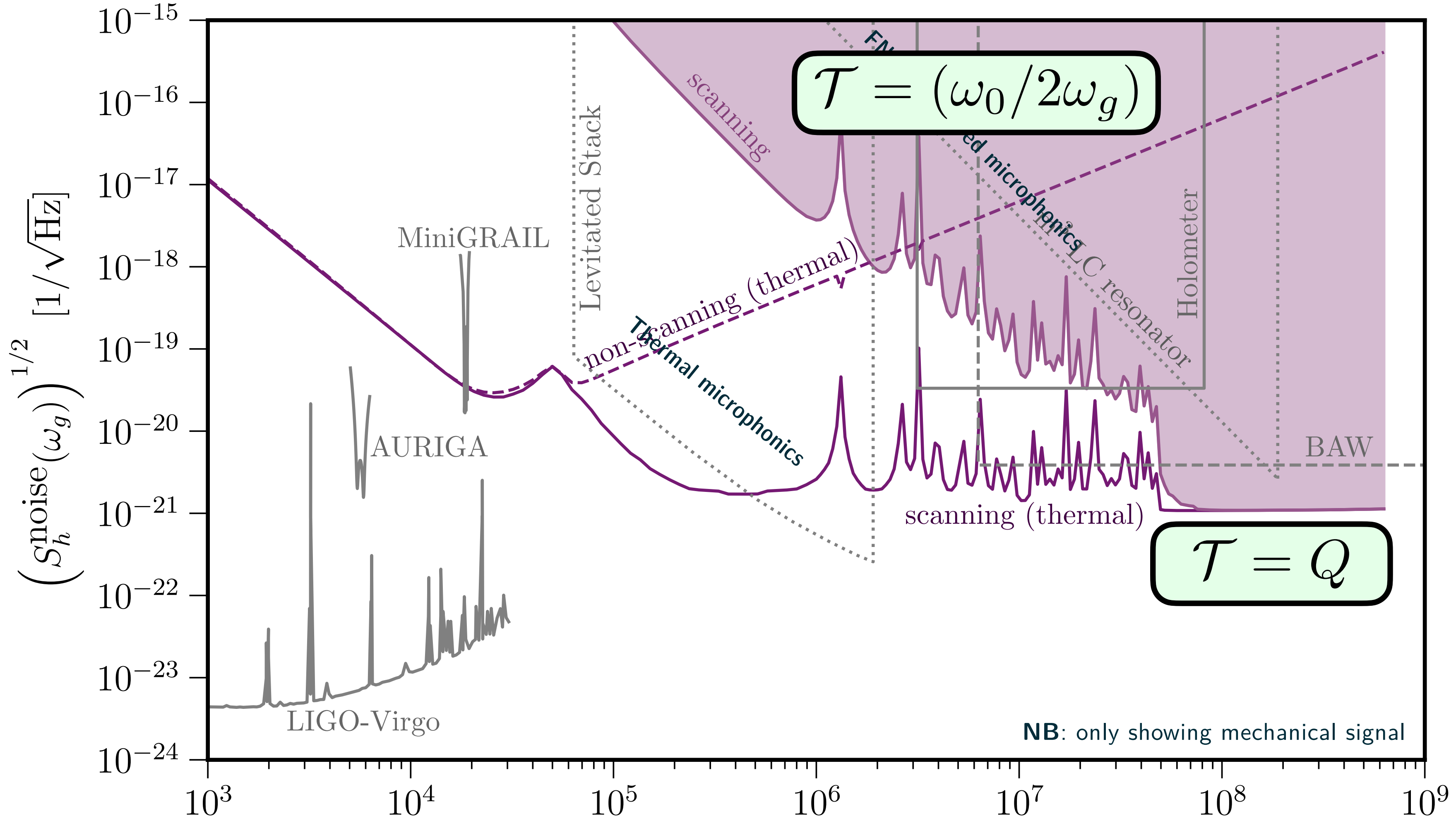
# Noise in MAGO 2.0



**NB:** missing radiation damping effect studied in Löwenberg, Moortgat-Pick: 2307.14379

# Noise-equivalent strain

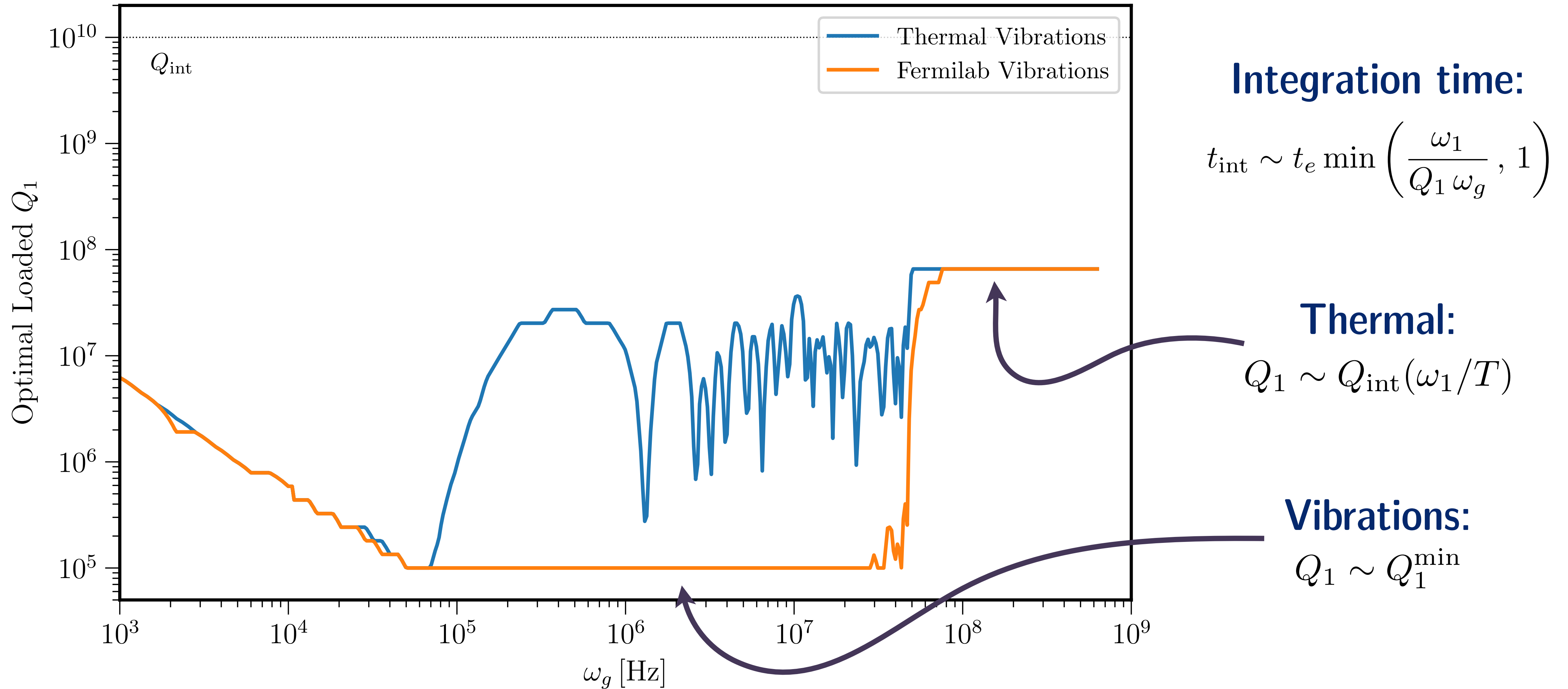
Berlin, Blas, D'Agnolo, SARE, Harnik, Kahn, Schutte-Engel, Wentzel (PRD 2023)



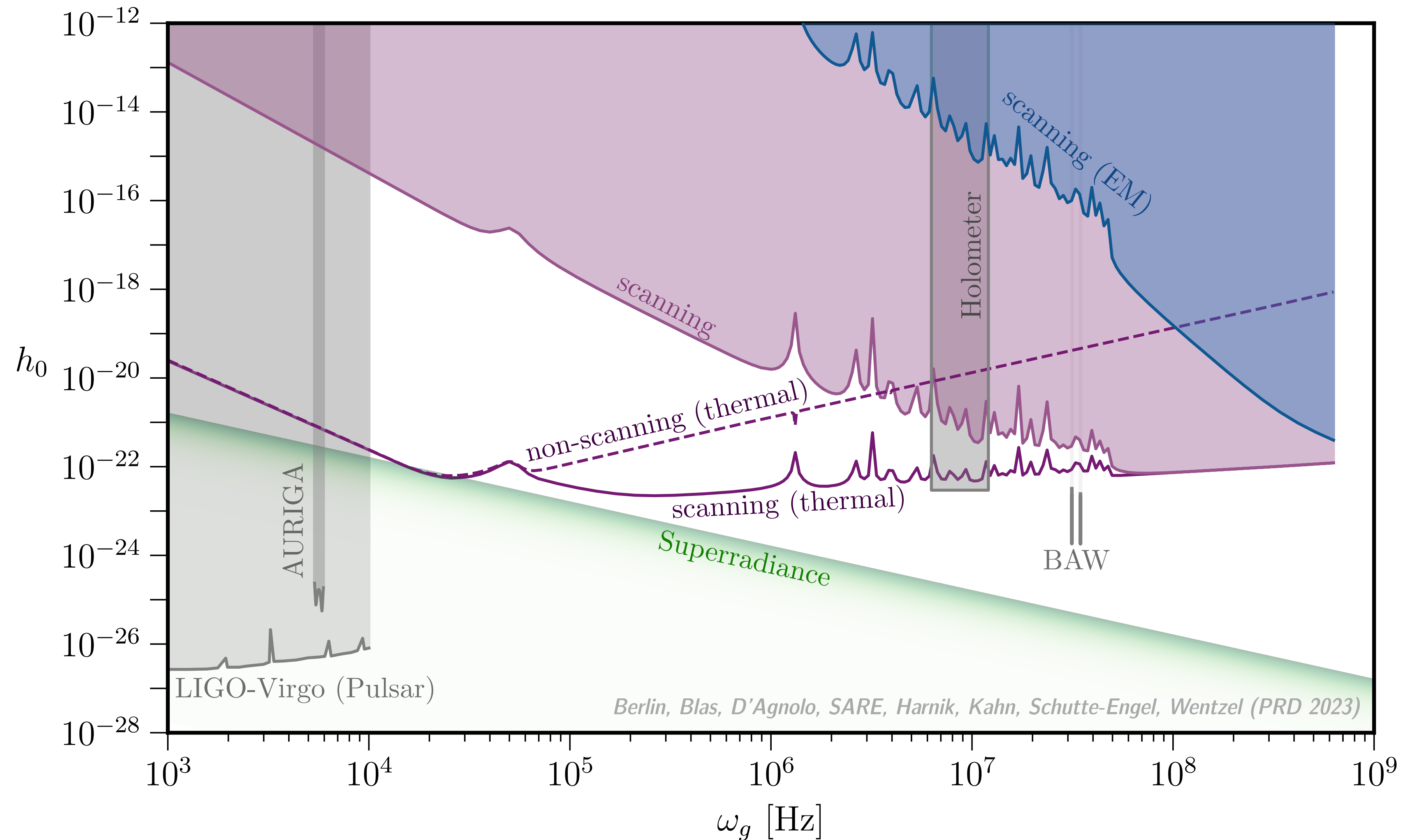
NB: missing radiation damping effect studied in Löwenberg, Moortgat-Pick: 2307.14379  $\omega_g [\text{Hz}]$



# Optimal Scanning



# MAGO 2.0 sensitivity to coherent GWs

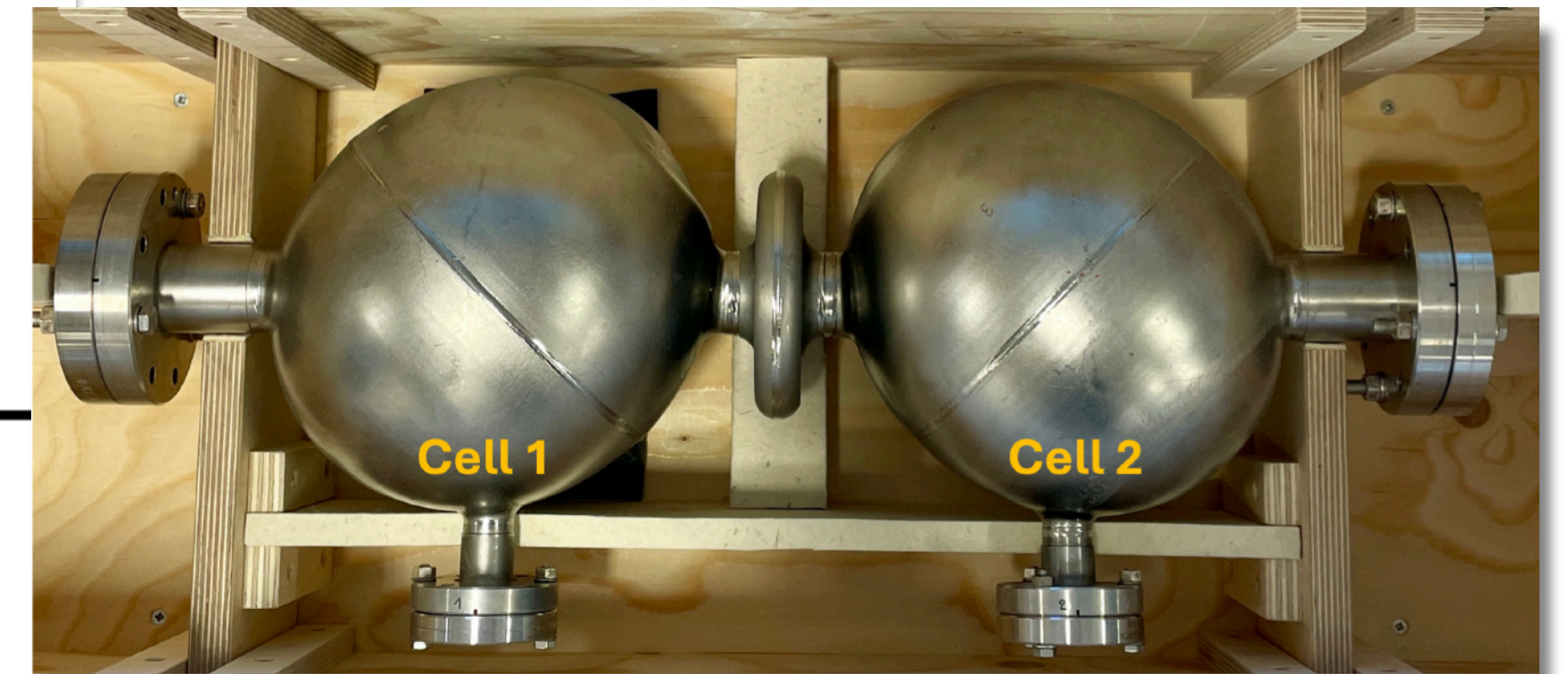


# From Proposal to Practice

**First characterisation of the MAGO cavity, a superconducting RF detector for kHz-MHz gravitational waves**

Lars Fischer<sup>a</sup> Bianca Giaccone<sup>b</sup> Ivan Gonin<sup>b</sup> Anna Grassellino<sup>b</sup> Wolfgang Hillert<sup>a</sup>  
Timergali Khabiboulline<sup>b</sup> Tom Krokotsch<sup>a</sup> Gudrid Moortgat-Pick<sup>a,c</sup> Andrea Muhs<sup>c</sup>  
Yuriy Orlov<sup>b</sup> Krisztian Peters<sup>c</sup> Sam Posen<sup>b</sup> Oleg Pronitchev<sup>b</sup> Marc Wenskat<sup>a,c</sup>

<https://arxiv.org/abs/2411.18346>

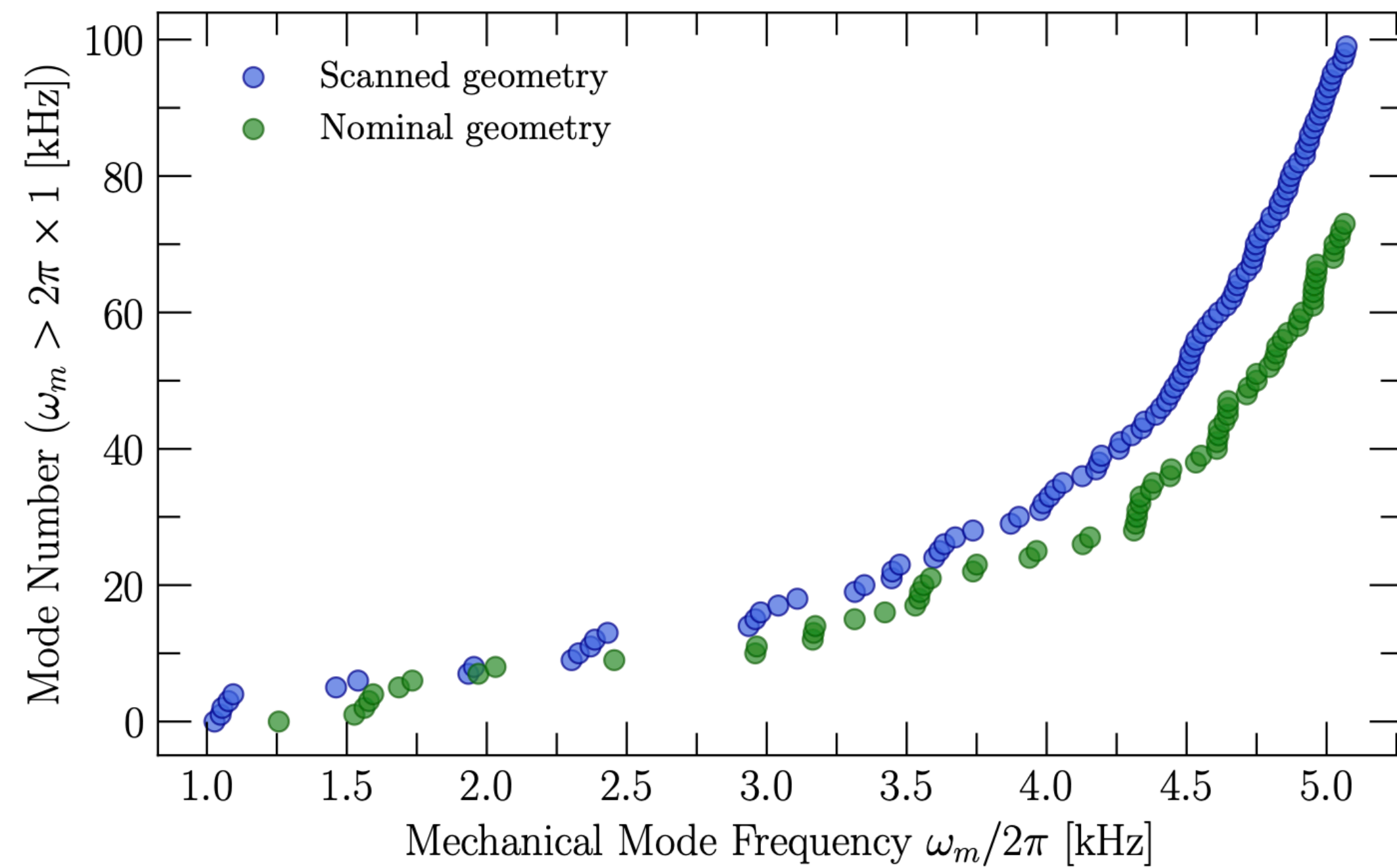


**Ongoing collaboration between INFN/DESY/FNAL**



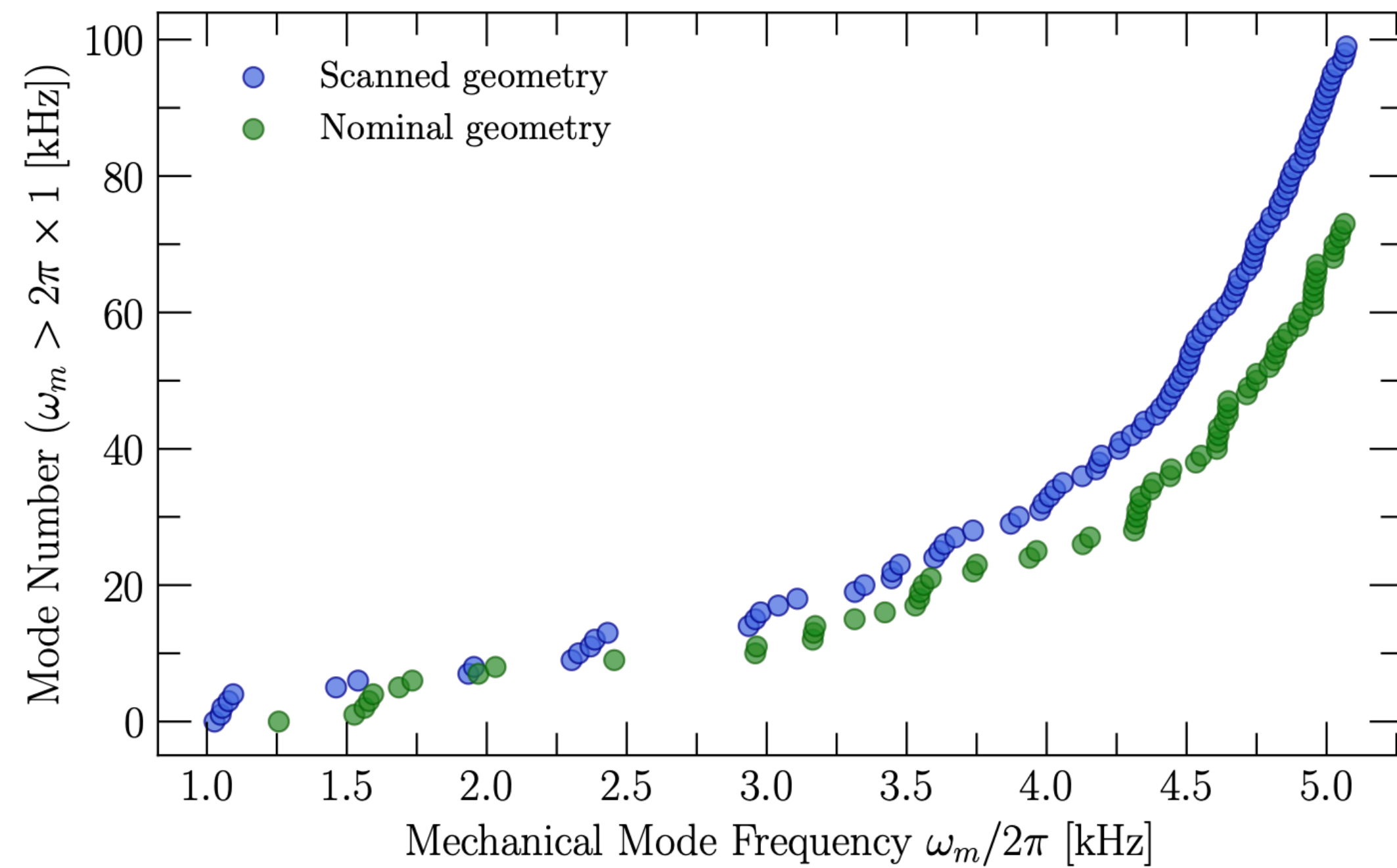
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# Characterising the Prototype

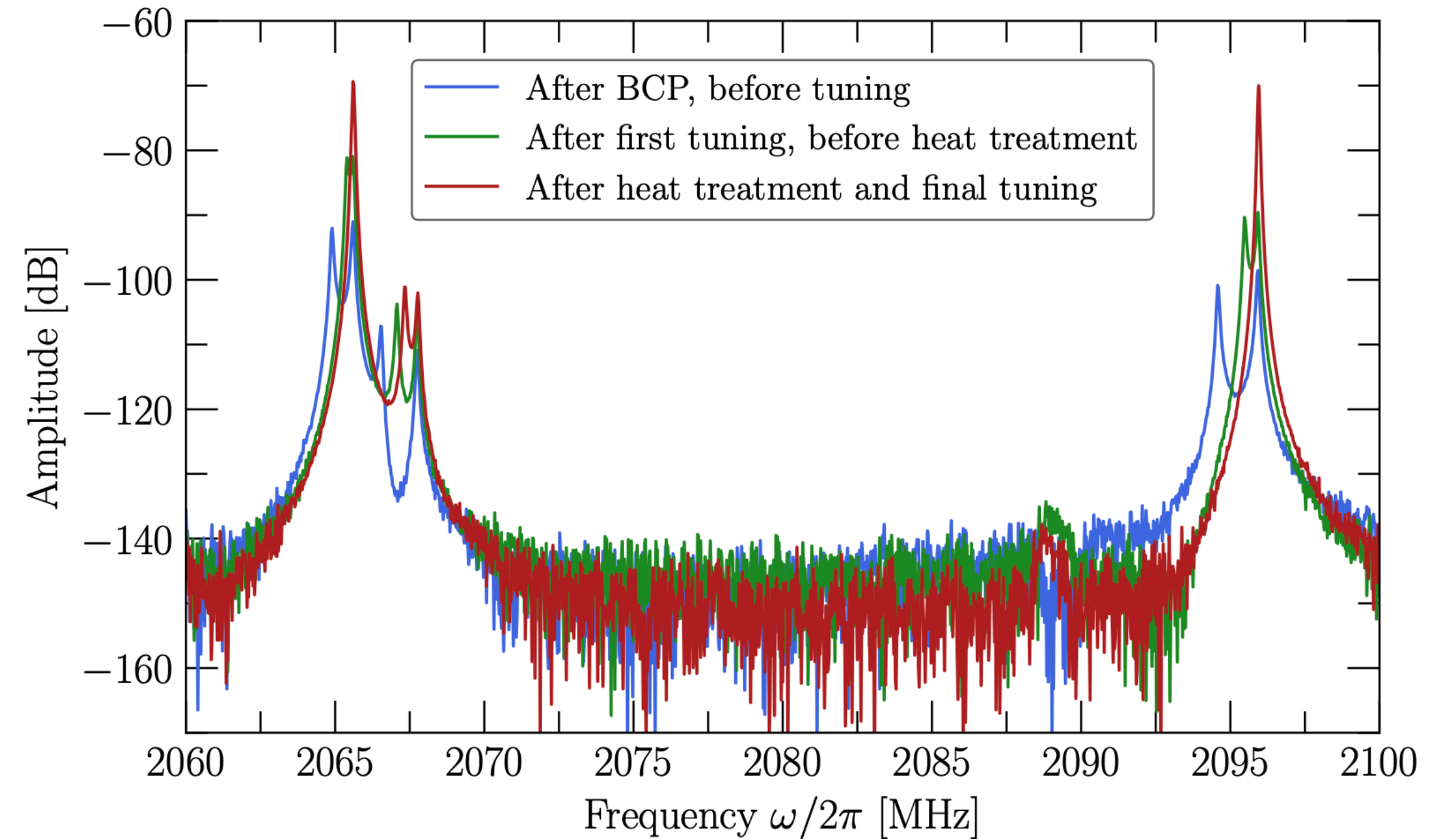


Mechanical modes

# Characterising the Prototype



Mechanical modes



RF spectrum



# Conclusion

**MAGO operates in:**

- Low-frequency regime:  $1/L \gg \omega_g$  — use PDF and account for current & boundary changes

**Renewed collaboration between INFN/DESY/  
FNAL**

**First data expected in 2025**

