# **Development of FORTVNE: Forward TeV Neutrino Detector**

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# **Background: FASER Run 4 Upgrade**

- FASER and FASERv have proved the ability to discover and measure TeV neutrinos from the LHC
- FASER has been approved for Run 4
  - an upgraded detector will collect 680 fb<sup>-1</sup>, and quite possibly up to 3 ab<sup>-1</sup> if run through the end of HL-LHC
- Emulsion is expensive, cannot envision having emulsion in place for all 680 fb<sup>-1</sup> of Run 4
- If we replace FASERv for all or most of Run 4, we should replace it with something that continues/enhances the neutrino program
- We have 1 or 2 boxes of tungsten that can be reused. Each box is ~1 tonne, consisting of 730 1 mm x 25 cm x 30 cm plates
- UCI group is pursuing a US NSF Major Research Infrastructure program
  - ~\$20 set aside to fund construction projects up to \$4M over 5 years (2025-30)
  - Each major University may submit 1 proposal, and we have been selected as UCI's proposal this year

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Leading Pls at UCI: Jonathan Feng Jianming Bian **David Casper** Michael Smy



### FORTVNE

- Tungsten/Scintillator bars
- 99 cm x 25 cm x 60 cm
  - 2 FASERv stacked up, with emulsion replaced by scintillator bars
- 66 repeating modules, each is
  - 5 mm tungsten ( $1.5 X_0$ )
  - 5 mm horizontal scint. bars (YZ view)
  - 5 mm vertical scint. bars (XZ view)
  - For each scint. bar: 0.5 cm x 1 cm
- Total mass: 0.954 tonnes
- Extended coverage upwards:  $\eta > 7.6$ (FASERv:  $\eta > 8.5$ , FASER:  $\eta > 9.2$ )
- No CE required, no disruption to FASER spectrometer, re-uses 660 tungsten plates







# **Neutrino Physics**

- Segmentation is too coarse for  $v_{\tau}$  physics. Would have preferred SciFi, 0.2 mmdiameter fibers. However, it would exceed \$4M budget. It would be good to re-visit this if other funding comes through
- Still there is significant physics case from both ve and  $v_{\mu}$  neutrinos at TeV energies, and possibly some BSM searches (thanks to Felix, Max, Roshan, Toni)



FIG. 1. Neutrino yields and cross sections at the proposed detector. The expected precision of measurements of neutrino interaction cross sections (top, statistical errors only) and the number of neutrinos interacting in the detector consisting of  $0.25 \text{ m} \times 0.6 \text{ m} \times 0.3 \text{ m}$  of tungsten (bottom) as a function of energy for electron (left) and muon (right) neutrinos. Existing data from accelerator experiments [1], IceCube [2], and the recent FASER $\nu$  result [3] are also shown, together with the prospects for SHiP.





## **Proton Structure and Astroparticle Physics**

- FORTVNE extends to  $\eta > 7.6$ , overlaps SND's "off-axis" coverage
- puzzle), oscillations to sterile neutrinos or other states



#### Electron neutrino energy spectrum normalization and shape shed light on proton structure (intrinsic charm, low-x gluon pdf), astroparticle physics (cosmic muon





# **Detector Design Consideration**

- Primary goal is v<sub>e</sub> CC measurement: reasonable longitudinal granularity of the scintillating bars is necessary to capture electron shower development. Tungsten thickness should be 1-2X<sub>0</sub> at most. Tungsten  $X_0 = 3.5 \text{ mm}$
- Sufficient tungsten to ensure statistics: 1 tonnes
- Minimize modifications due to the budget cap
  - limit the detector length to 1 meter, reuse existing 25 cm x 30 cm tungsten plates, stack up to increase total mass
- 5 mm thickness of tungsten (~1.5 X<sub>0</sub>) for electron identification
  - Further thinning would require more scintillator channels, which is not feasible with budget constraint
- Use SiPM readout without cooling





### **Scintillator bars**

#### Baby-MIND style scintillator bar and SiPM system, proposed for FLArE HadCal



- BC408: a general-purpose plastic scintillator widely used high-energy physics experiments
- Geometry: 0.5 cm x 1 cm x 25 cm (60 cm) for horizontal (vertical)
- Each scintillator bar has two SIPMs coupled to it, one at each end, to enable photon detection in both lowenergy and high-energy ranges
- Wrapped with a 100-200 µm Tyvek reflector to improve light collection. The reflector will also limit optical crosstalk between scintillator bars
- Two options for coupling with the SiPM: embedding a WLS fiber in a groove and directing the fiber end with the SiPM, or directly attach the SiPM to the end of the scintillator bars

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#### **Photosensors - SiPM**



- Hamamatsu MPPC 13360-3025CS/PE SiPMs
  - 14,400 pixels within a 3 mm x 3 mm active area, PDE 40% at 450 nm
- Fiber connectors are glued to the ends of the bars to align the fiber ends with the SiPMs
- Each SiPM will be mounted on a mini PCB board to facilitate connections to the electronics via coaxial cables
- Two SiPMs for each scintillator bar, one with a neutral density optical filter to reduce light yield for high-energy measurements, and the other without a filter for low-energy measurements
  - High dynamic range: from MIP tracks to high-energy-density EM showers



BabyMIND Scintillator bar + SiPM





#### **Readout Electronics**







CITIROC (Cherenkov Imaging Telescope Integrated Readout Chip) used in the Baby-MIND detector of the WAGASCI experiment and the SuperFGD detector of the T2K experiment, 32 channels, 3 chips per FEB

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Baby MIND electronics readout, developed by UniGe https://doi.org/10.7566/JPSCP.27.011011



#### Light Yield



- ulletin the SiPMs to these deposits
- ulletper channel, based on BabyMIND results.
- ulletelectronics non-linearity, and background noise.

The detector simulation involves first simulating energy deposits in the scintillators and then the digitized response

In the simulation, the light yield of a single scintillator bar is set to 30 p.e./MIP, and the dark noise is set to 2 p.e.

Other effects considered include detector cell non-uniformity, light attenuation, SiPM saturation, readout

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## Simulation: Event Displays, Y-Z view



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## Simulation: Event Displays, Y-Z view



#### Each type of event exhibits clear topological signatures that can be distinguished from the others

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#### ve Identification with CNN





- Train a CNN PID for v<sub>e</sub> CC Identification ullet
- With the event composition of  $v_{\mu}$ :  $v_e$ :  $v_{\tau}$  = 131:24:1
  - v<sub>e</sub> CC Eff: 86%, Pur: 83%
  - v<sub>μ</sub> CC Eff: 99%, Pur: 72%
  - v<sub>τ</sub> CC Eff: 8%, Pur: 48%

#### ve CC can be identified with good efficiency and purity

The relatively high impurity in  $v_{\mu}$  CC and low efficiency in  $v_{\tau}$  CC could be improved by combining FASER's spectrometer and calorimeter



## ve CC Calorimetric Energy Resolution



resolution as a function of the true neutrino energy.

- •
- energy resolution, ~20%
- energy of hadronic interactions in  $v_e CC$  events.

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FIG. 10: Simulated  $\nu_e$  CC energy resolution in FORTVNE. Left: The overall energy resolution. Right: Energy

Reconstructing Calorimetric Energy by adding up deposit energies in scintillator bars, no shower/track clustering Shift the peak of (RecoE–TrueE)/TrueE to 1. The RMS of the (RecoE–TrueE)/TrueE distribution is defined as the

• The long tail in the (RecoE–TrueE)/TrueE distribution is caused by the detector containment and the invisible



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## Summary

- - Identify v<sub>e</sub> CC from NC and other backgrounds
  - Measure  $v_e$  energy with a resolution of approximately ~20%
- lacksquarespectrometer for muon momentum reconstruction
- Single electron identification is understudy •
- recognition to identify  $v_{\tau}$  hadronic decay mode may have a chance
- Potential upgrades in the future  $\bullet$ 
  - Add SciFi between tungsten plates in FORTVNE to improve vertex resolution for  $v_{\tau}$
  - FORTVNE to improve hadron containment and muon momentum measurement

Tungsten target/scintillator bars combination is a well-established technology, with it FORTVNE can:

Due to smaller absorber thickness, the containment of  $v_{\mu}$  and  $v_{\tau}$  events is reduced, requiring the track

Due to reduced hadron/muon containment,  $v_{\tau}$  identification is challenging. Using CNN pattern

Physics cases: neutrino properties at TeV energies, QCD, proton structure, astroparticle connections

Replace the current Magnetized decay pipe with hadronic calorimeter/Muon Spectrometer after





#### **Backup Materials**

#### FORTVNE





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### **Supporting Structure**









#### **Search for New Particles**



recoils with energies from 300 MeV to 10 GeV.



FIG. 4: New Particle Searches at FORTVNE. The discovery reach for milli-charged particles [48] (left) and heavy neutral leptons coupling to the SM through the magnetic dipole operator (see text) [49] (right) for LHC Run 4 (680 fb<sup>-1</sup>) and the HL-LHC (3 ab<sup>-1</sup>). The contours are N = 3 signal contours, and the signal is single electron



### **Cost Estimate**

Scintillator + sipm system			
Number of x+y modules		66	
Channels per module (x + y)		85	
		cost in \$	
Scintillator material cost per channel		50	quote
Scintillator machining per channel		140	2 hou
Scintillator assembly per channel		45	3 hou
SiPM cost per channel		200	quote
Readout electronics cost per channel		100	Baby
Scintillator + sipm system cost		3001350	
Tungsten absorber		cost in USD	
Tungsten		0	from
Tungsten supporting structure		35000	50 ho
Personnel (26% overhead)	cost/year	cost in USD	
Technician	100000	500000	100k
Postdoc	80000	240000	80k/y
PhD students	70000	210000	70k/y
Total cost		3986350	

e from luxium

urs of machine shop @ 70/hour

urs of UG labor @ 15/hour

e from Hamamatsu

/MIND readout

current FASERnu

ours of machine shop @ \$70/hour

/year for 5 years, engineer design, prototyping, overlook production year for 3 years, simulation, Qa/Qc, analysis, detector assembly and commissio

year for 3 years, supervise UGs for scintllator production

oning



## **Construction/Management Plan**

- Simulation and Detector parameter optimization
- Engineering design
- Scintillator bar and SiPM procurement
- Scintillator bar machining, UCI Physical Sciences Machine Shop @ \$70/hours
- Scintillator bar + SiPM assembly @ UCI High Bay
- Readout electronics construction/procurement
- QA/QC
- Prototyping
- Beam test
- Detector assembly and commissioning

