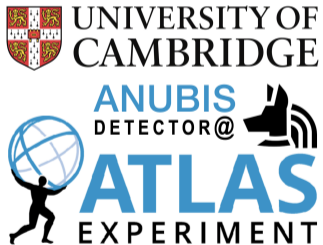


The ANUBIS detector proposal

An overview



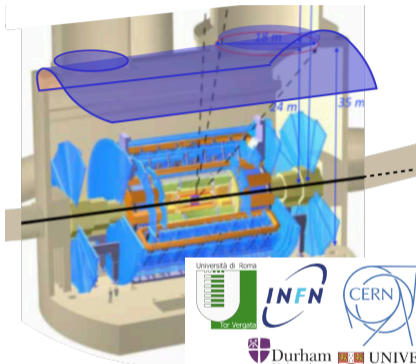
Paul Swallow (He/Him)

University of Cambridge

October 23, 2024

On behalf of the ANUBIS collaboration

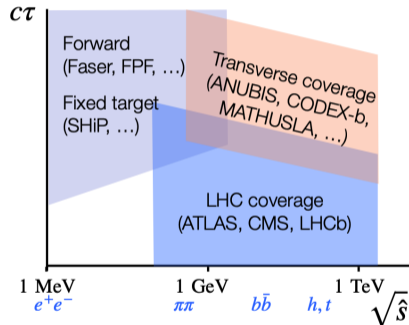
The ANUBIS Detector: Overview

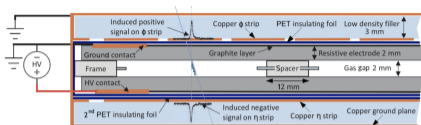


- AN Underground Belayed In-Shaft (ANUBIS) experiment.
 - Proposed in 2019 [1].
- Transverse LLP experiment.
 - $\mathcal{O}(20\text{ m})$ from Interaction Point (IP).
- Based in ATLAS Cavern (IP1) at CERN.
 - Incorporated as an official sub-project.
- Currently have $\mathcal{O}(10)$ Institutes involved.
 - Representing $\mathcal{O}(10)$ FTE.
 - Welcoming new collaborators in many areas.



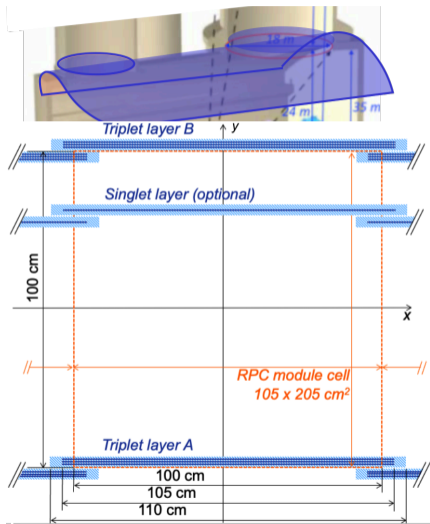
- **Unique sensitivity** to LLPs produced in high- Q^2 collisions.
 - With $m_{LLP} > 1$ GeV and $c\tau \gtrsim \mathcal{O}(10^2 \text{ m})$.
 - Not probed for GPDs or Forward/Beam dump experiments.
- **Unique opportunity** that other transverse experiments **don't** have.
 - Full integration with ATLAS trigger and DAQ system.
 - Get full information of LLP candidates measured in ATLAS **and** ANUBIS.
 - Can search for LLPs with associated production of prompt particles *e.g.* H , W , or Z .
 - Also an active veto for BG events.
- Also has **complementarity** with other experiments.



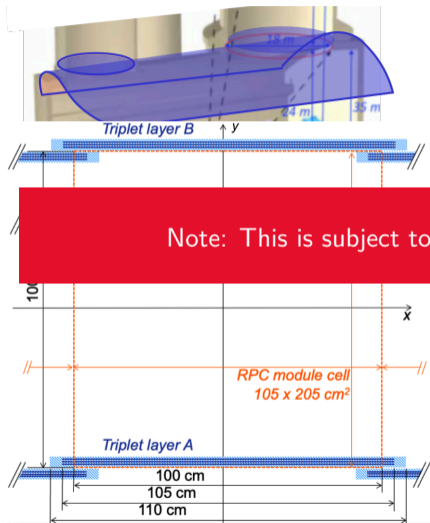


Parameter	Specification
Time resolution	$\delta t \lesssim 0.5 \text{ ns}$
Angular resolution	$\delta \alpha \lesssim 0.01 \text{ rad}$
Spatial resolution	$\delta x, \delta z \lesssim 0.5 \text{ cm}$
Per-layer hit efficiency	$\varepsilon \gtrsim 98\%$

- Instrument w/ ATLAS Phase II RPCs [1].
 - Gaseous detector with: $\text{C}_2\text{H}_2\text{F}_4$ (64%), CO_2 (30%), C_4H_{10} (0.3%) & SF_6 (1%).
 - Active R&D on Eco-gas alternatives.
 - Cheap, large Area, and validated technology.
 - Timing Resolution $\mathcal{O}(250 \text{ ps})$, and Spatial Resolution $\mathcal{O}(3 \text{ mm})$.
- Base unit of the detector geometry is the RPC panel:
 - Aluminium frame for mechanical support: $\sim 1 \times 2 \text{ m}^2$.
 - Frame can contain up to three RPC singlets.
 - Each RPC singlet has readouts either side of the gas gap, divided into strips.



- For ANUBIS: cover ATLAS Cavern ceiling.
- Follow the arc to maximise distance from IP uniformly.
- Two potential designs:
 1. Two separated layers of triplet RPC panels.
 2. As above with additional singlet between to improve vertexing and disambiguation.
- Overall singlet area $\approx 9,800 \text{ m}^2$ and detector area $\approx 1633.3 \text{ m}^2$.
- Extends ATLAS' LLP fiducial volume from $\sim [1 \text{ mm} - 1 \text{ m}]$ to $\mathcal{O}(20 \text{ m})$.
- Covers a large solid angle: $\sim 2.1 \text{ sr}$.
- Smaller-scale prototype in place: **proANUBIS**.

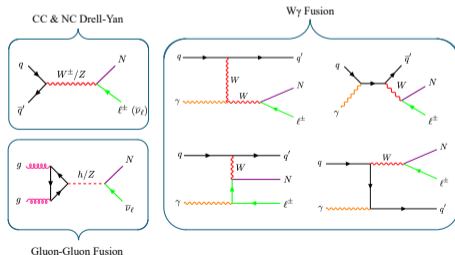
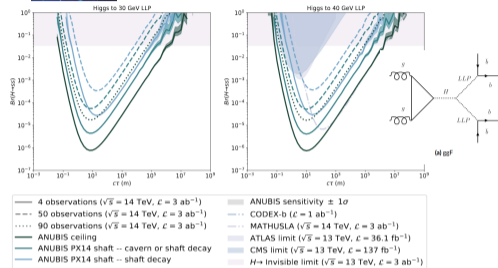


- For ANUBIS: cover ATLAS Cavern ceiling.
 - Follow the arc to maximise distance from IP uniformly.
 - Two potential designs:
 1. Two separated layers of triplet RPC panels.
- detector area $\approx 1633.3 \text{ m}^2$.
- Extends ATLAS' LLP fiducial volume from $\sim [1 \text{ mm} - 1 \text{ m}]$ to $\mathcal{O}(20 \text{ m})$.
 - Covers a large solid angle: $\sim 2.1 \text{ sr}$
 - Smaller-scale prototype in place: **proANUBIS**.

Ongoing Sensitivity Studies



- Studies have been performed previously that include ANUBIS [1-4].
- Internally have produced $H \rightarrow SS$ estimates [5].
- We want to cover the PBC benchmarks at a minimum.
 - Advanced study on HNL benchmarks BC6-8.
 - Developing a flexible, modular sensitivity study framework: **SET-ANUBIS**.



[1]: 2001.04750

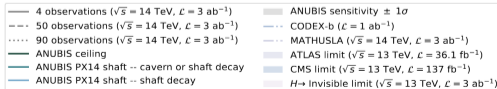
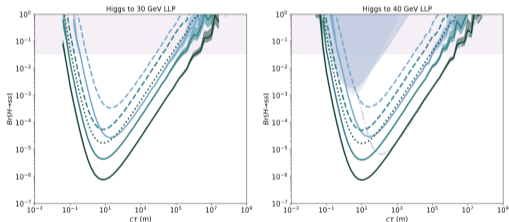
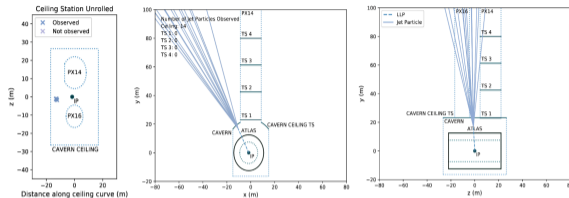
[2]: 2105.13851

[3]: 2010.07305

[4]: 2008.07539

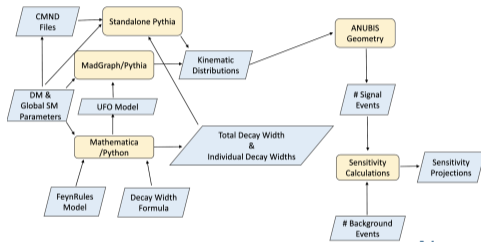
[5]: CDS:2839063

$H \rightarrow SS$ Sensitivity



- Initially looked at PBC model **BC5**.
 - Higgs portal to Dark Scalars.
 - Heavy mediators good targets for transverse detectors.
- Simulated signal events via MADGRAPH and applied cuts.
 - Require Jet and charged particle objects.
 - Simulate ATLAS active-veto and isolation requirements.
 - Acceptance cuts.
- Compared initial geometry (In PX14 service shaft) to current (ceiling).
- Dramatically improved sensitivity for ceiling configuration.
- $B(H \rightarrow SS) \lesssim \mathcal{O}(10^{-6})$, for $m_{LLP} \in [10 - 40]$ GeV.

Goal: Create a flexible framework to handle a variety of LLP models to determine ANUBIS' sensitivity.



$$N_{LLP} = \mathcal{L}_{HL-LHC} \cdot \sigma_{HNL} \cdot \mathcal{B}(HNL) \cdot \frac{N_{Obs}}{N_{gen}}$$

$$N_{LLP} \approx 4 \text{ (0 BG)}$$

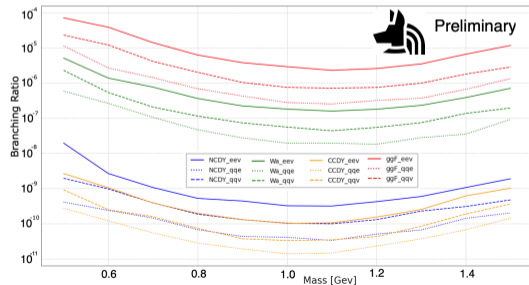
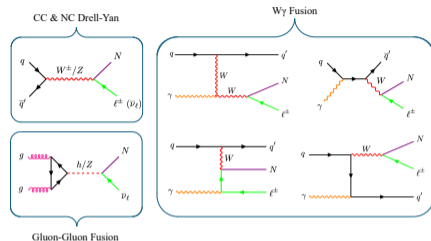
$$N_{LLP} \approx 90 \text{ (Conservative data-driven BG estimate based on [2])}$$

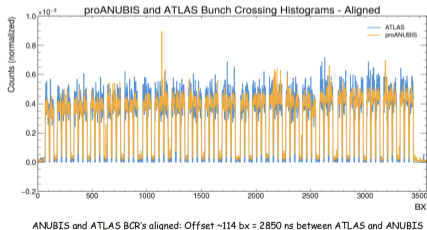
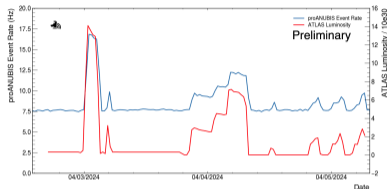
- Based on the work of T. Satterthwaite et al [1], $H \rightarrow SS$ model.
- Methodology:
 - Simulate LLP model (MADGRAPH, PYTHIA etc.).
 - Apply Selection: detector acceptance, background removal e.g. isolation requirements.
 - Calculate sensitivity: Number of LLP candidates (N_{LLP}) required to exceed BG expectations significantly.
- Planned Models:
 - **HNLs**, Dark Photon, Dark Scalar etc.
 - As outlined in PBC proposal [3].

Sensitivity Studies: HNLs

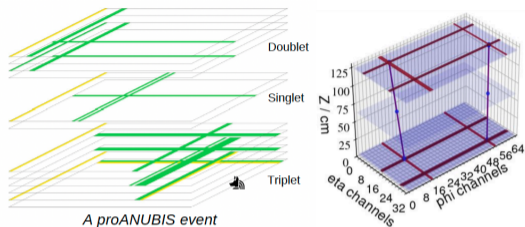


- Two generic production modes:
Hard Interaction Bosons (W, Z, H) and *Mesonic* (B, D).
 - Best expected sensitivity: boosted, high mass HNL produced by $W/Z/H$.
- We use MADGRAPH for **HIB** & PYTHIA for the **mesonic** production and HNL decays.
- Examining HNL decays to:
 $e^\pm qq'$, $\nu_e qq'$, $e^+ e^- \nu_e$.
- For $m_{LLP} \in [0.01, 10]$ GeV.
- Evaluated \mathcal{B} limit for each decay mode.
 - $\mathcal{B}(HNL) \sim \mathcal{O}(10^{-9})$ at 1 GeV.
 - Preliminary results, updated soon.
 - Plot and limit for coupling=1, for easier reinterpretation.

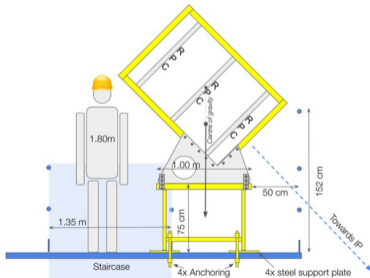




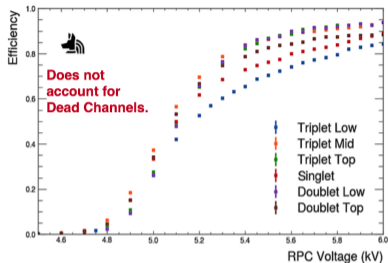
- Lots of ongoing work in analysing the proANUBIS data.
 - Initial studies performed to calibrate data e.g. Time walk corrections.
 - Working on synchronisation with ATLAS data.
 - Saw visual correlation initially.
 - Matching on a Bunch-Crossing level achieved: 114 bunch offset measured.
- Reconstruction methods being developed:
 - Clustering hits from the RPCs.
 - Reconstructing clusters as multiple tracks.
 - Vertexing studies are ongoing.
- Significant contribution from Cambridge students: T. Adolphus, P. Collins, Y. Wan, J. Dej.



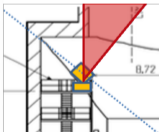
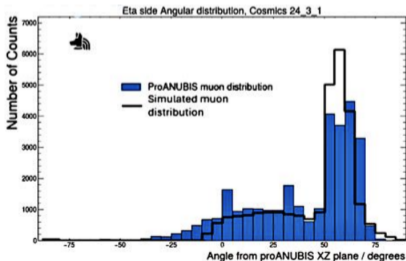
A proANUBIS event



- Lots of ongoing work in analysing the proANUBIS data.
 - Initial studies performed to align data hits e.g. Time walk corrections.
 - Working on alignment with ATLAS data.
 - Saw visual correlation initially.
 - Matching on a Bunch-Crossing level achieved: 114 bunch offset measured.
- Reconstruction methods being developed:
 - Clustering hits from the RPCs.
 - Reconstructing clusters as multiple tracks.
 - Vertexing studies are ongoing.
- Significant contribution from Cambridge students: T. Adolphus, P. Collins, Y. Wan, J. Dej.



- Also measured the efficiency of our RPCs in situ.
 - Nominally operates at 5.7 kV.
 - Get high efficiency of each individual RPC.
 - Triplet configuration has 99% efficiency.
- Performed a preliminary angular distribution analysis.
 - Want to disambiguate between up- and down-going tracks to reject Cosmics.
 - Simulated Cosmic distribution through sandstone and compared to proANUBIS measurements.
 - Peak aligns with Cosmics from the PX14 shaft.



Summary & Timeline



- ANUBIS is a proposed transverse LLP detector.
 - Sub-detector of ATLAS allowing for additional information and active vetoing.
- Instrumented with RPC detectors covering $\sim 9,800 \text{ m}^2$ & $\sim 2.1 \text{ sr}$.
- Demonstrator in place has taken $\sim 80 \text{ fb}^{-1} pp$ collision data.
 - Many ongoing studies to aid future full detector design and analysis.
- Sensitivity studies also ongoing, focusing initially on PBC benchmarks.
 - Developing modular framework: SET-ANUBIS.
- Welcoming new collaborators: [ANUBIS twiki](#).

▲ 2033+: FCC detector construction & exploitation

▲ 2035+: Run 5 full ANUBIS+ATLAS data taking

2033+: bulk ANUBIS deployment in cavern (LS4)

2030+: Run 4 partial ANUBIS data taking

2028+: partial ANUBIS deployment in cavern (LS3)

2026+: ANUBIS detector R&D (electronics, R/O) engineering for cavern deployment

2025: proANUBIS data analysis, Letter of Intent

2024: PBC model #7 (#8, #9), proANUBIS data taking

2023: finalise geometry, PBC model #6, proANUBIS

2022: seed funding for proANUBIS

2021: ANUBIS location & prototype conception

2020: proANUBIS sensitivity studies

2019: ANUBIS conception

Backup



Co-Spokesperson (theory)
Martin Bauer

martin.m.bauer@durham.ac.uk



Co-Spokesperson (expt)
Oleg Brandt

oleg.brandt@cern.ch



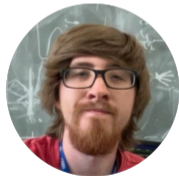
Outreach coordinator
Louie Corpe

l.corpe@cern.ch



Physics Coordinator
Michael Revering

M.Revering@cern.ch



Simulation Coordinator
Paul Swallow

paul.nathaniel.swallow@cern.ch



Technical Coordinator
Aashaq Shah

aashaq.shah@cern.ch



R&D Co-coordinator
Giulio Aielli

giulio.aielli@cern.ch



R&D Co-coordinator
Luca Pizzimento

luca.pizzimento@cern.ch

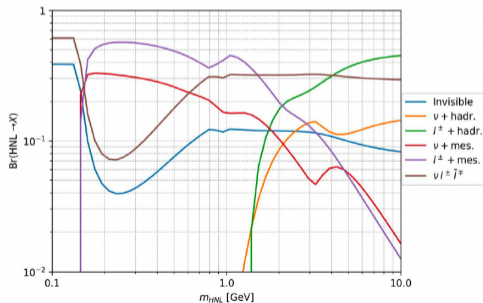


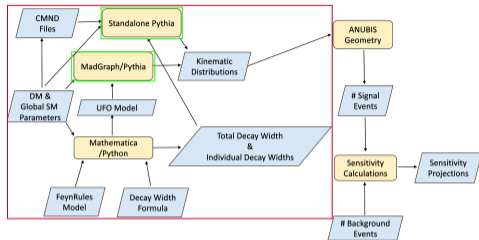
Decay mode of heavy neutrino
$N_4 \rightarrow \nu_{\ell_1} \nu_{\ell_2} \bar{\nu}_{\ell_2}$
$N_4 \rightarrow \nu_{\ell} e^- e^+$
$N_4 \rightarrow e^- \mu^+ \nu_m + c.c.$
$N_4 \rightarrow \mu^- e^+ \nu_e + c.c.$
$N_4 \rightarrow \nu_{\ell} \pi^0$
$N_4 \rightarrow e^- \pi^+ + c.c.$
$N_4 \rightarrow \nu_{\ell} \mu^- \mu^+$
$N_4 \rightarrow \mu^- \pi^+ + c.c.$
$N_4 \rightarrow e^- K^+ + c.c.$
$N_4 \rightarrow \nu_{\ell} \eta$
$N_4 \rightarrow \mu^- K^+ + c.c.$
$N_4 \rightarrow \nu_{\ell} \rho^0$
$N_4 \rightarrow e^- \rho^+ + c.c.$
$N_4 \rightarrow \nu_{\ell} \omega$
$N_4 \rightarrow \mu^- \rho^+ + c.c.$
$N_4 \rightarrow e^- K^{*+} + c.c.$
$N_4 \rightarrow \nu_{\ell} K^{*0}$
$N_4 \rightarrow \nu_{\ell} \bar{K}^{*0}$
$N_4 \rightarrow \nu_{\ell} \eta'$
$N_4 \rightarrow \mu^- K^{*+} + c.c.$
$N_4 \rightarrow \nu_{\ell} \phi$
$N_4 \rightarrow e^- \tau^+ \nu_{\tau} + c.c.$
$N_4 \rightarrow \tau^- e^+ \nu_e + c.c.$
$N_4 \rightarrow e^- D^+ + c.c.$

- Assumed a **minimal**, majorana HNL model (type I seesaw).
- Only introducing **one** HNL (PBC model BC6):
 - Currently only couple HNLs to ν_e 's.
- ANUBIS should see any charged final state:
 - $N \rightarrow e^{\pm} qq'$; $N \rightarrow \nu_e qq'$; $N \rightarrow e^+ e^- \nu_e$
 - But **not** $N \rightarrow \nu_e \nu \bar{\nu}$

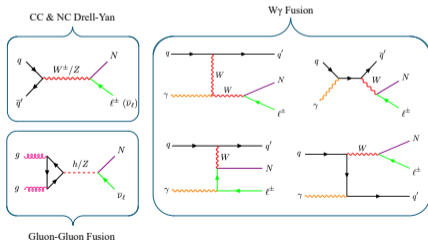
- For $m_{HNL} > m_{\pi}$ decays into mesons possible.
- More mesons are kinematically accessible for larger masses [1].

Majorana HNL, electron couplings only





- Use MATHEMATICA to derive decay widths [1,2].
- SET uses (for HNLs):
 - MADGRAPH: **HIB** production
 - PYTHIA: **Meson** production & HNL decays.
- Produce N_{gen} events in 12 combinations of:
 - Production modes (CCDY, NCDY, ggF, $W\gamma$).
 - Decay modes ($e^\pm qq'$, $\nu_e qq'$, $e^+ e^- \nu_e$).
- For 11 HNL masses $\in [0.5, 1.5]$ GeV.
 - Later: Expand to $\in [0.01, 10]$ GeV.
- Then apply low-level selections.
 - Representing BG removal.



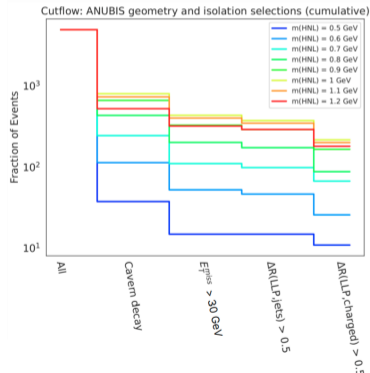
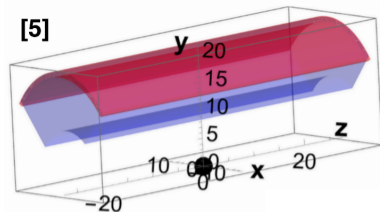
Exploring HNLs at ANUBIS: Selection

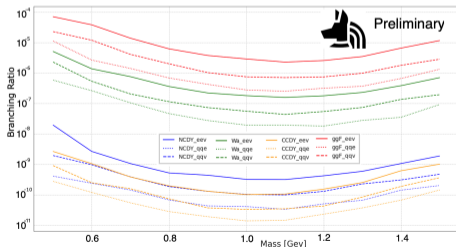
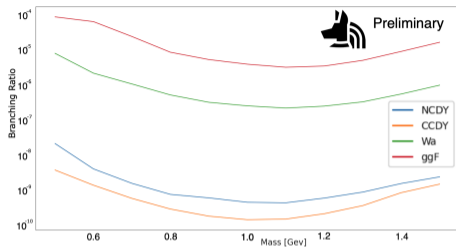
Expected Backgrounds [1]:

- **Neutral SM LLPs:** *e.g.* n/K decaying/scattering in cavern.
- **'Punch-through' jets/muons** that escape the ATLAS detector.
- **Cosmics:** reduced by rock shielding (~ 8 Hz).

Background Removal:

- Reduce most BGs by using ATLAS as an active veto and selections.
- ANUBIS geometric acceptance requirement.
- ATLAS isolation requirements (*e.g.* [2,3]) should eliminate collision BGs:
 - $E_T^{\text{miss}} > 30$ GeV; $\Delta R(\text{LLP}, \text{jets}) > 0.5$; $\Delta R(\text{LLP}, \text{charged}) > 0.5$
 - Isolation requirements effective at removing hadronic collision BGs.
 - Some discussion on older selection in [4].



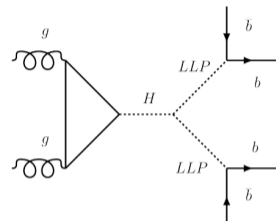


- Preliminary sensitivity limits:
 - (Top) Inclusive of decay mode (e.g. $e^\pm qq'$).
 - (Bottom) In each of the 12 combinations of production and decay.
- Also able to recast into:
 - \mathcal{B} vs $c\tau$.
 - $|U_{eN}|^2$ vs mass.
 - (See future paper.)
- The latter allows direct comparison to previous estimates [1-5].

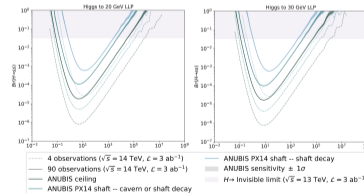
Future plans



- Finalise studies for HNLs:
 - Extend the mass range.
 - Include meson production.
 - Paper planned for the end of summer.
- Expand the framework to include additional models:
 - e.g. Dark Photon, Dark Scalar, Axions etc.
 - Framework is modular \rightarrow easy to add models.
 - Feel free to recommend models!
- Develop more detailed background estimation.
 - Developing GeoModel of (pro)ANUBIS: incorporated into ATLAS geometry.
 - proANUBIS data analysis – direct measurements of potential BGs.



(a) ggF



PBC model BC5.