

Commissioning of the CGEM Inner Tracker

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1) Project overview

2) Commissioning results

3) CGEM-IT installation



BEPCII @ IHEP (Beijing)



BEPCII @ BEPCII



Upgrade program



Replace the aging innermost layers with a new inner tracker: the CGEM-IT

Two steps accelerator upgrade: 1) 3x Increase luminosity at 4.7 GeV (now) 2) Extend the center of mass energy up to 5.6 GeV (in 2028)



The CGEM-IT: detector

The CGEM-IT info:

- three layers of cylindrical triple-GEM
- improve spatial resolution along the beam axis (< 300 µm), rate capability, and radiation hardness
- keep momentum and azimuthal spatial resolution (~150 μm and 0.5% at 1 GeV/c)
- 0.5% X_0 per layer material budget:
 - cathode and anode are the supporting structure, made by a sandwich of Rohacell-Kapton (L2), Honeycomb-carbon fiber (L1 and L3)
 - supporting permaglass rings at the two ends outside the active area





The CGEM-IT: layer 3 construction



Ferrara

The electrodes on the mandrels

Beijing (北京)

Assembly of the electrods with VIM



Gas and HV tests



Peek grids to prevent buckling



JINST 16 (2021) 08, P08065

The CGEM-IT: readout chain



On detector

Off detector

The CGEM-IT: TIGER and GUFI

TIGER (Torino Integrated GEM Electronics for Readout)





Photo Courtesy of M. Mignone (INFN TO)



64 channel mixed signal triggerless ASIC

A dedicated GUI (GUFI) has been prepared for standalone commissioning It controls the GEMROCs and the DAQ, plus other status function (e.g., noise extraction, threshold scan, ...)



Applied threshold on both T and E branch.

Change of paradigma in the noise estimation. Two steps procedure: 1) Inject a fixed number of analog test pulse of known amplitude 2) Fit the full shape to extract S-curve (from TP rise) and baseline



Commissioning with cosmic rays

November 2023 - March 2024

Results from 4M cosmics rays





https://inspirehep.net/files/6a82983db656b53db44e5293f8118f2d Journal of Physics: Conference Series 1498 (2020) 012036 MPGD 2019 proceedings



We are operating with Ar:Iso (91:9) gas mixture.

During the commissioning, the three detectors were connected in series

We have tested parallel gas flow for each layer and we saw improvements in the general.



We plan to have at least 4 vol/h gas flow in each layer 12

Measured noise on layers



Measured noise on layers



HV monitor

During the operations, modulation of current with humidity observed.



Time and charge



GEMROCs sends out data collected from TIGER in a time window of 1.6 us around the trigger. Expected signal length of about 140 ns

Strip and charge













Cluster Charge and size



Residual distribution



Clear dependence of the charge centroid (CC) residual with the incident angle

For results below 5°, for each layer residual ~ 200 um μ TPC algorithm under development using these results



Extracted from signals with 20σ of residual distribution (few mm)



CGEM-IT Installation



MDC was successfully removed in September by IHEP colleagues









CGEM-IT was uncabled and moved to the experimental hall on October 2...







...Lifted and adjusted on the insertion pole







Slowly pushed inside





It reached the nominal position on October 5!!



And now we are cabling!

We are working to power it on for the first time in BESIII

Next:

- Commissioning with cosmic ray until end of the year
- Commissioning with beam starting from January



Summary and outlook

- The CGEM-IT deploys 3 layers of cylindrical triple-GEMs to replace the inner tracker of BESIII
 - Italy-China Hybrid construction of L3 with peek grids
 - A dedicated readout chain was developed and tested
- The standalone commissioning of the CGEM-IT started in November
 - Collected more than 4M cosmic rays
 - Tested different grounding scheme
- Efficiency flat about 95% for all the layers on both views
 - Cluster charge and size matches the expectations, L2 slightly lower
- Resolution with charge centroid about 200 μm
 - μTPC under development
- Passed internal reviews, received the green light for the installation, now we are on floor!



科學院為能物理研究所



Thanks for your attention



Artwork @2009 Phil Allora

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Backup slides



Cosmics stand





December to February

March to May

(a) Nucl. Instrum. Meth. Phys.Res. Sec. A 924 (2019) 181; (b) JINST 14 (2019) 08, P08013

TIGER

Table 2

ENCLE

2800

2600

2400

2200

2000

1600

 χ^2 / ndf

slope

constant

Measured performance of the TIGER ASIC.

Parameters	Values
Input charge	5-55 fC
TDC resolution	30 ps RMS
Time-walk (5-55 fC range)	12 ns
Average gain	10.75 mV/fC
Nonlinearity (5-55 fC range)	0.5%
RMS gain dispersion	3.5%
Noise floor (ENC)	1500 e ⁻
Noise slope	$10 e^{-}/\text{pF}$
Maximum power consumption	12 mW/ch

64 channels mixed-mode signal ASIC Digital back-end from TOFPET-2 110 nm CMOS technology and operated at 1.2 V power supply ~60 kHz per channel Jitter < 2 ns





CGEMBOSS

Version of the BESIII offline code that is developed expecially for CGEM-IT to provide a smooth transition at the start of operation.



All the results of the commissioning obtained with this software



90

21.12

7.985

Incident angle (°)



Efficiency – some details



Micro sector effect:

GEM HV is segmented. This creates a small gap where there is less amplification.

For straight tracks the effect is larger

Efficiency – some details



Supporting grids effect:

Supporting grids are aligned in z direction (every 10 cm), but not in ϕ direction. This creates a small gap where there is less amplification.

For straight tracks the effect is larger

Cluster reconstruction



Contiguous fired strips on the anode form a **cluster**

50

Charge centroid & µTPC concepts



 $Q_{\mathrm{hit},i} \; x_{\mathrm{hit},i}$

38

- The CC averages the charge of all the strips of the cluster by weighting it by its charge
- The μ TPC instead considers the drift gap as a tiny TPC and with position and time information, it associates each strip with a bi-dimensional point and uses a linear fit to extrapolate the particle position

Federico Matias Melendi - ICHEP 2024 Operation and readout of CGEM Inner Tracker

Charge centroid & µTPC CGEM-IT state

Preliminary result, also need to take into account:

- time calibrations to be optimized (contribution of 200-250µm)
- the contribution of tracking must be subtracted (contribution of 100-200 μm)



PRELIMINARY PLOT

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HV stability from our experience

- L1-L2 on a stand (not suitable for cosmic performance studies) in 2020
 - Continuous run for more than 700 hours
 - Continuous data taking with restart scheme (run 17 of digitization studies)
- Full CGEM-IT
 - Daily power on (~10 hours per day) about total 500 hours
 - Data taking

Order few (3-4) trips in detector life so far, current dependence with humidity

HV monitor



On L3, Discharges of hudrends of nA that do not prevent the operations or the data taking

LV and DAQ monitor - I



Stable FEB temperature, voltages, currents during all the operations

FEB IA

FEB ID

LV and DAQ monitor – II

- During cosmic data taking, typical run integrates few hours of triggers, but we do reconfigure FEBs after 30 minutes (or 30k events)
- We have taken two runs of 60 (run 1044) and 75 (run 1045) minutes with no reconfiguration (closer to BESIII standard "decay mode" data taking) with no issue



