

Special Relativity

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Reading

- W. Rindler: Introduction to Special Relativity (OUP 1991)
- D. Lawden: An Introduction to Tensor Calculus and Relativity
- N.M.J. Woodhouse: Special Relativity (Springer 2002)
- A.P. French: Special Relativity, MIT Introductory Physics Series (Nelson Thomes)
- · Misner, Thorne and Wheeler: Relativity
- C. Prior: Special Relativity, CERN Accelerator School (Zeegse)



Overview

- The principle of special relativity
- · Lorentz transformation and consequences
- · Space-time
- 4-vectors: position, velocity, momentum, invariants, covariance.
- Derivation of F=mc²
- · Examples of the use of 4-vectors
- Inter-relation between β and γ, momentum and energy
- · An accelerator problem in relativity
- · Photons and wave 4-vector
- · Radiation from an Accelerating Charge
- Motion faster than speed of light



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Historical Background

- Groundwork of Special Relativity laid by Lorentz in studies of electrodynamics, with crucial concepts contributed by Einstein to place the theory on a consistent footing.
- Maxwell's equations (1863) attempted to explain electromagnetism and optics through wave theory
 - light propagates with speed c = 3×10⁸ m/s in "ether" but with different speeds in other frames
 - the ether exists solely for the transport of e/m waves
 - Maxwell's equations not invariant under Galilean transformations
- · To avoid setting e/m apart from classical mechanics, assume
 - light has speed c only in frames where source is at rest
 - the ether has a small interaction with matter and is carried along with astronomical objects such as the Earth



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Contradicted by Experiment

- Aberration of star light (small shift in apparent positions of distant stars)
- · Fizeau's 1859 experiments on velocity of light in liquids
- Michelson-Morley 1907 experiment to detect motion of the earth through ether
- Suggestion: perhaps material objects contract in the direction of their motion

 $L(v) = L_0 \left(1 - \frac{v^2}{c^2} \right)^{1/2}$

This was the last gasp of ether advocates and the germ of Special Relativity led by Lorentz, Minkowski and Einstein.



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Simultaneity

• Two clocks A and B are synchronised if light rays emitted at the same time from A and B meet at the mid-point of AB

- Frame F' moving with respect to F. Events simultaneous in F cannot be simultaneous in F'.
- Simultaneity is **not absolute** but frame dependent.



The Principle of Special Relativity

- A frame in which particles under no forces move with constant velocity is *inertial*.
- Consider relations between inertial frames where measuring apparatus (rulers, clocks) can be transferred from one to another: related frames.
- · Assume:
 - Behaviour of apparatus transferred from F to F' is independent of mode of transfer
 - Apparatus transferred from F to F', then from F' to F", agrees with apparatus transferred directly from F to F".
- The Principle of Special Relativity states that all physical laws take equivalent forms in related inertial frames, so that we cannot distinguish between the frames.



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The Lorentz Transformation

- Must be linear to agree with standard Galilean transformation in low velocity limit
- · Preserves wave fronts of pulses of light,



i.e.
$$P \equiv x^2 + y^2 + z^2 - c^2 t^2 = 0$$

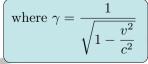
whenever $Q \equiv x'^2 + y'^2 + z'^2 - c^2 t'^2 = 0$

Solution is the Lorentz transformation from frame F
 (t,x,y,z) to frame F'(t',x',y',z') moving with velocity v
 along the x-axis:





$$t' = \gamma \left(t - \frac{vx}{c^2} \right)$$
$$x' = \gamma (x - vt)$$
$$y' = y$$
$$z' = z$$





Outline of Derivation

Set
$$t' = \alpha t + \beta x$$

 $x' = \gamma x + \delta t$
 $y' = \varepsilon y$
 $z' = \varsigma z$

Then
$$P = kQ$$

$$\Leftrightarrow c^2 t'^2 - x'^2 - y'^2 - z'^2 = k(c^2 t^2 - x^2 - y^2 - z^2)$$

$$\Rightarrow c^{2}(\alpha t + \beta x)^{2} - (\gamma x + \delta t)^{2} - \varepsilon^{2} y^{2} - \varsigma^{2} z^{2} = k(c^{2} t^{2} - x^{2} - y^{2} - z^{2})$$

Equate coefficients of x, y, z, t.

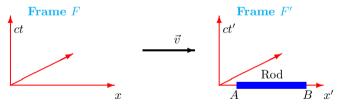
Isotropy of space $\Rightarrow k = k(\vec{v}) = k(|\vec{v}|) = \pm 1$

Apply some common sense (e.g. ε , ζ , k = +1 and not -1)



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Consequences: length contraction



A rod AB of length L', fixed in frame F' at x'_A , x'_B . What is its length measured in F?

Must measure position of ends in F at same time, so events in F are (ct, x_A) and $ct, x_B)$.

By Lorentz:

$$\begin{vmatrix} x'_A & = & \gamma(x_A - vt) \\ x'_B & = & \gamma(x_B - vt) \end{vmatrix} \implies \begin{aligned} L' & = & x'_B - x'_A \\ & = & \gamma(x_B - x_A) \\ & = & \gamma L > L \end{aligned}$$

Moving objects appear contracted in the direction of the motion



General 3D form of Lorentz Transformation:

$$\vec{x}' = \vec{x} - \vec{v} \left(\gamma t - (\gamma - 1) \frac{\vec{v} \cdot \vec{x}}{v^2} \right)$$

$$t' = \gamma \left(t - \frac{\vec{v} \cdot \vec{x}}{c^2} \right)$$



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Consequences: time dilation

• Clock in frame F at point with coordinates (x,y,z) at different times t_A and t_B



• In frame F' moving with speed v, Lorentz transformation gives

$$t'_A = \gamma \left(t_A - \frac{vx}{c^2} \right)$$
 $t'_B = \gamma \left(t_B - \frac{vx}{c^2} \right)$

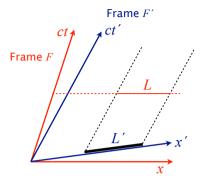
So

$$\Delta t' = t'_B - t'_A = \gamma \Big(t_B - t_A \Big) = \gamma \Delta t > \Delta t$$

Moving clocks appear to run slow

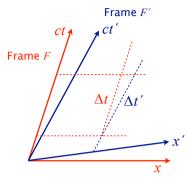


Schematic Representation of the Lorentz Transformation



Length contraction L < L'

Rod at rest in F'. Measurements in F at a fixed time t, along a line parallel to x-axis



Time dilation $\Delta t < \Delta t'$

Clock at rest in F. Time difference in F' from line parallel to t'-axis



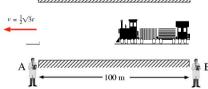
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Question 1



A's clock and the driver's clock read zero as the driver exits tunnel.

What does B's clock read when the guard goes in?

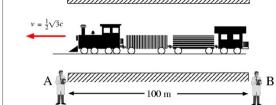


Moving train length 50m, so driver has still 50m to travel before he exits and his clock reads 0. A's clock and B's clock are synchronised. Hence the reading on B's clock is

$$-\frac{50}{v} = -\frac{100}{\sqrt{3}c} \approx -200 \,\mathrm{ns}$$



Example: High Speed Train



All clocks synchronised.

A's clock and driver's clock read 0 as front of train emerges from tunnel.

 Observers A and B at exit and entrance of tunnel say the train is moving, has contracted and has length

$$\frac{100}{\gamma} = 100 \times \left(1 - \frac{v^2}{c^2}\right)^{1/2} = 100 \times \left(1 - \frac{3}{4}\right)^{1/2} = 50 \text{m}$$

 But the tunnel is moving relative to the driver and guard on the train and they say the train is 100 m in length but the tunnel has contracted to 50 m



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Question 2



What does the guard's clock read as he goes in?

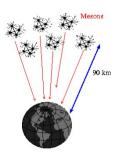


To the guard, tunnel is only 50m long, so driver is 50m past the exit as guard goes in. Hence clock reading is

$$+\frac{50}{v} = +\frac{100}{\sqrt{3}c} \approx +200 \,\text{ns}$$



Example: Cosmic Rays



- Muons are created in the upper atmosphere, 90km from earth. Their half life is τ =2 μ s, so they can travel at most 2 \times 10-6c=600 m before decaying. So how do more than 50% reach the earth's surface?
- Muons see distance contracted by γ , so

$$v\tau \approx \frac{90}{\gamma} \, km$$

• Earthlings say muons' clocks run slow so their half-life is $\gamma \tau$ and

$$v(\gamma \tau) \approx 90 \, km$$

Both give

$$\frac{\gamma v}{c} = \frac{90 \, km}{c\tau} = 150, \quad v \approx c, \quad \gamma \approx 150$$



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4-Vectors

The Lorentz transformation can be written in matrix form as



Lorentz matrix *L*

Position 4-vector X

An object made up of 4 elements which transforms like X is called a 4-vector (analogous to the 3-vector of classical mechanics)



Space-time

- An invariant is a quantity that has the same value in all inertial frames.
- Lorentz transformation is based on invariance of

$$c^{2}t^{2} - (x^{2} + y^{2} + z^{2}) = (ct)^{2} - \vec{x}^{2}$$

- 4D-space with coordinates (t,x,y,z) is called space-time and the point (t,x,y,z)=(t,x) is called an event.
- · Fundamental invariant (preservation of speed of light):

$$\Delta s^2 = c^2 \Delta t^2 - \Delta x^2 - \Delta y^2 - \Delta z^2 = c^2 \Delta t^2 \left(1 - \frac{\Delta x^2 + \Delta y^2 + \Delta z^2}{c^2 \Delta t^2} \right)$$
$$= c^2 \Delta t^2 \left(1 - \frac{v^2}{c^2} \right) = c^2 \left(\frac{\Delta t}{\gamma} \right)^2$$

 $au=\int rac{\mathrm{d}t}{\gamma}$ is called proper time, the time in the instantaneous rest-frame and an invariant. $\Delta \mathbf{s}$ is called the separation between two events.



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4-Vector Invariants

Basic invariant:

$$c^{2}t^{2}-x^{2}-y^{2}-z^{2}=(ct,x,y,z)\left(\begin{array}{cccc} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{array}\right)\left(\begin{array}{c} ct \\ x \\ y \\ z \end{array}\right)=X^{t}gX=X\cdot X$$

Inner product of two four vectors $A = (a_0, \vec{a}), B = (b_0, \vec{b})$:

$$A \cdot B = A^T q B = a_0 b_0 - a_1 b_1 - a_2 b_2 - a_3 b_3 = a_0 b_0 - \vec{a} \cdot \vec{b}$$

Invariance:

$$A' \cdot B' = (LA)^T g(LB) = A^T (L^T gL)B = A^T gB = A \cdot B$$



4-Vectors in S.R. Mechanics

• Velocity: $V=\frac{\mathrm{d}X}{\mathrm{d}\tau}=\gamma\frac{\mathrm{d}X}{\mathrm{d}t}=\gamma\frac{\mathrm{d}}{\mathrm{d}t}(ct,\vec{x})=\gamma(c,\vec{v})$

• Note invariant: $V\cdot V=\gamma^2(c^2-\vec{v}^2)=rac{c^2-\vec{v}^2}{1-\vec{v}^2/c^2}=c^2$

• Momentum: $P = m_0 V = m_0 \gamma(c, \vec{v}) = (mc, \vec{p})$

 $m = m_0 \gamma$ is the relativistic mass $p = m_0 \gamma \vec{v} = m \vec{v}$ is the relativistic 3-momentum



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Einstein's Relation: Energy and Mass

· Momentum invariant $P \cdot P = m_0^2 V \cdot V = m_0^2 c^2$

· Differentiate

$$P \cdot \frac{\mathrm{d}P}{\mathrm{d}\tau} \implies V \cdot \frac{\mathrm{d}P}{\mathrm{d}\tau} = 0 \implies V \cdot F = 0$$
$$\implies \gamma(c, \vec{v}) \cdot \gamma \left(c \frac{\mathrm{d}m}{\mathrm{d}t}, \vec{f} \right) = 0$$
$$\implies \frac{\mathrm{d}}{\mathrm{d}t} (mc^2) - \vec{v} \cdot \vec{f} = 0$$



 $\vec{v} \cdot \vec{f}$ = rate at which force does work = rate of change of kinetic energy

Therefore kinetic energy is

 $T = mc^2 + \text{constant} = m_0c^2(\gamma - 1)$

E=mc² is total energy



4-Force

From Newton's 2nd Law expect 4-Force given by

$$F = \frac{\mathrm{d}P}{\mathrm{d}\tau} = \gamma \frac{\mathrm{d}P}{\mathrm{d}t}$$

$$= \gamma \frac{\mathrm{d}}{\mathrm{d}t} (mc, \vec{p}) = \gamma \left(c \frac{\mathrm{d}m}{\mathrm{d}t}, \frac{\mathrm{d}\vec{p}}{\mathrm{d}t} \right)$$

$$= \gamma \left(c \frac{\mathrm{d}m}{\mathrm{d}t}, \vec{f} \right)$$

Note: 3-force equation: $\vec{f} = \frac{\mathrm{d}\vec{p}}{\mathrm{d}t} = m_0 \frac{\mathrm{d}}{\mathrm{d}t} (\gamma \vec{v})$



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Summary of 4-Vectors

Position $X = (ct, \vec{x})$

Velocity $V = \gamma(c, \vec{v})$

Momentum $P = m_0 V = m(c, \vec{v}) = \left(\frac{E}{c}, \vec{p}\right)$

Force $F = \gamma \left(c \frac{\mathrm{d}m}{\mathrm{d}t}, \vec{f} \right) = \gamma \left(\frac{1}{c} \frac{\mathrm{d}E}{\mathrm{d}t}, \vec{f} \right)$



Example of Transformation: Additionof Velocities

An object has velocity $\vec{u} = (u_x, u_y)$ in frame F', which moves with velocity $\vec{v} = (v, 0)$ with respect to frame F.

The 4-velocity $U = \gamma_u(c, u_x, u_y)$ has to be Lorentz transformed to F, resulting in a 4-velocity $W = \gamma_w(c, w_x, w_y)$:

$$\begin{pmatrix} c\gamma_w \\ \gamma_w w_x \\ \gamma_w w_y \end{pmatrix} = \begin{pmatrix} \gamma & \gamma v/c & 0 \\ \gamma v/c & \gamma & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} c\gamma_u \\ \gamma_u u_x \\ \gamma_u u_y \end{pmatrix}$$

$$\gamma_w = \gamma \gamma_u \left(1 + \frac{v u_x}{c^2} \right)$$

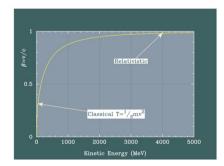
$$\gamma_w w_x = \gamma \gamma_u \left(v + u_x \right)$$

$$\gamma_w w_y = \gamma_u u_y$$

$$w_x = \frac{v + u_x}{\left(1 + \frac{vu_x}{c^2}\right)}$$
 $w_y = \frac{u_y}{\gamma\left(1 + \frac{vu_x}{c^2}\right)}$ Science & Technolog

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Velocity and Energy



$$T = m_0(\gamma - 1)c^2$$

$$\gamma = 1 + \frac{T}{m_0 c^2}$$

$$\beta = \sqrt{1 - \frac{1}{\gamma^2}}$$

$$p = m_0 \beta \gamma c$$

For
$$v \ll c$$
, $\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}} \approx 1 + \frac{1}{2}\frac{v^2}{c^2} + \frac{3}{8}\frac{v^4}{c^4} + \dots$
so $T = m_0 c^2 (\gamma - 1) \approx \frac{1}{2}m_0 v^2$

Basic Quantities usedin Accelerator Calculations

Relative velocity
$$\beta = \frac{1}{2}$$

Velocity
$$v = \beta c$$

Momentum
$$p = mv = m_0 \gamma v = m_0 \gamma \beta c$$

Kinetic energy
$$T = mc^2 - m_0c^2 = (\gamma - 1)m_0c^2 = (\gamma - 1)E_0$$

$$\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}} = \left(1 - \beta^2\right)^{-\frac{1}{2}}$$

$$\implies (\beta\gamma)^2 = \frac{\gamma^2 v^2}{c^2} = \gamma^2 - 1 \implies \beta^2 = \frac{v^2}{c^2} = 1 - \frac{1}{\gamma^2}$$



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Energy-Momentum

$$P \cdot P = m_0^2 V \cdot V = m_0^2 c^2 \quad \text{and} \quad P = \left(\frac{E}{c}, \vec{p}\right)$$

$$\downarrow \qquad \qquad \downarrow$$

$$\frac{E^2}{c^2} - \vec{p}^2 \qquad \qquad = m_0^2 c^2 = \frac{1}{c^2} E_0^2 \quad \text{where } E_0 \text{ is rest energy}$$

$$\implies p^2 c^2 = E^2 - E_0^2$$

$$= (E - E_0)(E + E_0)$$

$$= T(T + 2E_0)$$

Example: ISIS at RAL accelerates protons ($E_0 = 938 \,\mathrm{MeV}$) to $800 \,\mathrm{MeV}$

$$\implies pc = \sqrt{800 \times (800 + 2 \times 938)} \text{ MeV}$$

$$=1.463\,\mathrm{GeV}$$

$$\beta \gamma = \frac{m_0 \beta \gamma c^2}{m_0 c^2} = \frac{pc}{E_0} = 1.56$$

$$\gamma^2 = (\beta \gamma)^2 + 1 \Longrightarrow \gamma = 1.85$$

$$\beta = \frac{\beta \gamma}{\gamma} = 0.84$$



Relationships between small variations in parameters ΔE , ΔT , Δp , $\Delta \beta$, $\Delta \gamma$

$$(\beta\gamma)^2 = \gamma^2 - 1$$

$$\Rightarrow \beta\gamma\Delta(\beta\gamma) = \gamma\Delta\gamma$$

$$\Rightarrow \beta\Delta(\beta\gamma) = \Delta\gamma \qquad (1)$$

$$(\beta\gamma)^{2} = \gamma^{2} - 1$$

$$\Rightarrow \beta\gamma\Delta(\beta\gamma) = \gamma\Delta\gamma$$

$$\Rightarrow \beta\Delta(\beta\gamma) = \Delta\gamma$$

$$(1)$$

$$\frac{\Delta p}{p} = \frac{\Delta(m_{0}\beta\gamma c)}{m_{0}\beta\gamma c} = \frac{\Delta(\beta\gamma)}{\beta\gamma}$$

$$= \frac{1}{\beta^{2}} \frac{\Delta\gamma}{\gamma} = \frac{1}{\beta^{2}} \frac{\Delta E}{E}$$

$$\Rightarrow \frac{1}{\gamma^{3}} \Delta\gamma = \beta\Delta\beta$$

$$(2)$$

$$\frac{\Delta p}{p} = \frac{\Delta(m_{0}\beta\gamma c)}{m_{0}\beta\gamma c} = \frac{\Delta(\beta\gamma)}{\beta\gamma}$$

$$= \frac{1}{\beta^{2}} \frac{\Delta\gamma}{\gamma} = \frac{1}{\beta^{2}} \frac{\Delta E}{E}$$

$$= \gamma^{2} \frac{\Delta\beta}{\beta}$$

$$= \frac{\gamma}{\gamma+1} \frac{\Delta T}{T} \quad \text{(exercise)}$$

Note: valid to first order only



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4-Momentum Conservation

4-momentum

• Equivalent expression for
$$P=m_0\gamma(c,\vec{v})=(mc,\vec{p})=\left(\frac{E}{c},\vec{p}\right)$$
 4-momentum

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• Invariant
$$m_0^2 c^2 = P \cdot P = \frac{E^2}{c^2} - \vec{p}^2 \implies \boxed{\frac{E^2}{c^2} = \vec{p}^2 + m_0^2 c^2}$$

· Classical momentum conservation laws → conservation of 4momentum. Total 3momentum and total energy are conserved.

$$\sum_{\text{particles, i}} P_i = \text{constant}$$

$$\Rightarrow \sum_{\text{particles, i}} E_i \text{ and } \sum_{\text{particles, i}} \vec{p}_i \text{ constant}$$



	$\frac{\Delta \beta}{\beta}$	$\frac{\Delta p}{p}$	$\frac{\Delta T}{T}$	$\frac{\Delta E}{E} = \frac{\Delta \gamma}{\gamma}$
$\frac{\Delta \beta}{\beta} =$	$\frac{\Delta \beta}{\beta}$	$\frac{\frac{1}{\gamma^2} \frac{\Delta p}{p}}{\frac{\Delta p}{p} - \frac{\Delta \gamma}{\gamma}}$	$\frac{1}{\gamma(\gamma+1)} \frac{\Delta T}{T}$	$\frac{1}{\beta^2 \gamma^2} \frac{\Delta \gamma}{\gamma}$ $\frac{1}{\gamma^2 - 1} \frac{\Delta \gamma}{\gamma}$
$\frac{\Delta p}{p} = $	$\gamma^2 \frac{\Delta \beta}{\beta}$	$\frac{\Delta p}{p}$	$\frac{\gamma}{\gamma+1} \frac{\Delta T}{T}$	$\frac{1}{\beta^2} \frac{\Delta \gamma}{\gamma}$
$\frac{\Delta T}{T} = $	$\gamma(\gamma+1)\frac{\Delta\beta}{\beta}$	$\left(1+\frac{1}{\gamma}\right)\frac{\Delta p}{p}$	$\frac{\Delta T}{T}$	$\frac{\gamma}{\gamma - 1} \frac{\Delta \gamma}{\gamma}$
$\begin{array}{c c} \frac{\Delta E}{E} & = \\ \frac{\Delta \gamma}{\gamma} & = \\ \end{array}$	$(\beta\gamma)^2 \frac{\Delta\beta}{\beta}$	$\beta^2 \frac{\Delta p}{p}$	$\begin{pmatrix} 1 & 1 \end{pmatrix} \Delta T$	$\Delta\gamma$
$\frac{\Delta \gamma}{\gamma} = $	$(\gamma^2 - 1) \frac{\Delta \beta}{\beta}$	$\frac{\Delta p}{p} - \frac{\Delta \beta}{\beta}$	$\left(1 - \frac{1}{\gamma}\right) \frac{\Delta T}{T}$	$\frac{\Delta \gamma}{\gamma}$

Table 1: Incremental relationships between energy, velocity and momentum.



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A body of mass *M* disintegrates while at rest into two parts of rest masses M_1 and M_2 .

Show that the energies of the parts are given by

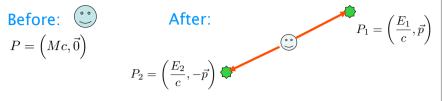
$$E_1 = c^2 \frac{M^2 + M_1^2 - M_2^2}{2M}, \quad E_2 = c^2 \frac{M^2 - M_1^2 + M_2^2}{2M}$$



Solution



$$P = \left(Mc, \vec{0}\right)$$



$$P_1 = \left(\frac{E_1}{c}, \vec{p}\right)$$

Conservation of 4-momentum:

$$P = P_1 + P_2 \implies P - P_1 = P_2$$

$$\Rightarrow (P - P_1) \cdot (P - P_1) = P_2 \cdot P_2$$

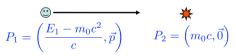
$$\Rightarrow P \cdot P - 2P \cdot P_1 + P_1 \cdot P_1 = P_2 \cdot P_2$$

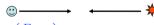
$$\Rightarrow M^2 c^2 - 2ME_1 + M_1^2 c^2 = M_2^2 c^2$$

$$\Rightarrow E_1 = \frac{M^2 + M_1^2 - M_2^2}{2M} c^2$$



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$$P_1 = \left(\frac{E_2}{2c}, \vec{p'}\right)$$

$$P_2 = \left(\frac{E_2}{2c}, -\bar{p}\right)$$

Total energy E₁

(Fixed target experiment)

Total energy E₂

(Colliding beams experiment)

 $P_2 \cdot (P_1 + P_2)$ Invariant:

$$m_0 c \times \frac{E_1}{c} - 0 \times p = \frac{E_2}{2c} \times \frac{E_2}{c} + p' \times 0$$

$$\Rightarrow 2m_0 c^2 E_1 = E_2^2$$



Example of use of invariants

- Two particles have equal rest mass m_o.
 - Frame 1: one particle at rest, total energy is E₁.
 - Frame 2: centre of mass frame where velocities are equal and opposite, total energy is E₂.

Problem: Relate E₁ to E₂



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Collider Problem

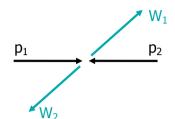
In an accelerator, a proton p₁ with rest mass m₀ collides with an anti-proton p₂ (with the same rest mass), producing two particles W₁ and W₂ with equal rest mass M₀=100m₀

- Experiment 1: p₁ and p₂ have equal and opposite velocities in the lab frame. Find the minimum energy of p₂ in order for W₁ and W₂ to be produced.
- Experiment 2: in the rest frame of p₁, find the minimum energy E' of p₂ in order for W₁ and W₂ to be produced.



$$\frac{E^2}{c^2} = \vec{p}^2 + m_0^2 c^2$$

Particles with same rest-mass and same momentum have same energies.



Total 3-momentum is zero before collision, so is zero afterwards

4-momenta before collision:

$$P_1 = \left(\frac{E}{c}, \vec{p}\right) \quad P_2 = \left(\frac{E}{c}, -\vec{p}\right)$$

4-momenta after collision:

$$P_1 = \left(\frac{E'}{c}, \vec{q}\right) \quad P_2 = \left(\frac{E'}{c}, -\vec{q}\right)$$







Photons and Wave 4-Vectors

- $\sin(\omega t \vec{k} \cdot \vec{x})$ • Monochromatic plane wave:
- \vec{k} is the wave vector, $|\vec{k}| = \frac{2\pi}{\lambda}$; ω is the angular frequency, $\omega = 2\pi\nu$
- The phase $\frac{1}{2\pi}(\omega t \vec{k} \cdot \vec{x})$ is the number of wave crests passing an observer

$$\omega t - \vec{k} \cdot \vec{x} = (ct, \vec{x}) \cdot \left(\frac{\omega}{c}, \vec{k}\right)$$
Position 4-vector Wave 4-vector

• 4-momentum $P = \left(\frac{E}{c}, p\right) = \hbar \left(\frac{\omega}{c}, k\right) = \hbar K$

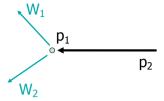


Exmarimant 2

Before collision

$$P_1 = (m_0 c, \vec{0}), \quad P_2 = \left(\frac{E'}{c}, \vec{p}\right)$$

Total energy is $E_1 = E' + m_0 c^2$



Use previous result $2m_0e^2E_1=E_2^2$ to relate E_1 to total energy E_2 in the centre of mass frame

$$2m_0c^2E_1 = E_2^2$$

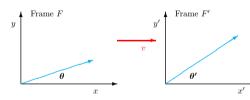
$$\implies 2m_0c^2(E' + m_0c^2) = (2E)^2 > (200m_0c^2)^2$$

$$\implies E' > (2 \times 10^4 - 1)m_0c^2 \approx 20,000 \, m_0c^2$$

Monday, 28 May 2012

Relativistic Doppler Shift

For light rays $\omega = c|\vec{k}|$ so $K = \left(\frac{\omega}{c}, \vec{k}\right)$ is a null vector and can be written $K = \frac{\omega}{c}(1, \vec{n})$ where $|\vec{n}| = 1$.



In F, $K = \frac{\omega}{c} (1, \cos \theta, \sin \theta)$

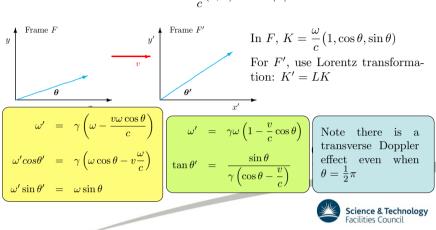
 $c^{(1), 3330, 33110)}$ For F', use Lorentz transformation: K' = LK

 $\begin{pmatrix} \omega'/c \\ (\omega'/c)\cos\theta' \\ (\omega'/c)\sin\theta' \end{pmatrix} = \begin{bmatrix} \gamma & -\gamma v/c & 0 \\ -\gamma v/c & \gamma & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{pmatrix} \omega/c \\ (\omega/c)\cos\theta \\ (\omega/c)\sin\theta \end{pmatrix}$



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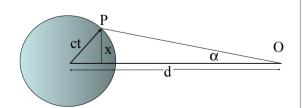
Motion faster than light

 Two rods sliding over each other. Speed of intersection point is v/ sinα, which can be made greater than c.



 Explosion of planetary nebula. Observer sees bright spot spreading out. Light from P arrives t=dα²/2c later.

$$\frac{d}{dx} = \frac{d\alpha^2}{2c} \approx \frac{x}{c} \frac{\alpha}{2} << \frac{x}{c}$$





Monday, 28 May 2012

Radiation from an accelerating charged particle

- Rate of radiation, R, known to be invariant and proportional to $|\vec{a}|^2$ in instantaneous rest frame.
- But in instantaneous rest-frame $A \cdot A = -|\vec{a}|^2$

• Deduce
$$R \propto A \cdot A = -\gamma^6 \left(\left(\frac{\vec{v} \cdot \vec{a}}{c} \right)^2 + \frac{1}{\gamma^2} \vec{a}^2 \right)$$

• Rearranged:

$$R = \frac{e^2}{6\pi\epsilon_0 c^3} \gamma^6 \left[|\vec{a}|^2 - \frac{(\vec{a} \times \vec{v})^2}{c^2} \right]$$

Relativistic Larmor Formula

If $\vec{a} \parallel \vec{v}$, $R \propto \gamma^6$, but if $\vec{a} \perp \vec{v}$, $R \propto \gamma^4$

