

*Proposal to the ISOLDE and Neutron Time-of-Flight Committee
77th Meeting - 12/11/24*

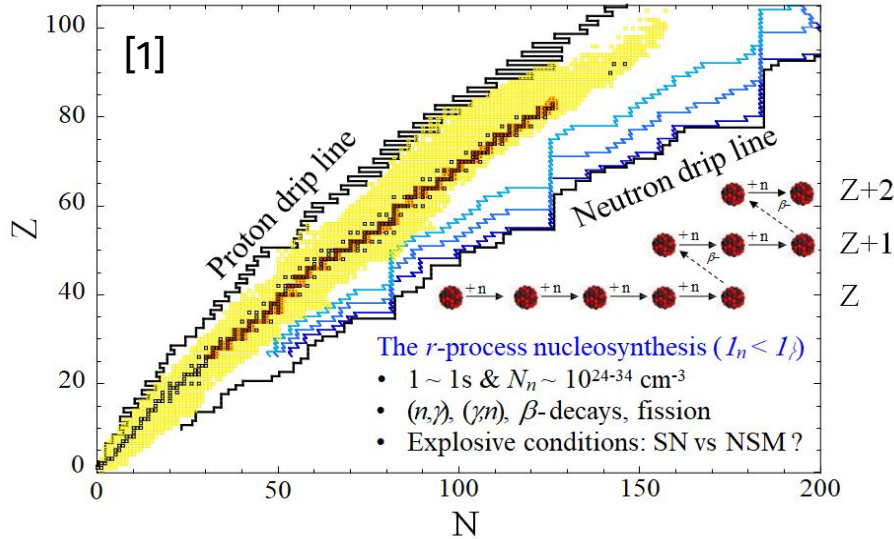
Mass measurements of neutron-rich Ag and In isotopes for r-process nucleosynthesis studies

INTC P-716

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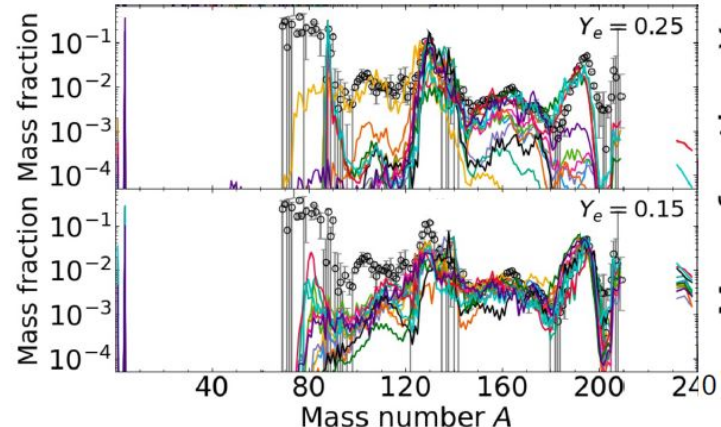


r - process



15 different “acceptable” sets of nuclear inputs (masses, β -decay, n-capture, fission) [2]

Different color line = Different r-process model:



neutron photo-disintegration rate:

$$\lambda_{(\gamma, n)} = \frac{2}{n_n} \frac{G'(T)}{G(T)} \left(\frac{\mu k_B T}{2\pi \hbar^2} \right)^{3/2} \lambda'_{(n, \gamma)} e^{-S_n/k_B T} \quad [3]$$

Uncertainties remaining:

- observed r-process nuclei in NSM, major site?
- process conditions and path
- competition between β -decay and fission

[1] S. Goriely, 3rd PhyNuBE meeting (2024) : Fission and Nuclear Astrophysics

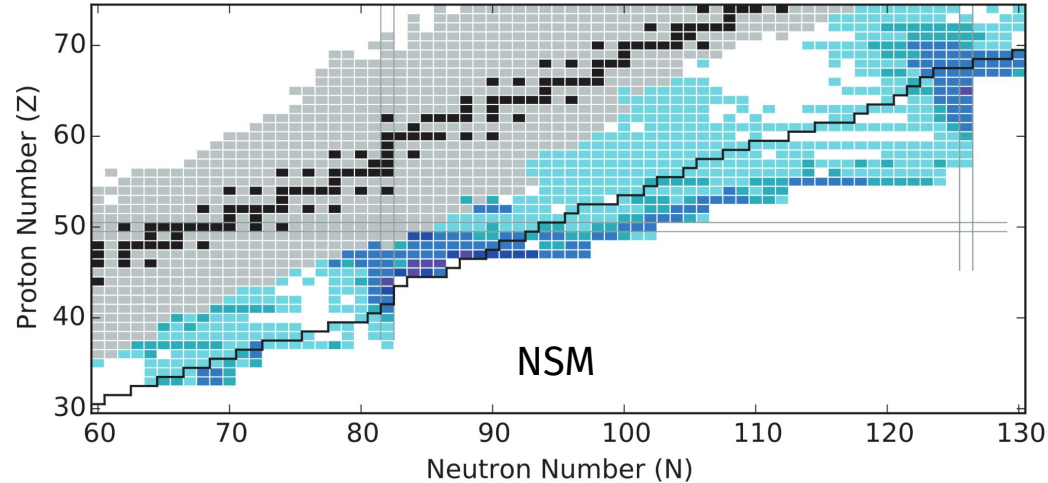
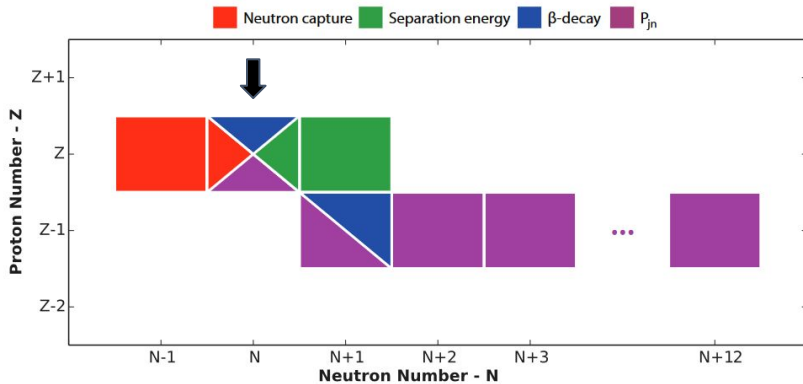
[2] I. Kullmann et al., 10.1093/mnras/stad1458 (2023)

[3] J. Clark et al., 10.1140/epja/s10050-023-01037-0 (2023)

Nuclear Masses

Nuclear masses to calculate important properties :

- neutron capture rates
- separation energy
- β -decay rates
- β -delayed emission probabilities



Nuclei **around** the **magic numbers** have **largest impact** on **r-process** abundances according to sensitivity studies
 study of isomers -> impact on reaction rates

Actinide Production

2017: observation of Sr and lanthanides in NSM kilonova but :

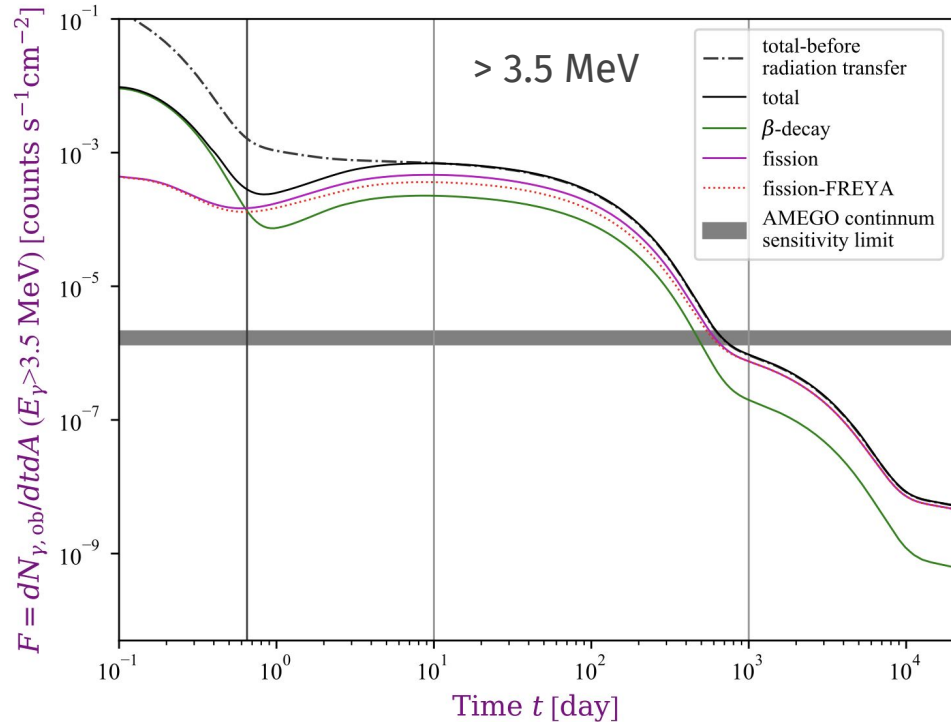
- ↳ no actinides observed yet
- ↳ close NSM not common

Prediction of **NSM light spectrum** 3.5-10 MeV:

- **β -decay** γ -rays dominant **earlier**
- **fission** γ -rays dominant **later**

γ -rays > 3.5 MeV \Rightarrow low flux :

- ↳ NSM observed only in close neighborhood



Isomeric yield ratios

β -decay of fission fragments emit γ -rays ~ 1 MeV:

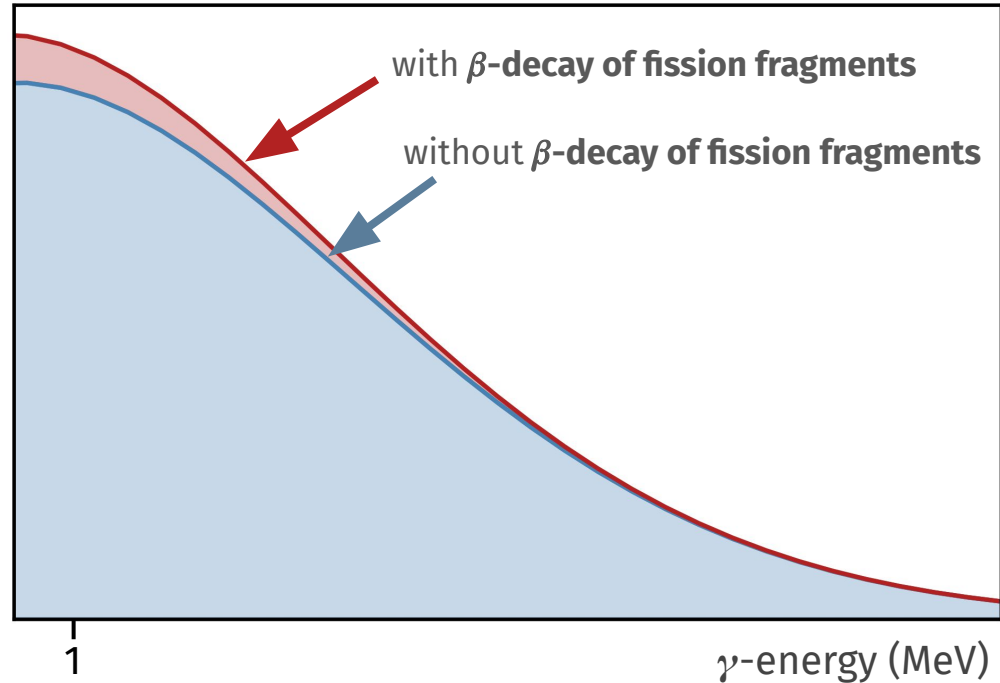
- ↳ light curve **intensity enhanced**
- ↳ **higher flux** at 1 MeV
- ↳ further distance **NSM** observations

⇒ **Isomeric Yield Ratios** to:

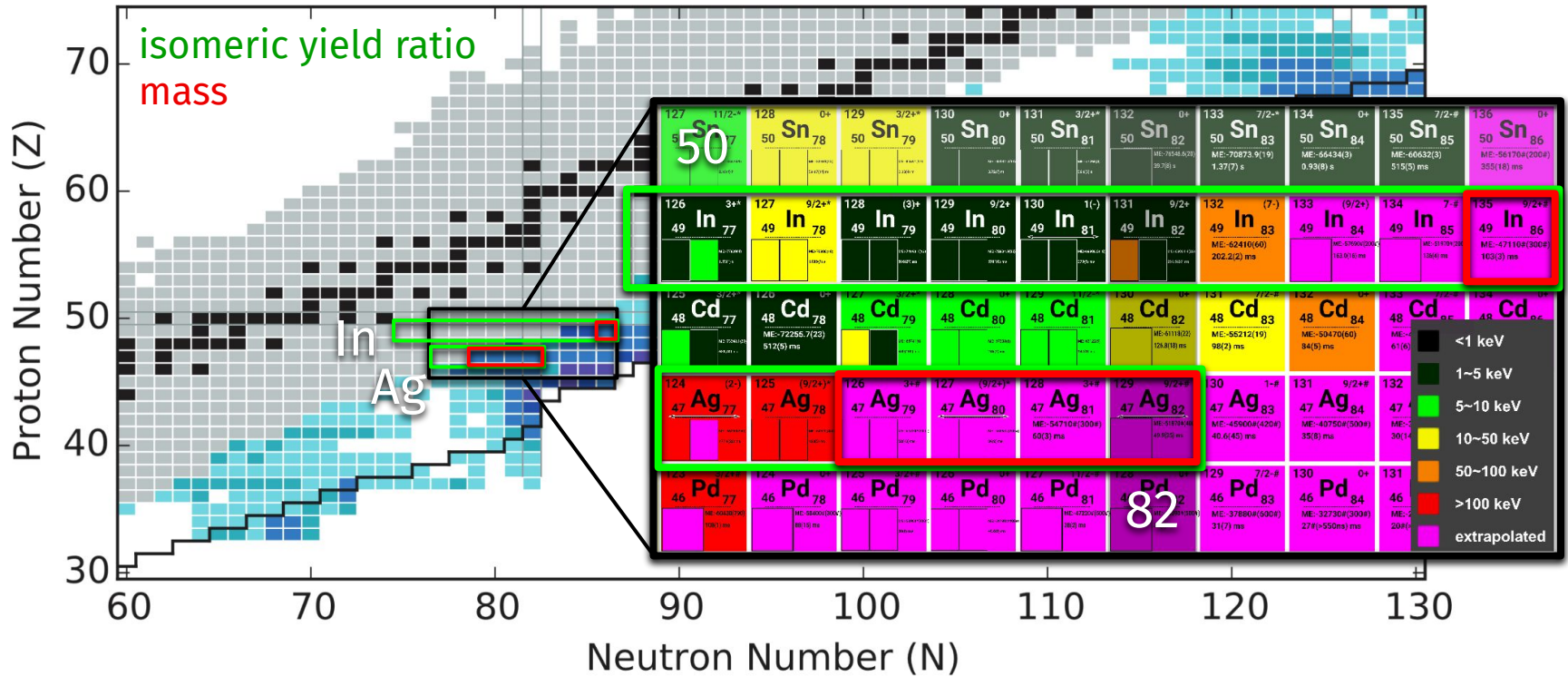
- ↳ quantify the **kilonova light curve intensity**
- ↳ identify Actinides in **multi-messenger astronomy** observations
- ↳ better knowledge of isomeric state **production at ISOLDE**

γ -spectra
(counts/MeV)

~ 10 days after merger



Nuclei of interest

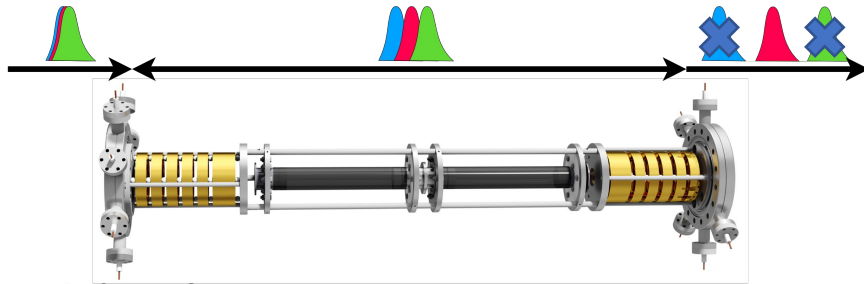


Experimental Techniques

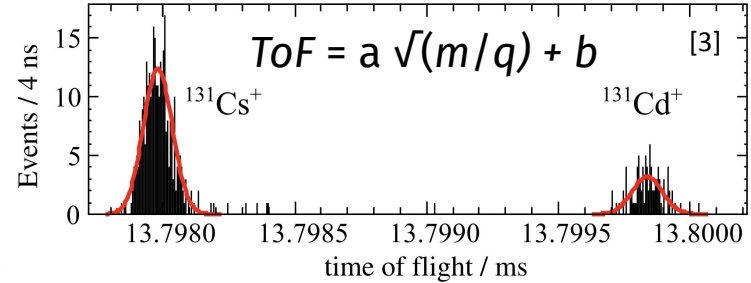
MR-ToF MS [1]:

versatile and fast

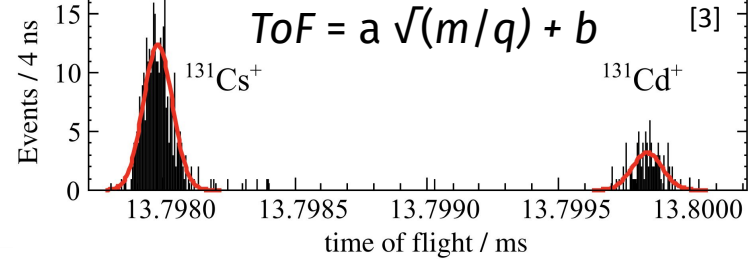
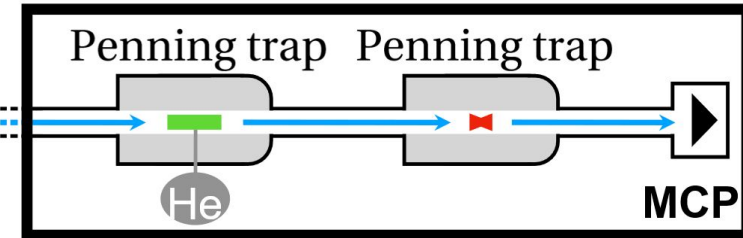
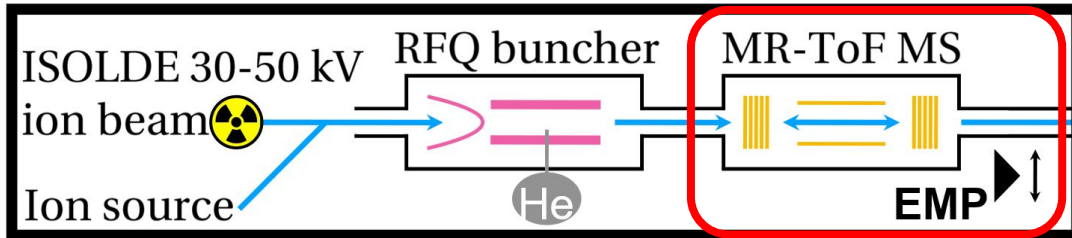
- **Mass separation** mode: mass selective ejection out of MR-ToF MS [2]
⇒ **removal of isobaric contaminants**
- mass spectrometry mode
⇒ up to $R = 5 \times 10^5$



Horizontal section



Vertical section

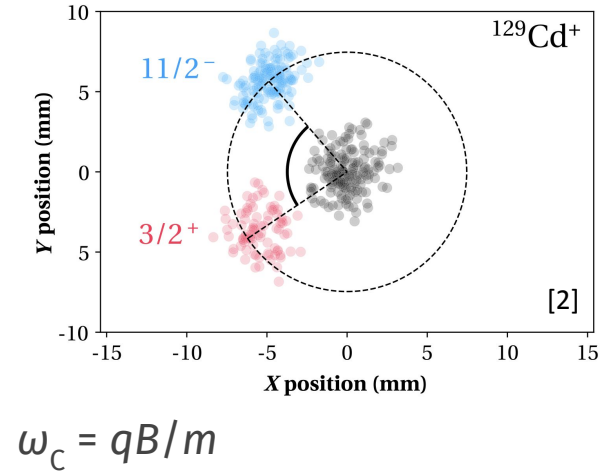
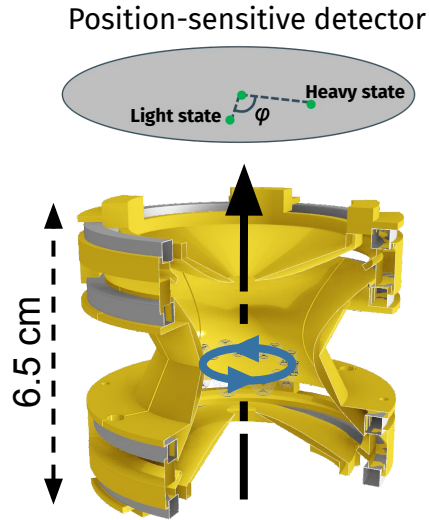


[1] R. N. Wolf *et al.*, [10.1016/j.iims.2013.03.0201](https://doi.org/10.1016/j.iims.2013.03.0201) (2013)
 [2] F. Wienholtz *et al.*, [10.1016/j.jiims.2017.07.016](https://doi.org/10.1016/j.jiims.2017.07.016) (2017)
 [3] D. Atanasov *et al.*, [10.1103/PhysRevLett.115.2325012](https://doi.org/10.1103/PhysRevLett.115.2325012) (2015)

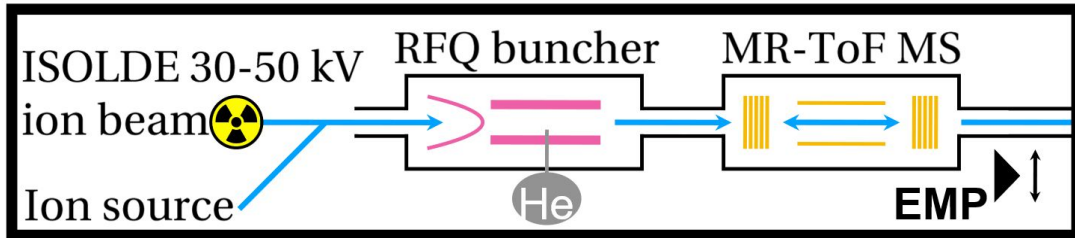
Experimental Techniques

PI-ICR [1]: high precision even with low yields

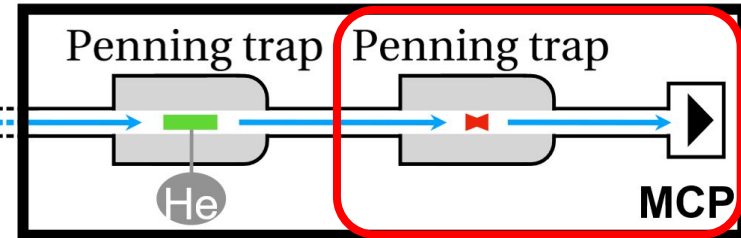
- Mass resolving power $R > 10^6$ in 100 ms at $A = 129$
- Longer storage times improve resolving power
- **Isomeric separation** capability demonstrated



Horizontal section



Vertical section



[1] S. Eliseev et al., [10.1103/PhysRevLett.110.082501](https://doi.org/10.1103/PhysRevLett.110.082501) (2013)

[2] V. Manea et al., [10.1103/PhysRevLett.124.092502](https://doi.org/10.1103/PhysRevLett.124.092502) (2020)

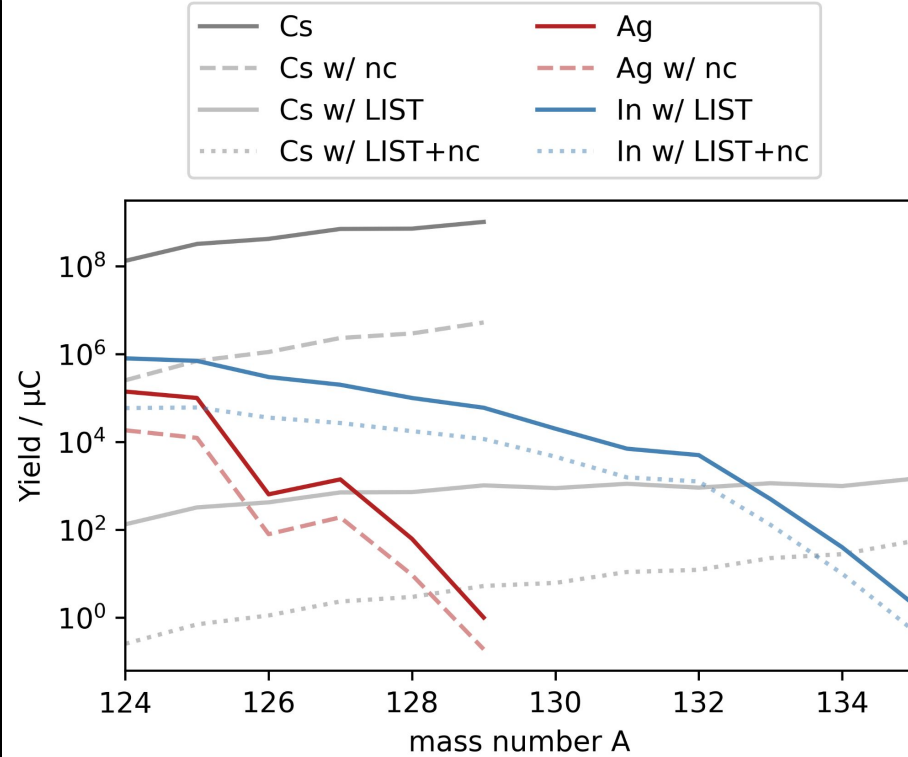
Yields and contamination

Yields:

- multiple shifts of data taking required for most exotic isotopes of Ag and In

Contamination: Cs and Ba

- **neutron converter** to improve IOI-to-contaminant ratio (also reduces IOI yield by order of magnitude)
- **LIST** for **In** to further reduce contaminants by $\approx 10^6$
 - both narrow and broadband (like in recent Hg measurement)
- **MR-ToF MS** allows separation of remaining contaminants in tens of milliseconds



Shift Summary

Indium shifts (UCx + nc & RILIS LIST)

Nuclide	$T_{1/2}$	Yield / μC	cts / h	Shifts
$^{124,125}\text{In}^{(m)}$	2-12 s	$\approx 1 \times 10^6$	$\approx 1 \times 10^7$	1
$^{126-131}\text{In}^{(m-q)}$	0.1-4 s	$\geq 7 \times 10^3$	$\geq 3 \times 10^4$	3
^{132}In	202 ms	$\approx 5 \times 10^3$	$\approx 2 \times 10^4$	2
^{133}In	163 ms	≈ 500	≈ 2100	2
$^{133}\text{In}^m$	167 ms			
^{134}In	136 ms	≈ 40	≈ 180	3
^{135}In	103 ms	≈ 2	≈ 2	4
Total with radioactive beam				15 + 2 = 17

Silver shifts (UCx + nc & RILIS)

Nuclide	$T_{1/2}$	Yield / μC	cts / h	Shifts
$^{124,125}\text{Ag}^{(m)}$	50#-178 ms	$\approx 1 \times 10^5$	$\geq 1 \times 10^5$	2
^{126}Ag	52 ms	$\approx 1 \times 10^3$	≈ 70	2
$^{126}\text{Ag}^m$	108 ms			
^{127}Ag	89 ms	$\approx 1 \times 10^3$	≈ 600	2
$^{127}\text{Ag}^m$	20# ms			
$^{127}\text{Ag}^n$	86 ms			
^{128}Ag	60 ms	≈ 60	≈ 10	3
^{129}Ag	50 ms	≈ 1	≈ 5	4
$^{129}\text{Ag}^m$	10# ms			
Total with radioactive beam				13 + 2 = 15

Summary

- **masses** for second abundance peak of **r-process** to pin down its **path** through the nuclear landscape
- **isomeric yield ratios** of neutron-rich nuclei as an alternative for **identifying actinides in NSMs**
- isotope separation in MR-ToF MS
- simultaneous mass and yield-ratio measurement with Penning trap (PI-ICR)

