

# The Alignment of the BESIII Drift Chamber Using Cosmic-ray Data

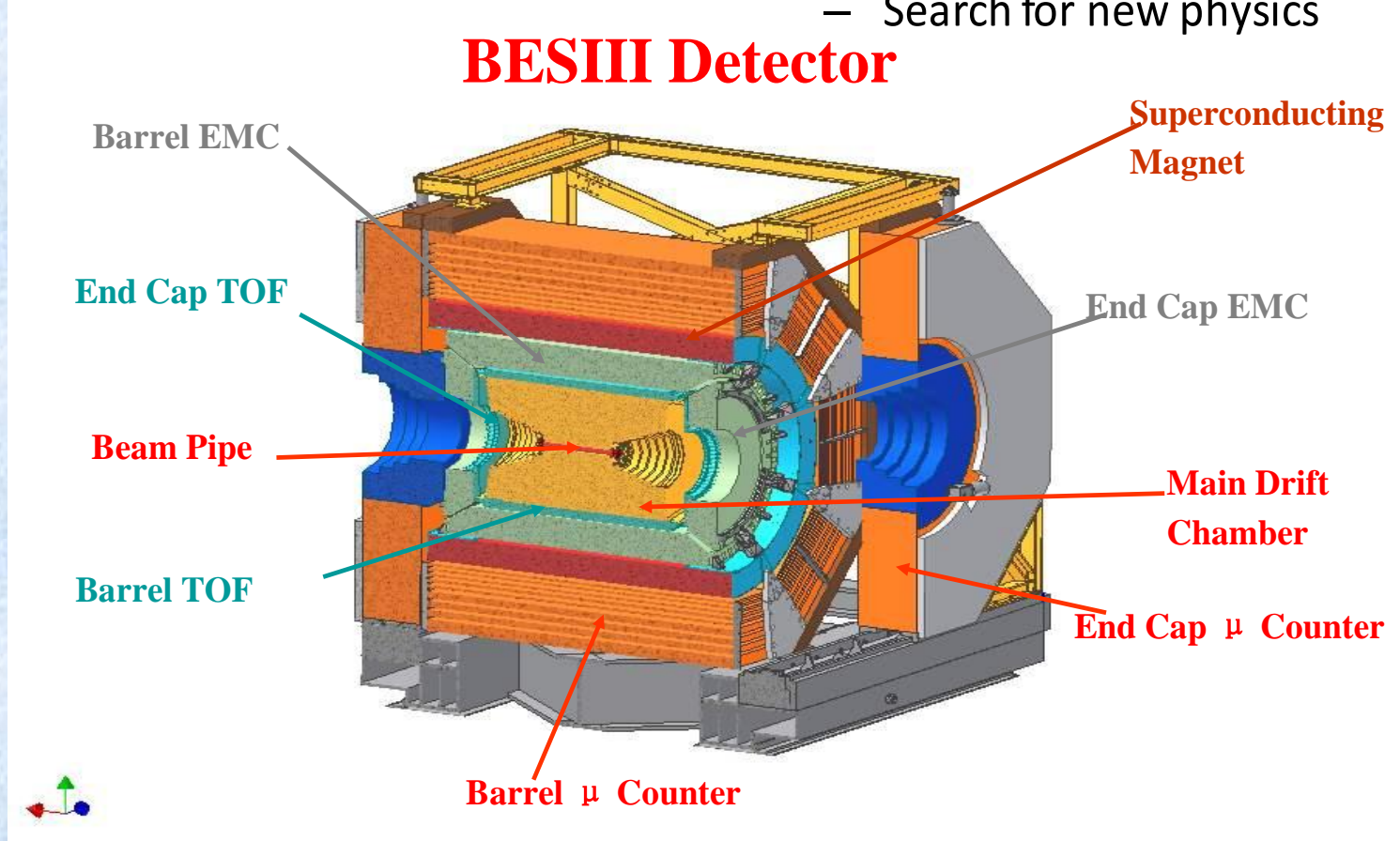
Wu Linghui

On behalf of BESIII Software Group

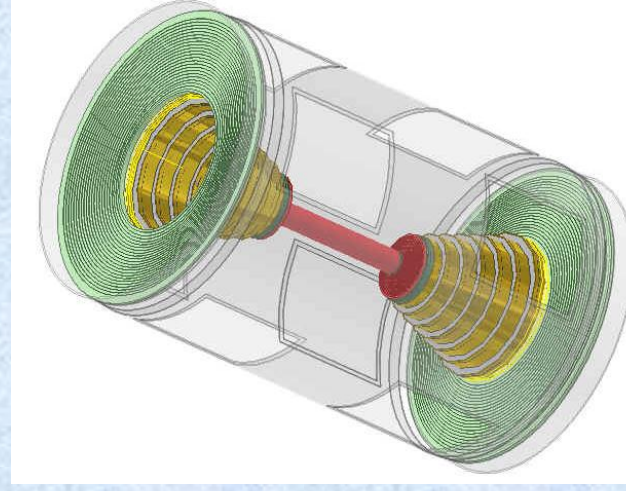
## BESIII Experiment

Physics goal:

- Precision measurement of CKM matrix
- Precision test of Standard Model
- QCD and hadron production
- Light hadron spectroscopy
- Charmonium physics
- Search for new physics

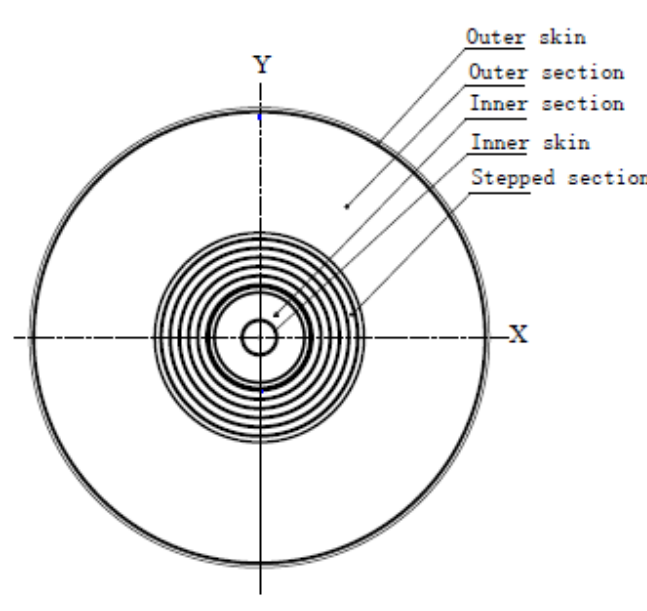
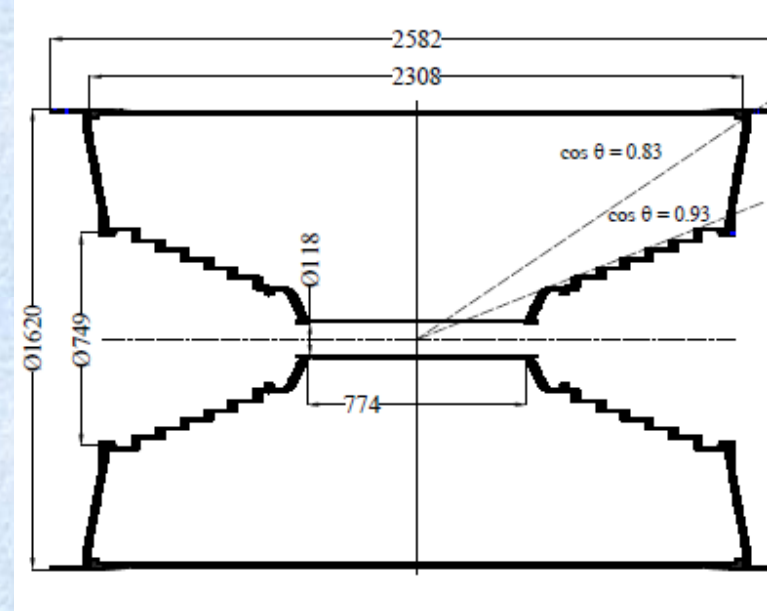
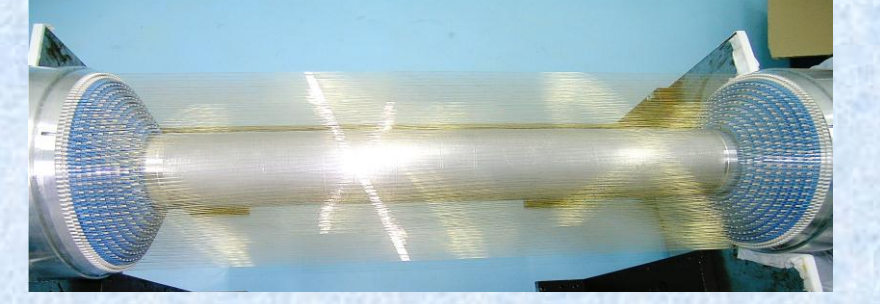


## Drift Chamber



- 2.6m long cylindrical chamber
- It consists of inner section, stepped section and outer section
- The stepped section is assembled from a set of 6 aluminum rings
- With an inner radius of 59mm and an outer radius of 810mm
- Filled with He/C<sub>3</sub>H<sub>8</sub>(60/40)

Inner section

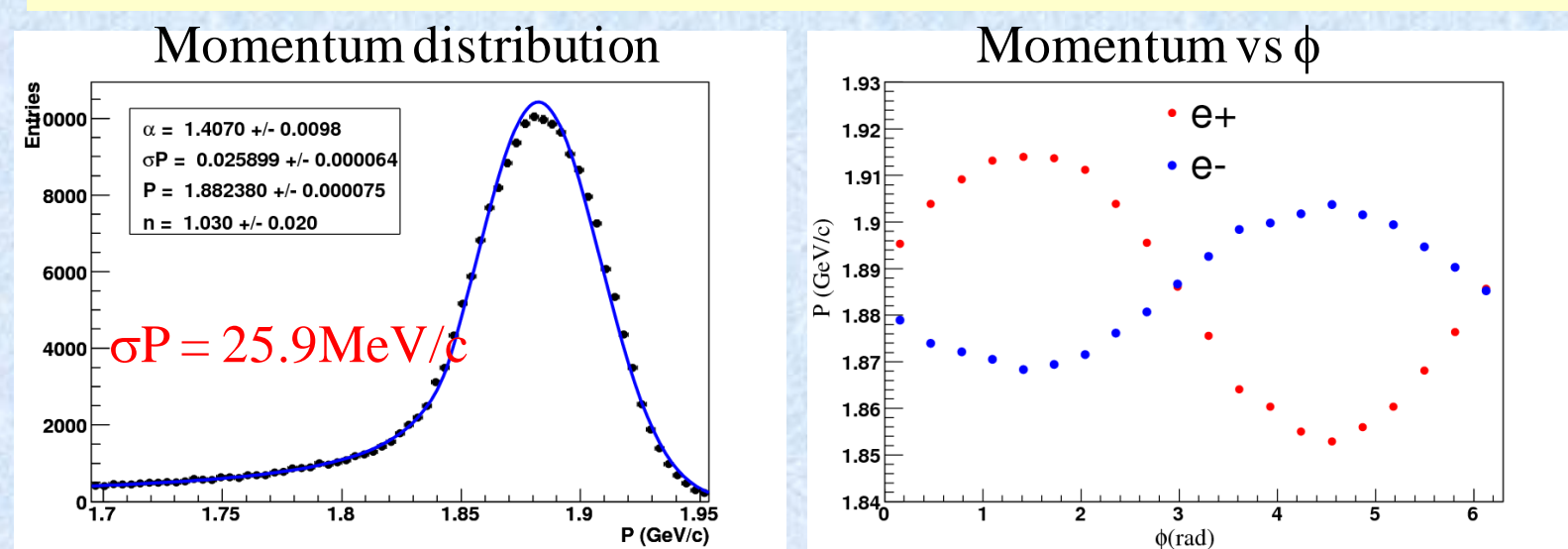


- Square cell
- 6796 sense wires and 21884 field wires.
- The average half-cell size is 6mm for the inner chamber and 8.1mm for the outer chamber
- Half cell staggering to resolve the left-right ambiguity

## Misalignment

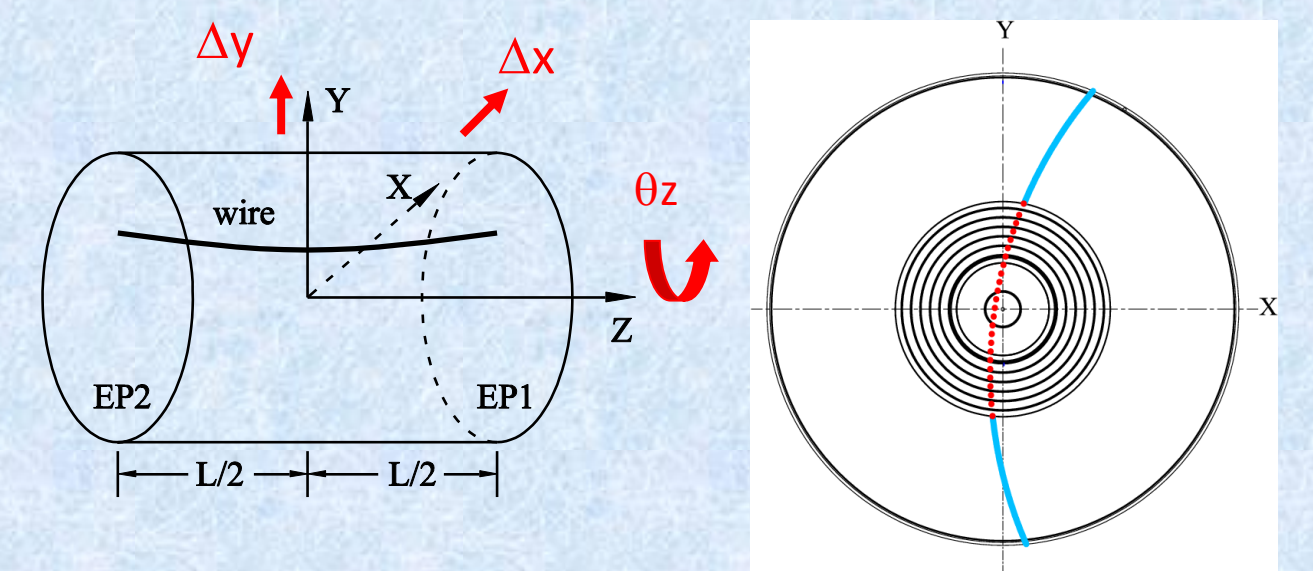
- Displacement of sub-endplates caused bad momentum resolution.
- Alignment with tracks is the only possible strategy to estimate positions and orientations of each component with sufficiently high precision.

### Misalignment in reconstruction of Bhabha events



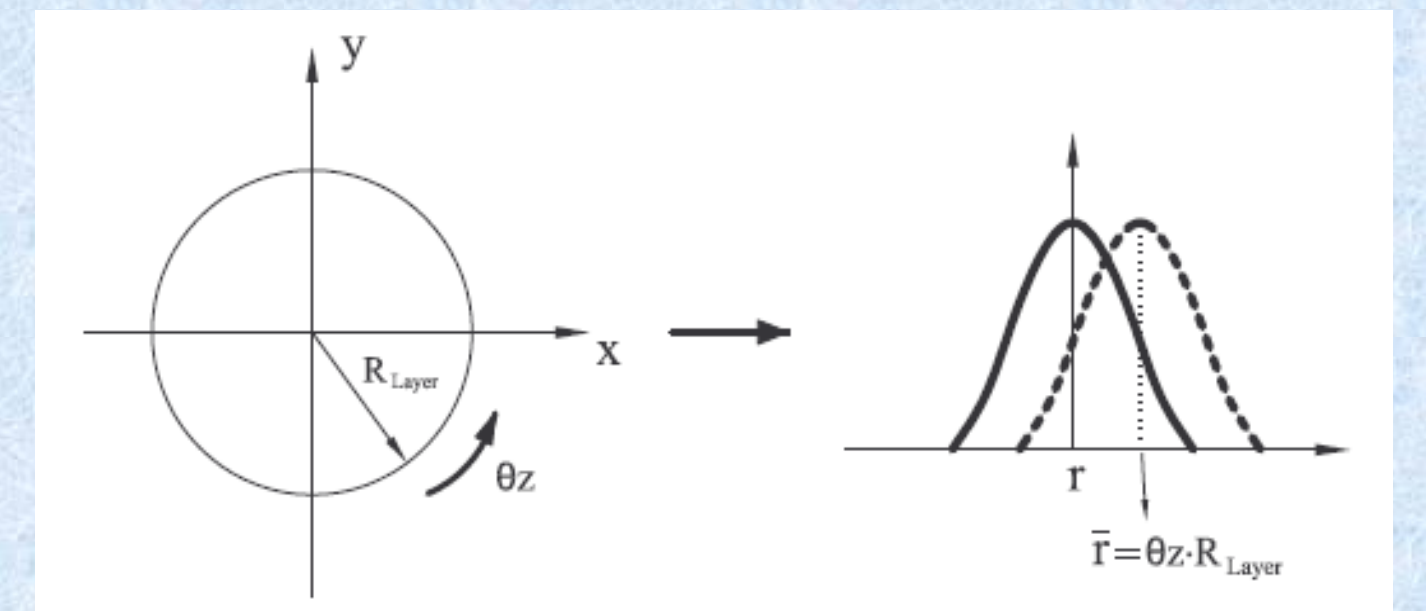
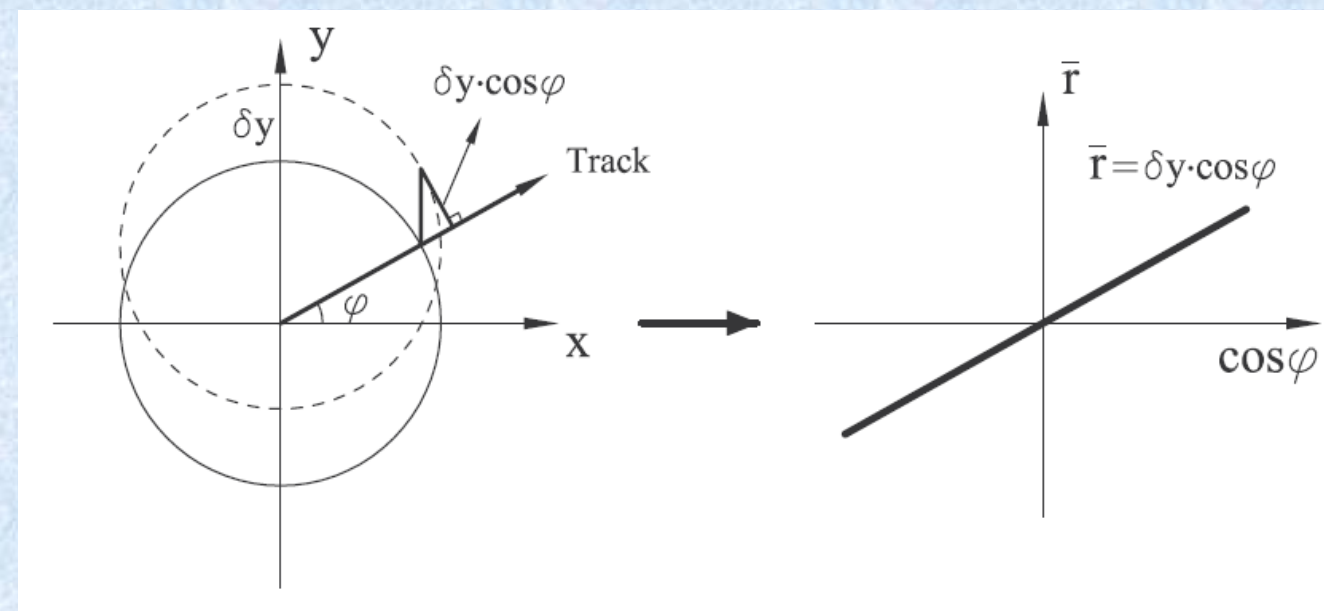
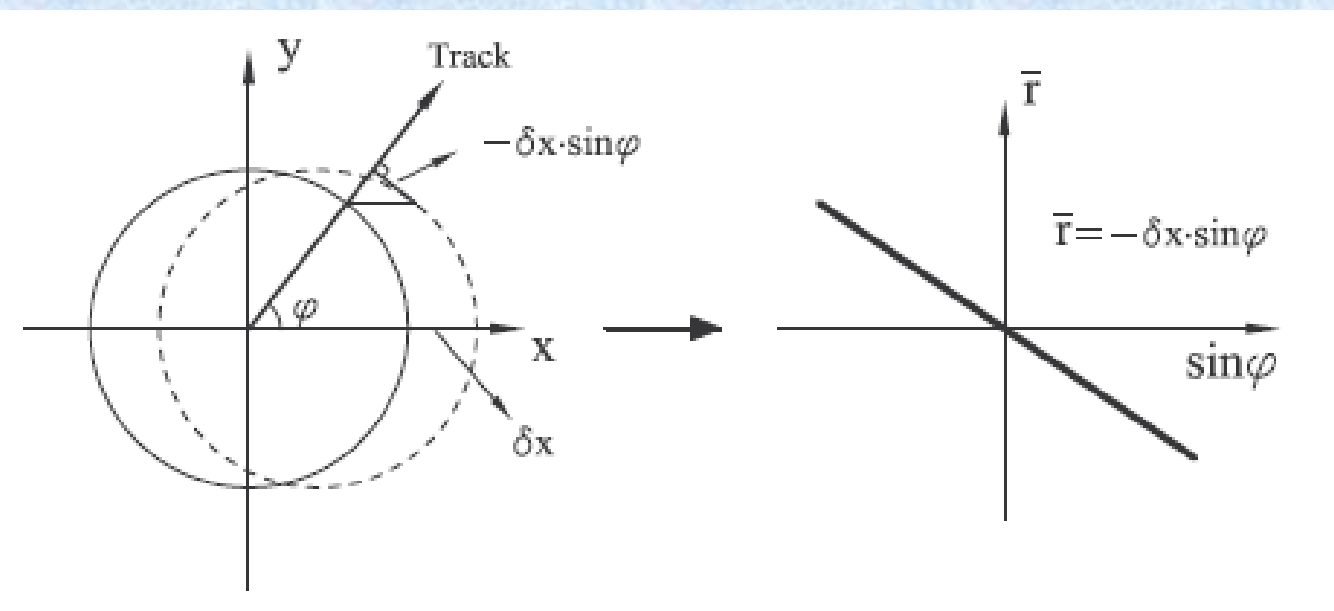
## Software Alignment

- Use cosmic-ray data to do preliminary alignment
- Alignment parameters
  - 16 independent sub-endplates
    - Inner section (x2)
    - Ring x 6 (x2)
    - Outer section (x2)
- For each component, 3 alignment parameters are considered
  - $\Delta x$ : Translation in x direction
  - $\Delta y$ : Translation in y direction
  - $\theta_z$ : Rotation around z axis
- Alignment methods
  - Use hits in the outer section to do track fit
  - Align the inner and stepped sections
  - Many iterations are necessary

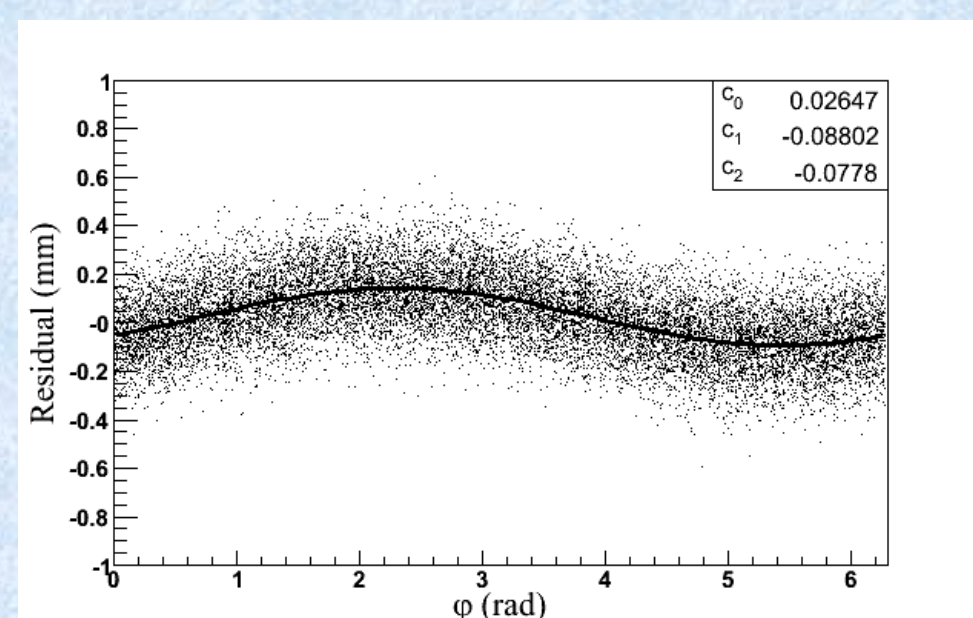


## Impact of Misalignment on Residuals

- Translation in x causes dependence of residual on  $\sin\phi$
- Translation in y causes dependence of residual on  $\cos\phi$
- Rotation in z causes shift of residuals which are independent of  $\phi$



### Residual vs phi (simulation)

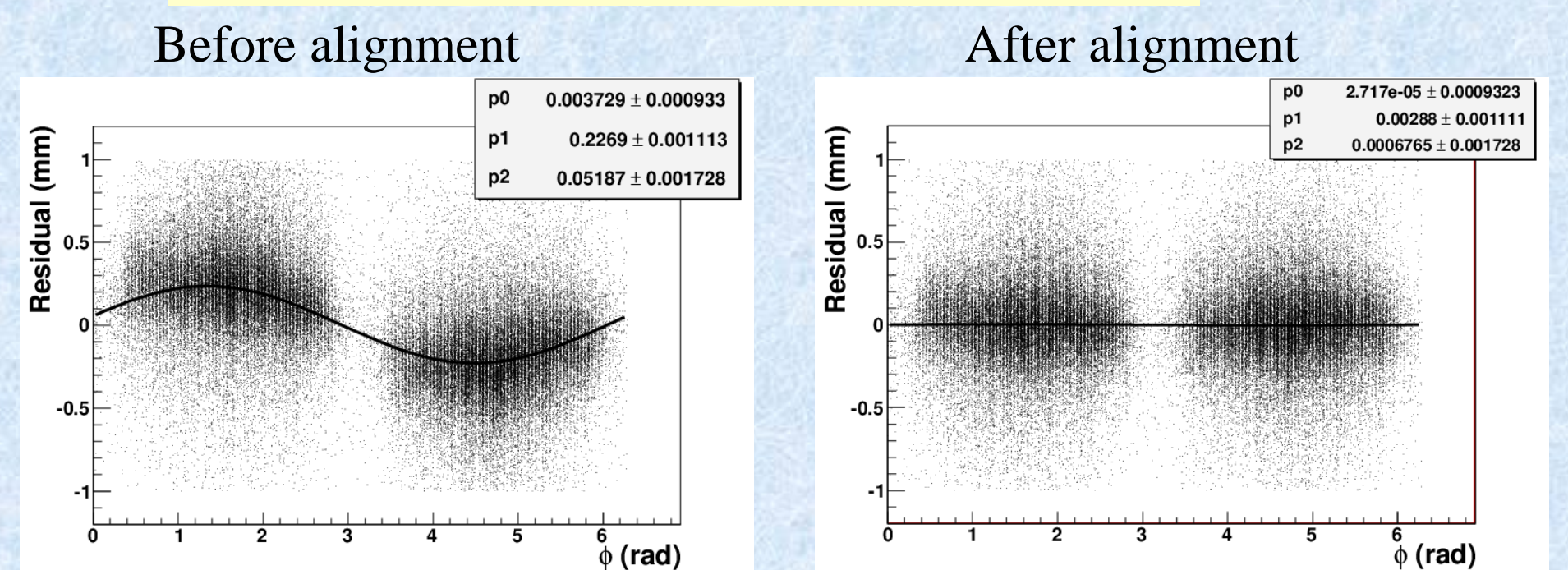


$$\text{So } r_{\text{mean}} = c_0 - c_1 \cdot \sin\phi + c_2 \cdot \cos\phi$$

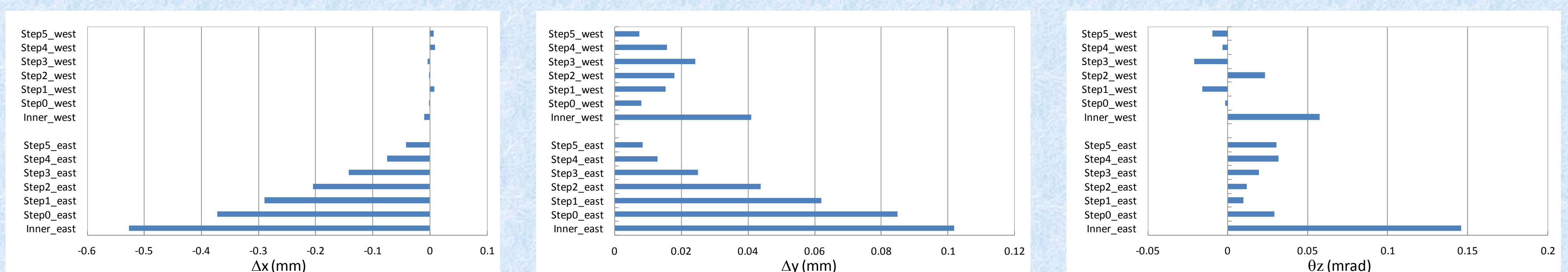
where

- $r_{\text{mean}}$  is the mean value of residuals
- $c_1$  and  $c_2$  are estimated values of  $\delta x$  and  $\delta y$ , respectively
- $\theta_z = c_0 / R_{\text{layer}}$  ( $R_{\text{layer}}$  is the radius of the layer)

### Residual vs phi (cosmic-ray data)

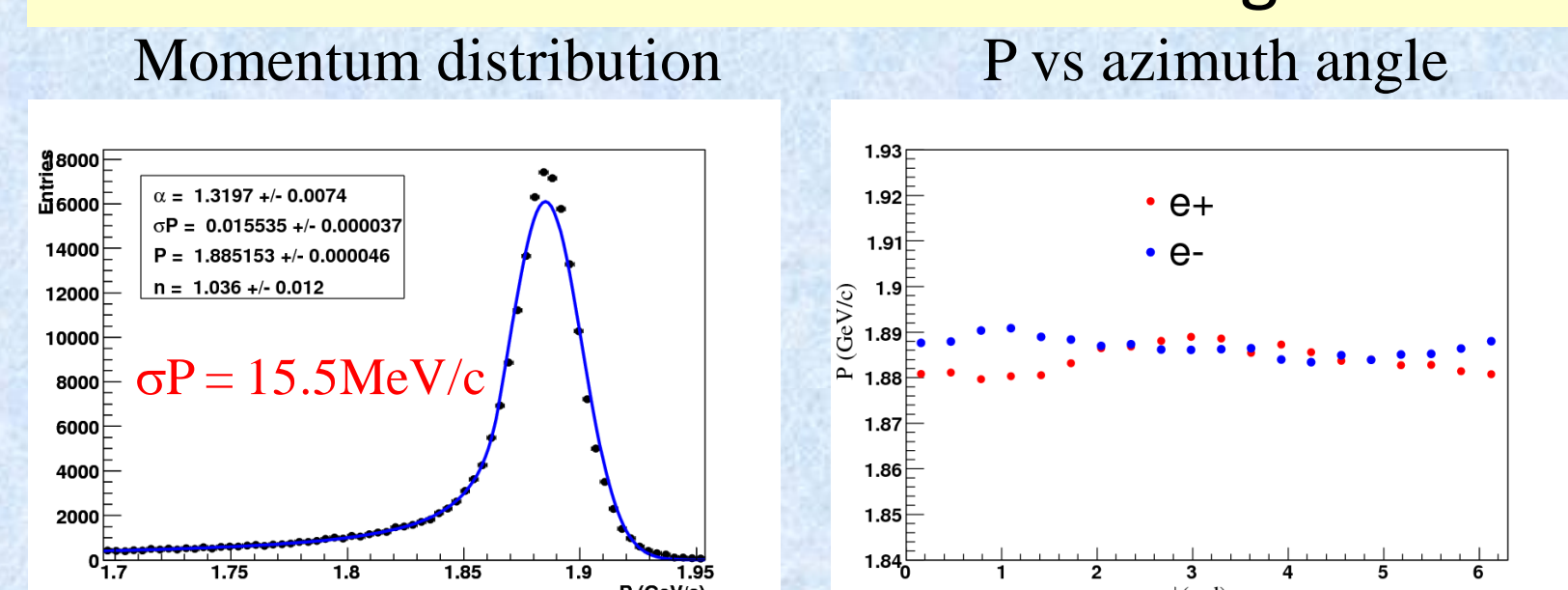


## Alignment Results

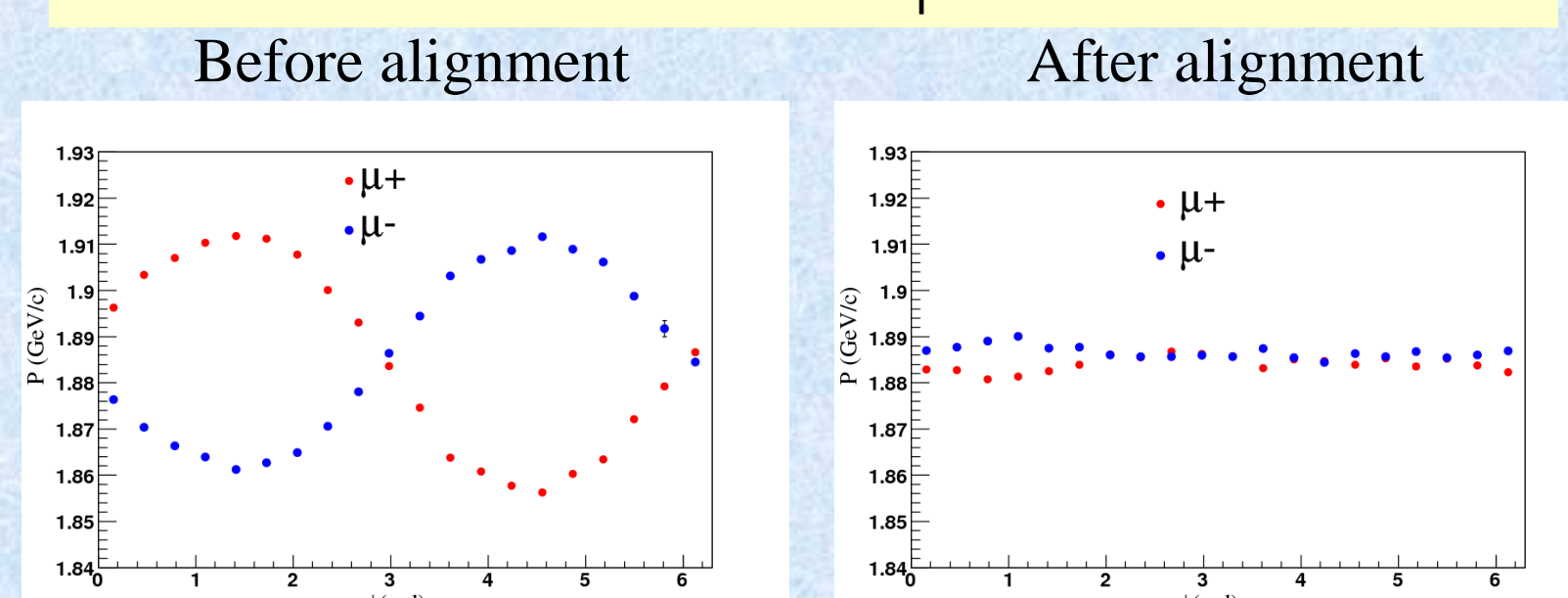


- The shift of the east sub-endplates in x direction is very large.

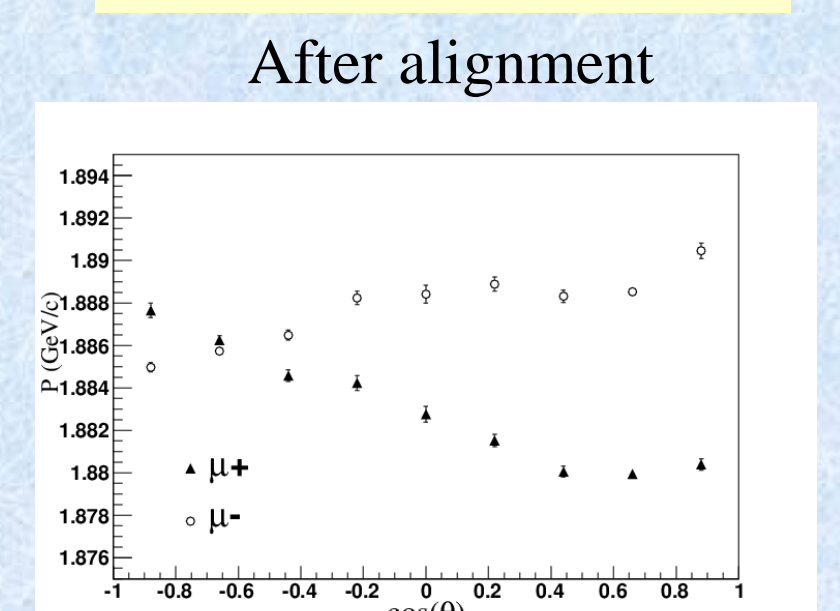
### Reconstruction of Bhabha after alignment



### P as a function of phi of dimuon



### P vs cos theta



## Conclusion and Outlook

- Use cosmic tracks to do preliminary alignment for the BESIII drift chamber. Estimate alignment parameters from the residual fits.
- Momentum resolution is improved significantly after alignment. But misalignment still exists.
- Begin to do alignment with high precision:
  - Use other alignment method: Millepede matrix method
  - Use other data samples: dimuon