

# Heavy Majorana neutrino effects on MSSM- $M_h$

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main results in JHEP05(2011)063



# Motivation

- Why **Majorana** neutrinos?
  - Neutrinos have mass and oscillate in flavor (LFV in  $\nu$  sector)
  - Simplest way to explain  $\nu$  masses: introduction of  $\nu_R$ 
    - Dirac masses:  $m_D \bar{\nu}_L \nu_R$  (non-violating lepton number)
    - Majorana masses:  $m_M \bar{\nu}_R^c \nu_R$  (violating lepton number)
  - Majorana masses  $\Rightarrow$  viable BAU via Leptogenesis
- Why **heavy** Majorana neutrinos?
  - **Seesaw** gives small  $m_\nu \sim m_D^2/m_M$  if  $m_D \sim m_{EW} \ll m_M$
  - $m_\nu^{\text{exp}} < 2$  eV requires very heavy  $m_M \sim 10^{13} - 10^{15}$  GeV
- If Dirac:  $Y_\nu \sim \mathcal{O}(10^{-12})$ , If heavy Majorana:  **$Y_\nu \sim \mathcal{O}(1)$  large!**  
 $\Rightarrow$  very important pheno implications: **large LFV** in  $l$  sector
- If large  $m_M$ , SUSY needed to avoid (extra) hierarchy problem  
 $\Rightarrow$  large effects on  $M_h$  from Maj- $\nu$ -RADCOR? (Our work)

## Our Model: MSSM-Seesaw. Part I: MSSM

### Enlarged Higgs sector: Two Higgs doublets

$$H_1 = \begin{pmatrix} H_1^1 \\ H_1^2 \end{pmatrix} = \begin{pmatrix} v_1 + (\phi_1 + i\chi_1)/\sqrt{2} \\ \phi_1^- \end{pmatrix}$$

$$H_2 = \begin{pmatrix} H_2^1 \\ H_2^2 \end{pmatrix} = \begin{pmatrix} \phi_2^+ \\ v_2 + (\phi_2 + i\chi_2)/\sqrt{2} \end{pmatrix}$$

$$V = m_1^2 H_1 \bar{H}_1 + m_2^2 H_2 \bar{H}_2 - m_{12}^2 (\epsilon_{ab} H_1^a H_2^b + \text{h.c.}) \\ + \underbrace{\frac{g'^2 + g^2}{8}}_{\text{gauge couplings, in contrast to SM}} (H_1 \bar{H}_1 - H_2 \bar{H}_2)^2 + \underbrace{\frac{g^2}{2}}_{\text{gauge couplings, in contrast to SM}} |H_1 \bar{H}_2|^2$$

physical states:  $h^0, H^0, A^0, H^\pm$       Goldstone bosons:  $G^0, G^\pm$

Input parameters: (to be determined experimentally)

$$\tan \beta = \frac{v_2}{v_1}, \quad M_A^2 = -m_{12}^2 (\tan \beta + \cot \beta)$$

RADCOR in the MSSM Higgs sector do contain info on NEW phys.

To lowest order, only 2 input parameters:  $\tan\beta$  and  $M_A$

To higher order, extra input parameters:  $m_{\text{SUSY}}$ , etc

$\Rightarrow M_h, M_H$ , mixing angle  $\alpha$ ,  $M_{H^\pm}$ : derived from input param.

With higher-order corrections:  $M_h^2 = m_h^2 + \Delta m_h^2$

$$m_h^2 = \frac{1}{2} \left[ M_A^2 + M_Z^2 - \sqrt{(M_A^2 + M_Z^2)^2 - 4M_Z^2 M_A^2 \cos^2 2\beta} \right]$$

$\Delta m_h^2$  test of the model and sensitivity to new mass scales!

Leading RADCOR is from 1-loop top/stop due to large  $Y_t \sim \mathcal{O}(1)$ :

$\Delta m_h^2 \sim G_F m_t^4 \log \frac{m_{\text{SUSY}}^2}{m_t^2}$ . With higher order corrs.:  $M_h < 135$  GeV

**Our goal:** Can one reach sensitivity to  $m_M$  in  $\Delta m_h$ ? Necessary:

- **discover** the Higgs(es) at the LHC (or at the ILC)
- **measure its mass/characteristics** at the LHC (or at the ILC)
- compare with **theory** :  $\Delta m_h(m_M)$  should be larger than precision expected precision on SM-like Higgs mass: LHC  $\sim 0.2$  GeV, ILC  $\sim 0.05$  GeV

## Our Model: MSSM-Seesaw. Part II: Seesaw

Present work: 1 generation  $\nu$  and  $\tilde{\nu}$ . Future work: 3 generations.

### The neutrino sector:

In the  $(\nu_L, \nu_R)$  basis, the  $2 \times 2$  neutrino mass matrix is given in terms of the Dirac mass  $m_D \equiv Y_\nu v_2$  and the Majorana mass  $m_M$  by:

$$M^\nu = \begin{pmatrix} 0 & m_D \\ m_D & m_M \end{pmatrix}$$

$\Rightarrow$  2 mass eigenstates

$\nu, N$  (Majorana fermions)

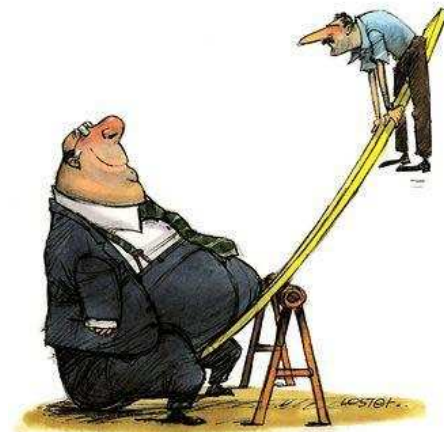
$$m_{\nu, N} = \frac{1}{2} \left( m_M \mp \sqrt{m_M^2 + 4m_D^2} \right)$$

Seesaw limit:  $\xi \equiv m_D/m_M \ll 1$ :

$$m_\nu = -m_D \xi + \mathcal{O}(m_D \xi^3) \simeq -\frac{m_D^2}{m_M} \quad \text{light, predom. } \nu_L$$

$$m_N = m_M + \mathcal{O}(m_D \xi) \simeq m_M \quad \text{heavy, predom. } \nu_R$$

Higgs-neutrino interactions driven by large  $Y_\nu = \sqrt{|m_\nu| m_N}/v_2 \sim \mathcal{O}(1)$



## The sneutrino sector:

$$V_{\text{soft}}^{\tilde{\nu}} = m_{\tilde{L}}^2 \tilde{\nu}_L^* \tilde{\nu}_L + m_{\tilde{R}}^2 \tilde{\nu}_R^* \tilde{\nu}_R + (Y_\nu A_\nu H_2^2 \tilde{\nu}_L \tilde{\nu}_R^* + m_M B_\nu \tilde{\nu}_R \tilde{\nu}_R + \text{h.c.})$$

Two  $2 \times 2$  mass matrices to describe the  $\mathcal{CP}$ -even (+) and  $\mathcal{CP}$ -odd (-) parts of the sneutrino sector:

$$\tilde{M}_{\pm}^2 = \begin{pmatrix} m_{\tilde{L}}^2 + m_D^2 + \frac{1}{2} M_Z^2 \cos 2\beta & m_D (A_\nu - \mu \cot \beta \pm m_M) \\ m_D (A_\nu - \mu \cot \beta \pm m_M) & m_{\tilde{R}}^2 + m_D^2 + m_M^2 \pm 2B_\nu m_M \end{pmatrix}$$

Diagonalization yields four mass eigenstates

$$\tilde{n}_1, \tilde{n}_2, \tilde{n}_3, \tilde{n}_4 \quad (\tilde{\nu}_+, \tilde{N}_+, \tilde{\nu}_-, \tilde{N}_-)$$

Seesaw limit:  $\xi \equiv m_D/m_M \ll 1$ :

$$m_{\tilde{\nu}_+, \tilde{\nu}_-}^2 = m_{\tilde{L}}^2 + \frac{1}{2} M_Z^2 \cos 2\beta \mp 2m_D (A_\nu - \mu \cot \beta - B_\nu) \xi \quad \text{light, predom. } \tilde{\nu}_L$$

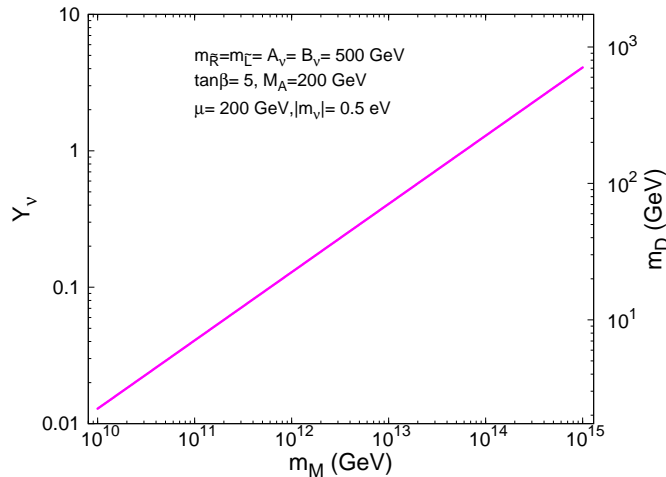
$$m_{\tilde{N}_+, \tilde{N}_-}^2 = m_M^2 \pm 2B_\nu m_M + m_{\tilde{R}}^2 + 2m_D^2 \quad \text{heavy, predom. } \tilde{\nu}_R$$

$m_M$  and  $m_{\tilde{R}}$  are the most relevant parameters (see later)

## Majorana versus Dirac couplings:

→ Same interactions in  $\mathcal{L}_{\nu H}$ , but very different size of couplings:

$$\mathcal{L}_{\nu H} = -\frac{gm_D}{2M_W s_\beta} \left[ (\bar{\nu}_L \nu_R + \bar{\nu}_R \nu_L)(H s_\alpha + h c_\alpha) - i(\bar{\nu}_L \nu_R - \bar{\nu}_R \nu_L) A c_\beta \right]$$



→ large couplings in Majorana case

$$Y_\nu^{\text{Maj}} = \sqrt{|m_\nu| m_M} / v_2 ; m_D^{\text{Maj}} = \sqrt{|m_\nu| m_M}$$

$$Y_\nu^{\text{Maj}} \sim 1 ; m_D^{\text{Maj}} \sim 200 \text{ GeV for:}$$

$$|m_\nu| = 0.5 \text{ eV, } m_M = 10^{14} \text{ GeV, } \tan \beta = 5$$

→ extremely small in Dirac case:

$$Y_\nu^{\text{Dirac}} \sim 10^{-12} ; m_D^{\text{Dirac}} \sim 0.1 \text{ eV}$$

→ New interactions in  $\mathcal{L}_{\tilde{\nu} H}$  with respect to the Dirac case:

$$\mathcal{L}_{\tilde{\nu} H}^{\text{new}} = -\frac{gm_D m_M}{2M_W s_\beta} \left[ (\tilde{\nu}_L \tilde{\nu}_R + \tilde{\nu}_L^* \tilde{\nu}_R^*)(H s_\alpha + h c_\alpha) - i(\tilde{\nu}_L \tilde{\nu}_R - \tilde{\nu}_L^* \tilde{\nu}_R^*) A c_\beta \right]$$

→ In the mass eigen. basis some couplings grow with  $m_M$ :  $\nu N h, \tilde{\nu} \tilde{N} h, \dots$

Full set of FRs in the physical basis in our paper JHEP05(2011)063

## Calculation

$M_h, M_H$ : higher-order corrected  $\mathcal{CP}$ -even Higgs masses in the MSSM

$M_h^{\nu/\tilde{\nu}}, M_H^{\nu/\tilde{\nu}}$ : masses in the MSSM-seesaw model

→ determined as poles of the propagator matrix

Inverse of the propagator matrix:

$$\left(\Delta_{\text{Higgs}}\right)^{-1} = -i \begin{pmatrix} p^2 - m_H^2 + \hat{\Sigma}_{HH}(p^2) & \hat{\Sigma}_{hH}(p^2) \\ \hat{\Sigma}_{hH}(p^2) & p^2 - m_h^2 + \hat{\Sigma}_{hh}(p^2) \end{pmatrix}$$

→ solve the equation:

$$\left[p^2 - m_h^2 + \hat{\Sigma}_{hh}(p^2)\right] \left[p^2 - m_H^2 + \hat{\Sigma}_{HH}(p^2)\right] - \left[\hat{\Sigma}_{hH}(p^2)\right]^2 = 0$$

with renormalized self-energies:

$$\hat{\Sigma}(p^2) = \hat{\Sigma}^{(1)}(p^2) + \hat{\Sigma}^{(2)}(p^2) + \dots$$

$$\Sigma(p^2) = \Sigma^{(1)}(p^2) + \Sigma^{(2)}(p^2) + \dots$$

⇒ calculation of  $\nu/\tilde{\nu}$  contributions to  $\hat{\Sigma}^{(1)}$  (one loop)



## Renormalization and formulas “as usual” :

$$\hat{\Sigma}_{hh}(p^2) = \Sigma_{hh}(p^2) + \delta Z_{hh}(p^2 - m_h^2) - \delta m_h^2$$

$$\hat{\Sigma}_{hH}(p^2) = \Sigma_{hH}(p^2) + \delta Z_{hH}(p^2 - \frac{1}{2}(m_h^2 + m_H^2)) - \delta m_{hH}^2$$

$$\hat{\Sigma}_{HH}(p^2) = \Sigma_{HH}(p^2) + \delta Z_{HH}(p^2 - m_H^2) - \delta m_H^2$$

$$\begin{aligned} \delta m_h^2 &= \delta M_A^2 c_{\beta-\alpha}^2 + \delta M_Z^2 s_{\alpha+\beta}^2 + \delta \tan\beta s_\beta c_\beta (M_A^2 s_{2\alpha-2\beta} + M_Z^2 s_{2\alpha+2\beta}) \\ &\quad + \frac{e}{2M_Z s_W c_W} (\delta T_H c_{\beta-\alpha} s_{\beta-\alpha}^2 - \delta T_h s_{\beta-\alpha} (1 + c_{\beta-\alpha}^2)) \end{aligned}$$

$$\begin{aligned} \delta m_{hH}^2 &= \frac{1}{2} (\delta M_A^2 s_{2\alpha-2\beta} - \delta M_Z^2 s_{2\alpha+2\beta} - \delta \tan\beta s_\beta c_\beta (M_A^2 c_{2\alpha-2\beta} + M_Z^2 c_{2\alpha+2\beta})) \\ &\quad + \frac{e}{2M_Z s_W c_W} (\delta T_H s_{\alpha-\beta}^3 - \delta T_h c_{\alpha-\beta}^3) \end{aligned}$$

$$\begin{aligned} \delta m_H^2 &= \delta M_A^2 s_{\alpha-\beta}^2 + \delta M_Z^2 c_{\alpha+\beta}^2 - \delta \tan\beta s_\beta c_\beta (M_A^2 s_{2\alpha-2\beta} + M_Z^2 s_{2\alpha+2\beta}) \\ &\quad - \frac{e}{2M_Z s_W c_W} (\delta T_H c_{\alpha-\beta} (1 + s_{\alpha-\beta}^2) + \delta T_h s_{\alpha-\beta} c_{\alpha-\beta}^2) \end{aligned}$$

$$\delta M_Z^2 = \text{Re}\Sigma_{ZZ}(M_Z^2), \quad \delta M_W^2 = \text{Re}\Sigma_{WW}(M_W^2), \quad \delta M_A^2 = \text{Re}\Sigma_{AA}(M_A^2)$$

$$\delta T_h = -T_h, \quad \delta T_H = -T_H$$

## Field and $\tan\beta$ renormalization:

“Normal”:  $\overline{\text{DR}}$ :

$$\delta Z_{\mathcal{H}_1}^{\overline{\text{DR}}} = - \left[ \text{Re}\Sigma'_{HH} |_{\alpha=0} \right]^{\text{div}}$$

$$\delta Z_{\mathcal{H}_2}^{\overline{\text{DR}}} = - \left[ \text{Re}\Sigma'_{hh} |_{\alpha=0} \right]^{\text{div}}$$

$$\delta \tan\beta^{\overline{\text{DR}}} = \frac{1}{2} \left( \delta Z_{\mathcal{H}_2}^{\overline{\text{DR}}} - \delta Z_{\mathcal{H}_1}^{\overline{\text{DR}}} \right)$$

$$[ ]^{\text{div}}: \propto \Delta \equiv 2/\epsilon - \gamma_E + \log(4\pi)$$

“More appropriate here:  $m\overline{\text{DR}}$ :

$$\delta Z_{\mathcal{H}_1}^{m\overline{\text{DR}}} = - \left[ \text{Re}\Sigma'_{HH} |_{\alpha=0} \right]^{m\text{div}}$$

$$\delta Z_{\mathcal{H}_2}^{m\overline{\text{DR}}} = - \left[ \text{Re}\Sigma'_{hh} |_{\alpha=0} \right]^{m\text{div}}$$

$$\delta \tan\beta^{m\overline{\text{DR}}} = \frac{1}{2} \left( \delta Z_{\mathcal{H}_2}^{m\overline{\text{DR}}} - \delta Z_{\mathcal{H}_1}^{m\overline{\text{DR}}} \right)$$

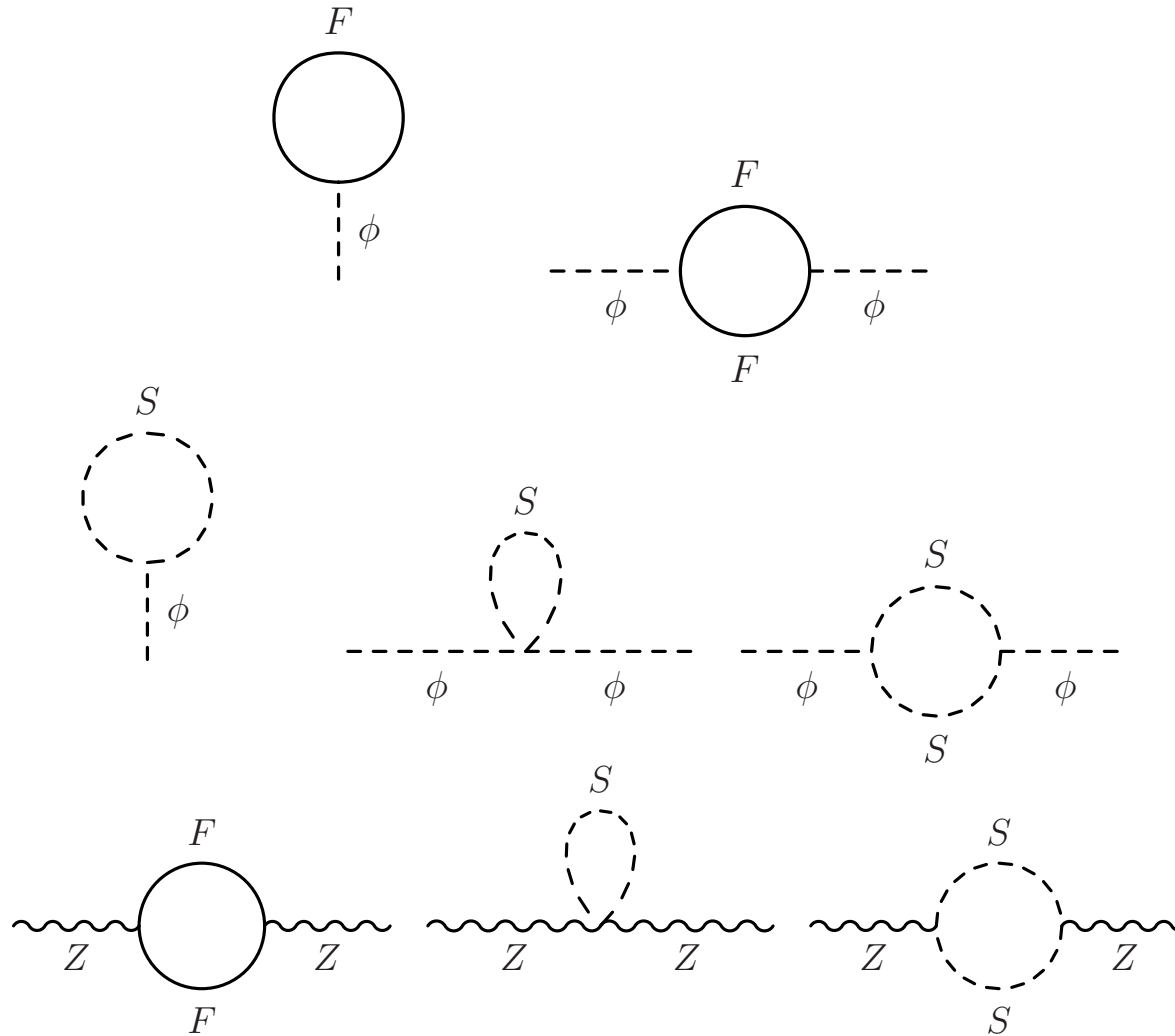
$$[ ]^{m\text{div}}: \propto \Delta_m \equiv \Delta - \log(m_M^2/\mu_{\overline{\text{DR}}}^2)$$

$\Rightarrow$  decoupling “by hand” of large logs (equivalent to set  $\mu_{\overline{\text{DR}}} = m_M$ )

## Calculation of Self-energies:

- all diagrams created with **FeynArts** → T
  - model file with all  $\nu/\tilde{\nu}$  interactions
  - further evaluation with **FormCalc**
  - Dimensional **RED**uction
  - all **UV** divergences cancel
  - results will be included into **FeynHiggs** ([www.feynhiggs.de](http://www.feynhiggs.de))
- example plots will focus on  $\widehat{\Sigma}_{hh}(p^2)$  and  $\Delta m_h^{\text{mDR}} := M_h^{\nu/\tilde{\nu}} - M_h$

# Feynman diagrams:



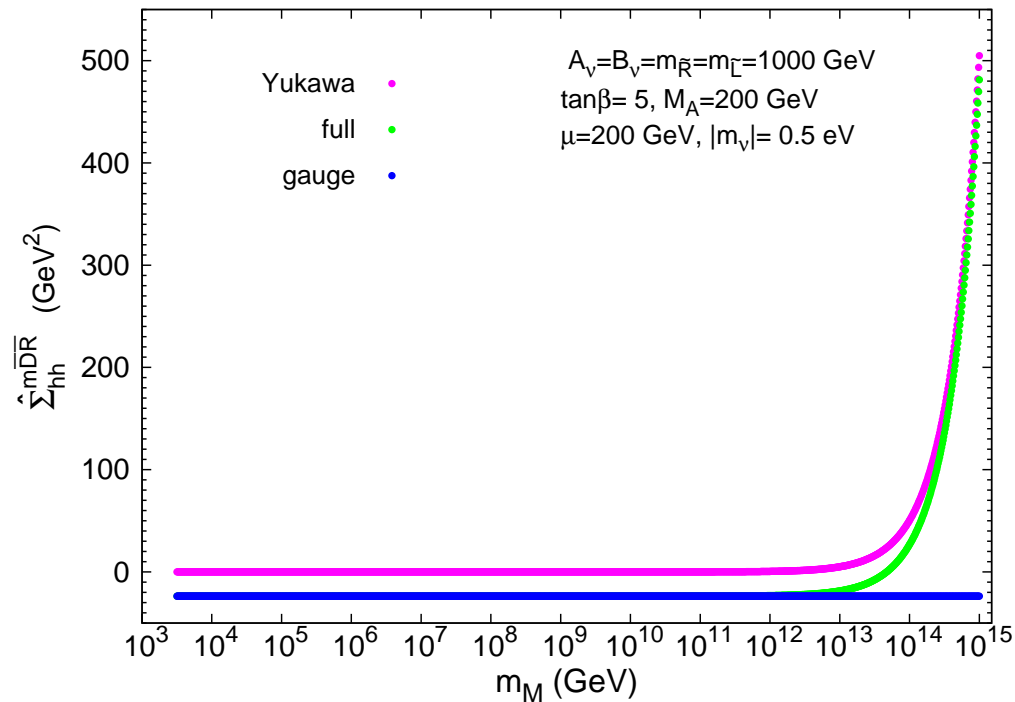
In mass eigenstate basis:

$$\phi = h, H; F = \nu, N; S = \tilde{\nu}_+, \tilde{\nu}_-, \tilde{N}_+, \tilde{N}_-$$

# Results

## Gauge part vs. Yukawa part:

$$\hat{\Sigma}(p^2)|_{\text{full}} = \underbrace{\hat{\Sigma}(p^2)|_{\text{gauge}}}_{\text{MSSM}} + \underbrace{\hat{\Sigma}(p^2)|_{\text{Yukawa}}}_{\text{extra}}$$

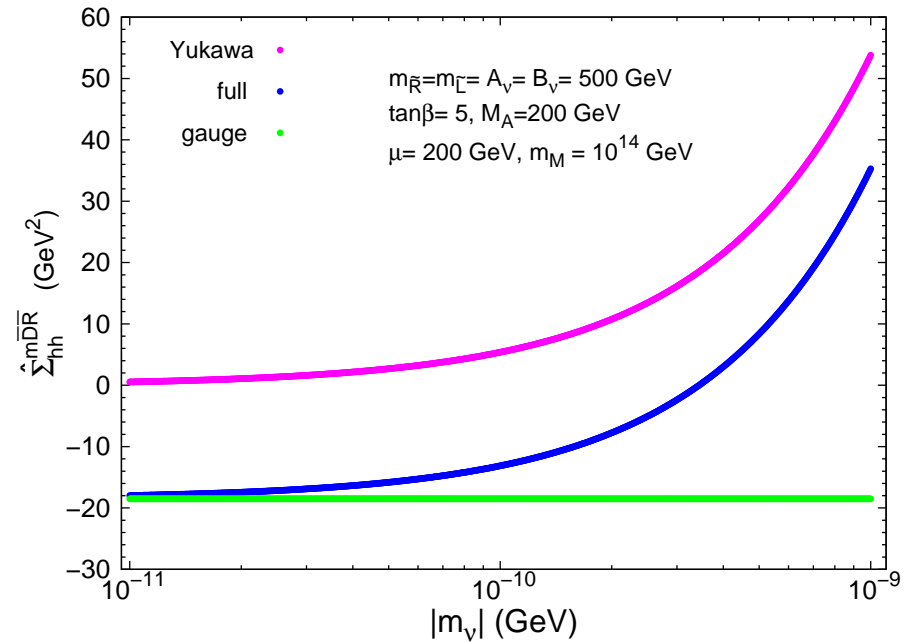
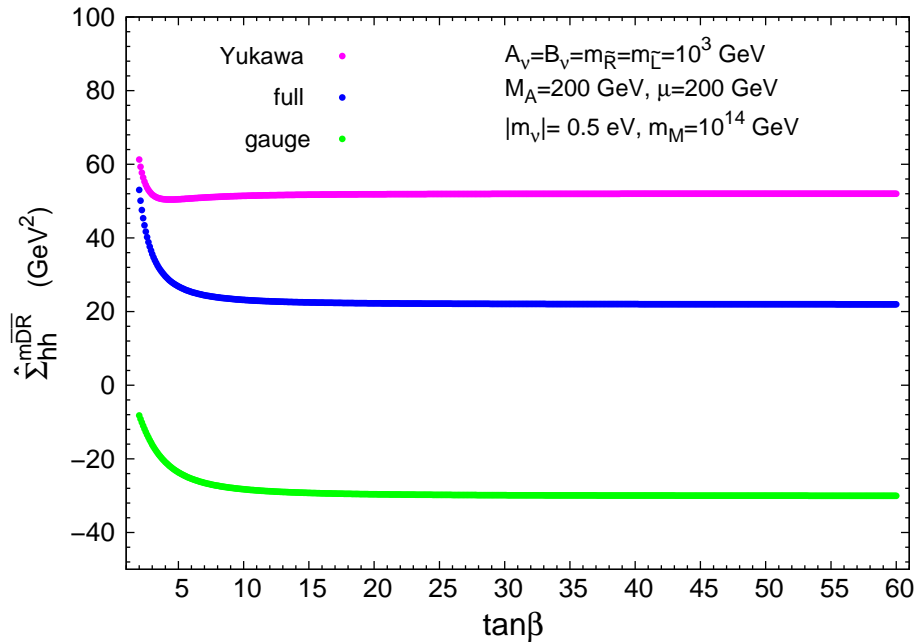


For  $m_M < 10^{12} \text{ GeV}$ , MSSM-seesaw  $\sim$  MSSM ( $\oplus$  Dirac neutrinos)

For  $m_M > 10^{13} \text{ GeV}$ , the Yukawa contribution dominates

$\hat{\Sigma}_{hh}^{m\overline{\text{DR}}}$  grows with  $m_M$ , sizeable RADCOR at  $10^{13} < m_M(\text{GeV}) < 10^{15}$

## Behaviour with $\tan\beta$ and $m_\nu$ :



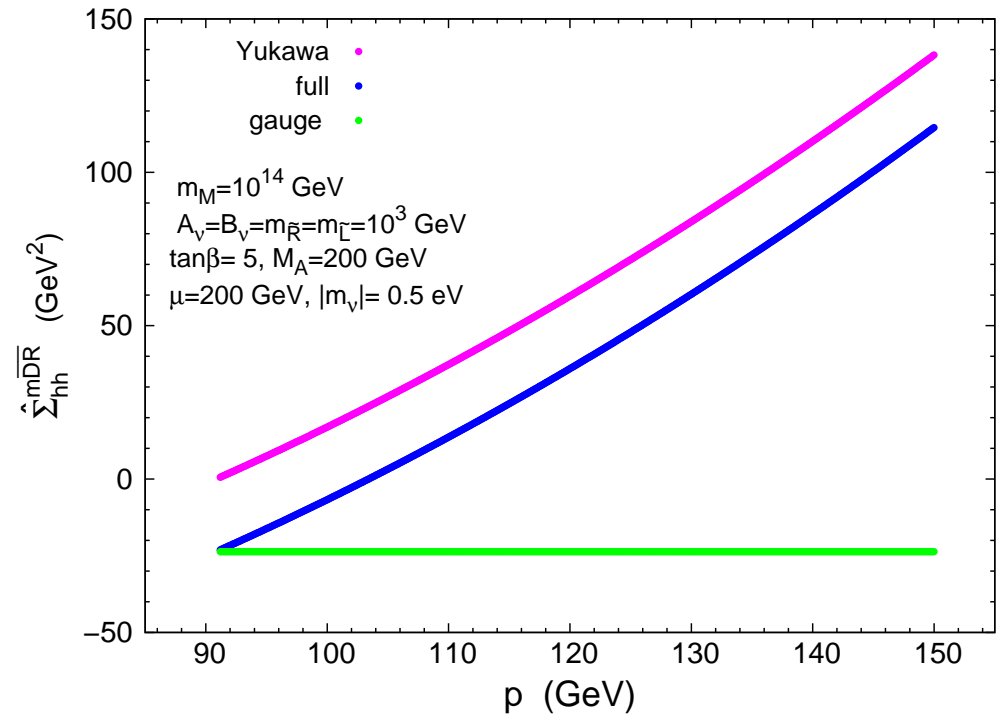
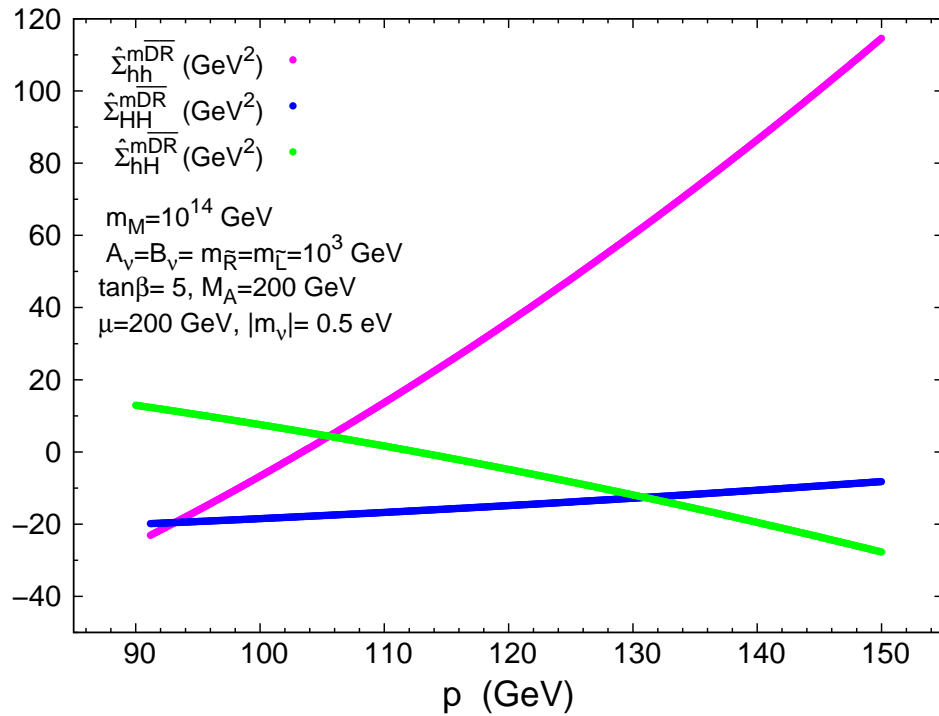
→ Not very relevant dependence on  $\tan\beta$  from  $\nu/\tilde{\nu}$  RADCOR:

Slightly larger RADCOR at low  $\tan\beta$

→ Relevant dependence on  $m_\nu$ :

$\hat{\Sigma}_{hh}^{mDR}$  grows with  $|m_\nu|$ , sizeable RADCOR at  $0.5 < |m_\nu(\text{eV})| < 1$

## Momentum dependence:



⇒ strong momentum dependence

⇒ only present in the Yukawa part

⇒ a mass correction estimate by setting  $p = 0$  is not a good approx. here

⇒ (contrary to  $\mathcal{O}(m_t^4)$  corrections)  $\mathcal{O}(m_D^2)$  term dominates (see next)



## Analysing the growing with $m_M$ : The seesaw expansion

For  $m_D \ll m_M$  a good approximation is to perform an expansion of  $\widehat{\Sigma}(p^2)$  in powers of  $\xi \equiv m_D/m_M$ :

$$\widehat{\Sigma}(p^2) = \underbrace{\left(\widehat{\Sigma}(p^2)\right)_{|m_D^0}}_{\text{gauge MSSM}} + \underbrace{\left(\widehat{\Sigma}(p^2)\right)_{|m_D^2} + \left(\widehat{\Sigma}(p^2)\right)_{|m_D^4} + \dots}_{\text{Yukawa}}$$

Our result: the  $\mathcal{O}(m_D^2)$  term dominates and provides a good estimate:

$$\left(\widehat{\Sigma}_{hh}^{m\overline{\text{DR}}}(p^2)\right)_{m_D^2} = \frac{g^2 m_D^2}{64\pi^2 M_W^2 \sin^2 \beta} \left[ -2M_A^2 \cos^2(\alpha - \beta) \cos^2 \beta + 2p^2 \cos^2 \alpha - M_Z^2 \sin \beta \sin(\alpha + \beta) \left( 2(1 + \cos^2 \beta) \cos \alpha - \sin 2\beta \sin \alpha \right) \right]$$

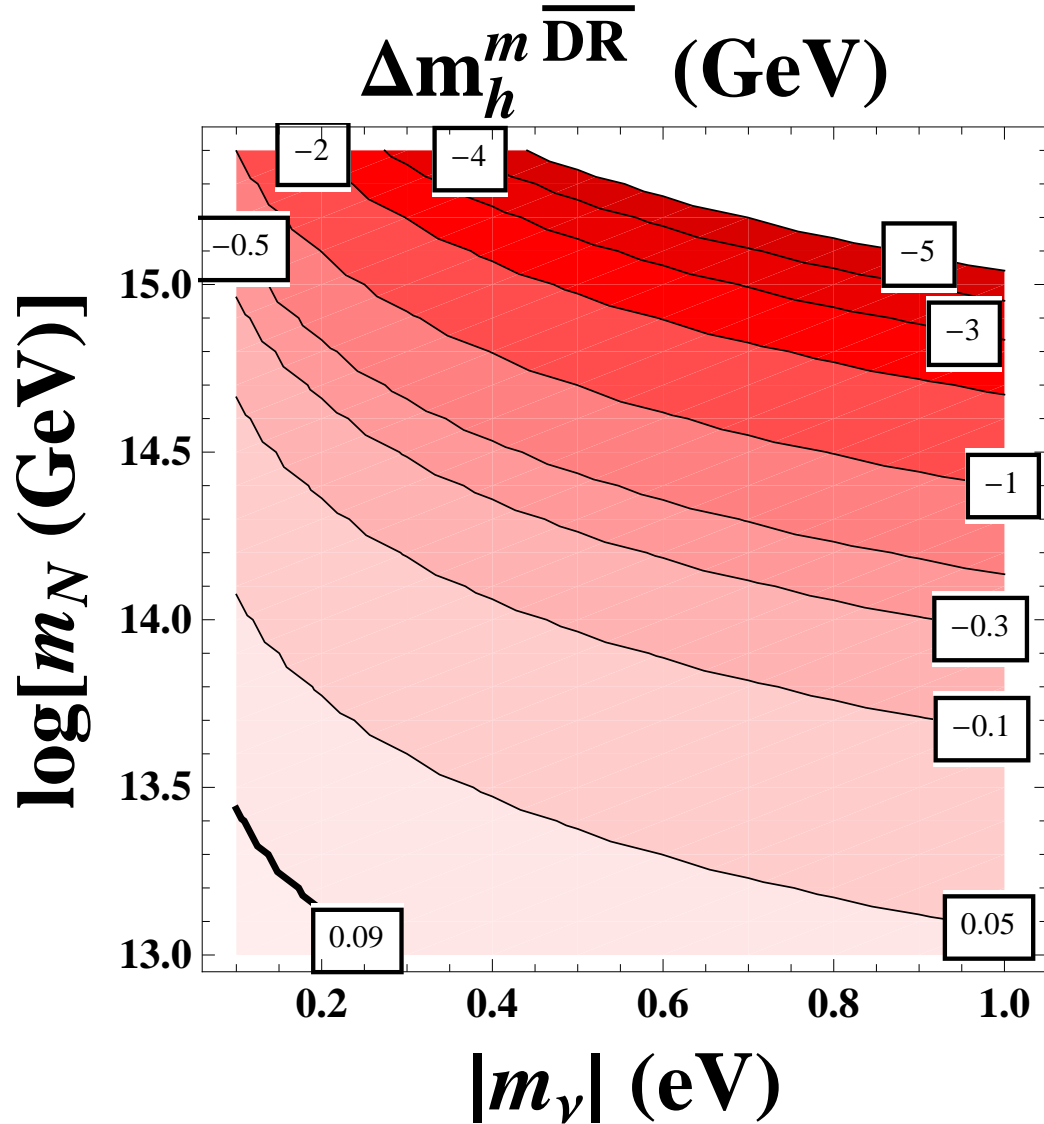
→ This is in contrast to the one-loop top/stop corrections where the dominant contributions are  $\mathcal{O}(m_t^4)$ , and setting  $p = 0$  is a reasonable approx.

→ This explains the growing with  $m_M$ , due to  $m_D^2 = |m_\nu| m_M$ .

This growing also appears at the physical Higgs mass correction (see next)

Main result:  $\Delta m_h^{m\overline{DR}}$ :

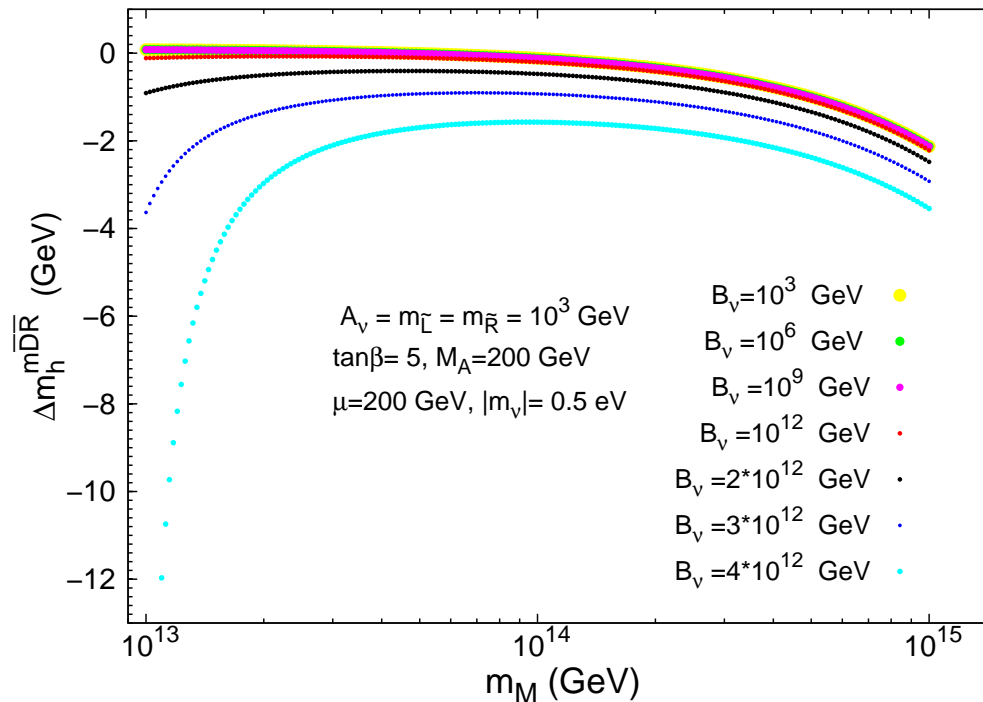
$$A_\nu = B_\nu = m_{\tilde{L}} = m_{\tilde{R}} = 1000 \text{ GeV}, M_A = \mu = 200 \text{ GeV}, \tan \beta = 5$$



⇒ large negative corrections possible for large  $|m_\nu|$  and  $m_N (\simeq m_M)$  (up to  $-5 \text{ GeV}$  for  $m_M = 10^{15} \text{ GeV}$  and  $|m_\nu| = 1 \text{ eV}$ )

Growing of  $\Delta m_h^{m\overline{DR}}$  with  $m_M$ :  
 due to  $Y_\nu = \frac{1}{v_2} \sqrt{m_M |m_\nu|}$

## Impact of $B_\nu$



⇒ we get large negative corrections for large  $B_\nu \sim m_M$

But so large  $B_\nu$  gives problems:

too large oscillations in the sneutrino sector ⇒

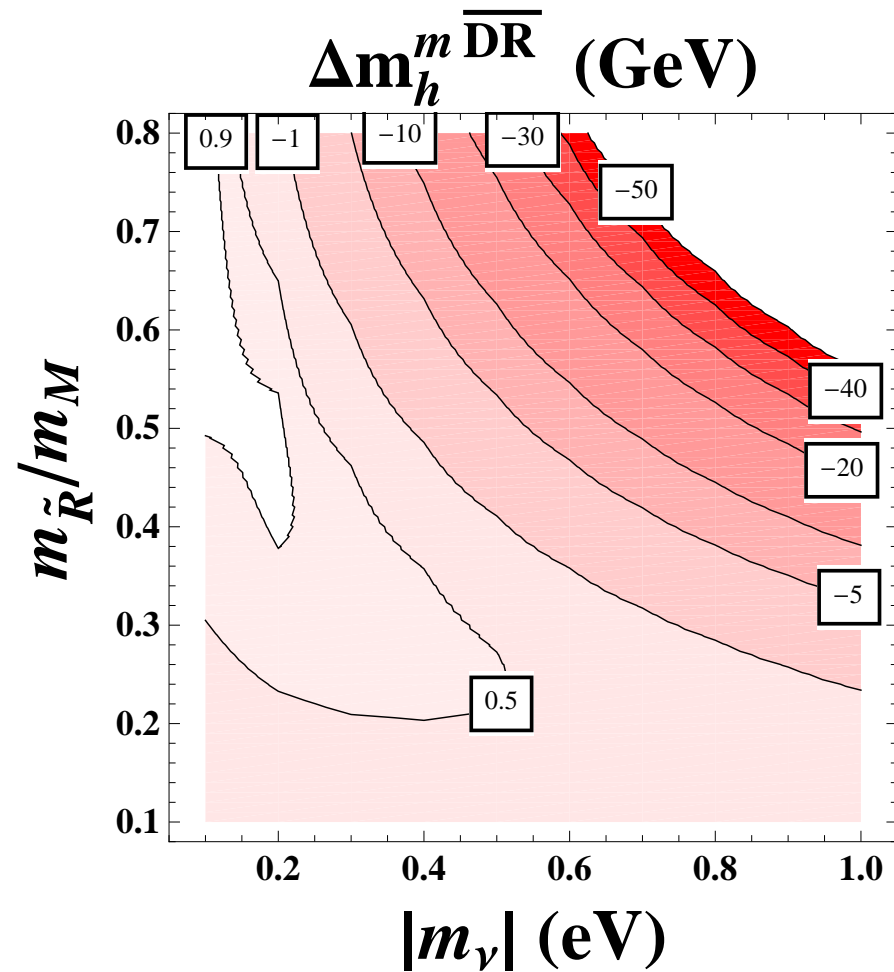
too large 1-loop generated  $m_\nu$ : Dedes, Haber, Rosiek JHEP11(2007)059

too large LFV rates: Kang, Morozumi, Yokozaki, JHEP11(2010)061

too large EDM: Farzan, PRD69(2004)073009;

Giudice, Paradisi, Strumia, PLB694(2010)26

The largest corrections  $\Delta m_h^{m\overline{\text{DR}}}$ :  $m_{\tilde{R}} \sim m_M$ :



⇒ very large and negative corrections!

$\Delta m_h^{m\overline{\text{DR}}} \sim -30 \text{ GeV}$  for  $m_M = 10^{14} \text{ GeV}$ ,  $m_{\tilde{R}}/m_M = 0.7$  and  $|m_\nu| = 0.6 \text{ eV}$

## An useful analytical result:

We have found a simple and accurate formula, valid for heavy  $m_M$ :

$$\Delta m_h^{\overline{\text{mDR}}} \simeq -\frac{\widehat{\Sigma}_{hh}^{\nu/\tilde{\nu}}(M_h^2)}{2M_h} \approx -\frac{\left(\widehat{\Sigma}_{hh}^{\overline{\text{mDR}}}(M_h^2)\right)_{m_D^2}}{2M_h} \approx$$
$$\frac{-g^2 m_D^2}{128\pi^2 M_W^2 M_h \sin^2 \beta} \left[ -2M_A^2 \cos^2(\alpha - \beta) \cos^2 \beta + 2M_h^2 \cos^2 \alpha \right. \\ \left. - M_Z^2 \sin \beta \sin(\alpha + \beta) \left( 2(1 + \cos^2 \beta) \cos \alpha - \sin 2\beta \sin \alpha \right) \right]$$

(if corrections are not too large ...

otherwise full pole determination necessary  $\Rightarrow$  FeynHiggs ...)

## Comments:

$\Rightarrow$  growing with  $m_M$  (and  $|m_\nu|$ ) explained by prefactor  $m_D^2 = |m_\nu| m_M$

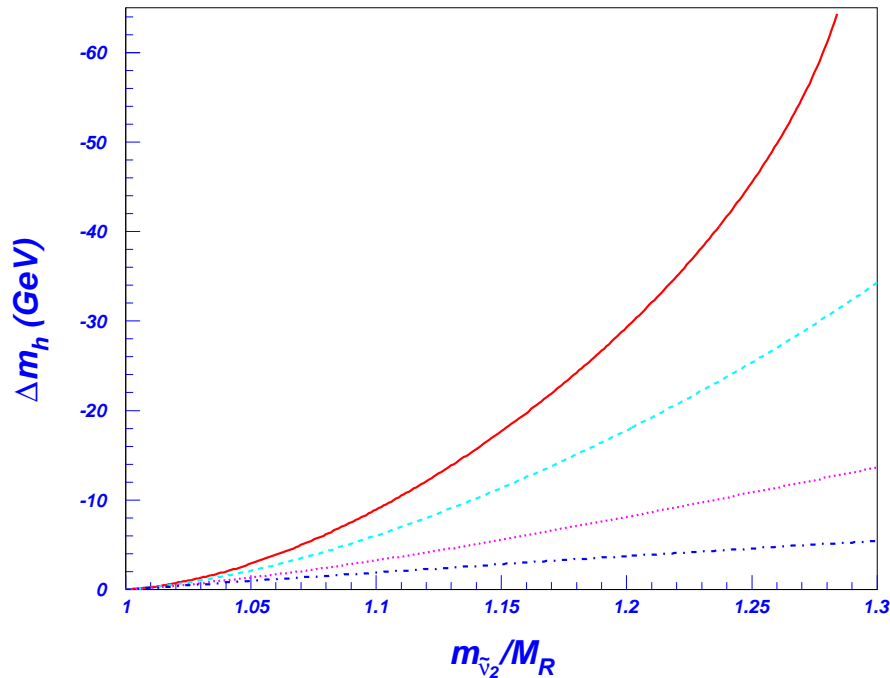
$\Rightarrow$  result not obtainable in the effective potential approach (it sets  $p = 0$ )

$\Rightarrow$  result not obtainable in the RGE approach (it mainly provides Llogs)

## Comparison with other works:

→ We agree with Cao, Yang (Phys.Rev.D71(2005)111701):

same size and sign of the corrections for large  $m_{\tilde{R}} \sim m_M$



Simplified diagrammatic estimate

$$m_M = 10^{14} \text{ GeV}, A_\nu = B_\nu = 0,$$

$$M_A = 200 \text{ GeV}, \tan \beta = 30$$

See **solid red line**,  $m_{\tilde{L}} = 1000 \text{ GeV}$

$$(m_{\tilde{\nu}_2}/M_R)_{\text{Cao}} = 1.2 \leftrightarrow m_{\tilde{R}}/m_M \sim 0.7$$

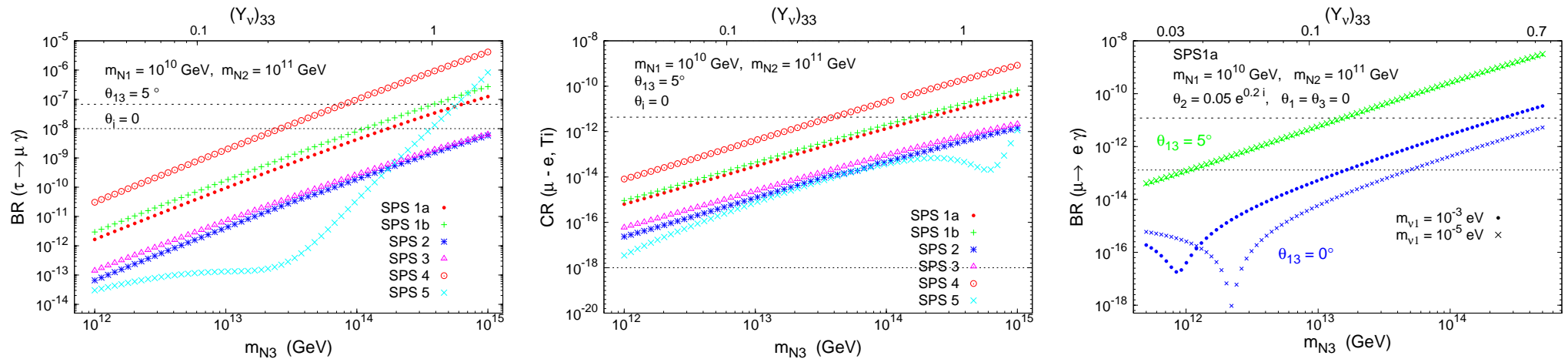
$\Delta m_h \sim -30 \text{ GeV}$ , **OK with us**.

→ We agree qualitatively with Kang, Morozumi, Yokozaki JHEP11(2010)061 on the big impact of large  $B_\nu$ . They use RGEs with heavy-sneutrino thresholds

→ The computation of Dedes, Haber, Rosiek (JHEP11(2007)059) is not comparable to ours. They use  $V_{\text{eff}}$  approach  $\Rightarrow$  do not capture our leading  $\mathcal{O}(m_D^2)$  contribution.

## Comparison with LFV:

The large RADCOR from Heavy Majorana neutrinos and the growing with  $m_M$  are well known in LFV processes:

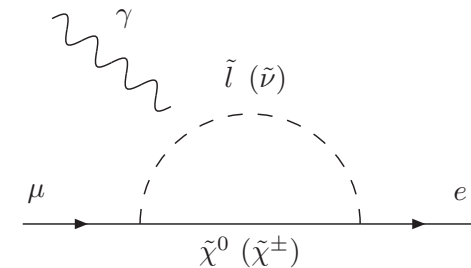


(Antusch, Arganda, Herrero, Teixeira, JHEP11(2006)090)

(Arganda, Herrero, Teixeira, JHEP10(2007)104)

$$\text{BR}(\mu \rightarrow e \gamma) = \frac{\alpha^3 \tan^2 \beta}{G_F^2 m_{\text{SUSY}}^8} \left| \frac{1}{8\pi^2} (3M_0^2 + A_0^2) (Y_\nu^\dagger L Y_\nu) \right|^2$$

$$v_2^2 (Y_\nu^\dagger L Y_\nu) \sim m_M m_\nu \log \left( \frac{m_M}{M_X} \right) \text{ from internal } \tilde{l}(\tilde{\nu})$$



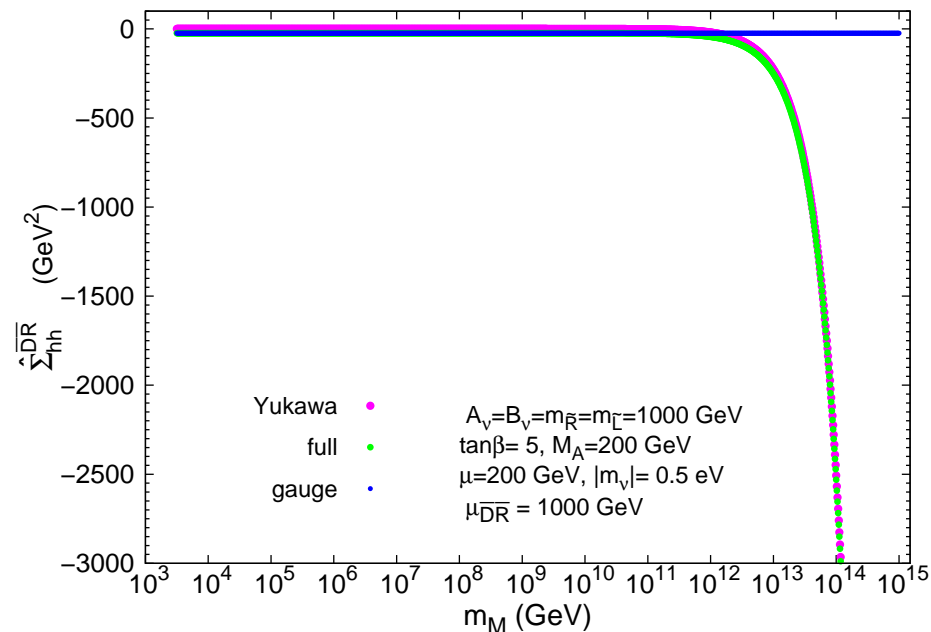
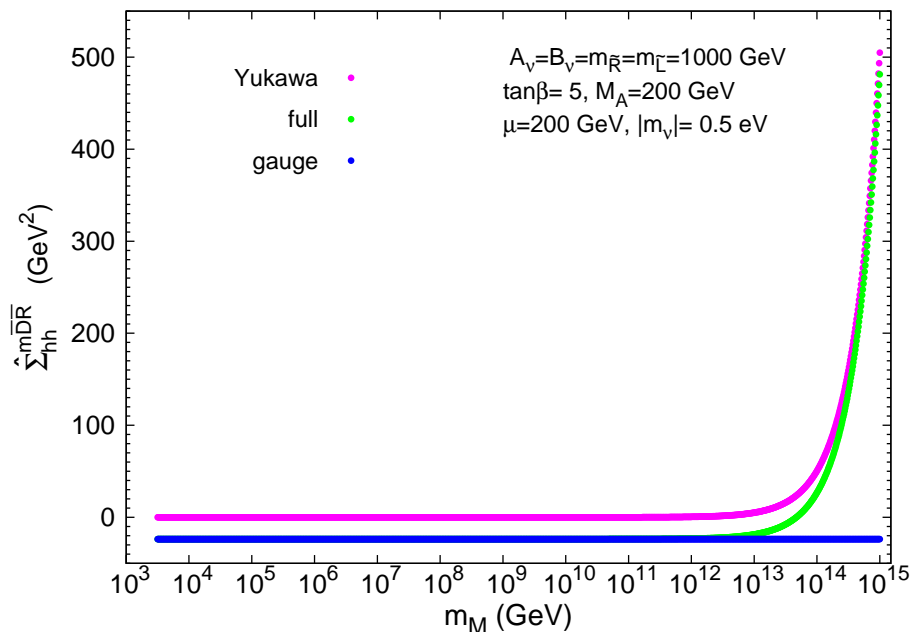
## Conclusions

- Neutrinos do have mass and oscillate  
⇒ good motivation for seesaw mass generation via heavy Majorana right-handed neutrinos
- The MSSM Higgs sector is sensitive to the heavy Majorana scale  $m_M$  via RADCOR from  $\nu/\tilde{\nu}$  sector, due to the large  $Y_\nu \sim \mathcal{O}(1)$
- $\Delta m_h$  grow with  $m_M$ . Similar to LFV:  $\tau \rightarrow \mu\gamma$ ,  $\mu \rightarrow e\gamma$ ,  $\mu - e$  conv.,...
- $\Delta m_h$  numerically relevant for  $m_M > 10^{13}$  GeV:
  - negative ⇒ push down the lightest Higgs mass
  - up to  $\sim -5$  GeV if  $m_{\tilde{R}} \ll m_M$
  - up to  $\sim (-10, -50)$  GeV if  $m_{\tilde{R}} \sim m_M$**RADCOR larger than the anticipated exp. precision!**  
(LHC:  $\sim 0.2$  GeV, ILC:  $\sim 0.05$  GeV)
- In progress: Generalization to three neutrinos/sneutrinos generations  
**Connection between neutrino physics/data (oscillations, masses etc) and Higgs physics/data (masses, couplings, etc)?**



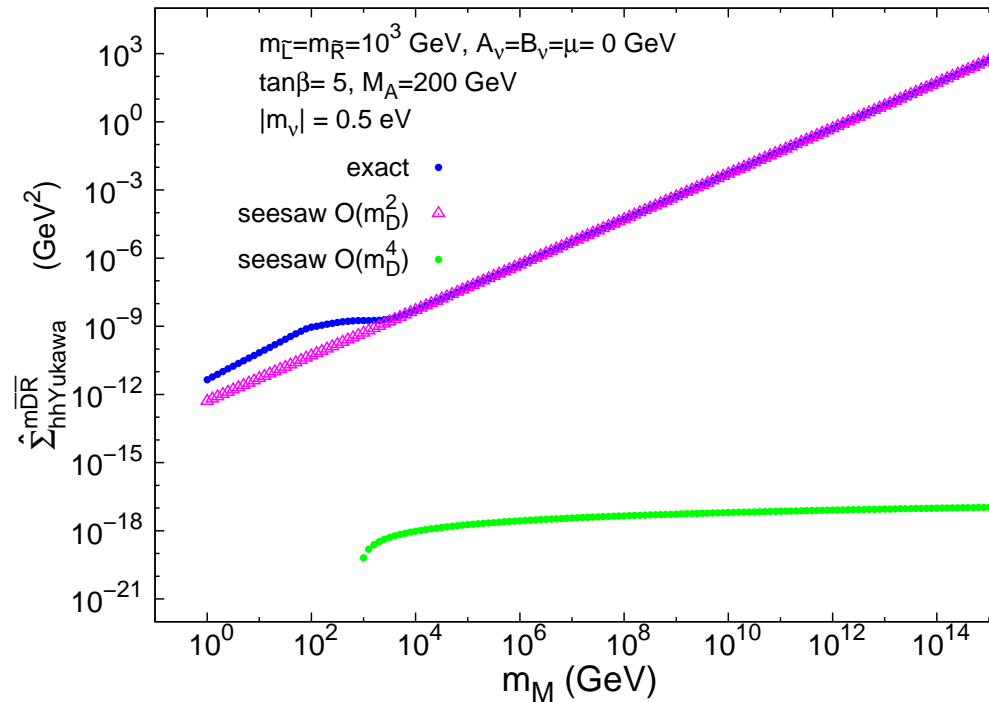
# Backup

## m $\overline{\text{DR}}$ versus $\overline{\text{DR}}$ :



- Large logs like  $\log\left(\frac{m_M}{\mu_{\overline{\text{DR}}}}\right)$  in  $\overline{\text{DR}}$ . No large logs in  $m\overline{\text{DR}}$
- Growing with  $m_M$  faster in  $\overline{\text{DR}}$  than in  $m\overline{\text{DR}}$
- Larger radiative corrections in  $\overline{\text{DR}}$  than in  $m\overline{\text{DR}}$ ,
- but  $\overline{\text{DR}}$  perturbatively more unstable than  $m\overline{\text{DR}}$ .

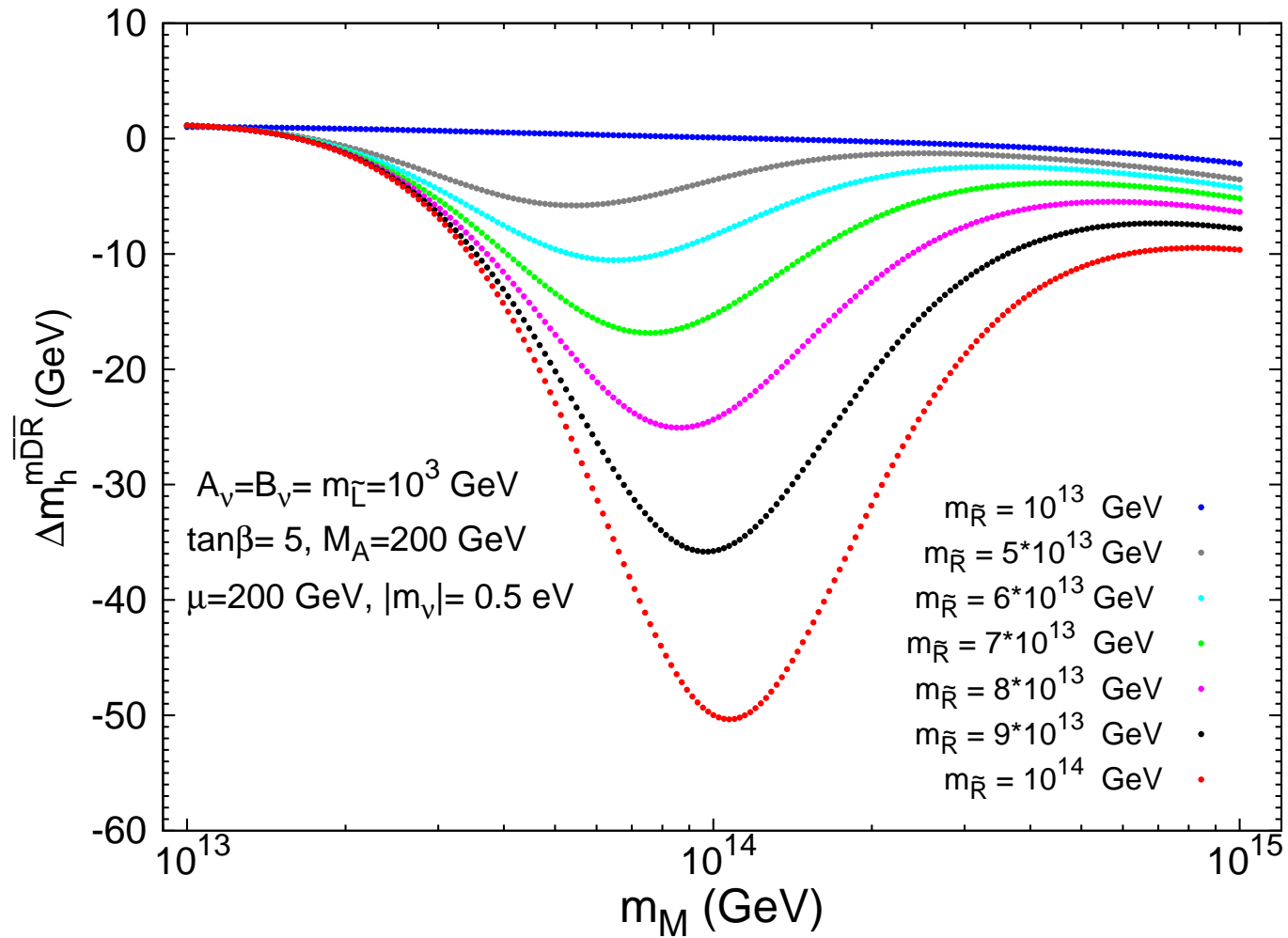
$\mathcal{O}(m_D^2)$  versus  $\mathcal{O}(m_D^4)$  :



→  $\mathcal{O}(m_D^2)$  dominates  $\mathcal{O}(m_D^4)$  by many orders of magnitude

→  $\mathcal{O}(m_D^2)$  term in seesaw expansion reproduces accurately the full result

## Dependence on $m_{\tilde{R}}$ :



⇒ large corrections for  $m_{\tilde{R}} \sim m_M$