

# Effects of Reactor Antineutrino Anomaly on Measurement of Theta-13

Animesh Chatterjee, Luis Serra Diaz-Cano,  
Aaron Osborn  
INSS Group Project  
July 29, 2011

# Overview

- Reactor experiments have been vital to the study of neutrino ( $\nu$ ) oscillations.
- Observed antineutrino rates have typically been lower than those expected from  $\nu$ -spectrum calculations
- Recent improvements to  $\nu$ -spectrum parameters result in an increase in the *expected*  $\nu$  rate
- However, this increase in the expected rate makes the ratio  $N_{\text{observed}}/N_{\text{predicted}}$  even lower
- This increased  $\nu$ -deficit is called the reactor antineutrino anomaly
- This anomaly changes the results of prior experiments and must be considered for current/future research

# Calculation of Neutrino Spectra

- Precise calculation of neutrino spectra is very complicated and uses *many* parameters

$$\Phi_{\nu}(E, t) = \frac{P_{th}(t)}{\sum_k \alpha_k(t) \times E_k} \times \sum_k \alpha_k(t) \times S_k(E)$$

- E is the neutrino energy
- $\phi_{\nu}(E,t)$  is reactor neutrino flux
- $P_{th}(t)$  is reactor thermal power at time t
- $\alpha_k(t)$  is fraction of all fissions at time t due to isotope k, where k is a major fuel isotope:  $^{235}\text{U}$ ,  $^{238}\text{U}$ ,  $^{239}\text{Pu}$ , or  $^{241}\text{Pu}$
- $E_k$  is energy per fission of isotope k
- $S_k(E)$  is neutrino spectrum per fission of isotope k
- The  $S_k(E)$  term is broken down even further...

# Calculation of Neutrino Spectra

- Expansion of  $S_k(E)$ :

$$S_k(E) = \sum_{fp}^{N_{fp}} A(T) \times \left( \sum_{b=1}^{N_b} BR_{fp}^b \times S_{fp}^b(Z, A, E_{0,fp}^b, E) \right)$$

- $A(T)$  is decay probability for fission product fp of isotope k
- $BR_{fp}^b$  is the branching ratio of decay possibility b of fission product fp of isotope k
- $S_{fp}^b(Z, A, E_{0,fp}^b, E)$  is the neutrino spectrum per decay for decay possibility b of fission product fp of isotope k
- Z and A are proton and mass number of nucleus of fp, respectively
- $E_{0,fp}^b$  is the end-point energy of decay along branch b
- And this is broken down even further...

# Calculation of Neutrino Spectra

- $S_{fp}^b(Z, A, E_{0,fp}^b, E)$  consists of the following:

$$S_{fp}^b(Z, A, E_{0,fp}^b, E) = K_{fp}^b \times \mathfrak{F}(Z, A, E) \times pE(E - E_{0,fp}^b)^2 \times \\ \times C_{fp}^b \times (1 + \delta_{fp}^b(Z_{fp}, A_{fp}, E))$$

- $K_{fp}^b$  is normalization factor
- $F(Z, A, E)$  is the Fermi function
- $pE(E - E_{0,fp}^b)^2$  is a phase-space factor
- $C_{fp}^b$  is the shape function
- the last term contains corrections,  $\delta_{fp}^b$  consists of:
  - $\delta_{\text{recoil}}$
  - $\delta_{\text{weak magnetism}} = E \times A_W = E \times (0.47\% / \text{MeV})$
  - $\delta_{\text{Coulomb}} = E \times A_C = E \times (-10Z\alpha R / 9\hbar c)$ ,  $R \approx (1.5 \text{ fm}) \times A^{1/3}$
- The recent improvement to these calculations then affects the calculation of cross-sections:

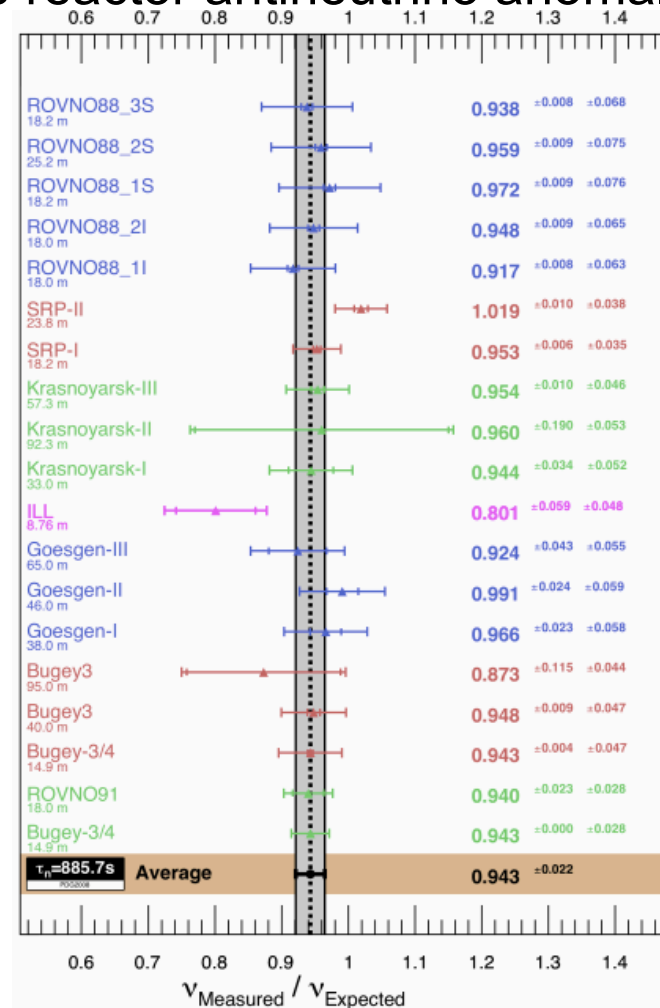
$$\sigma_f^{\text{predicted}} = \int_0^\infty S_{\text{tot}}(E_\nu) \sigma_{V-A}(E_\nu) dE_\nu = \sum_k f_k \sigma_{f,k}^{\text{predicted}}$$

# Enhancement of $\nu$ -spectra

- Previous  $\nu$ -spectrum calculations used  $\sim 30$  effective  $\beta$ -branches
- These calculations have recently been refined greatly, now including  $\sim 800$  daughter nuclei of fuel isotopes and  $\sim 10,000$   $\beta$ -branches!
- Some systematic errors have been fixed as well
  - for low ( $< 4$  MeV) energies, new  $\delta_{\text{weak magnetism}}$  and  $\delta_{\text{Coulomb}}$  terms deviate from the linear correction used in older experiments
  - for high ( $> 4$  MeV) energies, previous calculations ignored dispersion of  $Z$ , using a mean value instead
- Using the newer, larger set of parameters:
  - for  $^{235}\text{U}$ ,  $\sigma_{f,k}^{\text{predicted,new}} = 1.025\sigma_{f,k}^{\text{predicted,old}}$
  - for  $^{239}\text{Pu}$ ,  $\sigma_{f,k}^{\text{predicted,new}} = 1.031\sigma_{f,k}^{\text{predicted,old}}$
  - for  $^{241}\text{Pu}$ ,  $\sigma_{f,k}^{\text{predicted,new}} = 1.037\sigma_{f,k}^{\text{predicted,old}}$
  - for  $^{238}\text{U}$ ,  $\sigma_{f,k}^{\text{predicted,new}} = 1.098\sigma_{f,k}^{\text{predicted,old}}$
- Average increase of 3.5%
- Increased cross-section *increases* the number of neutrinos an experiment expects to observe

# Reactor Antineutrino Anomaly

- This increase in  $N_{\text{exp}}$  leads to a *decrease* in the ratio  $N_{\text{obs}}/N_{\text{exp}}$
- Previous average was  $0.976 \pm 0.024$ , decreased to  $0.943 \pm 0.013$
- This effect is the reactor antineutrino anomaly



# Effect on Upcoming Experiments


- Evidence of this reactor anti- $\nu$  anomaly results in having to make a choice of the correct cross-section per fission to use for constraints
- Use of  $\sigma_{f,k}^{\text{predicted,new}}$ , if it is incorrect, will lead to overestimation of  $\theta_{13}$ , thereby “faking” a discovery
- However, if the 3 + sterile hypothesis is correct, use of the new cross-section is still possible
- If average reactor experiments are correct, it is possible to use the anomalous cross-section, an average from all short-baseline reactor experiments
- Using  $\sigma_f^{\text{ano}}$  can absorb either a small physical  $\nu$ -defecit or a small miscalculation of reactor fluxes, but will lead to a conservative constraint on  $\theta_{13}$
- This choice only affects Double Chooz during far-only phase, as a multi-detector configuration absorbs uncertainties in reactor antineutrino fluxes
- In the 3 + sterile case, upcoming experimental sensitivities will be unaffected due to large value of  $\Delta m_{\text{new}}^2$
- Oscillations into  $\nu_{\text{sterile}}$  are an energy-independent suppression of anti- $\nu_e$  rate by a factor  $(1/2)\sin^2(2\theta_{\text{new}})$ , as first approximation



# Impact on CHOOZ

Because  $\sigma^{\text{old}} \sim \sigma^{\text{bugey}}$

	old [3]	new
$\sigma_{f, 235U}^{\text{pred}}$	$6.39 \pm 1.9\%$	$6.61 \pm 2.11\%$
$\sigma_{f, 239Pu}^{\text{pred}}$	$4.19 \pm 2.4\%$	$4.34 \pm 2.45\%$
$\sigma_{f, 238U}^{\text{pred}}$	$9.21 \pm 10\%$	$10.10 \pm 8.15\%$
$\sigma_{f, 241Pu}^{\text{pred}}$	$5.73 \pm 2.1\%$	$5.97 \pm 2.15\%$



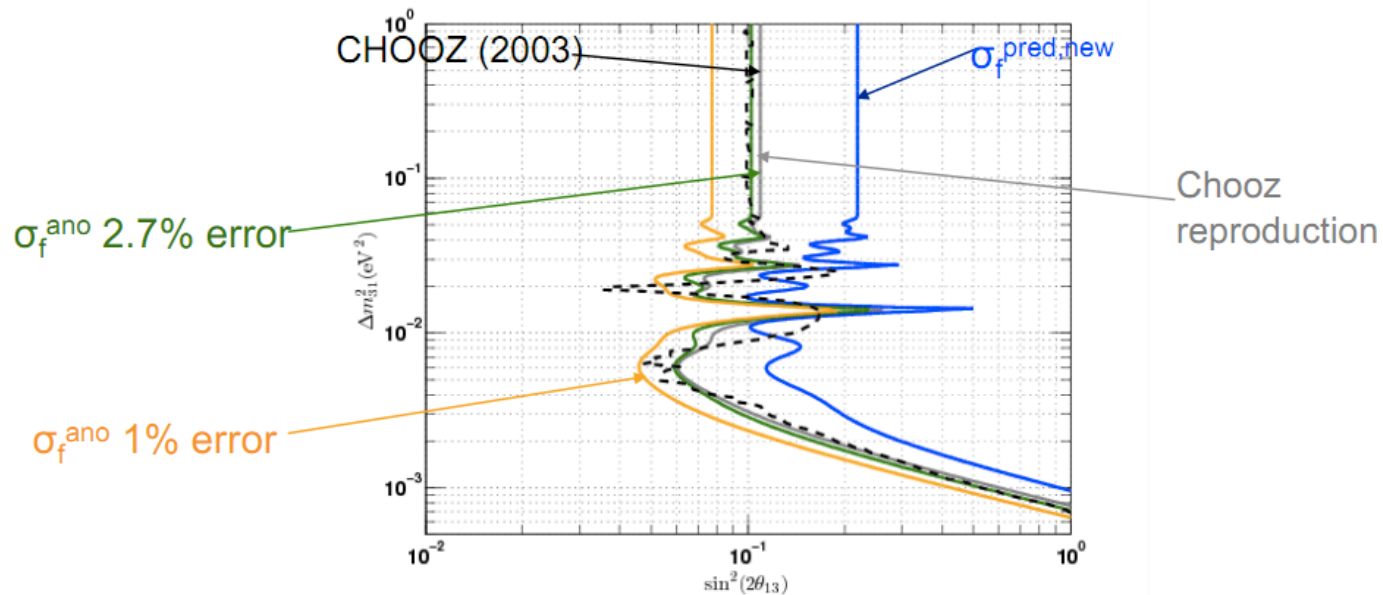
Now  $\sigma^{\text{new}} = 1.05 \sigma^{\text{old}}$ , now CHOOZ cannot use Bugey data

Revisited results:  $\Delta m_{13}^2 = 2.4 \cdot 10^{-3} \text{ eV}^2$

- Old result of CHOOZ
- Same analysis, new cross section
- Deficit of  $\nu_e$  due to  $\theta_{13}$
- 4<sup>th</sup> family of neutrinos. Sterile

$\sin^2(2\theta_{13}) < 0.14$  (1dof)  
 $\sin^2(2\theta_{13}) < 0.22$  (1dof)  
 $\sin^2(2\theta_{13}) < 0.11$  (1dof)  
 $\sin^2(2\theta_{13}) < 0.10$  (1dof)

# Impact on CHOOZ



They couldn't reproduce the results from 2003 -> Different statistics.

The result of  $\theta_{13}$  with this new  $\sigma$  doesn't change dramatically, but  $\theta_{13}$  compatible with 0 now

$$\Delta m_{13}^2 = 2.4 \cdot 10^{-3} \text{ eV}^2$$

# Arguments for re-evaluation

- Re-evaluation of cross-section used .
- Increase the systematic error.
- Check effects with near and far detector :

$$\text{near/far} \equiv (N_{\text{near}}/N_{\text{near}}^0)/(N_{\text{far}}/N_{\text{far}}^0)$$

$$\text{so } \frac{\text{near}}{\text{far}} \sim \frac{\overline{P}_{ee}(\text{near})}{\overline{P}_{ee}(\text{far})}$$

this will give effects of  $\Theta_{13}$  and  $\Theta_{14}$  with the different detector.

- Forth-neutrino hypothesis

# 3+1 neutrino Hypothesis

- This may give some information about new non-standard neutrino, called sterile neutrino.
- The motivation is the explanation of the antineutrino deficit by an oscillation of electron neutrinos into a new neutrinos state with a large  $\Delta m^2_{\text{new}}$  &  $\sin^2(2\theta_{\text{new}})$
- *4<sup>th</sup> neutrino Hypothesis* :

We are adding to the 3 active neutrinos a 4<sup>th</sup> sterile state and hence faced with four neutrino mass eigenstate. Then electron neutrino(or anti-neutrino) survival probability as

$$P_{ee} = \left| \sum_{i=1,2,3,4} |U_{ei}|^2 \exp\left(i \frac{\Delta m_{i1}^2 L}{2E_\nu}\right) \right|^2 ,$$
$$= 1 - \sum_{i < j} 4 |U_{ei}|^2 |U_{ej}|^2 \sin^2 \left( \frac{\Delta m_{ij}^2 L}{4E_\nu} \right)$$

*arXiv:0809.5076v1[hep-ph]*

# Cntd..

- where  $L$  is the propagation distance(baseline) and  $E_\nu$  is the neutrino energy.
- For reactor neutrino experiment and in the limit of  $L < 2 \text{ km}$  &  $E_\nu > 2 \text{ MeV}$
- Oscillation Probability (neglecting solar oscillation)

$$P_{ee} = 1 - \cos^4 \theta_{\text{new}} \sin^2(2\theta_{13}) \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E_{\bar{\nu}_e}} \right) - \sin^2(2\theta_{\text{new}}) \sin^2 \left( \frac{\Delta m_{\text{new}}^2 L}{4E_{\bar{\nu}_e}} \right).$$

- So by 1<sup>st</sup> approximation could be seen as a suppression of the electron anti-neutrino rate by  $0.5 * \sin^2(2\theta_{\text{new}})$
- This will give information about new mixing angle and  $\Delta m_{\text{new}}^2$ .
- Thus 3% shift of the reactor neutrino flux can give some information about sterile neutrino.