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PMT Analysis for <u>3D reconstruction</u> & <u>Negative Ion Drift</u>

David J. G. Marques* on behalf of the <u>CYGNO</u> collaboration:



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CYGNO in The DM paradigm



In the WIMP model, DM forms a halo within our galaxy

Solar system rotates around galaxy towards Cygnus constellation

Earth susceptible to an apparent WIMP wind from Cygnus direction! CYGNO wants to observe it through direct detection of nuclear recoils







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In the WIMP model, DM forms a halo within our galaxy +

Solar system rotates around galaxy towards Cygnus constellation

- Earth susceptible to an apparent WIMP wind from Cygnus direction! CYGNO wants to observe it through direct detection of nuclear recoils
- ENERGY ⇒ Excess would result in <u>falling</u> <u>exponentials</u>.
- TIME ⇒ Results
 in a <u>few %</u>
 <u>annual</u>
 <u>modulation.</u>



WIMP wind June v=220 km/s Cygnus galactic plane December ★ DIRECTION ⇒ Results in a characteristic anisotropy in the angular distribution of nuclear recoils ⇒ No background can mimic!

Directional discrimination is a striking feature to positively identify Dark Matter!



CEvNS produces NRs identical to the DM-induced ones, but, below 10 GeV/c², we have mostly Solar neutrinos ⇒ with directionality*, these never superimpose with Cygnus, allowing us to venture into the neutrino fog.

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CYGNO - What's the setup?





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LIME - Detector and Data





- → 50 L & 50 cm drift gaseous TPC
- → He:CF4, 60:40, 1 Atm (910 mbar), 293 K
- → 4 PMTs + 1 sCMOS camera

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 - 1 CMOS pic: R_{AF} = 500 ms
 - **X PMT WFs** = N_{triggers} * N_{PMTs} * N_{digitiz}
 - R_{\Deltat} = **1.3 ns & 4 ns**



1000

1500

2000

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- → <u>The information needs to be matched!</u>





3D Event Reconstruction



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- CZGNO G S
- → To fully reconstruct the information of one event we need to merge the CMOS and PMTs information.
 - We developed a <u>1-to-1 association</u> to merge the *CMOS clusters* to *PMT triggers*.
 - 1. Light seen by each PMT depends

on their relative positions.



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 - By integrating the charge in a time window, it's possible to retrieve the
 (x, y, L) information by performing a multi-variable Bayesian fit.



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4. The final goal is to implement this technique at the detector front-end level. The <u>efficiencies</u> are promising! (The reference is the 33x33cm² GEM plane)



→ **Optimization undergoing** concerning effects like

saturation, lens barreling, gain inter-calibration, etc.



3D Event Reconstruction



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3D Events - CMOS analysis



- The analysis of the CMOS images starts with a <u>directional iDBSCAN algorithm</u> which clusters groups of pixels belonging to the same ionization event.
 - a. For <u>PID</u>, each <u>cluster</u> can be selected through its: *light integral, length, slimness, photon density, dE/dx, etc.*



Directional iDBSCAN to detect cosmic-ray tracks for the CYGNO experiment - IOPscience

 A second layer analysis is used to determine other more dedicated variables such as <u>2D direction</u>



Reconstructed info here:

- ΔΧΥ
- X-Y angle (ϕ)
- 2D direction



3D Event Reconstruction



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3D Events - PMT analysis

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CZGNO G S Experiment S I

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3D Events - PMT analysis

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- **1.** Time-over-Threshold ⇒ Traveled Z
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Additional reconstructed info:

- ΔΖ
- Sign of θ



3D Events - PMT analysis

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3D Event Reconstruction

Dual sensor 3D analysis

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3D Events - 3D analysis

- → With the sensors' individual and merged information, we can now fully analyze the events and perform <u>particle</u>
 ID, reject backgrounds from known sources, and fully <u>characterize the 3D direction</u> ⇒ <u>Directionality</u>
 - The first studies were focused on <u>Alpha Particles</u>





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- Alphas at ~4 cm length (²²²Rn) see a preferential emission towards the cathode instead, although their origin is still being investigated.



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- Alphas at ~4 cm length (²²²Rn) see a preferential emission towards the cathode instead, although their origin is still being investigated.
- → Remaining alphas at lower (< 4 cm) length (see <u>*α*-lengths</u>) are thought to come from the U/Th chains, and emitted uniformly in angle and Z.

→ The transverse diffusion allows us to have some idea about the distribution of absolute Z of these alphas, but this methodology could be improved...





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Negative lons - Concept







Absolute Z from Δt between minority charge carriers







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and σ_{da,SF_x} - (10⁻²⁰

Table 2

Collisional processes that originate the SF₆ ion species of interest for this work. Adapted from Ref. [46].

Reaction	Process	Energy
$e^- + SF_6 \rightarrow SF_6^{-*}$	Electron attachment	<1 meV
$SF_6^{-*} \rightarrow SF_6 + e^-$	Autodetachment	(Metastable: > 1 µs)
$SF_6^{-*} + SF_6 \rightarrow SF_6^- + SF_6$	Collisional stabilization	
$e^- + SF_6 \rightarrow SF_5^- + F$	Dissociative electron attachment	0-2 eV
e^- + $SF_6 \rightarrow SF_4^-$ + $2F$		3-8 eV
$e^-~+~SF_6~\rightarrow~F^-~+~SF_5$		1-14 eV
$e^-~+~SF_6~\rightarrow~F_2^-~+~SF_4$		1–14 eV
$SF_{5/6}^- + SF_6 \rightarrow SF_{5/6} + SF_6 + e^-$	Collisional detachment	>90 eV
$SF_6^- + SF_6 \rightarrow SF_6 + SF_6^-$	Charge transfer	
$SF_6^- + SF_6 \rightarrow SF_5^- + F + SF_6$	Dissociative charge transfer	>1 eV

https://doi.org/10.1016/j.nima.2022.1661416

5 10 Electron Energy (eV)

Negative Ions - Analysis

< CMOS >

ED: He:CF₄ 60:40 NID: He:CF₄:SF₆ 59:39.6:1.6



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Negative Ions - Analysis

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< PMT >

$\star \star$ First ever PMT analysis for optical NID signals $\star \star$

- 1. Threshold ➡ Only individual peaks above 6*RMS
- 2. **Rebin** ➡ Selected peaks put into histogram
- 3. Delimitation → Start (end) when 2 bins are above (below) 10 mV



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Negative lons - Ion mobility

Final measurement: Ion Mobility

- 1. Tilted <u>ED</u> alpha tracks
 - a. Distribution ➡ Get mean value (△t)
 - b. Knowing electrons' velocity in gas (v)
 - i. <u>Z travelled by track ⇒ 1.5 cm</u>





- 2. Tilted <u>NID</u> alpha tracks
 - a. Average time window (<u>At</u>)
 - i. $Z / \Delta t = v_{ion}$



- Data *self-consistent* with at 650 and 900 mbar
- Charge carriers' *mobility* consistent with *SF6-* in literature



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Achievements:

- <u>CMOS</u> shows a transverse diffusion <u>~3 times</u> smaller!
- **<u>PMT</u>** undoubtedly <u>proves NID regime</u> with **optical readout!**

<u>Next steps:</u>

- Longitudinal diffusion
 - Primary ionization cluster counting
- Fiducialization with SFX-
- CMOS + PMT NID 3D analysis

Conclusions



- → The <u>CYGNO</u> collaboration is developing a high-precision <u>gaseous TPC</u> at atmospheric pressure with <u>optical readout</u>.
- → The main focus is the <u>directional direct search</u> of <u>DM WIMP-like particles</u> in the <u>low mass range</u> (0.5-10 GeV).
- → Through <u>nuclear recoil direction</u>, backgrounds can be rejected and <u>unambiguous confirmation of DM</u> is possible.
- → The merging of CMOS (X-Y) with PMT (Z) information is performed with a Bayesian fit.
 - The ability of <u>reconstructing</u> ionization events in <u>3 dimensions</u> greatly improves our <u>spatial resolution</u> and our <u>PID capabilities</u>.
 - From the suspicion of the presence of Rn, the <u>3D analysis</u> allows us to almost confirm its presence and, through its <u>emission direction</u>, its origin became more clear.
- → The addition of SF6 in the mixture would put the detector in the <u>Negative Ion Drift (NID)</u> regime.
 - Preliminary studies prove its plausibility at <u>atmospheric pressure</u>.
 - Upcoming studies will show how the <u>diffusion</u> can be improved with NID and how <u>minority carriers</u> could become a technique for detector <u>Z-fiducialization</u>.

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The CYGNO Experiment - Instruments

Thank you for

your attention!

The CYGNO Project counts with the collaboration of several international researchers, coming from:

Some cool pictures























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+ × ×



Backup

& more

details

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CYGNO-30 - Prospects



- → Low mass (0.5 10 GeV) directional DM searches
- → > 2027
- → 30 100 m3 detector
- → 0.5 1 keV_{ee} energy threshold
- → 30° angular resolution





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The CYGNUS project





Within the CYGNUS collaboration, several approaches are being studied.

The Italian group, <u>CYGNO</u>, is developing a **gaseous TPC** based on the setup:

GEMs + sCMOS + PMT to test Optical Readout

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CYGNO - Roadmap



Several ongoing efforts in different fronts:

- 3D reconstruction
- Directionality
- ER vs. NR (+ML)
- Shielding optimization
- Background data vs. MC
- DM Sensitivity
- Design and Commissioning of CYGNO_04
- Enhancement of the light yield
- Negative lon drift



CYGNO-Dark Matter paradigm





Low Density @ atm pressure

- Allows tracks of up to
 - millimetres at few keV without

compromising exposure.

< 10 GeV/c²

- → To observe lower WIMP masses:
 - Low thresholds are necessary, since lower mx originate lower energy recoils.
 - **Light nuclei** used to maximize energy transfer.

<u>Helium (He)</u>

→ Light target for SI in low mass range.



Fluorine (F)

- Heavier target to intermediate WIMP masses.
- One of the highest sensitivity to SD coupling.

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