

Revisiting Universal Extra-Dimension Model with Gravity Mediated Decays

in collaboration with Prof. Katri Huitu and Rameswar Sahu

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Universal Extra Dimension (UED)

UED scenarios are characterized by a **single flat universal** (accessible to all the SM particles) extra dimension (y), compactified on a S_1/Z_2 orbifold with radius R .

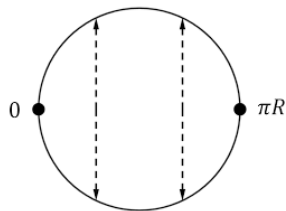
$$\mathcal{S}_{UED} = \mathcal{S}_{\text{gauge}} + \mathcal{S}_{\text{fermions}} + \mathcal{S}_{\text{scalar}}$$

$$\mathcal{S}_{\text{gauge}} = \int d^4x \int_0^{\pi R} dy \left\{ -\frac{1}{4} G_{MN}^a G^{aMN} + \right.$$

$$\left. \left(\delta(y) + \delta(y - \pi R) \right) \left[-\frac{r_G}{4} G_{\mu\nu}^a G^{a\mu\nu} \right] \right\}$$

① Bulk terms

② Boundary localized terms (BLTs).

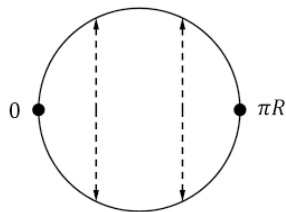
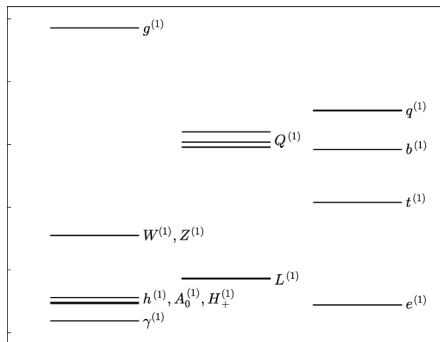


After Compactification



- ① Kaluza-Klein modes.
- ② Conserved KK-number.
- ③ Conserved KK-parity.
 - ① A dark matter candidate

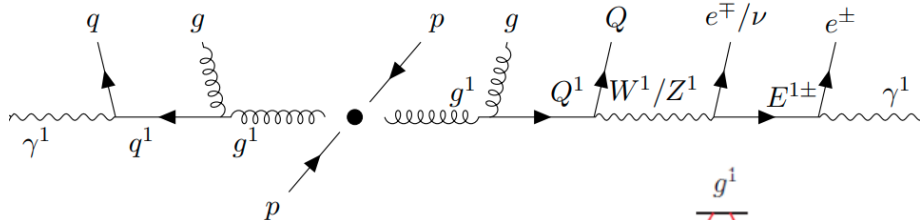
minimal Universal Extra Dimension (mUED)



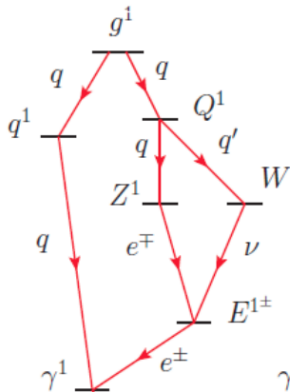
- 1 In the *minimal* version of UED (mUED), all BLTs are assumed to vanish at the cut-off scale (Λ).
- 2 The phenomenology of mUED is determined by only two additional parameters:

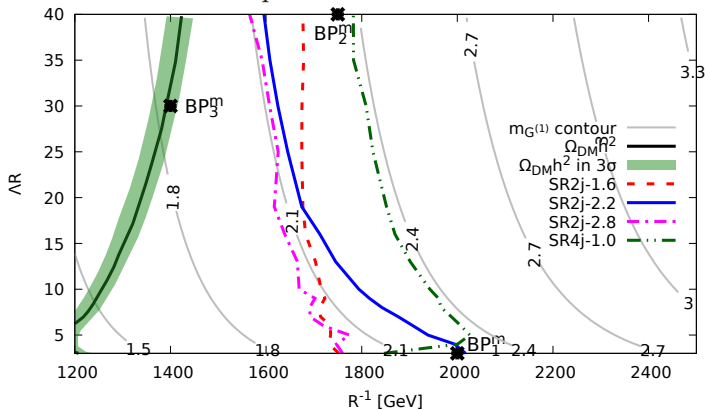
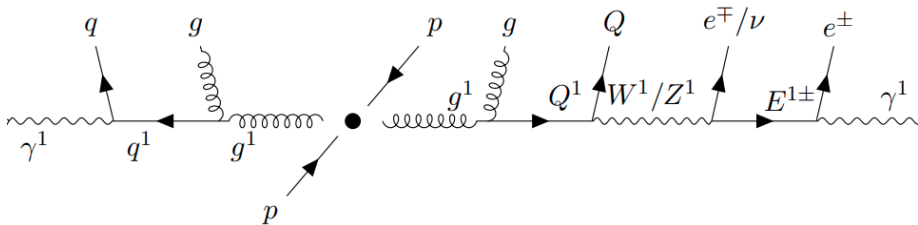
- 1 The radius of compactification, R
- 2 The cut-off scale, Λ

BPs	R^{-1} [TeV]	ΛR	$m_{G^{(1)}}$ [TeV]	$m_{Q^{(1)}}$ [TeV]	$m_{q^{(1)}}$ [TeV]	$m_{W^{(1)}/Z^{(1)}}$ [TeV]	$m_{L^{(1)}}$ [TeV]	$m_{B^{(1)}}$ [TeV]
BP_1^m	2.0	3	2.222	2.143	2.124	2.045	2.022	1.998
BP_2^m	1.75	40	2.341	2.171	2.122	1.877	1.814	1.748
BP_3^m	1.4	30	1.840	1.668	1.710	1.495	1.448	1.399

mUED after ATLAS search (ATLAS-CONF-2019-040) for multijets + E_T with $\mathcal{L} = 139 \text{ fb}^{-1}$


	SR2j-1.6	SR2j-2.2	SR2j-2.8	SR4j-1.0	SR4j-2.2
n_j	≥ 2	≥ 2	≥ 2	≥ 4	≥ 4
$p_{T}(j_1) [\text{GeV}]$	> 250	> 600	> 250	> 200	> 200
$p_{T}(j_{i=2, \dots, n_{\min}}) [\text{GeV}]$	> 250	> 50	> 250	> 100	> 100
$ \eta(j_{1, \dots, n_{\min}}) $	< 2.0	< 2.8	< 1.2	< 2.0	< 2.0
$\Delta\phi(j_{1,2,(3)}, p_{T}^{\text{mis}})_{\text{min}}$	> 0.8	> 0.4	> 0.8	> 0.4	> 0.4
$\Delta\phi(j_{i>3}, p_{T}^{\text{mis}})_{\text{min}}$	> 0.4	> 0.2	> 0.4	> 0.4	> 0.4
Aplanarity	-	-	-	> 0.04	> 0.04
$\frac{E_T}{\sqrt{H_T}} [\sqrt{\text{GeV}}]$	> 16	> 16	> 16	> 16	> 16
$m_{\text{eff}} [\text{TeV}]$	> 1.6	> 2.2	> 2.8	> 1.0	> 2.2
$\langle \epsilon\sigma \rangle_{\text{obs}}^{95} [\text{fb}]$	1.46	0.78	0.13	0.54	0.14
<i>mUED</i> predictions [fb]					
BP_1^m	0.35	0.92	0.03	0.28	0.03
BP_2^m	0.96	0.29	0.05	0.58	0.11
BP_3^m	5.25	2.96	0.33	6.41	0.75





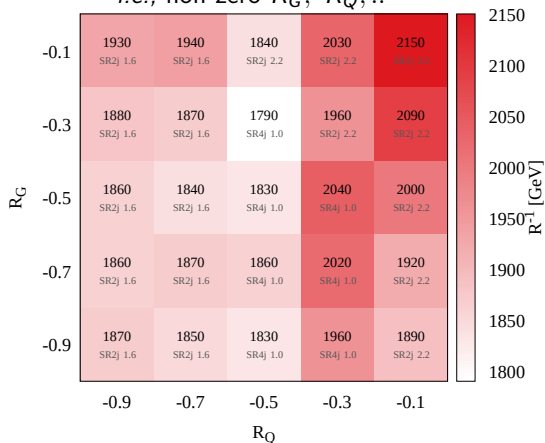
mUED
is
Ruled Out
!!

[Avnish, Ghosh, Jha,
Niyogi PRD 2021;
PDG 2023]

Options beyond mUED

1 Non-minimal UED:

- Non-vanishing BLTs
i.e., non-zero R_G , R_Q , ..



Next-to-minimal UED:

e.g., 'Fat-brane' realization of mUED



mUED is embedded in a $(4 + N)$ -dimensional bulk.



Gravity spans across N large extra dimensions $\sim eV^{-1}$, while the propagation of matter is confined to a length scale $\sim TeV^{-1}$

- 1 $M_{Pl}^2 = M_D^{N+2} \left(\frac{r}{2\pi}\right)^N$
- 2 Gravity-Matter interaction
- 3 Gravity mediated decay of LKP

'Fat-brane' realization of mUED: Matter-Gravity Interaction

(4 + N)-D Action

$$S = \int dx^{4+N} \delta(x^5) \dots \delta(x^{4+N}) \sqrt{-\hat{g}} \mathcal{L}_m$$



(4 + N)-D Gravity-Matter Interaction Action

$$S_{\text{int}} = -\frac{\hat{k}}{2} \int d^{N+4}x \delta(x^5) \dots \delta(x^N) \hat{h}^{\hat{\mu}\hat{\nu}} T_{\hat{\mu}\hat{\nu}}$$



4-D Gravity-Matter Interaction Action

$$S_{\text{int}} = -\frac{k}{2} \int d^4x \sum_n \left\{ \left[\tilde{h}^{\vec{n}}_{\mu\nu} + \omega \left(\eta_{\mu\nu} + \frac{\partial_\mu \partial_\nu}{m_n^2} \right) \tilde{\phi}^{\vec{n}} \right] T_{n5}^{\mu\nu} - 2\tilde{A}_{\mu 5}^{\vec{n}} T_{n5}^\mu + \left(\sqrt{2}\tilde{\phi}_{55}^{\vec{n}} - 3\omega a \left(1 - \frac{n_5^2}{n^2} \right) \tilde{\phi}^{\vec{n}} \right) T_{55}^{n5} \right\}$$

where, $T_{MN}^{n5}(x) = \int_0^{\pi R} dy T_{MN}(x, y) e^{\frac{2\pi i n_5 y}{r}}$

$$\hat{g}_{\hat{\mu}\hat{\nu}} = \eta_{\hat{\mu}\hat{\nu}} + \hat{k} \hat{h}_{\hat{\mu}\hat{\nu}}, \quad \hat{k}^2 = 16\pi G^{4+N}$$

$$\hat{h}_{\hat{\mu}\hat{\nu}}(x, y) = \sum_{\vec{n}} \hat{h}_{\hat{\mu}\hat{\nu}}^{\vec{n}}(x) \exp\left(i \frac{2\pi \vec{n} \cdot \vec{y}}{r}\right)$$

$$\hat{h}_{\hat{\mu}\hat{\nu}}^{\vec{n}} = V_N^{-1/2} \begin{pmatrix} h_{\mu\nu}^{\vec{n}} + \eta_{\mu\nu} \phi^{\vec{n}} & A_{\mu i}^{\vec{n}} \\ A_{\nu j}^{\vec{n}} & 2\phi_{ij}^{\vec{n}} \end{pmatrix}$$

The KK-gravity excitations:

1 Gravitons:

$$\tilde{h}_{\mu\nu}^{\vec{n}} = h_{\mu\nu}^{\vec{n}} - \omega \left(\frac{\partial_\mu \partial_\nu}{m_n^2} - \frac{1}{2} \eta_{\mu\nu} \right) \tilde{\phi}^{\vec{n}}$$

2 Graviphotons:

$$\tilde{A}_{\mu i}^{\vec{n}} = A_{\mu i}^{\vec{n}}$$

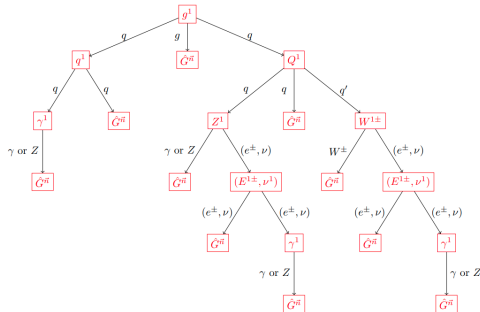
3 Graviscalars:

$$\frac{1}{\sqrt{2}} \tilde{\phi}_{ij}^{\vec{n}} = \phi_{ij}^{\vec{n}} + \frac{3\omega a}{2} \left(\delta_{ij} - \frac{n_i n_j}{n^2} \right) \tilde{\phi}^{\vec{n}}$$

where, $\omega = \sqrt{2/(3N+6)}$ and a satisfies $3(N-1)a^2 + 6a - 1 = 0$

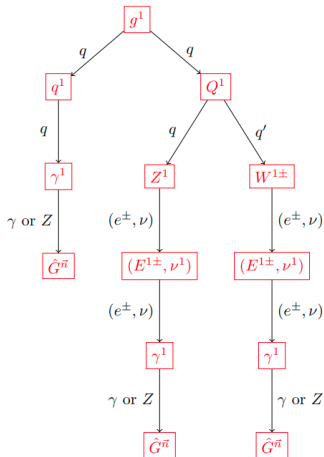
Matter field can decay into any kinematically allowed gravity excitations.

- 1 The decay width into individual graviton modes is suppressed by $1/M_{\text{Pl}}^2$
- 2 Since the mass difference between individual gravity excitations is of the order eV^{-1} , the number of allowed gravity-mediated decay modes is large.
 - Ensure prompt decay of LKP



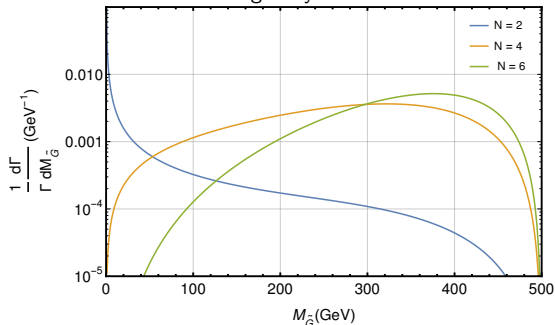
$$\Gamma = \sum_{\vec{n}} \Gamma_{\vec{n}} = \sum \Gamma_{h\vec{n}} + \Gamma_{A\vec{n}} + \Gamma_{\phi\vec{n}}$$

$$\Gamma = \frac{M_{\text{Pl}}^2}{M_D^{N+2}} \int dm d\Omega \Gamma_{\vec{n}} m_{\vec{n}}^{N-1}$$



Event Simulation

The normalized partial decay width for level-1 KK-gluon into a gluon and gravity excitations as a function of the mass of the gravity excitations.



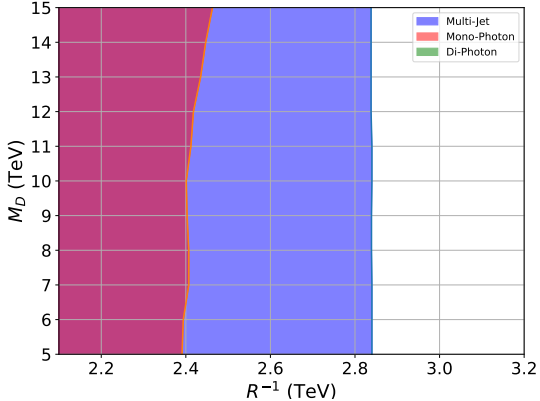
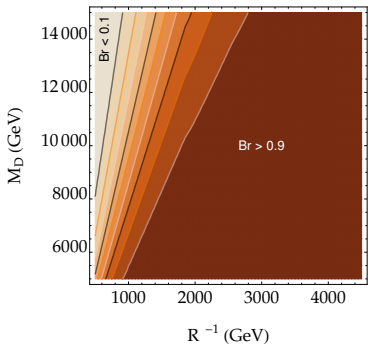
- 1 Production: MG5_AMC@NLO
- 2 Decays: PYTHIA8
 - 1 PYTHIA8 does not consider gravity-mediated decay (GMD)
 - 2 GMDs cannot be trivially incorporated by modifying the decay table in the SLHA file.
 - 3 Modified the PYTHIA PYWIDTH subroutine according to our requirements.
- 3 Showering/hadronization: PYTHIA8

(Re)interpretation of existing LHC searches

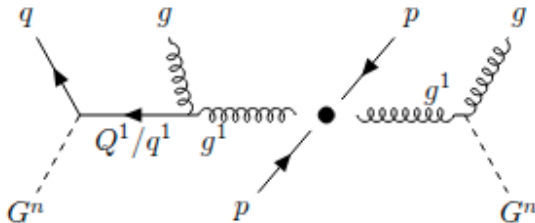
- 1 **ATLAS mono-photon plus missing transverse energy search**
"Search for new phenomena in final states with photons, jets and missing transverse momentum in pp collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector", 2206.06012
- 2 **ATLAS multi-jet plus missing transverse energy search**
"Search for squarks and gluinos in final states with jets and missing transverse momentum using 139 fb^{-1} of $\sqrt{s} = 13$ TeV pp collision data with the ATLAS detector", ATLAS-CONF-2019-040
- 3 **ATLAS di-Photon plus missing transverse energy search**
"Search for photonic signatures of gauge-mediated supersymmetry in 13 TeV pp collisions with the ATLAS detector", 1802.03158

M_{Y_1} M_{G_1}	2.2	2.4	2.6	2.8	3.0	3.2
	2.55	2.78	3.01	3.24	3.48	3.71

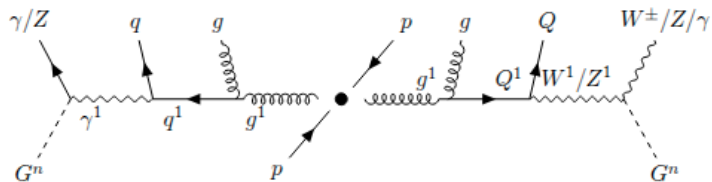
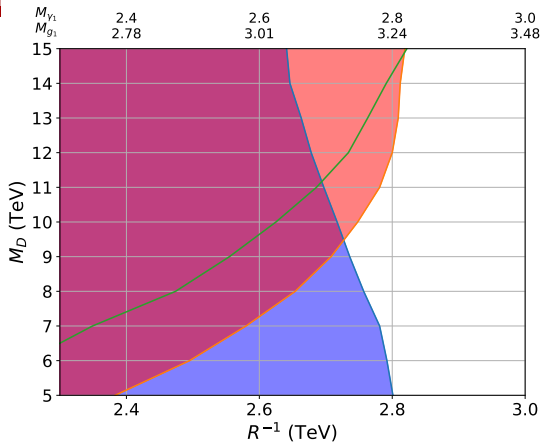
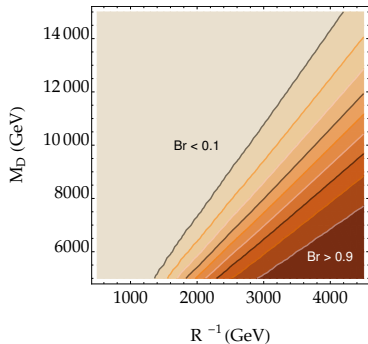
$N = 2$



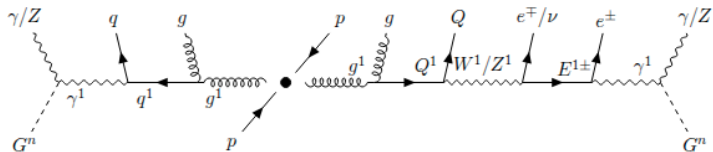
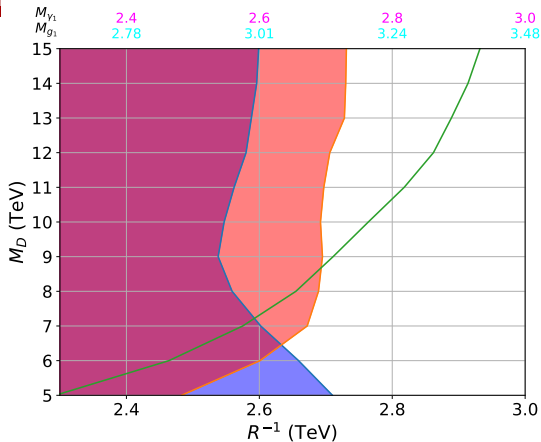
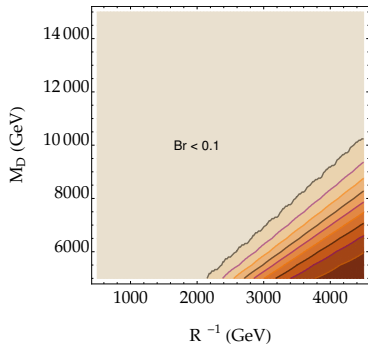
$Br = Br(g^1 \rightarrow GMD)$



$$N = 4$$



$$N = 6$$



Summary

- 1 We studied the fat-brane Minimal Universal Extra Dimensions (MUED) scenario, extending MUED with additional dimensions (eV^{-1} to KeV^{-1}) accessible only via gravity.
- 2 Studied the signatures of the level-1 particles at the LHC, leading to three distinct final states:
 - Multijet + missing transverse energy
 - Mono-photon + missing transverse energy
 - Di-photon + missing transverse energy
- 3 ATLAS has already conducted searches for these final states in MSSM and Gauge Mediated Supersymmetry Breaking (GMSB) models.
- 4 We utilized three ATLAS searches to constrain the fat-brane MUED parameter space.
- 5 Key results for different values of N (number of extra dimensions):
 - $N = 2$: Multijet search provides the best sensitivity, excluding $R^{-1} < 2975$ GeV, independent of M_D .
 - $N = 4$: Multijet and mono-photon searches probe complementary regions:
 - Multijet excludes $R^{-1} < 2898$ GeV for $M_D = 500$ GeV.
 - Mono-photon excludes $R^{-1} < 2958$ GeV for $M_D = 15000$ GeV.
 - $N = 6$: The di-photon search is most effective, setting a lower limit of $R^{-1} > 3002$ GeV for $M_D = 15000$ GeV.

Thank You