

# Trilepton Signatures in the General MSSM

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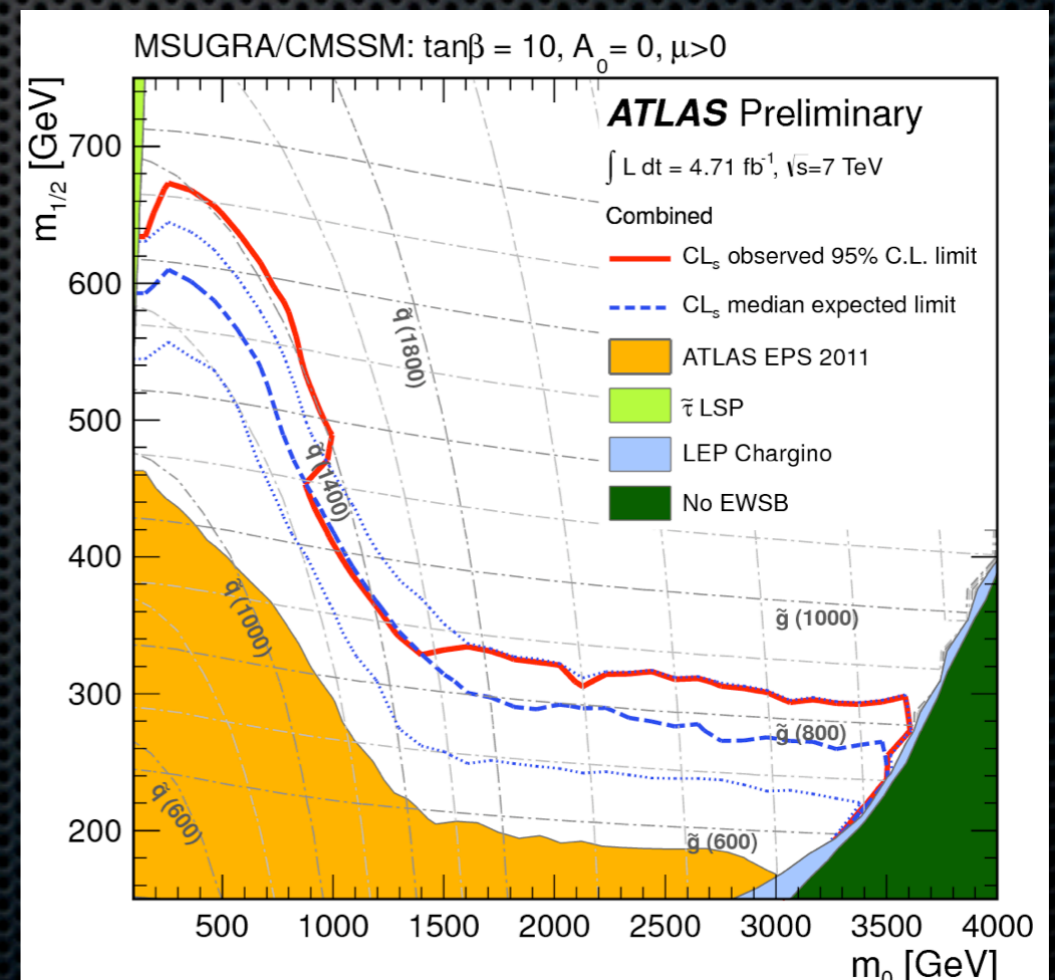
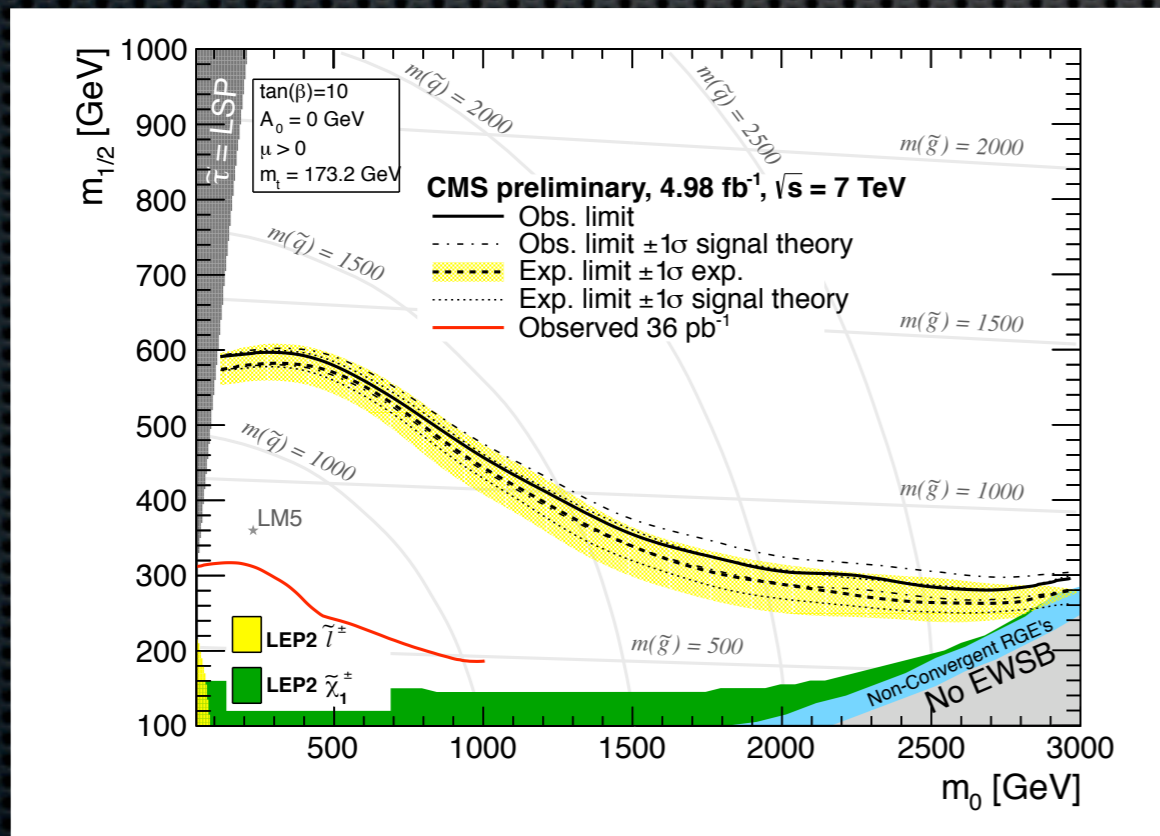
PHENO 2012

May 8, 2012

Work with Ed Berger and Zack Sullivan

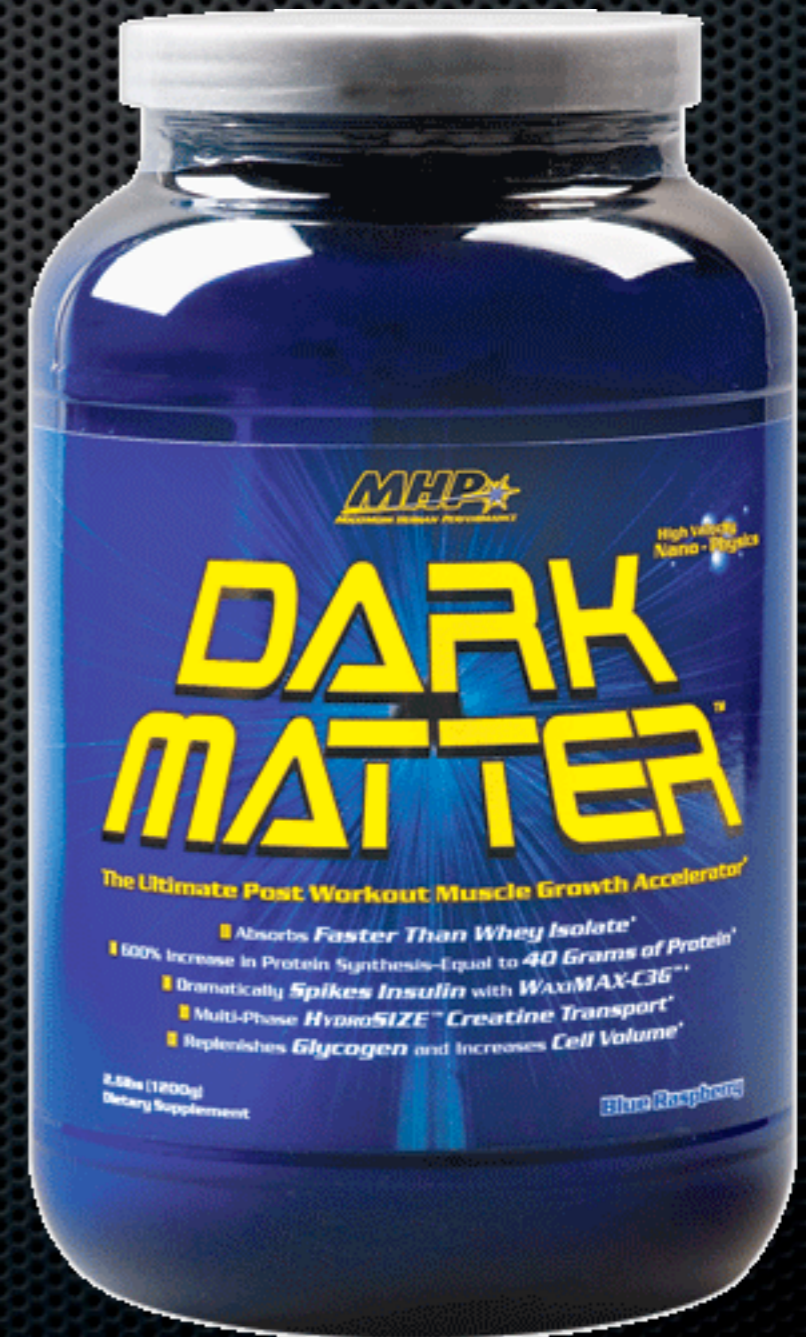
# SUSY Is Motivated!

- ✦ Otherwise you wouldn't be attending the "SUSY IV" session!
- ✦ But we haven't found it yet.



# Focus on Electroweak

- ✦ LHC has greatest sensitivity to colored sparticles
- ✦ Limits on uncolored sparticles are much weaker
- ✦ But important- an LSP (e.g.) neutralino could be the dark matter



# Scalars May Be Heavy

- ✦ Jet plus missing energy limits suggest heavy squarks
- ✦  $B_s \rightarrow \mu \mu$ ,  $B \rightarrow \tau \nu$ ,  $b \rightarrow s \gamma$ , etc. suggest Higgses other than the lightest one are heavy
- ✦ Sleptons and staus could still be light, but inspired (?) by the apparent heaviness of squarks, we assume they are heavy as well
  - ✦ Maybe split SUSY, but really we just assume these states are heavy enough to ignore in the analyses I'll describe.

# The pMSSM

- ✦ A general SUSY parameter space (not derived from a particular SUSY breaking scheme)
  - ✦ Some assumptions to keep number of parameters quasi-tractable
- ✦ Gaugino Masses:  $M_1, M_2, M_3$
- ✦ Trilinears:  $A_t, A_b, A_\tau$
- ✦  $\mu, \tan \beta$
- ✦  $m_A$
- ✦ Sfermion masses (10)

# Our pMSSM

- ✦ With the assumption of heavy scalars (and gluino) the only parameters we need to describe the physics we are interested in are
  - ✦  $M_1$ ,  $M_2$ ,  $\mu$ ,  $\tan \beta$
- ✦ We set the light Higgs mass to 125 GeV.
- ✦ Assume light Higgs is decoupled  $\Rightarrow \beta - \alpha = \pi/2$

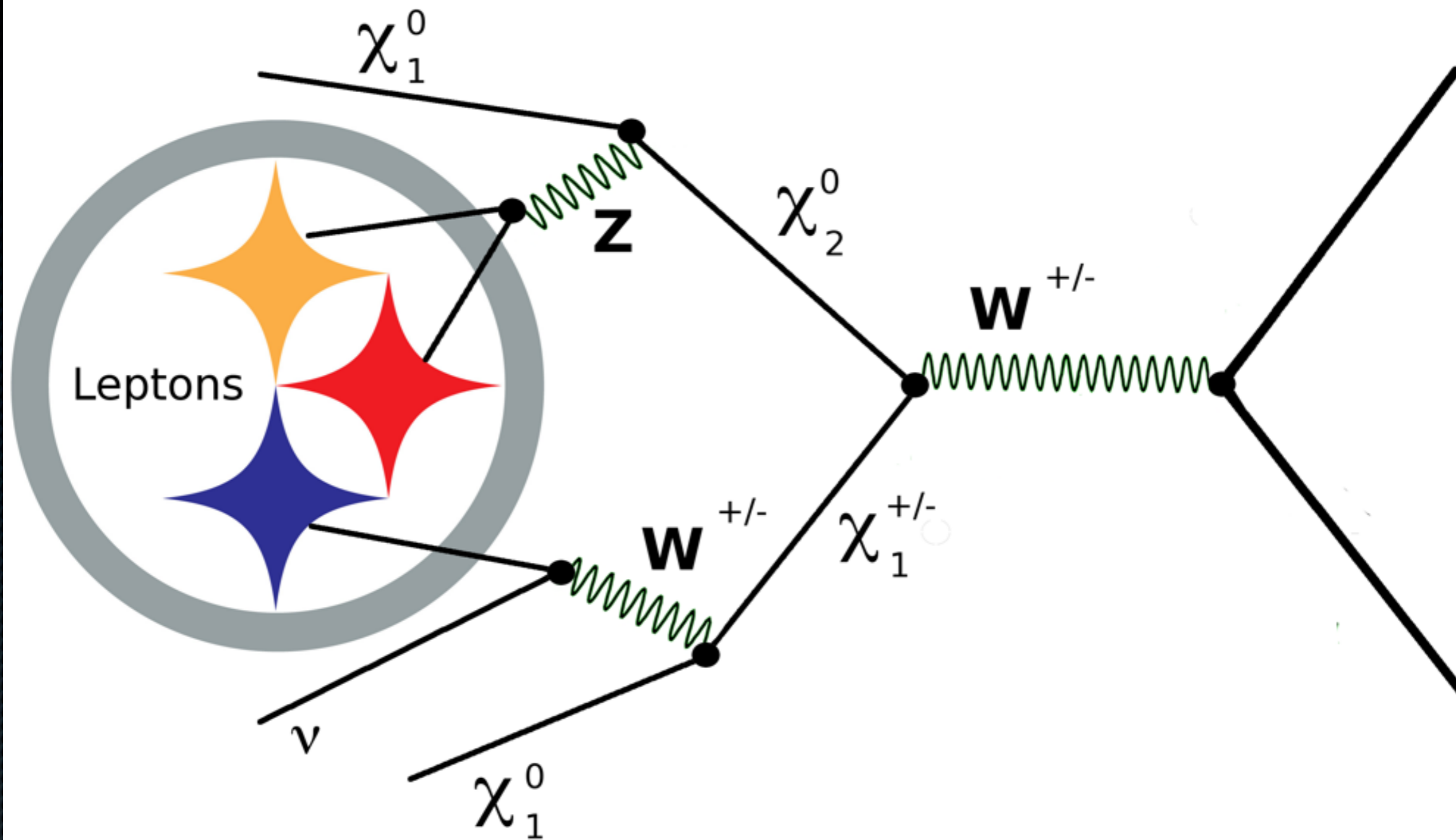
# Our pMSSM

- ✦ The SUSY spectrum is (generally)
  - ✦ A Bino-like neutralino with mass  $|M1|$
  - ✦ A Wino-like neutralino and a Wino-like chargino with masses  $|M2|$
  - ✦ Two Higgsino-like neutralinos and a Higgsino-like chargino with masses  $|\mu|$

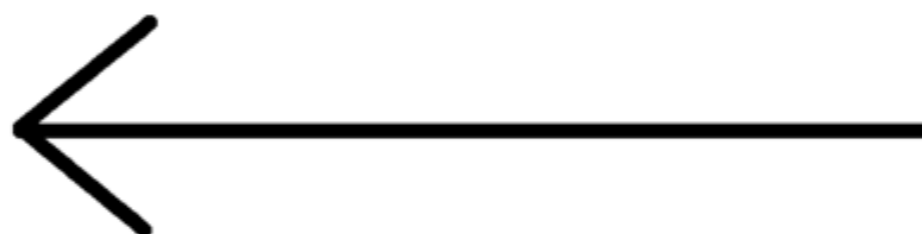
# Trileptons

- ✦ As a start on investigating the current and future ability of the LHC to discover SUSY in or rule out the scenario of light charginos plus neutralinos **and nothing else** we consider the (well-studied- apologies for not mentioning your work in this short talk) **trilepton channel (3l + MET)**





**Final State**

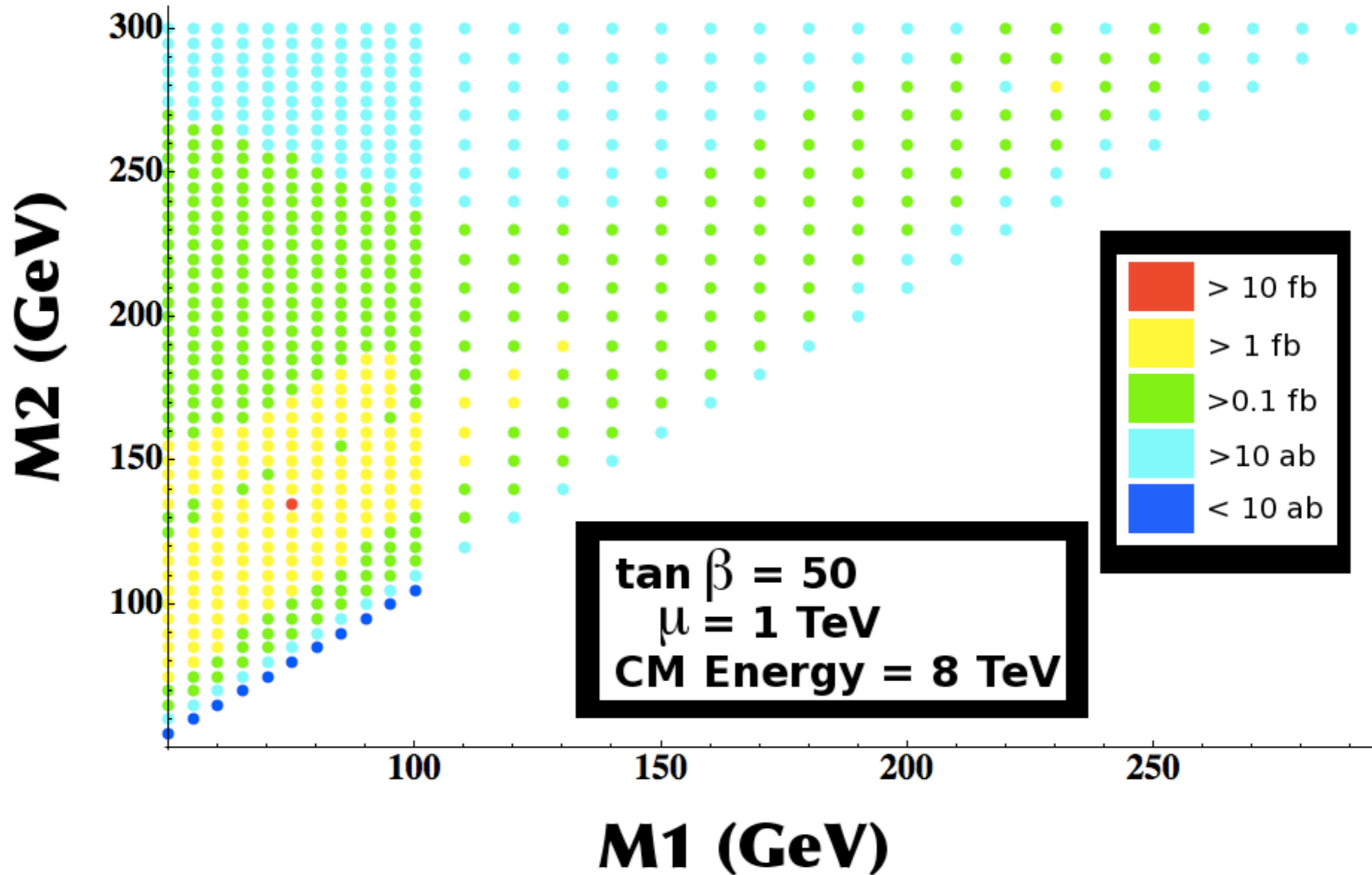


**Initial State**

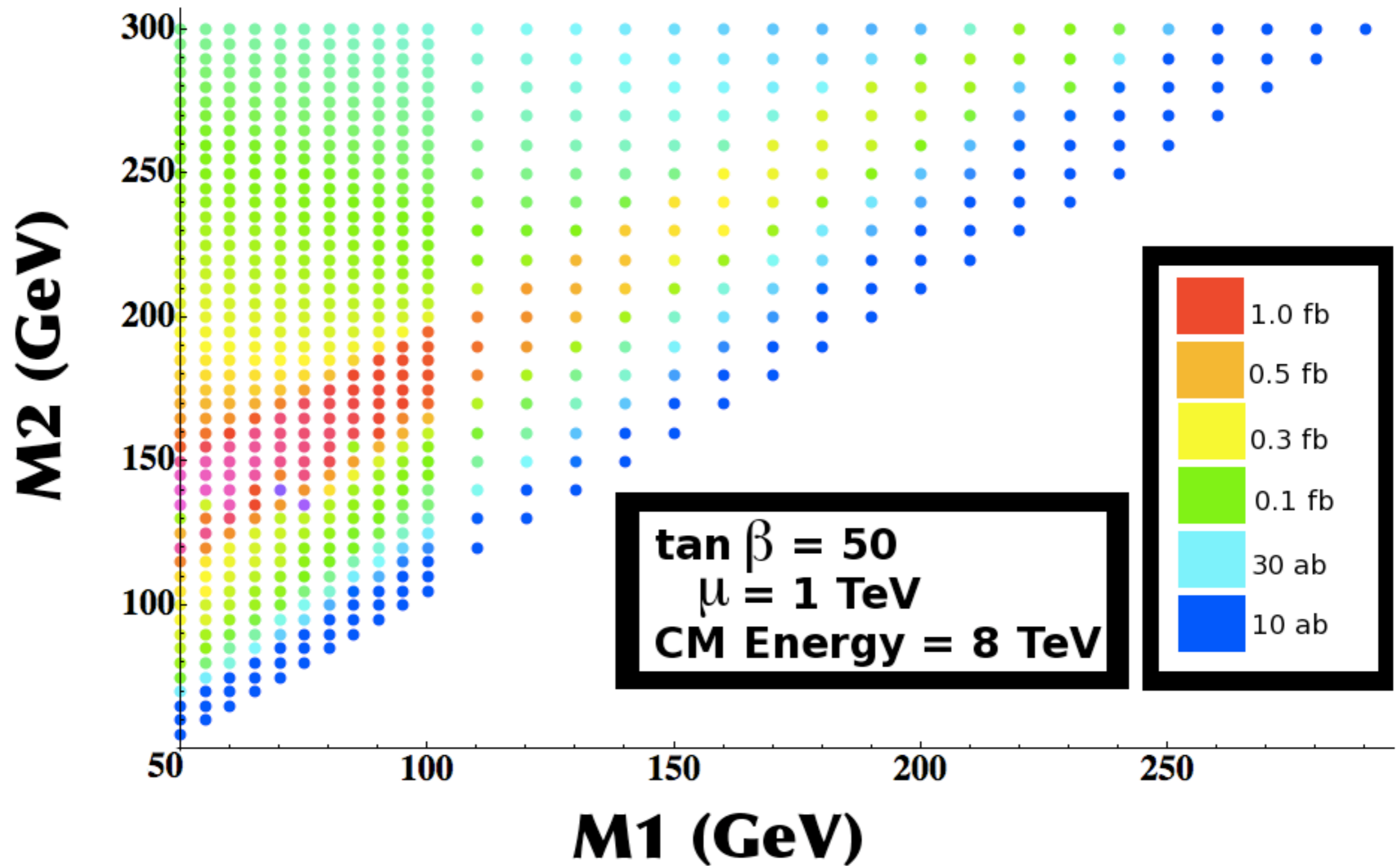
# Signal

- ✦ As an example of signal cross sections, I will show signal cross sections (at 8 TeV) for events passing the “Z-depleted” and “Z-enhanced” (7 TeV) ATLAS analyses
- ✦ Only contribution from positively charged chargino

# Trilepton Cross Section x BR for ATLAS Z Depleted Analysis



# Trilepton Cross Section x BR for ATLAS Z Enhanced Analysis



# Backgrounds

- Reducible backgrounds include:

$t\bar{t}$   
 $Wt$   
 $t b$   
 $b Z/\gamma^*$   
 $b\bar{b} Z/\gamma^*$   
 $Wb\bar{b}$   
 $b\bar{b}b\bar{b}$

where the lepton(s)  
come from the b(s)

TABLE I. Expected numbers of events from SM backgrounds (Bkg.) and observed numbers of events in data, for  $2.06 \text{ fb}^{-1}$ , in control regions VR1 and VR2, and in signal regions SR1 and SR2. Both statistical and systematic uncertainties are included.

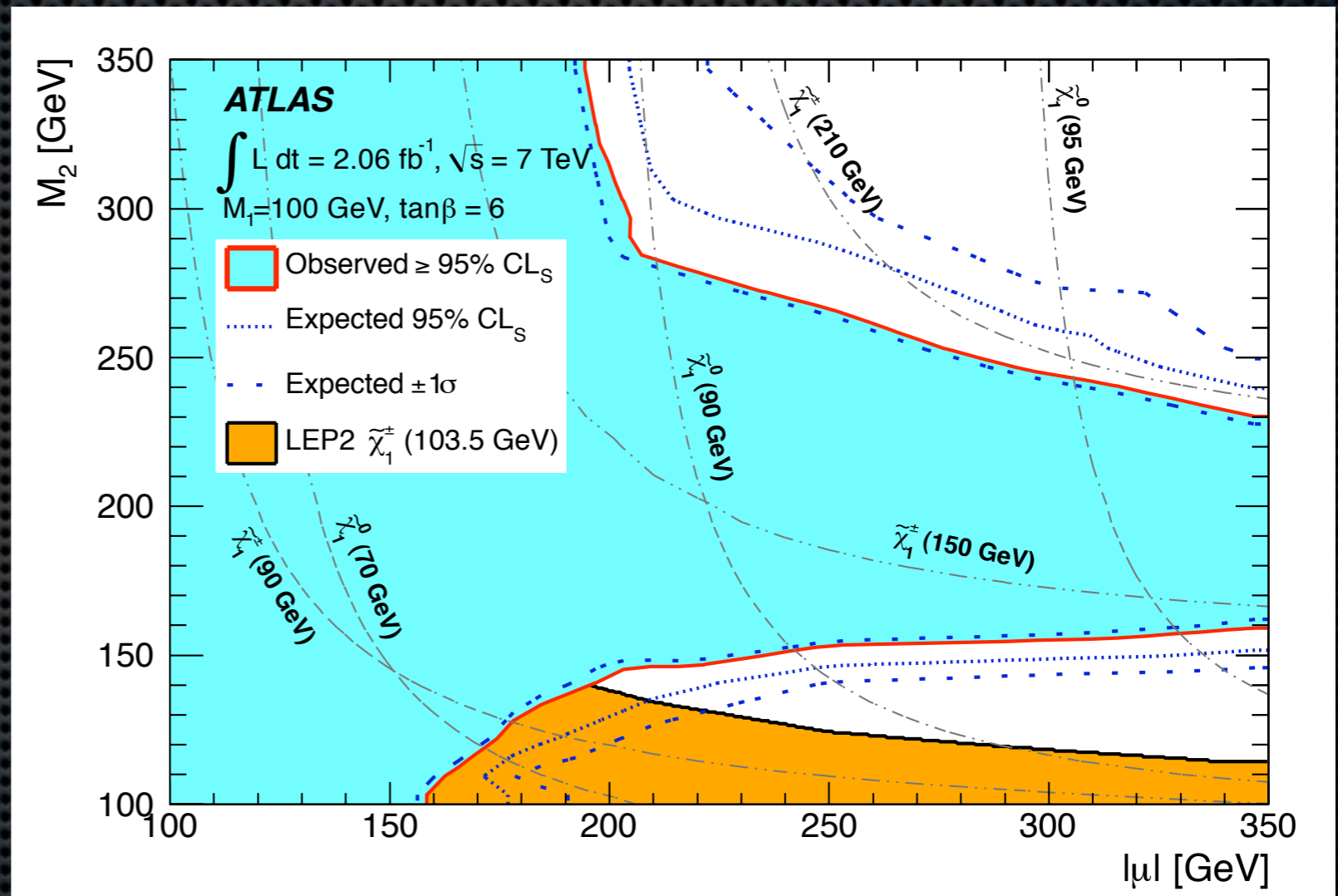
Selection	VR1	VR2	SR1	SR2
$t\bar{t}W^{(*)}/Z^{(*)}$	$1.4\pm 1.1$	$0.7\pm 0.6$	$0.4\pm 0.3$	$2.7\pm 2.1$
$ZZ^{(*)}$	$6.7\pm 1.5$	$0.03\pm 0.04$	$0.7\pm 0.2$	$3.4\pm 0.8$
$WZ^{(*)}$	$61\pm 11$	$0.4\pm 0.2$	$11\pm 2$	$58\pm 11$
Reducible Bkg.	$56\pm 35$	$14\pm 9$	$14\pm 4$	$7.5\pm 3.9$
Total Bkg.	$125\pm 37$	$15\pm 9$	$26\pm 5$	$72\pm 12$
Data	122	12	32	95

SR1: Z Depleted  
SR2: Z Enhanced

ATLAS: 1204.5368 (7 TeV)

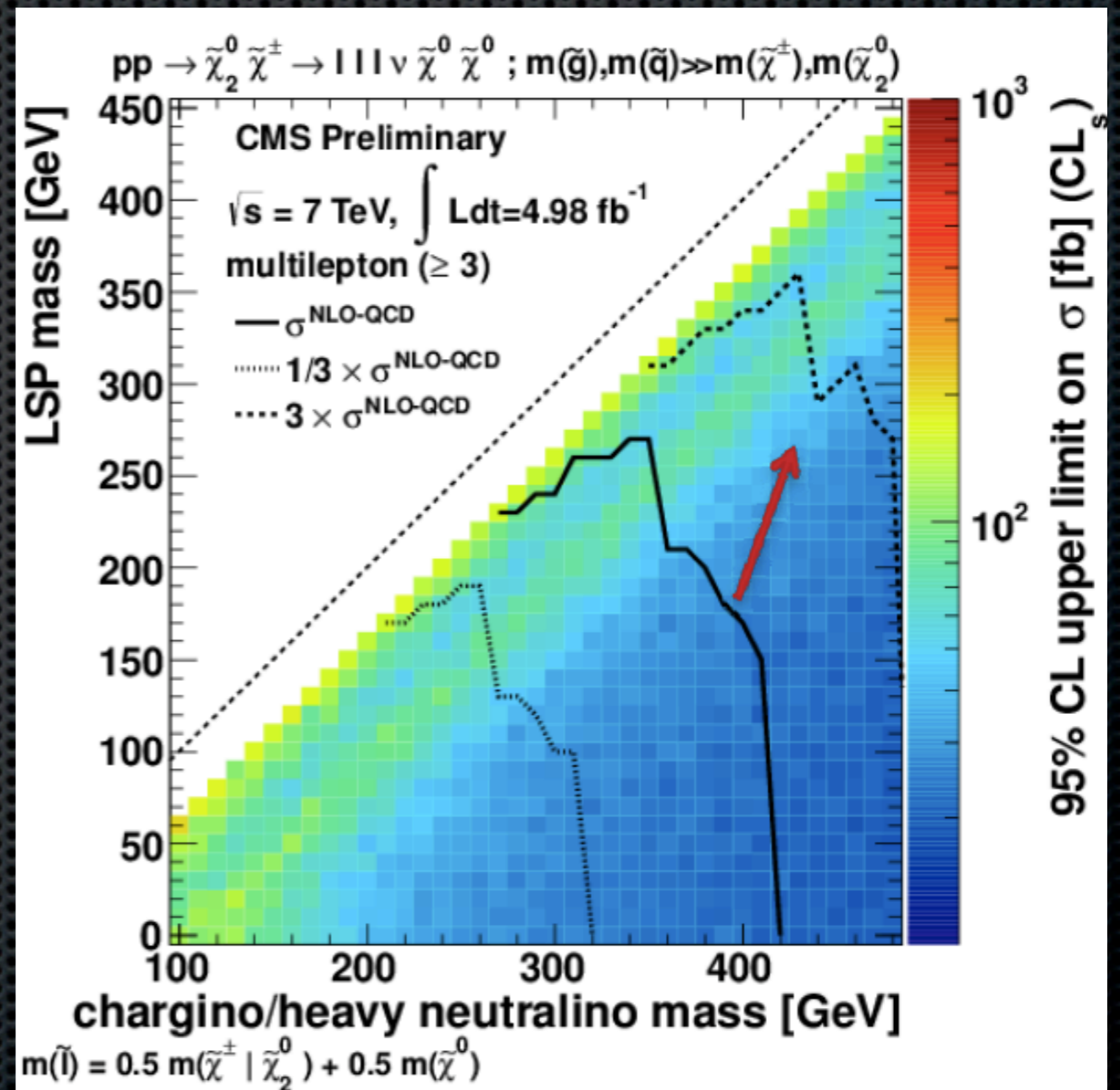
# What Experiments Have Done: pMSSM

- Both ATLAS and CMS have interpreted their trilepton results in terms of limits on a pMSSM parameter space
- They set the right-handed slepton mass to be intermediate between the LSP mass and neutralino...



# What Experiments Have Done: pMSSM

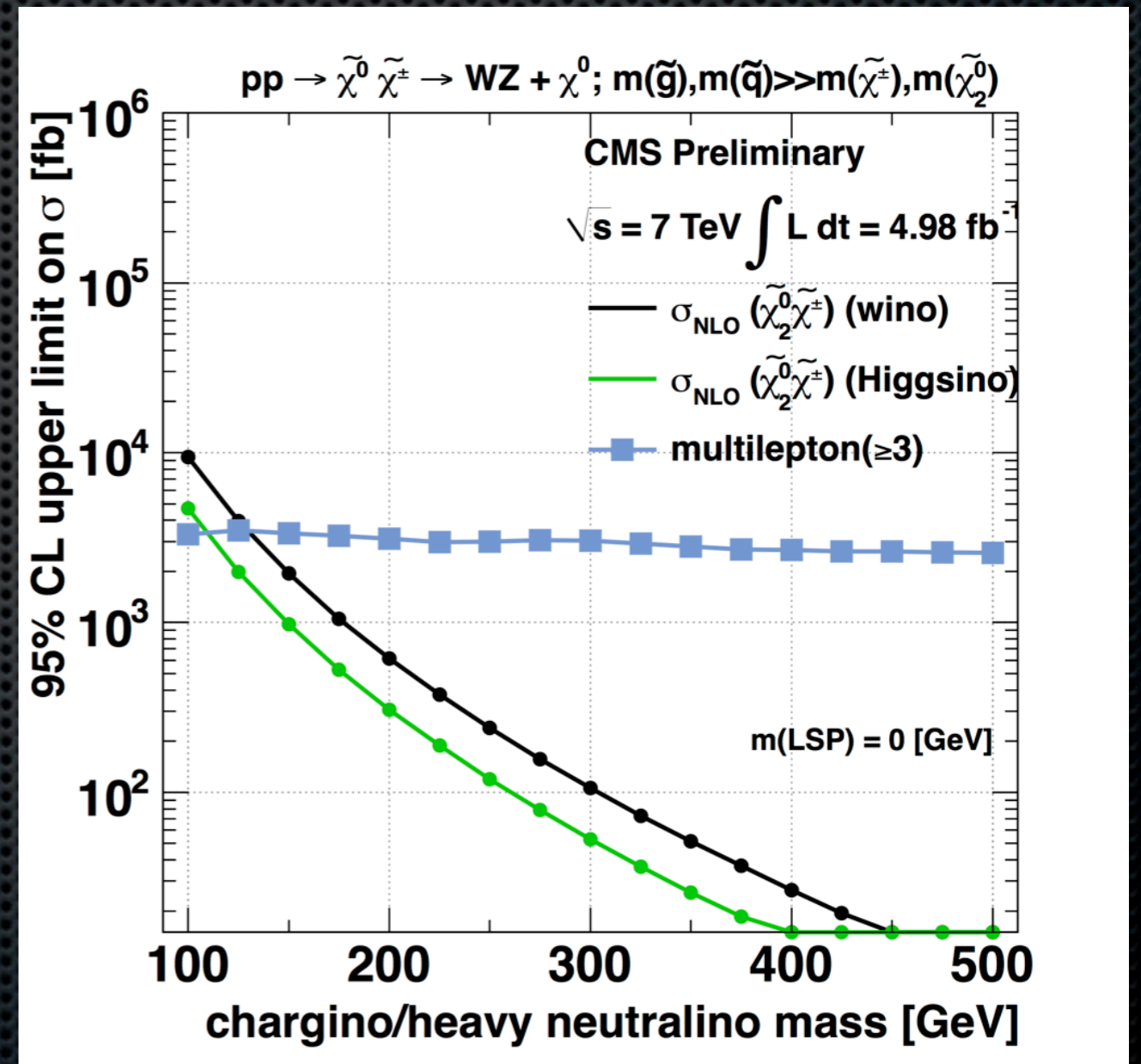
- ✦ Much higher cross section times branching ratio to trileptons due to decays to sleptons
- ✦ So this is a very different scenario—much more optimistic. Worthwhile to have an idea what can be done if sleptons are heavier.



From Sunil Somalwar's talk at U of C LHC Workshop

# What Experiments Have Done: Simplified Models

- When  $M1 < M2$  and  $|\mu| \gg |M1|, |M2|$ , our pMSSM space maps to the simplified model “TChiwz”.





# Why Do More?

- Not hard to recast these simplified models limits as limits in our pMSSM space when there is a **Bino-like LSP** and only **Winos** (or only **Higgsinos**) are also accessible.

# Why Do More?

- ✦ However other hierarchies of  $M_1$ ,  $M_2$ , and  $\mu$  are more complicated:
  - ✦ Additional light neutralinos and charginos with different masses may be relevant.
  - ✦ If the LSP is **Wino** ( $|M_2|$  small) or **Higgsino** ( $|\mu|$  small), we can replace the LSP in our diagram with a nearly degenerate chargino or neutralino which may decay softly enough to avoid vetoing the event.

# Why Do More?

- ✦ This particular **pMSSM** space is a good reflection of the **MSSM** without any optimistic assumptions
  - ✦ Additional e.g. sleptons make discovery in trileptons easier.
  - ✦ Light squarks would reduce the trilepton cross section, but would (hopefully!) show up in other analyses.
- ✦ A **pMSSM** analysis like this may be as close as one can reasonably get to obtaining limits on chargino and neutralino masses in the **MSSM** as a whole.
- ✦ Certainly a major step in that direction.

# Conclusions

- The importance of **SUSY** together with its current non-observation at the **LHC** motivate studying scenarios where charginos and neutralinos are light, but other sparticles are heavier.
- **Trilepton** events provide a probe of this scenario, though for much of the parameter space **other channels/ 14 TeV** are definitely needed (and even then...)
- It would be good to understand the consequences of observing (or not observing) excesses in this channel (and channels in general!) for as general of an **MSSM** parameter space as possible.
- Both experiments have made steps in this direction; more can and should be done to increase generality.