

Double-Beta Decay Experiments

Outline

- Double-Beta Decay Basics
- Experiment Classifications
- Another 10 years
- Open for Discussion

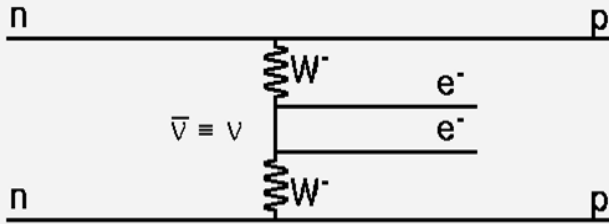
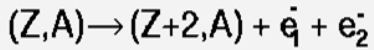
Most Recent Reviews

- J.J. Gomez-Cadenas et al., arXiv:1109.5515
- A.S. Barabash, arXiv:1104.2714
- A.S. Barabash, arXiv:1109.6423
- TAUP11 conference: <http://taup2011.mpp.mpg.de>

Physics topics in double-beta decay

- Theory of Flavour
 - Neutrino mass, Majorana or Dirac nature
 - Mass mechanism = physics at high energy scales
- Exotic interactions = Beyond SM interactions
 - (Scalar and tensor contributions to V-A from any BSM theory, Lepton number violation → Leptogenesis, ...)
- See the previous presentations

Double Beta Decay Experiments



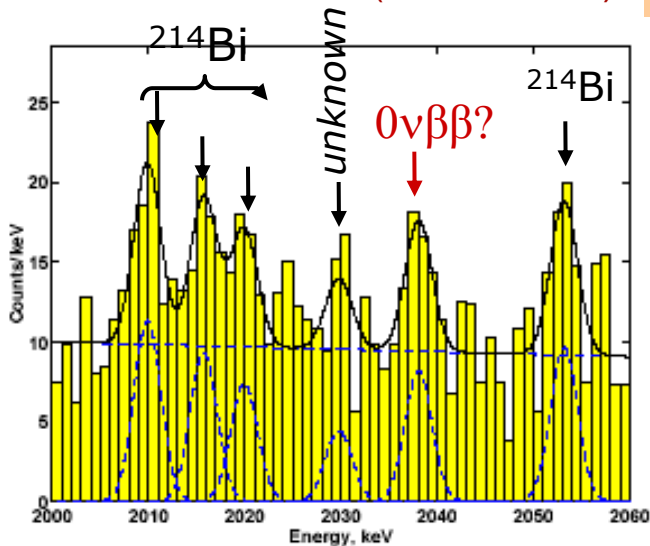
$$T_{1/2} \propto \epsilon \frac{a}{A} \sqrt{\frac{M \times t}{\Delta E \times B}}$$

$$\langle m_{\nu_e} \rangle \propto 1 / \sqrt{T_{1/2}}$$

Source = detector (calorimeters)

Two approaches

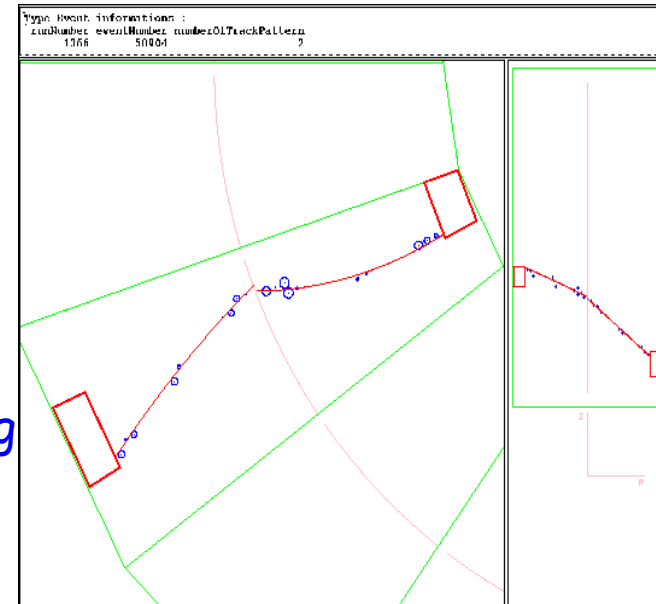
Source ≠ detector
(foil+tracking+calorimetry)



Isotope flexibility
"smoking" gun
Topological bkg suppression

e.g. Heidelberg-Moscow

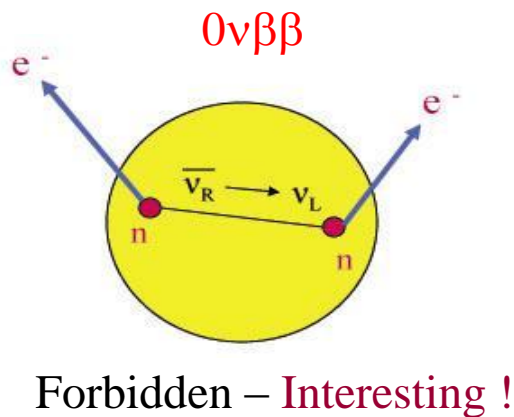
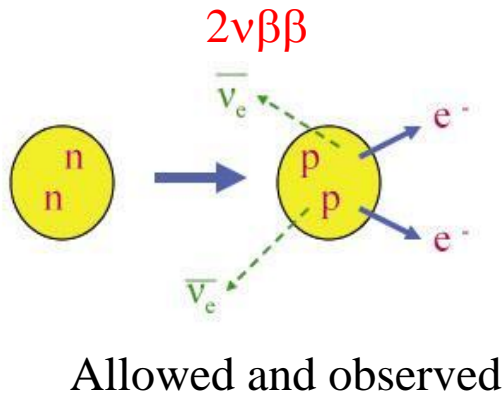
Great $\Delta E/E$
high efficiency



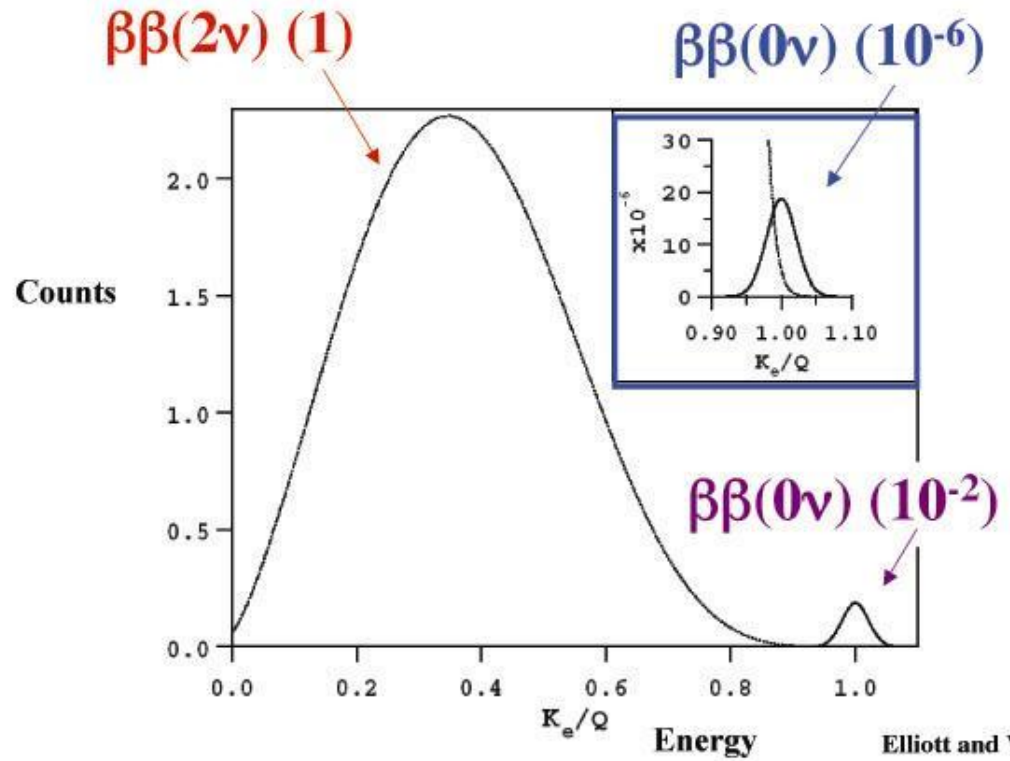
e.g. NEMO-III

Double Beta Decay

Double Beta Decay:
$$|\langle m \rangle| = \left| \sum_k U_{ek}^2 m_k \right|$$



Experimental signature



Neutrinos must be massive Majorana particles

Success and further improvements: HowTo

$$T_{1/2} \propto \varepsilon \frac{a}{A} \sqrt{\frac{M \times t}{\Delta E \times B}}$$

$$\langle m_{\nu_e} \rangle \propto 1 / \sqrt{T_{1/2}}$$

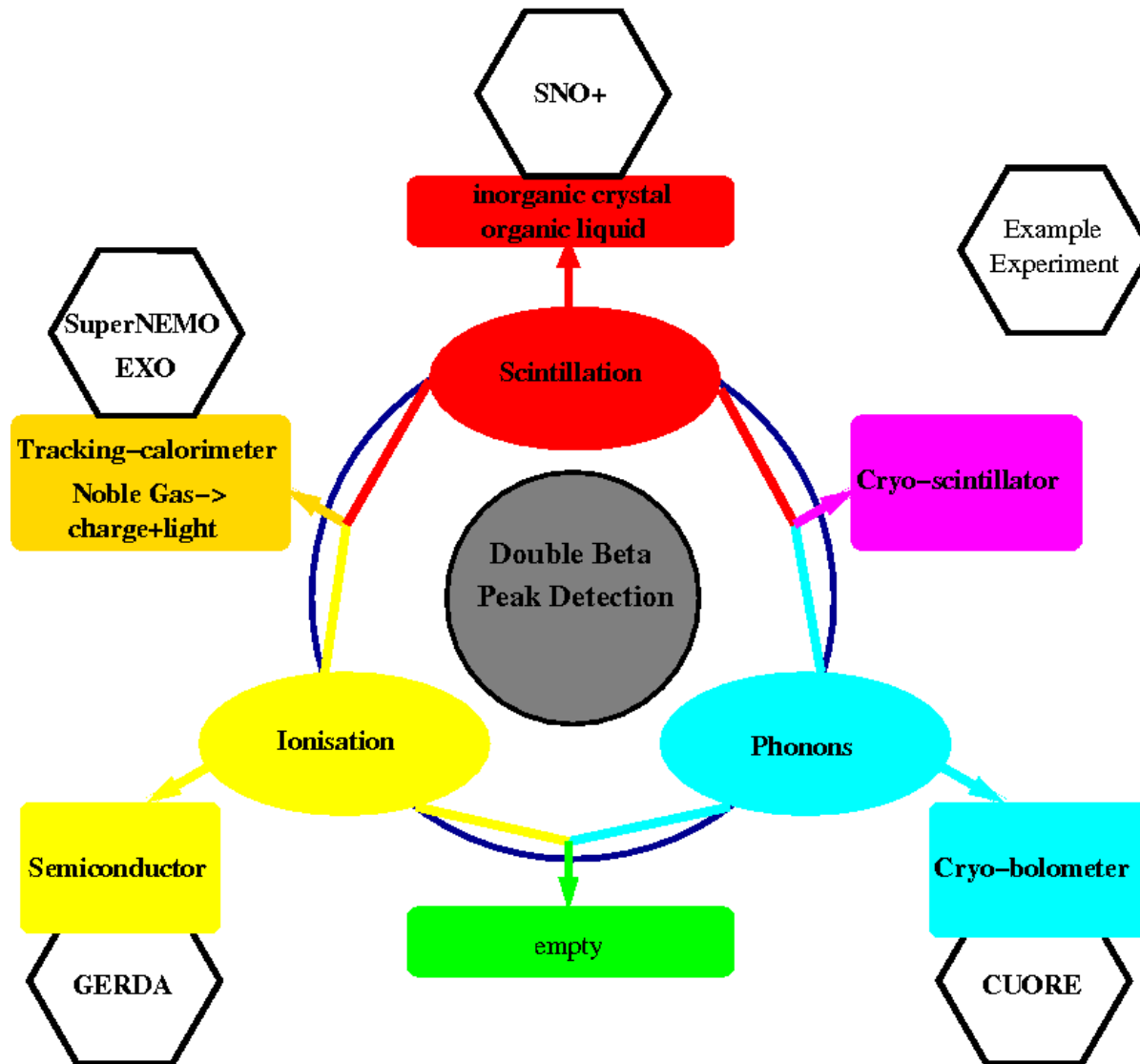
On the way to 100meV: Missing factor 2-5 gained by

- a) 16-625 fold increase in exposure (Mt) and/or
- b) 16-625 fold reduction of background, B

On the way to 10meV : Extrapolate existing experiments over
5-6 orders of magnitude !!!

Not at all hopeless, but a challenge !

Techniques



Experiment List

TABLE 3. Seven most developed and promising projects. Sensitivity at 90% C.L. for three (1-st step of GERDA and MAJORANA, SNO+, and KamLAND-Xe) five (EXO, SuperNEMO and CUORE) and ten (full-scale GERDA and MAJORANA) years of measurements is presented. M - mass of isotopes.

Experiment	Isotope	M, kg	Sensitivity $T_{1/2}, y$	Sensitivity $ \langle m_\nu \rangle , \text{meV}$	Status
CUORE [20]	^{130}Te	200	2.1×10^{26}	35–90	in progress
GERDA [21]	^{76}Ge	40	2×10^{26}	70–300	in progress
		1000	6×10^{27}	10–40	R&D
MAJORANA [22, 23]	^{76}Ge	30–60	$(1-2) \times 10^{26}$	70–300	in progress
		1000	6×10^{27}	10–40	R&D
EXO [24]	^{136}Xe	200	6.4×10^{25}	95–220	in progress
		1000	8×10^{26}	27–63	R&D
SuperNEMO [25, 26, 27]	^{82}Se	100–200	$(1-2) \times 10^{26}$	40–110	R&D
KamLAND-Xe [28]	^{136}Xe	400	4.5×10^{26}	40–80	in progress
		1000	$\sim 10^{27}$	25-50	R&D
SNO+ [29]	^{150}Nd	56	$\sim 4.5 \times 10^{24}$	100–300	in progress
		500	$\sim 3 \times 10^{25}$	40-120	R&D

A.S. Barabash, arXiv:1109.6423

A hypothetical $\beta\beta$ -experiment

All reviews agree on one point: There is no optimal $\beta\beta$ -experiment!

So, pick out what is good and compromise for a practical solution:

A highly subjective procedure

(A) Simple

detector technology

(best possible energy res.
highest efficiency)

$$T_{1/2} \propto \frac{\varepsilon a}{A} \sqrt{\frac{M \times t}{\Delta E \times B}}$$

(B) Best practical $\beta\beta$ -isotope

(high-Q value, enrichment costs)

(C) Lowest possible B

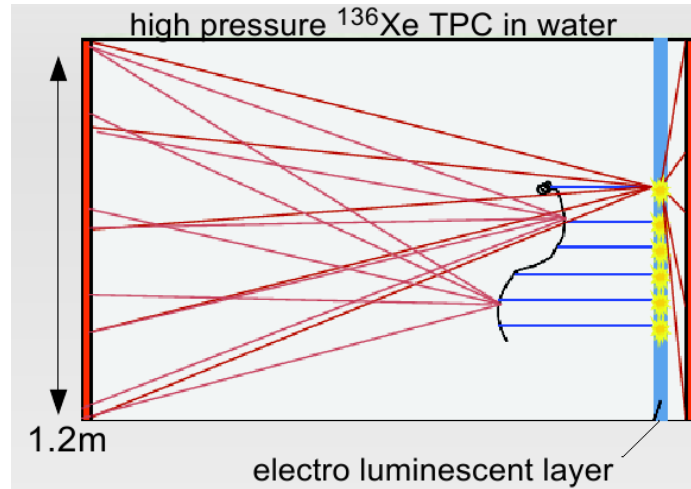
- Tracking
SuperNEMO, NEXT, COBRA
- Intrinsic cleanliness
GERDA, CUORE, EXO
- Active veto shielding
GERDA, EXO, SNO+

More than finding DBD

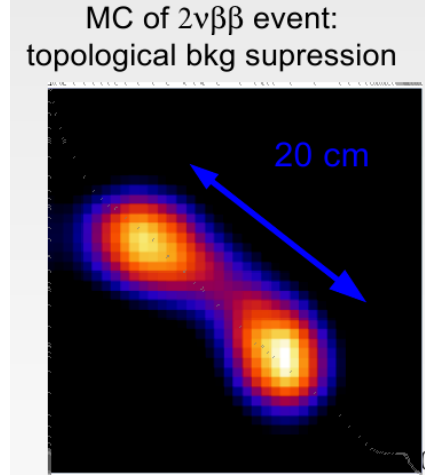
Tracking+
Calorimeter

TPC

NEXT



NEXT, arXiv:1106.3630



SuperNEMO



COBRA

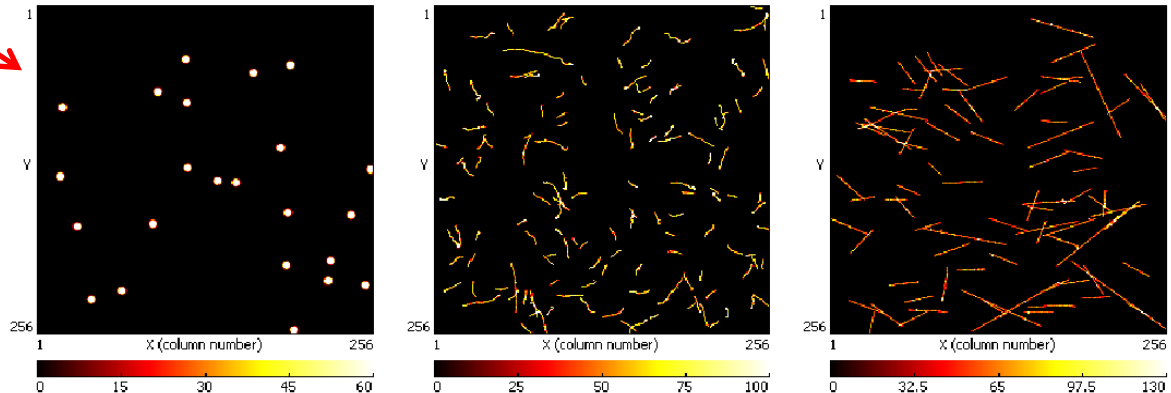


Fig. 4: Preselected sample of alpha (left) and beta particles (middle) as well as muons (right) obtained by COBRA with a $55\mu\text{m}$ pixel detector. Evident is the particle identification and the potential tracking option.

Energy resolution issues?

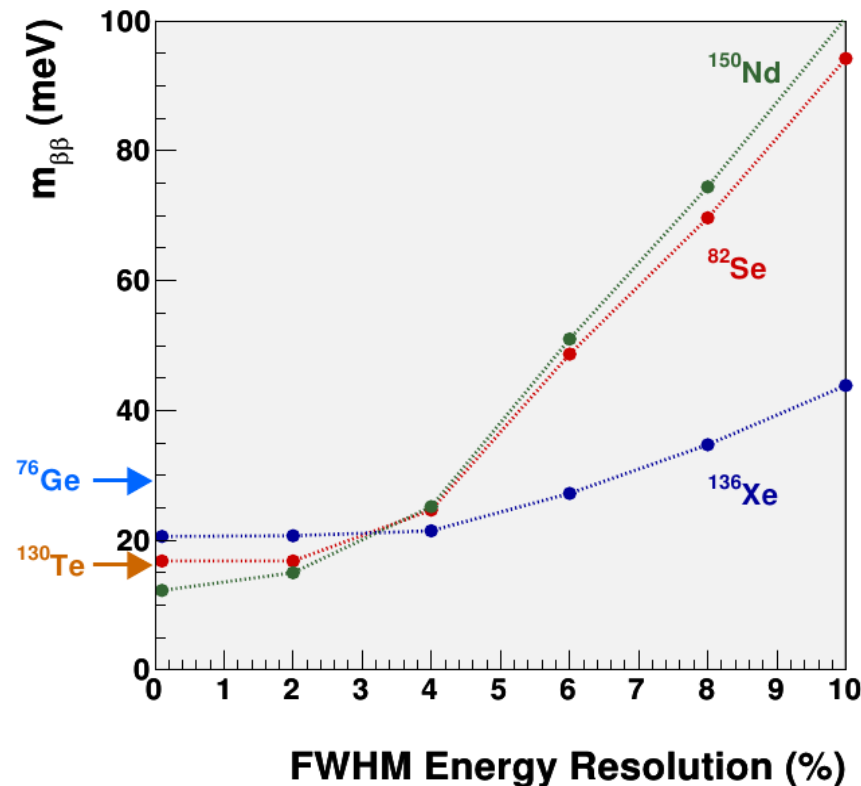
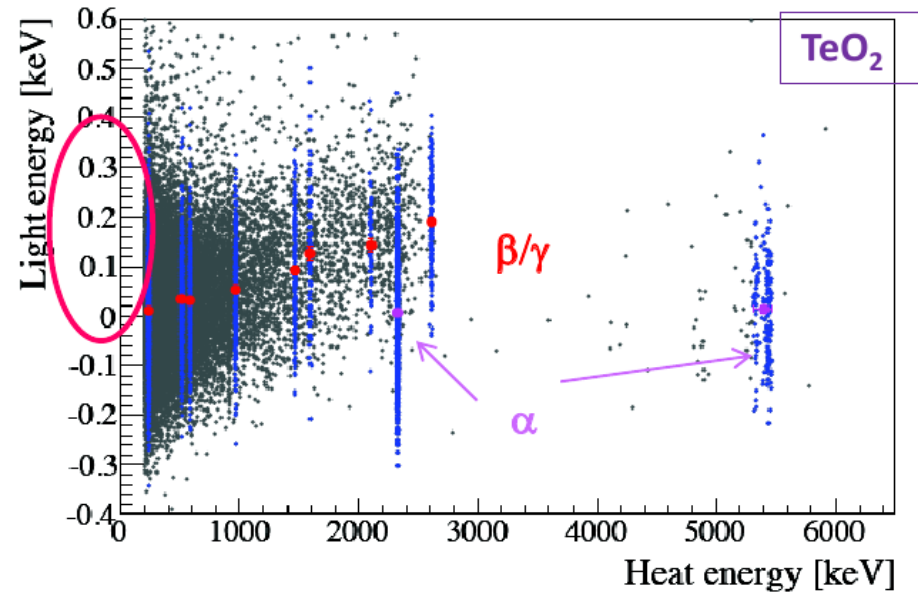


Fig. 18. – Sensitivity to $m_{\beta\beta}$ at 90% CL as a function of FWHM energy resolution, for ideal experiments using five different isotopes, each with $100 \text{ kg}_{\beta\beta}$ of $\beta\beta$ emitter mass and 5 years of data-taking. The experiments are assumed to have perfect efficiency and to be affected only by $\beta\beta 2\nu$ backgrounds. In practice, experiments using ^{76}Ge and ^{130}Te always feature an excellent energy resolution and are therefore not affected by $\beta\beta 2\nu$ backgrounds, hence only the background-free sensitivity limit is shown in those cases, with an arrow.

J.J. Gomez-Cadenas et al., arXiv:1109.5515

Exotic Backgrounds



Cardani, TAUP2011

- Rare α decays
- Rare β decays from short-lived daughter isotopes, see GERDA talk
- Neutrinos (de Barros, Zuber, arXiv:1103.5757)

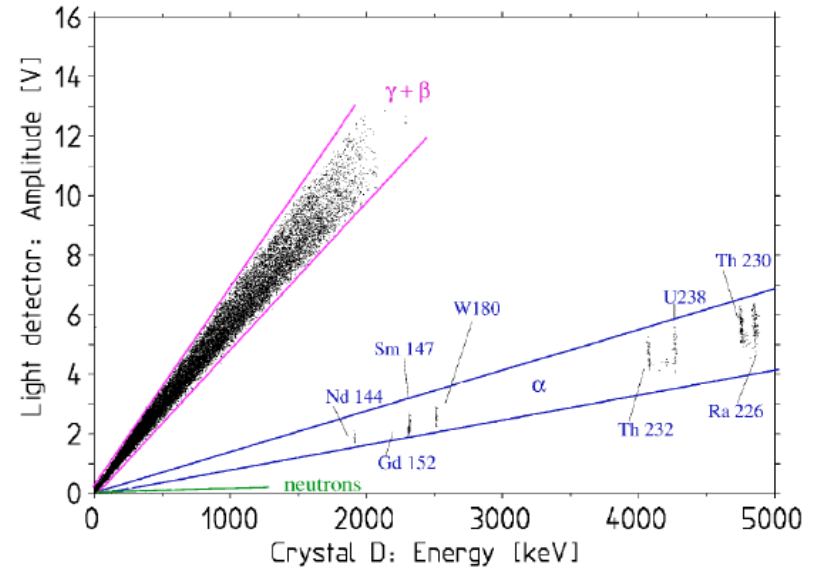
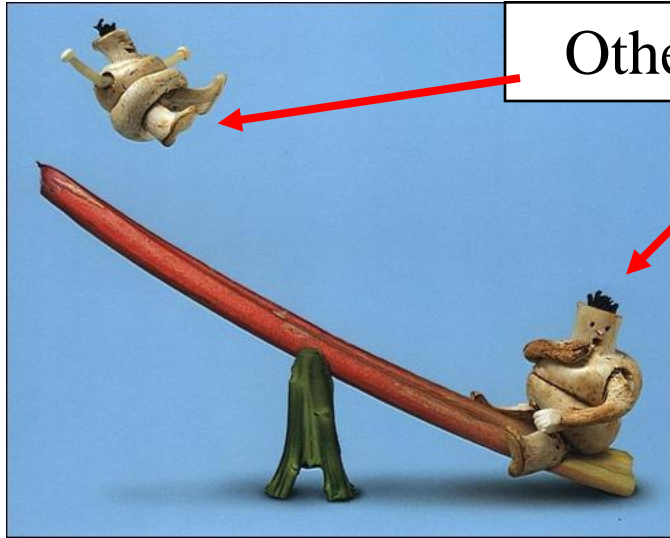


FIG. 1. Pulse amplitude in the light channel vs energy in the CaWO_4 crystal D from run 28. Three clearly separated event populations appear. The solid lines border the three different bands: the upper band is due to γ 's and β 's, the lowest one to neutrons and the middle one is due to α 's. Each peak in the α band has been identified (see text) and is here labeled accordingly.

Cozzini et al, Phys. Rev. C70, 064606

An X-ton device?

Other constraints: Background, ΔE , tracking



1 ton enriched
Double Beta
Material cost?

Least amount of compromises currently(!) results in an:
Enriched Xenon TPC (high pressure gas)
NEXT and(?) EXO

An X-ton device?

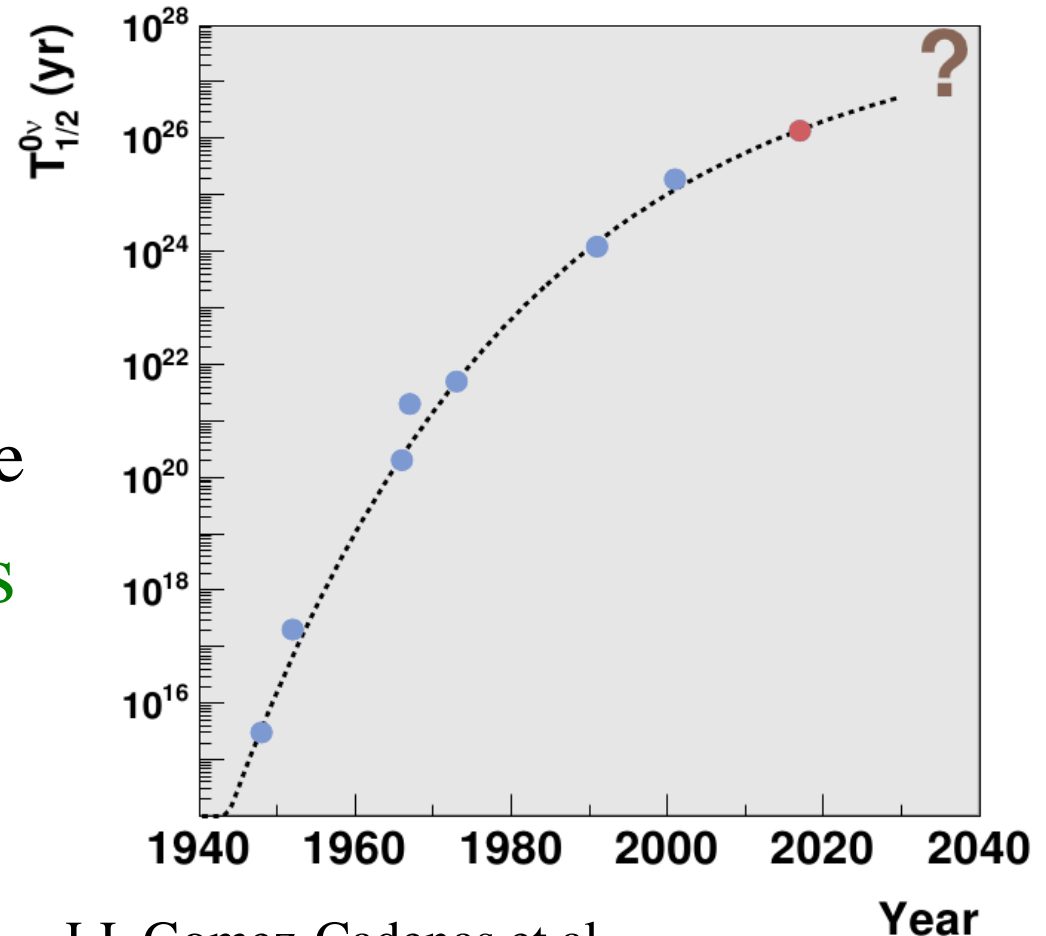
TABLE 7. Approximate price of 2β isotopes obtained by centrifugation. *) Taking into account 20% reduction for mass production case.

Isotope	Abundance	Price per kg, kS	Cost of 10 t, \$M
^{76}Ge	7.61	~ 80	800 (640)*)
^{82}Se	8.73	~ 120	1200 (1000)*)
^{100}Mo	9.63	~ 80	800 (640)*)
^{116}Cd	7.49	~ 180	1800 (1440)*)
^{130}Te	34.08	~ 20	200 (160)*)
^{136}Xe	8.87	$\sim 5-10$	50-100 (40-80)*)
^{150}Nd (?)	5.6	> 200	> 2000

A.S. Barabash, arXiv:1109.6423

Summary

- Exciting times ahead
 - New experiments starting
 - New backgrounds
 - Test of DBD evidence
- ‘Going beyond’ starts now – ideas most welcome



J.J. Gomez-Cadenas et al.,
arXiv:1109.5515