

Exact Amplitude-Based Resummation in Quantum Field Theory: Recent Results

BFL Ward

Baylor University

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w S. Majhi, S. Yost

Preface

** NEW PHYSICS AT LHC

Must Distinguish from Higher
Order SM Processes AND Must
Probe Precisely to Specify Uniquely

⇒ Precision QCD for the LHC

** UV LIMIT OF EINSTEIN'S THEORY

Can QFT Handle It?

⇒ Exact Residual Control in Resummation for
UV



The Paradigm:

*An Approach to Precision LHC
Physics Theory: Amplitude-Based
QED \otimes QCD Resummation Realized
by MC Methods --

$$d\sigma = \sum_{i,j} \int dx_1 dx_2 F_i(x_1) F_j(x_2) d\hat{\sigma}_{res}$$

$$\Rightarrow \Delta\sigma_{th} = \Delta F \oplus \Delta\hat{\sigma}_{res} = \Delta\sigma_{th}(tech) \oplus \Delta\sigma_{th}(phys)$$

\Rightarrow



$$\begin{aligned}
d\hat{\sigma}_{res} &= \sum_n d\hat{\sigma}_n \\
&= e^{SUM_{IR}(QCED)} \sum_{m,n=0}^{\infty} \frac{1}{m!n!} \int \prod_{j_1=1}^m \frac{d^3 k_{j_1}}{k_{j_1}} \prod_{j_2=1}^n \frac{d^3 k_{j_2}}{k_{j_2}} \int \frac{d^4 y}{(2\pi)^4} e^{iy(p_1+q_1-p_2-q_2-\sum_{j_1} k_{j_1}-\sum_{j_2} k_{j_2})+D_{QCED}} \\
&\quad * \tilde{\beta}_{m,n}(k_1, \dots, k_m; k_1^{'}, \dots, k_n^{'}) \frac{d^3 p_2}{p_2^0} \frac{d^3 q_2}{q_2^0} \tag{1}
\end{aligned}$$



The Paradigm: *An Approach to Feynman's Formulation of Einstein's Theory : Amplitude-Based Resummation of the Feynman Propagators in the Theory

$$\begin{aligned} i\Delta_F'(k) &= \frac{i}{k^2 - m^2 - \Sigma_s(k) + i\varepsilon} \\ &= \frac{ie^{B_g''(k)}}{k^2 - m^2 - \Sigma_s'(k) + i\varepsilon} \\ &\equiv i\Delta_F'(k)|_{\text{Re summed}} \end{aligned}$$



Precision QCD for the LHC

- Contact with Standard Resummation (Abyat et al. , Phys. Rev. D 74 (2006) 074004):
2→n process -

$$\begin{aligned}\mathcal{M}_{\{r_i\}}^{[f]} &= \sum_{L=1}^C \mathcal{M}_L^{[f]}(c_L)_{\{r_i\}} \\ &= J^{[f]} \sum_{L=1}^C S_{LI} H_I^{[f]}(c_L)_{\{r_i\}}\end{aligned}$$

- Noting relation of J and SUM_{IR} , in our formula

$$d\hat{\mathcal{O}}^m = \frac{e^{2\alpha_s \text{Re } B_{QCD}}}{m!} \int \prod_{j=1}^m \frac{d^3 k_j}{(k_j^2 + \lambda^2)^{1/2}} \delta(p_1 + q_1 - p_2 - q_2 - \sum_{i=1}^m k_i)$$
$$\bar{\rho}^{(m)}(p_1, q_1, p_2, q_2, k_1, \dots, k_m) \frac{d^3 p_2 d^3 q_2}{p_2^0 q_2^0},$$



we get the identification:

$$\begin{aligned}\overline{\rho}^{(m)}(p_1, q_1, k_1, \dots, k_m) &= \sum_{\text{colors, spin}} |\mathcal{M}_{\{r_i\}}^{[f]}|^2 \\ &\equiv \sum_{\text{spins}, \{r_i\}, \{r'_i\}} h_{\{r_i\} \{r'_i\}}^{\text{cs}} |\bar{\mathbf{J}}^{[f]}|^2 \sum_{L, L'=1}^C S_{LL'}^{[f]} H_I^{[f]}(\mathbf{c}_L)_{\{r_i\}} \left(S_{L'L'}^{[f]} H_F^{[f]}(\mathbf{c}_{L'})_{\{r'_i\}} \right)^*\end{aligned}$$



- **Shower/ME Matching:**

$\tilde{\beta}_{m,n} \rightarrow \hat{\beta}_{m,n}$, shower - subtracted residuals

- **IR-Improved DGLAP-CS Theory(PRD81(2010)076008):**
New resummed scheme for P_{AB} , reduced cross section derived from (1) applied to splitting process --

$F_j, \hat{\sigma} \rightarrow F'_j, \hat{\sigma}'$ for

$$P_{qq} \rightarrow P_{qq}^{\text{exp}} = C_F F_{YFS}(\gamma_q) e^{\frac{1}{2}\delta_q} \frac{1+z^2}{1-z} (1-z)^{\gamma_q}, \text{etc.},$$

giving the same value for σ , with improved MC stability
-- no need for IR cut - off (k_0) parameter

- Complete Set:

$$\begin{aligned}
 P_{qq}^{exp}(z) &= C_F F_{YFS}(\gamma_q) e^{\frac{1}{2}\delta_q} \left[\frac{1+z^2}{1-z} (1-z)^{\gamma_q} - f_q(\gamma_q) \delta(1-z) \right], \\
 P_{Gq}^{exp}(z) &= C_F F_{YFS}(\gamma_q) e^{\frac{1}{2}\delta_q} \frac{1+(1-z)^2}{z} z^{\gamma_q}, \\
 P_{GG}^{exp}(z) &= 2C_G F_{YFS}(\gamma_G) e^{\frac{1}{2}\delta_G} \left\{ \frac{1-z}{z} z^{\gamma_G} + \frac{z}{1-z} (1-z)^{\gamma_G} \right. \\
 &\quad \left. + \frac{1}{2} (z^{1+\gamma_G}(1-z) + z(1-z)^{1+\gamma_G}) - f_G(\gamma_G) \delta(1-z) \right\}, \\
 P_{qG}^{exp}(z) &= F_{YFS}(\gamma_G) e^{\frac{1}{2}\delta_G} \frac{1}{2} \{ z^2 (1-z)^{\gamma_G} + (1-z)^2 z^{\gamma_G} \},
 \end{aligned}$$

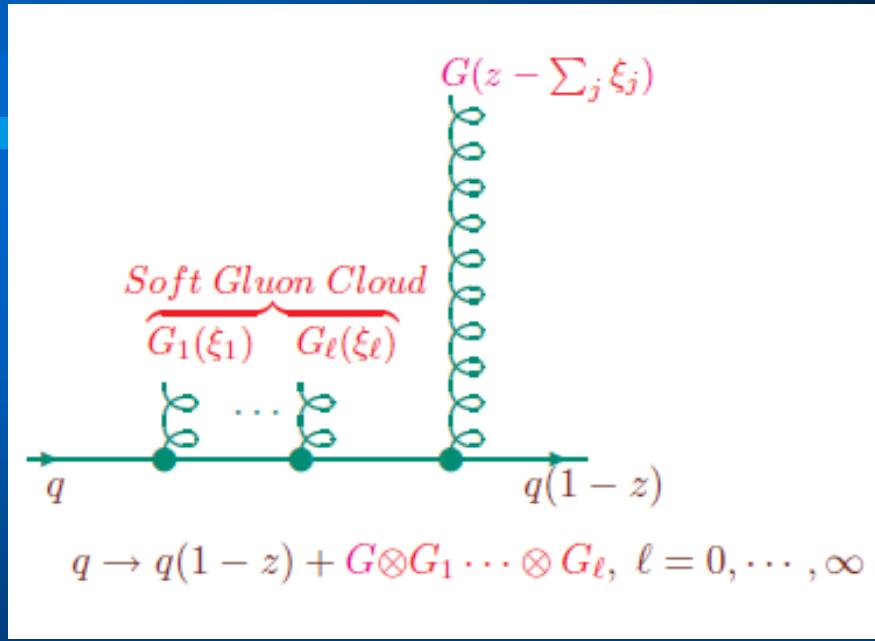
$$\begin{aligned}
 \gamma_q &= C_F \frac{\alpha_s}{\pi} t = \frac{4C_F}{\beta_0}, & \delta_q &= \frac{\gamma_q}{2} + \frac{\alpha_s C_F}{\pi} \left(\frac{\pi^2}{3} - \frac{1}{2} \right), \\
 f_q(\gamma_q) &= \frac{2}{\gamma_q} - \frac{2}{\gamma_q+1} + \frac{1}{\gamma_q+2}, \\
 \gamma_G &= C_G \frac{\alpha_s}{\pi} t = \frac{4C_G}{\beta_0}, & \delta_G &= \frac{\gamma_G}{2} + \frac{\alpha_s C_G}{\pi} \left(\frac{\pi^2}{3} - \frac{1}{2} \right), \\
 f_G(\gamma_G) &= \frac{n_f}{6C_G F_{YFS}(\gamma_G)} e^{-\frac{1}{2}\delta_G} + \frac{2}{\gamma_G(1+\gamma_G)(2+\gamma_G)} + \frac{1}{(1+\gamma_G)(2+\gamma_G)} \\
 &\quad + \frac{1}{2(3+\gamma_G)(4+\gamma_G)} + \frac{1}{(2+\gamma_G)(3+\gamma_G)(4+\gamma_G)}, \\
 F_{YFS}(\gamma) &= \frac{e^{-C\gamma}}{\Gamma(1+\gamma)}, & C &= 0.57721566...
 \end{aligned}$$



Basic Physical Idea: Bloch-Nordsiek –

Accelerated Charge \Rightarrow Coherent State of Soft
Gluons (Photons)

\Rightarrow More Physical View of Splitting Process:



•Illustration of Calculation

$$\begin{aligned}
 \int \frac{\alpha_s(t)}{2\pi} P_{BA} dt dz &= e^{\text{SUM}_{\text{IR}}(\text{QCD})(z)} \int \left\{ \bar{\beta}_0 \int \frac{d^4 y}{(2\pi)^4} e^{(iy \cdot (p_1 - p_2) + \int^{k < K_{\max}} \frac{d^3 k}{k} \tilde{S}_{\text{QCD}}(k) [e^{-iy \cdot k} - 1])} \right. \\
 &\quad + \int \frac{d^3 k_1}{k_1} \bar{\beta}_1(k_1) \int \frac{d^4 y}{(2\pi)^4} e^{(iy \cdot (p_1 - p_2 - k_1) + \int^{k < K_{\max}} \frac{d^3 k}{k} \tilde{S}_{\text{QCD}}(k) [e^{-iy \cdot k} - 1])} \\
 &\quad + \dots \} \frac{d^3 p_2}{p_2^0} \frac{d^3 q_2}{q_2^0} \\
 &= e^{\text{SUM}_{\text{IR}}(\text{QCD})(z)} \int \left\{ \bar{\beta}_0 \int_{-\infty}^{\infty} \frac{dy}{(2\pi)} e^{(iy \cdot (E_1 - E_2) + \int^{k < K_{\max}} \frac{d^3 k}{k} \tilde{S}_{\text{QCD}}(k) [e^{-iy \cdot k} - 1])} \right. \\
 &\quad + \int \frac{d^3 k_1}{k_1} \bar{\beta}_1(k_1) \int_{-\infty}^{\infty} \frac{dy}{(2\pi)} e^{(iy \cdot (E_1 - E_2 - k_1^0) + \int^{k < K_{\max}} \frac{d^3 k}{k} \tilde{S}_{\text{QCD}}(k) [e^{-iy \cdot k} - 1])} \\
 &\quad + \dots \} \frac{d^3 p_2}{p_2^0 q_2^0}
 \end{aligned}$$



$$\begin{aligned}
I_{YFS}(zE, 0) &= \int_{-\infty}^{\infty} \frac{dy}{2\pi} e^{[iy(zE) + \int^{k < zE} \frac{d^3 k}{k} \tilde{S}_{QCD}(k)(e^{-iyk} - 1)]} \\
&= F_{YFS}(\gamma_q) \frac{\gamma_q}{zE} \\
I_{YFS}(zE, k_1) &= \int_{-\infty}^{\infty} \frac{dy}{2\pi} e^{[iy(zE - k_1) + \int^{k < zE} \frac{d^3 k}{k} \tilde{S}_{QCD}(k)(e^{-iyk} - 1)]} \\
&= \left(\frac{zE}{zE - k_1} \right)^{1-\gamma_q} I_{YFS}(zE, 0)
\end{aligned}$$

$$\int \left(\bar{\beta}_0 \frac{\gamma_q}{zE} + \int dk_1 k_1 d\Omega_1 \bar{\beta}_1(k_1) \left(\frac{zE}{zE - k_1} \right)^{1-\gamma_q} \frac{\gamma_q}{zE} \right) \frac{d^3 p_2}{E_2 q_2^0} = \int dt \frac{\alpha_s(t)}{2\pi} P_{BA}^0 dz + \mathcal{O}(\alpha_s^2).$$

differentiation yields

$$P_{BA} = P_{BA}^0 z^{\gamma_q} F_{YFS}(\gamma_q) e^{\frac{1}{2}\delta_q}$$

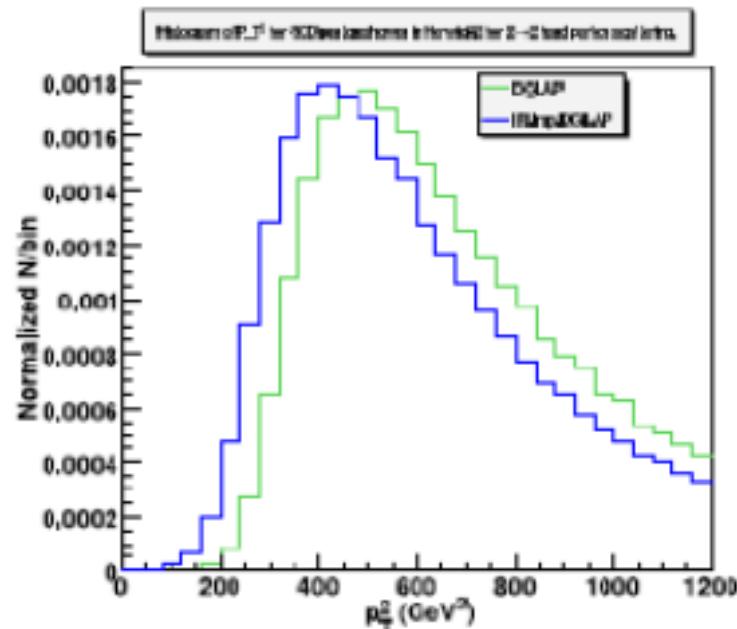
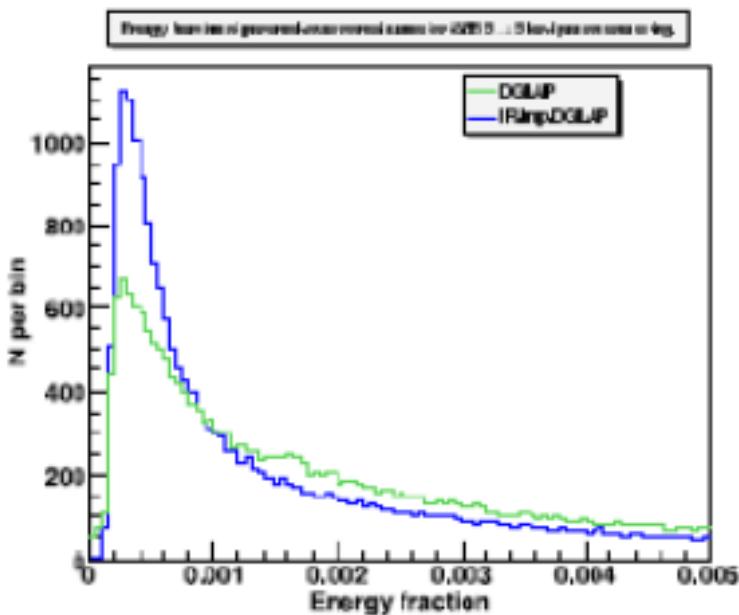


- Herwiri1.031(PRD81(2010)076008):
w consultation from Bryan Webber, Stefano Frixione, and Mike Seymour, implementation of IR-improved kernels in Herwig 6.5 environment to get Herwiri1.031,
MC@NLO/Herwiri1.031.

Observations:

1. SUM_{IR} is an IR effect – It contains as designed only the IR part of the LL, the rest of the LL is in D and the residuals $\hat{\tilde{\beta}}_m$, as we show in $\bar{\rho}^m$.
 2. Herwiri is just as general as Herwig6.5, as they run the same set of processes
- Some illustrative results follow.





Energy fraction distribution of parton shower for single Z production.

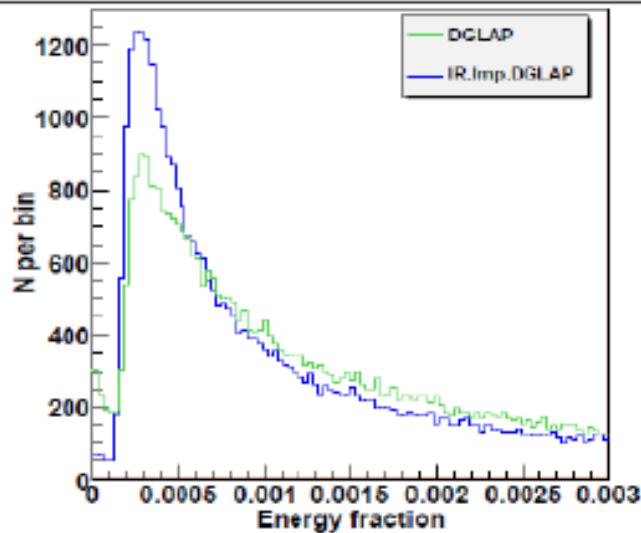


FIG.3: The z-distribution(ISR parton energy fraction) shower comparison
in HERWIG6.5.

Generated Z transverse momentum.

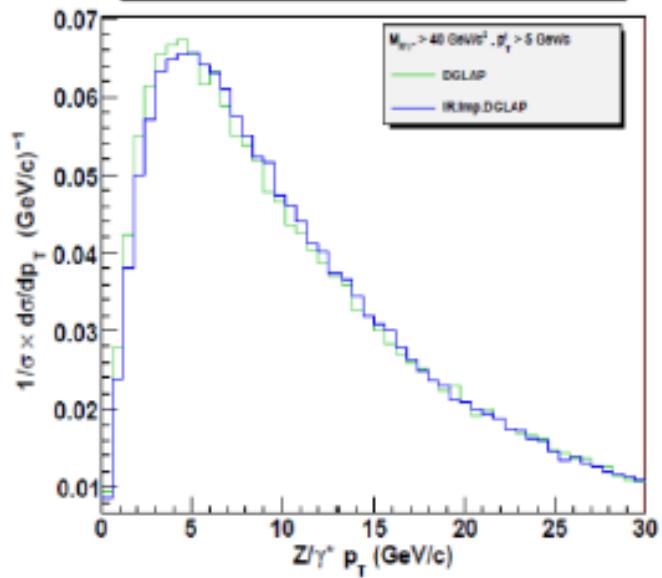


FIG.4 : The Z p_T -distribution(ISR parton shower effect) comparison
in HERWIG6.5.

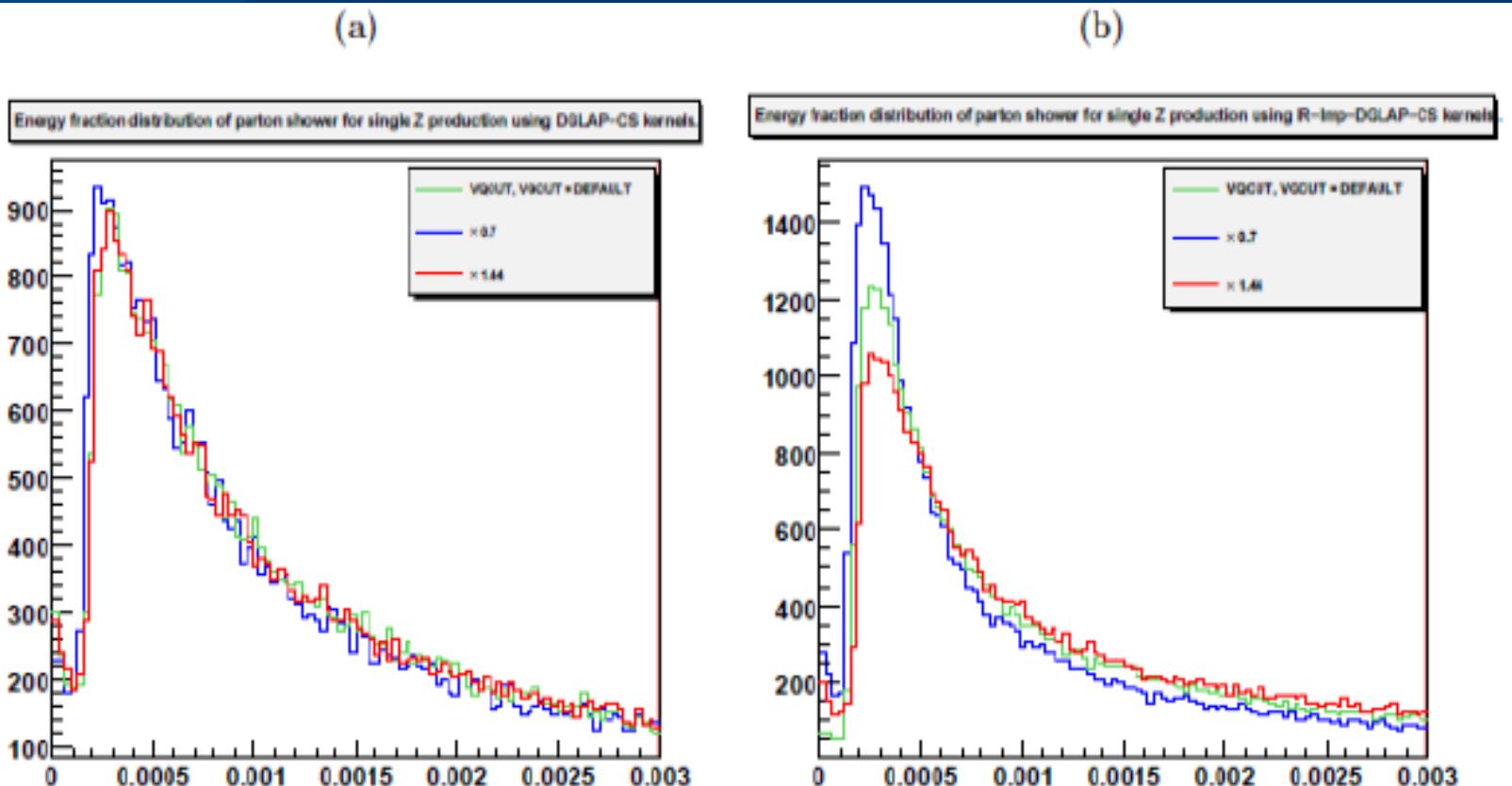


FIG.5: IR-cut-off sensitivity in z-distributions of the ISR parton energy fraction: (a), DGLAP-CS; (b), IR-I DGLAP-CS.



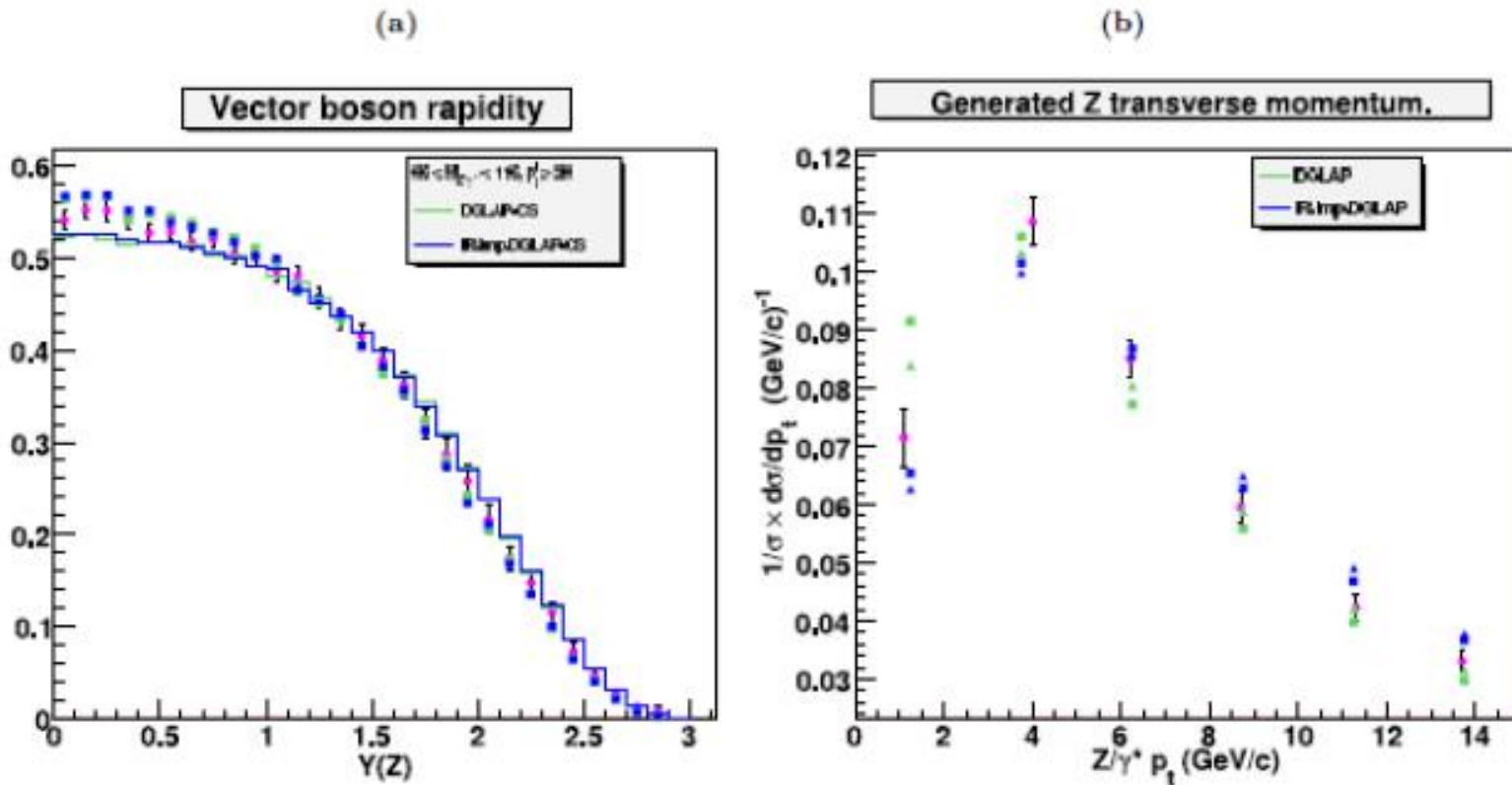


FIG. 6: Comparison with FNAL data: (a), CDF rapidity data on (Z/γ^*) production to e^+e^- pairs, the circular dots are the data, the green(blue) lines are HERWIG6.510(HERWIRI1.031); (b), D0 p_T spectrum data on (Z/γ^*) production to e^+e^- pairs, the circular dots are the data, the blue triangles are HERWIRI1.031, the green triangles are HERWIG6.510 – in both (a) and (b) the blue squares are MC@NLO/HERWIRI1.031, and the green squares are MC@NLO/HERWIG6.510. These are untuned theoretical results.



- HERWIRI1.031 -- PRD81 (2010) 076008:

- For the CDF rapidity data, HERWIRI1.031 is closer to the data than is HERWIG6.510 (1.54 vs 1.77 for $\chi^2/\text{d.o.f.}$ resp.);
for MC@NLO/HERWIRI1.031 and MC@NLO/HERWIG6.510
the $\chi^2/\text{d.o.f}$ are 1.42 and 1.40 resp., both are within 10% of the data
⇒ Need NNLO level, in progress.
- For the D0 p_T data, HERWIRI1.031 gives a better fit
to the data compared to HERWIG6.5 for low p_T ,
for $p_T < 12.5\text{GeV}$, the $\chi^2/\text{d.o.f.}$ are ~ 2.5 and 3.3 respectively
- we add the statistical and systematic errors,
showing that the IR-improvement makes a better representation
of QCD in the soft regime for a given fixed order in perturbation theory.

- HERWIRI1.031: PRD81 (2010) 076008 –
HERWIG++(SOON, thanks to Mike and Bryan)

Compare
CMS(Mans)
(soon) \Rightarrow

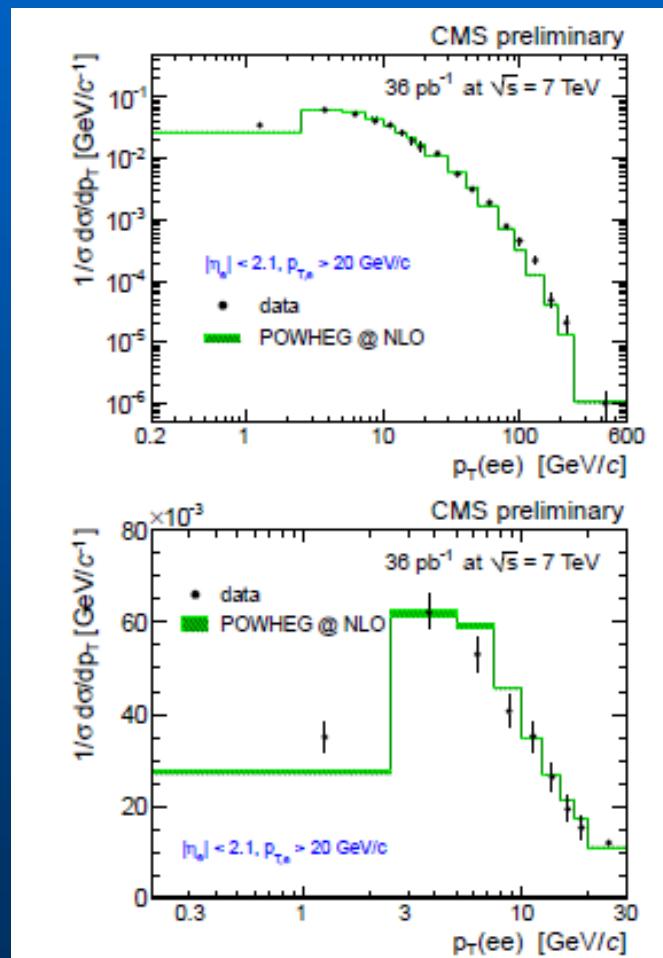
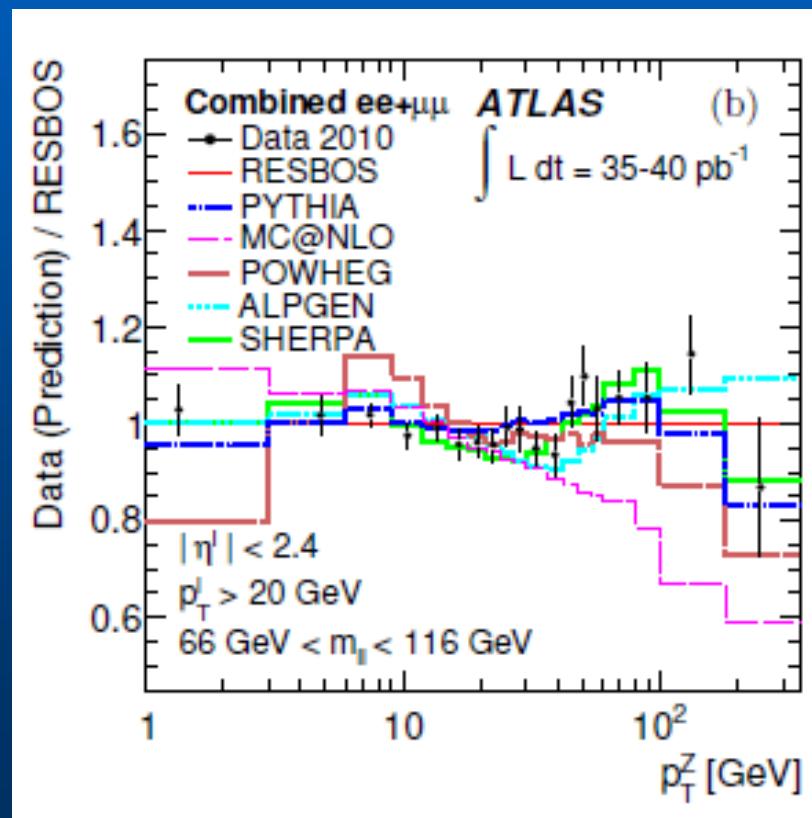


Figure 10: Comparison between the $Z \rightarrow e^+ e^-$ differential cross section shape to the POWHEG prediction for $|\eta_e| < 2.1$ and $p_{T,e} > 20 \text{ GeV}/c$. The same comparison is presented for the whole p_T range and zoomed in for the low p_T range ($< 30 \text{ GeV}/c$).

- HERWIRI1.031: PRD81 (2010) 076008 –
HERWIG++(SOON, thanks to Mike and Bryan)

Compare
ATLAS(1107.2381)
(soon) \Rightarrow



Near Future

- * Herwig++(soon, running , under cross checks)
- * Pyhtia 8,6 (w consultation from Peter Skands and Torbjorn Sjostrand)
- * Sherpa (w consultation from Jan Winter)



Resummed Quantum Gravity

- Recent Progress: Cosmological Constant Λ

In arXiv.org:1008.1048, using

$$B_g''(k) = \frac{\kappa^2 |k^2|}{8\pi^2} \ln \left(\frac{m^2}{m^2 + |k^2|} \right)$$

we show that we get the UV limit

$$k^2 G_N(k) \rightarrow .0442$$



and the scalar contribution to Λ as

$$\begin{aligned}\Lambda_s &= -8\pi G_N \frac{\int d^4k}{2(2\pi)^4} \frac{(2k_0^2)e^{-\lambda_c(k^2/(2m^2))\ln(k^2/m^2+1)}}{k^2 + m^2} \\ &\cong -8\pi G_N \left[\frac{1}{G_N^2 64\rho^2} \right], \quad \rho = \ln \frac{2}{\lambda_c}\end{aligned}$$

for $\lambda_c = \frac{2m^2}{M_{Pl}^2}$.

A Dirac fermion gives -4 times Λ_s .
 \Rightarrow UV limit

$$\begin{aligned}\Lambda(k) &\xrightarrow{k^2 \rightarrow \infty} k^2 \lambda_* \\ \lambda_* &= -\frac{c_{2eff}}{2880} \sum_j (-1)^{F_j} n_j / \rho_j^2 \\ &\cong 0.0817\end{aligned}$$



Comparison with EFRG(Reuter et al., Percacci et al,
Litim ,....):
Illustration(Laucsher&Reuter(PRD65(2002)025013))--

UV Fixed Point:

$$\begin{aligned}\beta_\lambda(\lambda_k, g_k; \alpha, d) = & -2\lambda_k + \nu_d d g_k + \left[2d(d-1+2\alpha)(4\pi)^{1-\frac{d}{2}} \Phi_{d/2}^2(0) - (d-2)\omega_d \right] \lambda_k g_k \\ & + \frac{1}{2} d(d+1)(d-2)(4\pi)^{1-\frac{d}{2}} \omega_d \Phi_{d/2}^1(0) g_k^2 + \mathcal{O}(g^3) ,\end{aligned}$$

$$\beta_g(\lambda_k, g_k; \alpha, d) = (d-2) g_k - (d-2)\omega_d g_k^2 + \mathcal{O}(g^3)$$



For $d=4$, cut-off profile

$$R^{(0)}(y) = y/(e^y - 1),$$

$$g_* \cong \pi/(13\pi^2/144 + 55/24 + \alpha)$$

$$\lambda_* \cong 3\zeta(3)/(13\pi^2/144 + 19/24)$$

Evidently, for appropriate α and $R^{(0)}(y)$ we can have qualitative agreement with our pure gravity results

$$g_* \cong 0.0533$$

$$\lambda_* \cong -0.000189$$



An Estimate of Λ : Planck Scale Cosmology -- (Bonanno&Reuter(J.Phys.Conf.Series140(2008)012008)) Transition between Planck regime and classical FRW regime at

$$t_{\text{tr}} \simeq 25 t_{\text{Pl}}$$

\Rightarrow

$$\begin{aligned}\rho_{\Lambda}(t_{\text{tr}}) &\equiv \frac{\Lambda(t_{\text{tr}})}{8\pi G_N(t_{\text{tr}})} \\ &= \frac{-M_{\text{Pl}}^4(k_{\text{tr}})}{64} \sum_j \frac{(-1)^F n_j}{\rho_j^2}\end{aligned}$$



For t_{eq} = time of radiation matter equality

we get (see Branchina&Zappala (G.R.Grav.42(2010)141))

$$\begin{aligned}\rho_{\Lambda}(t_0) &\cong \frac{-M_{Pl}^4(1 + c_{2,\text{eff}}k_{tr}^2/(360\pi M_{Pl}^2))^2}{64} \sum_j \frac{(-1)^F n_j}{\rho_j^2} \\ &\quad \times \frac{t_{tr}^2}{t_{\text{eq}}^2} \times \left(\frac{t_{\text{eq}}}{t_0^{2/3}}\right)^3 \\ &\cong \frac{-M_{Pl}^2(1.0362)^2(-9.197 \times 10^{-3})}{64} \frac{(25)^2}{t_0^2} \\ &\cong (2.400 \times 10^{-3} \text{eV})^4.\end{aligned}$$

Compare:

$$\rho_{\Lambda}(t_0)|_{\text{expt}} \cong (2.368 \times 10^{-3} \text{eV}(1 \pm 0.023))^4.$$

CONCLUSIONS

- * Herwiril.031 Just as General Herwig6.5, No Tweaking, Should Be Better in IR Due to Bloch-Nordsiek Effect -- Try it!
- * Real Progress on Λ in QFT
(Resummed Quantum Gravity Realization of Feynman's Approach)

