

Chemical freeze-out of hadrons within the advanced Hadron Resonance Gas Model

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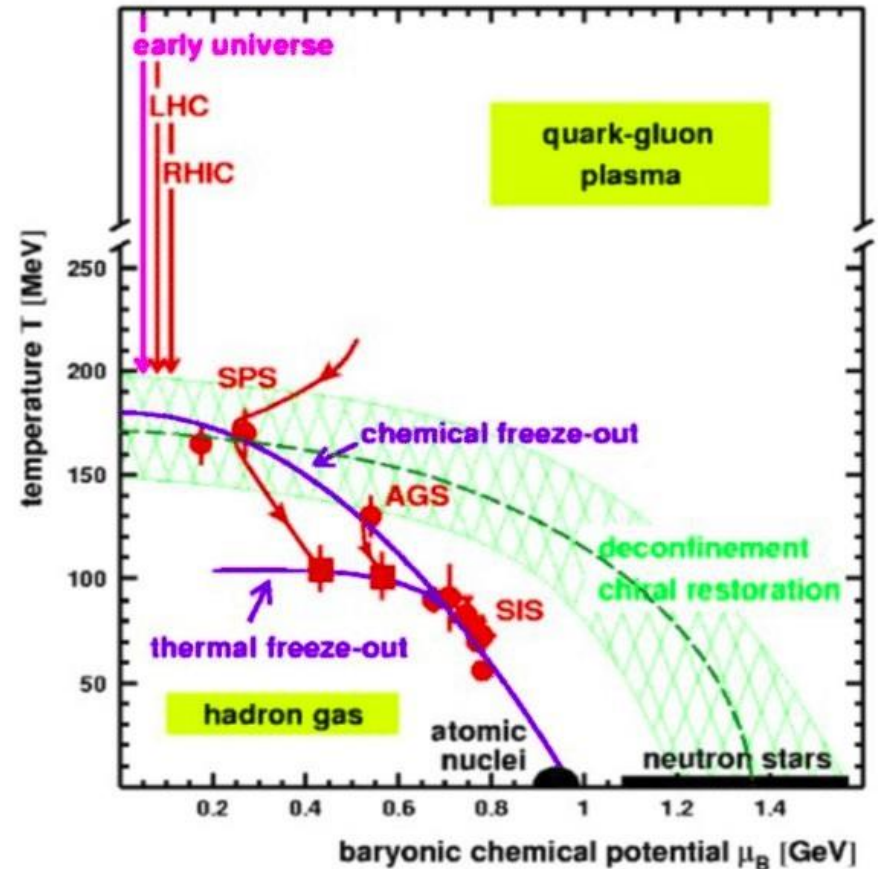


Motivation

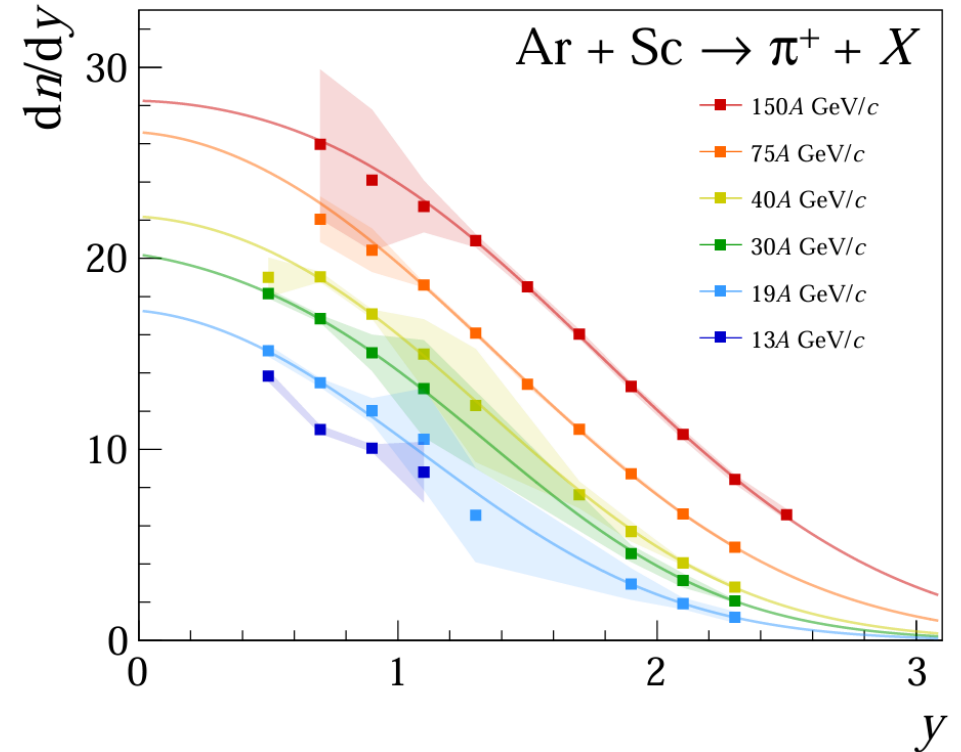
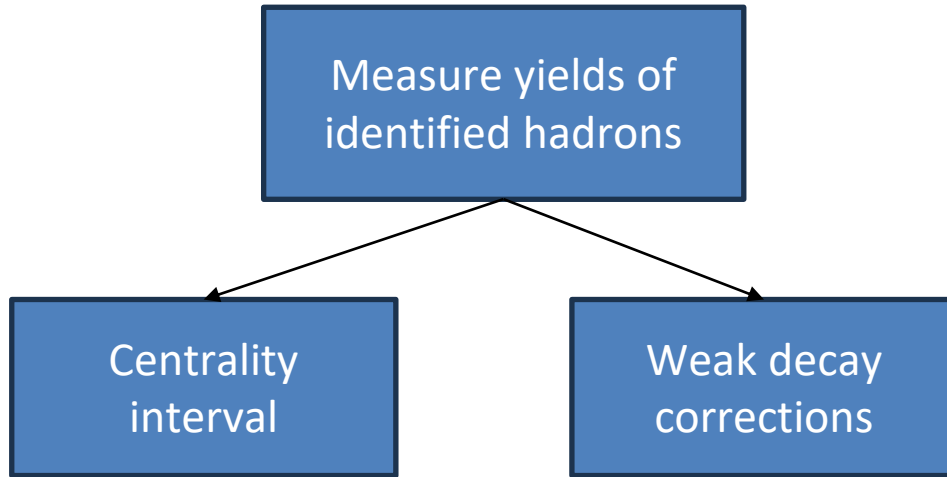
Exploring of the QCD phase diagram:

- detect signals of colour deconfinement
- detect signals of chiral symmetry restoration
- locate critical endpoint of QCD phase diagram

In order to resolve these tasks we need a good tool to analyze the data.



Motivation



- Theoretical model should make feed-down corrections consistently with experimental analysis
- Particle ratios should be taken within the same centrality interval

Induced Surface Tension EOS

System of coupled equations between the pressure p and the induced surface tension coefficient Σ :

$$\begin{cases} p = \sum_{k=1}^N p_k^{Id}(T, v_k^P) \\ \Sigma = \sum_{k=1}^N R_k p_k^{Id}(T, v_k^S) \end{cases}$$

$$n_k^{Id}(T, \mu) = \frac{g_k}{2\pi^2 \hbar^3} \int_{0_{\infty}}^{\infty} \frac{p^2 dp}{\exp[(E - \mu)/T] \pm 1}$$

$$p_k^{Id}(T, \mu) = \frac{g_k}{2\pi^2 \hbar^3} \int_0^{\infty} \frac{p^4 dp}{3E} \frac{1}{\exp[(E - \mu)/T] \pm 1}$$

Effective chemical potentials:

$$v_k^S = \mu_k - pV_k - \alpha \Sigma S_k$$

$$v_k^P = \mu_k - pV_k - \Sigma S_k$$

Induced Surface Tension EOS

Particle number density of kth sort:

$$n_k = \frac{a_{22}n_k^{Id}(T, v_k^{Id}) - a_{12}R_k n_k^{Id}(T, v_k^S)}{a_{11}a_{22} - a_{12}a_{21}}$$

$$a_{11} = 1 + \sum_{k=1}^N V_k n_k^{Id}(T, v_k^P)$$

$$a_{22} = 1 + \alpha \sum_{k=1}^N S_k R_k n_k^{Id}(T, v_k^S)$$

$$a_{12} = \sum_{k=1}^N S_k n_k^{Id}(T, v_k^P)$$

$$a_{21} = \sum_{k=1}^N V_k R_k n_k^{Id}(T, v_k^S)$$

Advantages

- 2 equations, its number does not depend on the number of different hard core radii
- Allow one to go beyond the Van der Waals approximation, since it reproduces 2nd, 3rd and 4th virial coefficients of the gas of hard spheres for $\alpha = 1.245$

[V. V. Sagun et al., EPJ Web of Conferences 137 \(2017\) 09007](#)

[K. A. Bugaev et al., Nucl. Phys. A 970 \(2018\) 133-155](#)

IST EOS settings

Experimental data: STAR Collaboration

Energies: **7.7 — 200 GeV**

Local fit parameters: **T , μ_B , μ_{I3} , μ_S , γ_S**

Global fit parameters: **R_π , R_K , R_{mesons} , R_{baryons} , R_Λ**

Global parameters were fixed as:

$$R_\pi = 0.15 \text{ fm},$$

$$R_K = 0.395 \text{ fm},$$

$$R_{\text{mesons}} = 0.42 \text{ fm},$$

$$R_{\text{baryons}} = 0.365 \text{ fm},$$

$$R_\Lambda = 0.085 \text{ fm}$$

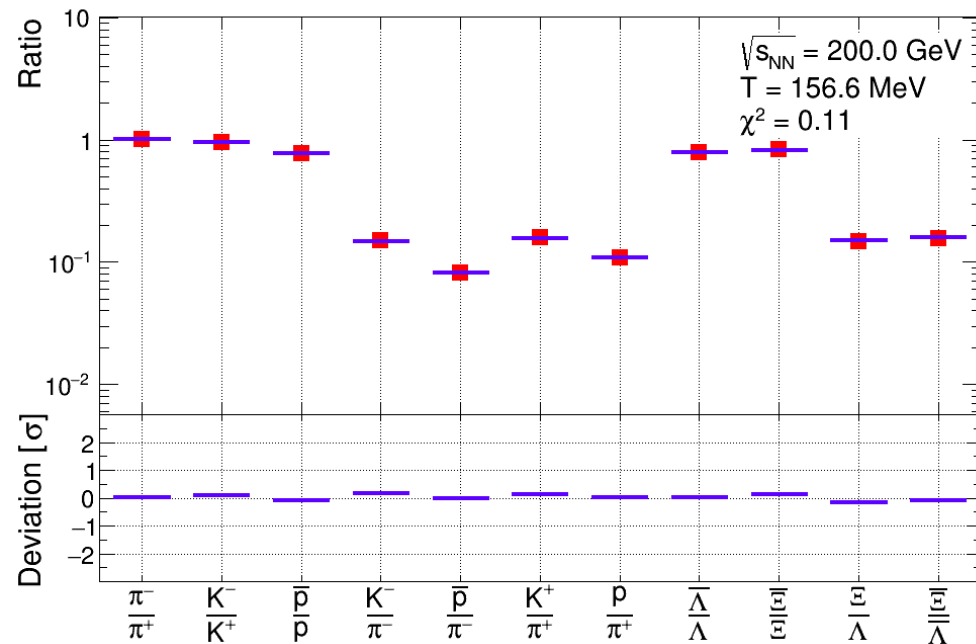
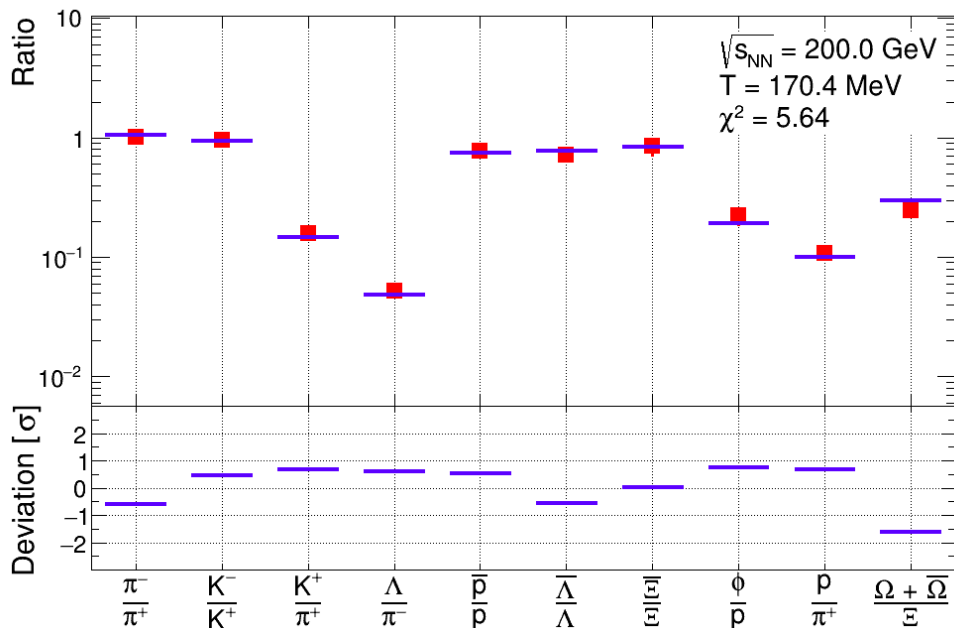
[A. Andronic et al., Nucl. Phys. A 834 \(2010\) 237c](#)

[K. A. Bugaev et al., Ukr. J. Phys. 61 \(2016\) no. 8, 659](#)

[J. Cleymans et al., Phys. Rev. C 73 \(2006\) 034905](#)

[S. Borsanyi et al., Phys. Rev. Lett. 125 \(2020\) 052001](#)

IST EOS fit result of STAR data at 200 GeV



- Inclusion of weak decays **greatly** improves the description of particle ratios in the experimental data (**50 times better** in this case).

Inclusion of weak decays for STAR data in the IST EOS

Fit by STAR Collaboration for $\sqrt{s} > 200$ GeV has $T \sim 154$ MeV

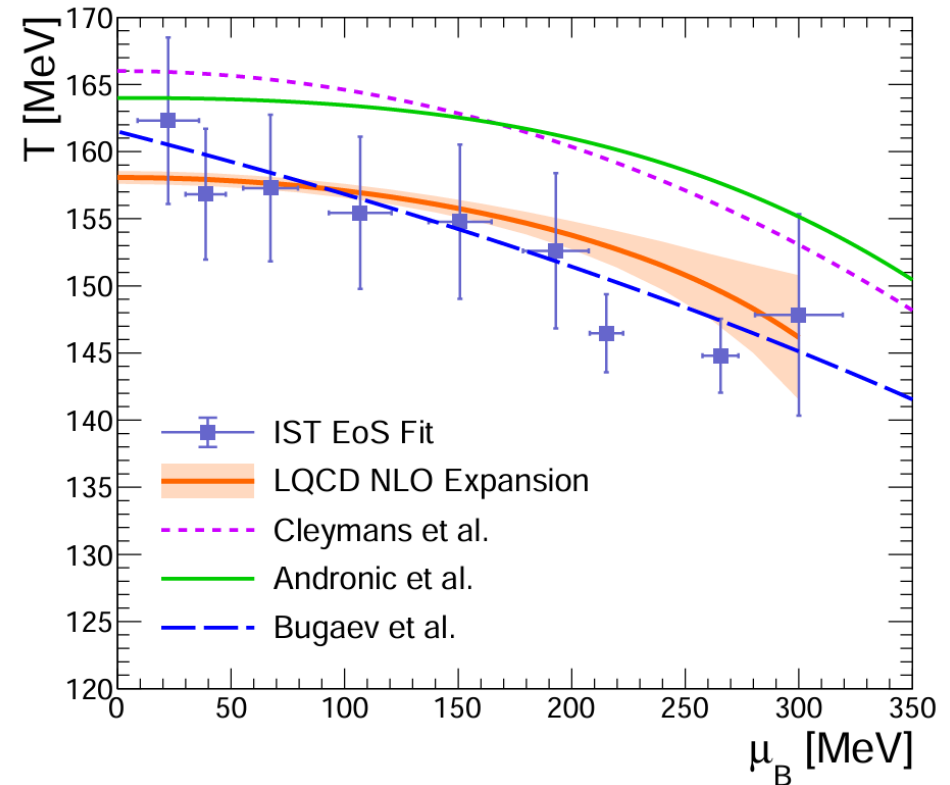
Inclusion of weak decays **decrease** temperature of chemical freeze-out **on 10 MeV**

[A. Andronic et al., Nucl. Phys. A 834 \(2010\) 237c](#)

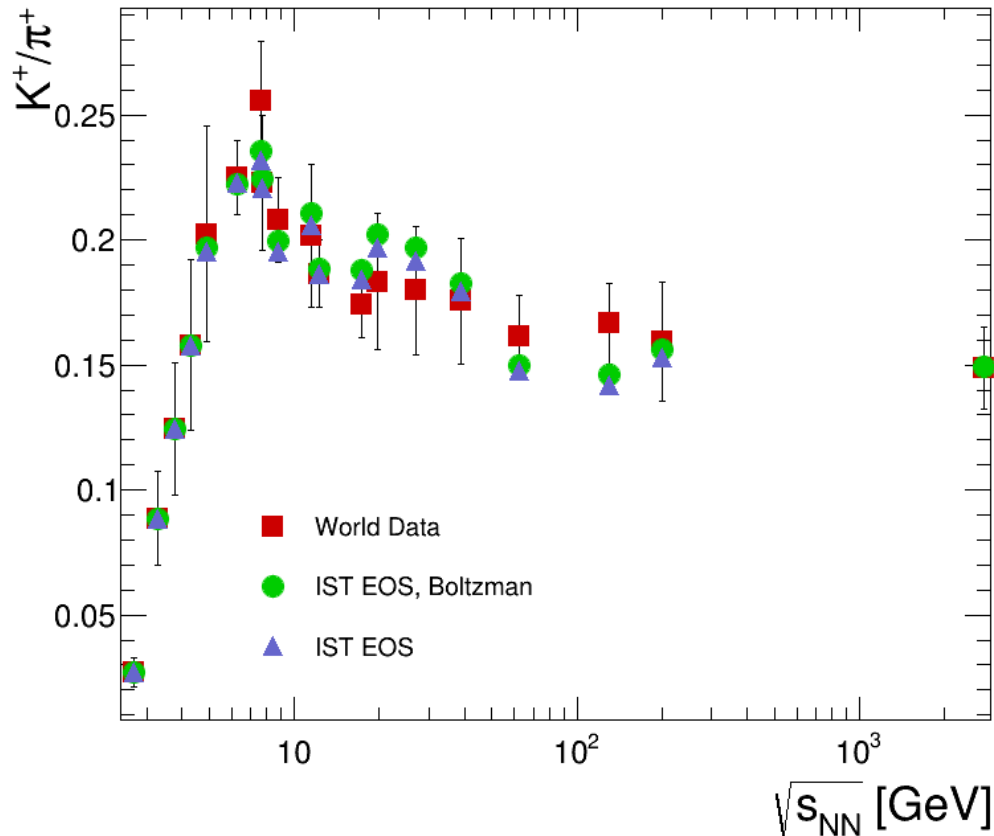
[K. A. Bugaev et al., Ukr. J. Phys. 61 \(2016\) no. 8, 659](#)

[J. Cleymans et al., Phys. Rev. C 73 \(2006\) 034905](#)

[S. Borsanyi et al., Phys. Rev. Lett. 125 \(2020\) 052001](#)



Results of the IST EOS with weak decays for K^+/π^+ ratio



- K^+/π^+ is the most problematic ratio for description by different models
- Inclusion of weak decays greatly improves the description of particle ratios in the experimental data

Conclusions

- ❑ IST EOS is a good tool to describe particle yields and to get chemical freeze-out parameters
- ❑ An updated version of this model allows the fitting of ratios, taking into account both inclusive and exclusive feed-down corrections consistently with experimental analysis
- ❑ Brings the chemical freeze-out temperature to the right track. It gets lower than LQCD predictions for pseudocritical T
- ❑ Provides a good description of the particle ratios from the existing experimental data
- ❑ The chemical freeze-out parameters from the IST EOS fits for STAR and NA49 data are close to the LQCD calculations.

Back up

Resonances width

- The resonance width is taken into account in thermal densities as it is crucial in a thermal model
- For instance, description of pion yields cannot be achieved without its inclusion: $m_\sigma = 484 \pm 24$ MeV, width $\Gamma_\sigma = 510 \pm 20$ MeV

$$n_X^{tot} = n_X^{thermal} + n_X^{decay} = n_X^{th} + \sum_Y n_Y^{th} Br(Y \rightarrow X)$$

$Br(Y \rightarrow X)$ – decay branching of Yth hadron into X

- Fit of the particle ratios gives smaller systematic uncertainties than fitting of yields