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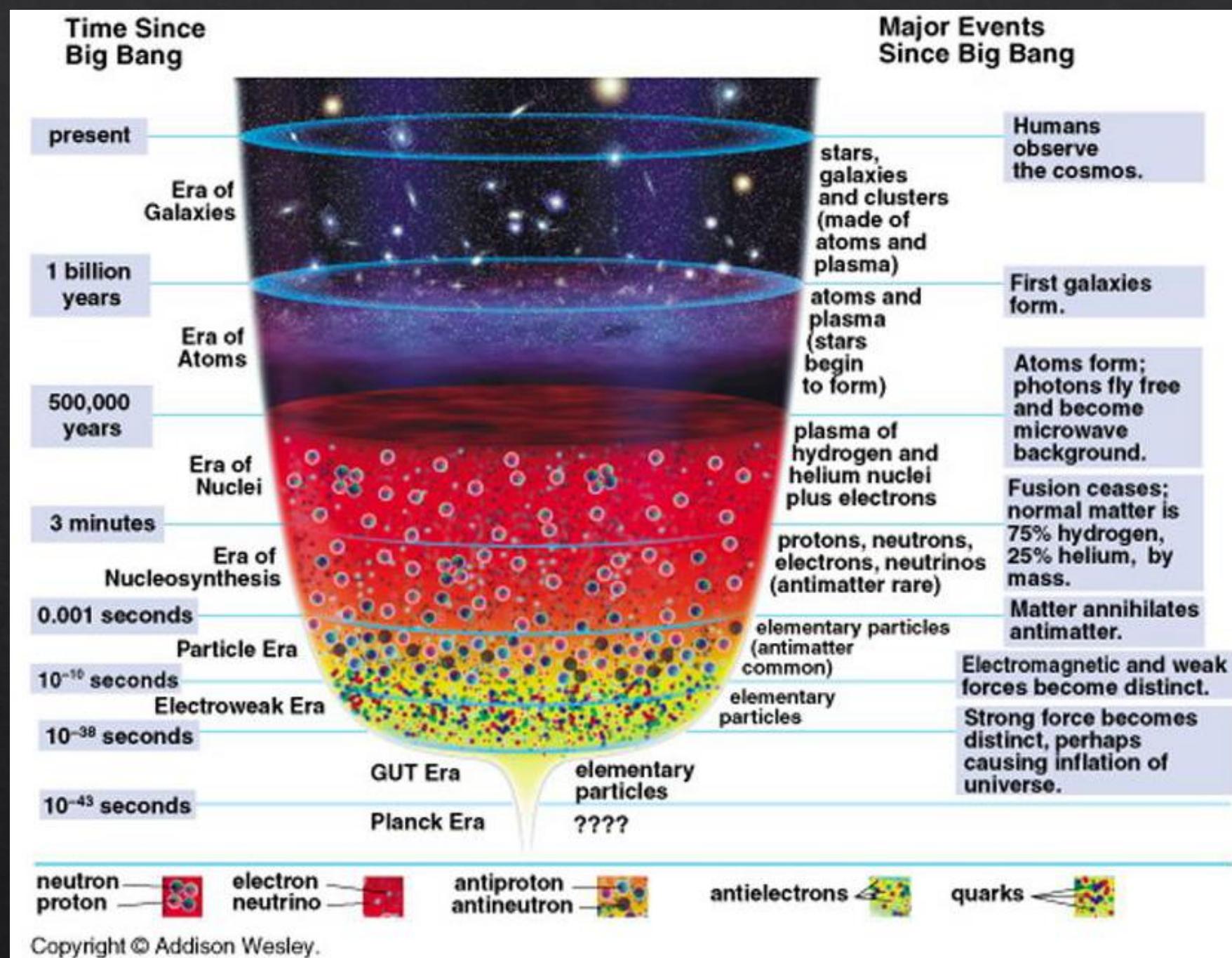
# Introducere în fizica particulelor la energii înalte II

“What is the Universe made of?”  
High Energy Physics summer school

3 septembrie 2024

# Întrebări

- ? Din ce este constituită materia?
- ? Din ce este compus universul?
- ? Care este originea universului și cum a evoluat?
- ? De ce se comportă aşa universul?
- ? Cum va evoluă?
- ? ....



# Care sunt elementele din care este constituita materia?



Empedocles 492-432 BC

By convention there is color,  
by convention sweetness,  
by convention bitterness,  
but in reality there are  
atoms and space.

Democritus 400 BC

Periodic Table of the Elements																																																																																																																																																					
IA		IIA		IIIB		IVB		VB		VIB		VII		IB		IIB		0																																																																																																																																			
1	H	2	Li	3	Na	19	K	37	Rb	55	Cs	87	Fr	5	B	13	Al	31	Ga	53	In	81	Tl	1	He																																																																																																																												
2	Li	3	Mg	12	Mg <th>20</th> <td>Ca</td> <th>38</th> <td>Sr</td> <th>56</th> <td>Ba</td> <th>88</th> <td>Ra</td> <th>4</th> <td>Sc</td> <th>21</th> <td>Ti<th>39</th><td>Zr</td><th>72</th><td>Hf</td><th>73</th><td>Ta</td><th>74</th><td>W</td><th>75</th><td>Re</td><th>76</th><td>Os</td><th>77</th><td>Ir</td><th>78</th><td>Pt</td><th>79</th><td>Au</td><th>80</th><td>Hg</td><th>81</th><td>Tl</td><th>82</th><td>Pb</td><th>83</th><td>Bi</td><th>84</th><td>Po</td><th>85</th><td>At</td><th>86</th><td>Rn</td></td>	20	Ca	38	Sr	56	Ba	88	Ra	4	Sc	21	Ti <th>39</th> <td>Zr</td> <th>72</th> <td>Hf</td> <th>73</th> <td>Ta</td> <th>74</th> <td>W</td> <th>75</th> <td>Re</td> <th>76</th> <td>Os</td> <th>77</th> <td>Ir</td> <th>78</th> <td>Pt</td> <th>79</th> <td>Au</td> <th>80</th> <td>Hg</td> <th>81</th> <td>Tl</td> <th>82</th> <td>Pb</td> <th>83</th> <td>Bi</td> <th>84</th> <td>Po</td> <th>85</th> <td>At</td> <th>86</th> <td>Rn</td>	39	Zr	72	Hf	73	Ta	74	W	75	Re	76	Os	77	Ir	78	Pt	79	Au	80	Hg	81	Tl	82	Pb	83	Bi	84	Po	85	At	86	Rn																																																																																																				
4	K	5	Rb	17	Cl	35	Br	53	I	71	Fr	89	Ac	90	Th	91	Pa	92	U	93	Np	94	Pu	95	Am	96	Cm	97	Bk	98	Cf	99	Es	100	Fm	101	Md	102	No	103	Lr																																																																																																												
5	Cl	6	Si	7	P	8	S	9	Cl	10	Ar	18	Kr	36	Kr	37	Br	38	Se	39	Br	40	Xe	41	At	42	At	43	At	44	At	45	At	46	At	47	At	48	At	49	At	50	At	51	At	52	At	53	At	54	At	55	At	56	At	57	At	58	At	59	At	60	At	61	At	62	At	63	At	64	At	65	At	66	At	67	At	68	At	69	At	70	At	71	At	72	At	73	At	74	At	75	At	76	At	77	At	78	At	79	At	80	At	81	At	82	At	83	At	84	At	85	At	86	At	87	At	88	At	89	At	90	At	91	At	92	At	93	At	94	At	95	At	96	At	97	At	98	At	99	At	100	At	101	At	102	At	103	At

Mendeleev, 1869

# Scurt istoric

- sfârșitul secolului XIX:
  - mecanică clasică;
  - electromagnetism;
  - termodinamică.
- începutul secolului XX:
  - **domeniul relativist** (mecanica Newtoniană nu poate fi folosită la viteze foarte mari)
  - **domeniul microscopic** (fizica clasică nu poate fi folosită la nivel microscopic – e.g. pentru descrierea atomilor și moleculelor, a interacției cu câmpul electromagnetic etc.)

**Ce legi folosim?**

**Ce mecanică folosim?**

**Legea atracției universale**



**Legile de mișcare ale lui  
Newton – mecanica clasică**



# Câmpul electromagnetic ca sistem fizic

Fenomenele electrostaticii

Fenomenele magnetostaticii

Fenomene legate de curentii electrici

Fenomene electromagnetice cu variația rapidă în timp

Fenomene optice

**Câmpul electromagnetic ca sistem fizic  
CONCEPUTUL DE SARCINA ELECTRICA**

**ELECTRODINAMICA CLASICA**

QED – Quantum Electrodynamics

**TEORIA RELATIVITATII  
RESTRINSE**

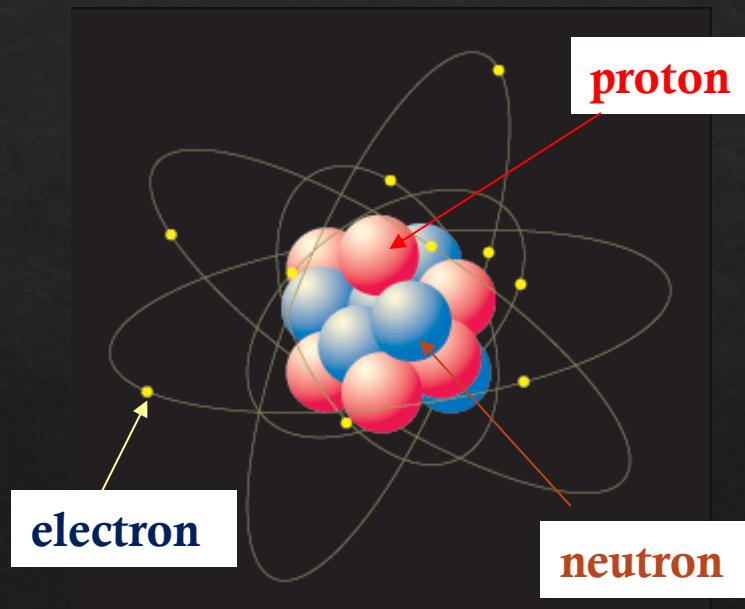
**ELECTRODINAMICA CUANTICA**

**TEORII DE ETALONARE (NE)ABELIENE**

# Atomii

Atomii:

- protoni și neutroni în nucleu
- electroni



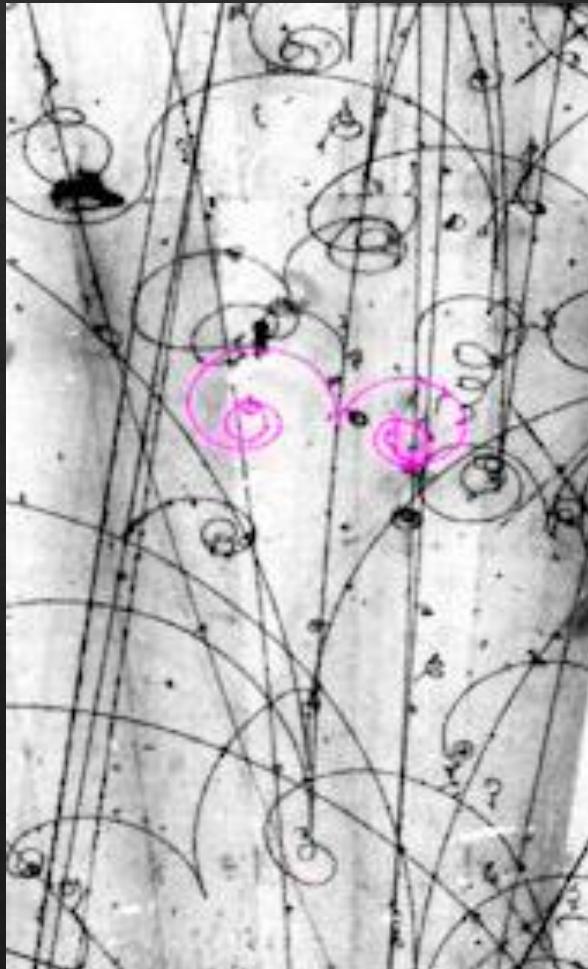
Sunt protonii și neutronii particule elementare?

# Fizica particulelor elementare

1. Care sunt particulele elementare (ce proprietăți au – masă, sarcină electrică, spin, ...)?
2. Cum interacționează?
3. Cum producem particule elementare?
4. Cum detectăm particule elementare?

# Dirac – particulă - antiparticulă

sarcină electrică  
de semn opus



- Pereche electron-positron creată din fotonii într-o cameră cu bule.
- Energia fotonului este transformată în  
•materie și anti-materie.
- Energia și impulsul se conservă (dar nu și masa de repaus)

# Yukawa – 1934

- Ce ține protonii și neutronii în nucleul?

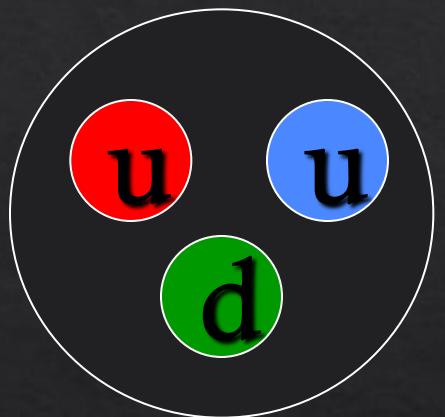
## INTERACȚIA TARE

- De ce nu o experimentăm în viața de zi cu zi?

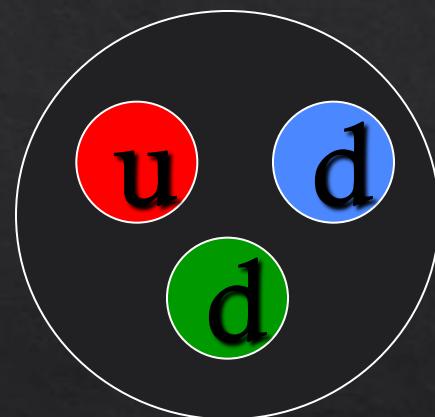
acționează la distanțe scurte

# Protonii și neutronii – modelul cuarcilor (1964)

**proton** (sarcină +1)



**neutron** (sarcină 0)



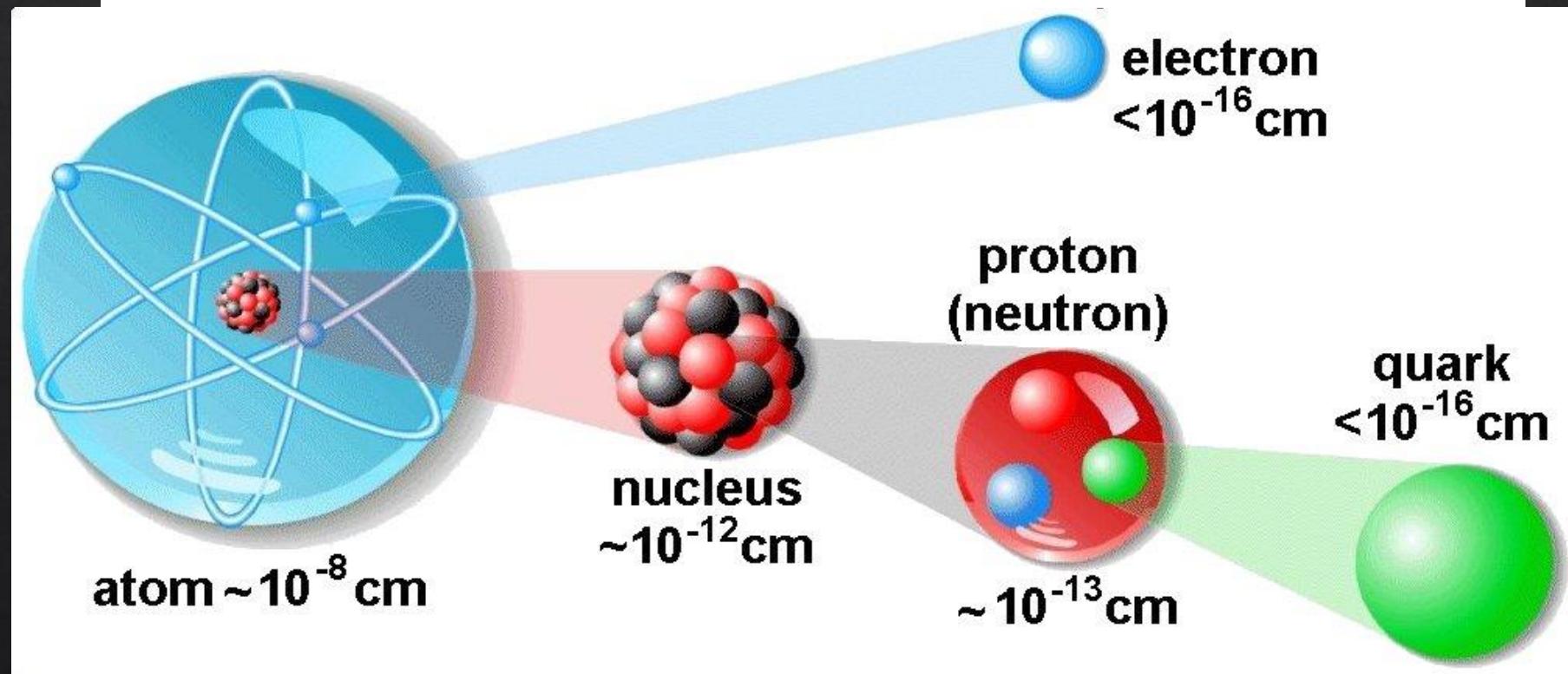
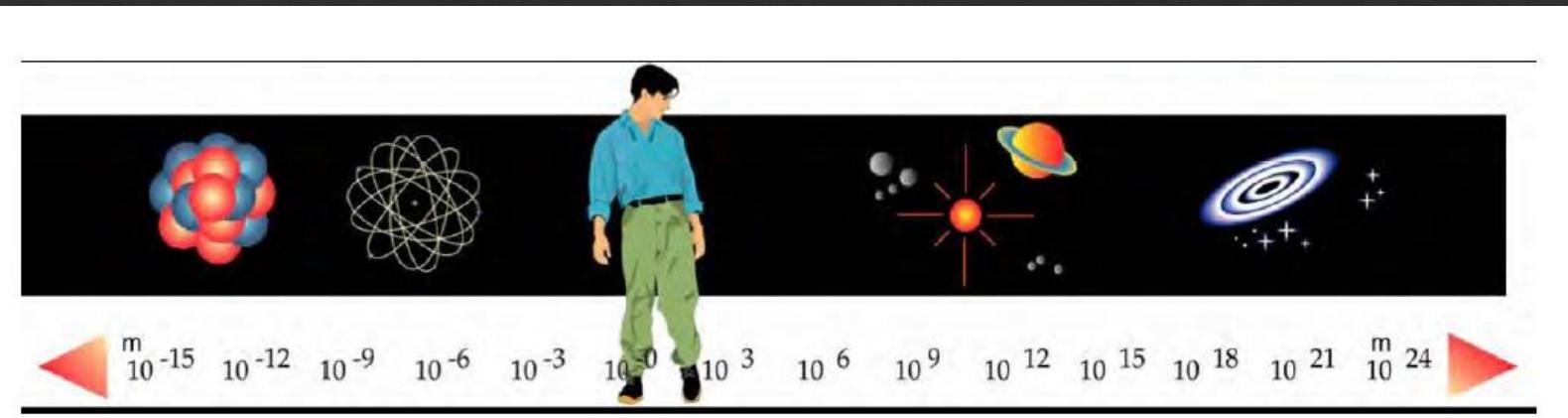
*Cuarci au sarcini electrice fractionare*

u - cuarcul up – sarcină electrică + 2/3

d - cuarcul down – sarcină electrică -1/3

$$u\left(+\frac{2}{3}\right)u\left(+\frac{2}{3}\right)d\left(-\frac{1}{3}\right) = p(+1) \quad u\left(+\frac{2}{3}\right)d\left(-\frac{1}{3}\right)d\left(-\frac{1}{3}\right) = n(0)$$

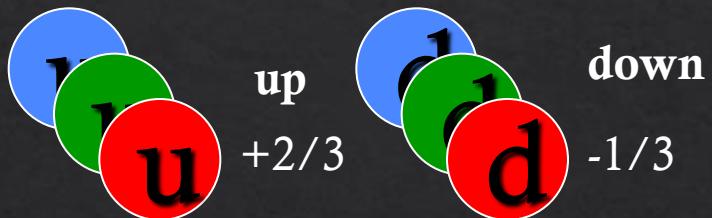
# Structura materiei (astăzi!)



# Cuarcii și culorile

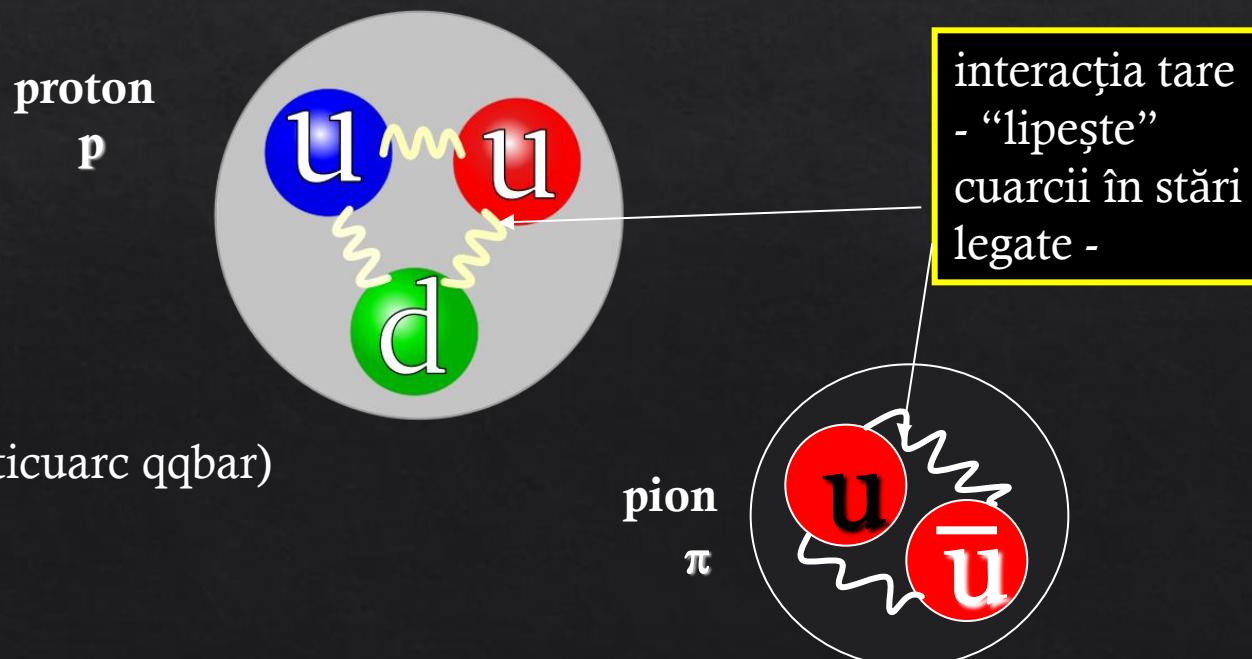
QCD – Quantum chromodynamics

Fiecare cuarc poate avea 3 “culori”



Cuarcii se combină în aşa fel încât să formeze particule „incolore” (confinare).

-**Barioni** (3 cuarci qqq: roşu + verde + albastru = alb)



-**Mezoni** (cuarc-anticuarc qqbar)

# PDG – Particle Data Group

STRANGE MESONS		CHARMED MESONS		$c\bar{c}$ MESONS		LIGHT BARYONS		BOTTOM BARYONS	
$K_L^0$	130	$D^+$	411	$\eta_c(1S)$	441	$p$	2212	$A_b^0$	5122
$K_S^0$	310	$D^0$	421	$\chi_{c0}(1P)$	10441	$n$	2112	$\Sigma_b^-$	5112
$K^0$	311	$D_0^*(2400)^+$	10411	$\eta_c(2S)$	100441	$\Delta^{++}$	2224	$\Sigma_b^0$	5212
$K^+$	321	$D_0^*(2400)^0$	10421	$J/\psi(1S)$	443	$\Delta^+$	2214	$\Sigma_b^+$	5222
$K_0^*(800)^0$	9000311*	$D^*(2010)^+$	413	$h_c(1P)$	10443	$\Delta^0$	2114	$\Sigma_b^{*-}$	5114
$K_0^*(800)^+$	9000321*	$D^*(2007)^0$	423	$\chi_{c1}(1P)$	20443	$\Delta^-$	1114	$\Sigma_b^{*0}$	5214
$K_0^*(1430)^0$	10311	$D_1(2420)^+$	10413	$\psi(2S)$	100443	STRANGE BARYONS		$\Sigma_b^{*+}$	5224
$K_0^*(1430)^+$	10321	$D_1(2420)^0$	10423	$\psi(3770)$	30443	$\Lambda$	3122	$\Xi_b^-$	5132
$K(1460)^0$	100311	$D_1(H)^+$	20413	$\psi(4040)$	9000443	$\Sigma^+$	3222	$\Xi_b^0$	5232
$K(1460)^+$	100321	$D_1(2430)^0$	20423	$\psi(4160)$	9010443	$\Sigma^0$	3212	$\Xi_b'^-$	5312
$K(1830)^0$	9010311*	$D_2^*(2460)^+$	415	$\psi(4415)$	9020443	$\Sigma^-$	3112	$\Xi_b'^0$	5322
$K(1830)^+$	9010321*	$D_2^*(2460)^0$	425	$\chi_{c2}(1P)$	445	$\Sigma^{*+}$	3224 <sup>d</sup>	$\Xi_b^{*-}$	5314
$K_0^*(1950)^0$	9020311*	$D_s^+$	431	$\chi_{c2}(2P)$	100445*	$\Sigma^{*-}$	3114 <sup>d</sup>	$\Xi_b^{*0}$	5324
$K_0^*(1950)^+$	9020321*	$D_{s0}^*(2317)^+$	10431			$\Xi^0$	3322	$\Omega_b^-$	5332
$K^*(892)^0$	313	$D_s^+$	433	$b\bar{b}$ MESONS		$\Xi^-$	3312	$\Omega_b^0$	5334
$K^*(892)^+$	323	$D_{s1}(2536)^+$	10433	$\eta_b(1S)$	551	$\Xi^{*0}$	3324 <sup>d</sup>	$\Omega_b^{*-}$	5142
$K_1(1270)^0$	10313	$D_{s1}(2460)^+$	20433	$\chi_{b0}(1P)$	10551	$\Xi^{*-}$	3314 <sup>d</sup>	$\Xi_{bc}^0$	5242
$K_1(1270)^+$	10323	$D_{s2}^*(2573)^+$	435	$\eta_b(2S)$	100551	$\Omega^-$	3334	$\Xi_{bc}^+$	5412
$K_1(1400)^0$	20313			$\chi_{b0}(2P)$	110551	CHARMED BARYONS		$\Xi_{bc}'^0$	5422
$K_1(1400)^+$	20323	BOTTOM MESONS		$\eta_b(3S)$	200551	$A_c^+$	4122	$\Xi_{bc}'^+$	5414
$K^*(1410)^0$	100313	$B^0$	511	$\chi_{b0}(3P)$	210551	$\Sigma_c^{++}$	4222	$\Xi_{bc}^{*0}$	5424
$K^*(1410)^+$	100323	$B^+$	521	$\Upsilon(1S)$	553	$\Sigma_c^+$	4212	$\Xi_{bc}^{*+}$	5342
$K_1(1650)^0$	9000313*	$B_0^{*0}$	10511	$h_b(1P)$	10553	$\Sigma_c^0$	4112	$\Omega_{bc}^0$	5432
$K_1(1650)^+$	9000323*	$B_0^{*+}$	10521	$\chi_{b1}(1P)$	20553	$\Sigma_c^{*++}$	4224	$\Omega_{bc}'^0$	5432
$K^*(1680)^0$	30313	$B^{*0}$	513	$\Upsilon_1(1D)$	30553	$\Sigma_c^{*+}$	4214		

# Universul este alcătuit numai din cuarci și electroni?

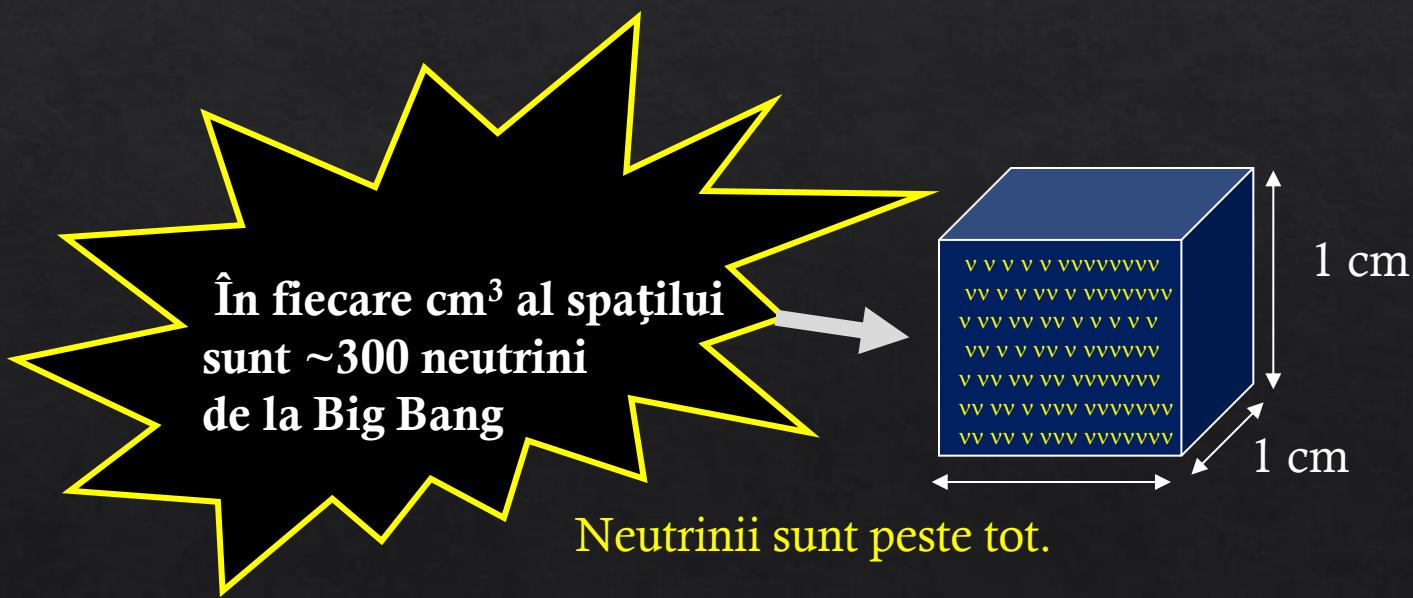
*Există și neutrini!*



Electronul, protonul și neutronul sunt rari!

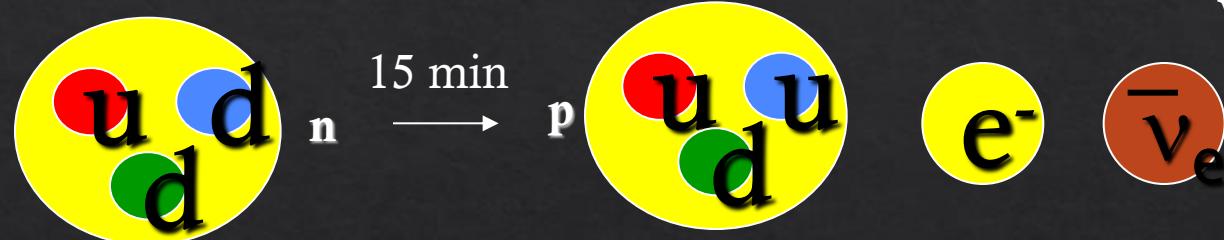
Pentru fiecare dintre ei, există 1 billion neutrini.

*Neutrinii sunt cele mai abundente particule ale materiei în univers.*



# Dezintegrarea $\beta$

La nivelul cuarcilor:  $d \rightarrow u e^- \bar{\nu}_e$

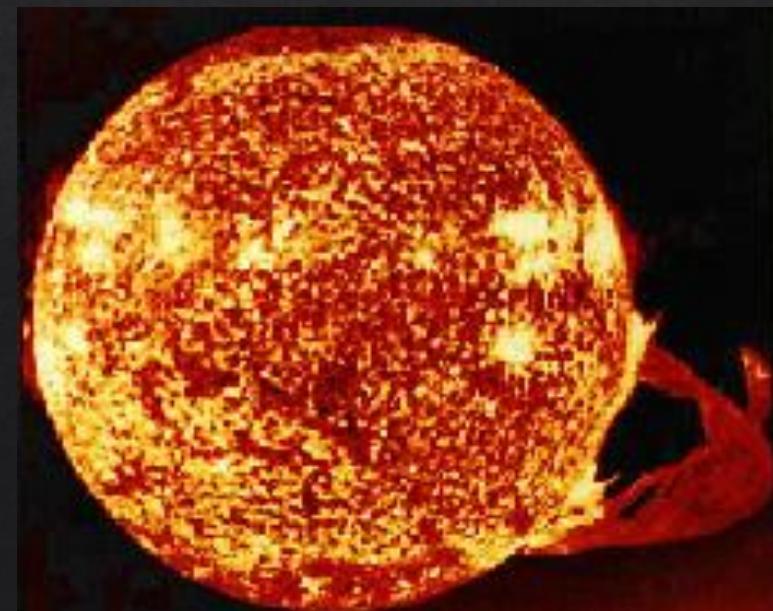
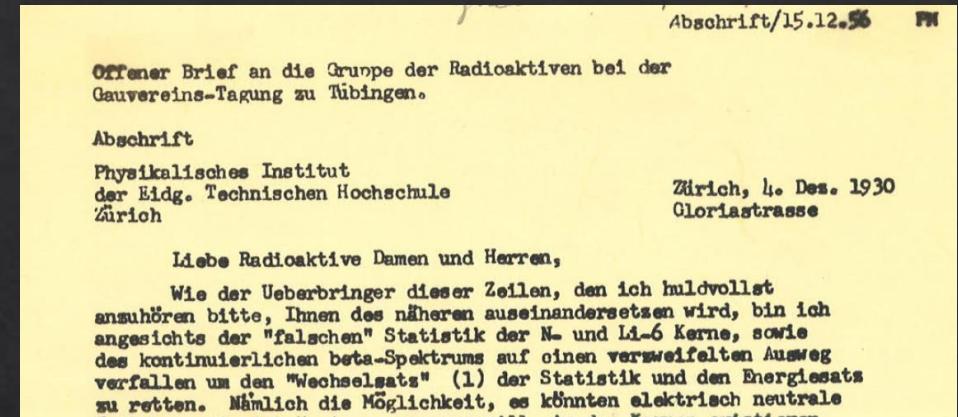


Un neutron se dezintegrează în 15 minute.

interacție „slabă”!

QFD – Quantum flavor dynamics

!

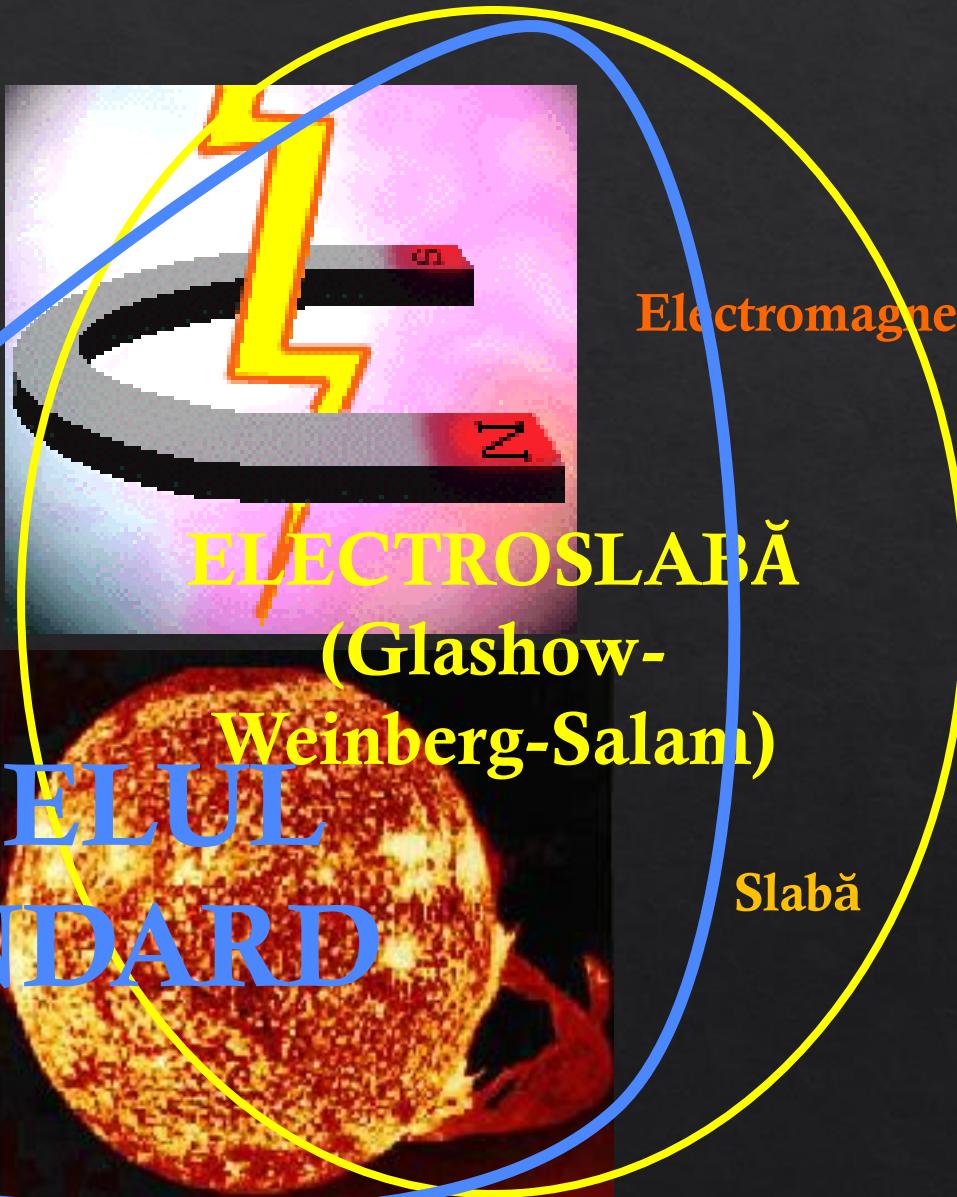
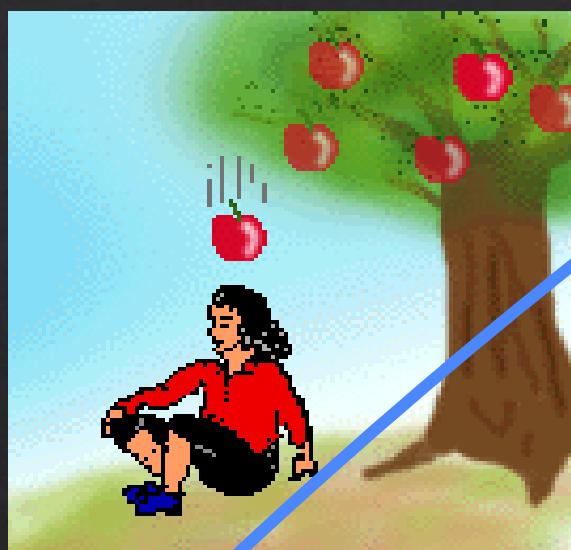


# Tipuri de interacție

Gravitațională

masa

Tare



# Cine mediază interacțiile?

Gravitațională  
? graviton ?

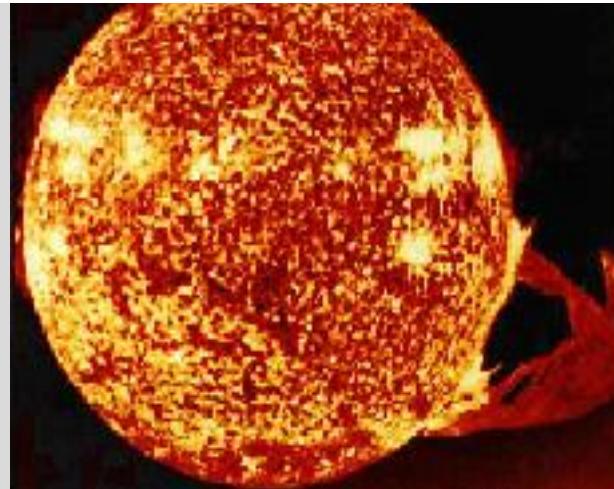
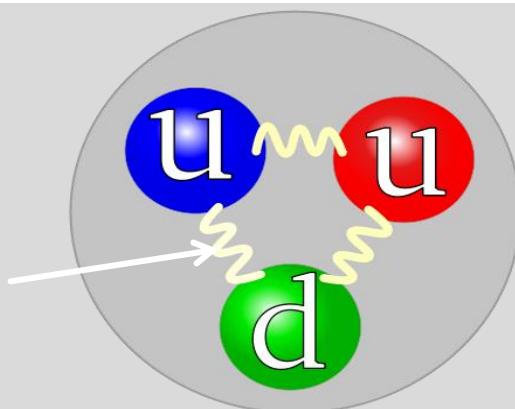


QED – Quantum  
Electrodynamics

Electromagnetică  
 $\gamma$ - fotonul

$$G_{SM} = SU(3)_C \times SU(2)_L \times U(1)_Y \rightarrow SU(3)_C \times U(1)_Q$$

Tare  
gluonii  
QCD – Quantum  
chromodynamics



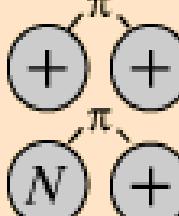
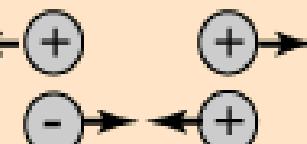
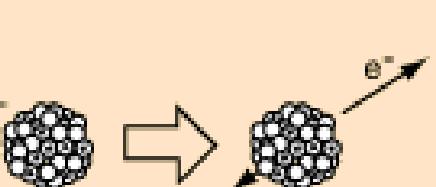
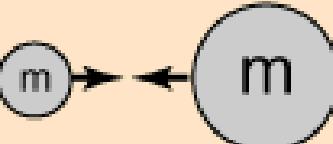
QFD – Quantum  
flavordynamics

Slabă

$W^+$ ,  $W^-$ ,  $Z$

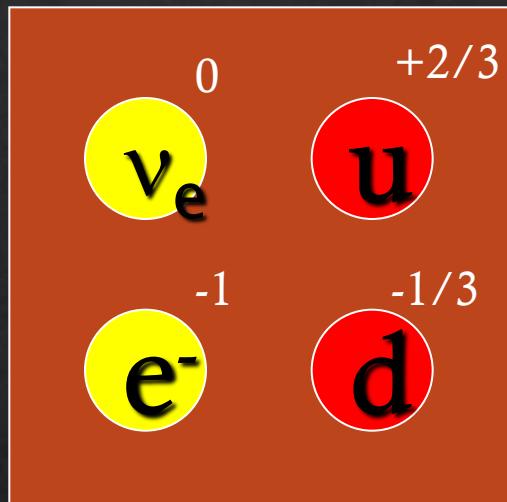
# Fundamental Forces

<http://hyperphysics.phy-astr.gsu.edu/hbase/Forces/funfor.html>

		Strength	Range (m)	Particle
<i>Strong</i>	 Force which holds nucleus together	1	$10^{-15}$ (diameter of a medium sized nucleus)	gluons, $\pi$ (nucleons)
<i>Electro-magnetic</i>		$\frac{1}{137}$	Infinite	photon mass = 0 spin = 1
<i>Weak</i>	 neutrino interaction induces beta decay	$10^{-6}$	$10^{-18}$ (0.1% of the diameter of a proton)	Intermediate vector bosons $W^+$ , $W^-$ , $Z_0$ , mass > 80 GeV spin = 1
<i>Gravity</i>		$6 \times 10^{-39}$	Infinite	graviton ? mass = 0 spin = 2

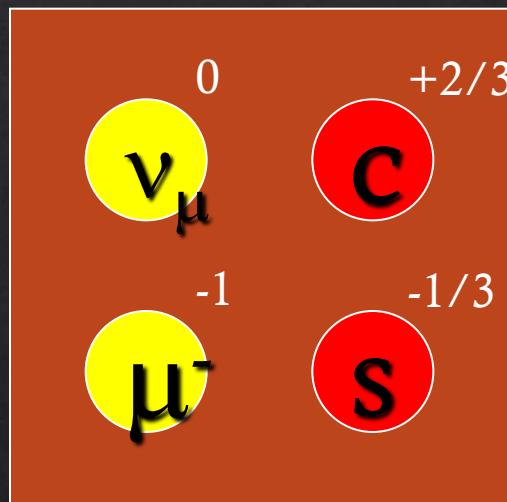
# 3 FAMILII (ASTĂZI!)

prima generație



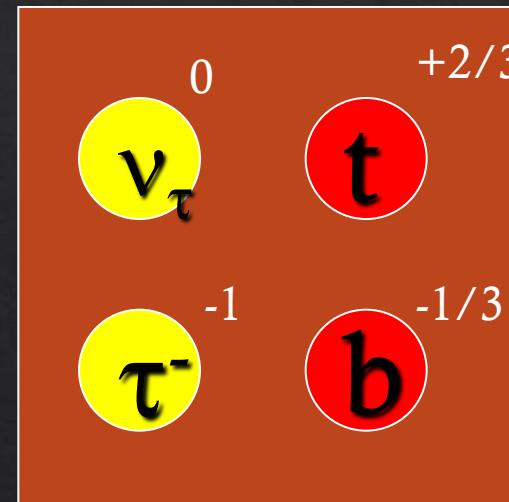
materie „obișnuită”

a doua generație



radiație cosmică

a treia generație



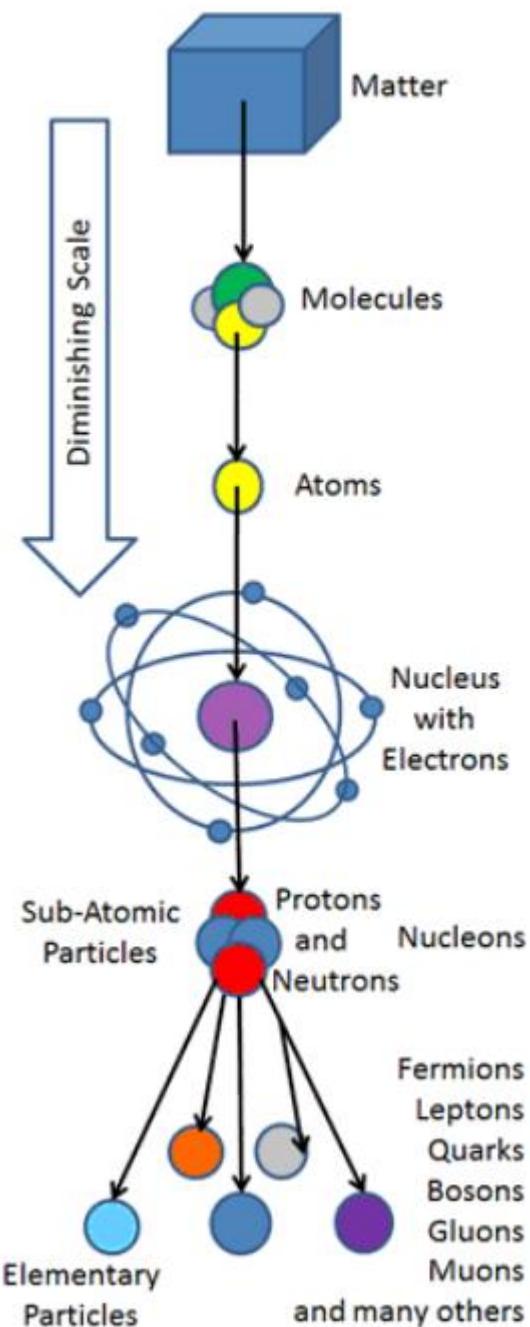
acceleratori

*cele 3 generații diferă prin masă!*

## Three Generations of Matter (Fermions)

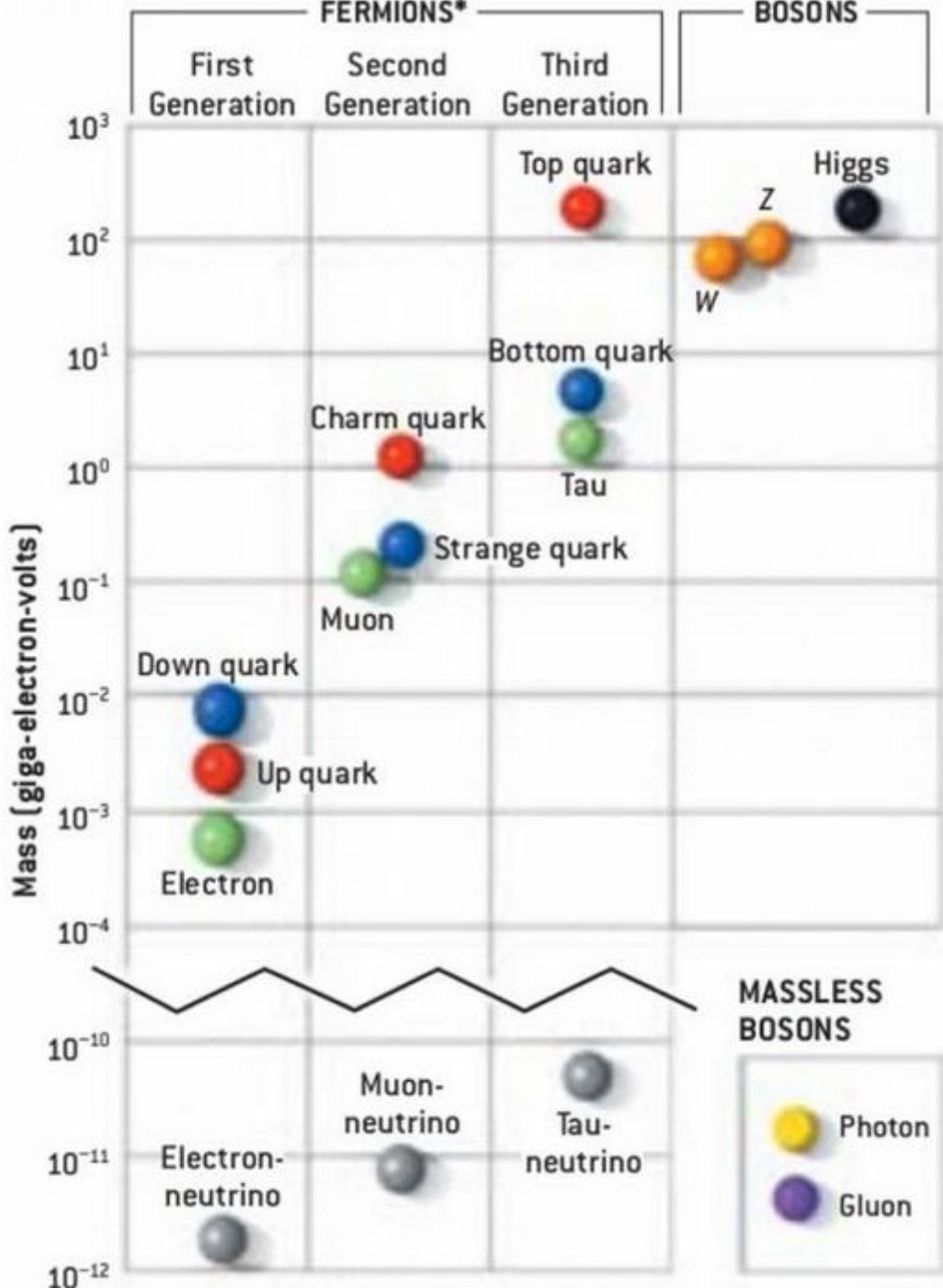
	I	II	III	
mass→	2.4 MeV	1.27 GeV	171.2 GeV	0
charge→	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0
spin→	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1
name→	up	charm	top	photon
Quarks	<b>u</b>  down	<b>c</b>  strange	<b>t</b>  bottom	<b>γ</b>  gluon
Leptons	$<2.2$ eV $0$ $\frac{1}{2}$ <b>v<sub>e</sub></b>  electron neutrino	$<0.17$ MeV $0$ $\frac{1}{2}$ <b>v<sub>μ</sub></b>  muon neutrino	$<15.5$ MeV $0$ $\frac{1}{2}$ <b>v<sub>τ</sub></b>  tau neutrino	$91.2$ GeV $0$ $1$ <b>Z<sup>0</sup></b>  weak force
Bosons (Forces)	0.511 MeV $-1$ $\frac{1}{2}$ <b>e</b>  electron	105.7 MeV $-1$ $\frac{1}{2}$ <b>μ</b>  muon	1.777 GeV $-1$ $\frac{1}{2}$ <b>τ</b>  tau	$80.4$ GeV $\pm 1$ $1$ <b>W<sup>±</sup></b>  weak force

## Particle Hierarchy



FERMIONS\*

BOSONS



număr asociat sarcinii electrice (Q)

număr leptonic L ( $L_e$ ,  $L_\mu$ ,  $L_\tau$ )

număr barionic B



	Gravity	Weak (Electroweak)	Electromagnetic	Strong
Carried By	Graviton (not yet observed)	$W^+$ $W^-$ $Z^0$	Photon	Gluon
Acts on	All	Quarks and Leptons	Quarks and Charged Leptons and $W^+$ $W^-$	Quarks and Gluons

interacție slabă

(cuarci și leptonii) interacție electromagnetică

(particule cu sarcină electrică: cuarci,  
leptoni cu sarcină electrică)

interacție tare (cuarci)

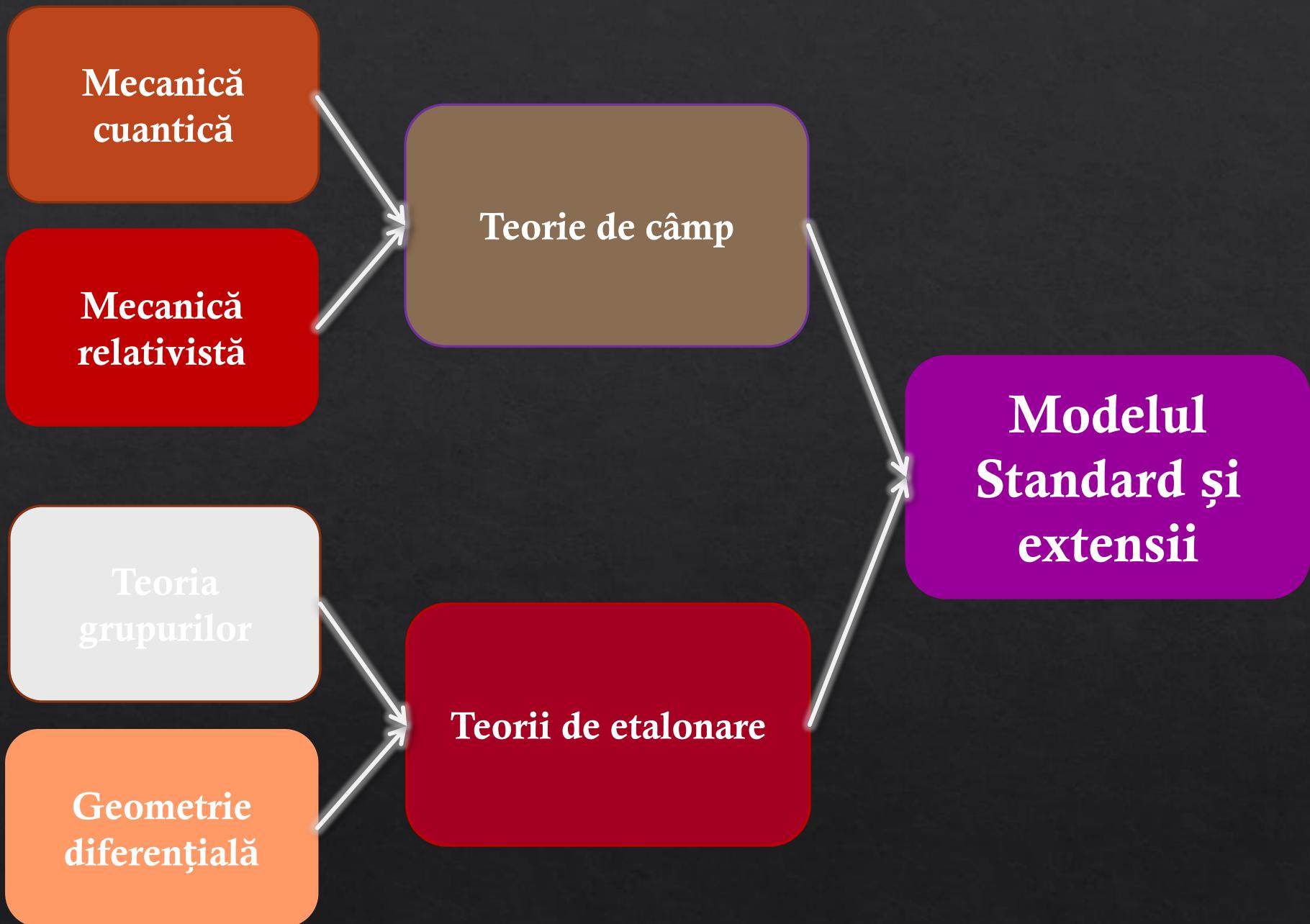
indiscernabilitatea  
particulelor identice

statistică cuantică

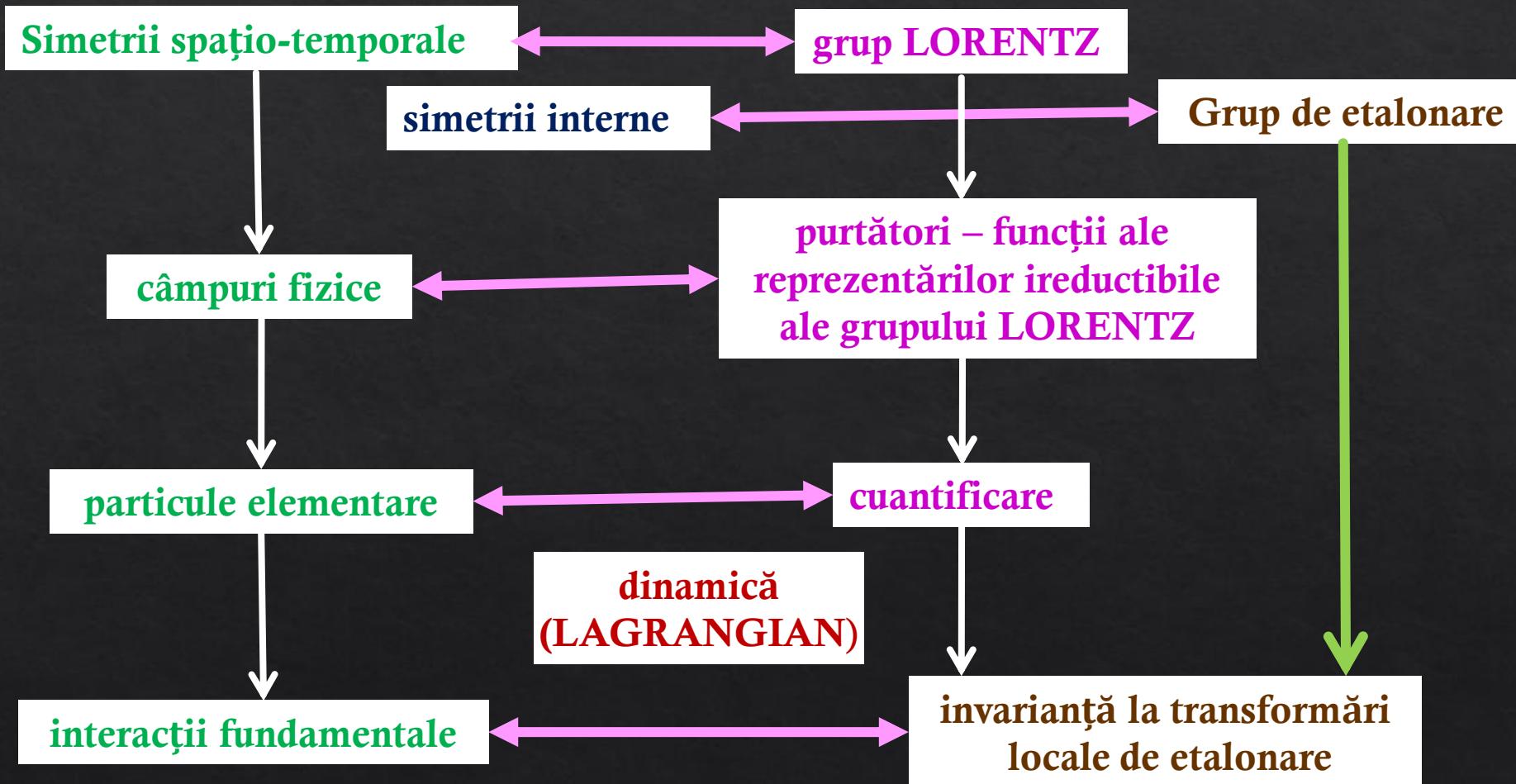
existența antiparticulelor

procese de creare și  
anihilare a particulelor

interacție cu schimb de particule



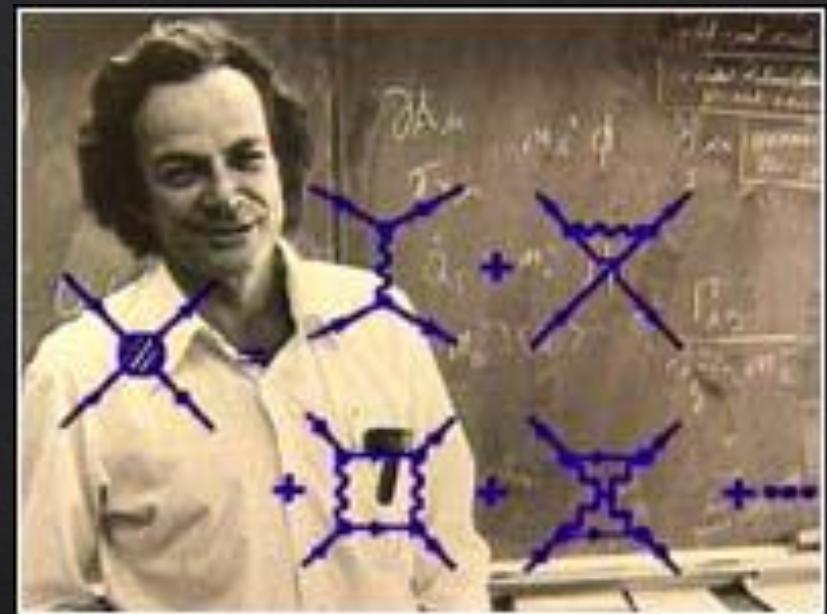
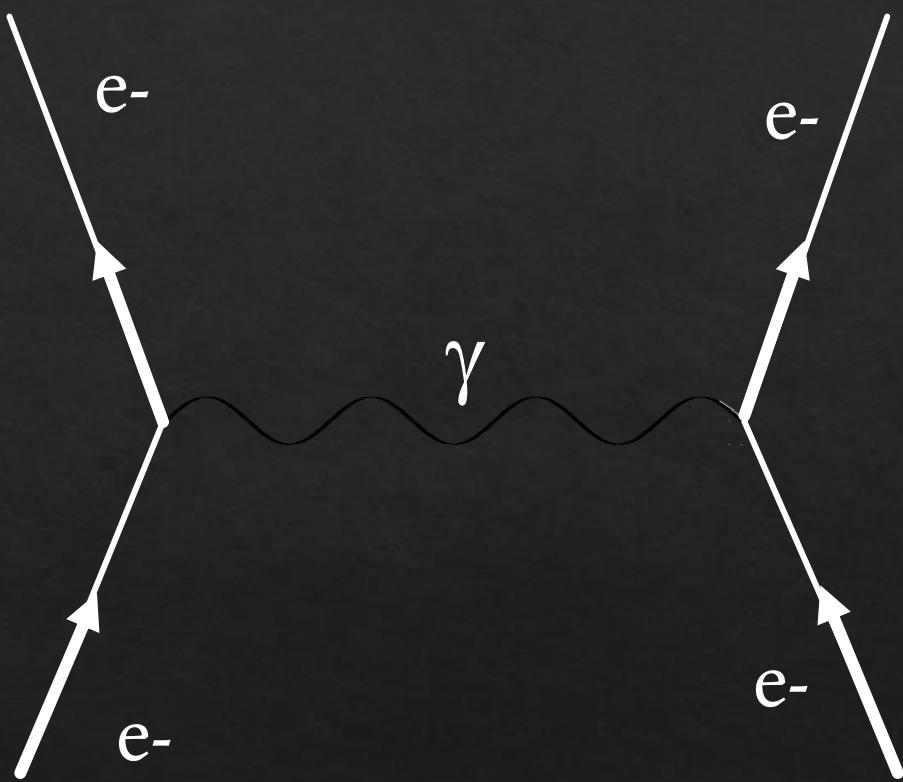
# Schemă teoretică: de la simetrii spațio-temporale la particule elementare și interacții



# Lagrangian – Glashow-Weinberg-Salam (interacție electro-slabă)

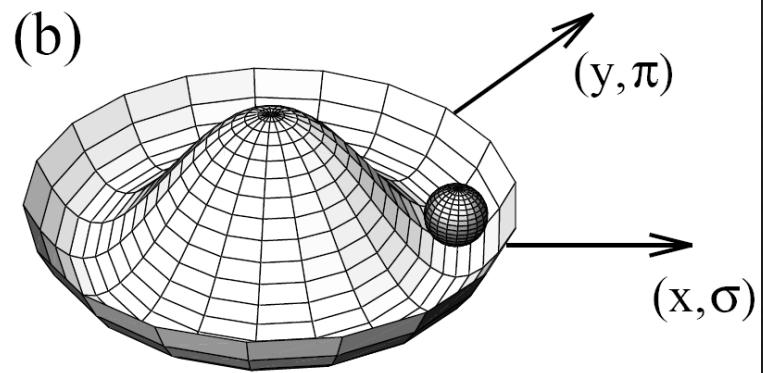
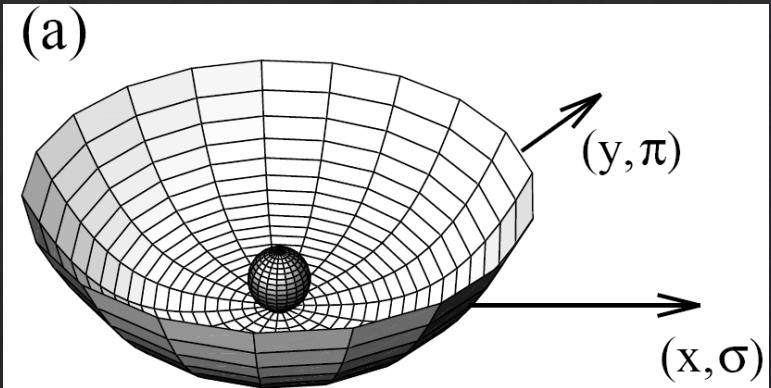
$$\begin{aligned}
\mathcal{L}_{GWS} = & \sum_f (\bar{\Psi}_f (i\gamma^\mu \partial_\mu - m_f) \Psi_f - e Q_f \bar{\Psi}_f \gamma^\mu \Psi_f A_\mu) + \\
& + \frac{g}{\sqrt{2}} \sum_i (\bar{a}_L^i \gamma^\mu b_L^i W_\mu^+ + \bar{b}_L^i \gamma^\mu a_L^i W_\mu^-) + \frac{g}{2c_w} \sum_f \bar{\Psi}_f \gamma^\mu (I_f^3 - 2s_w^2 Q_f - I_f^3 \gamma_5) \Psi_f Z_\mu + \\
& - \frac{1}{4} |\partial_\mu A_\nu - \partial_\nu A_\mu - ie(W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 - \frac{1}{2} |\partial_\mu W_\nu^+ - \partial_\nu W_\mu^+ + \\
& - ie(W_\mu^+ A_\nu - W_\nu^+ A_\mu) + ig' c_w (W_\mu^+ Z_\nu - W_\nu^+ Z_\mu)|^2 + \\
& - \frac{1}{4} |\partial_\mu Z_\nu - \partial_\nu Z_\mu + ig' c_w (W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 + \\
& - \frac{1}{2} M_\eta^2 \eta^2 - \frac{g M_\eta^2}{8 M_W} \eta^3 - \frac{g'^2 M_\eta^2}{32 M_W} \eta^4 + |M_W W_\mu^+ + \frac{g}{2} \eta W_\mu^+|^2 + \\
& + \frac{1}{2} |\partial_\mu \eta + i M_Z Z_\mu + \frac{ig}{2c_w} \eta Z_\mu|^2 - \sum_f \frac{g}{2} \frac{m_f}{M_W} \bar{\Psi}_f \Psi_f \eta
\end{aligned}$$

# Feynman – diagrame și reguli

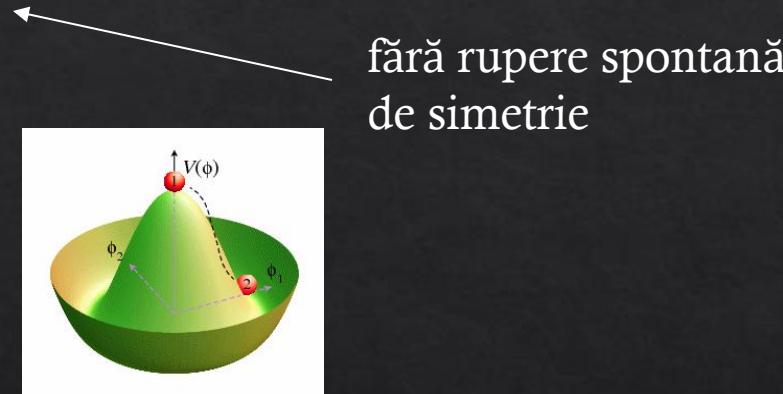


Feynman earned his Nobel for creating  
these diagrams  
(Courtesy Auckland University)

# De unde apare masa particulelor în teorie?



**W<sup>+</sup>, W<sup>-</sup>, Z – au masă**  
- rezultă **bozonul Higgs**



fără rupere spontană  
de simetrie

rupere spontană de simetrie

# Quarks

<i>u</i>	<i>c</i>	<i>t</i>
up	charm	top

<i>d</i>	<i>s</i>	<i>b</i>
down	strange	bottom

<i>e</i>	$\mu$	$\tau$
electron	muon	tau

$\nu_e$	$\nu_\mu$	$\nu_\tau$
electron neutrino	muon neutrino	tau neutrino

# Leptons

# Forces

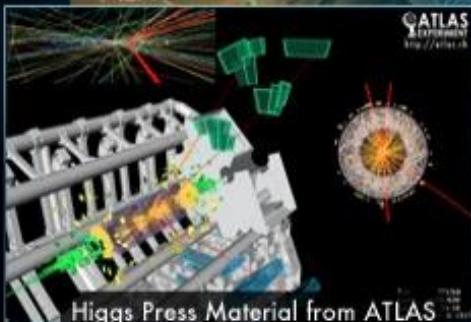
$Z$	$\gamma$
Z boson	photon

$W$	$g$
W boson	gluon

H  
Higgs  
boson

Congratulations to Professors  
**François Englert & Peter Higgs**  
for the  
**2013 Nobel Prize in Physics**



The ATLAS and CMS experiments at CERN congratulate Professors François Englert and Peter Higgs for their pioneering work in identifying the electro-weak-symmetry-breaking mechanism. CMS and ATLAS independently announced the discovery of a new particle on 4 July 2012, later identified as a Higgs boson, confirming the predictions of Professors Higgs, Englert and others in seminal papers published in 1964. We join in this celebration of the triumph of human curiosity and ingenuity.

Higgs boson discovery was the culmination of the decades of dedicated and intense work by so many collaborators in designing, building and operating ATLAS, and in understanding and analysing the data. None of it would have been possible without the huge dedication also of the LHC accelerator team, the worldwide distributed computing teams, and the continuing support of the governments and funding agencies of the 38 countries home to our 177 member institutes.

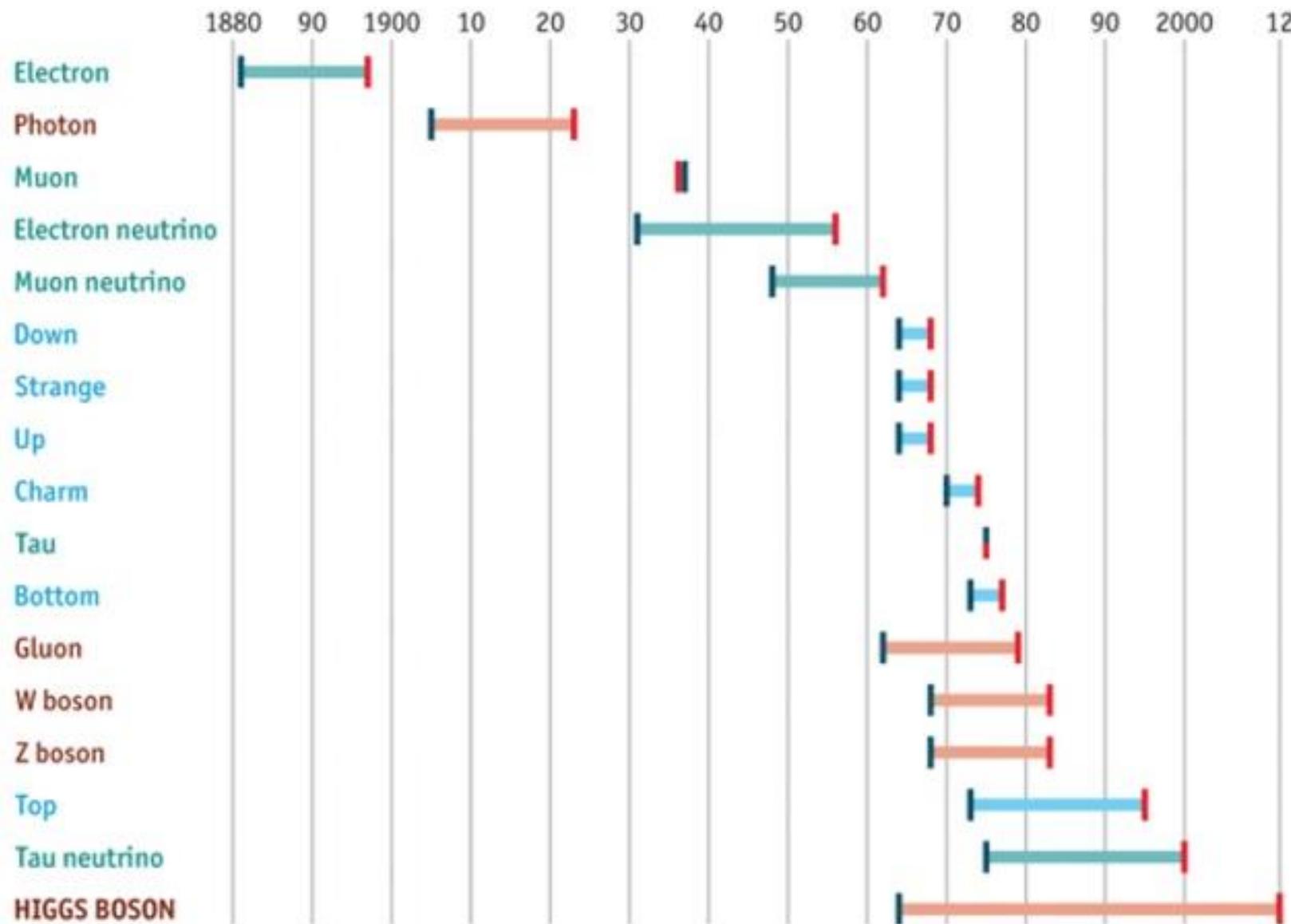
We can all feel proud that our experimental observations demonstrated that the insights rewarded by the Nobel prize are realised in nature.

## The Standard Model of particle physics

Years from concept to discovery

Leptons  
Bosons  
Quarks

Theorised/explained  
Discovered



# Modelul Standard

mass → $\approx 2.3 \text{ MeV}/c^2$	charge → 2/3	spin → 1/2	mass → $\approx 1.275 \text{ GeV}/c^2$	charge → 2/3	spin → 1/2	mass → $\approx 173.07 \text{ GeV}/c^2$	charge → 2/3	spin → 1/2	mass → 0	charge → 0	spin → 0	mass → $\approx 126 \text{ GeV}/c^2$	charge → 0	spin → 0
up	charm	top	gluon	Higgs boson										
down	strange	bottom	photon											
electron	muon	tau	Z boson											
electron neutrino	muon neutrino	tau neutrino	W boson											

+ antiparticule  
(Q, L, B opus)!

# Standard Model of Elementary Particles

three generations of matter (elementary fermions)			three generations of antimatter (elementary antifermions)			interactions / force carriers (elementary bosons)	
I	II	III	I	II	III		
mass	=2.2 MeV/c <sup>2</sup>	=1.28 GeV/c <sup>2</sup>	=173.1 GeV/c <sup>2</sup>	=2.2 MeV/c <sup>2</sup>	=1.28 GeV/c <sup>2</sup>	=173.1 GeV/c <sup>2</sup>	0
charge	2/3	2/3	2/3	-2/3	-2/3	-2/3	0
spin	1/2	1/2	1/2	1/2	1/2	1/2	1
QUARKS	up	charm	top	antiup	anticharm	antitop	gluon
	u	c	t	ū	c̄	t̄	g
	down	strange	bottom	antidown	antistrange	antibottom	higgs
	d	s	b	d̄	s̄	b̄	H
LEPTONS	electron	muon	tau	positron	antimuon	antitau	Z <sup>0</sup> boson
	e	μ	τ	e <sup>+</sup>	μ̄	τ̄	Z
	<2.2 eV/c <sup>2</sup>	<0.17 MeV/c <sup>2</sup>	<18.2 MeV/c <sup>2</sup>	<2.2 eV/c <sup>2</sup>	<0.17 MeV/c <sup>2</sup>	<18.2 MeV/c <sup>2</sup>	W <sup>+</sup> boson
	0	0	0	0	0	0	W <sup>-</sup> boson
	1/2	1/2	1/2	1/2	1/2	1/2	1
	electron neutrino	muon neutrino	tau neutrino	electron antineutrino	muon antineutrino	tau antineutrino	
	v <sub>e</sub>	v <sub>μ</sub>	v <sub>τ</sub>	v̄ <sub>e</sub>	v̄ <sub>μ</sub>	v̄ <sub>τ</sub>	
GAUGE BOSONS	vector bosons	SCALAR BOSONS					

# Fizica particulelor elementare

1. Care sunt particulele elementare (ce proprietăți au – masă, sarcină electrică, spin, ...)?
2. Cum interacționează? - De unde obținem informații?
3. Cum producem particule elementare?
4. Cum detectăm particule elementare?

# The Review of Particle Physics (2023)

R.L. Workman *et al.* (Particle Data Group), Prog. Theor. Exp. Phys. 2022, 083C01 (2022) and 2023 update

<b>Gauge &amp; Higgs Bosons</b>	<a href="#">reviews</a>	<b>Leptons</b>	<a href="#">reviews</a>	<b>Quarks</b>	<a href="#">reviews</a>
$\gamma$ gluon graviton $W$ $Z$ $H$ Neutral Higgs Bosons, Searches for Charged Higgs Bosons ( $H^\pm, H^{\pm\pm}$ ) Heavy Bosons Axions		$e$ $\mu$ $\tau$ Heavy Charged Lepton Neutrino Properties Number of Neutrino Types Double $\beta$ -Decay Neutrino Mixing Heavy Neutral Leptons		Light quarks ( $u, d, s$ ) $c$ $b$ $t$ $b'$ $t'$ Free quark	
<b>Mesons</b>	<a href="#">reviews</a>	<b>Baryons</b>	<a href="#">reviews</a>	<b>Other Searches</b>	<a href="#">reviews</a>
Light Unflavored Strange Charmed Charmed, Strange (incl. possibly non- $q\bar{q}$ states) Bottom Bottom, Strange Bottom, Charmed $c\bar{c}$ (incl. possibly non- $q\bar{q}$ states) $b\bar{b}$ (incl. possibly non- $q\bar{q}$ states) Other Mesons		$N$ Baryons $\Delta$ Baryons $\Lambda$ Baryons $\Sigma$ Baryons $\Xi$ Baryons $\Omega$ Baryons Charmed Baryons Doubly-Charmed Bottom Baryons Exotic Baryons		Magnetic Monopole Supersymmetric Particles Technicolor Quark and Lepton Compositeness Extra Dimensions WIMPs Other Particle Searches	
				<b>Conservation Laws</b>	<a href="#">reviews</a>
				Discrete Space-Time Symm. Number Conservation Laws	

Quantity	Symbol, equation	Value	Uncertainty (ppb)
speed of light in vacuum	$c$	299 792 458 m s <sup>-1</sup>	exact
Planck constant	$h$	6.626 070 15×10 <sup>-34</sup> J s (or J/Hz) <sup>‡</sup>	exact
Planck constant, reduced	$\hbar \equiv h/2\pi$	1.054 571 817... × 10 <sup>-34</sup> J s = 6.582 119 569... × 10 <sup>-22</sup> MeV s	exact*, exact*
electron charge magnitude	$e$	1.602 176 634×10 <sup>-19</sup> C	exact
conversion constant	$\hbar c$	197.326 980 4... MeV fm	exact*
conversion constant	$(\hbar c)^2$	0.389 379 372 1... GeV <sup>2</sup> mbarn	exact*
electron mass	$m_e$	0.510 998 950 00(15) MeV/c <sup>2</sup> = 9.109 383 7015(28)×10 <sup>-31</sup> kg	0.30
proton mass	$m_p$	938.272 088 16(29) MeV/c <sup>2</sup> = 1.672 621 923 69(51)×10 <sup>-27</sup> kg = 1.007 276 466 621(53) u = 1836.152 673 43(11) $m_e$ 0.053, 0.060	0.31
neutron mass	$m_n$	939.565 420 52(54) MeV/c <sup>2</sup> = 1.008 664 915 95(49) u 0.57, 0.48	
deuteron mass	$m_d$	1875.612 942 57(57) MeV/c <sup>2</sup>	0.30
unified atomic mass unit**	$u = (\text{mass } {}^{12}\text{C atom})/12$	931.494 102 42(28) MeV/c <sup>2</sup> = 1.660 539 066 60(50)×10 <sup>-27</sup> kg	0.30
permittivity of free space	$\epsilon_0 = 1/\mu_0 c^2$	8.854 187 8128(13) × 10 <sup>-12</sup> F m <sup>-1</sup>	0.15
permeability of free space	$\mu_0/(4\pi \times 10^{-7})$	1.000 000 000 55(15) N A <sup>-2</sup>	0.15
fine-structure constant	$\alpha = e^2/4\pi\epsilon_0\hbar c$	7.297 352 5693(11)×10 <sup>-3</sup> = 1/137.035 999 084(21) <sup>†</sup>	0.15
classical electron radius	$r_e = e^2/4\pi\epsilon_0 m_e c^2$	2.817 940 3262(13)×10 <sup>-15</sup> m	0.45
( $e^-$ Compton wavelength)/ $2\pi$	$\lambda_e = \hbar/m_e c = r_e \alpha^{-1}$	3.861 592 6796(12)×10 <sup>-13</sup> m	0.30
Bohr radius ( $m_{\text{nucleus}} = \infty$ )	$a_\infty = 4\pi\epsilon_0\hbar^2/m_e e^2 = r_e \alpha^{-2}$	0.529 177 210 903(80)×10 <sup>-10</sup> m	0.15
wavelength of 1 eV/c particle	$hc/(1 \text{ eV})$	1.239 841 984... × 10 <sup>-6</sup> m	exact*
Rydberg energy	$hcR_\infty = m_e e^4 / 2(4\pi\epsilon_0)^2 \hbar^2 = m_e c^2 \alpha^2 / 2$	13.605 693 122 994(26) eV	1.9×10 <sup>-3</sup>
Thomson cross section	$\sigma_T = 8\pi r_e^2/3$	0.665 245 873 21(60) barn	0.91
Bohr magneton	$\mu_B = e\hbar/2m_e$	5.788 381 8060(17)×10 <sup>-11</sup> MeV T <sup>-1</sup>	0.3
nuclear magneton	$\mu_N = e\hbar/2m_p$	3.152 451 258 44(96)×10 <sup>-14</sup> MeV T <sup>-1</sup>	0.31
electron cyclotron freq./field	$\omega_{\text{cycl}}^e/B = e/m_e$	1.758 820 010 76(53)×10 <sup>11</sup> rad s <sup>-1</sup> T <sup>-1</sup>	0.30
proton cyclotron freq./field	$\omega_{\text{cycl}}^p/B = e/m_p$	9.578 833 1560(29)×10 <sup>7</sup> rad s <sup>-1</sup> T <sup>-1</sup>	0.31

gravitational constant <sup>‡</sup>	$G_N$	$6.674\ 30(15) \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$ $= 6.708\ 83(15) \times 10^{-39} \hbar c (\text{GeV}/c^2)^{-2}$	$2.2 \times 10^4$ $2.2 \times 10^4$
standard gravitational accel.	$g_N$	$9.806\ 65 \text{ m s}^{-2}$	exact
Avogadro constant	$N_A$	$6.022\ 140\ 76 \times 10^{23} \text{ mol}^{-1}$	exact
Boltzmann constant	$k$	$1.380\ 649 \times 10^{-23} \text{ J K}^{-1}$ $= 8.617\ 333\ 262\dots \times 10^{-5} \text{ eV K}^{-1}$	exact exact*
molar volume, ideal gas at STP	$N_A k (273.15 \text{ K})/(101\ 325 \text{ Pa})$	$22.413\ 969\ 54\dots \times 10^{-3} \text{ m}^3 \text{ mol}^{-1}$	exact*
Wien displacement law constant	$b = \lambda_{\max} T$	$2.897\ 771\ 955\dots \times 10^{-3} \text{ m K}$	exact*
Stefan-Boltzmann constant	$\sigma = \pi^2 k^4 / 60 \hbar^3 c^2$	$5.670\ 374\ 419\dots \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$	exact*
Fermi coupling constant <sup>‡‡</sup>	$G_F/(\hbar c)^3$	$1.166\ 378\ 7(6) \times 10^{-5} \text{ GeV}^{-2}$	510
weak-mixing angle	$\sin^2 \hat{\theta}(M_Z) (\overline{\text{MS}})$	0.231 22(4) <sup>††</sup>	$1.7 \times 10^5$
$W^\pm$ boson mass	$m_W$	80.379(12) $\text{GeV}/c^2$	$1.5 \times 10^5$
$Z^0$ boson mass	$m_Z$	91.1876(21) $\text{GeV}/c^2$	$2.3 \times 10^4$
strong coupling constant	$\alpha_s(m_Z)$	0.1179(10)	$8.5 \times 10^6$
$\pi = 3.141\ 592\ 653\ 589\ 793\ 238\dots$		$e = 2.718\ 281\ 828\ 459\ 045\ 235\dots$	$\gamma = 0.577\ 215\ 664\ 901\ 532\ 860\dots$
1 in $\equiv 0.0254 \text{ m}$	1 G $\equiv 10^{-4} \text{ T}$	$1 \text{ eV} = 1.602\ 176\ 634 \times 10^{-19} \text{ J}$ (exact)	$kT$ at 300 K $= [38.681\ 740(22)]^{-1} \text{ eV}$
1 Å $\equiv 0.1 \text{ nm}$	1 dyne $\equiv 10^{-5} \text{ N}$	$(1 \text{ kg})c^2 = 5.609\ 588\ 603\dots \times 10^{35} \text{ eV}$ (exact*)	0 °C $\equiv 273.15 \text{ K}$
1 barn $\equiv 10^{-28} \text{ m}^2$	1 erg $\equiv 10^{-7} \text{ J}$	$1 \text{ C} = 2.997\ 924\ 58 \times 10^9 \text{ esu}$	1 atmosphere $\equiv 760 \text{ Torr} \equiv 101\ 325 \text{ Pa}$

<sup>‡</sup> CODATA recommends that the unit be J/Hz to stress that in  $h = E/\nu$  the frequency  $\nu$  is in cycles/sec (Hz), not radians/sec.

<sup>\*</sup> These are calculated from exact values and are exact to the number of places given (*i.e.* no rounding).

<sup>\*\*</sup> The molar mass of  $^{12}\text{C}$  is 11.999 999 9958(36) g.

<sup>†</sup> At  $Q^2 = 0$ . At  $Q^2 \approx m_W^2$  the value is  $\sim 1/128$ .

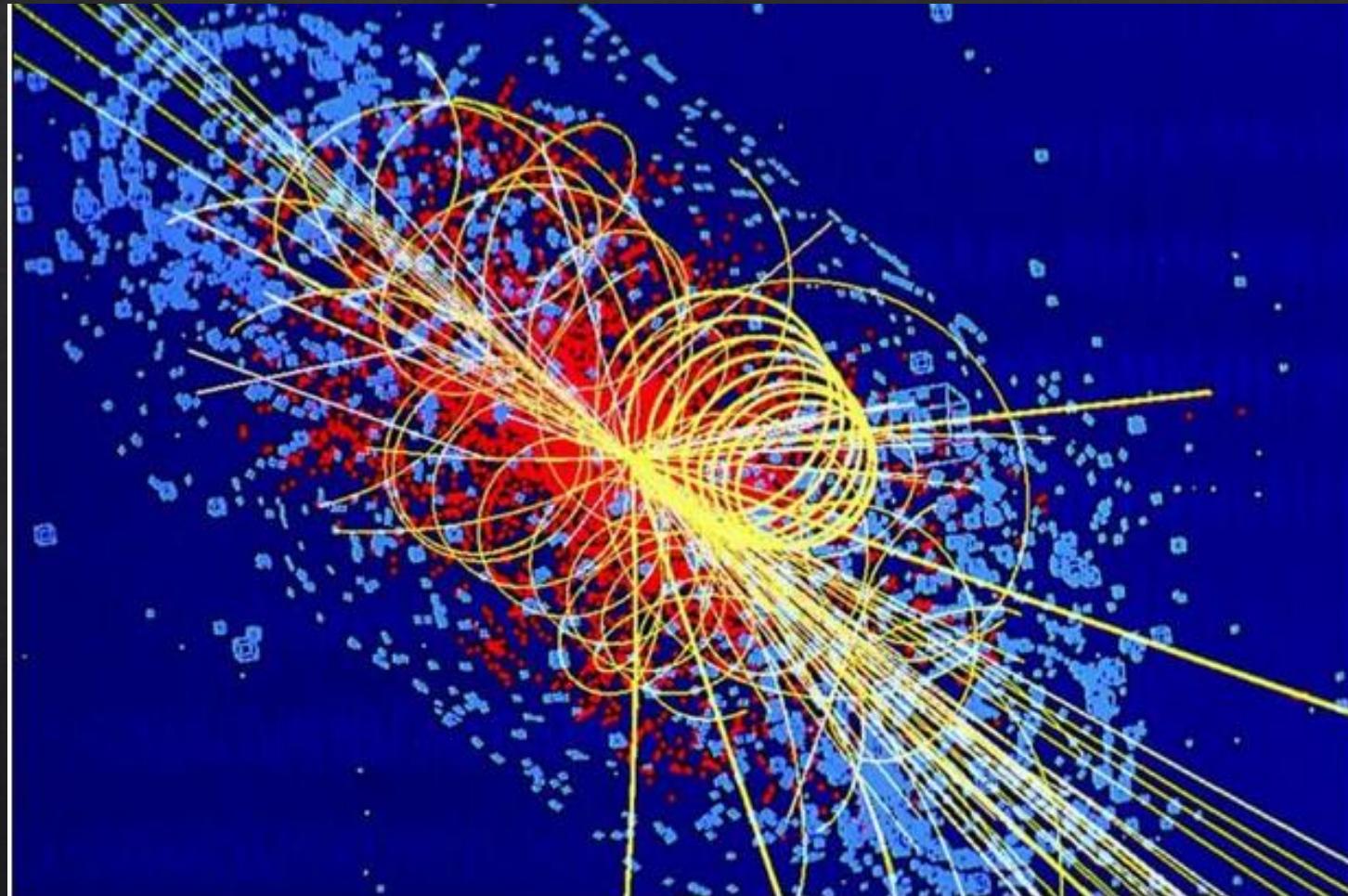
<sup>‡‡</sup> Absolute laboratory measurements of  $G_N$  have been made only on scales of about 1 cm to 1 m.

<sup>#</sup> See the discussion in Sec. 10, “Electroweak model and constraints on new physics.”

<sup>††</sup> The corresponding  $\sin^2 \theta$  for the effective angle is 0.23155(4).

## 2. De unde obținem informații despre particule?

- ❖ ciocniri
- ❖ dezintegrări
- ❖ stări legate



Simulation of a particle collision in which a Higgs boson is produced (Image: Lucas Taylor/CMS)

particle physics collisions: cross section  $\sigma_{fi}$  =  $\frac{W_{fi}}{(initial\ flux)}$  (*number of final states*)  
 ciocniri: secțiune eficace de împrăștiere

$$d\sigma = \frac{|M|^2}{F} dQ \quad \text{M - amplitudine de tranziție}$$

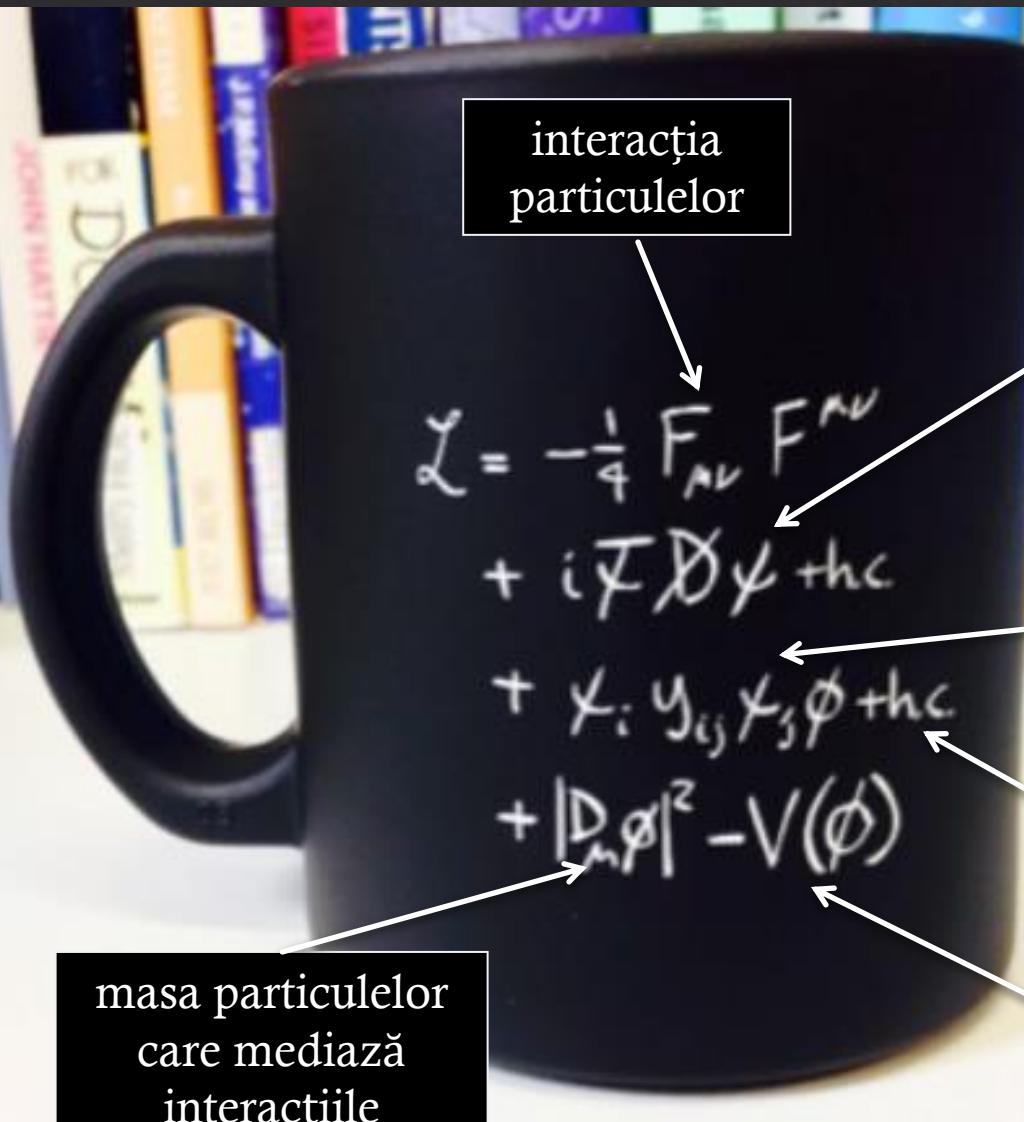
$$\begin{aligned} \sigma_{tot}(a_1 + a_2 \rightarrow b'_1 + \dots + b'_m) &= \frac{1}{2w(s, m_1^2, m_2^2)} \\ &\cdot \int \prod_{i=1}^m \left( \frac{dp'_i}{(2\pi)^3} \delta_+(p_i'^2 - m_i'^2) \right) (2\pi)^4 \delta(p'_1 + \dots + p'_m - p_1 - p_2) \\ &\cdot |\langle b_1(p'_1) \dots b_m(p'_m) | T | a_1(p_1) a_2(p_2) \rangle|^2. \end{aligned}$$

dezintegrări: rata de dezintegrare

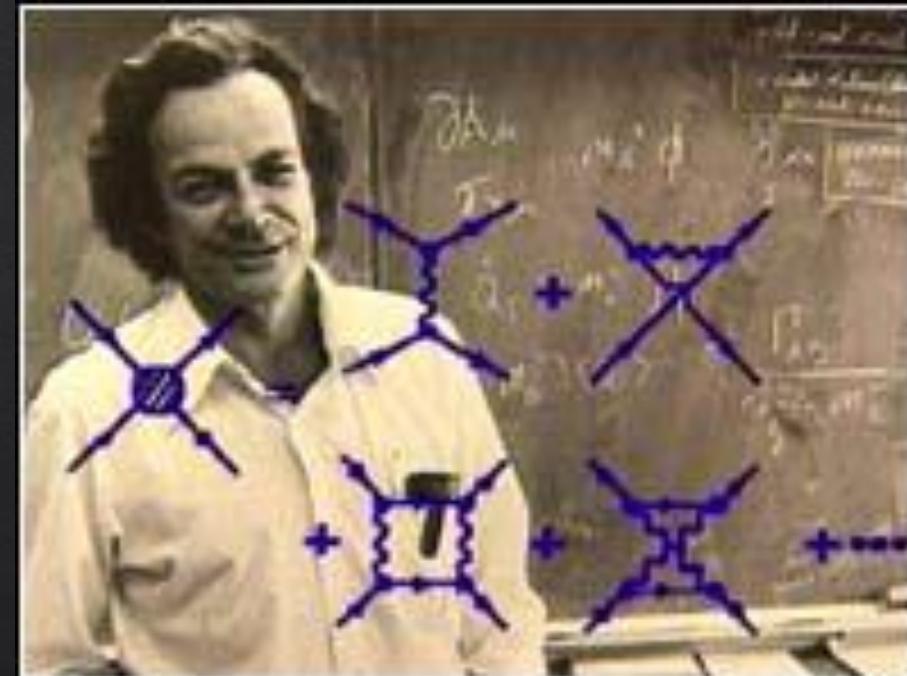
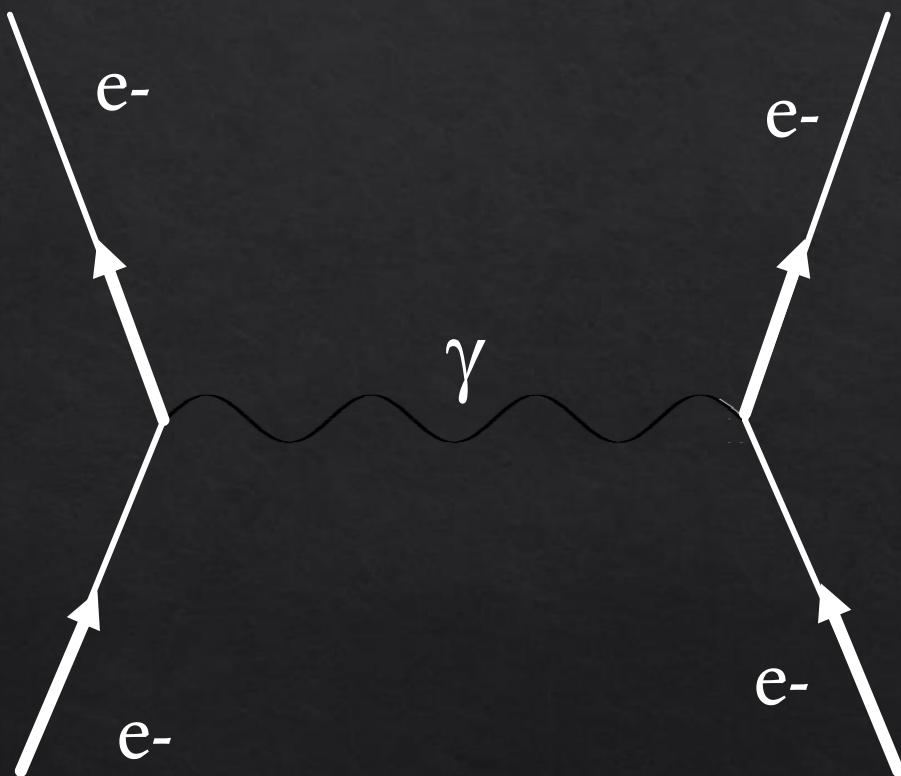
$$\begin{aligned} d\Gamma(a(p) \rightarrow b_1(p'_1) + \dots + b_m(p'_m)) &= \frac{1}{2m_a} (2\pi)^4 \delta(p'_1 + \dots + p'_m - p) \\ &\cdot \prod_{i=1}^m \frac{d^3 p'_i}{(2\pi)^3 2p_i'^0} |\langle b_1(p'_1) \dots b_m(p'_m) | T | a(p) \rangle|^2. \end{aligned}$$

# Cum interacționează?

$$\begin{aligned}
& -\frac{1}{2}\partial_\nu g_\mu^a \partial_\mu g_\nu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4}g_s^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e + \\
& \frac{1}{2}ig_s^2 (\bar{q}_i^\sigma \gamma^\mu q_j^\sigma) g_\mu^a + \bar{G}^a \partial^2 G^a + g_s f^{abc} \partial_\mu G^a G^b g_\mu^c - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- - \\
& M^2 W_\mu^+ W_\mu^- - \frac{1}{2}\partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2} M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2}\partial_\mu A_\mu \partial_\mu A_\nu - \frac{1}{2}\partial_\mu H \partial_\mu H - \\
& \frac{1}{2}m_h^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \frac{1}{2c_w^2} M \phi^0 \phi^0 - \beta_h [\frac{2M^2}{g^2} + \\
& \frac{2M}{g} H + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-)] + \frac{2M^4}{g^2} \alpha_h - ig c_w [\partial_\nu Z_\mu^0 (W_\mu^+ W_\nu^- - \\
& W_\nu^+ W_\mu^-) - Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\mu^0 (W_\nu^+ \partial_\nu W_\mu^- - \\
& W_\nu^- \partial_\nu W_\mu^+)] - ig s_w [\partial_\nu A_\mu (\bar{W}_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - \\
& W_\mu^- \partial_\nu W_\mu^+) + A_\mu (W_\nu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\nu W_\mu^+)] - \frac{1}{2}g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \\
& \frac{1}{2}g^2 W_\mu^- W_\mu^- W_\nu^+ W_\nu^- + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - Z_\mu^0 Z_\nu^0 W_\mu^+ W_\nu^-) + \\
& g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w [A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\
& W_\nu^+ W_\mu^-) - 2A_\mu Z_\mu^0 W_\nu^+ W_\nu^-] - g\alpha [H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-] - \\
& \frac{1}{8}g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2] - \\
& g M W_\mu^+ W_\mu^- H - \frac{1}{2}g \frac{M}{c_w^2} Z_\mu^0 Z_\mu^0 H - \frac{1}{2}ig [W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - \\
& W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)] + \frac{1}{2}g [W_\mu^+ (H \partial_\mu \phi^- - \phi^- \partial_\mu H) - W_\mu^- (H \partial_\mu \phi^+ - \\
& \phi^+ \partial_\mu H)] + \frac{1}{2}g \frac{1}{c_w} (Z_\mu^0 (H \partial_\mu \phi^0 - \phi^0 \partial_\mu H) - ig \frac{s_w^2}{c_w} M Z_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \\
& ig s_w M A_\mu (W_\mu^+ \phi^- - W_\mu^- \phi^+) - ig \frac{1-2c_w^2}{2c_w} Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + \\
& ig s_w A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \frac{1}{4}g^2 W_\mu^+ W_\mu^- [H^2 + (\phi^0)^2 + 2\phi^+ \phi^-] - \\
& \frac{1}{4}g^2 \frac{1}{c_w} Z_\mu^0 Z_\mu^0 [H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2}g^2 \frac{s_w^2}{c_w} Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + \\
& W_\mu^- \phi^+) - \frac{1}{2}ig^2 \frac{s_w^2}{c_w} Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2}g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\
& W_\mu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{s_w}{c_w} (2c_w^2 - 1) Z_\mu^0 A_\mu \phi^+ \phi^- - \\
& g^1 s_w^2 A_\mu A_\mu \phi^+ \phi^- - \bar{e}^\lambda (\gamma \partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda \gamma \partial \nu^\lambda - \bar{u}_j^\lambda (\gamma \partial + m_u^\lambda) u_j^\lambda - \bar{d}_j^\lambda (\gamma \partial + \\
& m_d^\lambda) d_j^\lambda + ig s_w A_\mu [-(\bar{e}^\lambda \gamma e^\lambda) + \frac{2}{3}(\bar{u}_j^\lambda \gamma u_j^\lambda) - \frac{1}{3}(\bar{d}_j^\lambda \gamma d_j^\lambda)] + \frac{ig}{4c_w} Z_\mu^0 [(\bar{\nu}^\lambda \gamma^\mu (1 + \\
& \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - 1 - \gamma^5) u_j^\lambda) + \\
& (\bar{d}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 - \gamma^5) d_j^\lambda)] + \frac{ig}{2\sqrt{2}} W_\mu^+ [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (1 + \\
& \gamma^5) C_{\lambda\kappa} d_j^\kappa)] + \frac{ig}{2\sqrt{2}} W_\mu^- [(\bar{e}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger \gamma^\mu (1 + \gamma^5) u_j^\lambda)] + \\
& \frac{ig}{2\sqrt{2}} \frac{m_\lambda^2}{M} [-\phi^+ (\bar{e}^\lambda (1 - \gamma^5) e^\lambda) + \phi^- (\bar{e}^\lambda (1 + \gamma^5) \nu^\lambda)] - \frac{g}{2} \frac{m_\lambda^2}{M} [H (\bar{e}^\lambda e^\lambda) + \\
& i\phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda)] + \frac{ig}{2M\sqrt{2}} \phi^+ [-m_d^\kappa (\bar{u}_j^\lambda C_{\lambda\kappa} (1 - \gamma^5) d_j^\kappa) + m_u^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 + \\
& \gamma^5) d_j^\kappa)] + \frac{ig}{2M\sqrt{2}} \phi^- [m_d^\lambda (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 + \gamma^5) u_j^\kappa) - m_u^\kappa (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 - \gamma^5) u_j^\kappa)] - \\
& \frac{g}{2} \frac{m_\lambda^2}{M} H (\bar{u}_j^\lambda u_j^\lambda) - \frac{g}{2} \frac{m_\lambda^2}{M} H (\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2} \frac{m_\lambda^2}{M} \phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \frac{ig}{2} \frac{m_\lambda^2}{M} \phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \\
& \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \frac{M^2}{c_w^2}) X^0 + \bar{Y} \partial^2 Y + \\
& ig c_w W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \partial_\mu \bar{X}^- X^0) + ig s_w W_\mu^+ (\partial_\mu \bar{Y} X^- - \partial_\mu \bar{X}^- Y) + \\
& ig c_w W_\mu^- (\partial_\mu \bar{X}^- X^0 - \partial_\mu \bar{X}^0 X^+) + ig s_w W_\mu^- (\partial_\mu \bar{X}^- Y - \partial_\mu \bar{Y} X^+) + \\
& ig c_w Z_\mu^0 (\partial_\mu \bar{X}^+ X^- - \partial_\mu \bar{X}^- X^-) + ig s_w A_\mu (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) - \\
& \frac{1}{2}g M [\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w^2} \bar{X}^0 X^0 H] + \frac{1-2c_w^2}{2c_w} ig M [\bar{X}^+ X^0 \phi^+ - \\
& \bar{X}^- X^0 \phi^-] + \frac{1}{2c_w} ig M [\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + ig M s_w [\bar{X}^0 X^- \phi^+ - \\
& \bar{X}^0 X^+ \phi^-] + \frac{1}{2}ig M [\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0]
\end{aligned}$$



# Feynman – diagrame și reguli



Feynman earned his Nobel for creating  
these diagrams  
(Courtesy Auckland University)

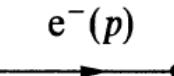
# Feynman – diagrame și reguli

## Electrodinamică cuantică

electron in initial state

 $u(p)$ 

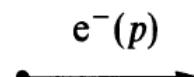
incoming electron line



electron in final state

 $\bar{u}(p)$ 

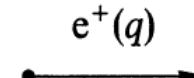
outgoing electron line



positron in initial state

 $\bar{v}(q)$ 

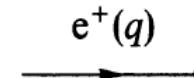
outgoing electron line



positron in final state

 $v(q)$ 

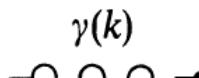
incoming electron line



photon in initial state

 $\epsilon^\mu$ 

incoming photon line



photon in final state

 $\epsilon^{\mu*}$ 

virtual photon

$$\frac{-ig_{\mu\nu}}{k^2 + i\epsilon}$$

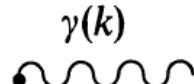
virtual electron

$$i \frac{\not{p} + m}{p^2 - m^2 + i\epsilon}$$

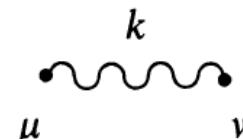
elementary process

$$ie\gamma^\mu$$

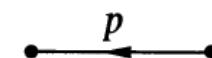
outgoing photon line



internal photon line



internal electron line



vertex



# Feynman – diagrame și reguli

$W^-$  initial state

$\varepsilon(k)$

$W^-$  in final state

$\varepsilon^*(k)$

$W^+$  in initial state

$\varepsilon(k)$

$W^+$  in final state

$\varepsilon^*(k)$

incoming  $W$  line

$W^-(k)$

outgoing  $W$  line

$W^-(k)$

outgoing  $W$  line

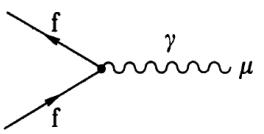
$W^+(k)$

incoming  $W$  line

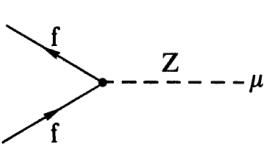
$W^+(k)$

Fermion–Boson vertices:

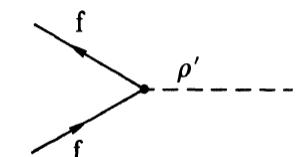
$$-ieQ_f \gamma^\mu$$



$$-i \frac{e}{\sin \theta_W \cos \theta_W} \left\{ T_3^f \gamma^\mu \frac{1 - \gamma_5}{2} - \sin^2 \theta_W Q_f \gamma^\mu \right\}$$



$$-i \frac{m_f}{\rho_0}$$

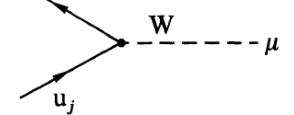
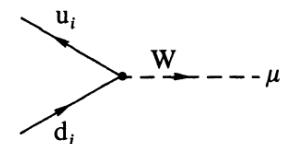
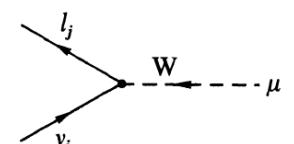
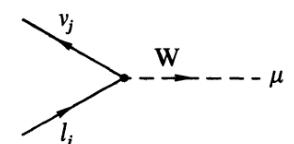


$$-i \frac{e}{\sqrt{2} \sin \theta_W} \gamma^\mu \frac{1 - \gamma_5}{2}$$

$$-i \frac{e}{\sqrt{2} \sin \theta_W} \gamma^\mu \frac{1 - \gamma_5}{2}$$

$$-i \frac{e}{\sqrt{2} \sin \theta_W} V_{ij} \gamma^\mu \frac{1 - \gamma_5}{2}$$

$$-i \frac{e}{\sqrt{2} \sin \theta_W} V_{ji}^* \gamma^\mu \frac{1 - \gamma_5}{2}$$



$Z$  in initial (final) state

$\varepsilon(k)(\varepsilon^*(k))$

Higgs particle in initial (final) state

1

virtual  $W$  boson

$$\frac{i \left( -g^{\mu\nu} + \frac{k^\mu k^\nu}{m_W^2} \right)}{k^2 - m_W^2 + i\epsilon}$$

virtual  $Z$  boson

$$\frac{i \left( -g^{\mu\nu} + \frac{k^\mu k^\nu}{m_Z^2} \right)}{k^2 - m_Z^2 + i\epsilon}$$

virtual Higgs particle

$$\frac{i}{k^2 - m_{\rho'}^2 + i\epsilon}$$

external  $Z$  line

$Z(k)$

external  $\rho'$  line

$\rho'(k)$

internal  $W$  line

$W(k)$

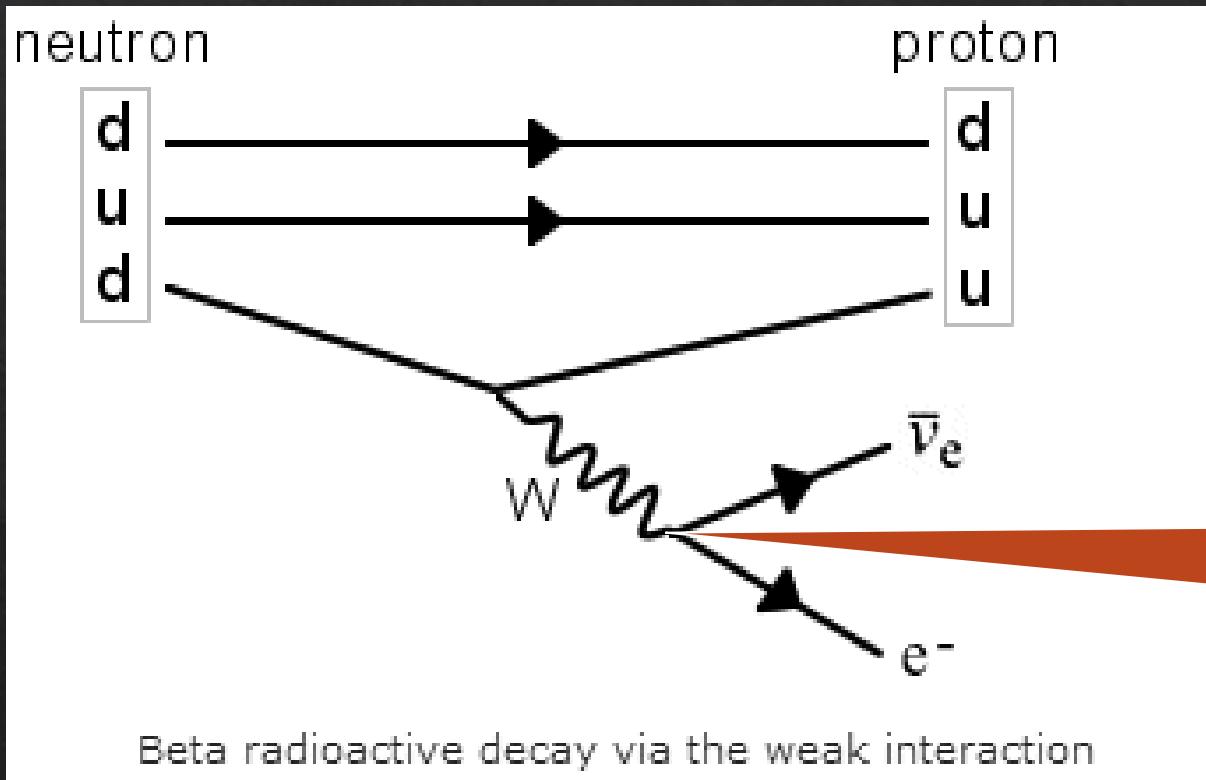
internal  $Z$  line

$Z(k)$

internal  $\rho'$  line

$\rho'(k)$

# Feynman – diagrame și reguli



Sarcina electrică  
se conservă la  
fiecare vertex

# Feynman – diagrame și reguli

$$-i \frac{e}{\sqrt{2} \sin \theta_W} \gamma^\mu \frac{1 - \gamma_5}{2}$$

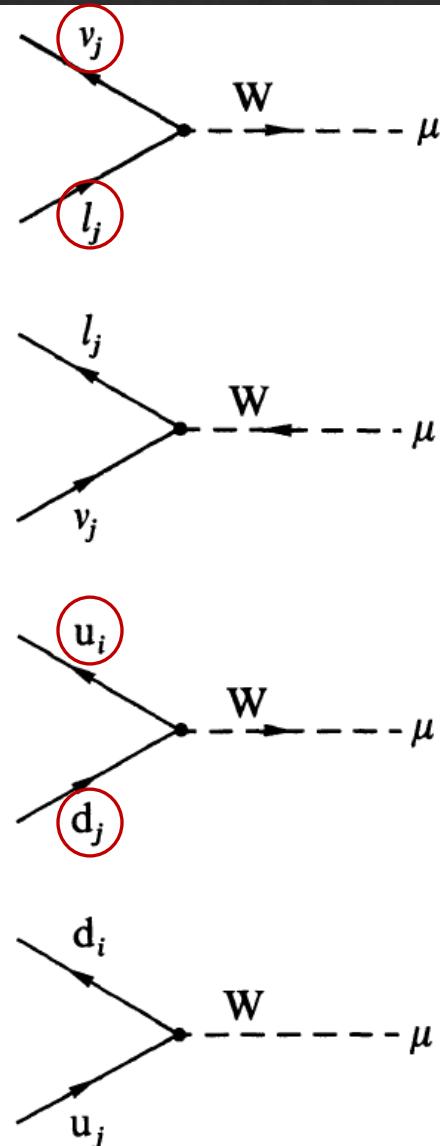
$$-i \frac{e}{\sqrt{2} \sin \theta_W} \gamma^\mu \frac{1 - \gamma_5}{2}$$

$$-i \frac{e}{\sqrt{2} \sin \theta_W} V_{ij} \gamma^\mu \frac{1 - \gamma_5}{2}$$

matricea de mixing a cuarcilor  
(Cabibbo–Kobayashi–Maskawa)

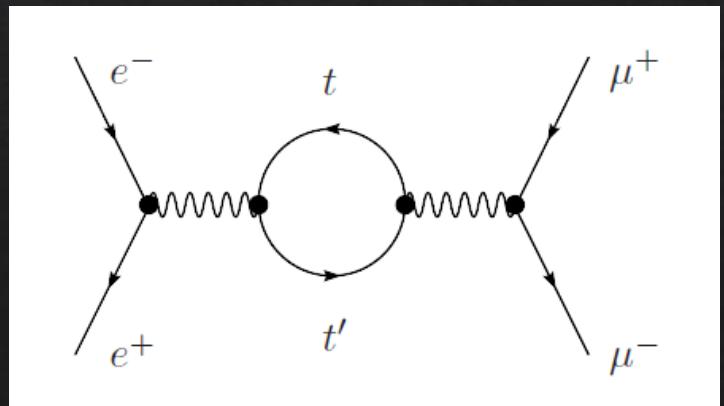
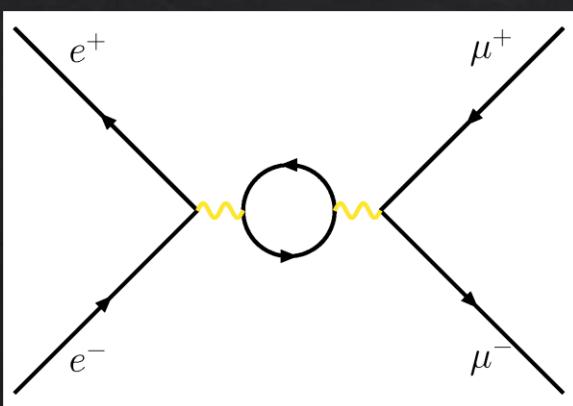
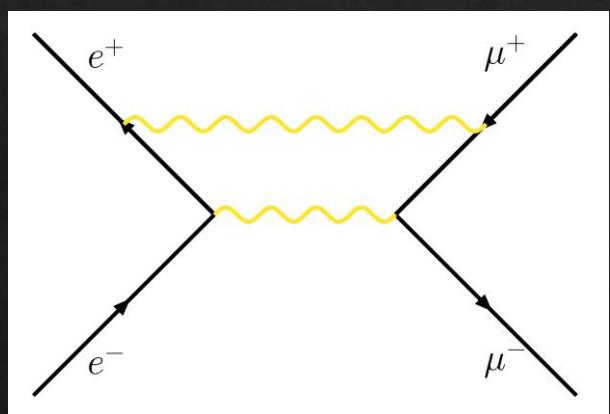
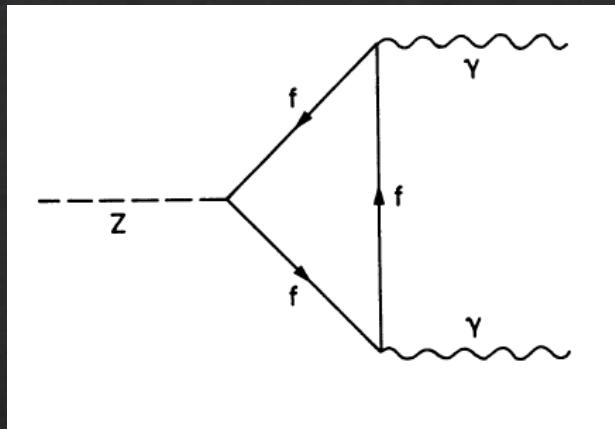
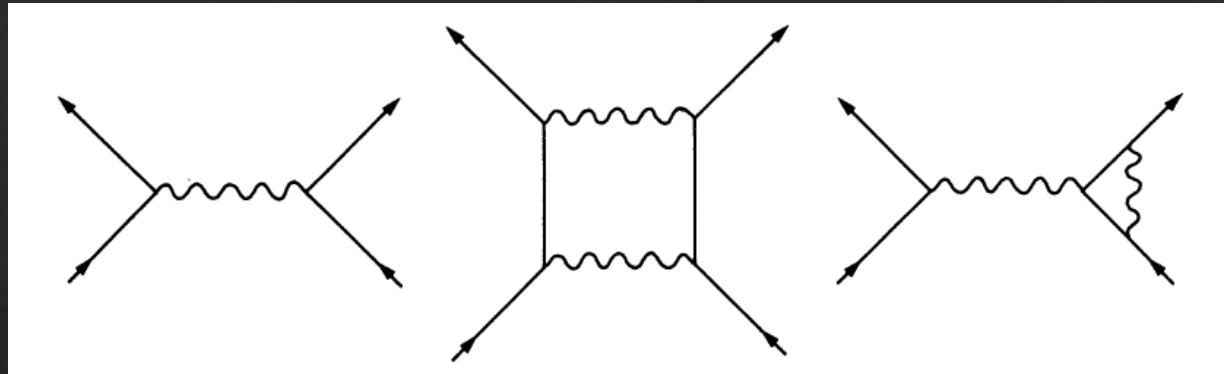
$$-i \frac{e}{\sqrt{2} \sin \theta_W} V_{ji}^* \gamma^\mu \frac{1 - \gamma_5}{2}$$

$$\begin{array}{lll} v_1 \equiv v_e, & v_2 \equiv v_\mu, & v_3 \equiv v_\tau; \\ \ell_1 \equiv e, & \ell_2 \equiv \mu, & \ell_3 \equiv \tau; \\ u_1 \equiv u, & u_2 \equiv c, & u_3 \equiv t; \\ d_1 \equiv d, & d_2 \equiv s, & d_3 \equiv b. \end{array}$$



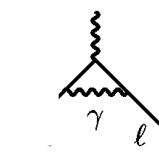
...și mai multe!

## Teoria perturbațiilor – corecții

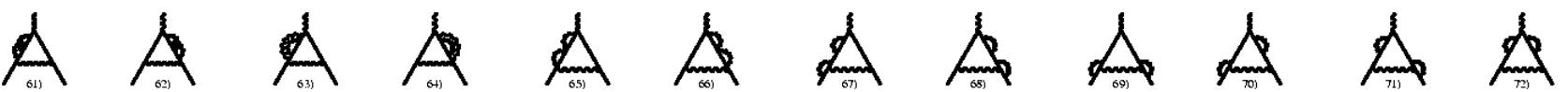
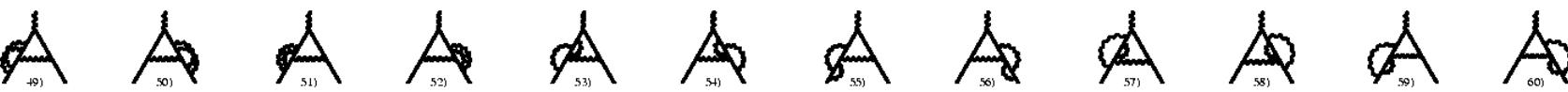
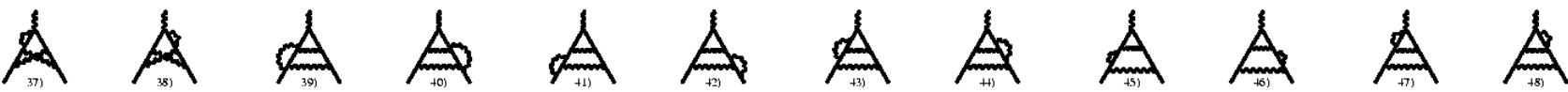
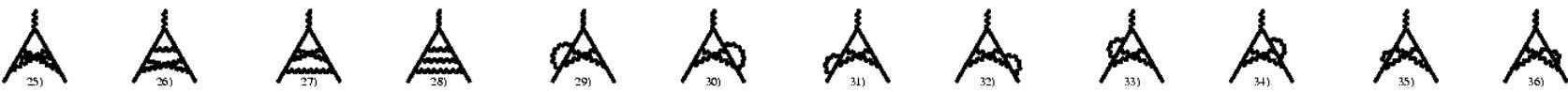


$(g-2)_\ell$

$\ell = e, \mu, \tau$



$$a_e = a_\mu = a_\tau = \frac{\alpha}{2\pi}$$

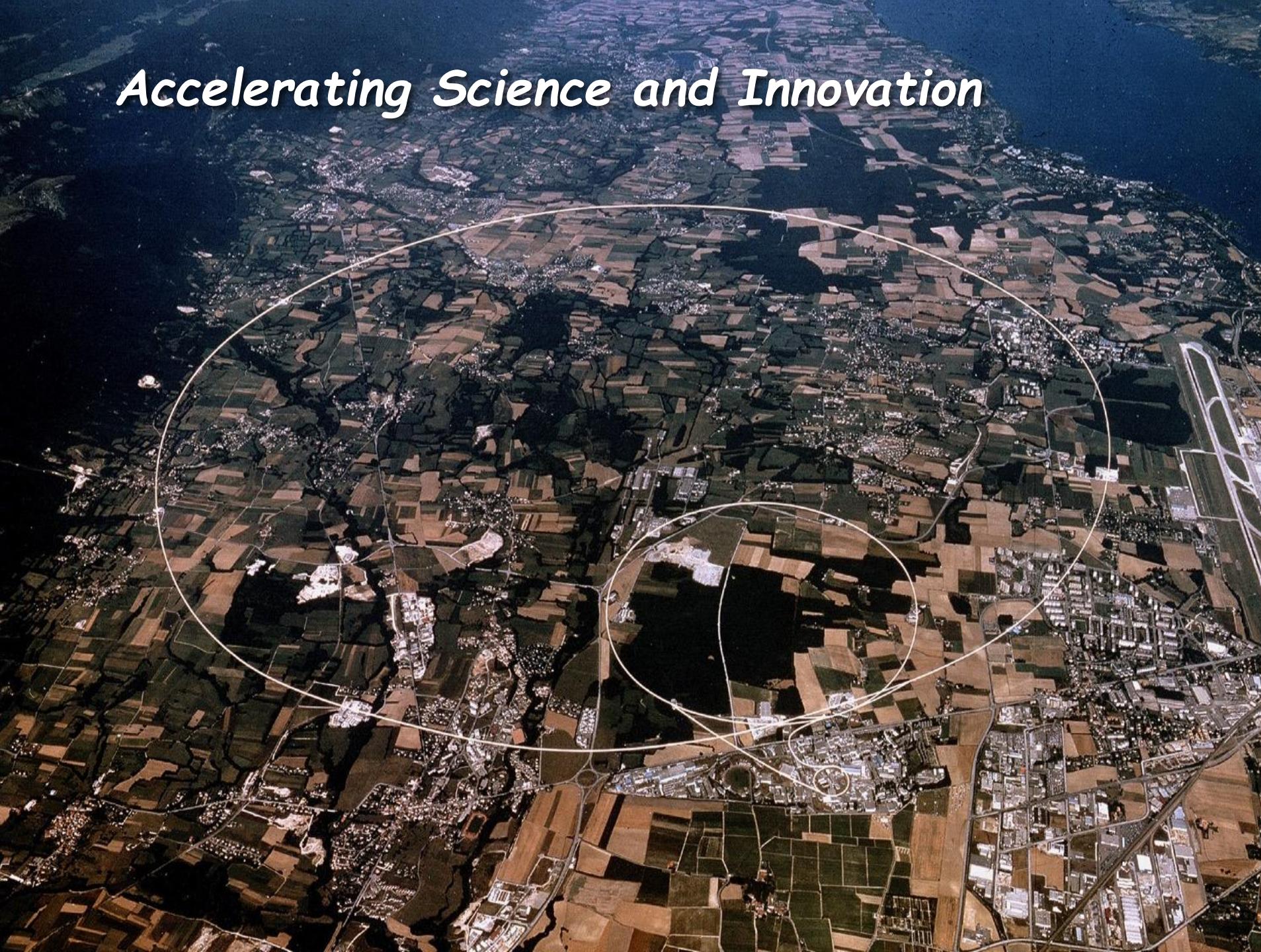


### 3. Cum producem particule elementare?

- ❖ metode simple pentru electroni, protoni (e.g. ionizări)
- ❖ radiații cosmice
- ❖ reactori nucleari
- ❖ acceleratori de particule



*Accelerating Science and Innovation*



CERN (*Conseil Européen pour la Recherche Nucléaire* 1954)  
is often referred to as the European Laboratory for Particle Physics.



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# Romania

Romania entered into direct collaboration with CERN in the early 1990s. Even before becoming a CERN member, Romania made significant contributions to the [ALICE](#), [ATLAS](#) and [LHCb](#) experiments.

Romania became CERN's 22nd Member State on 17 July 2016. The Institute of Atomic Physics is the funding agency covering the Romanian participation in the CERN experiments. In the current national plan, which started in 2016, the Romanian institutions contribute to the following experiments: ALICE, ATLAS, LHCb, [WLCG](#), [ISOLDE](#), [NA62](#), [n\\_TOF](#), [MoEDAL](#) and [WA105](#). The participations are evaluated yearly by an International Scientific Advisory Board. There are four national R&D institutes and six universities from six cities involved in the CERN collaborations, with IFIN-HH being the largest stakeholder. The number of scientists and engineers involved is over 100 and it has been increasing steadily in the last decade.

*This page was last updated on 12 May, 2020*

CERN contact(s): [P. Wells](#), [O. Capatina](#)

116 CERN users - [Overview of participation](#)

[Industrial Liaison](#) | [Knowledge Transfer](#) | [Scientific Computing](#)

[Teacher Student Forum](#) | [Communication](#) | [Outreach](#)

## Experiments

[ALICE](#), [ATLAS](#), [LHCb](#), [DIRAC](#), [ISOLDE](#), [NA62](#), [CLICdp](#)

## WLCG participation

Tier 2

## Member States

[Austria](#)

[Belgium](#)

[Bulgaria](#)

[Czech Republic](#)

[Denmark](#)

[Finland](#)

[France](#)

[Germany](#)

[Greece](#)

[Hungary](#)

[Israel](#)

[Italy](#)

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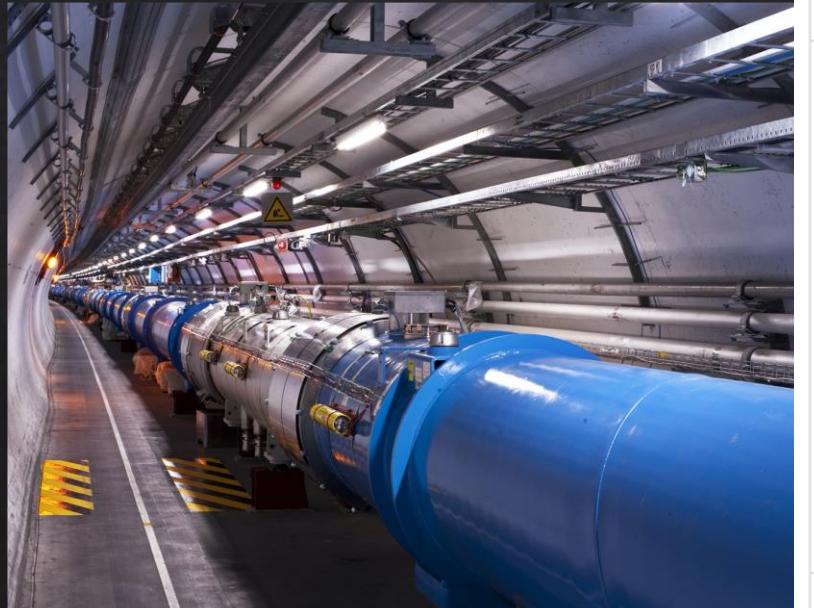
[Slovak Republic](#)

[Spain](#)

[Sweden](#)

[Switzerland](#)

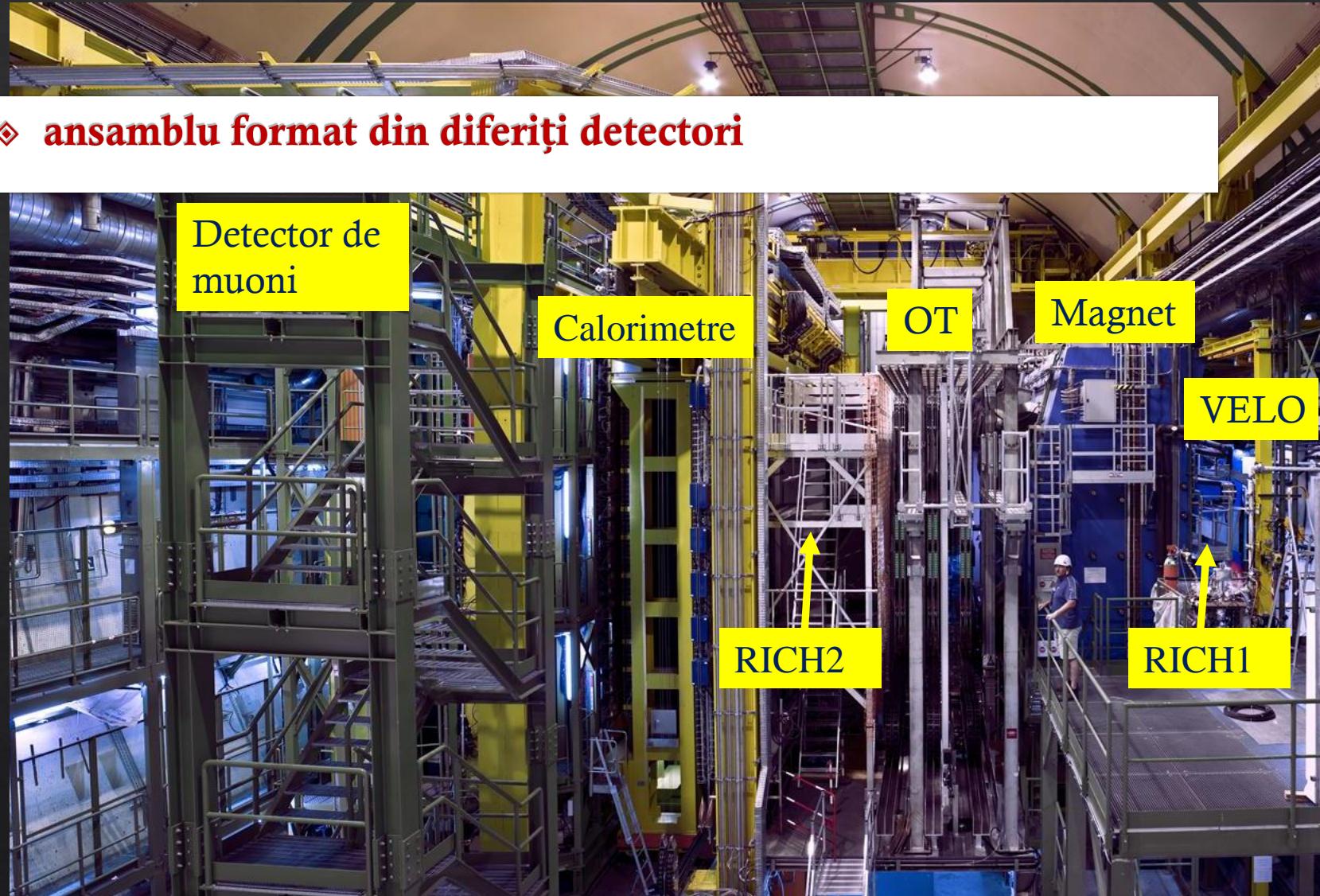
[United Kingdom](#)

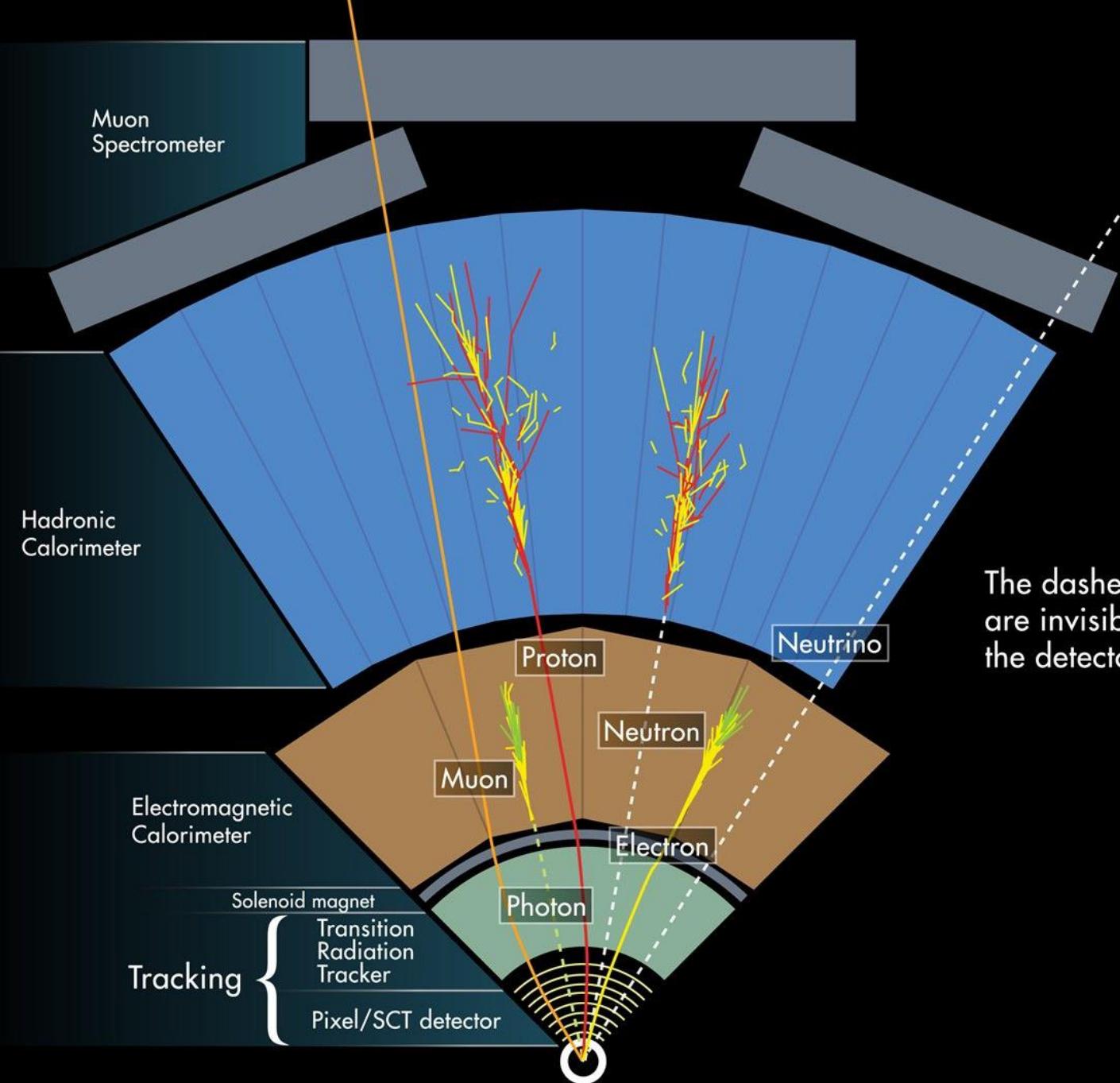


## Proiecte in derulare

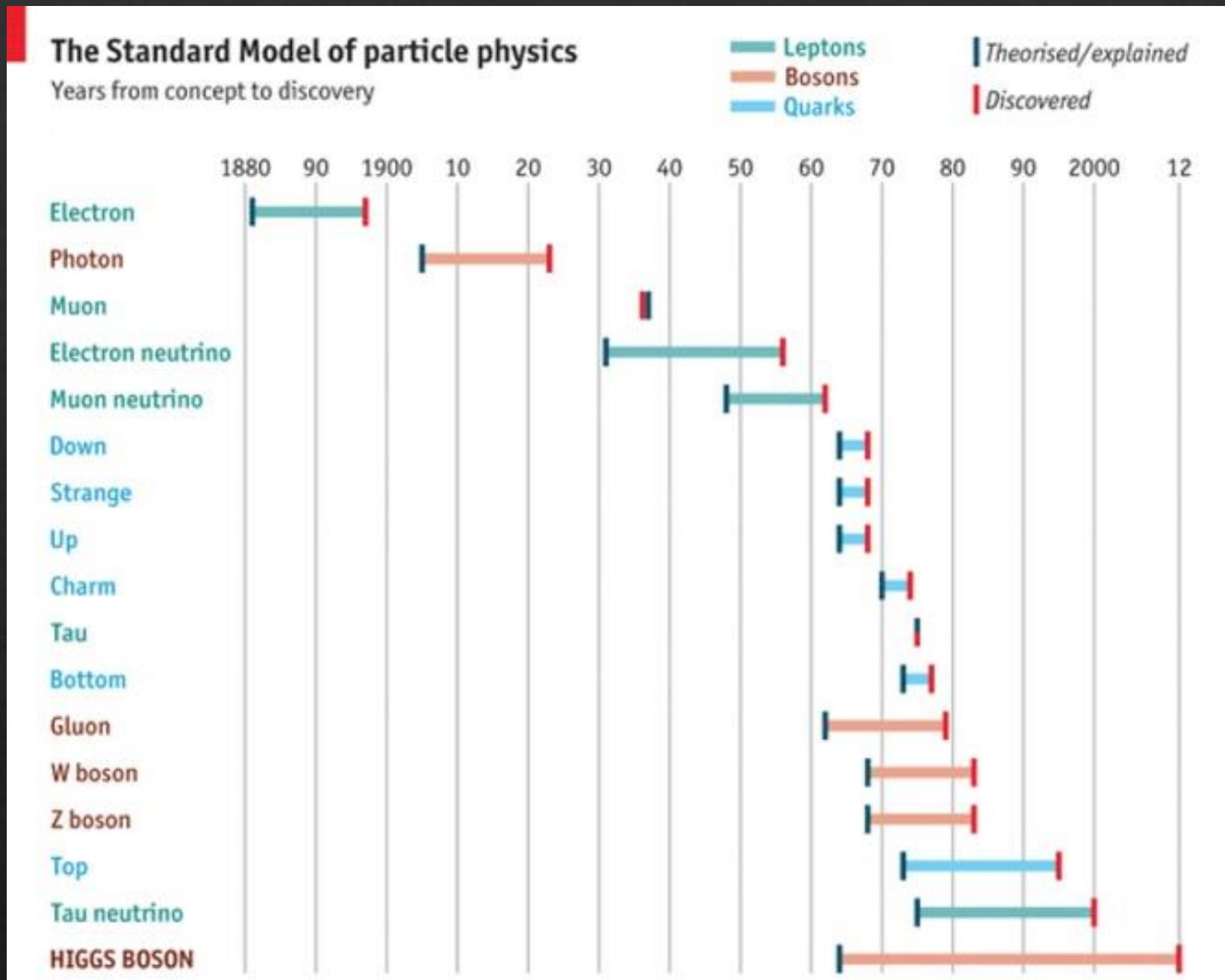
Experiment	Conducator proiect(CO) / Parteneri	Director proiect	WLCG	CO: IFIN-HH	Mihnea DULEA
ALICE	CO: IFIN-HH	Mihai PETROVICI <a href="mailto:mpetro@ifin.nipne.ro">mpetro@ifin.nipne.ro</a>	GOOLDE	CO: IFIN-HH	Constantin MIHAI <a href="mailto:constantin.mihai@nipne.ro">constantin.mihai@nipne.ro</a>
	CO: ISS	Alexandru Florin DOBRIN <a href="mailto:alexandru.florin.dobrin@cern.ch">alexandru.florin.dobrin@cern.ch</a>	TOF	CO: IFIN-HH	Alexandru NEGRET <a href="mailto:alnegret@tandem.nipne.ro">alnegret@tandem.nipne.ro</a>
and RD51	CO: IFIN-HH P1: ITIM-CJ P2: UPB P3: UAIC P4: UVT P5: UTB P6: UB	Calin ALEXA <a href="mailto:calin.alexa@cern.ch">calin.alexa@cern.ch</a>	NA62	CO: IFIN-HH	Alexandru-Mario BRAGADIREANU <a href="mailto:mario.bragadireanu@nipne.ro">mario.bragadireanu@nipne.ro</a>
LHCb	CO: IFIN-HH	Florin MACIUC <a href="mailto:florin.maciuc@cern.ch">florin.maciuc@cern.ch</a>	RD50	CO: INCDFM	Ioana PINTILIE <a href="mailto:ioana@infim.ro">ioana@infim.ro</a>
	P1: USV		MoEDAL	CO: ISS	Vlad POPA <a href="mailto:vpopa@spacescience.ro">vpopa@spacescience.ro</a>
			WA105	CO:UB-FF	Ionel LAZANU <a href="mailto:ionel.lazanu@g.unibuc.ro">ionel.lazanu@g.unibuc.ro</a>
			DsTau <sup>NA65</sup>	CO:ISS	Elena FIRU <a href="mailto:elena.firu@spacescience.ro">elena.firu@spacescience.ro</a>

# 4. Cum detectăm particule elementare?





The dashed tracks  
are invisible to  
the detector



# Fizica particulelor

1. Care sunt particulele elementare (ce proprietăți au – masă, sarcină electrică, spin, ...)?
  2. Cum interacționează? - De unde obținem informații?
  3. Cum producem particule elementare?
  4. Cum detectăm particule elementare?
- **Particle Physics**
  - **Particle Detectors**

- ? Din ce este constituită materia?
- ? Din ce este compus universul?
- ? Care este originea universului și cum a evoluat?
- ? De ce se comportă aşa universul?
- ? Cum va evoluă?
- ? ....

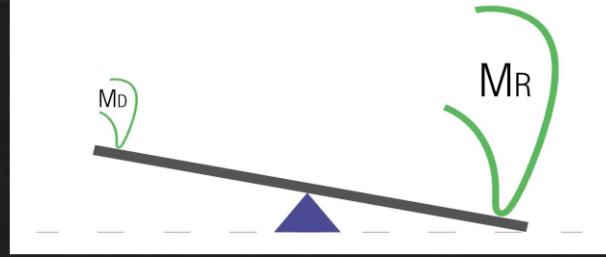
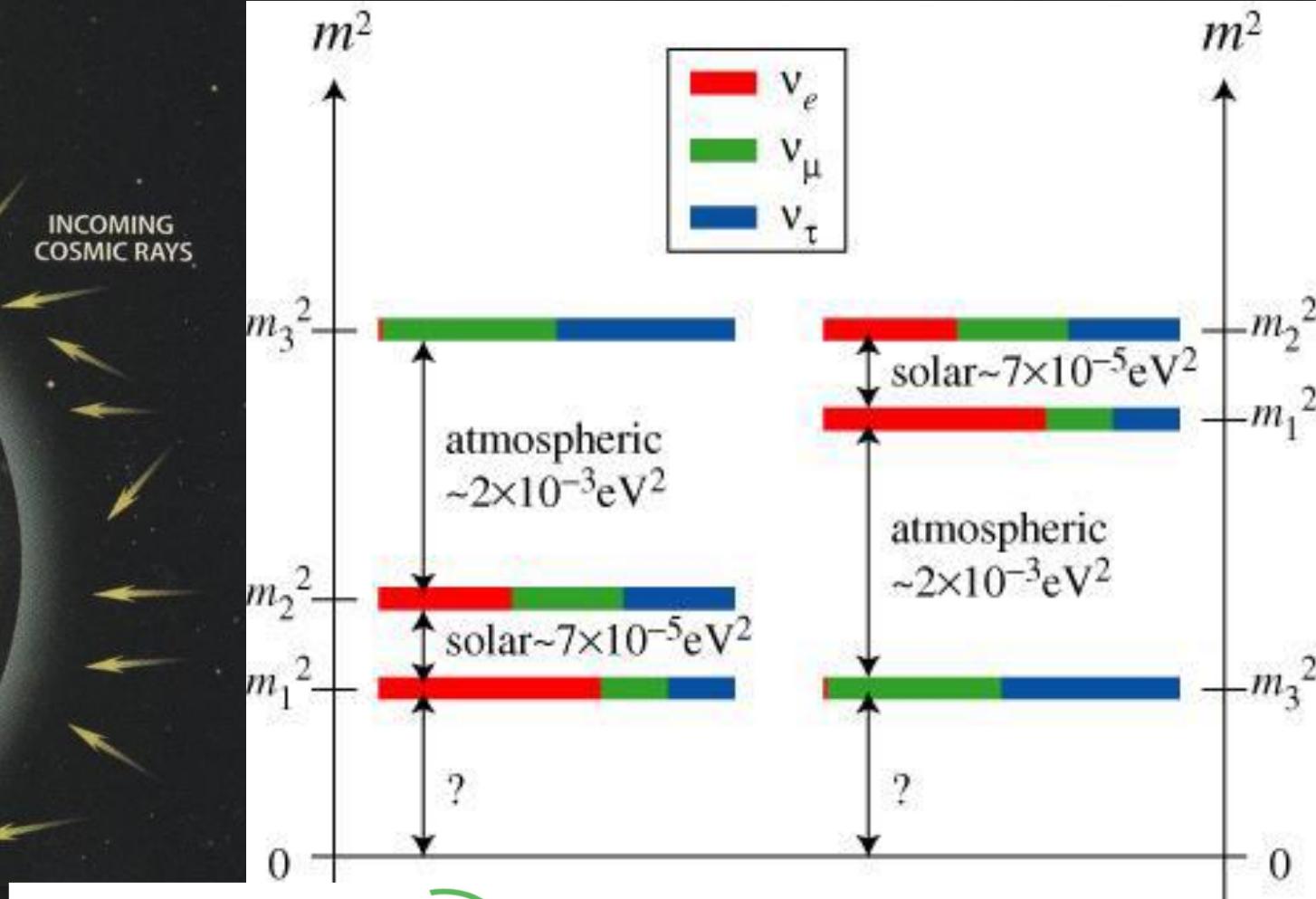
Modelul Standard –  
răspunsul la toate întrebările?

❖ **NU!**

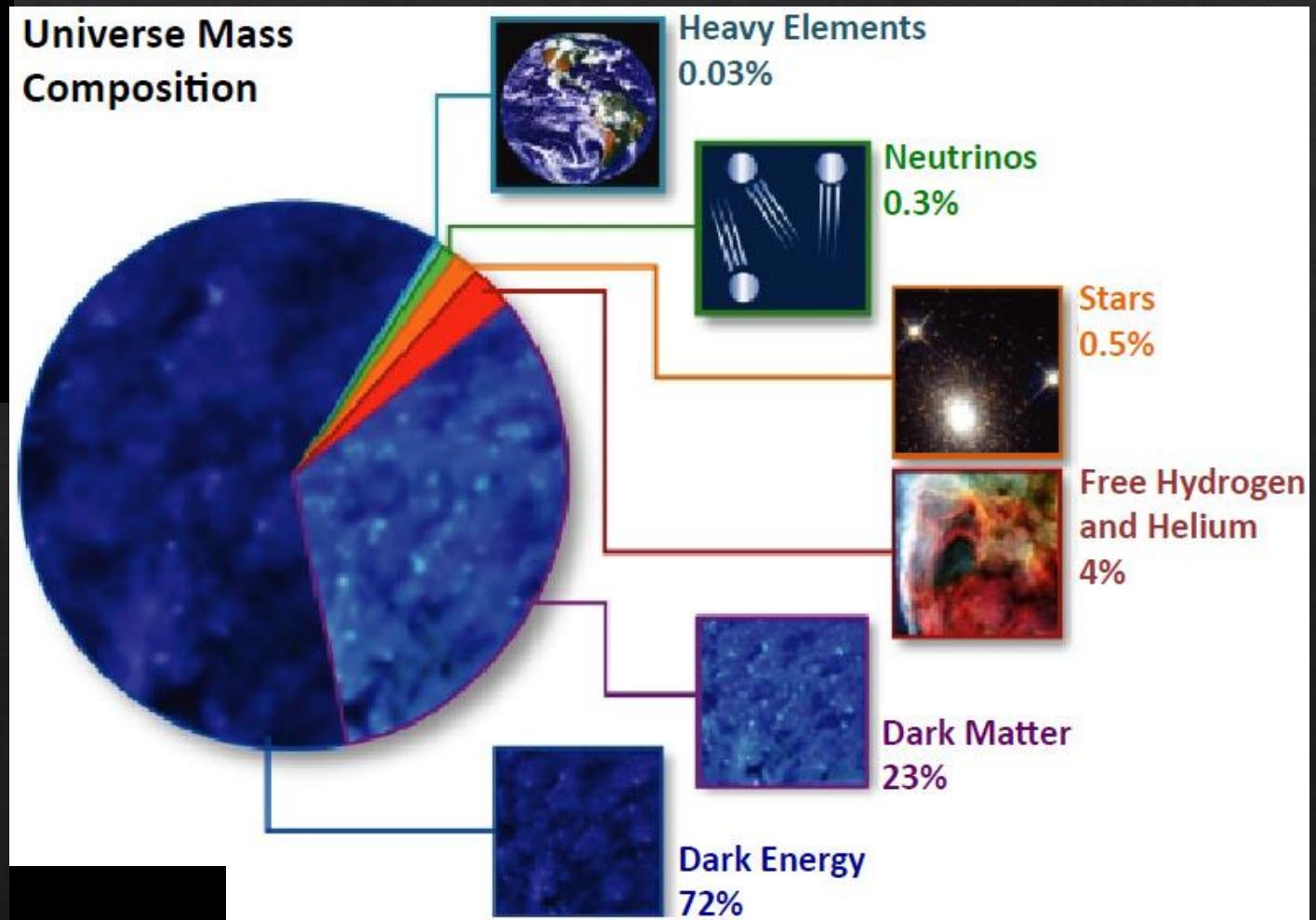
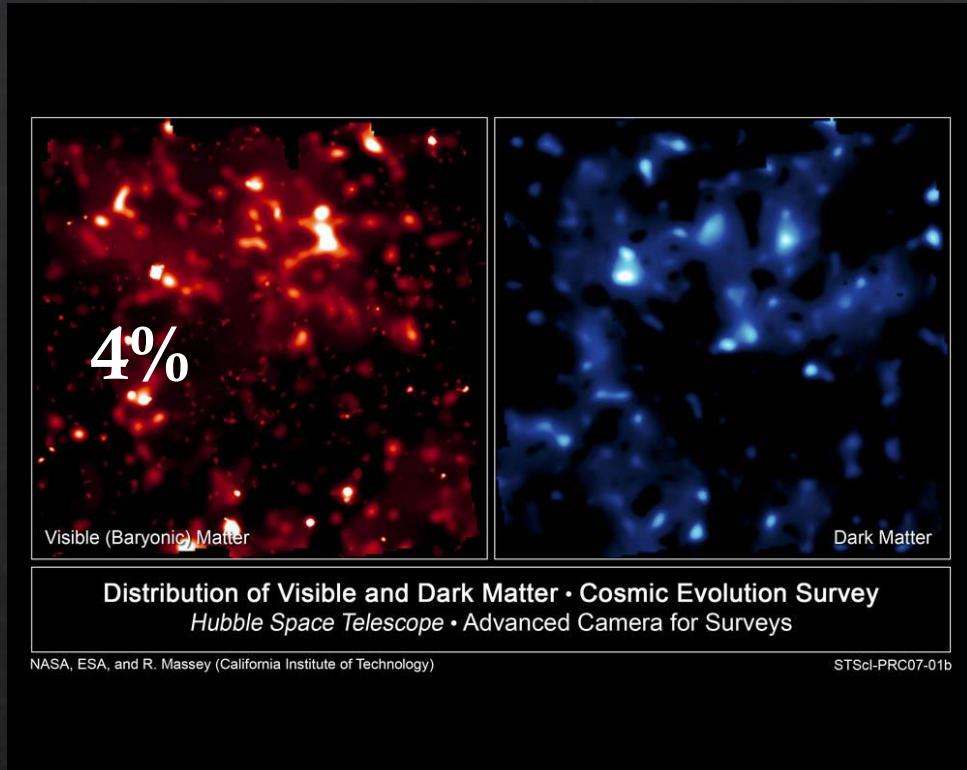
# Cum acomodăm în teorie masa neutrinilor?



matricea de mixing a neutrinilor?



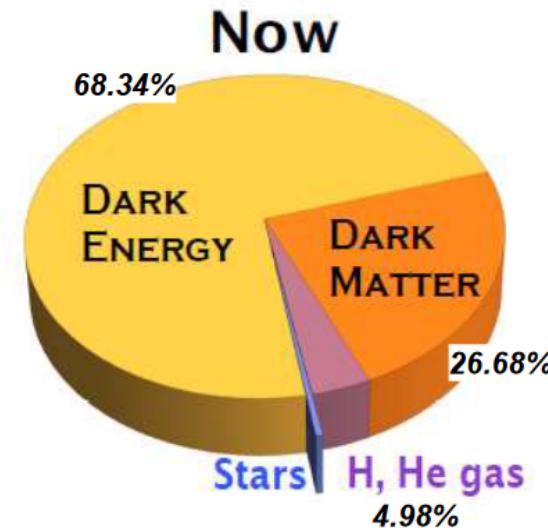
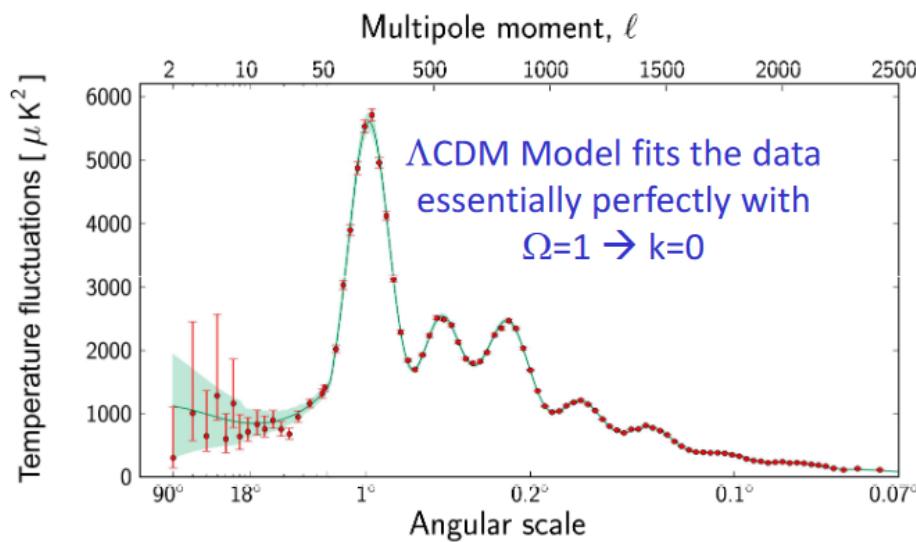
# Ce este „dark matter”?



# Urmează: Pascal Pralavorio's on Dark matter

## Cosmic microwave background

- Use this photon flux to know the composition of the Universe
  - Decompose data in spherical harmonics
  - Amplitude and position of “acoustic” peaks gives the composition of the Universe



Cosmology is making precision measurements ... and  
we don't know 95% of the Universe and 85% of the matter !

# De ce în univers există mai multă materie decât antimaterie?

- ◊ Există cantități mari de materie, dar nu și dovada unor cantități mari de antimaterie.

violarea conservării sarcinii și parității  
CP – charge-parity

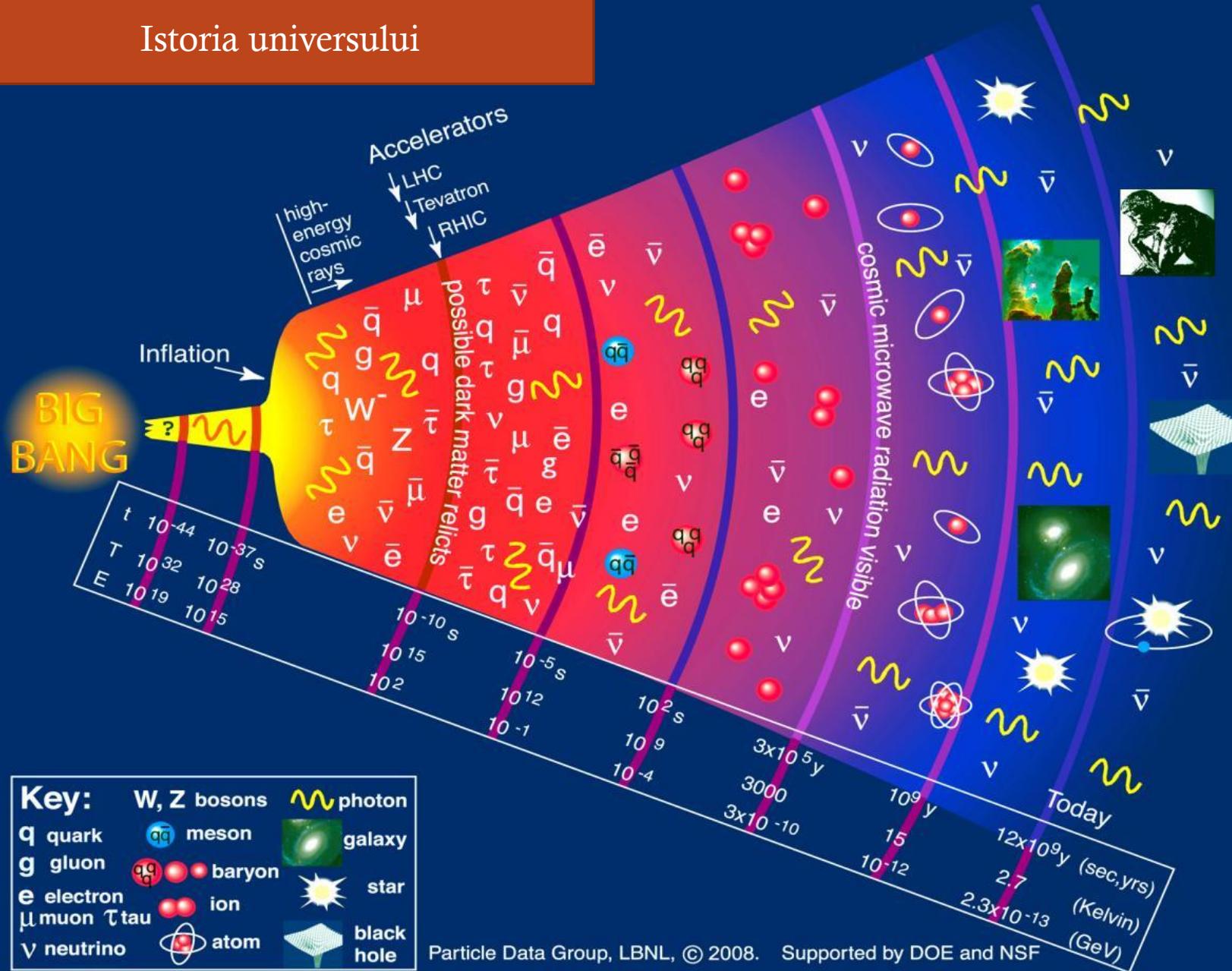


beauty  
 $B^0$

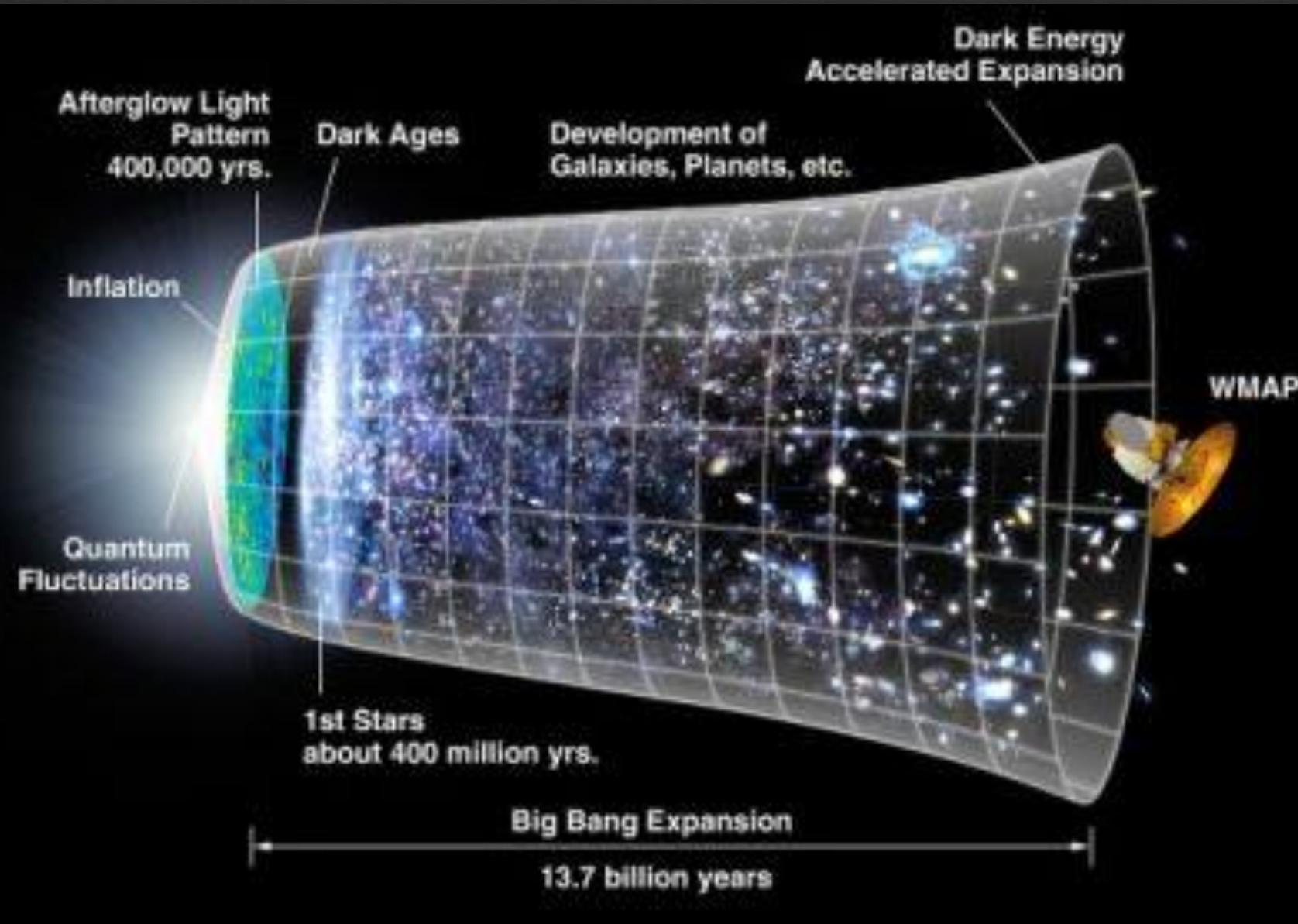


anti-beauty  
 $\bar{B}^0$

## Istoria universului



# Ce face universul astăzi și cum va evoluă?



NASA/WMAP Science Team

# Modelul Standard – răspunsul la toate întrebările?

◆ **NU!**

- Extensions of the Standard Model of elementary particles

# Particle Listings

R.L. Workman *et al.* (Particle Data Group), Prog. Theor. Exp. Phys. **2022**, 083C01 (2022) and 2023 update

Cut-off date for Listings/Summary Tables was Jan. 15, 2023. Files can be downloaded directly by clicking on the icon: [PDF](#). For a key to the listings [click here](#).

Expand/Collapse All

## Gauge & Higgs Bosons

gamma

g (gluon)

graviton

W boson

Z boson

H

Neutral Higgs Bosons, Searches for

Charged Higgs Bosons ( $H^\pm$  and  $H^{\pm\pm}$ ), Searches for

Heavy Bosons, Other Than Higgs Bosons, Searches for

Axions ( $A^0$ ) and Other Very Light Bosons, Searches for

## Leptons (e, mu, tau, neutrinos, heavy leptons ...)

electron

[PDF](#) [pdgLive](#)

muon

[PDF](#) [pdgLive](#)

tau

[PDF](#) [pdgLive](#)

### Heavy Charged Lepton Searches

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### Neutrino Properties

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### Number of Neutrino Types

[PDF](#) [pdgLive](#)

### Double-beta Decay

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### Neutrino Mixing

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### Heavy Neutral Leptons, Searches for

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## Quarks (u,d,s,c,b,t...)

### LIGHT QUARKS --- u, d, s

[PDF](#) [pdgLive](#)

c quark

[PDF](#) [pdgLive](#)

b quark

[PDF](#) [pdgLive](#)

t quark

[PDF](#) [pdgLive](#)

b' quark (4\*\*\*th Generation)

[PDF](#) [pdgLive](#)

t' quark (4\*\*\*th Generation)

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Free Quark Searches

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# Fizica particulelor

1. Care sunt particulele elementare (ce proprietăți au – masă, sarcină electrică, spin, ...)?
2. Cum interacționează? - De unde obținem informații?
3. Cum producem particule elementare?
4. Cum detectăm particule elementare?

- Particle Physics
- Particle Detectors