Eiichiro Komatsu (Max Planck Institute for Astrophysics) *EuCAPT Colloquium, November 12, 2024*

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MAX-PLANCK-INSTITUT FÜR ASTROPHYSIK

Parity Violation in Cosmology *Does the Universe distinguish between left and right?*

Overarching Theme *Let's find new physics!*

• The current cosmological model (*ΛCDM*) **requires** new physics beyond the

- standard model of elementary particles and fields.
	- What is dark matter (*CDM*)?
	- What is dark energy (*Λ*)?

Overarching Theme There are many ideas, but how can we make progress?

- The current cosmological model (*ΛCDM*) **requires** new physics beyond the standard model of elementary particles and fields.
	- What is dark matter (*CDM*)? => CDM, WDM, FDM, ...
	- What is dark energy (Λ)? => Dynamical field, modified gravity, quantum gravity, …

New in cosmology! **Violation of parity symmetry** may hold the answer to these fundamental questions.

Available also at nature > nature reviews physics > review articles > article arXiv:2202.13919

Review Article | Published: 18 May 2022 New physics from the polarized light of the cosmic microwave background *Key Words:*

Eiichiro Komatsu

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> 1. Cosmic Microwave Background (CMB) **Polarization** 3. **Parity Symmetry**

1. Parity

Probing Parity Symmetry Definition

- **• Parity transformation = Inversion of all spatial coordinates**
	- (x, y, z) –> $(-x, -y, -z)$

- Parity symmetry in physics states:
	-

• The laws of physics are invariant under inversion of all spatial coordinates.

• Violation of parity symmetry = The laws of physics are **not** invariant under…

But, who cares about coordinates? The key is the coordinate *transformation*

• You may say, "*Coordinates are just a convenient mathematical tool. Physics*

- *should not depend on how we chart the world with coordinates.*"
	- Yes, that is absolutely correct.
- contains useful information.

• Coordinate **transformations** are different. The underlying physical principle does not depend on the choice of coordinates. However, "**how a physical system appears to change from one coordinate system to another**" often

Continuous Coordinate Transformation - 1 Spatial translation and homogeneity

• This is evidence for **invariance under spatial translation**. We shift spatial coordinates by a constant vector $c, x \rightarrow x + c$, and the physics relevant to the

• We do an experiment in Sendai, and repeat it in Munich. We find the same

- answer (to within the uncertainty).
- experiment does not change.
	- There is no special location in space => **homogeneity**.
	- This even implies that the total momentum is conserved!
		- Noether's theorem

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Continuous Coordinate Transformation - 2 Spatial rotation and isotropy

• This is evidence for **invariance under spatial rotation**. We rotate spatial coordinates by *x* –> R*x*, where R is a 3-dimensional rotation matrix, and the

- apparatus at different angles. We find the same answer (to within the uncertainty).
- physics relevant to the experiment does not change.
	- There is no special direction in space => **isotropy**.
	- This even implies that the total angular momentum is conserved!
		- Noether's theorem

• We do an experiment. We repeat it a few times after rotating the experimental

https://mathshistory.st-andrews.ac.uk

Parity: Discrete Coordinate Transformation

• We ask, "*When we observe a certain phenomenon in nature, do we also*

• (*) "Mirror image" is an ambiguous word. A parity transformation is $(x, y, z) \rightarrow$ $(-x, -y, -z)$, whereas a "mirror image" often refers to, e.g., $(x, y, z) \rightarrow (-x, y, z)$,

observe its mirror image() with equal probability?*"

where only one of (x,y,z) is flipped.

Do we also observe this with equal probability?

Note that this is not full parity transformation, as only one axis is flipped.

Parity and Rotation

- Parity transformation $(x \rightarrow x)$ and 3d rotation $(x \rightarrow Rx)$ are different.
	- R is a continuous transformation and the determinant of R is det(R) $= +1$.
	- Parity is a discrete transformation and the **determinant is –1**, as

 $\pmb{\mathcal{X}}$ \boldsymbol{y} $\tilde{}$

\boldsymbol{z} **Parity = Mirror + 2d Rotation**• One may think of parity transformation as a mirror in one of the coordinates (e.g., $z \rightarrow -z$) and **2d** rotation by π in the others.

-
- Let's demonstrate it!

Rotation In x-y

2. Pseudovector, Pseudoscalar

Parity Transformation: Vector E.g., momentum, electric field

- **p** is the same vector, written using two different basis vectors.
-

$$
\hat{e}_y' + p_z' \hat{e}_z'
$$

$$
p_y' \hat{e}_y - p_z' \hat{e}_z
$$

• Therefore, **p**'s components are transformed as $(p'_x, p'_y, p'_z) = (-p_x, -p_y, -p_z)$

Parity Transformation: Pseudovector E.g., angular momentum, magnetic field

- - change sign. Thus, their products do not change, e.g.,

 $= (-Y)(-p_z) - (-Z)$

 $L'_x = Y'p'_z - Z'p'_y$

 $= L_x$

• Orbital angular momentum, $L = r \times p$, is a *pseudovector*. Its *components* do not change under parity transformation: $(L'_x, L'_y, L'_z) = (L_x, L_y, L_z)$

• Both $\mathbf{r} = (X, Y, Z)$ and $\mathbf{p} = (p_x, p_y, p_z)$ are vectors whose components

Parity Transformation: Pseudoscalar How to test parity symmetry?

- A dot product of a vector and a pseudovector is a **pseudoscalar**.
	- Like a scalar, a pseudoscalar is invariant under rotation.
	- But, a pseudoscalar changes sign under parity transformation.
- **Experimental test of parity symmetry**: Construct a pseudoscalar and see if the average value is zero. If not, the system violates parity symmetry!
	- *Example*: a dot product of particle A's momentum and particle B's angular momentum: $\mathbf{p}_\mathrm{A} \cdot \mathbf{L}_\mathrm{B}$. Measure this and average over many trials. Does the average vanish, $\langle \mathbf{p}_A \cdot \mathbf{L}_B \rangle = 0$?

3. Discovery of Parity Violation in β-decay (weak interaction)

Physical Review, 105, 1413 (1957) $\widehat{\blacktriangleright}$ 50 $\overline{}$ Letters the Editor, S 5 etters to
I Review, \overline{c} NOVCIO

Experimental Test of Parity Conservation in Beta Decay*

C. S. Wu, Columbia University, New York, New York

AND

E. AMBLER, R. W. HAYWARD, D. D. HOPPES, AND R. P. HUDSON, National Bureau of Standards, Washington, D. C. (Received January 15, 1957)

 \mathbf{T} N a recent paper¹ on the question of parity in weak I interactions, Lee and Yang critically surveyed the experimental information concerning this question and reached the conclusion that there is no existing evidence either to support or to refute parity conservation in weak interactions. They proposed a number of experiments on beta decays and hyperon and meson decays which would provide the necessary evidence for parity conservation or nonconservation. In beta decay, one could measure the angular distribution of the electrons coming from beta decays of polarized nuclei. If an asymmetry in the

The Wu Experiment of β-decay 60Co –> 60Ni + e- + ν^e + 2γ

• Electrons must be emitted with equal probability in all directions relative to **J**, if

• This was not observed: $\langle \mathbf{p}_\mathrm{e} \cdot \mathbf{J} \rangle \neq 0$. Parity symmetry is violated in β-decay!

- parity symmetry is respected in β-decay.
	-

Wu et al. (1957)

• Wolfgang Pauli said, "*Ich glaube aber nicht, daß der Herrgott ein schwacher Linkshänder ist*" (I do not believe that the Lord is a weak left-hander).

- To Lee and Yang's theoretical paper on parity violation in β-decay:
	-
- To Wu's discovery paper:
	- exciting. How sure is this news?)
- **left and right!**
-

Initial reaction Many physicists did not believe it initially. "This Month in Physics History", APS News, October 2022

• Wolfgang Pauli said, "*Sehr aufregend. Wie sicher ist die Nachricht?*" (Very

• This was shocking news. The weak interaction distinguishes between

• In this talk we ask, "*Does the Universe distinguish between left and right?"*

4. Parity Violation in the Cosmic Microwave Background (CMB)

Credit: WMAP Science Team

26 **Visible -> Near Infrared -> Far Infrared -> Submillimeter -> MicrowaveThe sky in various wavelengths**

Where did the CMB we see today come from?

Credit: ESA

Credit: ESA

Temperature (smoothed) + Polarisation

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 1.11177

Pseudoscalar: EB correlation

• The observed pattern of the CMB polarization can be decomposed into eigenstates of parity, called "E modes" and "B modes".

- E and B modes are transformed differently under the parity transformation. Therefore, the product of the two, **the "EB correlation", is a pseudoscalar**.
- **• The full-sky average of the EB correlation must vanish (to within the measurement uncertainty), if there is no parity violation!** 30

 $\langle T_{\boldsymbol{\ell}} E_{\boldsymbol{\ell}'}^* \rangle = \langle T_{\boldsymbol{\ell}}^* E_{\boldsymbol{\ell}'} \rangle = (2\pi)^2 \delta_D^{(2)}(\boldsymbol{\ell} - \boldsymbol{\ell}') C_{\ell}^{TE}$

 $\langle B_{\ell} B_{\ell'}^* \rangle = (2\pi)^2 \delta_D^{(2)}(\ell \langle T_{\ell} E_{\ell'}^* \rangle = \langle T_{\ell}^* E_{\ell'} \rangle = (2\pi)^2 \delta_D^{(2)}(\ell - \ell') C_{\ell}^{++}$

Lercon
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sensitive doscalars and sensitive to parity violation! The other combinations, <TB> and <EB>, are pseudoscalars and

Spherical Harmonics Decomposition *aℓmY^m ^ℓ* (*n*)̂

0.1 0.07
Angular size $0.2₁$ 01

CMB Power Spectra Progress over 30 years

- This is the typical figure seen in talks and lectures on the CMB.
	- The temperature and the E- and B-mode polarization power spectra are well measured.
- **• Parity violation appears in the TB and EB power spectra, not shown here.**

This is the EB power spectrum (WMAP+Planck) Galactic plane removed (62% of the sky) Eskilt, EK (2022)

5. Cosmic Birefringence: Rotation of the Plane of Linear Polarization

How does the EM wave of the CMB propagate?

The surface of "last scattering" by electrons

(Scattering generates *polarization*!)

Credit: WMAP Science Team

How does the EM wave of the CMB propagate?

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If the plane of linear polarization of the CMB is rotated uniformly by β, it is the strategies sign of parity violation!

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Credit: ESA

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Temperature (smoothed) + Polarisation

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 $\frac{1}{2}$

"*Cosmic Birefringence*"

E-B mixing by rotation of the plane of linear polarization Lue, Wang, Kamionkowski (1999); Feng et al. (2005, 2006)

- Observed E- and B-mode polarization, E_lº and B_lº, are related to those before rotation as
- $E^{\rm o}_\ell \pm iB^{\rm o}_\ell = (E_\ell \pm iB_\ell)e^{\pm 2i\beta}$
- which gives
- $E^{\rm o}_{\ell}=E_{\ell}\cos(2\beta)-B_{\ell}\sin(2\beta)$
- $B_{\ell}^{\circ} = E_{\ell} \sin(2\beta) + B_{\ell} \cos(2\beta)$

CMB Power Spectra

 $\frac{1}{2}$

BB.

- Rotation of the plane of linear polarization **mixes** E and B modes.
- Therefore, the EB correlation will be given by the difference between the EE and BB correlations.
- Observed EE is much greater than BB. We expect EB to look like EE!

$$
C_{\ell}^{EB,\text{o}} = \frac{\tan(4\beta)}{2} \left(C_{\ell}^{EE,\text{o}} - C_{\ell}^{BB,\text{c}} \right)
$$

Cosmic Birefringence fits well(?) $\left(C_\ell^{EE,\mathrm{o}}-C_\ell^{BB,\mathrm{o}} \right)$ **Nearly full-sky data (92% of the sky)**

Eskilt, EK (2022)

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- -

$\tan(4\beta)$ **Cosmic Birefringence fits well(?) Galactic plane removed (62% of the sky)**

Eskilt, EK (2022)

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The Biggest Problem: Miscalibration of detectors

• Is the plane of linear polarization rotated by the genuine cosmic birefringence effect, or simply because the polarization-sensitive directions of the detectors are rotated with

- respect to the sky coordinates (and we did not know it)?
-

Impact of miscalibration of polarization angles Wu et al. (2009); Miller, Shimon, Keating (2009); EK et al. (2011) **Cosmic or Instrumental?**

• If the detectors are rotated by α, it seems that we can measure only the **sum α+β**.

(but we do not know it)

The past measurements The quoted uncertainties are all statistical only (68%CL)

- $\alpha + \beta = -6.0 \pm 4.0$ deg (Feng et al. 2006)
- $\alpha + \beta = -1.1 \pm 1.4$ deg (WMAP Collaboration, Komatsu et al. 2009; 2011)
- $\alpha + \beta = 0.55 \pm 0.82$ deg (QUaD Collaboration, Wu et al. 2009)
- …
- $\alpha + \beta = 0.31 \pm 0.05$ deg (Planck Collaboration 2016)
- $\alpha + \beta = -0.61 \pm 0.22$ deg (POLARBEAR Collaboration 2020)
- $\alpha + \beta = 0.63 \pm 0.04$ deg (SPT Collaboration, Bianchini et al. 2020)
- $\alpha + \beta = 0.12 \pm 0.06$ deg (ACT Collaboration, Namikawa et al. 2020)
- $\alpha + \beta = 0.07 \pm 0.09$ deg (ACT Collaboration, Choi et al. 2020)

first measurement

The past measurements Now including the estimated systematic errors on α • β = –6.0 ± 4.0 ± **??** deg (Feng et al. 2006)

-
- $\beta = -1.1 \pm 1.4 \pm 1.5$ deg (WMAP Collaboration, Komatsu et al. 2009; 2011)
- $\beta = 0.55 \pm 0.82 \pm 0.5$ deg (QUaD Collaboration, Wu et al. 2009)
- …
- $\beta = 0.31 \pm 0.05 \pm 0.28$ deg (Planck Collaboration 2016)
- $\beta = -0.61 \pm 0.22 \pm ?$? deg (POLARBEAR Collaboration 2020)
- $\beta = 0.63 \pm 0.04 \pm$ **??** deg (SPT Collaboration, Bianchini et al. 2020)
- $\beta = 0.12 \pm 0.06 \pm ?$? deg (ACT Collaboration, Namikawa et al. 2020)
- $\beta = 0.07 \pm 0.09 \pm$ **??** deg (ACT Collaboration, Choi et al. 2020)

Uncertainty in the calibration of a has been the major limitation

The Key Idea: The polarized Galactic foreground emission as a calibrator

Minami et al. (2019); Minami, EK (2020)

Credit: ESA

Directions of the magnetic field inferred from polarization of the thermal dust emission in the Milky Way

Emitted "right there" - it would not be affected by the cosmic birefringence.

Polarized dust emission within our Milky Way!

ESA's Planck

- The angles are all over the place, and are well within the quoted calibration uncertainty of instruments.
	- 1.5 deg for WMAP
	- 1 deg for Planck
- They cancel!
	- The power of adding independent datasets.

Miscalibration angles (WMAP and Planck) Nearly full-sky data (92% of the sky) Minami, EK (2020); Diego-Palazuelos et al. (2022); Eskilt, EK (2022)

LFI HFI WMAP Ø

-
-
-

-
-
-
- $0.33^{\circ} \pm 0.10^{\circ}$ 353
	- Light traveling in a uniform magnetic field also experiences a rotation of the plane of linear polarization, called "**Faraday rotation**". However, the rotation angle depends on the frequency, $as \ \beta(\nu) \propto \nu^{-2}.$
	- No evidence for frequency dependence is found!
		- For $\beta \propto \nu^n$, $n = -0.20^{+0.41}_{-0.39}$ (68% CL)
		- **Faraday rotation** $(n = -2)$ **is disfavoured.**

No frequency dependence is found It is not due to Faraday rotation. Eskilt (2022); Eskilt, EK (2022)

-0.39

Is *β* **caused by non-cosmological effects? We need to measure it in independent experiments.** Diego-Palazuelos et al. (2022, 2023); Eskilt et al., arXiv:2305.02268

- The **known** instrumental effects of the WMAP and Planck missions are shown to have negligible effects on *β*.
	- However, we can never rule out **unknown** instrumental effects… We need to measure *β* in independent experiments.
- The polarized Galactic foreground emission was used to calibrate the instrumental polarization angles, *α*. The intrinsic EB correlations of the Galactic foreground emission (**polarized dust and synchrotron emission**) could affect the results.
	- We need to measure *β* without relying on the foreground by calibrating *α* well, e.g., Murata et al. (Simons Observatory), arXiv:2309.02035; Murphy et al. (ACT), arXiv:2403.00763; Cornelison et al. (BICEP3), arXiv:2410.12089; Ritacco et al. (COSMOCal), arXiv:2405.12135.

Implications DM = Dark Matter; DE = Dark Energy

$$
I = \int d^4x \sqrt{-g} \left[-\frac{1}{2} (\partial \chi)^2 - V(\chi) - \frac{1}{4} F^2 - \frac{\alpha}{4f} \chi F \right]
$$

EK, Nature Rev. Phys. **4**, 452 (2022)

This term exists for a pion. What if DM/DE is "pion-like particle"

• This **rotates** the plane of linear polarization of light by *β*

Implications DM = Dark Matter; DE = Dark Energy

$$
I = \int d^4x \sqrt{-g} \left[-\frac{1}{2} (\partial \chi)^2 - V(\chi) - \frac{1}{4} F^2 - \frac{\alpha}{4f} \chi F \right]
$$

• The measured angle, β, implies that the field has evolved by

$$
\Delta \chi = \chi(\tau_{\rm obs}) - \chi(\tau_{\rm em})
$$

- If it is due to DE: this measurement rules out DE being a cosmological constant.
- If it is due to DM: at least a fraction of DM violates parity symmetry.

EK, Nature Rev. Phys. **4**, 452 (2022)

This term exists for a pion. What if DM/DE is "pion-like particle"

$$
\simeq \frac{10^{-2}}{\alpha} f
$$

- Violation of parity symmetry is a new topic in cosmology.
- It may hold the answers to fundamental questions, such as
	- *• What is Dark Matter?*
	- *• What is Dark Energy?*
- Since parity is violated in the weak interaction, it seems natural to expect that **parity is also violated in the Dark Sector**.
	- **3.6σ hint of Cosmic Birefringence**: Space may be filled with parityviolating DM and DE fields?
- **• What else should we be looking? New and exciting research topics.**

Summary *Let's find new physics!*

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