

# *NuMI $\nu_{\mu}$ CC Inclusive Analysis Status*

Guadalupe Moreno, Promita Roy

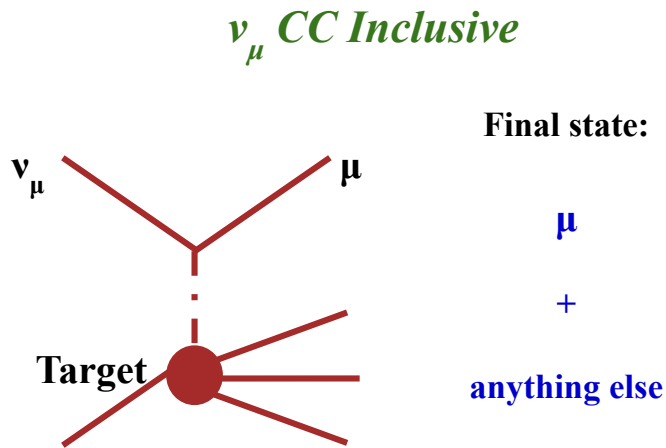
*ICARUS Collaboration Meeting*  
**October 14, 2024**

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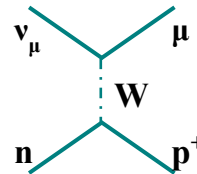
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# Overview $\nu_\mu$ CC Inclusive Selection

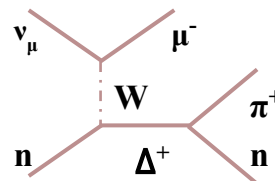
As part of the efforts to study the NuMI  $\nu$ -Ar cross-section at ICARUS, *we are analyzing the  $\nu_\mu$  inclusive charged current channel, which consists of studying the kinematics of the  $\mu$  at the final state, without imposing constraints on the hadronic system.*



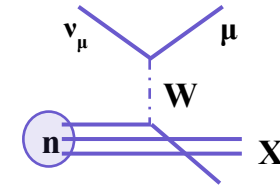
Quasi-elastic scattering (QE)



Resonance production (RES)



Deep Inelastic scattering (DIS)



Considering that we are studying the  $\nu_\mu$  CC inclusive channel, we must consider all interactions at the final state, i.e.

- *Quasi-elastic scattering (QE)*
- *Resonance production (RES)*
- *Deep Inelastic scattering (DIS)*
- *2p-2h (MEC)*

❖ *In this channel the dominant background is due to cosmic ray interactions.* This background will be constrained with cosmic data i.e OFF-Beam NuMI data.

# Revisiting $\nu_{\mu}$ CC inclusive selection criteria

Pre-selection  
(for cosmic background  
rejection)

- Remove everything that is a **Pandora clear cosmic**
- **Fiducial Volume cut**: the vertex should be contained in the Fiducial Volume:
  - $x = (-358.49+25, -61.94-25) \text{ cm} \cup (61.94+25, 358.49-25) \text{ cm}$
  - $y = (-181.86+25, 134.96-25) \text{ cm}$
  - $z = (-894.951+30, 894.951-50) \text{ cm}$
- **CRLongestTrackDirY** > - 0.7 : Longest track's Y-direction (under Cosmic Ray hypothesis)
- **Barycenter Flash Matching** < 1m

$\mu$ - selection  
(for muon  
signals)

- Muon track
  - **Distance from Vertex** < 10 cm
  - **Contained**
    - Length > 50 cm
    - $\chi^2_{\mu} < 30$  &  $\chi^2_p > 60$  [which tells us basically that the track is consistent with muon and inconsistent with proton]
  - **Exiting**
    - Length > 100cm

# Review

## $\nu_{\mu}$ CC Inclusive Selection using BFM

- The Barycenter flash matching tool developed by **J. Smedley** matches the **optical flashes (PMT OpFlashes)** to the TPC slices by minimizing the distance between the **weighted charge center of the slice and the weighted center of the flash**, thus rejecting out-of-time slices based on the time of the flash.
- This slide and the following slides show
  - a. Current status of the inclusive selection,
  - b. Initial results of application of 3-fold matching (CRT-TPC-PMT) to improve background rejection,
  - c. Data-MC comparisons
  - d. Glance at the muon momentum
  - e. Preliminary study of systematic errors in this  $\nu_{\mu}$  CC Inclusive channel.

- **Efficiency = selected signal / all signal**

- **Purity = selected signal / (selected signal + background)**

- **Selected signal is  $\nu_{\mu}$  CC**

Considering all events	Purity $\nu_{\mu}$ CC	Purity ( $\nu_e$ CC + NC)	Purity (out-of-time Cosmics)	Purity (In-time cosmics)	Efficiency $\nu_{\mu}$ CC
<b>Fiducial</b>	0.016	0.005	0.502	0.478	0.56
<b>+ clear cosmic</b>	0.071	0.022	0.448	0.458	0.54
<b>+ Track cut</b>	0.094	0.011	0.447	0.448	0.50
<b>+ Longest track in Y direction</b>	0.229	0.038	0.327	0.406	0.46
<b>+ Barycentre flash matching</b>	0.674	0.077	0.055	0.20	0.40

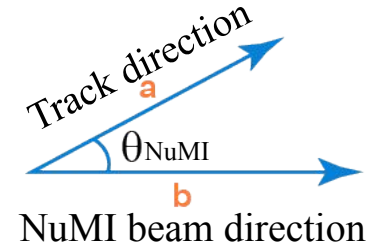
# 3-fold Matching

## *CRT-PMT-TPC matching*

- Previous studies made by **Andrea S. (doc-27838)**, **Francesco P. (doc-28827)** and **Anna H. (doc 32538, 32375)** had shown the possibility of using the time-matched CRT-hit for each OpFlash to identify entering/exiting tracks.
- This CRT-PMT(flash) matching is performed by matching an optical flash with one or more CRT hits and could be useful in rejecting in cosmics which is the principal background in our  **$\nu_{\mu}$  CC inclusive analysis**.
- So, we wanted to explore CRT-PMT matching tool along with all the other  $\nu_{\mu}$  CC inclusive selection criteria and investigate if we can do a better background rejection.
- *For this purpose a CRT-PMT matched spill-cut is used on top of Barycentre flash matched slices which investigates if there is any match between CRT system and PMT flash within the NuMI beam gate and accordingly rejects or selects the slices (more info on sbn docdb [38233-v1](#)).*
- *We are using Run2 reprocessed onbeam and offbeam NuMI data and latest NuMI MC sample.*

# CRT-PMT-TPC matching

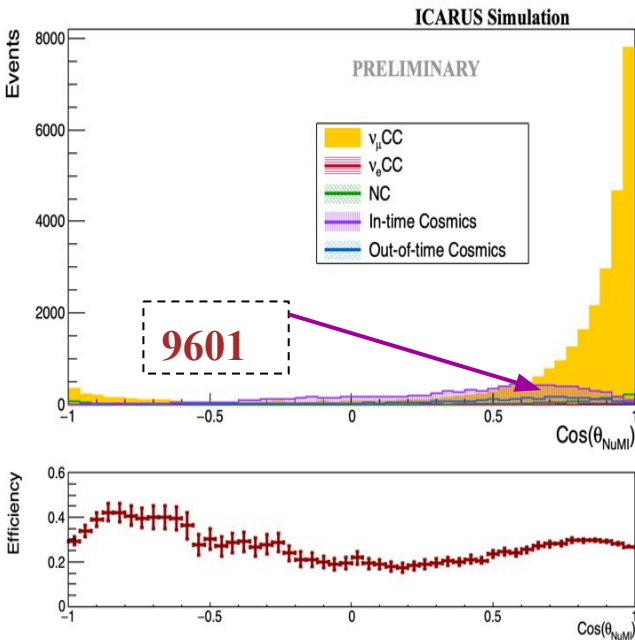
$$\text{Cos } \theta_{\text{NUMI}}$$



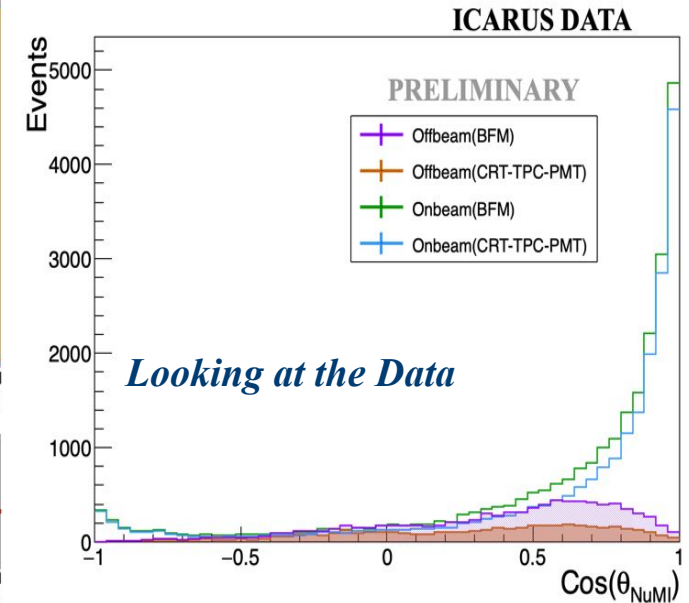
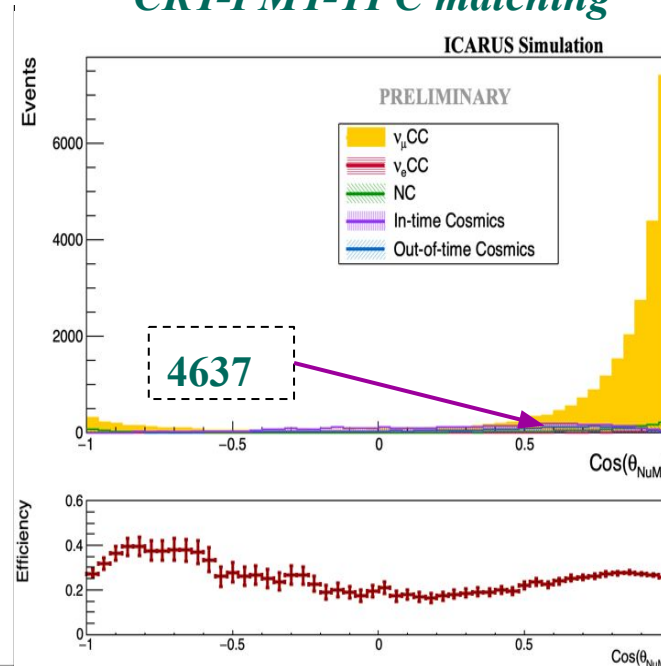
$$\cos(\theta_{\text{NuMI}}) = \frac{\vec{a} \cdot \vec{b}}{|\vec{a}| |\vec{b}|}$$

**Pre-BFM selection:** Fiducial volume cut, Not a Pandora Clear Cosmic, Longest track Y-direction cut, Track cut, Split muons cut

## Barycentre flash matching



## CRT-PMT-TPC matching

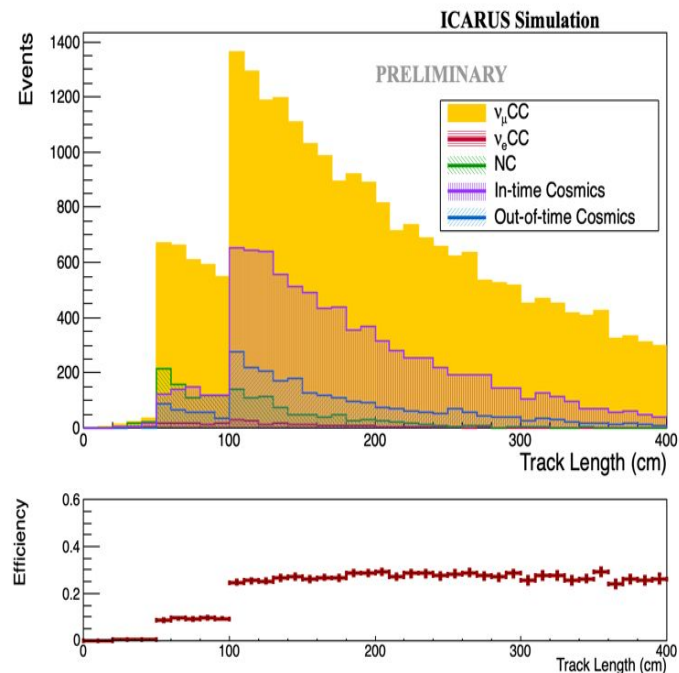


*On using triple matching (CRT-PMT-TPC), in-time cosmics (offbeam data) is reduced considerably.*

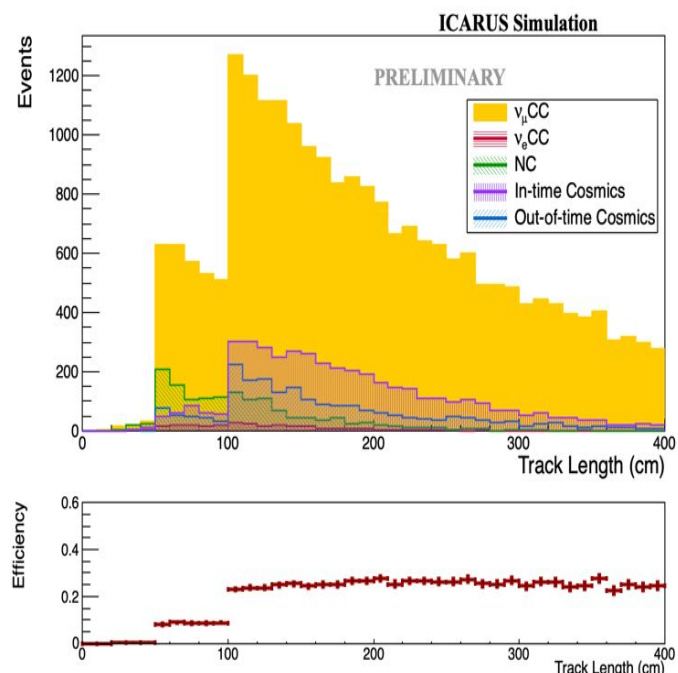
# CRT-PMT-TPC matching *Track length*

*Pre-BFM selection:* Fiducial volume cut, Not a Pandora Clear Cosmic, Longest track Y-direction cut, Track cut, Split muons cut

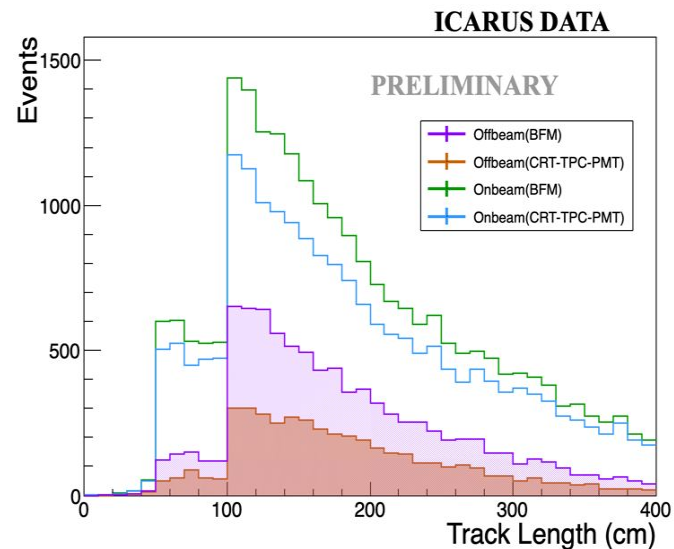
## Barycentre flash matching



## CRT-PMT-TPC matching



## Looking at the Data



- *On using triple matching (CRT-PMT-TPC), in-time cosmics is reduced considerably.*
- *For out-of-time cosmics, performance of triple matching is similar as barycentre flash matching*

# Efficiency and Purity *BFM and CRT-PMT-TPC matching*

- *Efficiency* = selected signal / all signal
- *Purity* = selected signal / (selected signal + background)
- Selected signal is  $\nu_{\mu} CC$

Efficiency  
 CRT-PMT-TPC match w splittrk cut: 26%  
 BFM w splittrk cut: 27.5%

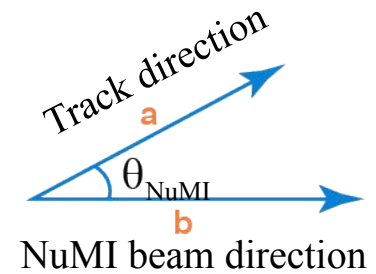
Considering all events	Purity $\nu_{\mu} CC$	Purity ( $\nu_e CC + NC$ )	Purity (out-of-time Cosmics)	Purity (In-time cosmics)	Efficiency $\nu_{\mu} CC$
<b>Fiducial</b>	0.016	0.005	0.502	0.478	0.56
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<b>+ Barycentre flash matching</b>	<b>0.674</b>	0.077	0.055	<b>0.20</b>	0.40
<b>+ CRT-PMT spillcut (CRT-PMT-TPC matching)</b>	<b>0.761</b>	0.087	0.052	<b>0.11</b>	0.371



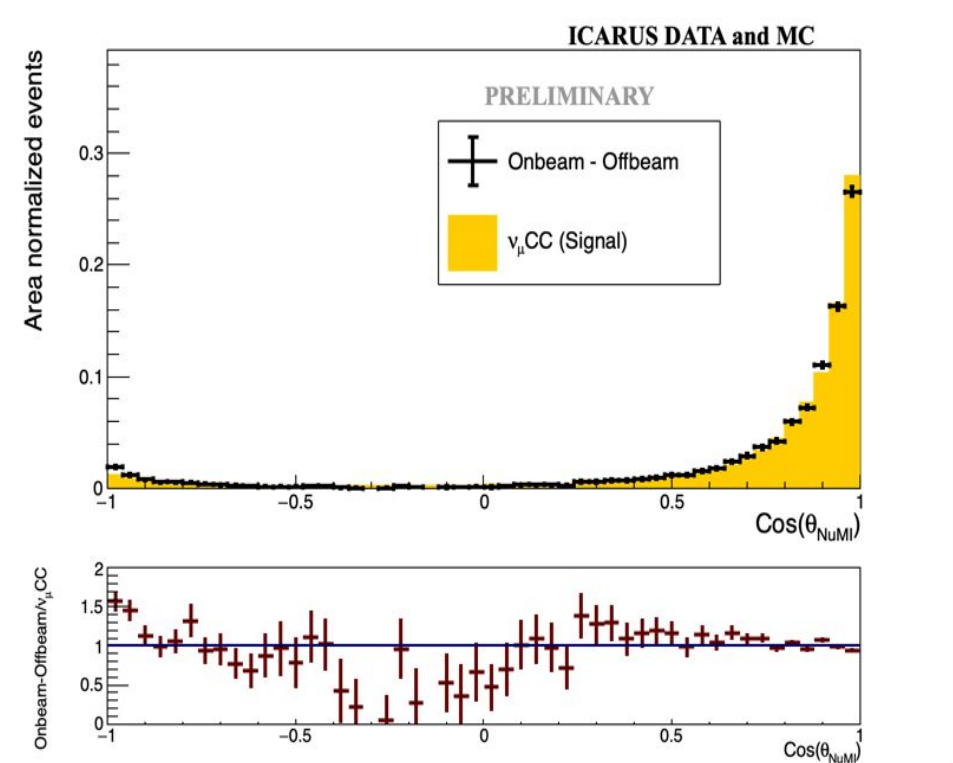
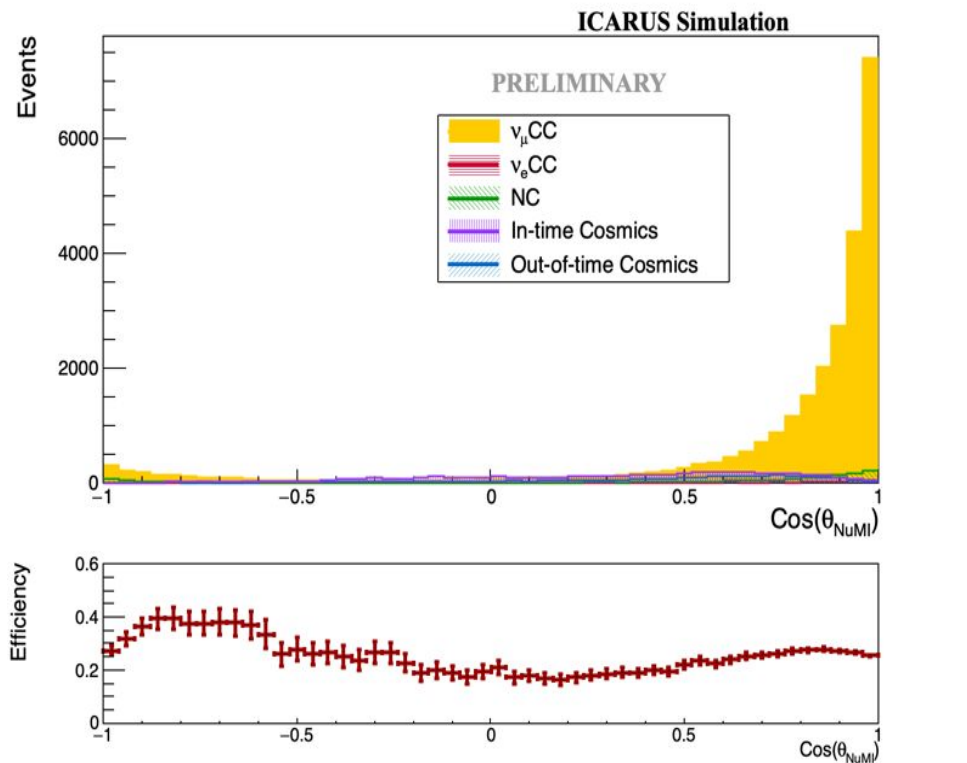
# Inclusive Selection *Data vs MC*

- Inclusive selection criteria with triple matching is used
- “*OnBeam - OffBeam*” is matched with the signal ( $\nu_\mu$  CC) MC
- The Data-MC plot is area normalized

$$\cos(\theta_{NuMI}) = \frac{\vec{a} \cdot \vec{b}}{|\vec{a}| |\vec{b}|}$$



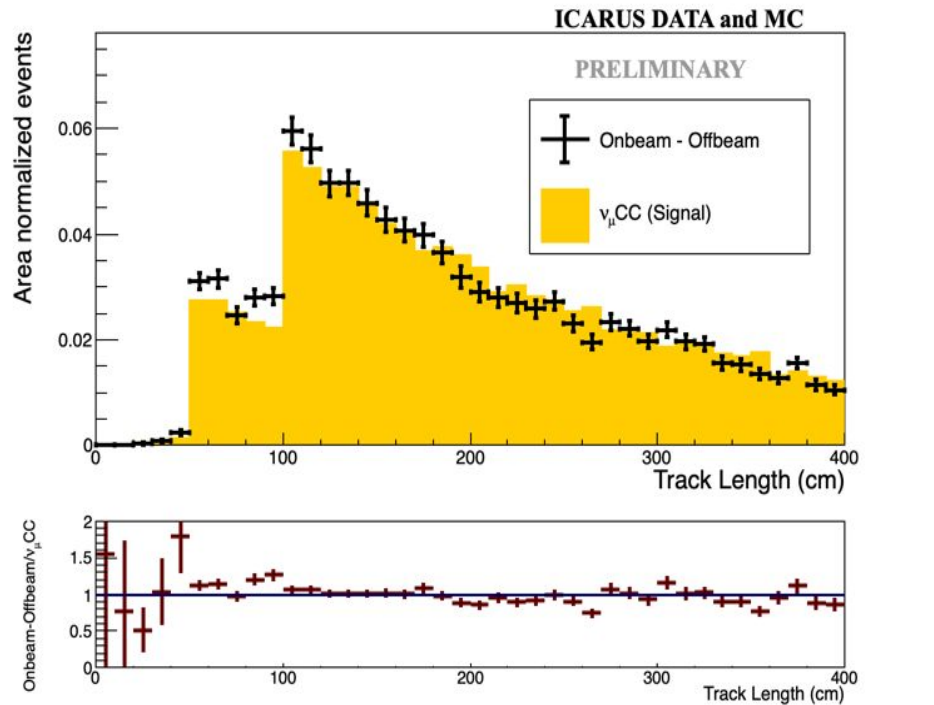
← **Cosine  $\theta_{NuMI}$**  →



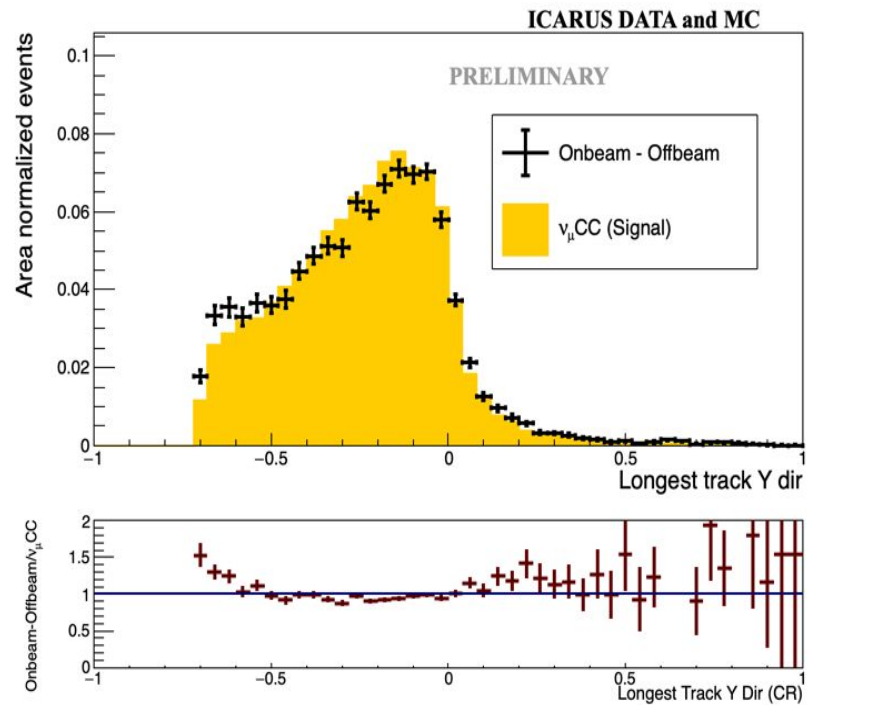
# Inclusive Selection *Data vs MC*

- Inclusive selection criteria with triple matching is used
- “*OnBeam - OffBeam*” is matched with the signal ( $\nu_\mu$  CC) MC
- The Data-MC plots are area normalized

## Track length



## Longest track in Y- direction



# Muon Momentum *MC (Inclusive selection)*

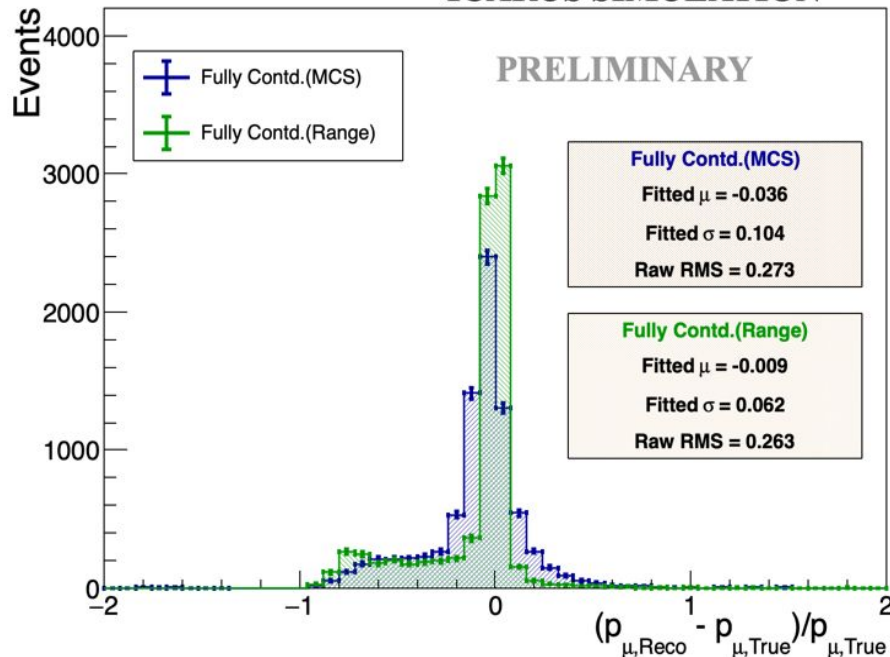
Sbn-docdb 31487-v2 by  
Moon

- *Range based momentum*
- *MCS: Occurs when a charged particle enters a medium and undergoes electromagnetic scattering with the atoms. This scattering deviates the original trajectory of the particle within the material. The deflection of muons from the original trajectory within the material form a Gaussian distribution with mean at 0 and standard deviation varying with the track momentum.*

Split track cut is applied for both  
(Thanks to Bruce & Harry)

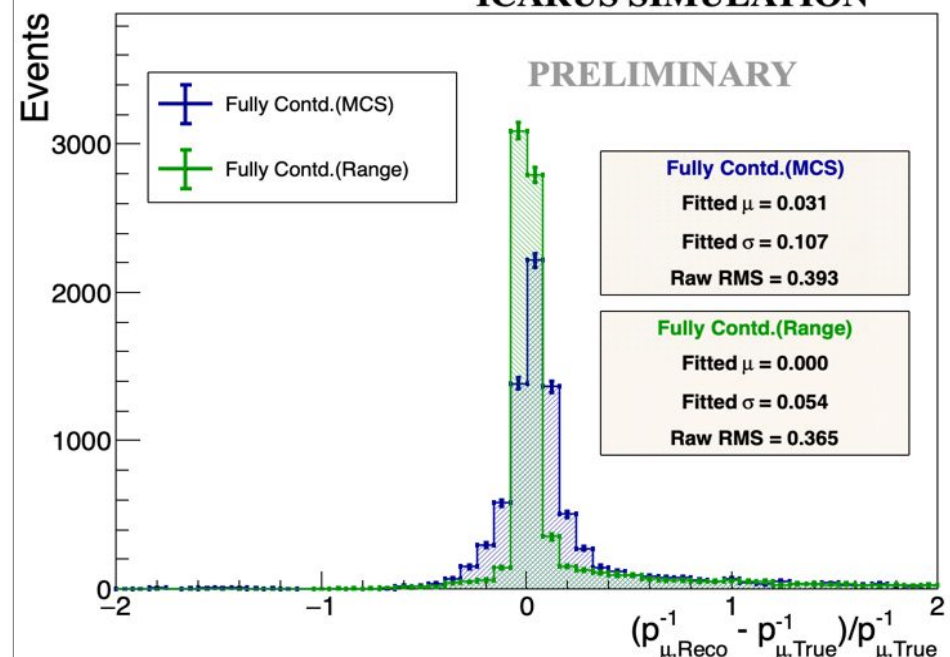
## Momentum resolution

### ICARUS SIMULATION



## Inverse Momentum resolution

### ICARUS SIMULATION



# Muon Momentum *MC (Inclusive selection)*

*Fully Contd.(Range)*

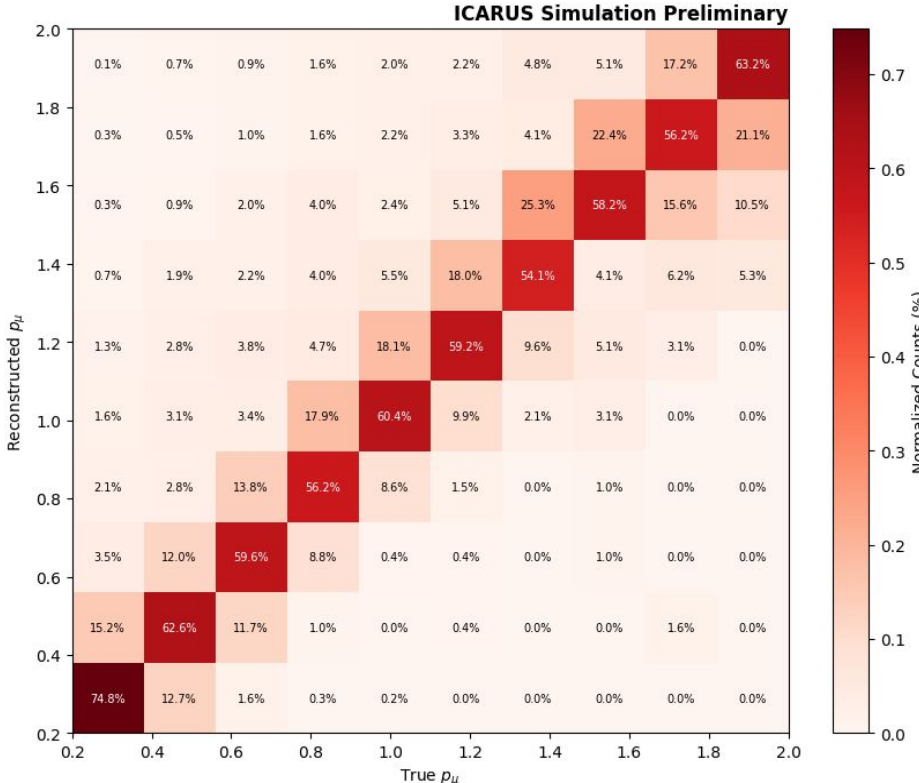
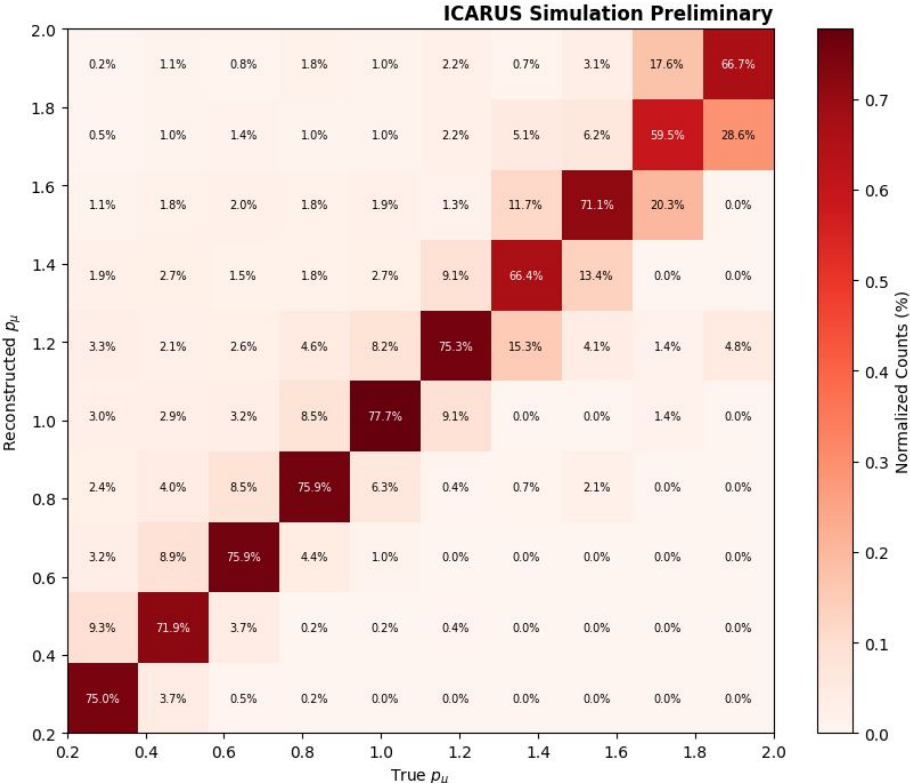
- Fractional bias( $\mu$ ) < 1%
- Fractional resolution( $\sigma$ ) ~ 5-6%

*Fully Contd.(MCS)*

- Fractional bias( $\mu$ ) < 4%
- Fractional resolution( $\sigma$ ) ~ 10%

*Fully contained tracks (range)*

*Fully contained tracks (MCS)*



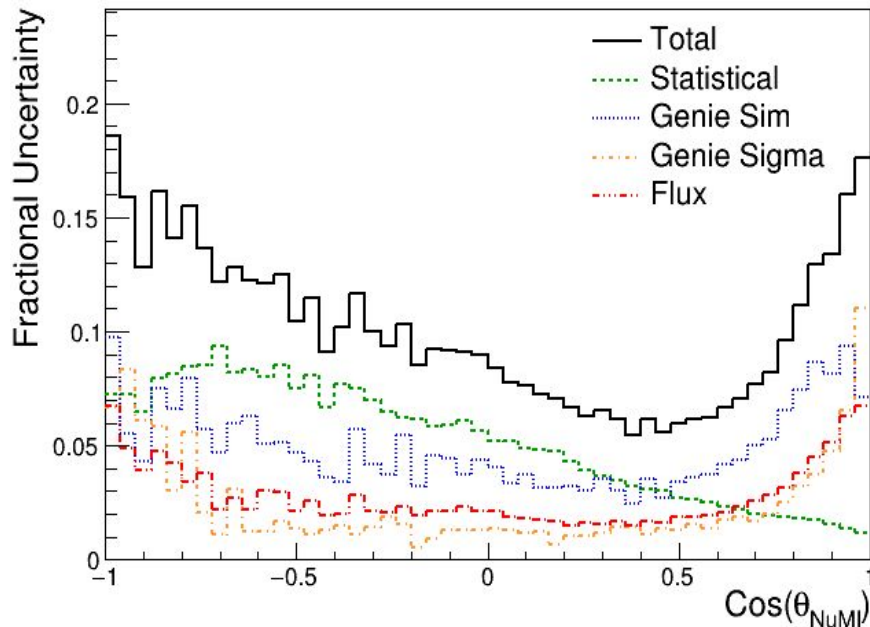
# Systematics (to do)

## *Inclusive selection*

In this study we only consider the systematics due to the **Flux and the  $\nu$  interactions**.

The systematic errors can be categorized as

- Detector
- **Flux**: we use a combination of all beamline systematics and the Npc (number of principal components) hadron production components.
- **Neutrino interaction (GENIE)**: we utilize two kind of GENIE weights
  - Genie Sigma (Multisigma) weights are assigned to parameters that we assume operate independently from others. This means we can adjust these parameters individually without affecting the rest.
  - Genie Sim (Multisim) is employed for parameters that are interdependent. To see the impact of these variations, we generate event weights for numerous “universes,” each representing different settings of these parameters.

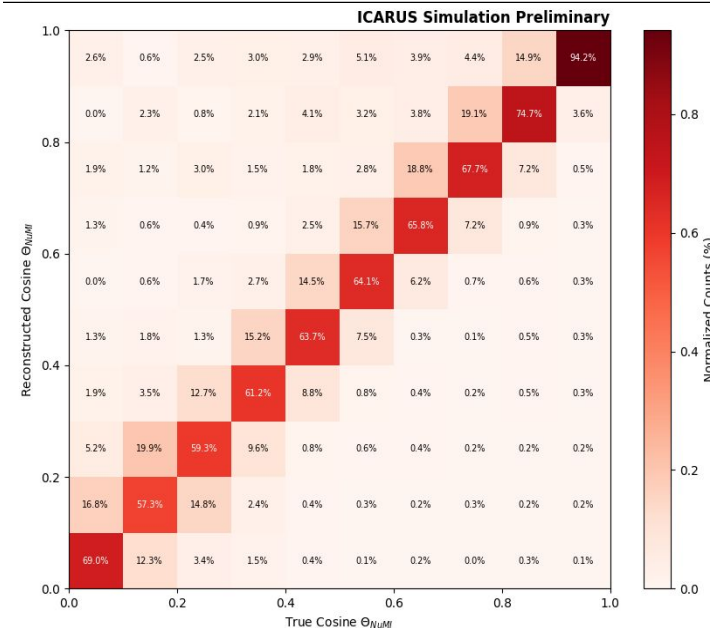
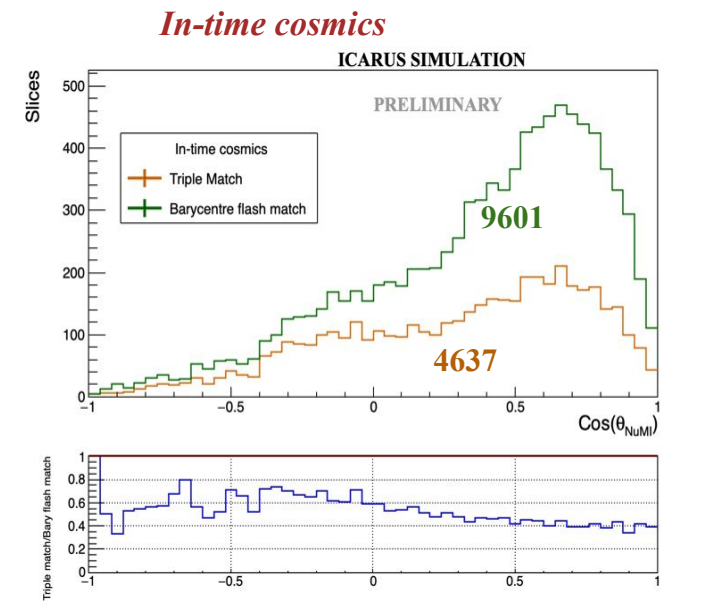


### *Detector systematics:*

We have started producing small samples for detector systematics and started looking at them following the way **Jacob Z.** is doing (**sbn-docdb 35850-v1, 36027-v2, 36350-v1** and many others)

# Summary *and Next To Do's*

- CRT-PMT-TPC matching comes out to be an effective tool rejecting for in-time cosmics.
  - Final purity of the signal increases by  $\sim 9-10\%$  and final efficiency reduces by  $\sim 1.5\%$ .
  - More investigations at a deeper level is going on, so further improvements are anticipated.
- The initial area normalized Data and MC comparison plots look decent with the newly produced MC and Run2 reprocessed datasets.
  - Need to add POT to these plots and look at them closely.
- Muon momentum studies show that applying Split track cut clears out a lot of mis-reconstructed tracks which otherwise affects the momentum estimation. For fully contained sample MCS performs almost as good as range based reconstruction. The reco vs true matrix looks almost identical for both.
- Recently we have started looking at Flux, GENIE and Detector systematics. We will continue the systematics study and have a breakdown of each of them.
- Next step is to extract cross-section using unfolding matrix (reco-true cosine  $\theta_{NuMI}$ )

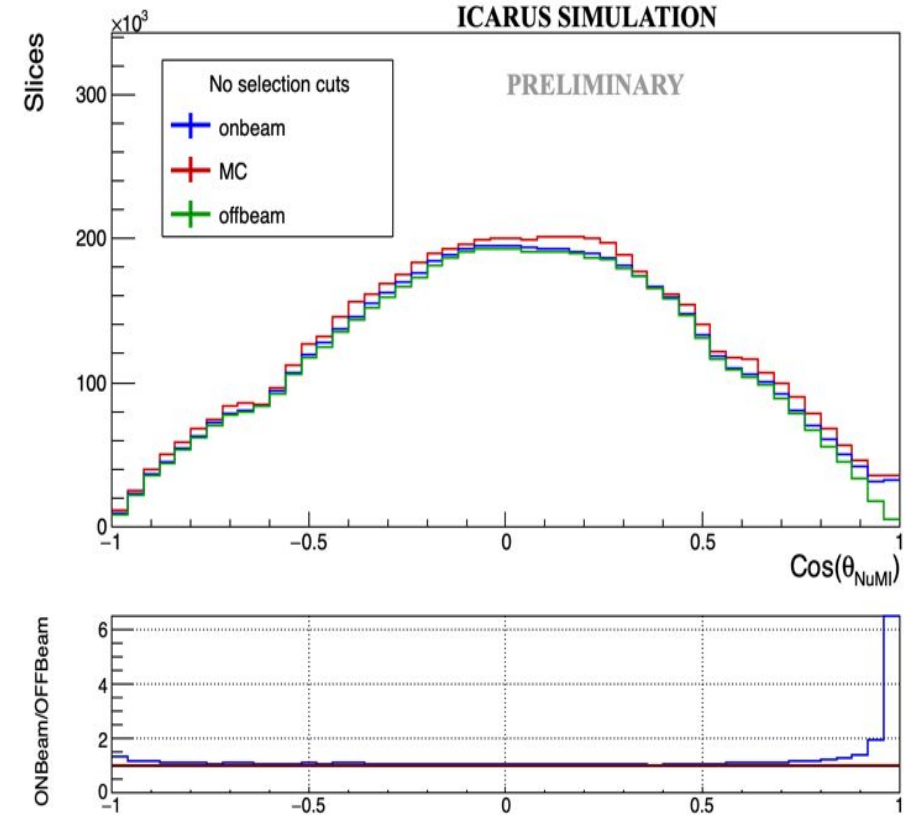




Back up

# Using data (onbeam and offbeam)

Considering all events	Purity $\nu_{\mu}$ CC
Pre-bary selection cuts + Barycentre flash matching	0.65
+ CRT-PMT spillcut (CRT-PMT-TPC matching)	0.8



sbndocdb: 32106-v1 (Jacob Larkin,  
Joseph Z.)



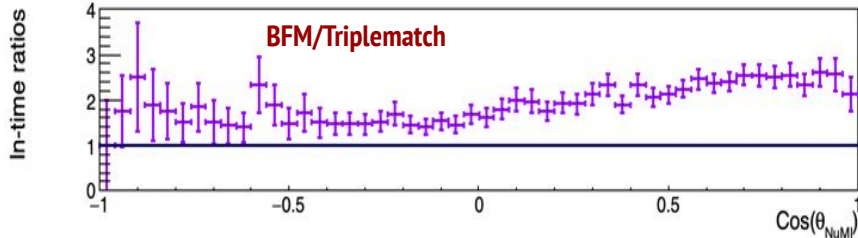
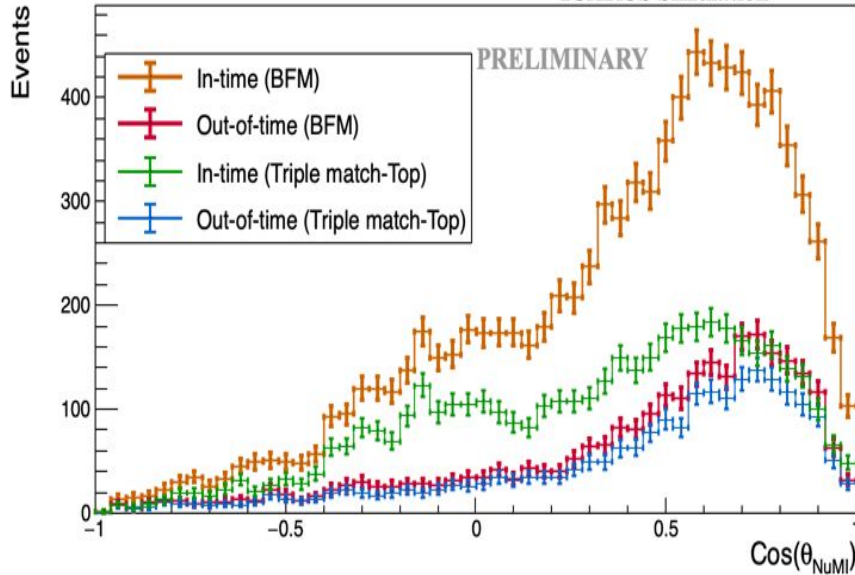
# Looking closely at the cosmics *In-time & Out-of-time*

Only Top CRT

- In-time cosmics is reduced considerably after using triple matching
- In-time ratios (BFM/Triple match) shows the extent of reduction

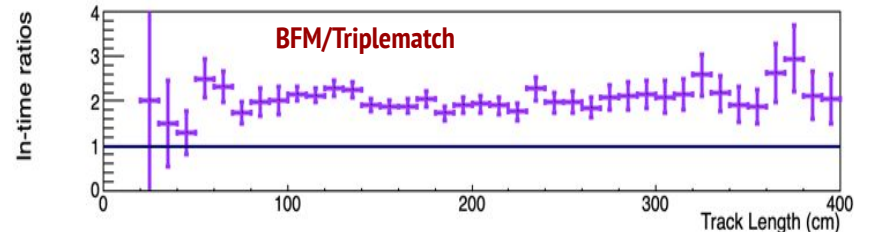
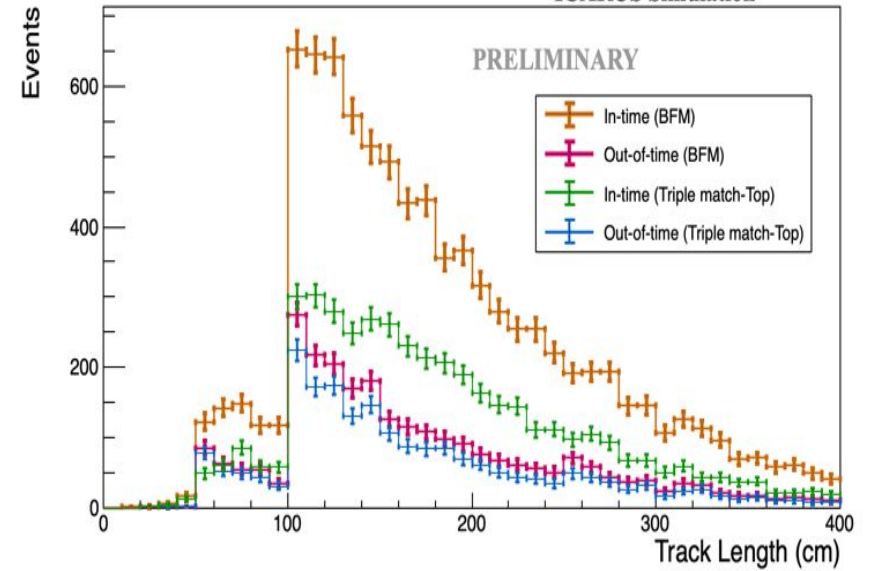
**Cosine  $\theta_{NUMI}$**

ICARUS Simulation



**Track length**

ICARUS Simulation



## MCS momentum estimation

Original highland formula:

$$\sigma_o^{\text{HL}} = \frac{S_2}{p\beta c} z \sqrt{\frac{\ell}{X_0}} \left[ 1 + \epsilon \times \ln \left( \frac{\ell}{X_0} \right) \right],$$

MicroBOONE used:

$$\sigma_o = \sqrt{(\sigma_o^{\text{HL}})^2 + (\sigma_o^{\text{res}})^2}.$$

where the second term is inherent detector angular resolution  $\sim 3\text{mrad}$

MicroBOONE tuned the highland formula and replaced  $S_2$  with  $k(p)$

$$\kappa(p) = \frac{0.105 \text{ MeV}}{(p(\text{GeV}))^2} + 11.004 \text{ MeV}.$$

Thus, the tuned standard deviation (of the gaussian distribution) used by MicroBOONE (at  $l = X_0$ ) becomes the following:

$$\sigma_o^{\text{RMS}} = \sqrt{(\sigma_o)^2 + (\sigma_o^{\text{res}})^2} = \sqrt{\left( \frac{\kappa(p)}{p\beta c} \right)^2 + (\sigma_o^{\text{res}})^2}.$$

# Muon Momentum MC (Inclusive selection)

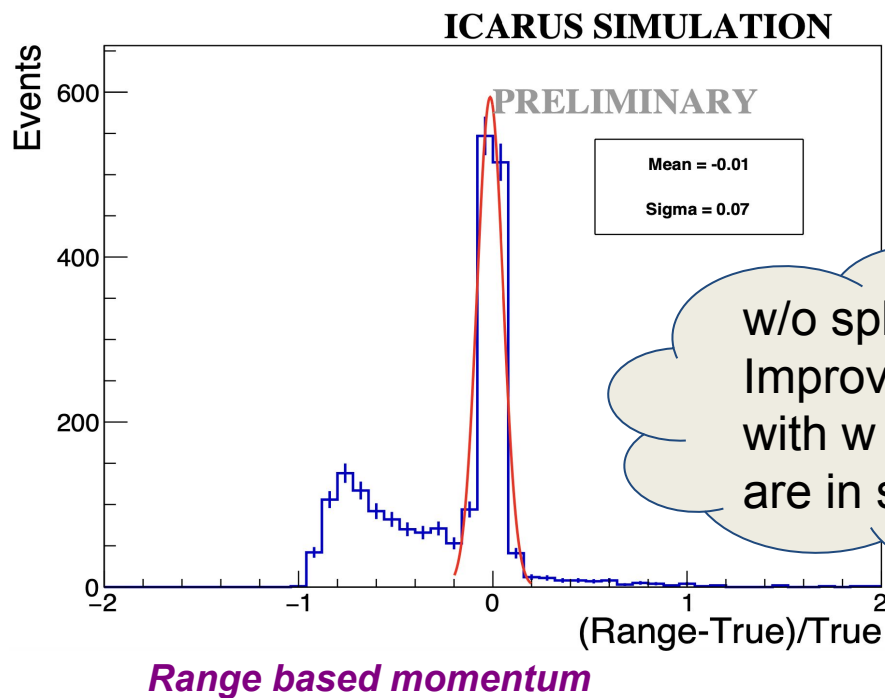
## Muon momentum resolution ( $p_\mu$ )

The collection of the deflection of muons from the original trajectory within the material form a Gaussian distribution with mean at 0 and standard deviation given by the Highland formula:

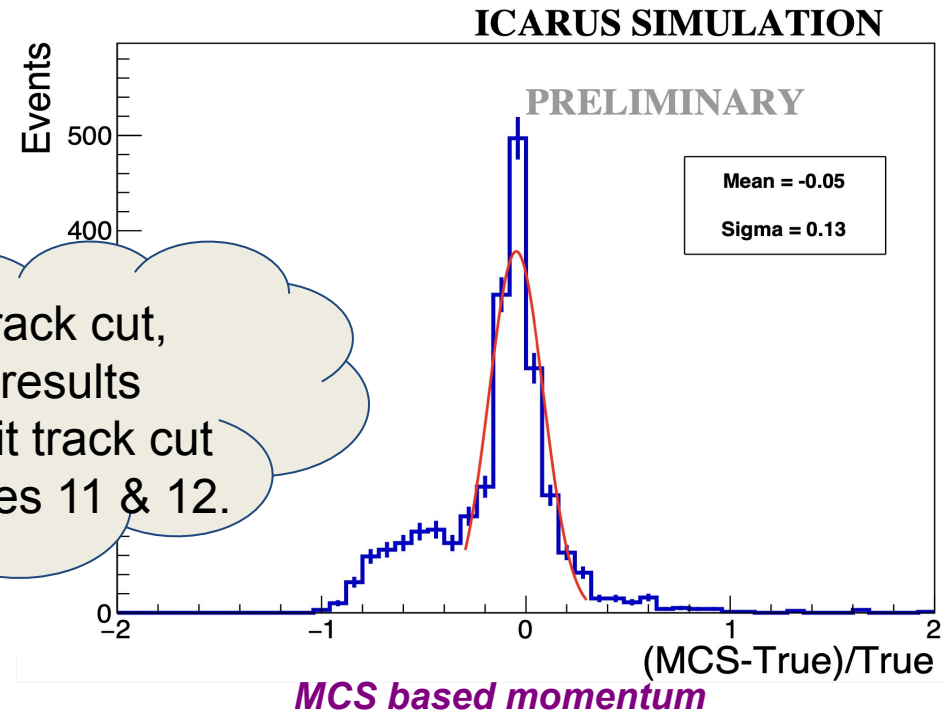
$$\theta_0 = \frac{13.6 \text{ MeV}}{p\beta c} z \sqrt{\frac{\ell}{X_0}} \left[ 1 + 0.0038 \ln\left(\frac{\ell}{X_0}\right) \right]$$

- $\beta$ :  $v/c$
- $\ell$ : is the distance traveled inside the material
- $X_0$ : Radiation length

Fully contained



w/o split track cut,  
Improved results  
with w split track cut  
are in slides 11 & 12.



# Muon Momentum *MC (Inclusive selection)*

*Contained + exiting sample*

*Contd.(Range), Uncontd.(MCS)*

- *Fractional bias( $\mu$ )* < 4%
- *Fractional resolution( $\sigma$ )* ~ 8-10%

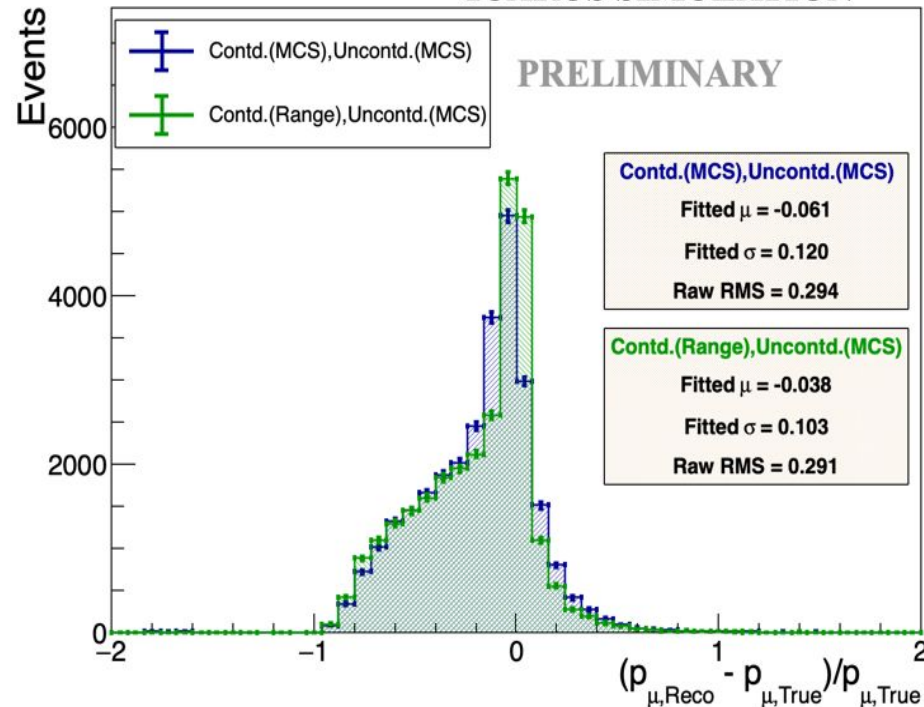
*Contd.(MCS), Uncontd.(MCS)*

- *Fractional bias( $\mu$ )* ~ 4-6%
- *Fractional resolution( $\sigma$ )* ~ 12-13%

**Split track cut is applied for both**

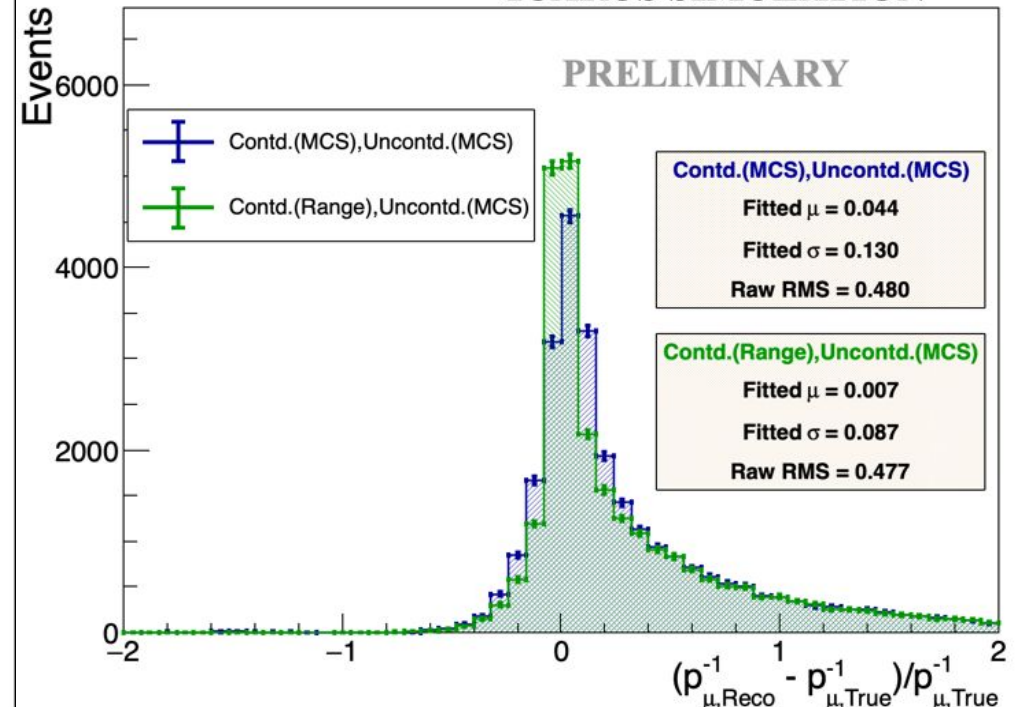
## *Momentum resolution*

ICARUS SIMULATION



## *Inverse Momentum resolution*

ICARUS SIMULATION

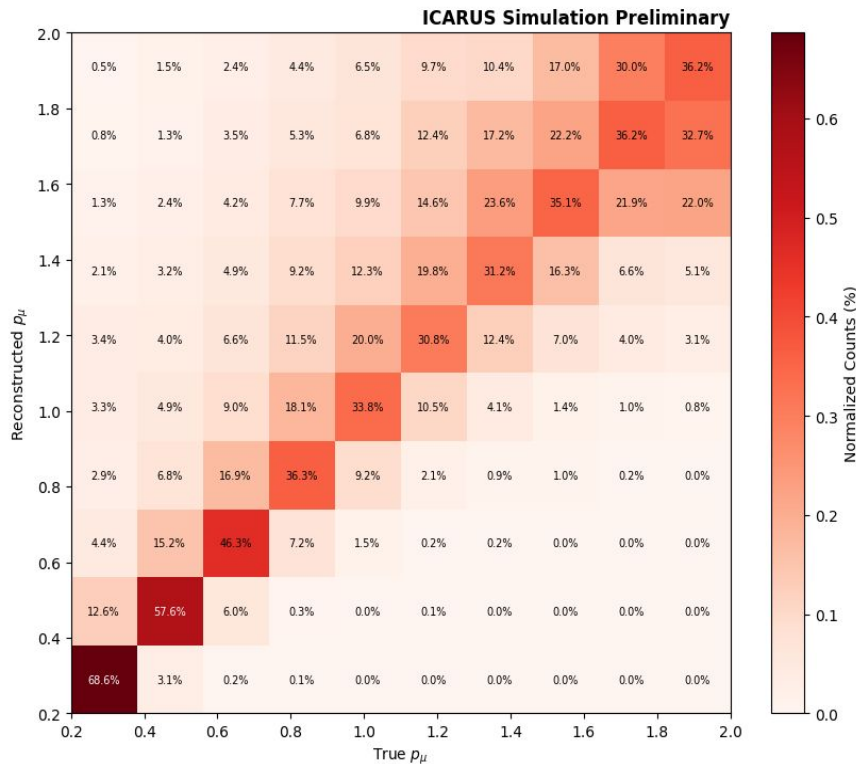


# Muon Momentum *MC (Inclusive selection)*

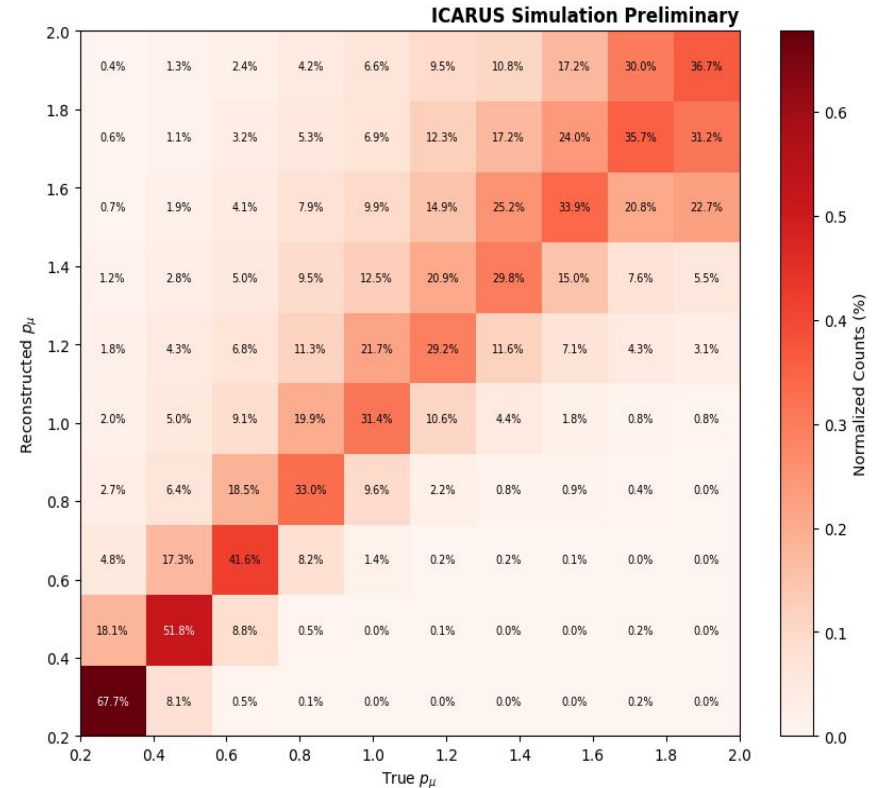
*Contained + exiting sample*

- Underestimation of the MCS momentum due to track mis-reconstruction
- These mis-reconstruction distort the MCS angle distribution for the track and push the MCS momentum to lower values, especially for higher momentum tracks (that should have small scattering angles)
- Applying **split-track cut**, we could clear out a chunk of mis-reconstructed tracks and that improved the resolution.

*Contained(Range), Uncontained (MCS)*



*Contained(MCS), Uncontained (MCS)*





*New MC and  
data samples*

- ❖ **icaruspro\_production\_v09\_89\_01\_01p01\_2024A\_ICARUS\_NuMI\_MC\_NuMI\_MC\_flatcaf**  
**(official production in July, 2024)**
- ❖ **Run2 reprocessed NUMIMajority (v09\_89\_01)**  
**(official production in September, 2024)**
- ❖ **Run2 reprocessed OFFBEAMNUMIMajority (v09\_89\_01)**  
**(official production in September, 2024)**

# Flash classification IDs: Type of match between a TPC object (e.g. PMT flash) and CRT system

## A. With CRT hit

**noMatch = 0** < No CRT hit/match >

*This will actually cut both entering and exiting tracks so it would not work if we want to keep neutrinos with exiting muons.*

Events in the last bin of Cosine  $\theta_{NUMI}$

Cosmic veto	<b>4500</b>
Cosmic tagger	<b>7000</b>

## B. With CRT-PMT matching

- ★ A spill is considered as "*Cosmic*" when all intime flashes are entering
- ★ If a spill has anything not entering, that should not be considered as a cosmic
- **enTop = 1**, ///< Matched with Top CRT hit before optical flash >
- **enSide = 2**, ///< Matched with Side CRT hit before optical flash >
- **enTop\_exSide = 3**, ///< Matched with one Top CRT hit before the optical flash and matched with one Side CRT hit after the optical flash >
- **enTop\_mult = 6**, ///< Matched with multiple Top CRT hits before the optical flash >
- **enTop\_exSide\_mult = 7**, ///< Matched with multiple Top CRT hits before the optical flash and more than 1 side CRT hits after the optical flash >

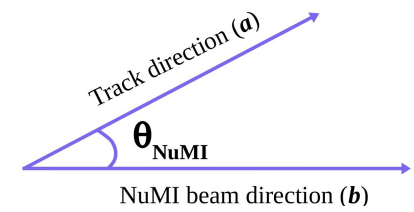
# Inclusive Selection *MC Unfolding matrix*

The  $U_{\mu a}$  is the *unfolding matrix* that basically takes us from the measured variable to the true variable. The purpose of this matrix is to tell us the probability that an event observed in bin  $a$  (measured) actually occurred in bin  $\mu$  (true). We can use our MC simulation to perform a migration matrix indicating what fraction of events generated in each true bin  $\mu$  were observed in each reconstructed bin  $a$ .

Unfolding

$$\left\langle \frac{d^n \sigma}{dx} \right\rangle_{\mu} = \frac{\sum_a U_{\mu a} (D_a - B_a)}{\Phi T \Delta x_{\mu}}$$

$$\cos(\theta_{NuMI}) = \frac{\vec{a} \cdot \vec{b}}{|\vec{a}| |\vec{b}|}$$



*Cosine  $\theta_{NuMI}$  : Reconstructed vs True*

