
Coherent synchrotron radiation: theory and simulations.

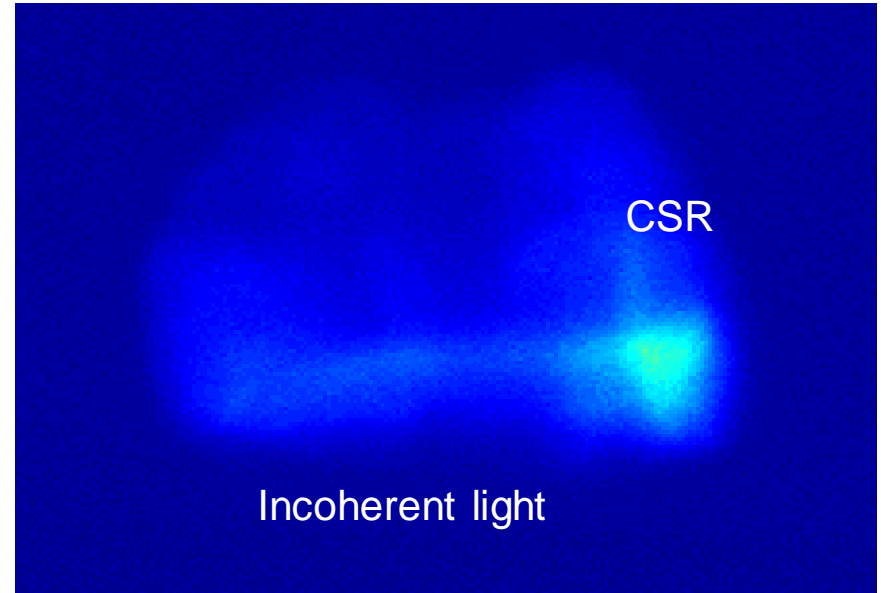
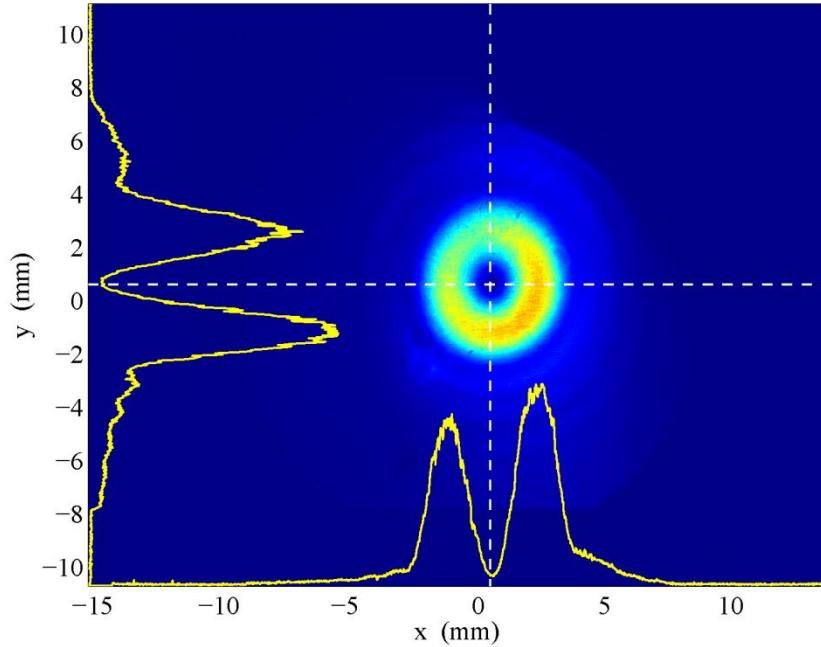
Sasha Novokhatski

ICFA Beam Dynamics Mini Workshop on
Low Emittance Rings 2011
October 3-5, 2011 Heraklion, Crete, Greece

CSR images



Profile Monitor YAGS:DMP1:500 07-Apr-2009 12:58:20



LCLS: CSR light on the YAG screen,
when the electron bunch is bent down

Radiation from a horizontal bend



Theory of CSR



In 1949 J.Schwinger published a paper “On the Classical radiation of Accelerated Electrons” (*Phys. Rev. v. 75, Num 12, 1920*), where he presented his approximation for the spectrum of the synchrotron radiation for the circular trajectory of an electron

$$P(\omega) = \frac{3^{2/3}}{4\pi} \frac{e^2}{\rho} \left(\frac{E}{mc^2} \right)^4 \frac{\omega_0 \omega}{\omega_c^2} \int_{\omega/\omega_c}^{\infty} K_{5/3}(\eta) d\eta$$

introducing the critical frequency

$$\omega_c = \frac{3}{2} \omega_0 \left(\frac{E}{mc^2} \right)^3 = \frac{3}{2} \frac{c}{\rho} \left(\frac{E}{mc^2} \right)^3$$

He verified that his approximation gives the total power

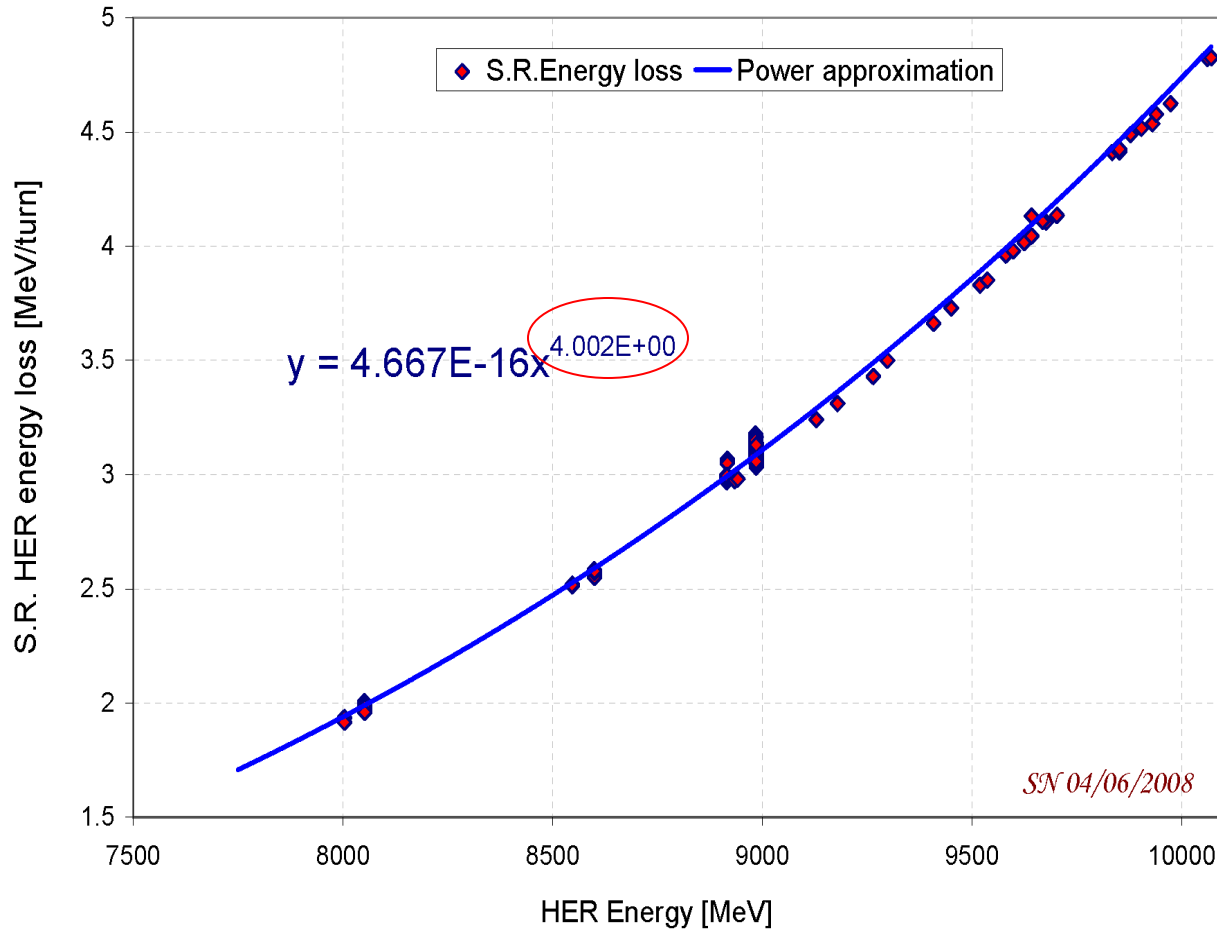
$$P(t) = \int_0^{\infty} P(\omega) d\omega = \frac{2}{3} \omega_0 \frac{e^2}{\rho} \left(\frac{E}{mc^2} \right)^4$$



I had a chance to check E^4



Measurement at PEP-II HER during the energy scan in 2008



SR spectrum



For small frequencies
power is independent of
the electron energy

$$\omega \ll \omega_c$$

$$P(\omega) \approx \frac{3^{1/6} \Gamma\left(\frac{2}{3}\right)}{4\pi} \frac{e^2}{\rho} \left(\frac{\omega}{\omega_0}\right)^{1/3}$$

CSR-impedance ?

For high frequencies
power exponentially
decreases

$$\omega \gg \omega_c$$

$$P(\omega) \approx \frac{3}{4} \left(\frac{3}{2\pi}\right)^{1/2} \frac{e^2}{\rho} \left(\frac{E}{mc^2}\right)^4 \frac{\omega_0}{\omega_c} \left(\frac{\omega}{\omega_c}\right)^{1/2} \exp\left(-\frac{\omega}{\omega_c}\right)$$



Critical frequency and power loss



We know that a bunch of electrons with a finite length may coherently excite electromagnetic fields if their frequencies less than a bunch (critical) frequency

$$\omega \ll \omega_b = \frac{c}{\sigma}$$

Comparing the critical frequencies we may introduce an equivalent bunch length for SR fields

$$\omega_c = \omega_b \quad \sigma_{SR} = \frac{3}{2} \frac{\rho}{\left(\frac{E}{mc^2}\right)^3}$$

Now we rewrite the formula for the SR power

$$P(t) = \frac{2}{3} \omega_0 \frac{e^2}{\rho} \left(\frac{E}{mc^2}\right)^4 = \frac{2}{3} \omega_0 \frac{Q^2}{\rho} \left(\frac{3}{2} \frac{\rho}{\sigma_{SE}}\right)^{4/3} = 2 \cdot 3^{1/3} \frac{Q^2}{\rho^2} \left(\frac{\rho}{2\sigma_{SE}}\right)^{4/3}$$

Schiff, Nordvick, Saxon, Murphy, Derbenev, ...



CSR shielding in 1954



PHYSICAL REVIEW

VOLUME 96, NUMBER 1

OCTOBER 1, 1954

Suppression of Coherent Radiation by Electrons in a Synchrotron*

JOHN S. NODVICK† AND DAVID S. SAXON
University of California, Los Angeles, California
(Received May 25, 1954)

$$K_{pp}(\rho, h, \sigma) = \frac{2}{3} \frac{Z_0 c}{\pi \rho} \times \frac{\rho}{h} \sum_{n=1} \left(\frac{\sin n\sqrt{3} \frac{\sigma}{\rho}}{n\sqrt{3} \frac{\sigma}{\rho}} \right)^2 \times \sum_{j=1,3,\dots}^{\gamma_j < n} \frac{\gamma_j^4}{n^3} \times \left[K_{1/3}^2 \left(\frac{\gamma_j^3}{3n^2} \right) + K_{2/3}^2 \left(\frac{\gamma_j^3}{3n^2} \right) \right]$$

$$\gamma_j = j\pi \frac{\rho}{h}$$

Murphy, Warnock, ...



It can be a very strong shielding

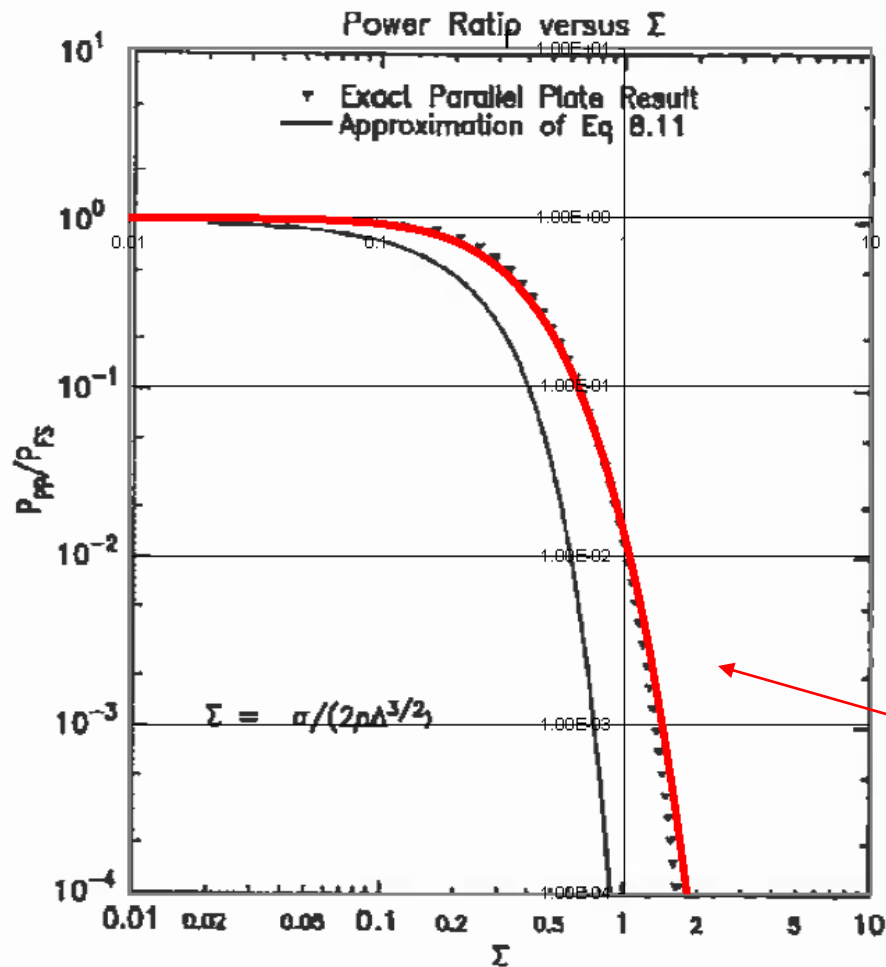


LONGITUDINAL WAKEFIELD FOR AN ELECTRON MOVING ON A CIRCULAR ORBIT

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^a National Synchrotron Light Source, Brookhaven National Laboratory, Upton, NY 11973, USA; ^b Physics Department, University of Maryland, College Park, MD 20742, USA

(Received 6 May 1996; Revised 28 January 1997; In final form 10 March 1997)



$$\tilde{\Sigma} = 2\sqrt{12} \times \Sigma = \frac{\sqrt{12}\sigma}{h} \sqrt{\frac{\rho}{h}}$$

$$\frac{P_{pp}}{P_{FS}} = \frac{\tilde{\Sigma}}{\sinh(\tilde{\Sigma})}$$

hyperbolic sine



CSR simulations



- R. Li. Nucl. Instrum. Meth. Phys. Res. A, 429, 310, 1998.
- G. Bassi et al., Nuc. Instrum. Methods Phys. Res. A, 557, pp. 189–204 (2006).
- M. Borland, Phys. Rev. ST Accel. Beams 4, 070701 (2001).
- G.V. Stupakov and I. A. Kotelnikov, Phys. Rev. ST Accel. Beams, 12, 104401 (2009).
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- ...



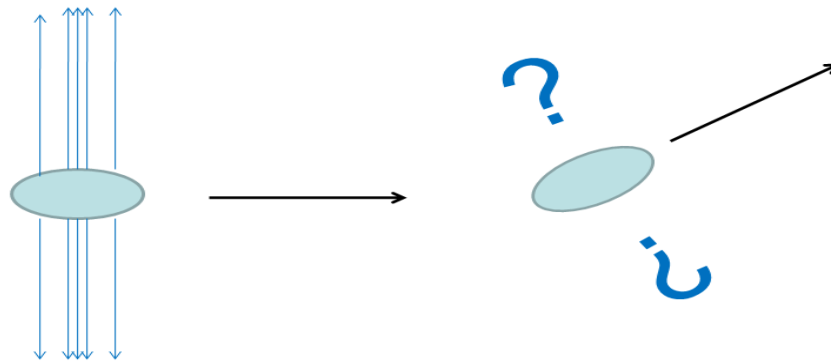
Time-domain presentation



A bunch and radiation are moving together for a long time. How do they separate? A chamber wall cuts the field?

A bunch retarding causes the field radiation?

How does the bunch field change when the bunch has been rotated in a magnetic field?



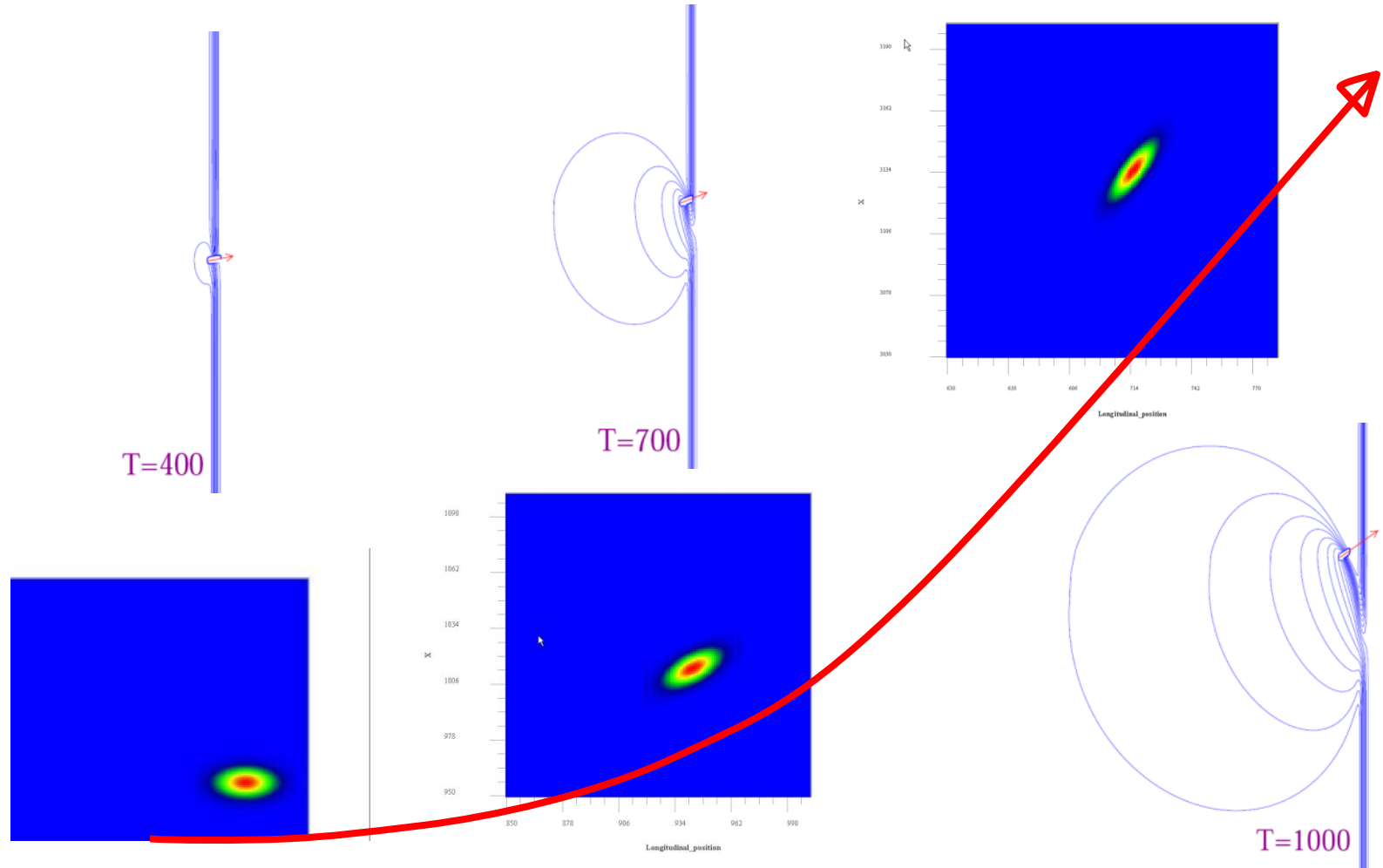
Recently a new method was developed at SLAC



- The method is based on an implicit scheme for solving the electromagnetic equations, Maxwell's equations.
- This algorithm is free of frequency dispersion effects which means that all propagating waves will have their natural phase velocity, completely independent of simulation parameters like mesh size or time step.
- Other known methods, usually explicit, have “mesh driven” dispersion and because of this they need a much smaller mesh size which slows down calculations and can sometimes cause unstable solutions.
- An implicit scheme is a self-consistent method that allows us to calculate fields of much shorter bunches.



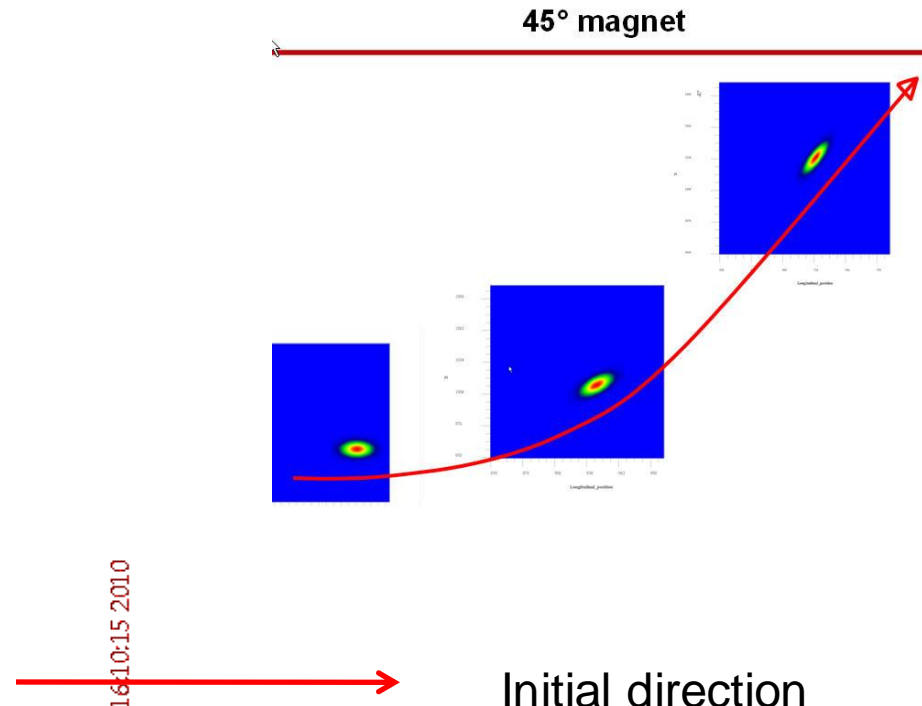
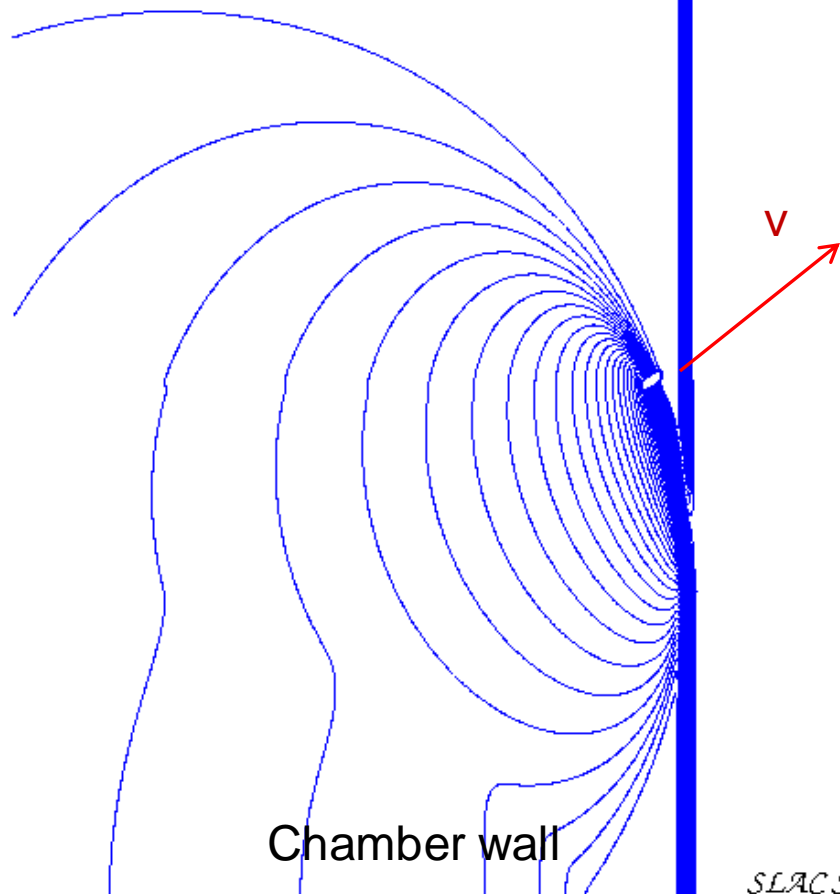
First results: Field dynamics in a magnet



Chamber wall

Time = 1600

Movie show



NOVO:Tue Dec 7 16:10:15 2010

SN

SLAC Stanford

Sasha Novokhatski October 4, 2011

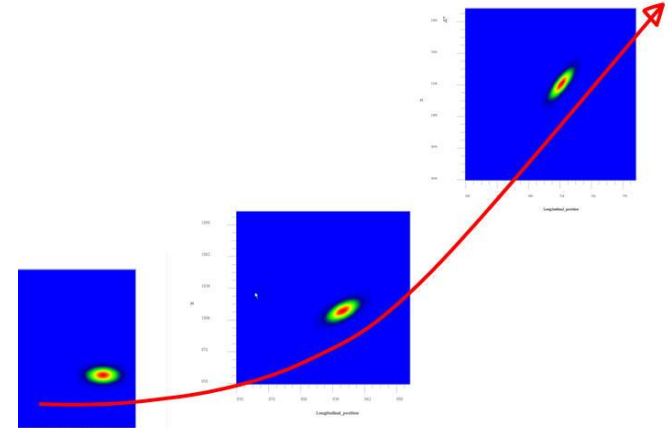


Time = 40

Movie show



45° magnet



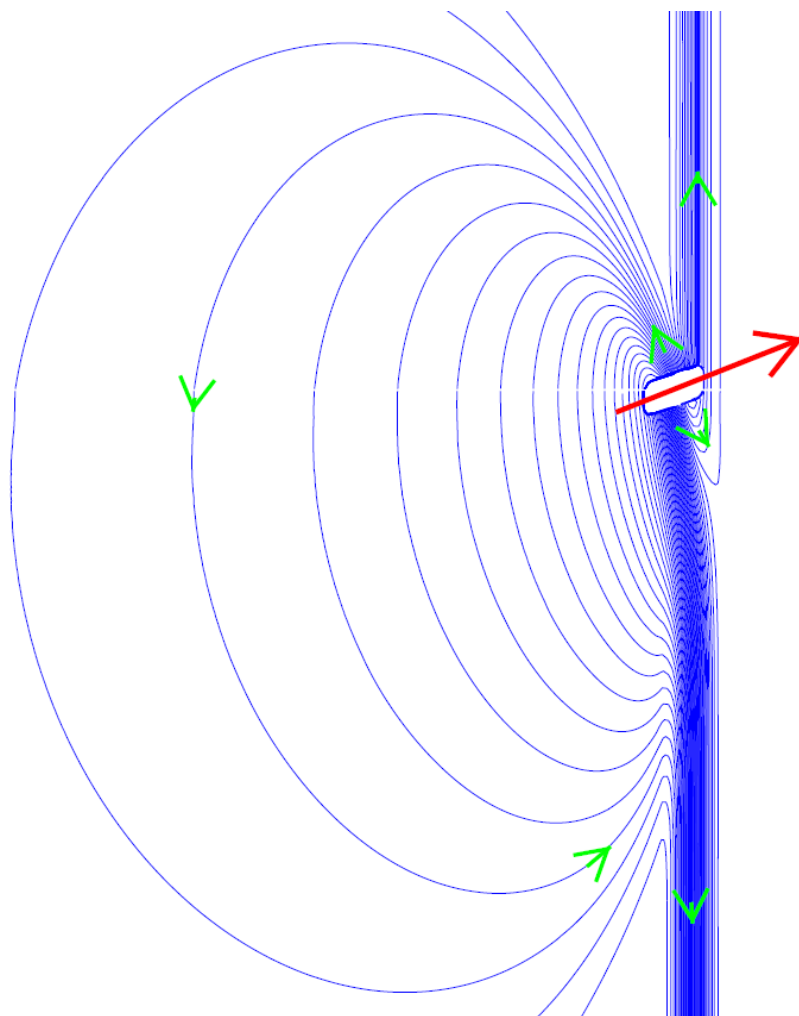
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Initial direction



Bunch self-field remakes itself moving in a magnetic field



The upper field lines take the position of the lower lines

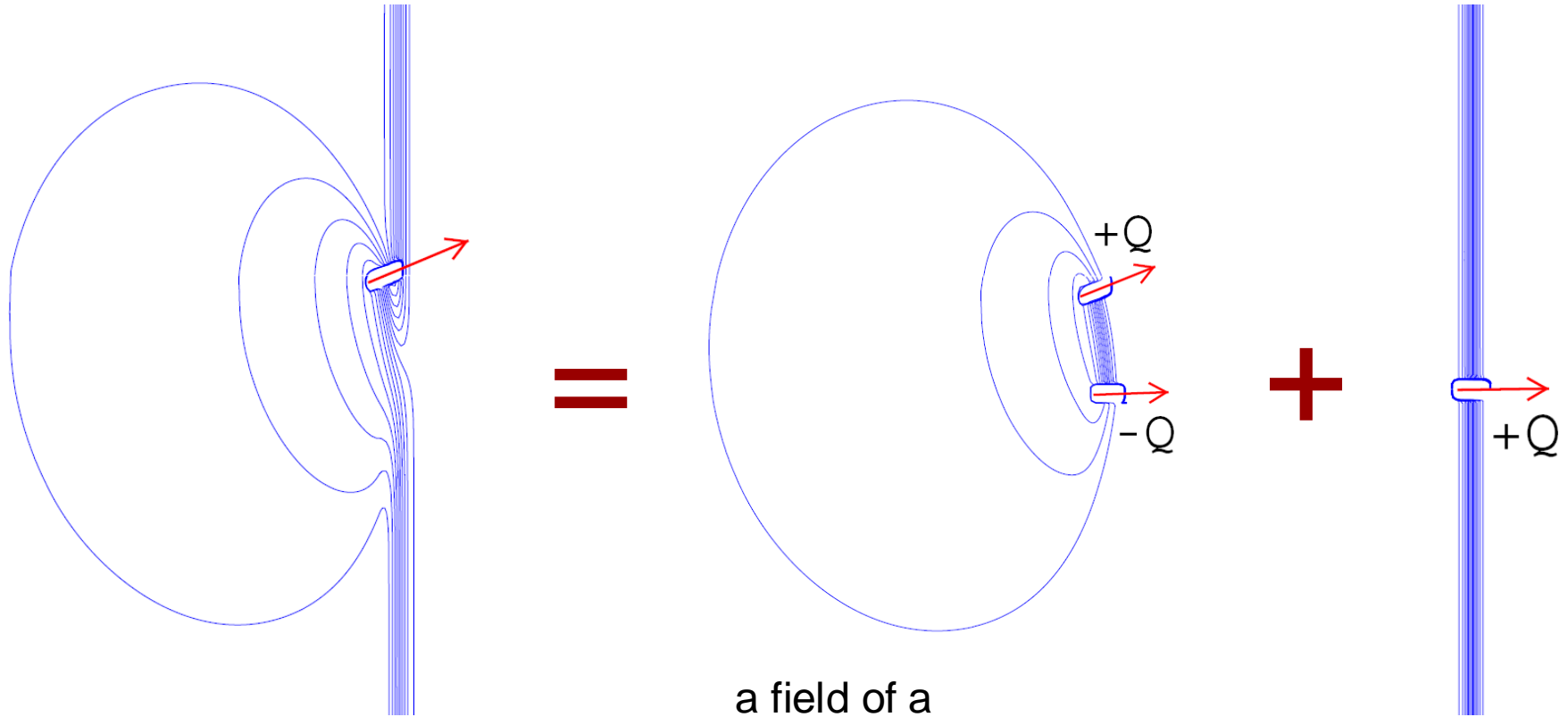
The white box shows the bunch location

The red arrow shows the bunch velocity vector

Green arrows show field line directions

The lower field lines take the place of the upper lines

The picture becomes clear if we decompose the field



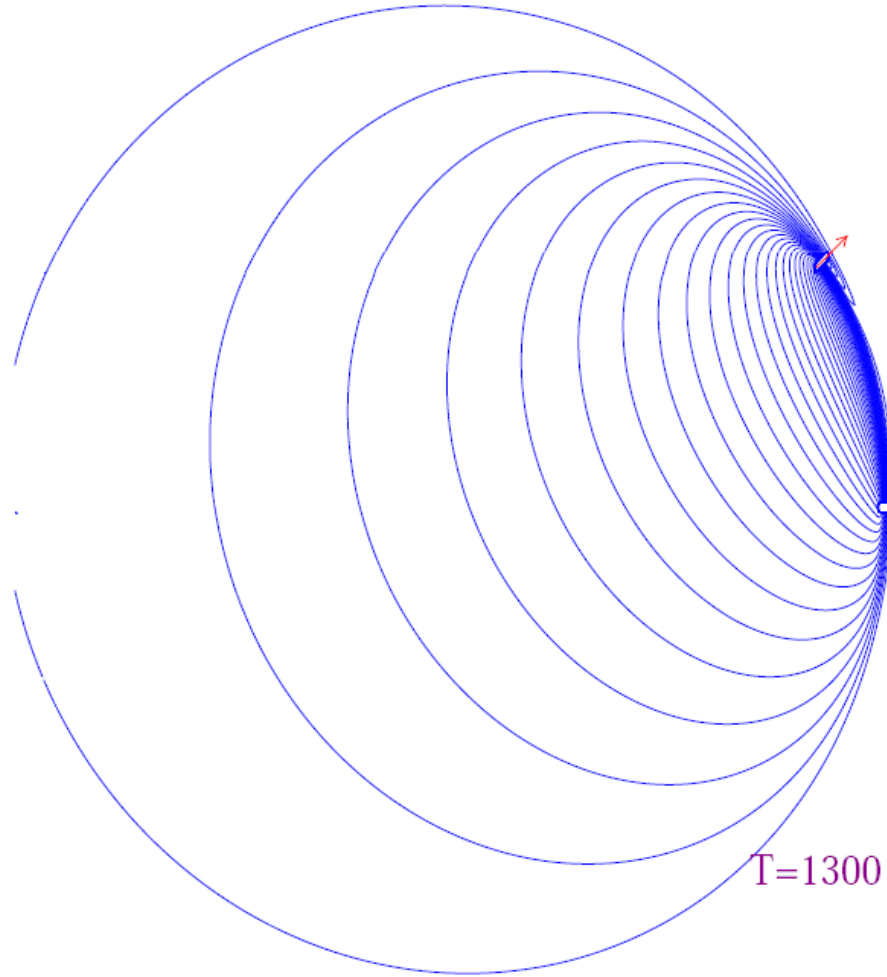
Decomposition of the field of a bunch moving in a magnetic field into two fields:

a field of a moving dipole

a field of a bunch moving straight in initial direction



Detailed plot of a dipole field

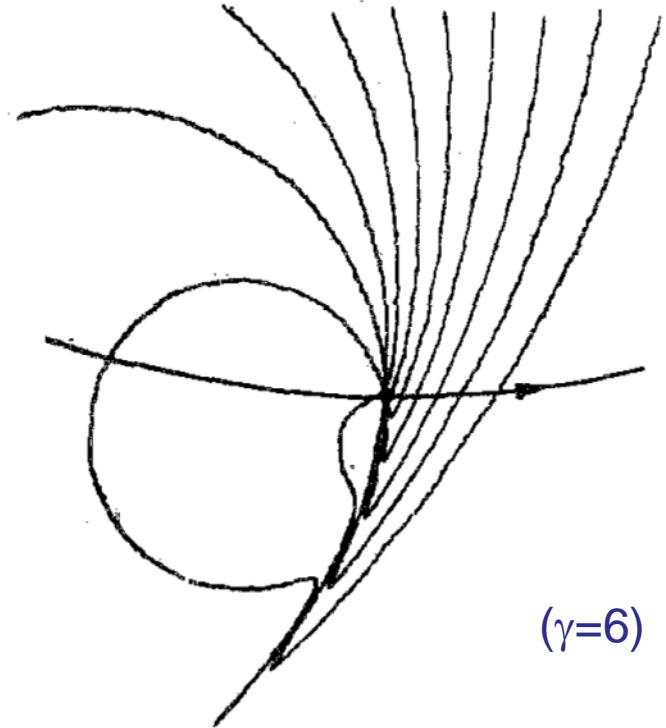
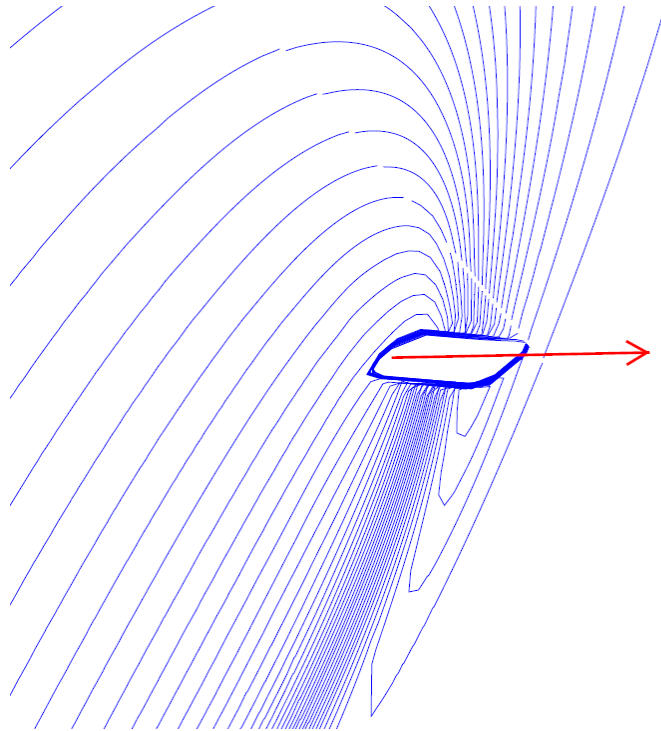


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Comparison with a classical synchrotron radiation

γ - region in front of a particle



($\gamma=6$)

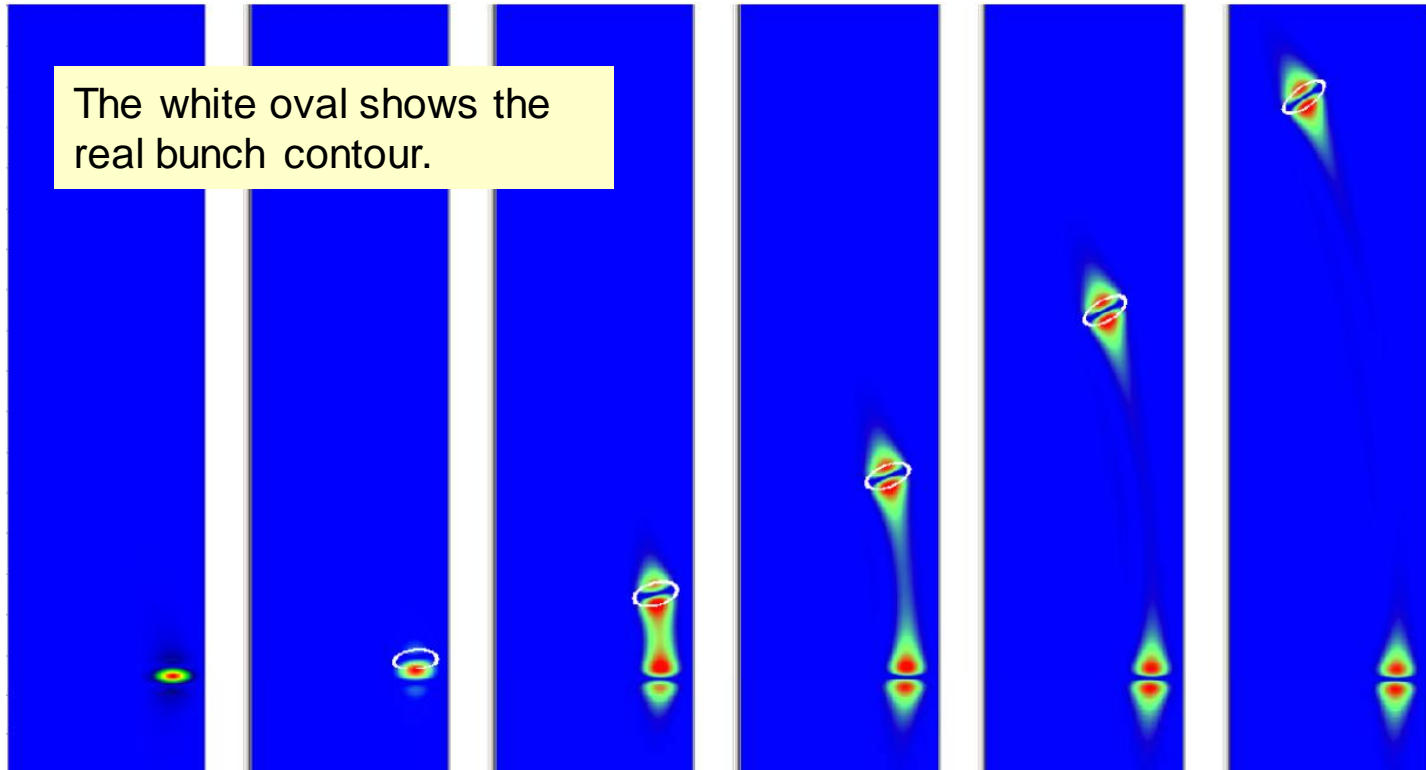
equivalent bunch length
for a bending radius ρ

$$\sigma \approx \frac{2}{3} \frac{\rho}{\gamma^3}$$

FORCE LINES OF ELECTRIC AND MAGNETIC FIELDS
OF AN ARBITRARILY MOVING CHARGE

S. G. Arutyunyan

An absolute dipole electric field in time



The white oval shows the real bunch contour.

When a dipole is created an electric field appears between a real bunch and a virtual bunch. This field increases in value and reaches a maximum value when the bunches are completely separated and then it goes down as the bunches move apart leaving fields only around the bunches.



Electrical forces inside a bunch

Bunch shape

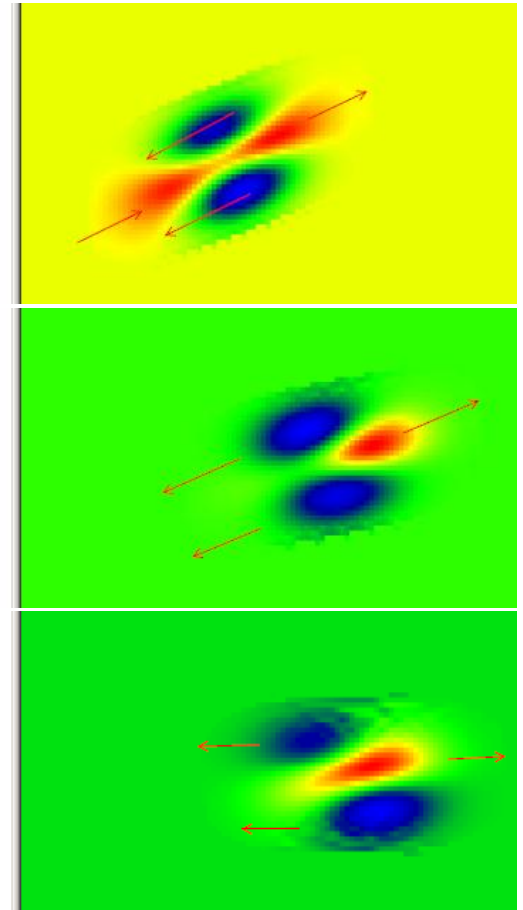
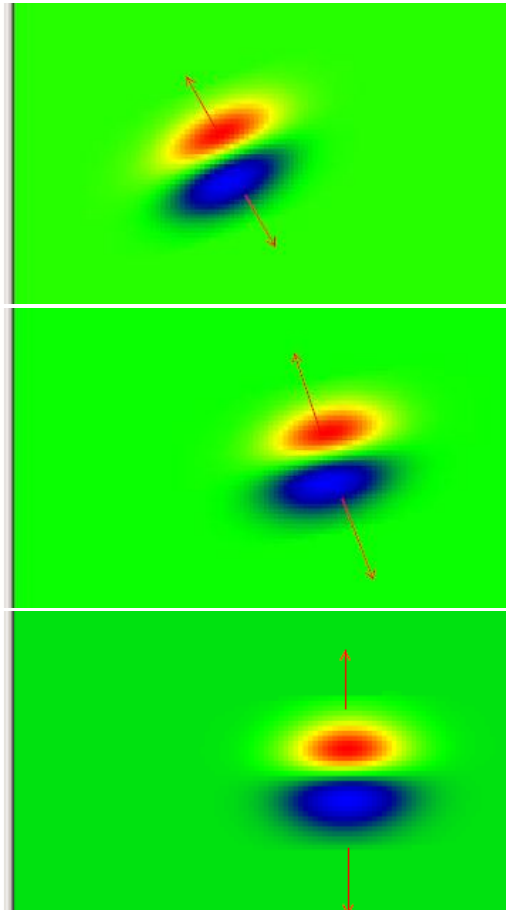
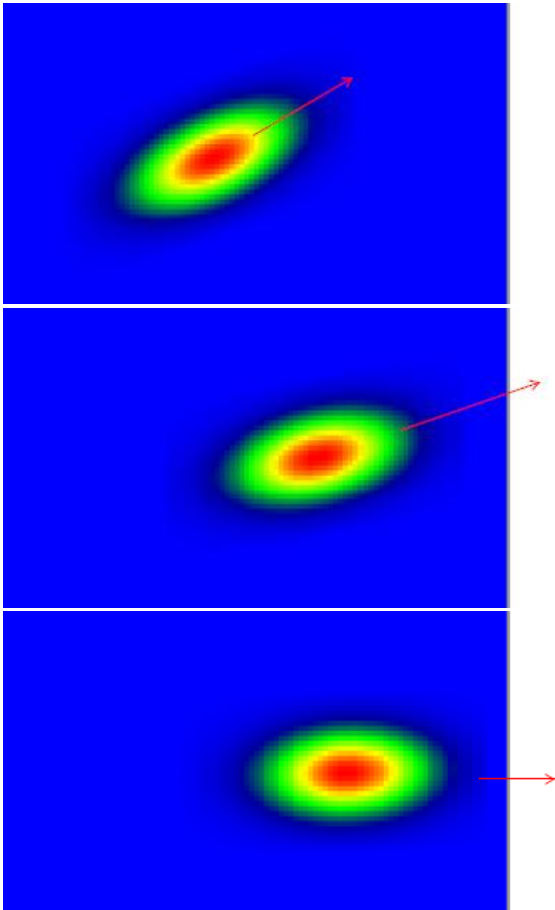
transverse force

collinear force

$$\vec{F}_{\perp} = \frac{\vec{J}}{|\vec{J}|} \times [\vec{E} \times \vec{J}]$$

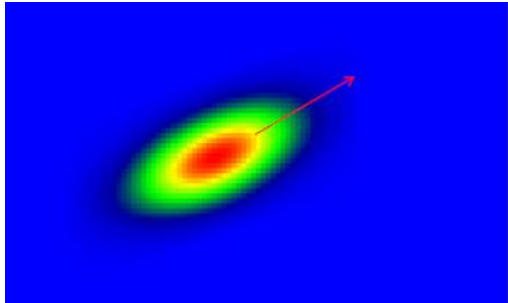
$$F_{\parallel}^e = (\vec{J} \cdot \vec{E})$$

↑
t
i
m
e



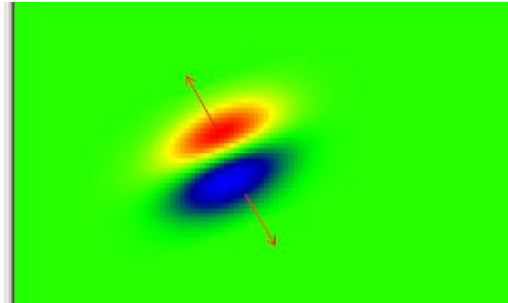
Electrical forces inside a bunch

Bunch shape



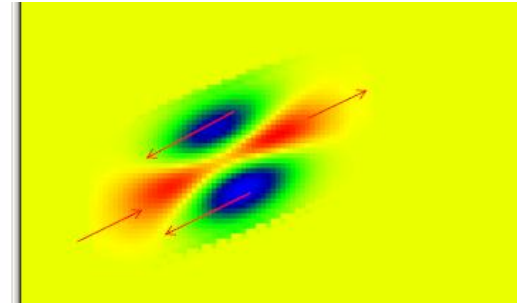
transverse force

$$\vec{F}_{\perp} = \frac{\vec{J}}{|\vec{J}|} \times [\vec{E} \times \vec{J}]$$



collinear force

$$F_{\parallel}^e = (\vec{J} \cdot \vec{E})$$

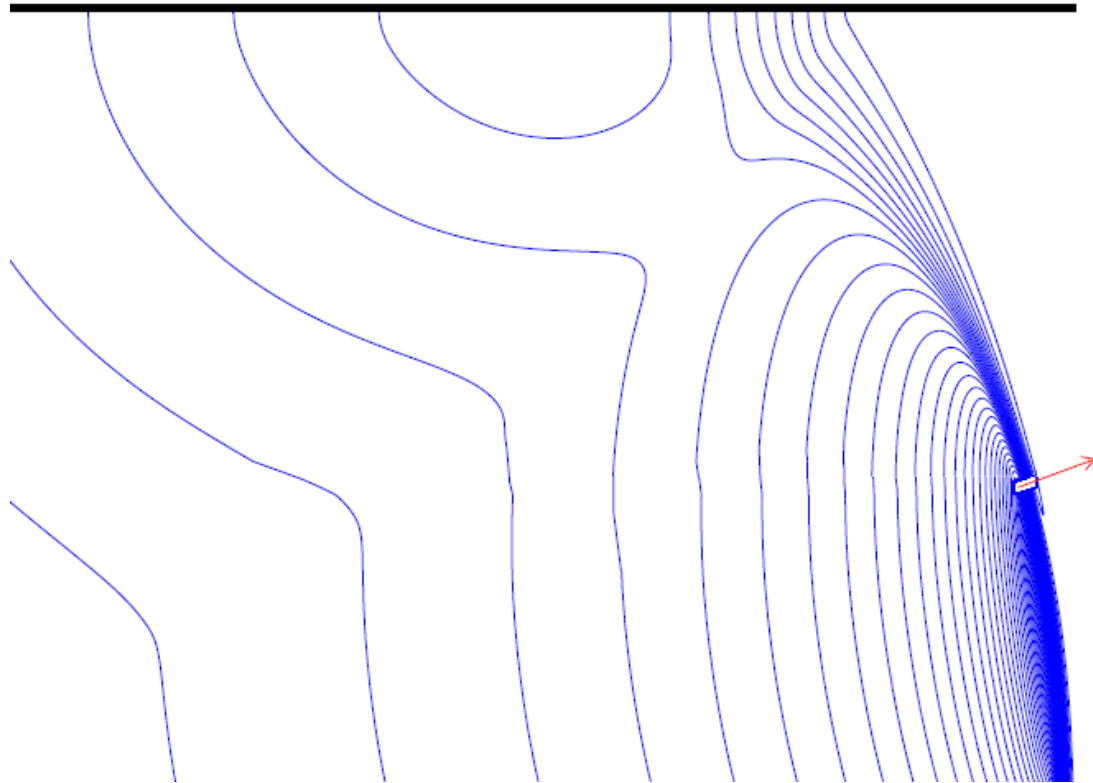


The transverse force is the well known space-charge force, which probably is compensated by a magnetic force in the ultra-relativistic case.

The collinear force is responsible for an energy gain or an energy loss. The particles, which are in the center, in front and at the end of the bunch are accelerating, whereas the particles at the boundaries are decelerating.

The total effect is deceleration and the bunch loses energy, **however the bunch gets an additional energy spread in the transverse direction.**

A new bunch field

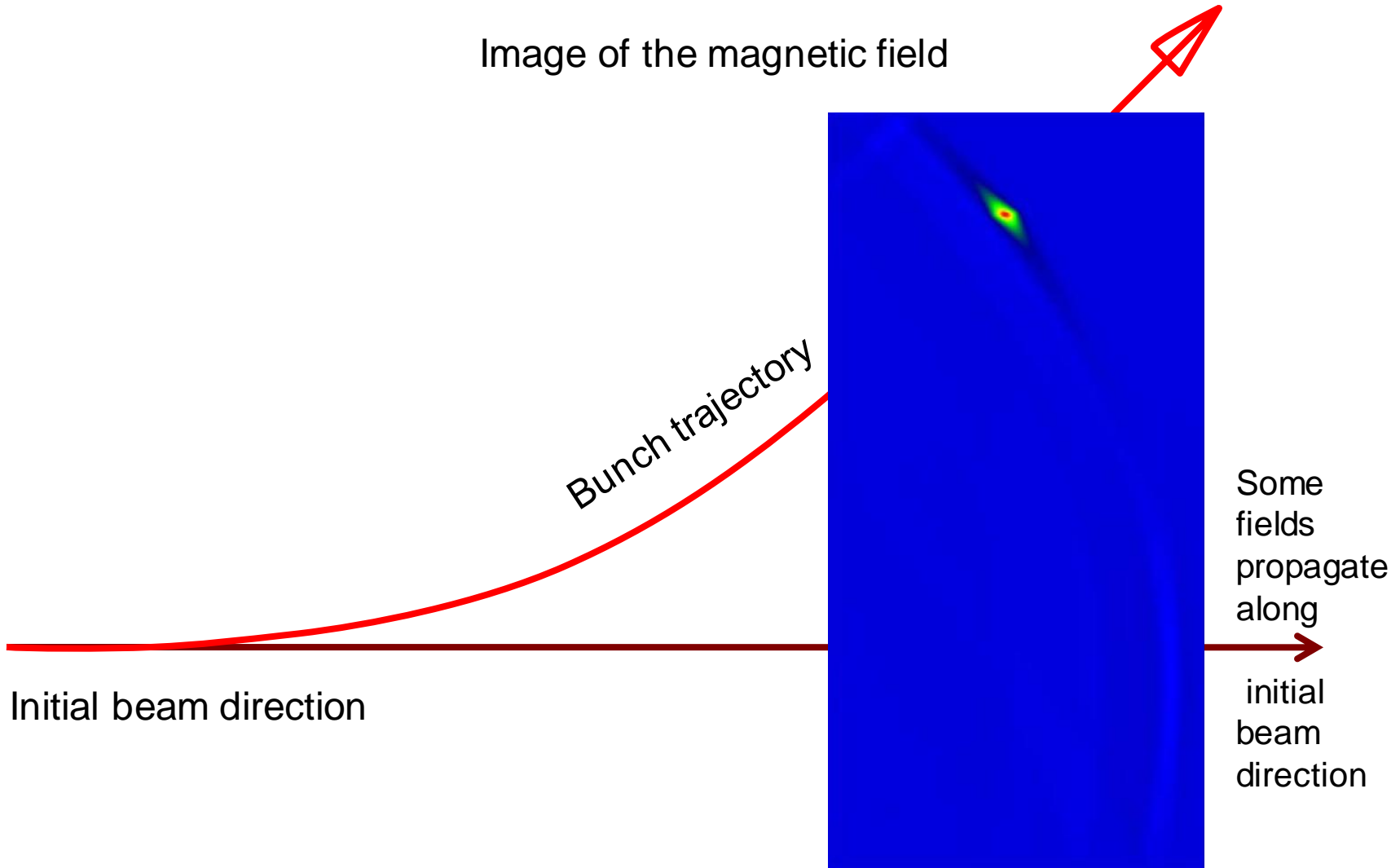


A long way to a steady-state regime

Coherent edge radiation

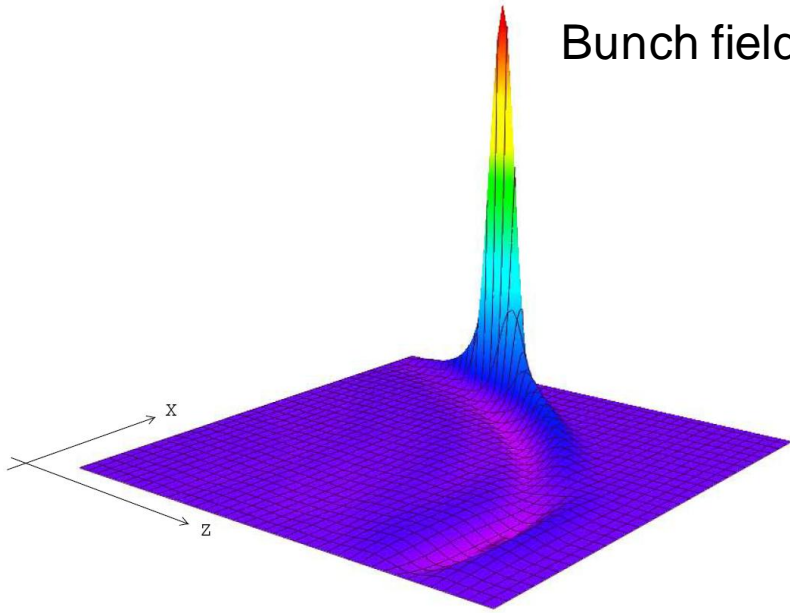


Image of the magnetic field

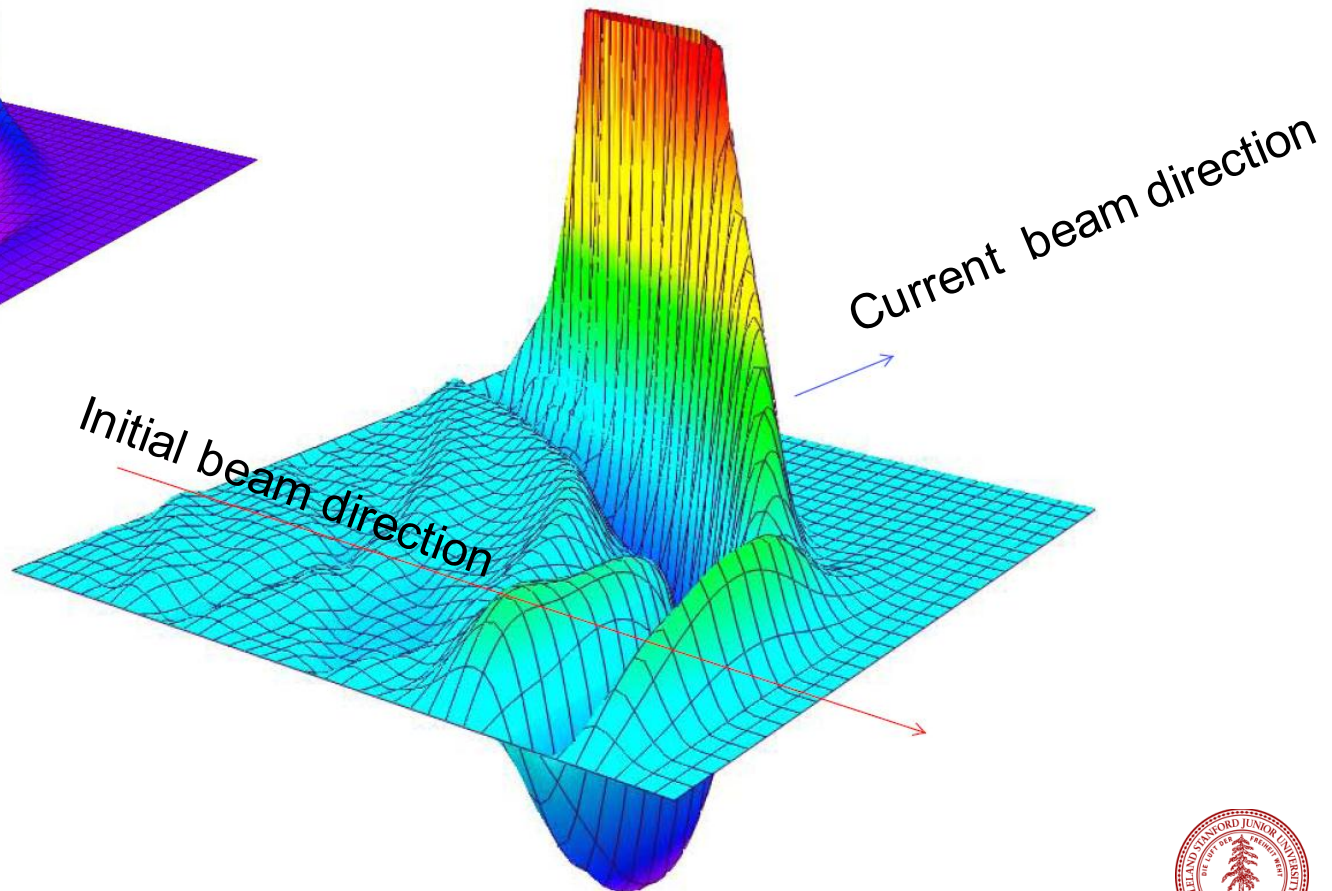


Magnetic field plots

Bunch field



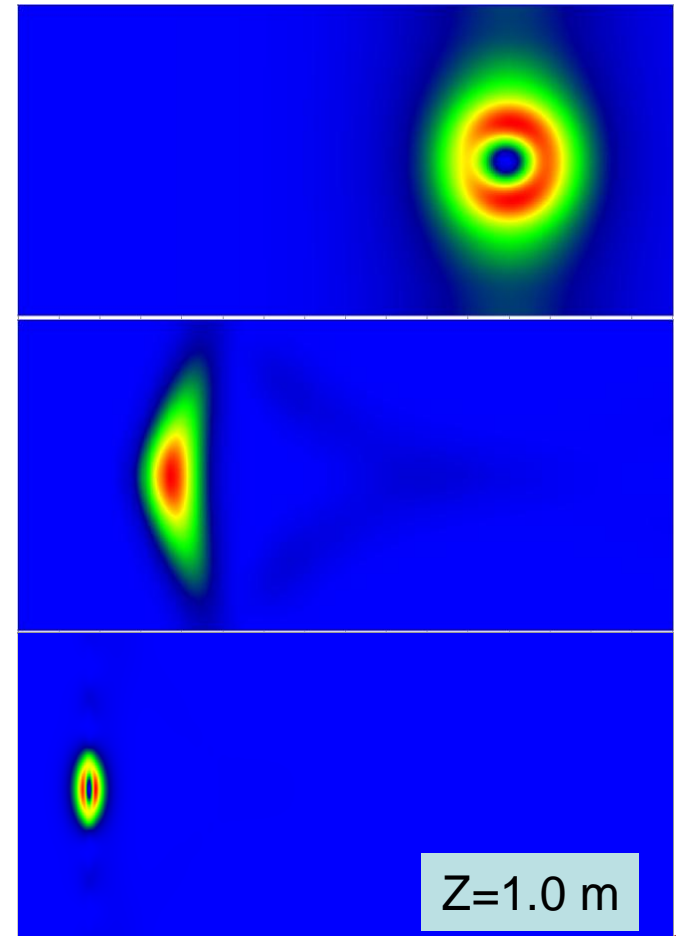
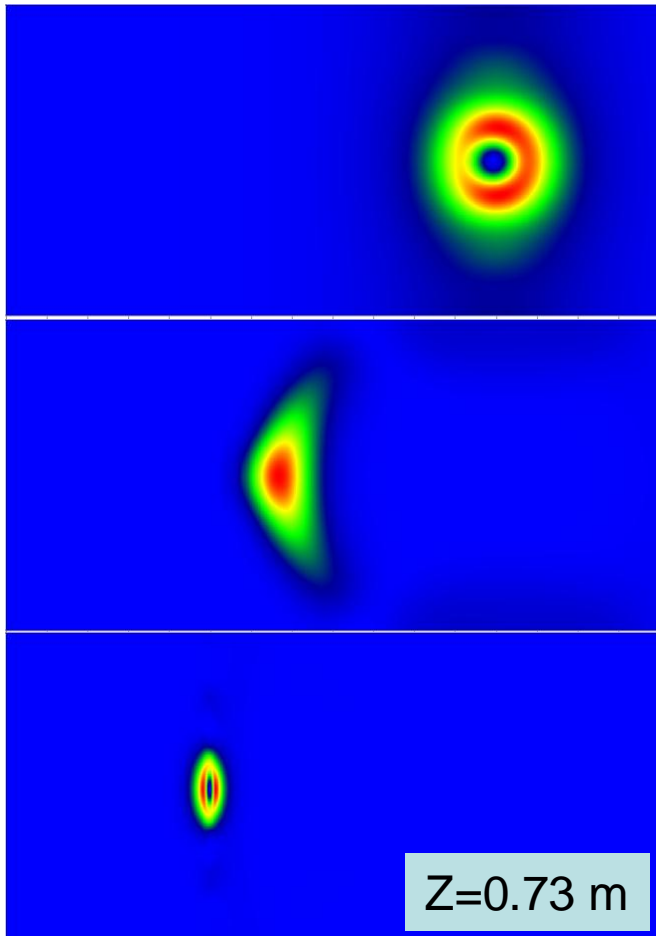
Magnified



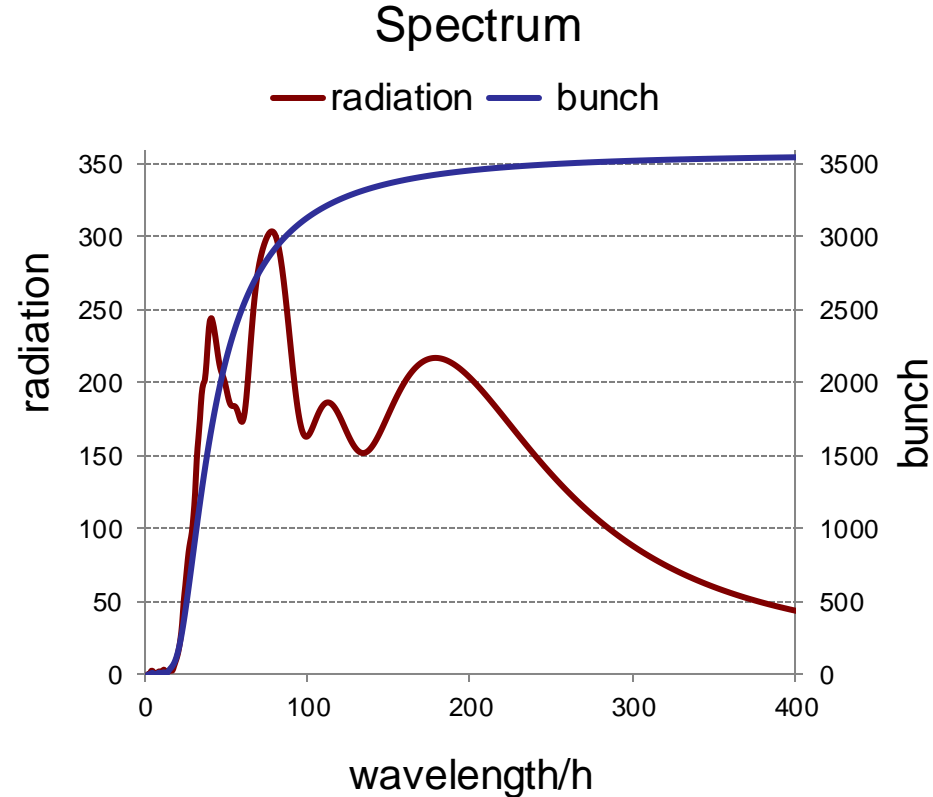
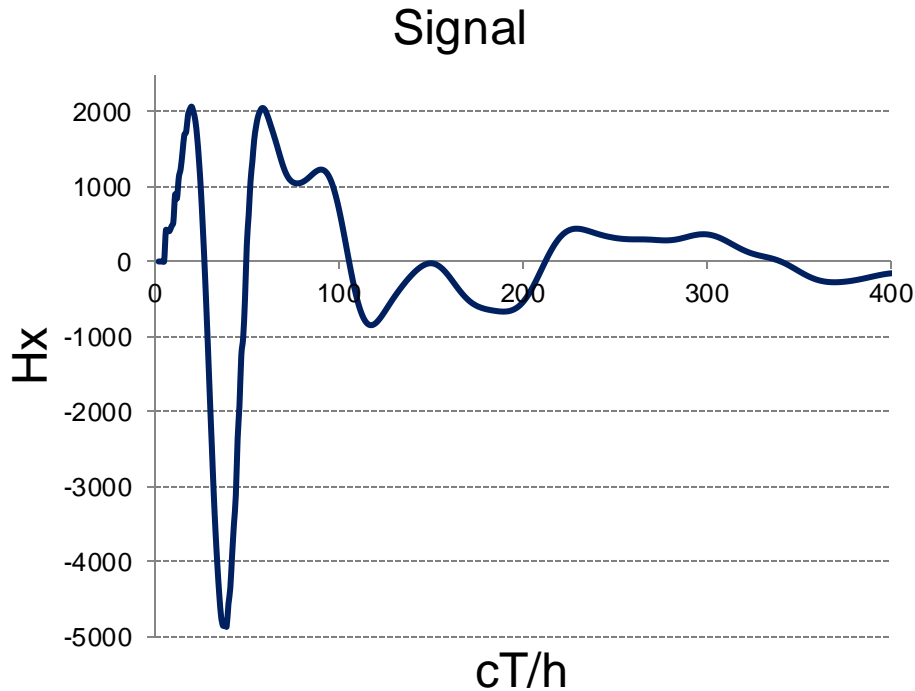
Images of radiation (transverse magnetic field)



very similar to the images, which we have seen on the YAG screen after the dump magnets, which bend down the beam at LCLS.



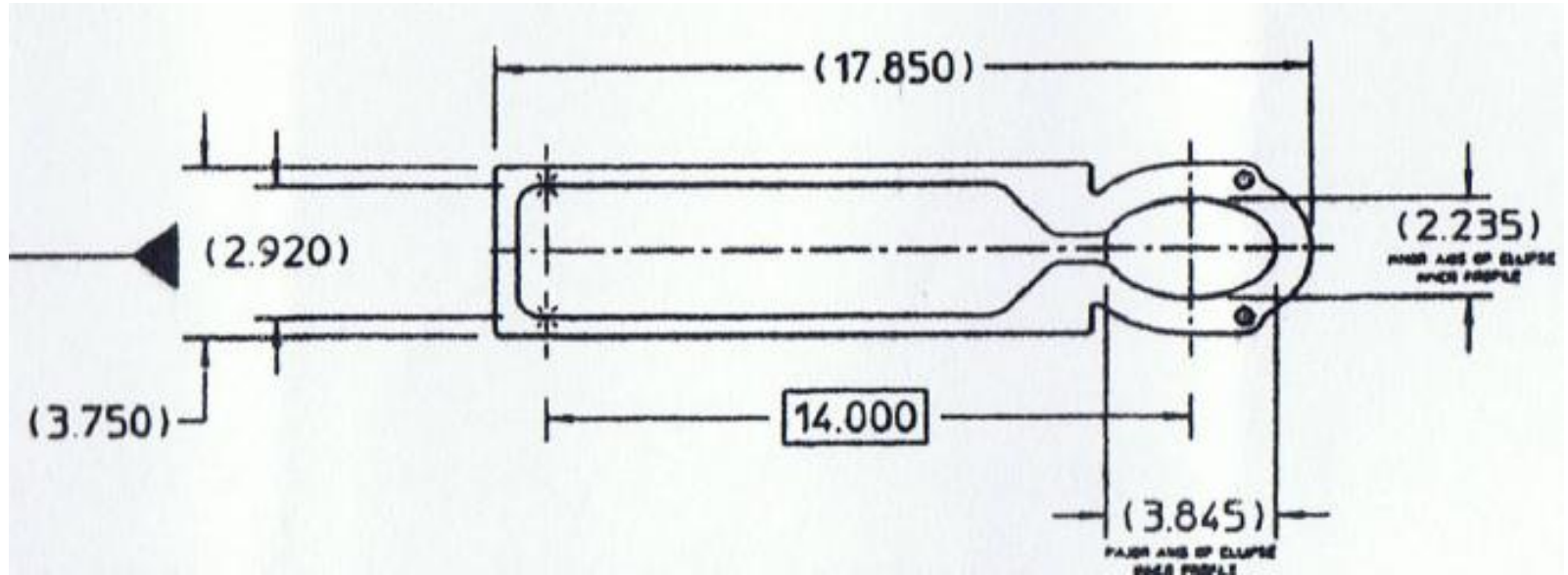
Signal and spectrum of edge radiation



First peak in spectrum corresponds to and is a little bit less than the bunch wavelength = $2\pi\sigma$

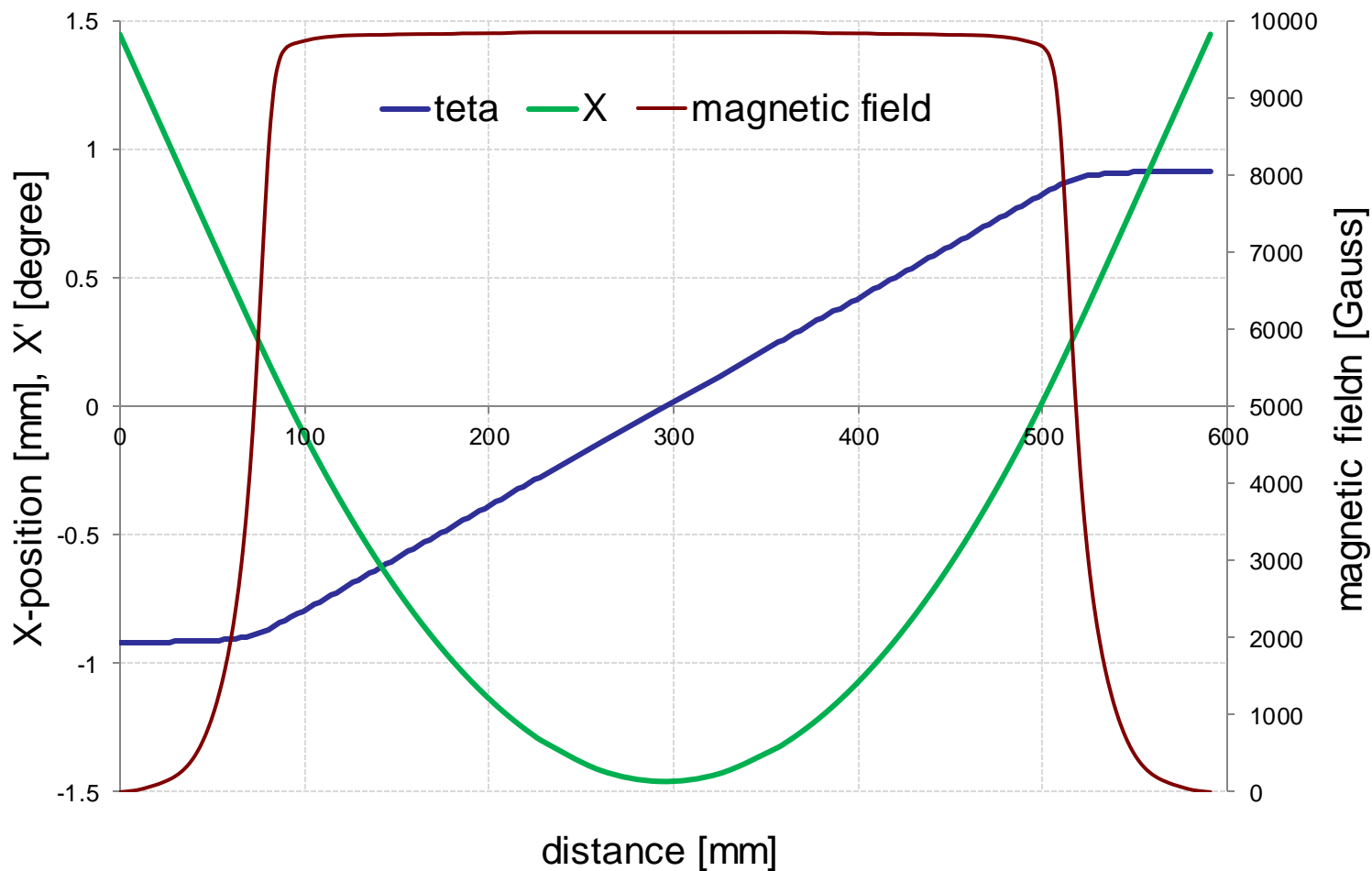


PEP-II LER pumping chamber

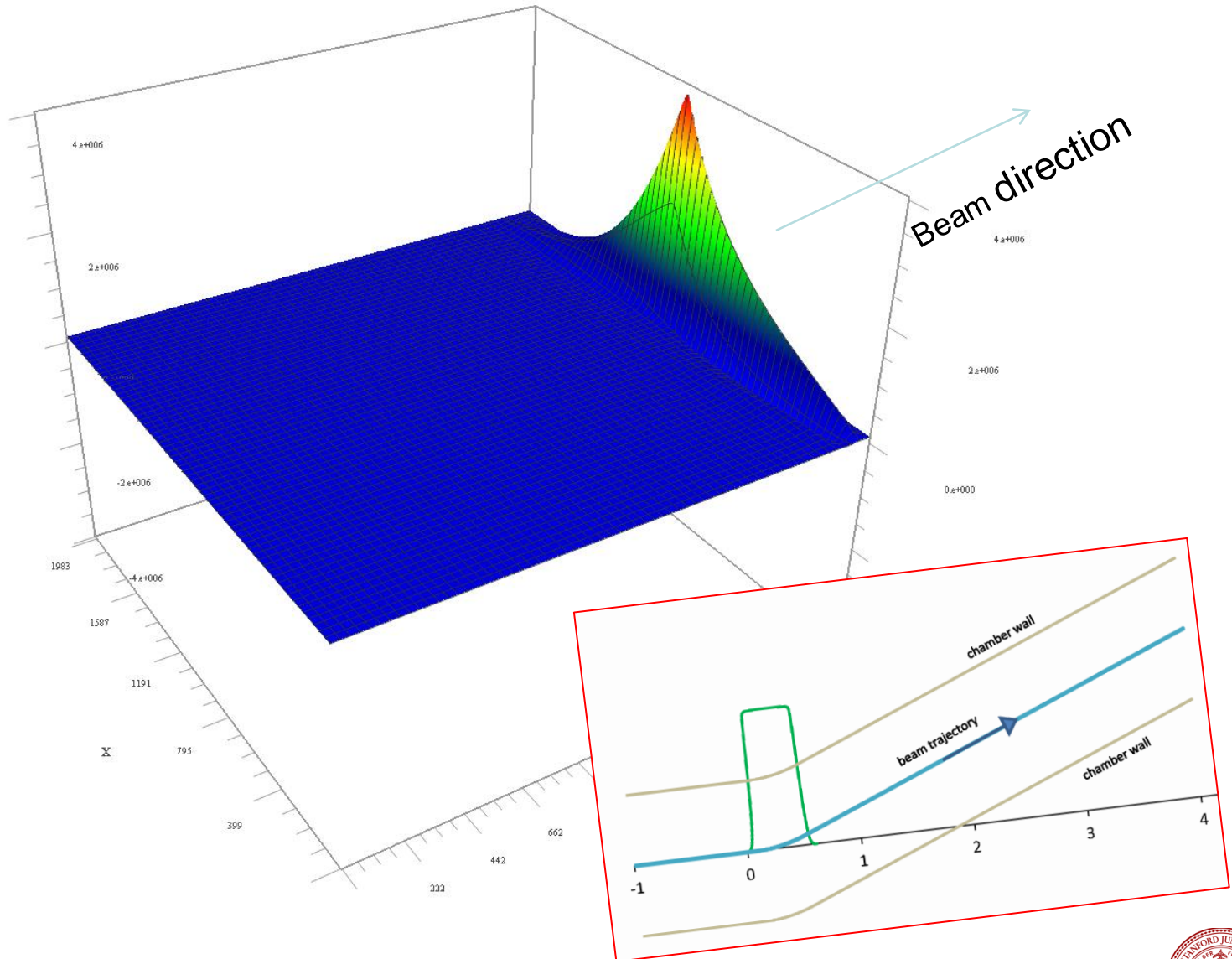


PEP-II LER dipole (a model)

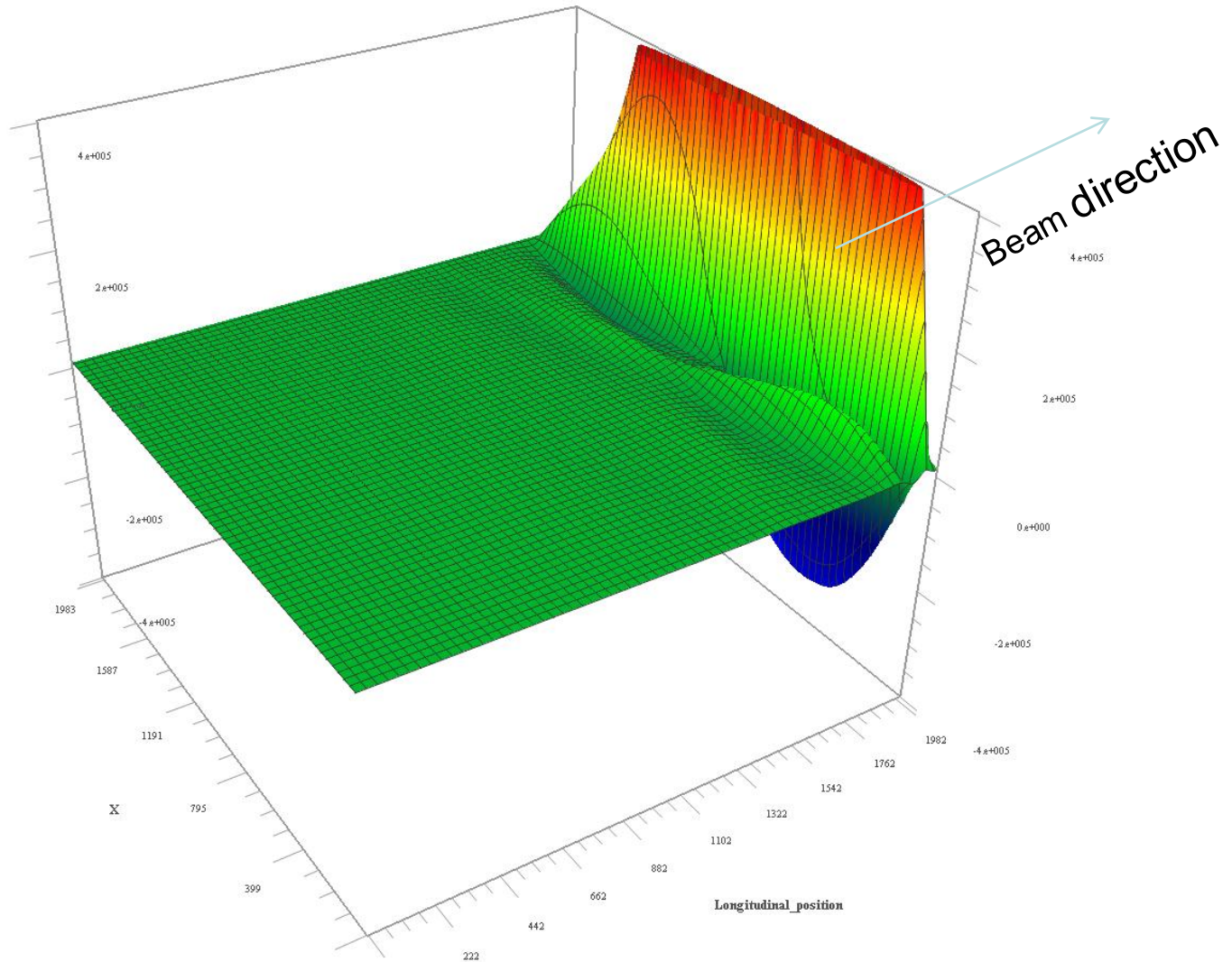
X beam position and magnetic field



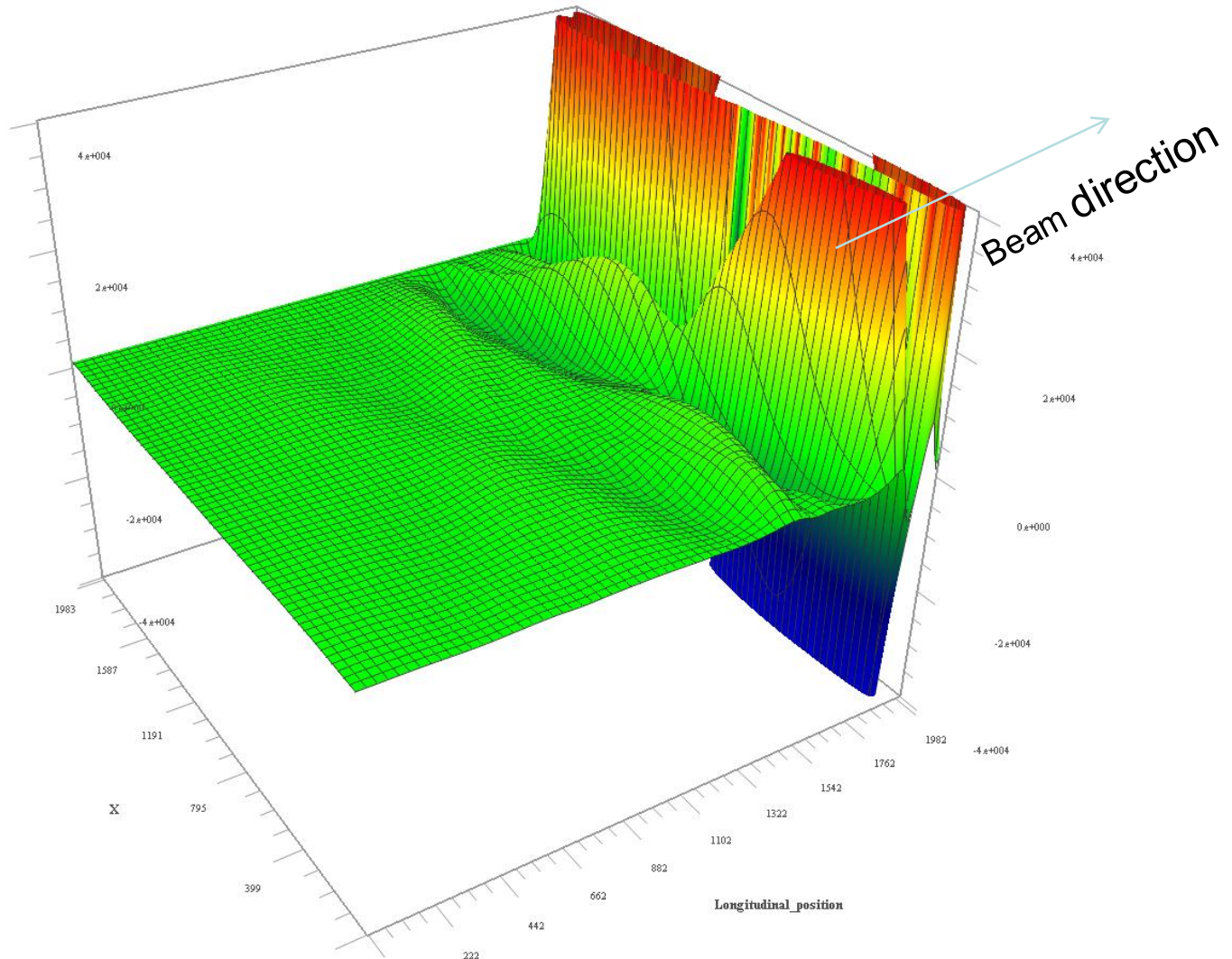
Beam and chamber fields



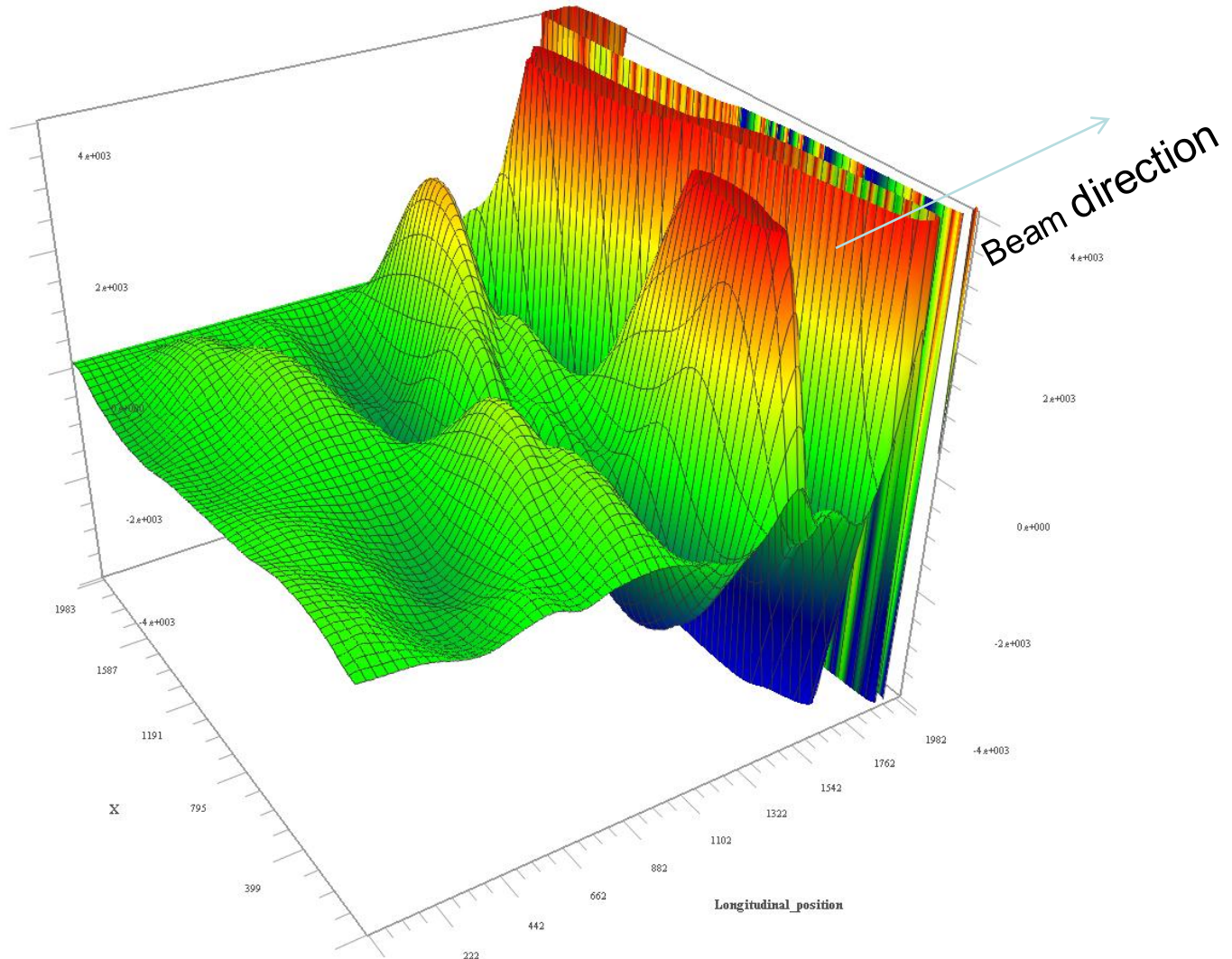
Magnified by 10 times



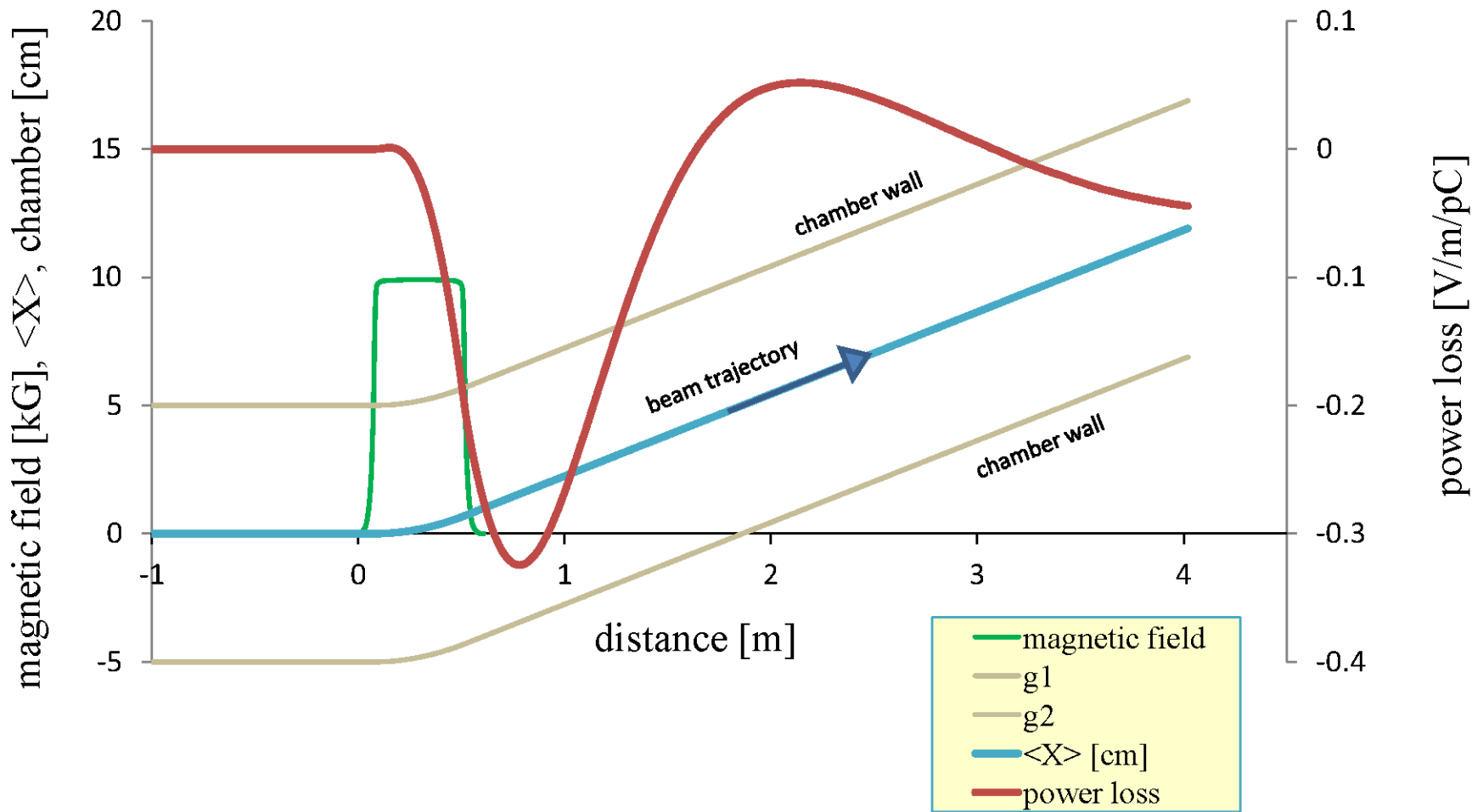
Magnified by 100 times



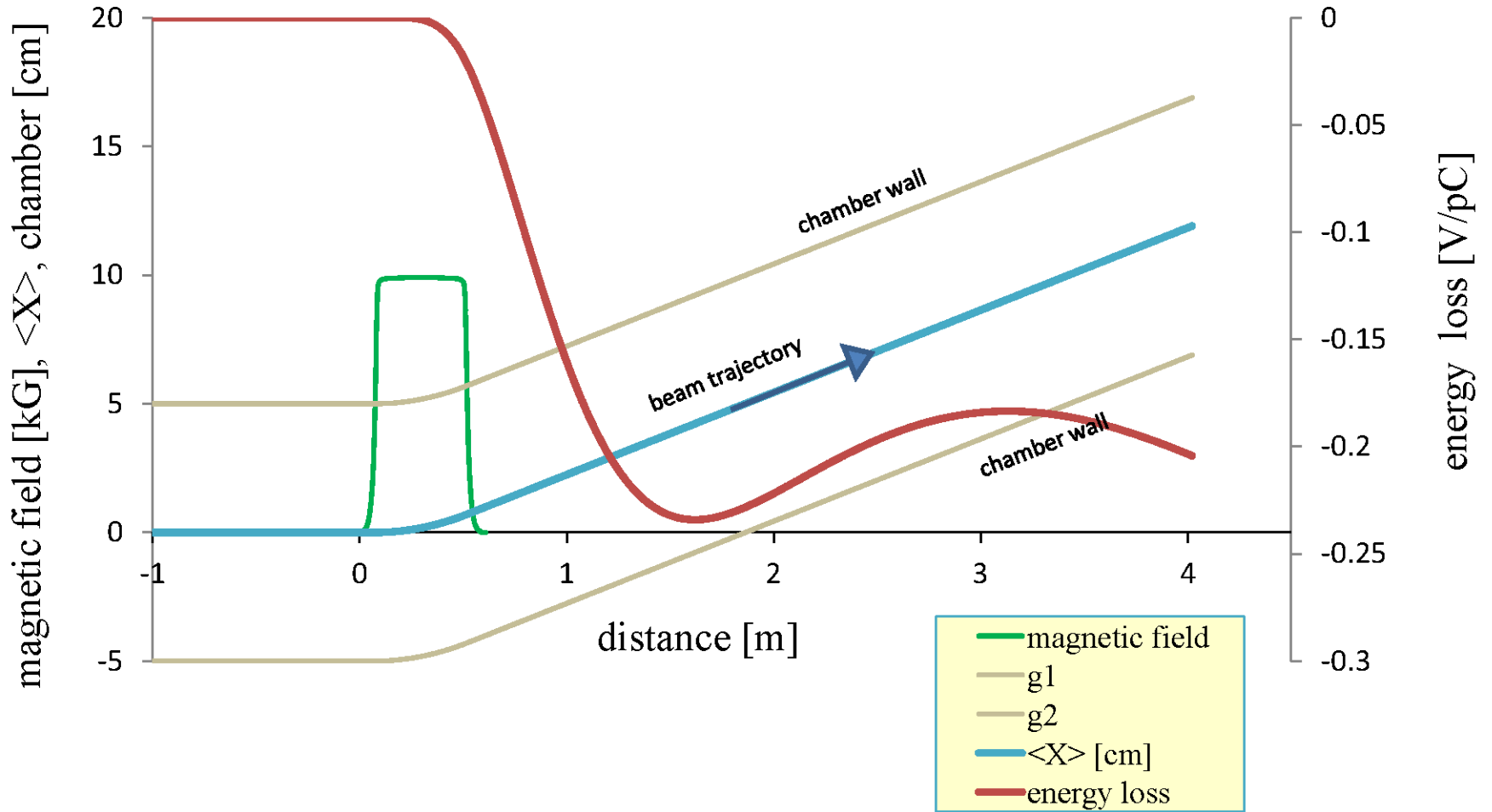
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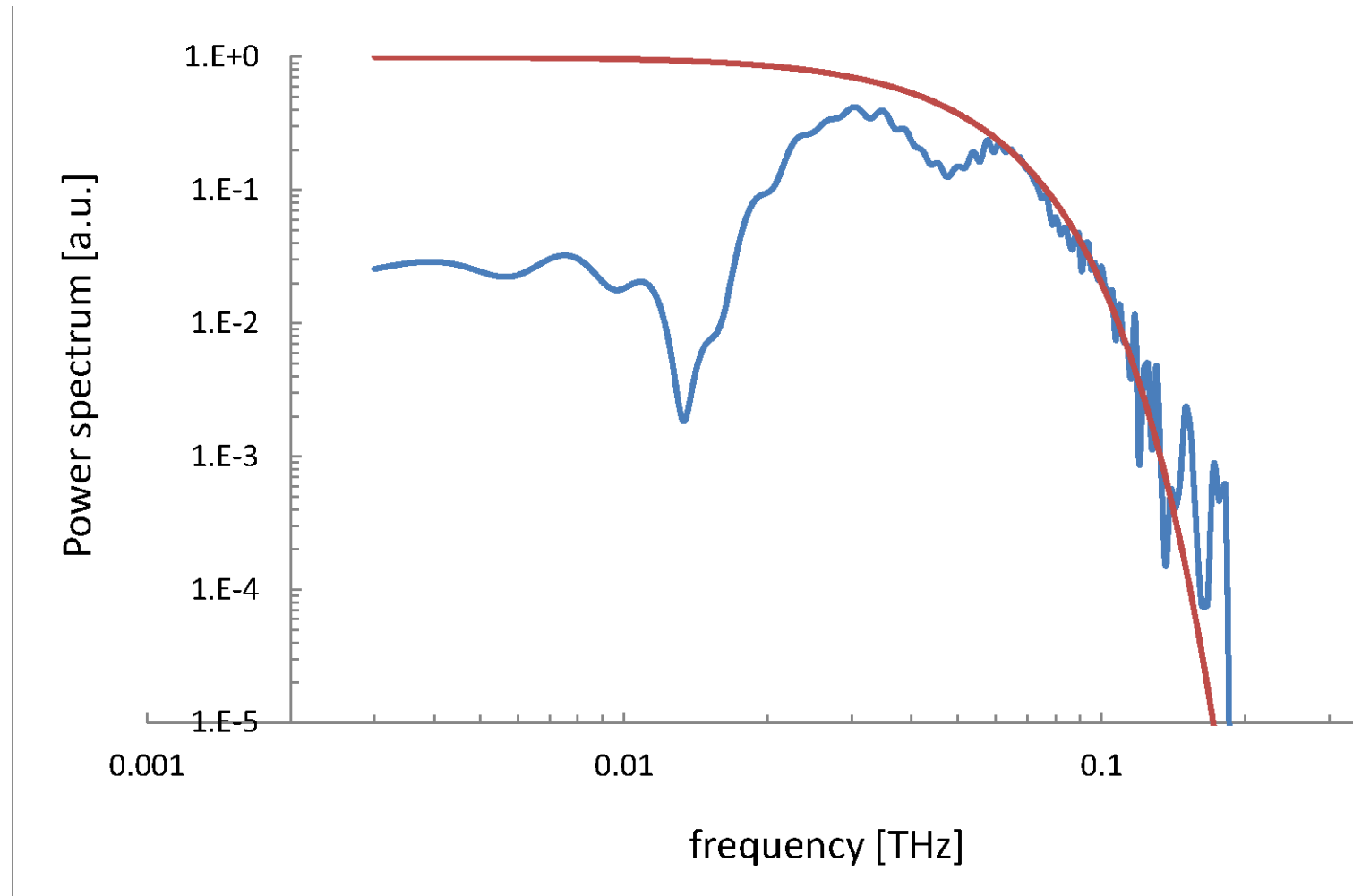
Beam power loss



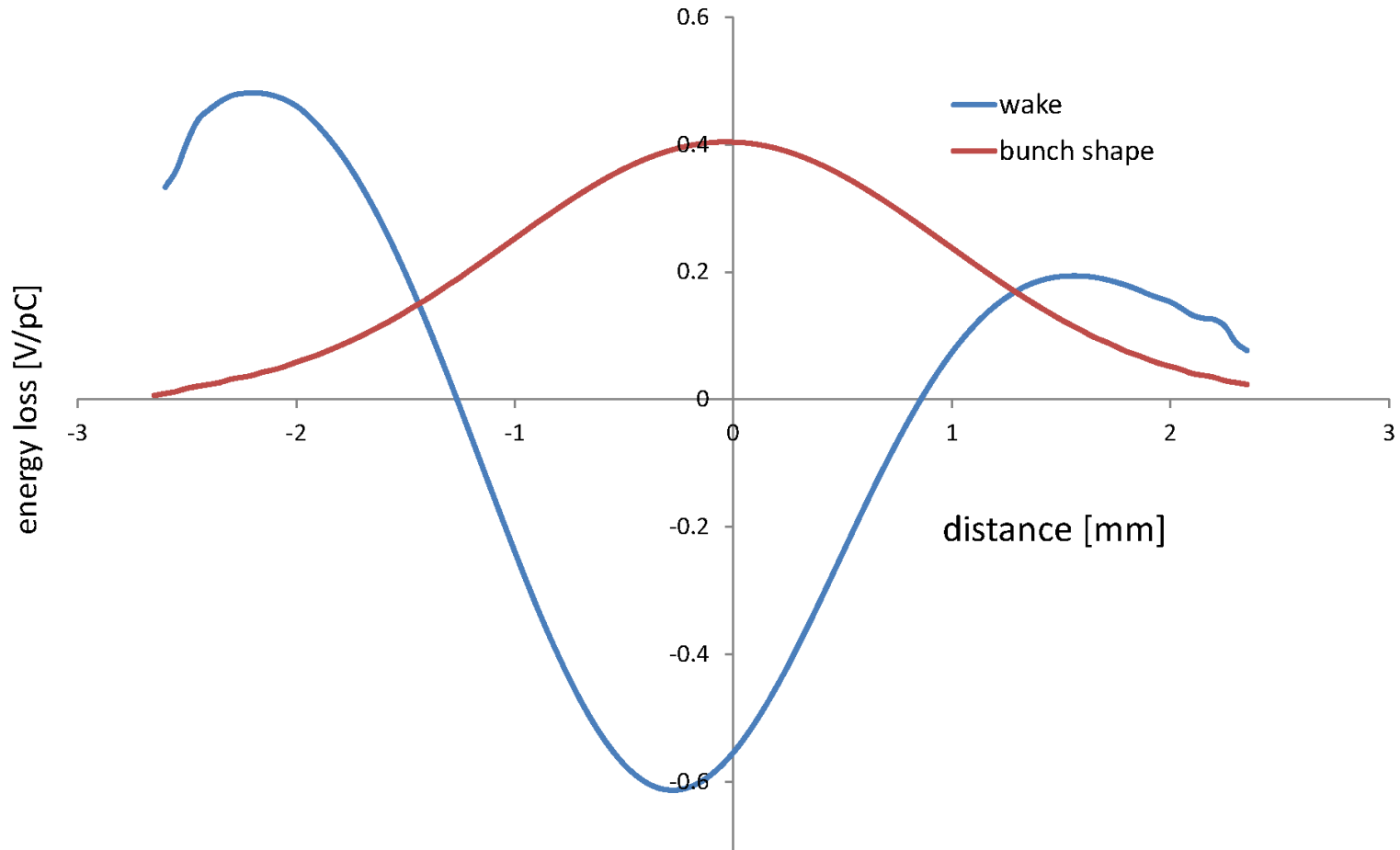
Beam energy loss



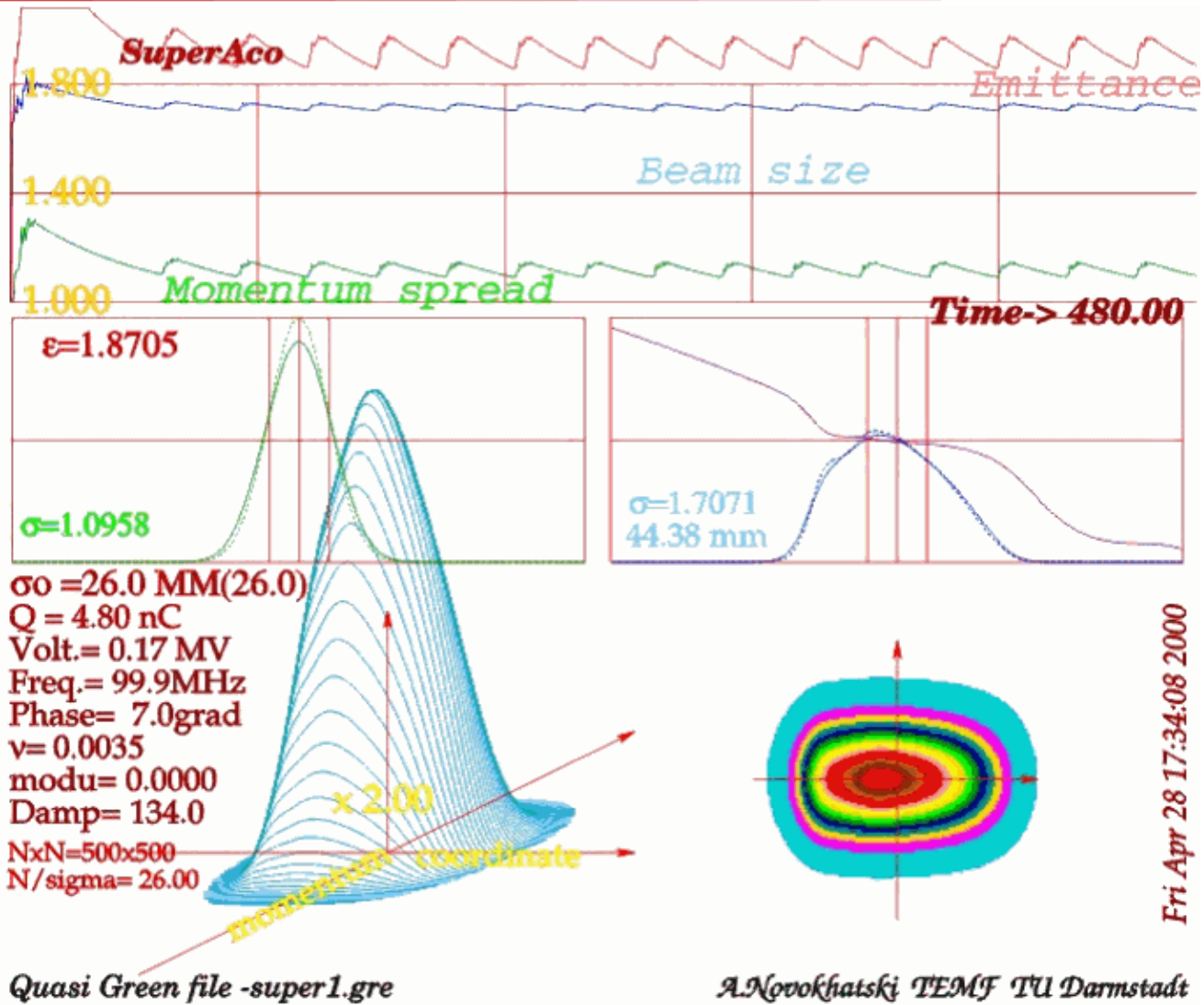
Spectrum of a 1mm bunch



Energy loss along a bunch



Next step: combine CSR and Fokker-Planck calculations.



Acknowledgments



The author would like to thank

Mike Sullivan and R. Clive Field

for help and valuable comments;

Franz-Josef Decker, Paul J. Emma and Yunhai Cai

for support and interest in this work;

Physicists of the SLAC Beam Physics Department

for useful discussions.

