Chemical Freeze-Out in the QCD Phase Diagram

XVII Polish Workshop on Relativistic Heavy-Ion Collisions

Phase diagram and Equation of State of strongly interacting matter

Chemical Freeze-Out in the QCD Phase Diagram

David Blaschke

In collaboration with:



CASUS











HELMHOLTZ ZENTRUM DRESDEN ROSSENDORF

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The Goal: Proving the phase structure!

Lattice QCD results only for pseudocritical temperature Tc near $\mu \sim 0$ (sign problem)

Liquid-gas PT indicated in experiment

Other structures are so far model dependent conjectures!!

G. Endrödi, SQM2024





V. Dexheimer, QCHS2024



Fig. 1 Regions of the QCD phase diagram where constraints from heavy-ion collisions (HIC), lattice QCD (LQCD), perturbative QCD (pQCD), lowenergy heavy-ion collisions (LENP), chiral effective field theory (χ EFT), and astrophysics (neutron stars, NS) are available

Living Reviews in Relativity (2024)27:3 https://doi.org/10.1007/s41114-024-00049-6



This Talk: Chemical Freeze-Out @ T ~ 20 – 100 MeV

- Statistical Model Fit, T µ diagram
- CFO in the T n diagram
- Mott dissociation for light clusters
- CFO as inverse Mott dissociation
- Summary & Outlook



Statistical Model Fit for CFO, T – µ Diagram



Statistical Model Fit for CFO, T – n Diagram



Statistical Model Fit for CFO, T – n Diagram



Correlation of CFO line with chiral restoration/ deconfinement gets lost at T < 140 MeV

Is CFO in the baryon-dominated region correlated with Mott lines for dissociation of light clusters?

Blaschke et al., Springer Proc. Phys. 250, 183 (2020); arXiv:2001.02156 (SQM 2019)

Mott dissociation for bound states in a plasma

Chemical picture:

Ideal mixture of reacting components Mass action law



Interaction between the components internal structure: Pauli principle

Physical picture:

"elementary" constituents and their interaction



Quantum statistical (QS) approach, quasiparticle concept, virial expansion

Mott dissociation for bound states in a plasma

Effective wave equation for deuterons in nuclear matter

In-medium two-particle wave equation in mean-field approximation $\left(\frac{p_1^2}{2m_1} + \Delta_1 + \frac{p_2^2}{2m_2} + \Delta_2\right)\Psi_{d,P}(p_1, p_2) + \sum_{p_1', p_2'}(1 - f_{p_1} - f_{p_2})V(p_1, p_2; p_1', p_2')\Psi_{d,P}(p_1', p_2')$ Add self-energy $= E_{d,P}\Psi_{d,P}(p_1, p_2)$

Fermi distribution function

$$f_p = \left[e^{(p^2/2m-\mu)/k_B T} + 1 \right]^{-1}$$

Thouless criterion $E_d(T,\mu) = 2\mu$

BEC-BCS crossover: Alm et al.,1993

Binding energies for light clusters in T – n plane



Vanishing binding energies Indicate Mott effect for the Light clusters!

Mott-lines in the T-µ plane can be extracted, where the Binding energy vanishes

Here lower temperatures: 0 < T[MeV] < 20

S. Typel et al., PRC 81, 015803 (2010)

Binding energies for light clusters in T – n plane



Pauli blocking: phase space occupation



momentum space

The deformation is maximal at P = 0. It leads to the weakening of the interaction (disintegration of the bound state).

Momentum-dependent binding energies



CFO in the Temperature – density plane



CFO in the Temperature – density plane



Main result:

Chemical freeze-out may be interpreted as "inverse" Mott transition:

Strong localization effect of nucleon-nucleon correlations in bound states (clusters) entails freeze out of the nuclear composition

D. Blaschke, B. Dönigus, S. Liebing, G. Röpke, arXiv:2408.01399, PLB accepted

Outlook: Primordial CFO of heavy elements?



Outlook: Primordial CFO of heavy elements?



Can the primordial evolution of the Universe lead to these freeze-out Parameters (red star):

T=5 MeV, μ_n=940.317 MeV, μ_p= 845.069 MeV

Maybe inhomogeneous Big Bang?

The freeze-out point lies in the domain of supernova explosions and binary neutron star mergers

G. Röpke, D. Blaschke, F. Röpke, arXiv:2411.00535

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Phase diagram and Equation of State of strongly interacting matter

Mott-Anderson localization model for sudden CFO

The basic idea: Localization of (certain) multiquark states ("cluster") = hadronization; Reverse process = delocalization by quark exchange between hadrons

Freeze-out criterion:

Povh-Huefner law, PRC 46 (1992) 990 → total x-section



(7



Hippe & Klevansky, PRC 52 (1995) 2172



Mott-Anderson localization model for sudden CFO

Model results:

 $\tau_{exp}(T, \mu) = \tau_{coll}(T, \mu)$

Collision time strongly T, mu dependent !

Schematic resonance gas: dn pions, dN nucleons



Expansion time scale from entropy conservation:

$$s(T, \mu) V(\tau_{exp}) = \text{const}$$

 $\tau_{\exp}(T,\mu) = as^{-1/3}(T,\mu),$

Thermodynamics consistent with phenomenological Freeze-out rules:



A) Chiral condensate for the full hadron resonance gas model \rightarrow radii of hadrons

- nonstrange hadrons: $\begin{aligned} \langle r_{\pi}^2 \rangle_{T,\mu} &= \frac{3}{4\pi^2 f_{\pi}^2} \qquad f_{\pi}^2(T,\mu) = \frac{-m_q \langle \bar{q}q \rangle_{T,\mu}}{m_{\pi}^2} , \\ \langle r_{\pi}^2 \rangle_{T,\mu} &= \frac{3m_{\pi}^2}{4\pi^2 m_q} |\langle \bar{q}q \rangle_{T,\mu}|^{-1} \qquad \langle r_{\rm N}^2 \rangle_{T,\mu} = r_0^2 + \langle r_{\pi}^2 \rangle_{T,\mu} \end{aligned}$

- strange hadrons:
$$f_K^2 m_K^2 = -\frac{\langle \bar{q}q \rangle_{T,\mu} + \langle \bar{s}s \rangle_{T,\mu}}{2} (m_q + m_s)$$
$$\langle r_K^2 \rangle_{T,\mu} = \frac{3}{4\pi^2 f_K^2} = \frac{3}{2\pi^2} \frac{m_K^2}{|\langle \bar{q}q \rangle_{T,\mu} + \langle \bar{s}s \rangle_{T,\mu}|(m_q + m_s)} \qquad \langle r_\Lambda^2 \rangle_{T,\mu} = r_0^2 + \langle r_K^2 \rangle_{T,\mu}$$

B) Chemical freeze-out: only "reactive" cross section, flavor equilibration

Some flavor changing processes involve reaction thresholds and need activation energy, like in the Eyring theory of chemical processes with activation:

$$\sigma_{ij}^{\star}(T,\mu) = \exp\left(-\frac{\Delta E}{T}\right) \sigma_{ij}(T,\mu) \qquad \qquad \sigma_{ij}(T,\mu) = \lambda \langle r_i^2 \rangle_{T,\mu} \langle r_j^2 \rangle_{T,\mu}$$

Assumption: average activation threshold for reactive processes: $\Delta E = 0.2 \text{ GeV}$ (to be refined, account for all individual processes, e.g., SMASH)



$$\begin{split} \langle \bar{q}q \rangle_{T,\mu} &= \langle \bar{q}q \rangle_{T,\mu}^{MF} + \sum_{h=M,B} \frac{\sigma_q^h}{m_q} n_h(T,\mu) \ ,\\ n_h(T,\mu) &= \frac{d_h}{2\pi^2} \int_0^\infty dk k^2 \frac{m_h}{E_h} \frac{1}{\mathrm{e}^{(E_h - \mu_h)/T} \mp 1} \\ \tau_{\mathrm{coll},i}^{-1}(T,\mu) &= \sum_j \sigma_{ij}^{\star} v n_j(T,\mu) \ ; \ \sigma_{ij} = \lambda \langle r_i^2 \rangle \langle r_j^2 \rangle \\ \langle r_\pi^2 \rangle_{T,\mu} &\simeq \frac{3}{4\pi^2} f_\pi^{-2}(T,\mu) = \frac{3M_\pi^2}{4\pi^2 m_q} |\langle \bar{q}q \rangle_{T,\mu}|^{-1} \\ \langle r_K^2 \rangle_{T,\mu} &\simeq \frac{3M_K^2}{\pi^2(m_q + m_s)} |\langle \bar{q}q \rangle_{T,\mu} + \langle \bar{s}s \rangle_{T,\mu}|^{-1} \end{split}$$

The factor a stands for the inverse system size in the formula

 $\tau_{exp}(T, \mu) = \tau_{coll}(T, \mu)$

for the 3D expansion time scale assuming entropy conservation

(2013) Rev Full HRG model condensate; Phys a et ankowski

Inelastic collision rate $\tau_{\rm coll} \propto T^{\kappa}, \kappa \gtrsim 20$ from fit to STAR data

U. Heinz and G. Kestin, PoS CPOD 2006, 038 (2006) [nucl-th/0612105]





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for the 3D expansion time scale assuming entropy conservation

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arxiv:1

M. Naskret

Jankowski,

B

CFO in the Temperature – density plane



CFO in the Temperature – density plane

