

# An Invitation to the New Variables with Possible Applications

Norbert Bodendorfer and Andreas Thurn (work by NB, T. Thiemann, AT [arXiv:1106.1103])

FAU Erlangen-Nürnberg

QG Colloquium 6, 5 October 2011

## Plan of the talk

- 1 Why Higher Dimensional Loop Quantum (Super-)Gravity?
- 2 Review: Hamiltonian Formulations of General Relativity
  - ADM Formulation
  - Extended ADM I
  - Ashtekar-Barbero Formulation
  - Extended ADM II
- The New Variables
  - Hamiltonian Viewpoint
  - Comparison with Ashtekar-Barbero Formulation
  - Lagrangian Viewpoint
  - Quantisation, Generalisations
- 4 Possible Applications of the New Variables
  - Solutions to the Simplicity Constraint
  - Canonical = Covariant Formulation?
  - Supersymmetry Constraint
  - Black Hole Entropy
  - Cosmology
  - AdS / CET Corresponder
  - AdS / CFT Correspondence
  - Conclusion

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## Why Higher Dimensional Loop Quantum (Super-)Gravity? Quantum Gravity:

- Perturbative: Superstring theory / M-theory (ST / MT), require
  - Additional particles
  - Supersymmetry
  - Higher dimensions
- Non-perturbative: Loop Quantum Gravity
  - Various matter couplings & SUSY possible
  - ▶ 3+1 dimensions (Ashtekar Barbero variables) [however, Melosch, Nicolai '97; Nieto '04, '05]
  - What if LHC finds evidence for higher dimensions?
- ightarrow Make contact between them? [Thiemann '04; Fairbairn, Noui, Sardelli '09, '10]
  - Compare results in 3+1 dimensions:
     Landscape problem: Dimensional reduction of ST / MT highly ambiguous
  - Compare results in higher dimensions: Starting points:
    - Higher dimensional Supergravities
      - ★ are considered as the low-energy limits of ST / MT
      - $\star$  have action of the type  $S_{GR}$  + more
    - Symmetry reduced models (higher dim. & SUSY black holes or cosmology)
- → Extend loop quantisation programme to higher dimensions and Supergravities

  [Jacobson '88: Fülöp '93; Armand-Ugon, Gambini, Obrégon, Pullin '95; Ling, Smolin 'ஐ9, Sayaggychi 'இ1, Smolin', Obrégon, Pullin '95; Ling, Smolin', 'ஐ9, Sayaggychi 'இ1, Smolin', 'இ9, Sayaggychi 'இ1, Smolin', 'B9, Sayaggychi '@1, Smolin', 'B9, Sayaggychi 'B1, Smolin', 'B9, Sayaggychi '@1, Smolin', 'B9, Sayaggychi 'B1, Smolin', 'B1, Smoli

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## ADM Formulation [Arnowitt, Deser, Misner '62]

#### D+1 split

• Foliation of  $\mathcal{M}$ :

$$\mathcal{M}$$
 top.  $\mathbb{R} \times \sigma$ ,  $\Sigma_t = X_t(\sigma)$ ,  $X_t : \sigma \to \mathcal{M}$ 

• Important fields on  $\sigma$ : Lapse, Shift: N, N<sup>a</sup>

Spatial metric 
$$q_{ab} = (X^*g)_{ab}$$
,

Extrinsic curvature 
$$K_{ab} = (X^* \mathcal{L}_n q)_{ab}$$
  
=  $\frac{1}{N} (\dot{q}_{ab} - (\mathcal{L}_{\vec{N}} q)_{ab})$ 

$$\Sigma_{t_{d}}$$

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$$\Rightarrow S_{EH} = \int dt \int_{\sigma} d^{D}x \ N\sqrt{\det q} \left( R^{(D)} \pm \left[ K_{ab} K^{ab} - \left( K_{a}^{a} \right)^{2} \right] \right) \quad [a, b = 1, ..., D]$$

- Canonical variables:  $q_{ab}$ ,  $P^{ab}$  ( $\sim$  extrinsing curvature  $K_{ab}$ )
- Poisson brackets:  $\{q_{ab}(x), P^{cd}(y)\}_{ADM} = \delta^c_{(a}\delta^d_{b)}\delta^{(D)}(x,y)$
- 1<sup>st</sup> class constraints:

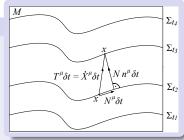
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Totally constrained Hamiltonian:  $H = \int_{\sigma} d^{D}x(N\mathcal{H} + N^{a}\mathcal{H}_{a})$ Spatial diffeomorphism constraint  $\mathcal{H}_a(q, P)$ 

Hamiltonian constraint  $\mathcal{H}(q, P) = \pm \sqrt{\det q} \ R^{(D)} + \frac{1}{\sqrt{\det q}} [P_{ab}P^{ab} - \frac{1}{D-1}(P_a^a)^2]$ 

### Extended ADM I

#### Extension of ADM phase space I

• Introduce SO(D)-valued vielbein:

- Poisson bracket relations:  $\{E^{ai}, K_{bj}\} = \delta^a_b \delta^i_j$
- Increased number of degrees of freedoms ⇒ new constraint needed:

$$K_{[ab]} = 0 \quad \Leftrightarrow \quad K_{[a}{}^{i}e_{b]i} = 0 \quad \Leftrightarrow \quad G_{ij} := K_{a[i}E^{a}{}_{j]} = 0 \tag{2}$$

#### Valid extension?

ADM Possion bracket relations reproduced on extended phase space

$$\{q_{ab}(E), P^{cd}(E, K)\}|_{G=0} = \{q_{ab}, P^{cd}\}_{ADM} = \delta^{c}_{(a}\delta^{d}_{b)}\delta^{(D)}(x, y)$$
(3)

- New constraints close amongst themselves:  $\{G,G\} \sim G$
- $q_{ab}(E), P_{cd}(E, K)$  (and in particular  $\mathcal{H}, \mathcal{H}_a$ ) are Dirac observables w.r.t. new constraint  $G_{ii}$
- $\Rightarrow~\mathcal{H},\mathcal{H}_{\mathsf{a}}$  and  $\mathit{G}_{ij}$  constitute  $1^{\mathit{st}}$  class constraint algebra by construction

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$$q_{ab} = e_a{}^i e_b{}^j \delta_{ij}$$
  $K_{ab} = K_{ai} e_b^i$   $E^{ai} = \sqrt{\det q} e^{ai}$   $i, j, ... \in \{1, ..., D\}$  (1)

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## Ashtekar-Barbero Formulation

#### Canonical transformation to Ashtekar-Barbero variables [Sen; Ashtekar; Immirzi; Barbero]

- Introduce spin connection  $\Gamma^{SPIN}_{aij}[e]$  s.t.  $\partial_a e_{bi} \Gamma^c_{ab} e_{ci} + \Gamma^{SPIN}_{aij}[e]$   $e_b{}^j = 0$
- Crucial: Defining and adjoint representation of SU(2) equivalent!
- Only in D = 3: Canonical transformation

$$\{E^{ai}, K_{bj}\} \longrightarrow \{\frac{1}{\gamma}E^{ai}, A_{bj} := 1/2 \epsilon_j^{kl} \Gamma_{bkl}^{SPIN}[e] + \gamma K_{bj}\} \qquad \gamma \in \mathbb{R}/\{0\}: \text{ Immirzi Parameter}$$
 (4)

- $\Rightarrow$  Simple Poisson algebra  $\{A,E\}\sim 1$  and  $1^{st}$  class constraint algebra
- Canonicity of the above transformation non-trivial
- New constraint  $G_{ij} = K_{a[i}E^{a}{}_{j]} \Rightarrow SU(2)$  Gauß law constraint:

$$G_{ij} = \gamma K_{a[i} \frac{1}{\gamma} E^{a}{}_{j]} + \frac{1}{2\gamma} \epsilon_{ij}{}^{k} (\partial_{a} E^{a}{}_{k} + \Gamma^{SPIN}_{akl}[e] E^{al})$$

$$= \frac{1}{2\gamma} \epsilon_{ij}{}^{k} (\partial_{a} E^{a}{}_{k} + \epsilon_{k}{}^{lm} A_{al} E^{a}{}_{m})$$
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### Higher dimensions?

No obvious way of combining  $K_{ai}$  and  $\Gamma_{aij}^{SPIN}[e]$  to a connection conjugate to  $E^{bj}$  in a mathematically sensible way!

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## Extended ADM II

#### Extension of ADM phase space II

• Introduce SO(D+1) or SO(1,D) "hybrid" vielbein:

$$q_{ab} = e_a{}^I e_b{}^J \eta_{IJ}$$
  $K_{ab} = K_{aJ} e_b^J$   $E^{aJ} = \sqrt{\det q} e^{aJ}$   $I, J, ... \in \{0, 1, ..., D\}$  (6)

- Motivation: 2<sup>nd</sup> order Palatini formulation of General Relativity
- Poisson bracket relations:  $\{E^{al}, K_{bJ}\} = \delta^a_b \delta^l_J$
- New constraints:  $K_{[ab]} = 0 \Leftrightarrow K_{[a}{}^{I}e_{b]I} = 0$  insufficient! Use  $\Leftrightarrow G^{IJ} := K_{a}{}^{[I}E^{aJ]}$  (7)
- Proof of validity of extension II analogous to extension I case

#### Connection formulation?

- "Hybrid" spin connection [Peldan '94]  $\Gamma_{aIJ}^{HYB}[e]$  s.t.  $\partial_a e_{bI} \Gamma_{ab}^c e_{cI} + \Gamma_{aIJ}^{HYB}[e]$   $e_b{}^J = 0$
- BUT: No obvious way of combining  $K_{aJ}$  and  $\Gamma_{aIJ}^{HYB}[e]$  to a connection conjugate to  $E^{bJ}$  in a mathematically sensible way (if  $D \neq 2$ )!



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## The New Variables - Hamiltonian Viewpoint

#### Extension of ADM phase space III

• Introduce "generalised" vielbein, transforming in the adjoint representation of SO(D+1) or SO(1,D):

$$q_{ab} = e_{alJ}e_b^{\ IJ}$$
  $K_{ab} = K_{alJ}e_b^{\ IJ}$   $\pi^{alJ} = \sqrt{\det q} \ e^{alJ}$   $I, J, ... \in \{0, 1, ..., D\}$  (8)

- Motivation: 1<sup>st</sup> order Palatini formulation of General Relativity (cf. next to next slide)
- Poisson bracket relations:  $\{\pi^{alJ}, K_{bKL}\} = \delta^a_b \ \delta^I_{[K} \delta^J_{L]}$
- New constraints: Gauß and simplicity constraint

$$G^{IJ} := K_a^{[I|K} \pi^a_{K}^{J]}$$
 and  $S^{aIJ\ bKL} := \pi^{a[IJ|} \pi^{b|KL]}$  (9)

Proof of validity of extension analogous to extension I and II case

#### Canonical transformation to new connection formulation

•  $\Gamma^{HYB}_{alJ}[\pi]$ : Extension of  $\Gamma^{HYB}_{alJ}[e]$  off the simplicity constraint surface

$$S = 0 \Leftrightarrow \pi^{alJ} = n^{[l} E^{a|J]}$$
 [Freidel, Krasnov, Puzio '99] (10)

• Canonical transformation (non-trivial):

$$\{\pi^{\mathsf{a}IJ}, K_{\mathsf{bKL}}\} \longrightarrow \{\frac{1}{\beta}\pi^{\mathsf{a}IJ}, A_{\mathsf{bKL}} := \Gamma^{\mathsf{HYB}}_{\mathsf{bKL}}[\pi] + \beta K_{\mathsf{bKL}}\} \quad \beta \in \mathbb{R}/\{0\}, \neq \gamma!$$
 (11)

ullet G<sup>IJ</sup> becomes SO(D+1) or SO(1, D) Gauß law constraint:

$$^{IJ} = \partial_a \pi^{aIJ} + A_a^{[I]} \kappa \pi^{aK[J]}$$
 (12)

• Formulation works with SO(D+1) and SO(1,D) independent of spacetime signature!

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•  $G^{IJ}$  becomes SO(D+1) or SO(1,D) Gauß law constraint:

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## Comparison with Ashtekar-Barbero Formulation

#### Ashtekar-Barbero formulation

- Canonical variables  $A_{ai}^{LQG}$ ,  $E^{bk}$  are real
- Simple Poisson algebra  $\{A^{LQG}, E\} \sim 1$
- Compact gauge group SU(2)
- First class constraints  $\mathcal{H}, \mathcal{H}_a$  and  $G^i$
- Physical information:

$$A_{aij}^{LQG} - \Gamma_{aij}^{SPIN}[e] = \gamma \epsilon_{ij}^{k} K_{ak}$$
 (13)

• Relation to other formulations: AB  $\stackrel{G=0}{\longrightarrow}$  ADM

#### New formulation, D = 3

- Canonical variables  $A_{alJ}^{NEW}$ ,  $\pi^{bKL}$  are real
- Simple Poisson algebra  $\{A^{\it NEW},\pi\}\sim 1$
- Compact gauge group SO(4)
- First class constraints  $\mathcal{H}, \mathcal{H}_a, G^{IJ}$  and  $S^{aIJ\ bKL}$
- Physical information:

$$A_{aij}^{NEW} - \Gamma_{aij}^{HYB}[\pi] \approx S - gauge, \qquad A_{a0j}^{NEW} - \Gamma_{a0j}^{HYB}[\pi] \approx \beta K_{aj}$$
 (14)

 $\bullet \ \ \mathsf{NEW} \ \ \stackrel{S=0}{\longrightarrow} \ \ \mathsf{Ex.} \ \ \mathsf{ADM} \ \mathsf{II} \ \ \stackrel{\mathsf{time\ gauge}}{\longrightarrow} \ \ \mathsf{Ex.} \ \ \mathsf{ADM} \ \mathsf{I} \ \ \stackrel{G=0}{\longrightarrow} \ \ \mathsf{ADM}$ 

## The New Variables - Lagrangian Viewpoint

Canonical analysis of the 1st order Palatini action [Peldan '94]

$$S_{P} = \int \left( \pi^{aIJ} \ \dot{A}_{aIJ} - N\mathcal{H} - N^{a}\mathcal{H}_{a} - \Lambda \cdot G - c \cdot S \right) \tag{15}$$

- Gauß and simplicity constraint: Exactly like before
- ullet Dirac constraint analysis: Additional constraint D, second class partner to S
- A<sub>alJ</sub> not self-commuting w.r.t. corresponding Dirac bracket [Alexandrov '00]
- ⇒ Loop quantisation not (directly) applicable! [see, however: Alexandrov & Roche '10;

Geiller, Lachieze-Rey, Noui, Sardelli '11]

#### Gauge Unfixing [Mitra & Rajaraman '89 '90; Henneaux & Teitelboim '92; Anishetty & Vytheeswaran '93]

- Well defined procedure:  $2^{nd}$  class  $\Rightarrow$   $1^{st}$  class constrained system
- ullet Applied to GR: Drop D at the cost of a more complicated  ${\cal H}$
- Resulting theory coincides with result of Hamiltonian derivation iff
  - Internal and external signatures match
  - Free parameter  $\beta = 1$

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### Gauge Unfixing [Mitra & Rajaraman '89 '90; Henneaux & Teitelboim '92; Anishetty & Vytheeswaran '93]

- ullet Well defined procedure:  $2^{nd}$  class  $\Rightarrow$   $1^{st}$  class constrained system
- ullet Applied to GR: Drop D at the cost of a more complicated  ${\cal H}$
- Resulting theory coincides with result of Hamiltonian derivation iff
  - Internal and external signatures match
  - Free parameter  $\beta = 1$

## Quantisation, Generalisations

#### Quantisation [Rovelli, Smolin, Ashtekar, Isham, Lewandoski, Marolf, Mourao, Thiemann...]

- Most results of loop quantisation formulated independently of
  - Dimension of spacetime
  - Choice of compact gauge group
- Sole new ingredient for canonical theory: Implementation of simplicity constraint (but well-known from covariant approach, cf. below)

#### Generalisations

- Extension to diverse matter fields and supergravity:
  - Dirac, Weyl, Majorana fermions
  - Gauge fields with compact gauge groups
  - Scalar fields
  - Rarita-Schwinger fields (gravitinos)
  - Abelian higher *p*-form fields
- Not treatable so far:
  - Non-abelian higher *p*-form fields (higher gauge theory?)
  - Non-compact gauge groups
- ⇒ Includes, inter alia, supergravity theories in 4, 10 and 11 dimensions

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## Plan of the talk

- Why Higher Dimensional Loop Quantum (Super-)Gravity?
- 2 Review: Hamiltonian Formulations of General Relativity
  - ADM Formulation
  - Extended ADM I
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- The New Variables
  - Hamiltonian Viewpoint
  - Comparison with Ashtekar-Barbero Formulation
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  - Quantisation, Generalisations
- 4 Possible Applications of the New Variables
  - Solutions to the Simplicity Constraint
  - Canonical = Covariant Formulation?
  - Supersymmetry Constraint
  - Black Hole Entropy
  - Cosmology
  - AdS / CFT Correspondence
  - Conclusion

## Solutions to the Simplicity Constraint

There exist multiple, plausible suggestions for solving the simplicity constraint, e.g.

- Weak implementation [Engle, Pereirra, Rovelli '07; Livine, Speziale '07]
- Coherent states [Freidel, Krasnov '07]
- Holomorphic simplicity constraints [Dupuis, Freidel, Livine, Speziale '11]
- Maximally commuting subsets [NB, Thiemann, AT '11]
- ...

It is however in general unclear, if they lead to the same dynamics.

## Application of the new variables

- Test different implementations of the simplicity constraint within the new canonical framework for dynamical equivalence
- New requirement: Anomaly-freedom of the constraint algebra including the Hamiltonian constraint, i.e. implement

$$\{S[...], H[N]\} = S[...] \rightarrow \left[\hat{S}, \hat{H}\right] = \hat{S}$$
 (16)

Basic idea: The spinfoam provides a rigging map for the Hamiltonian constraint.

$$\langle \phi \mid \psi \rangle_{\text{phys}} = \sum_{\kappa: \psi \to \phi} Z[\kappa]$$
 (17)

### New question

For which canonical quantisation should we test the above equation?



Are the quantum theories based on the Ashtekar-Barbero and the newly proposed variables equivalent?

Ashtekar-Barbero	New variables
simplicity solved classically	simplicity can be quantised
$\Rightarrow$ Hilbert spaces have to be related	$\Rightarrow$ Hilbert spaces are the same
usual Hamiltonian constraint	Hamiltonian constraint more complicated
⇒ calculations "easier"	⇒ calculations "harder"

## Supersymmetry Constraint

Important open problem:

Understand the solution space of the Hamiltonian constraint, including matter.

[Teitelboim '77]

$$\{S,S\} = H + H_a + S,$$
 S: supersymmetry constraint (18)

Assuming an anomaly-free implementation of the super Dirac algebra:

Solution to the supersymmetry constraint operator

\$\\$\\$\\$\$
Solution to the Hamiltonian constraint operator

ightarrow Supergravity as a simplified version of General Relativity coupled to matter

Important progress with implementing the supersymmetry constraint has been made in the GSU(2) framework. [Armand-Ugon, Gambini, Obrégon, Pullin '95]

## Comparing LQG to other Approaches to Quantum Gravity

#### General considerations

- Supergravity has been extensively studied as a low energy limit of String- / M-theory
- A great deal of "technology" has been developed in order to deal with String- / M-theory and Supergravity

## Comparing LQG to String- / M-theory

- Dimensional reduction to 4 dimensions is not unique
  - ightarrow Work in the natural dimensions of String- / M-theory
- Generic calculations are hard both in LQG and String- / M-theory
  - → Work in symmetry reduced situations

## Black Hole Entropy

## Calculation of black hole entropy

- Thermodynamic analogy [Bekenstein '73]; QFTCS [Hawking '74]
- String theory [Strominger, Vafa; ... '96]
- Loop quantum gravity [Krasnov '96; Rovelli 96'; Ashtekar, Baez, Corichi, Krasnov '97-, ...]

 $\Rightarrow {\sf Calculation\ possible\ in\ different\ theories!}$ 

## Application of the new variables

- Calculate entropy of a supersymmetric extremal black hole in higher dimensions
- Compare to results coming from string theory

## Cosmology

## Cosmology from different points of view

- Wheeler-DeWitt quantum cosmology [Wheeler '64-; DeWitt '67; Misner '69]
- String cosmology [Veneziano; ... '91]
- Loop quantum cosmology [Bojowald '01-, Ashtekar, Kaminski, Lewandowski, Pawlowski, Singh, ... '02-]

 $\Rightarrow {\sf Calculation\ possible\ in\ different\ theories!}$ 

## Application of the new variables

- Investigate SLQC in higher dimensions
- Compare to results coming from string cosmology and possibly from experiments
  - ightarrow hints of higher dimensions and supersymmetry in cosmological observables?

## Conjectured exact equivalence

Type IIB String Theory on  $AdS^5 \times S^5$ 

String coupling  $g_s$ , String tension T



 $\mathcal{N}=$  4 Super Yang-Mills Theory in 4d

YM coupling gym, number of coulors N

- weak string coupling
- strong string tension (only massless states)

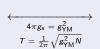


- weak YM-coupling
- strong 't-Hooft coupling (only planar diagrams)

## Well tested low energy equivalence

Type IIB Supergravity in AdS<sup>5</sup>xS<sup>5</sup>

$$g_s \to 0$$
,  $T \to \infty$ 



 $\mathcal{N}=4$  Super Yang-Mills Theory in 4dat strong 't Hooft coupling

$$g_{\mathsf{YM}} \to 0$$
,  $g_{\mathsf{YM}}^2 N \to \infty$ 

## Conjectured exact equivalence

Type IIB String Theory on AdS<sup>5</sup>xS<sup>5</sup>

String coupling  $g_s$ , String tension T

$$4\pi g_{S} = g_{YM}^{2}$$

$$T = \frac{1}{2\pi} \sqrt{g_{YM}^{2} N}$$

 $\mathcal{N}=$  4 Super Yang-Mills Theory in 4d

YM coupling  $g_{YM}$ , number of coulors N

## New non-perturbative limit?

Loop quantized Type IIB Supergravity (in AdS<sup>5</sup>xS<sup>5</sup>?)

$$g_s = ?, \quad T = ?$$



 $\mathcal{N}=$  4 Super Yang-Mills Theory in 4d

$$g_{YM} = ?$$
,  $g_{YM}^2 N = ?$ 

## Well tested low energy equivalence

Type IIB Supergravity in AdS<sup>5</sup>xS<sup>5</sup>

$$g_s \to 0$$
,  $T \to \infty$ 



 $\mathcal{N}=$  4 Super Yang-Mills Theory in 4d at strong 't Hooft coupling

$$g_{YM} \rightarrow 0$$
,  $g_{YM}^2 N \rightarrow \infty$ 

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### Conclusion

- ullet D + 1 dim. GR formulated on an SO(D + 1) Yang-Mills phase space
- ullet LQG methods apply o rigorous quantisation exists
- Extensions to interesting Supergravities exist
- Possible applications include
  - Better understanding the simplicity constraint
  - Supergravity as "simplified" matter coupled GR
  - ► Higher dimensional (supersymmetric) black hole entropy
  - Higher dimensional (supersymmetric) quantum cosmology
  - New tests / applications of the AdS/CFT correspondence?
    - Thank you for your attention!