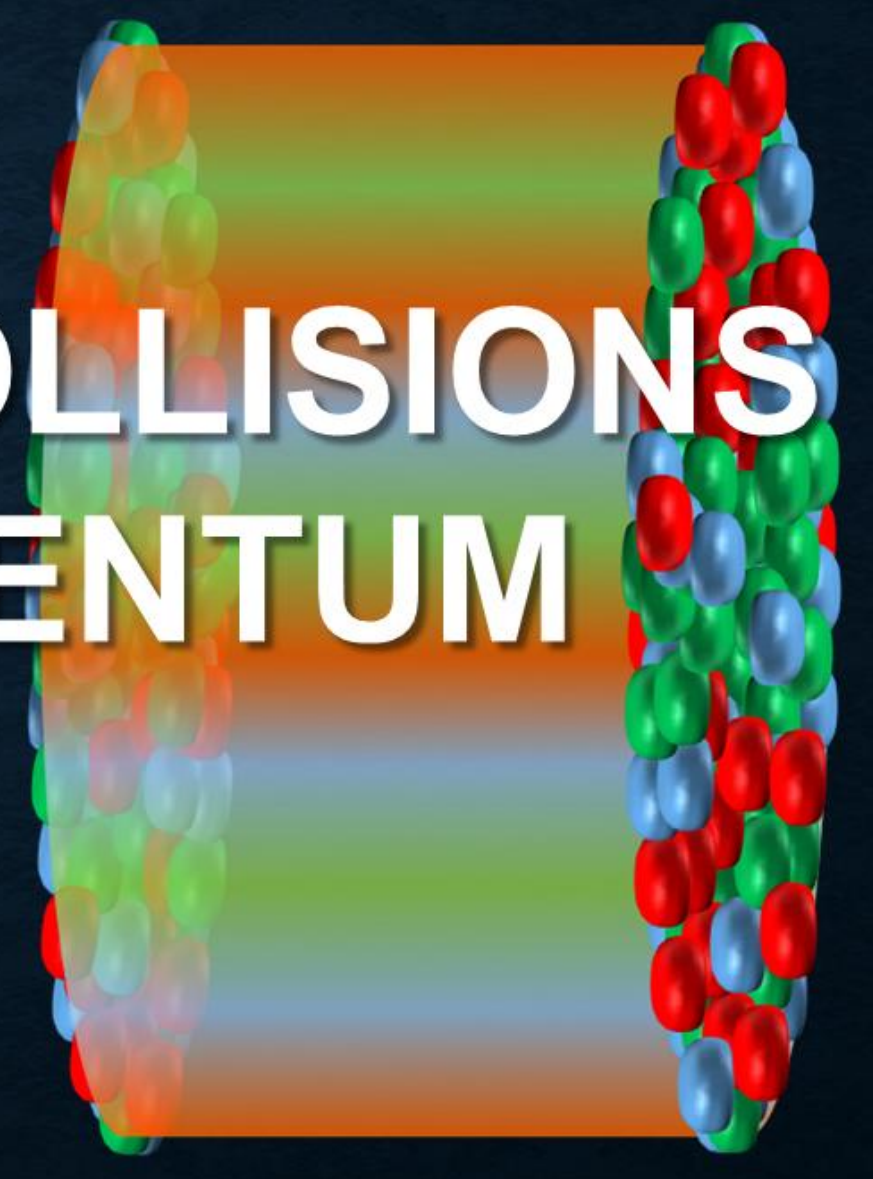


APPLICATION OF (3+1)D GLASMA SIMULATION TO AU-AU COLLISIONS AT $\sqrt{s_{NN}} = 200$ GEV: ECCENTRICITY, AND ANGULAR MOMENTUM



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<Abstract>

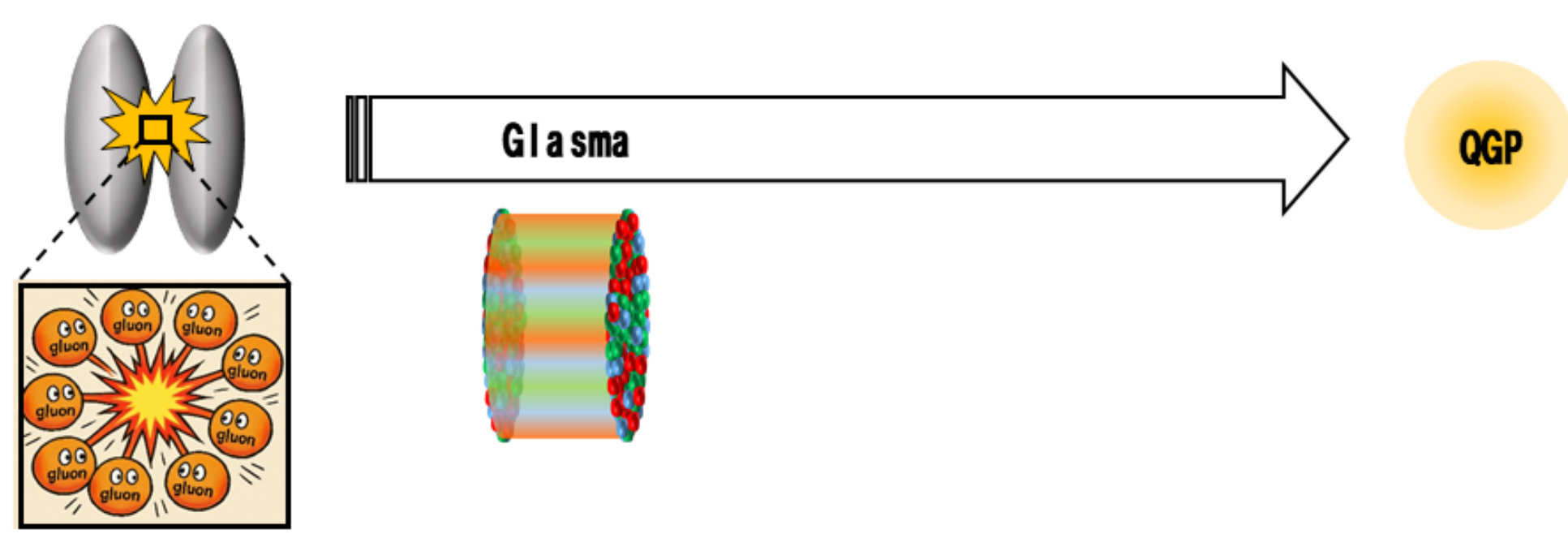
We perform a (3+1)D glasma simulation of the early stages of Au–Au collisions at $\sqrt{s_{NN}} = 200$ GeV over a wide range of impact parameters [1]. This approach goes beyond the usual boost-invariant approximation by incorporating the nuclei's longitudinal structure, including finite nucleon thickness and random positions along the beam axis. We study geometry-dependent observables such as eccentricity and angular momentum, finding that eccentricities grow with impact parameter. While angular momentum also increases, its generation at mid-rapidity is suppressed across all impact parameters, consistent with experimental observations of global polarization at $\sqrt{s_{NN}} = 200$ GeV.

[1] Phys. Rev. D **110**, 114032, (2024)

<Background>

IP-Glasma Model

B. Schenke, et al. (2012)



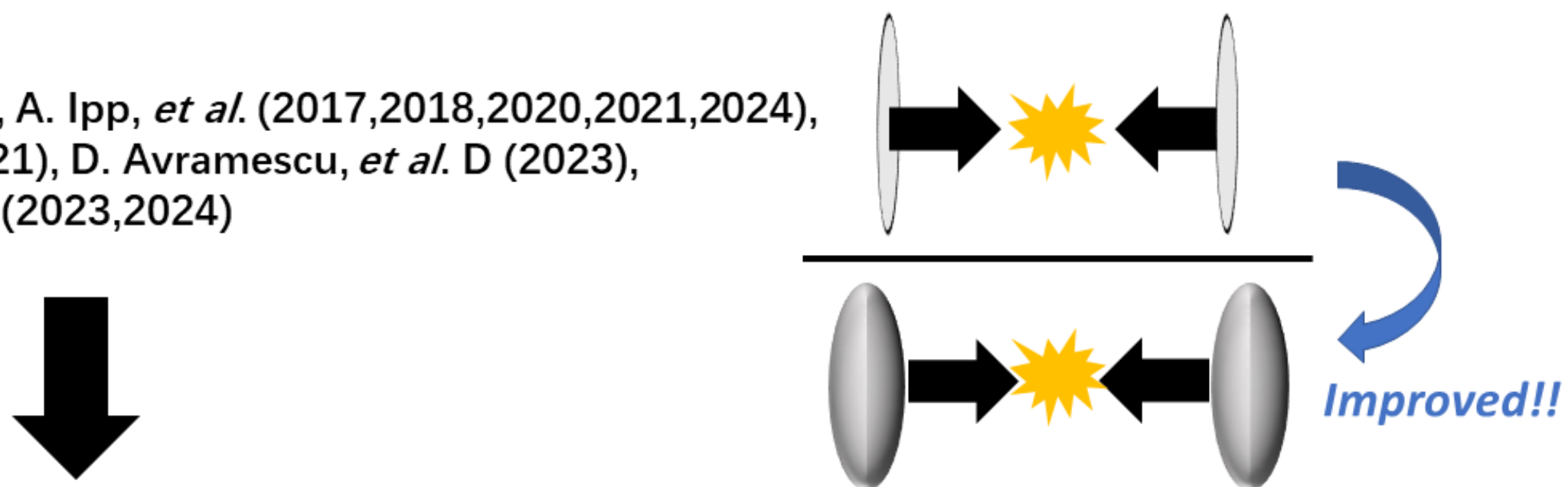
IP-Glasma Model: an "Initial State Model" focusing on early gluon dynamics

- Simulate generation and evolution of the glasma, based on CGC effective theory
- Assume physical quantities of the glasma are directly inherited by QGP, and interpret them as inputs for relativistic hydrodynamics
- Be widely used and successfully reproduce experimental results
- Neglect the longitudinal structure of the nucleus

3D Glasma: Finite Longitudinal Structure and Color Rotation

<Recent attempt>

D. Gelfand, et al. (2016), A. Ipp, et al. (2017,2018,2020,2021,2024),
S. Schlichting, et al. (2021), D. Avramescu, et al. D (2023),
H. M. and X.-G. Huang (2023,2024)



In this talk, we applied the 3D glasma simulation to Au+Au collisions and clarified the collision geometry dependence of various physical quantities

Phys. Rev. D **110**, 114032, (2024)

<Method>

Phys. Rev. D **108**, 114008, (2023)

Step 1: Set initial conditions of two nuclei on the lattice

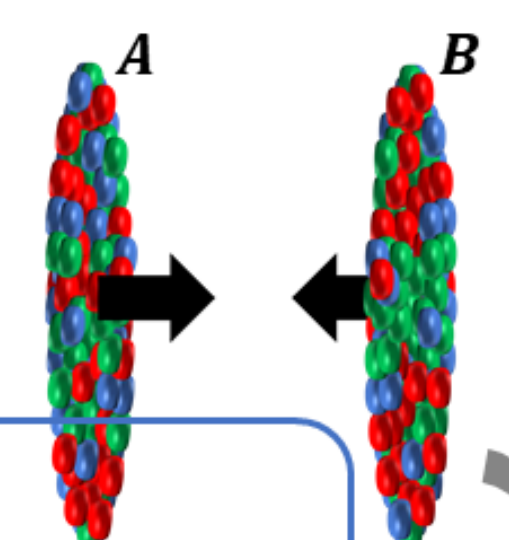
$$(\rho_{ini}^A, \rho_{ini}^B, A_{ini}, E_{ini})$$

Classical color charge density (ρ)

- Hard parton
- **Input of model**

Classical Yang-Mills field (A_μ)

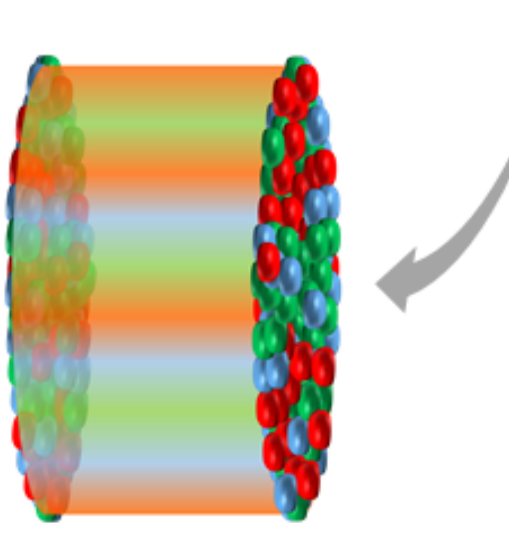
- Soft gluon
- **Determined by solving Classical E.O.M. with ρ**



Step 2: Evolve CYM fields and classical color charges numerically

$$[D_\mu, F^{\mu\nu}] = \delta^{v+} \rho^A + \delta^{v-} \rho^B$$

$$[D_\mu, \delta^\pm \rho^{A/B}] = 0$$



Step 3: Repeat steps 1 and 2, and take the event average of the observables of interest

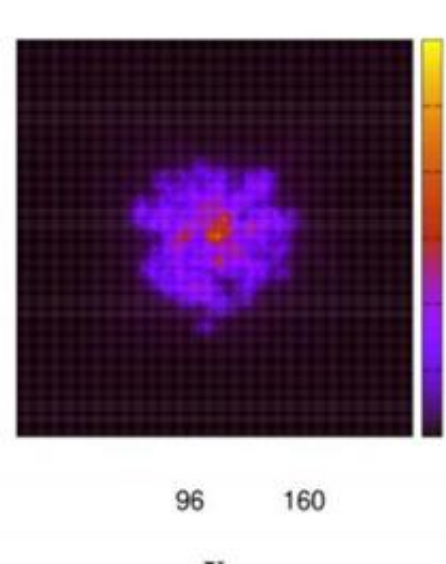
<Setup>

Gold Nucleus:

- $\rho(x)$: Incoherent sum of nucleon's color charge densities (ρ_i)
- $\rho_i(x)$: Gaussian Nucleon shape \times event-by-event random number (Γ_i)
- $\Gamma_i(x)$: Average of squared Γ_i is given by saturation scale (Q_s)
- $Q_s(x_\perp)$: Estimated by IP-sat model

Bartels, et al. (2002), H. Kowalski and D. Teaney (2003)

Simplicity: SU(2)



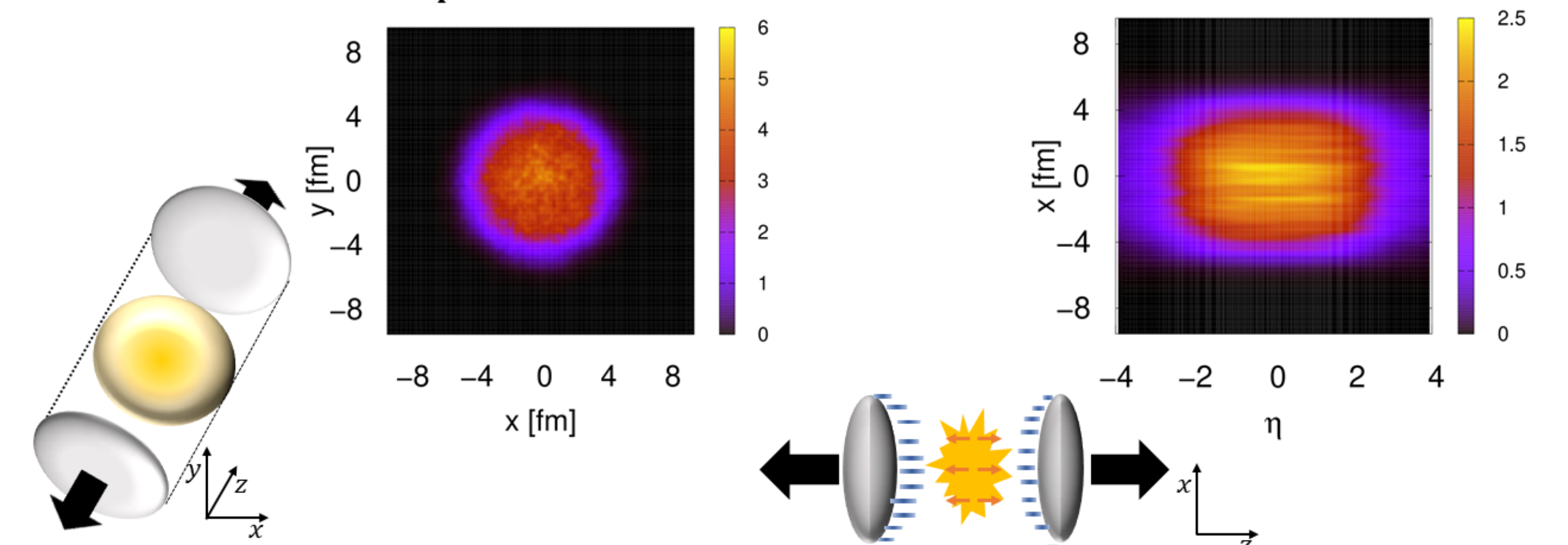
<Results>

Energy Density Profile

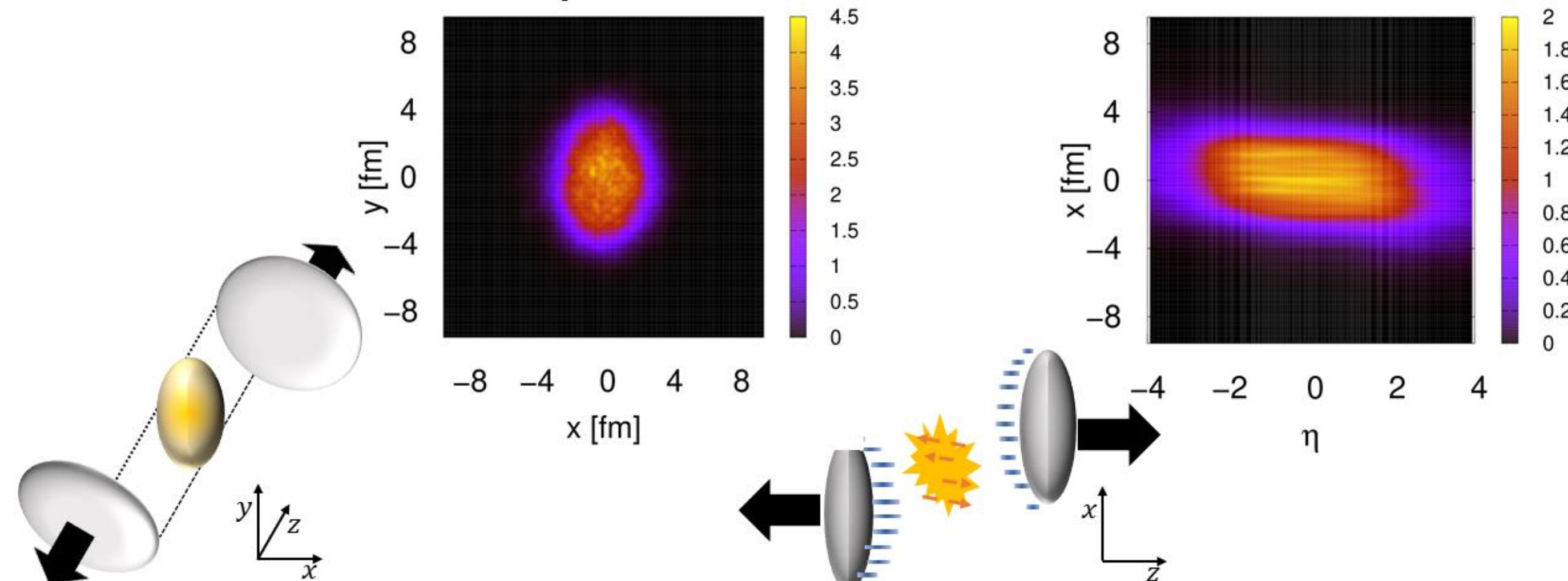
ε_{LRF} : Local rest frame energy density

- ✓ The ε_{LRF} deformation reflects the shape of the overlap region of the colliding nuclei and the difference in their thickness.

<Central Collision ($b_{imp} = 0$) at $\tau = 0$ >

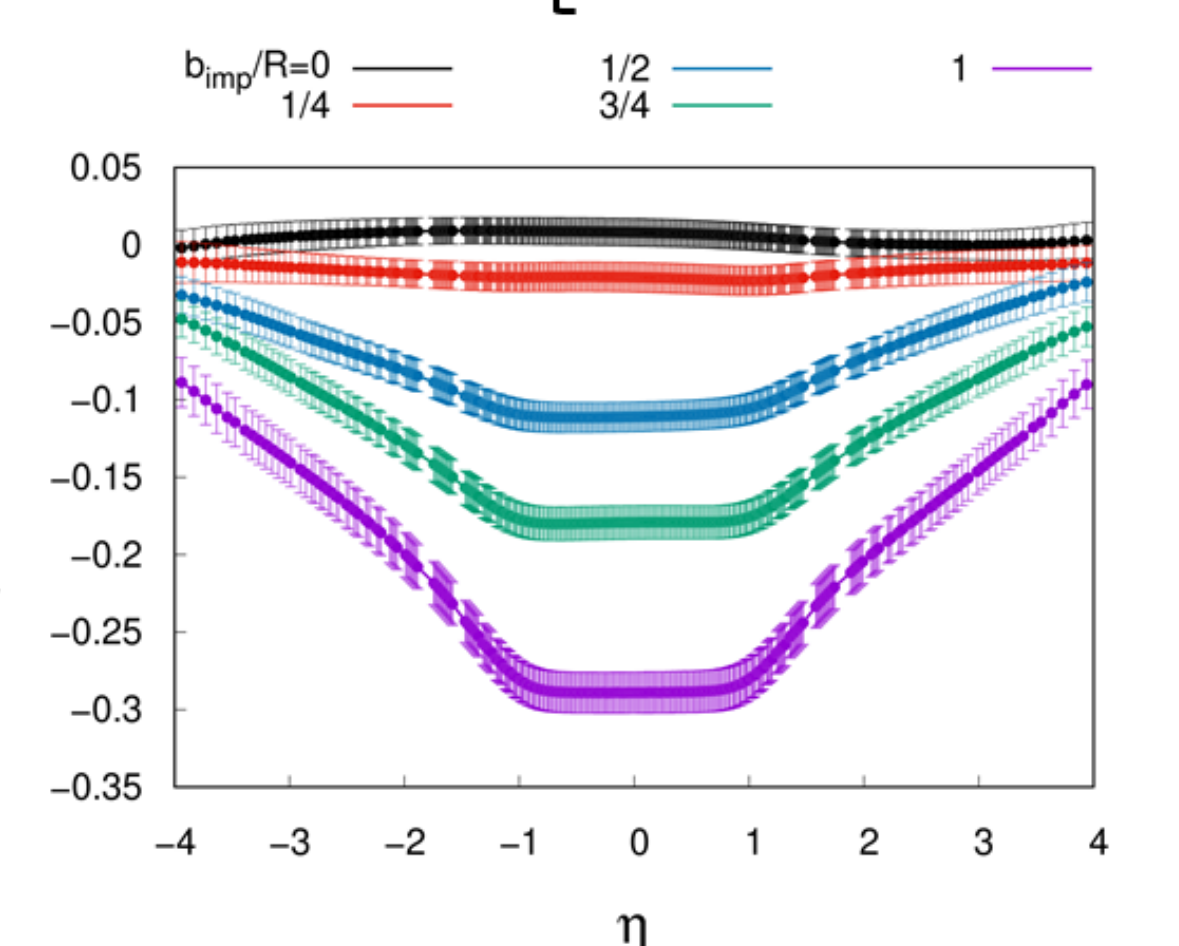


<Non-Central Collision ($b_{imp} = R$) at $\tau = 0$ >



Rapidity Profile of $\text{Re}[\varepsilon_2]$

$$\varepsilon_n = \frac{\int d^2x_\perp \varepsilon_{LRF} r_\perp^n e^{in\phi}}{\int d^2x_\perp \varepsilon_{LRF} r_\perp^n} \quad \begin{cases} \phi \equiv \arctan(x^2/x^1) \\ r_\perp \equiv \sqrt{(x^1)^2 + (x^2)^2} \end{cases}$$



- ✓ consistent with shapes of energy densities

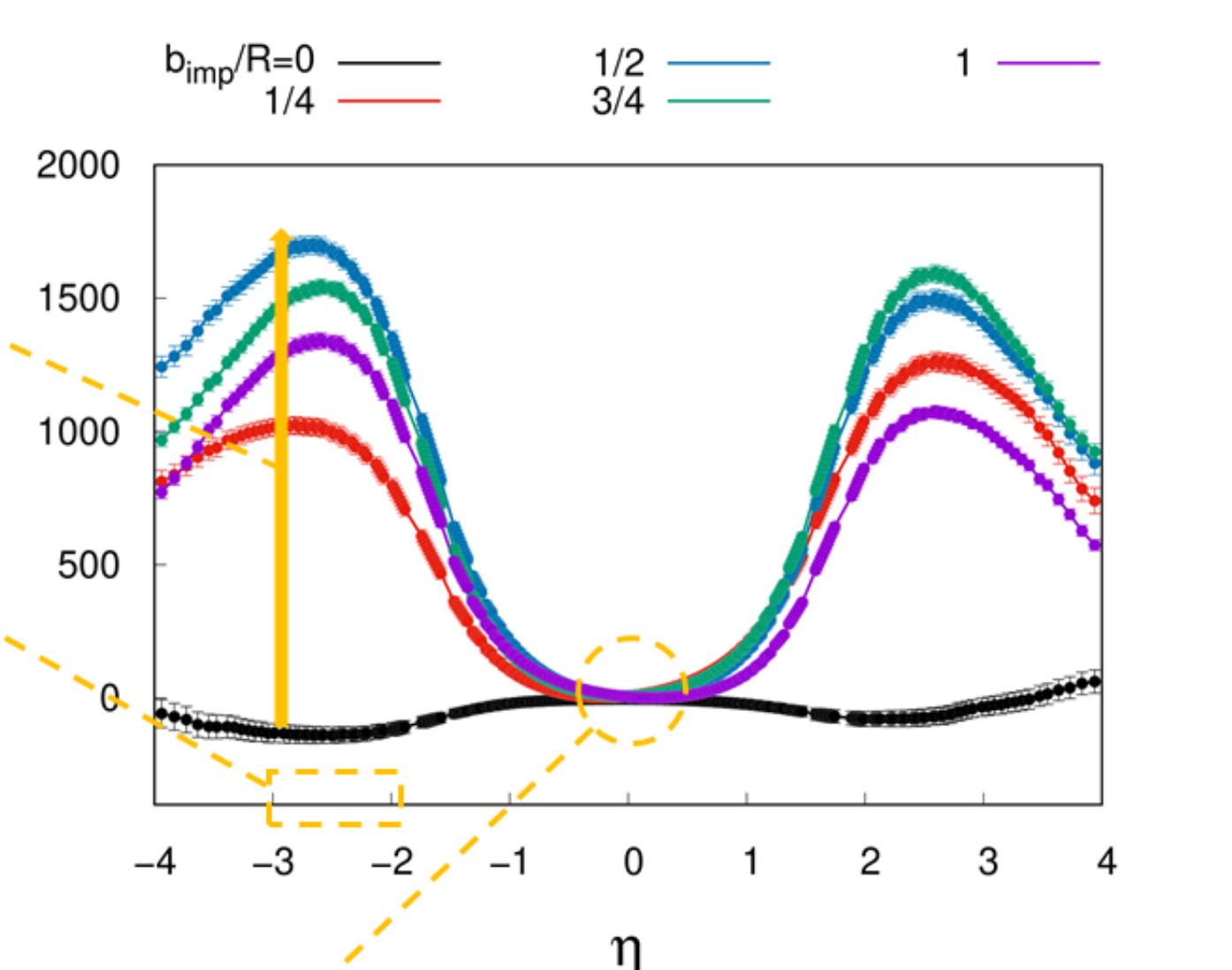
- ✓ consistent with experimental results of elliptic flow

Rapidity Profile of L^y

- ✓ Generation of L^y with increasing b_{imp}

- ✓ Emergence of a peak at $\eta = 2 - 3$

- ✓ Absence of L^y generation at $\eta = 0$



Comment on little angular momentum around midrapidity

Given that CGC descriptions are more reliable at high energies, these results imply that effects beyond the high-energy limit may be important for explaining spin polarization at mid-rapidity, which is consistent with the experimental trend that spin polarization is very small at $\sqrt{s_{NN}} = 200$ GeV, 2.76 TeV, and 5.02 TeV and increases as the beam energy decreases.

<Outlook>

Apply the 3D glasma model as the initial state for hydrodynamic simulations, and compare with experimental data.