# Guided tour through a reco distribution ready for fitting



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# Reference and history unfolded q3 < 1.2



# q3 bin chosen to match Naseem's region of interest



Next higher q3 bin chosen to match Naseem's check NDLAr has a blind spot NDGAr has good acceptance

But anyway, its a core panel of the inclusive distribution that we will fit.

not unfolding, so binned to give 1k to 10k events/bin (3% to 1% stat fluctuation)

### Calorimetric reconstructed energy with leakage



Leakage into HCAL is not negligible for q3 > 1.2, but ignored here.



Intuitions about variables pZ correlates with Enu

pT correlates with q3, sorta Q2

Ehad, Eavail correlates q0, W

Around Enu ~ 1 GeV correlation is less strong and also significant energy dependence

Q2 = 0 is the diagonal here Q2 =  $q3^2 - q0^2$ is far right in energy transfe<sup>6</sup>

#### Kinematics orientation part one





# Step back to GENIE truth

the last Ascencio slice = blue

Distribution on previous slide is the next slice higher in q3

Previous slide horizontal axis Is an estimator for this

Q2 = 0 is the diagonal here  $Q2 = q3^2 - q0^2$ is far right in energy transfer

Notice lower corner is empty<sup>8</sup>

#### Kinematics orientation slide repeated



# Map to proposed pT, pz, Eavail binning



So equivalent to one panel of our proposed 3D binning! Yay!

### Wait, does it look like this with the DUNE flux?



### But 2<sup>nd</sup> oscillation minimum is 2.5ish GeV?



#### What about 2<sup>nd</sup> oscillation maximum ~ 1 GeV?



Events are removed from right to left as you go down in Enu, p2

# GENIE doesn't produce events below QE peak



# Can we guess how DUNE resolution different?



Tracking protons and pions in NDGAr is probably higher resolution

# Why is 2p2h so small?



#### Whats sis1 and sis4 an acronym for?



#### Whats unsafe about Q2 < 1 GeV2 ?



If purple describes data, its probably a quark-hadron duality thing?

#### sis4 is easier to see at higher q3



I went up one more step SIS Q2 < 1 is half the rate its W ~ 1.8 GeV and is near Q2 ~ 0 (to the right!)

dark blue can be seen rising at higher Q2 (to the left!)

We are designing uncertainty priors for these based on the AMU model

#### This is what a fitter gets. What will it do? Don't know.



Parameters change rate and (multi-dimension) shape <sup>20</sup>

### GENIE error bands c. 2020 in the reco distribution



### The inputs have improved since Ascencio



Improved uncertainty priors most around 10% but are more numerous

Net error band priors will still be about 20% but better at describing model shifts and smears rather than normalizations

# So what would a new hand tuning give?



# Conclusions

Have shown a little-before seen slice of MINERvA MC the next q3 bins up from the MINERvA Ascencio paper especially one that touches the SIS region

demonstration is equivalent to proposed 3D binning many features of interest to models and the fitter stand out

not ready to show data yet (but see q3<1.2 in Ascencio paper) Undergraduate projects looking at it and up to q3 < 3.0 GeV

will be in MINERvA's upcoming data preservation product



