

Some possible interpretations from data of the CODALEMA experiment

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Study the of the radio-detection method in the energy range of 10^{16} - 10^{18} eV @Nançay

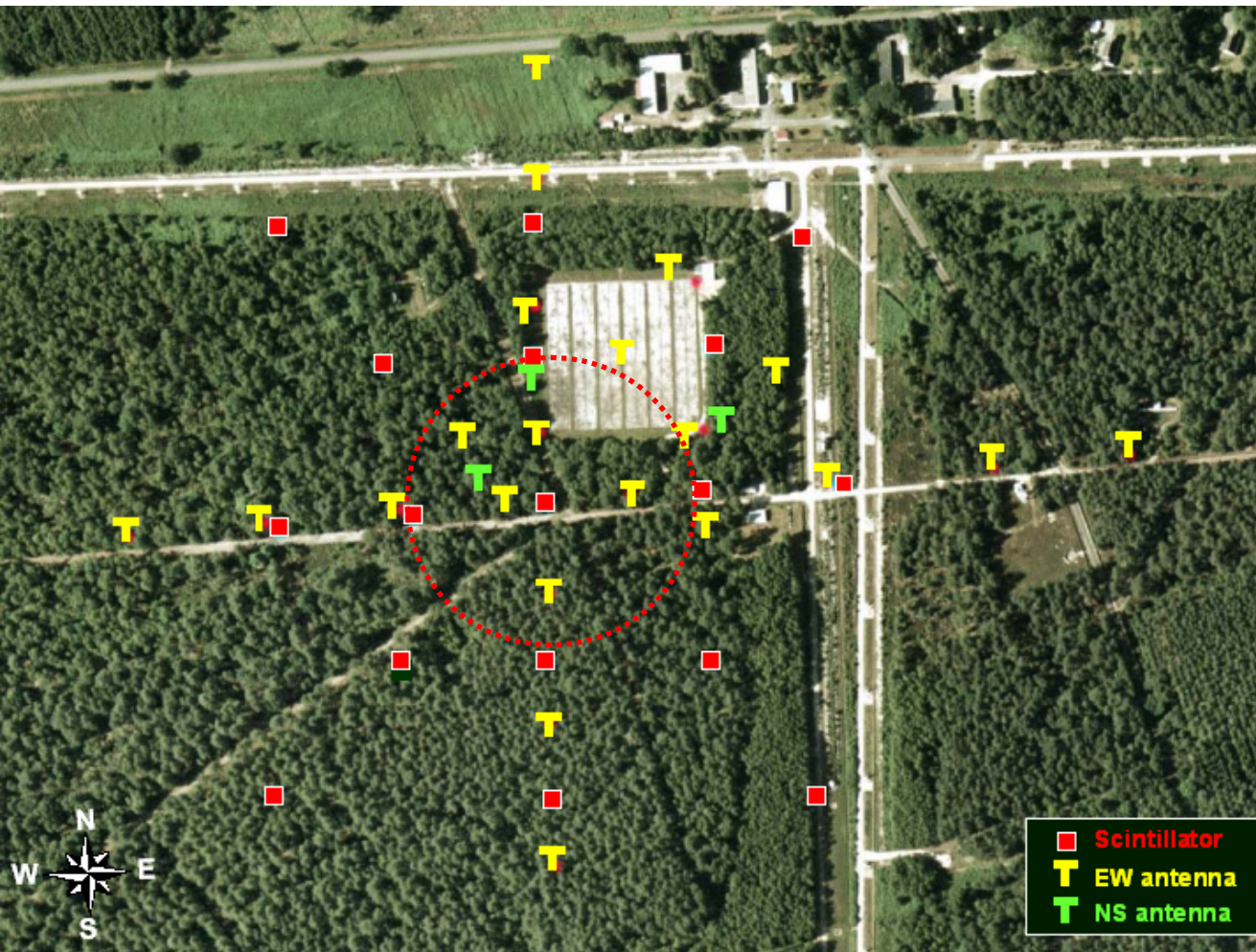
Purpose of these analyzes:

Try to go further on some previous studies (potential of some observables)

- **Study of an energy estimator** => internal correlations
(with the original setup)
- **Study of the wave fronts** (first step in retrieving information about the incoming wave) => source identification
(made necessary with the autonomous stations)

CODALEMA 2006-11 @ Nançay (0.25 km²)

Antenna array in slave mode of the scintillator array



Radio array

-24 dipole antennas

cross: 600 m x 500 m

21 ant. in E-W polarization

3 ant. in N-S polarization

Analyzes in 24-82 MHz

-Decametric Array

Cables + central acquisition

12 bits ADC

@ 1 Gsample/s, 2.5 μ s

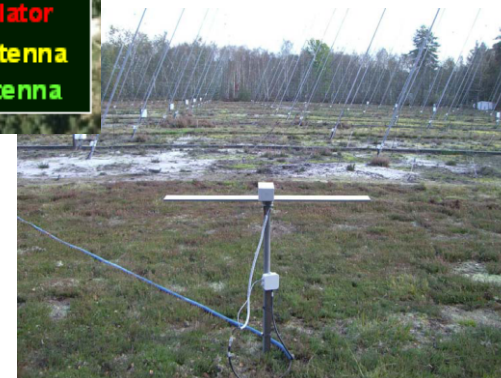
Particle array

17 scintillator stations : square 350 m x 350 m

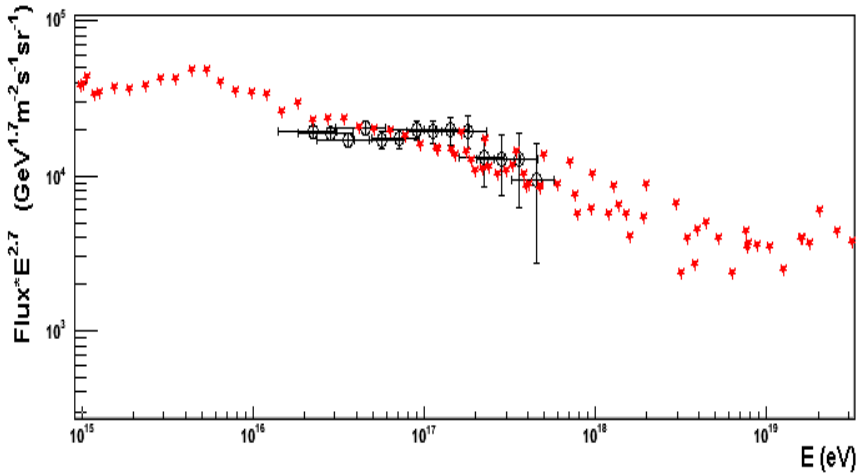
Trigger : the 5 central particle stations

Internal Showers : higher signal in central stations

Core Position + Direction + Energy (via CIC method)



Study of an energy estimator



Event overall features

Energy calibration of particle detector array

(for internal events)

=> NKG distribution

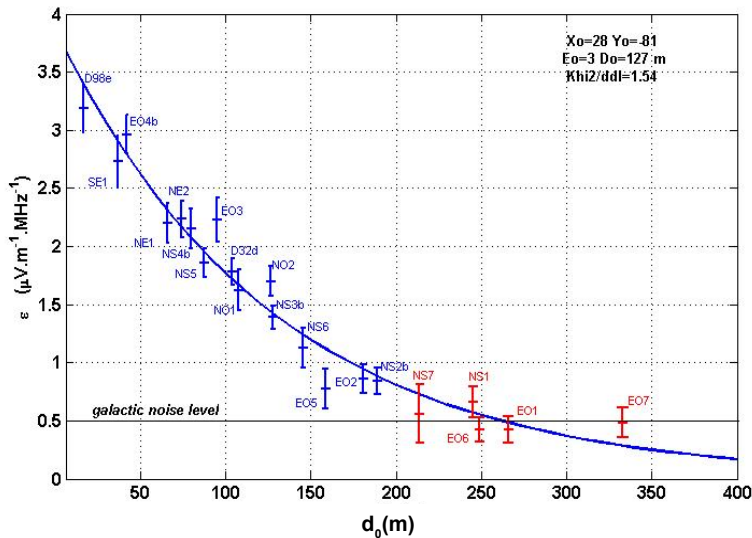
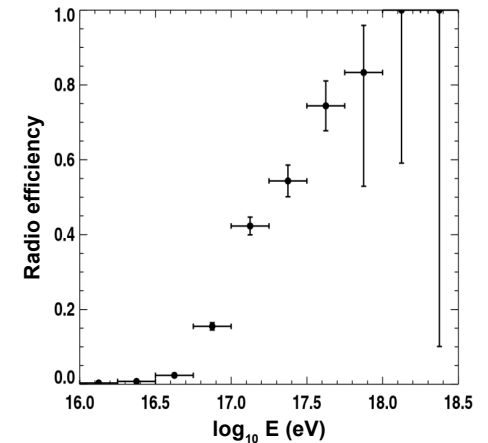
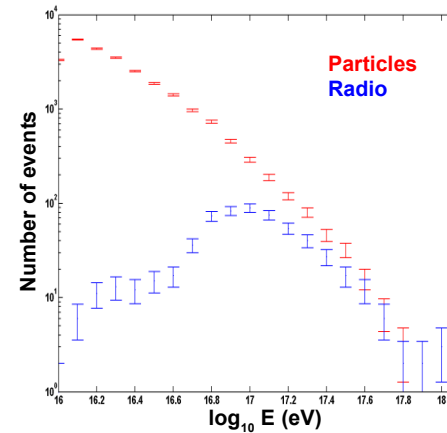
=> CIC method

=> $\sigma(E_p)/E_p = 30\%$

Coincidences in E-W polarization with internals

Criteria on: angle, time, core position
+ antennas > 4 + $\log_{10}(E_p) > 16.6$

= 315 events retained



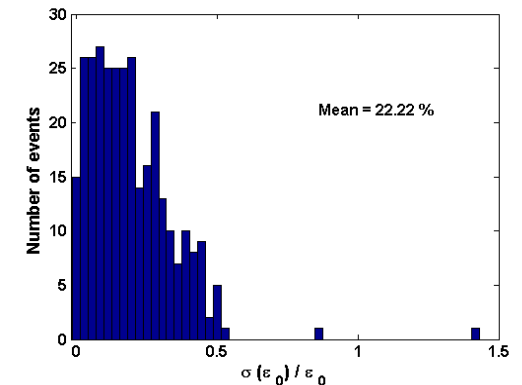
Radio Lateral Distribution Function

$$\varepsilon = \varepsilon_0 e^{-d/d_0}$$

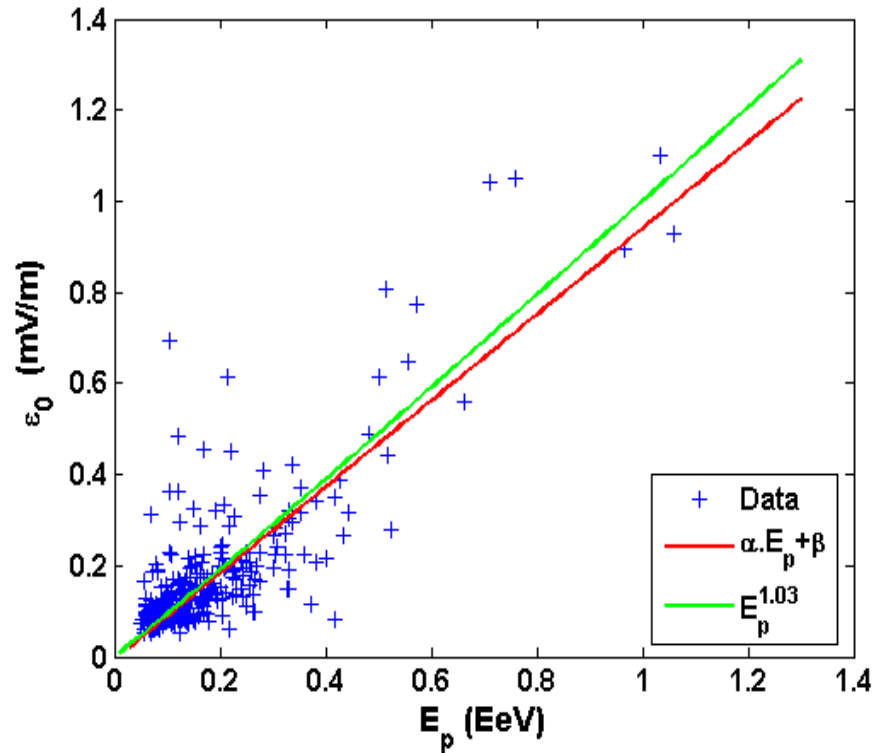
=> Extract the electric field at the shower core ε_0

Error estimation on ε_0 by

Monte-Carlo => $\sigma(\varepsilon_0)/\varepsilon_0 = 22\%$



Relationship between ε_0 and E_p



Compatible with linear dependence

Adjustment with $\varepsilon_0 = \alpha \cdot E_p + \beta$

=> Minimization of

$$\chi^2 = \sum_i [\varepsilon_0 - (\alpha \cdot E_p + \beta)]^2 / [\sigma^2(\varepsilon_0) + \alpha^2 \cdot \sigma^2(E_p)]$$

With $\sigma(\varepsilon_0)/\varepsilon_0 = 22\%$, $\sigma(E_p)/E_p = 30\%$

Inversion => Energy ratio: $E_0 = 1/\alpha \cdot \varepsilon_0 - \beta/\alpha$

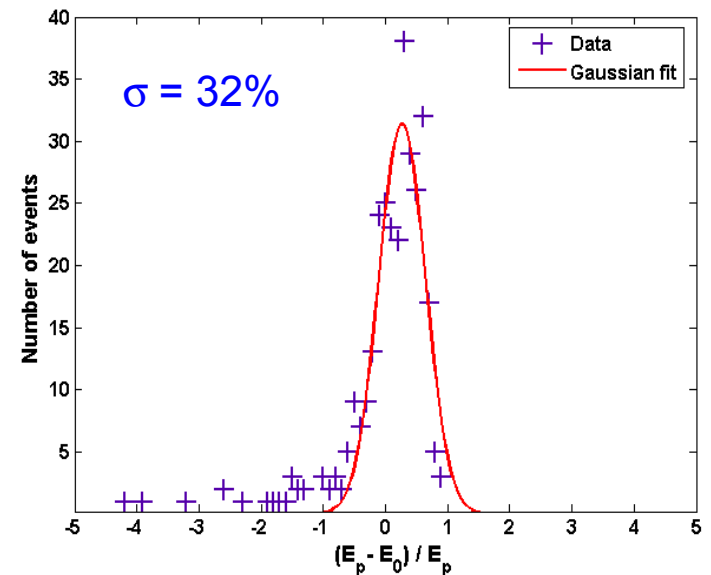
Characterization of the distribution (E_p, E_0)

=> Distribution of the relative differences $(E_p - E_0)/E_p$

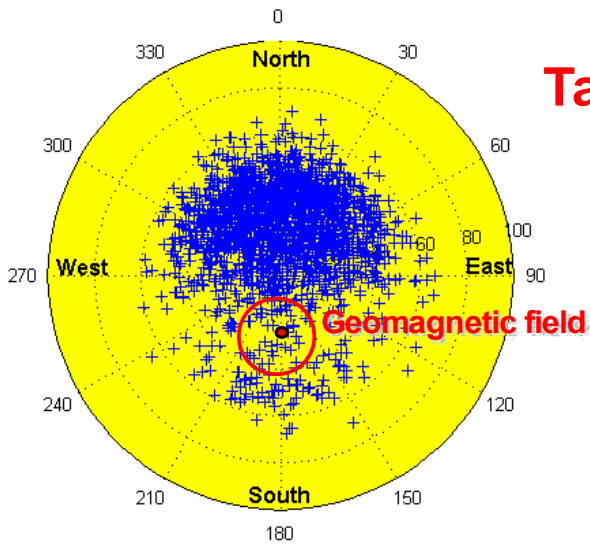
=> Analysis of $\sigma(E_p - E_0)/E_p$

(due to the tail, standard deviation used instead of Gaussian adjustment)

Error on $(\sigma(E_p - E_0)/E_p) = \sqrt{2 \cdot \sigma^2 / (N-1)}^{1/2} \sim 0.1\% - 1\%$



Taking into account the the geomagnetic effect

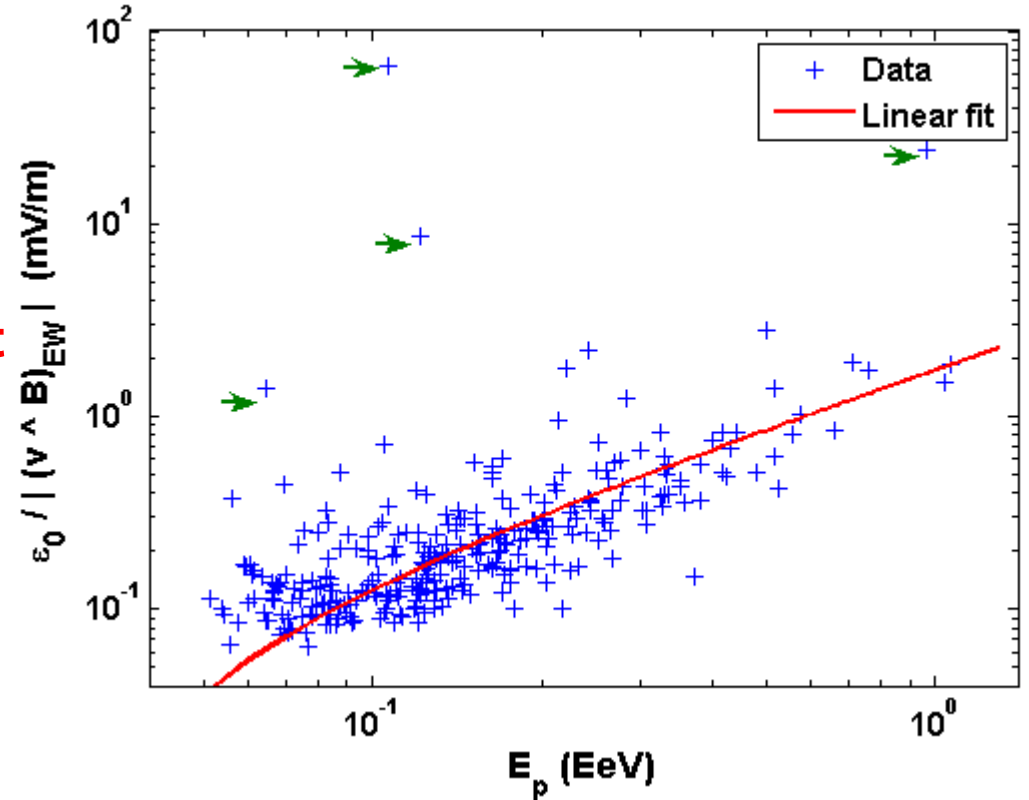


=> Correction by the geomagnetic effect

$$\varepsilon_0 \sim E \cdot |(\mathbf{v} \wedge \mathbf{B})_{EW}|$$

$$\varepsilon_0' \sim E \cdot |(\mathbf{v}' \wedge \mathbf{B})_{EW}|$$

$$\Rightarrow \varepsilon_0 \rightarrow \varepsilon_0 / |(\mathbf{v} \wedge \mathbf{B})_{EW}|$$



=> Overestimation of the energy for events near the geomagnetic field

- Relevant of ε_0 with the geomagnetic correction ?

=> Arguments of counting:

- Events are detected
- Uniform detection expected for higher energies (near the geomagnetic field)

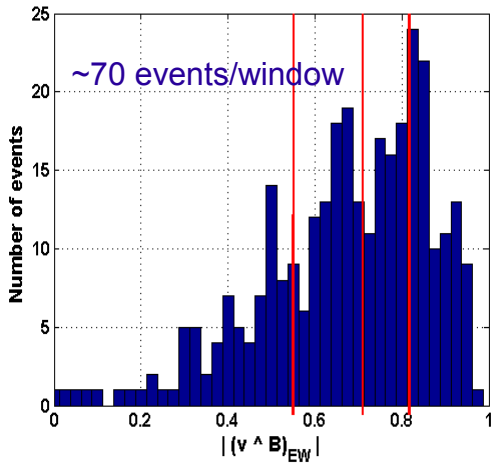
=> an additional contribution to the correction

Simplest assumption:
Additional contribution proportional to the total charge produced in the shower

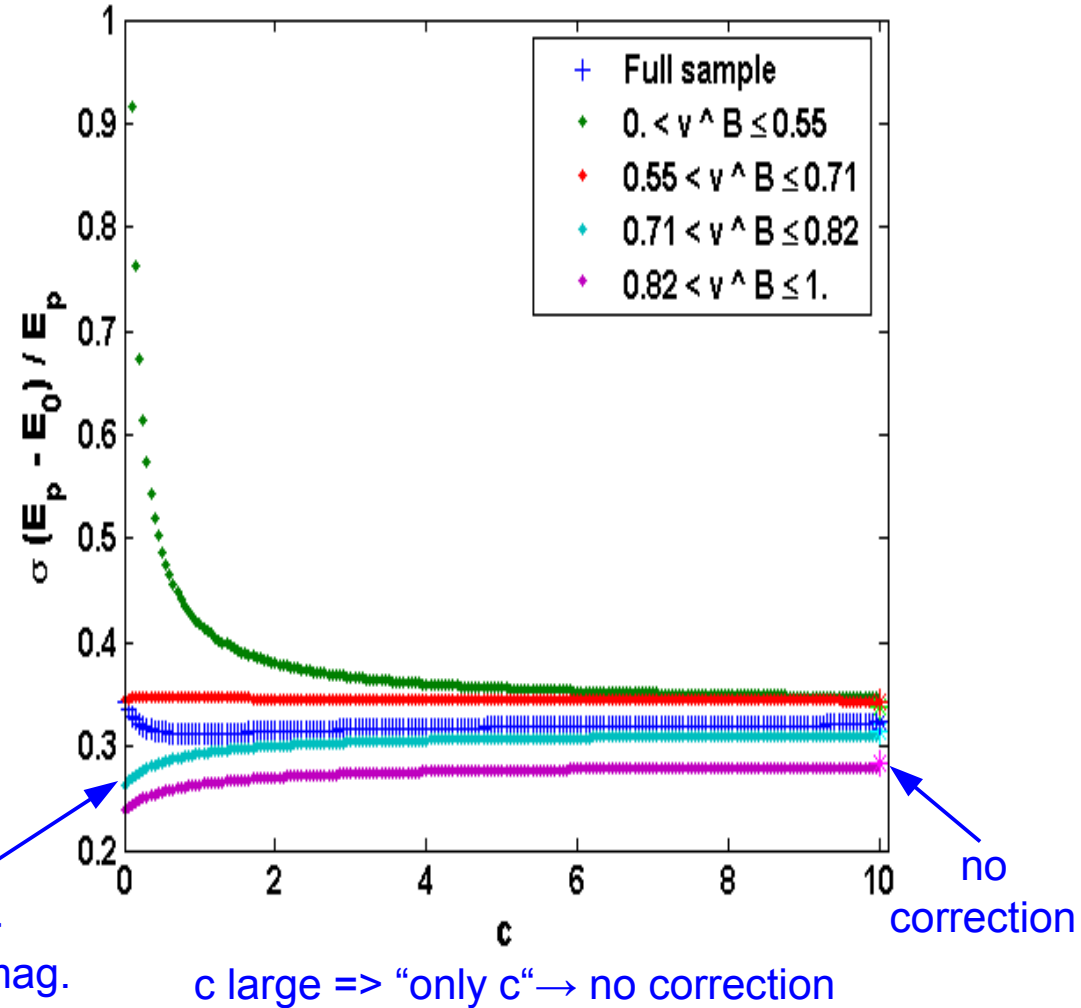
$$\varepsilon_0 \sim E \cdot |(\mathbf{v} \wedge \mathbf{B})_{EW}| + E \cdot c$$

$$\varepsilon_0 \rightarrow \varepsilon_0 / (|(\mathbf{v} \wedge \mathbf{B})_{EW}| + c)$$

Range of the parameters
 $0 < |(\mathbf{v} \wedge \mathbf{B})_{EW}| < 1$
 $0 < c < \dots$



Correction: $|(\mathbf{v} \wedge \mathbf{B})_{EW}| + c$



Full sample => min=0.31 for c=0.95
Sub-samples => exhibit angular dependencies

- best result obtained for $|v^B| > 0.82$ and $c=0$ => geomagnetic dominates
- for low $|v^B|$ => best result with large c => E.c dominates

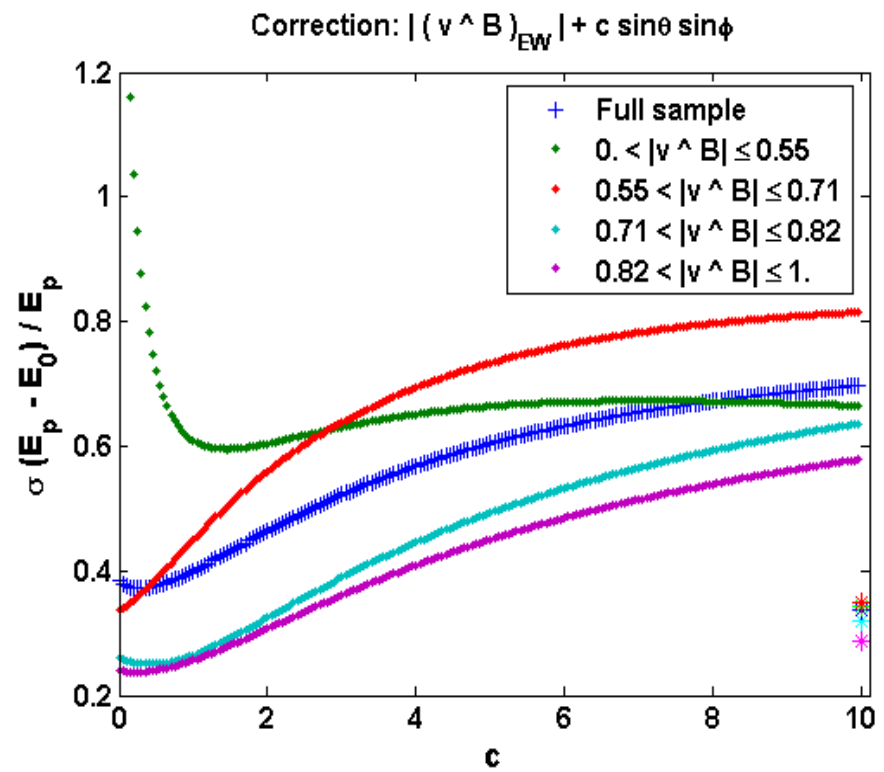
Association of the second term to an existing electric field

- Component \perp shower axis ?
- Component \parallel shower axis ?

$$\Rightarrow \varepsilon_0 \sim E \cdot |(\mathbf{v} \wedge \mathbf{B})_{EW}| + E \cdot c \cdot |\sin\theta \sin\phi|$$

(zenith θ , azimuth ϕ)

=> not favored



=> Possible interpretations (with the present statistic)

$$\varepsilon_0 \sim E \cdot |(\mathbf{v} \wedge \mathbf{B})_{EW}| + E \cdot c \Rightarrow \text{indicates a mix of effects ?}$$

Second term depends on the charge of the shower

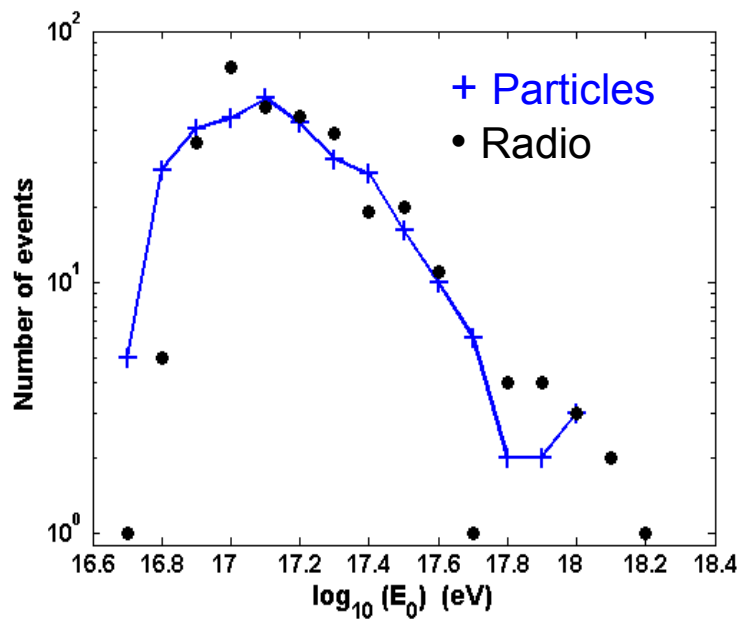
Greater weight of c when $|(\mathbf{v} \wedge \mathbf{B})_{EW}|$ decrease...

- Charge excess mechanism ?

$$\varepsilon_0 \sim E \cdot |(\mathbf{v} \wedge \mathbf{B})_{EW}| \cdot (1 + c / |(\mathbf{v} \wedge \mathbf{B})_{EW}|)$$

Analog to deflection with a magnetic dipole ? => charged particles deviate more when $|(\mathbf{v} \wedge \mathbf{B})_{EW}|$ increase ? => distances between the particles increase ? => decreases coherence ?

- Coherence effect ?



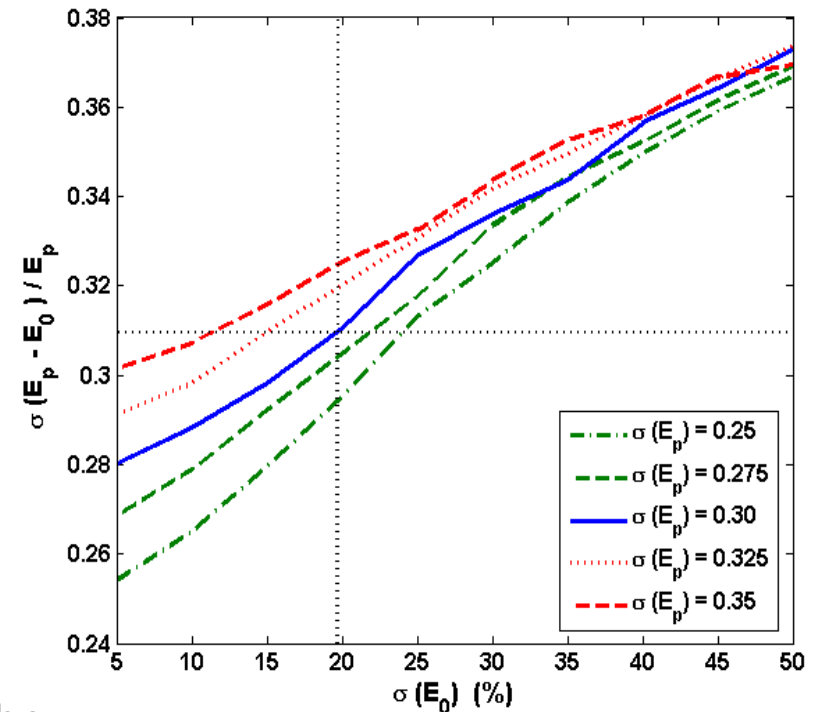
Result of the calibration
(without correction)

Attempt to estimate an energy resolution for radio

Monte Carlo => distribution in energy of the events in coincidence
=> random E
=> Random E_p in 3σ with $\sigma(E_p)/E_p$ fixed (25%, 30%,...)
=> Random E_0 in 3σ by varying $\sigma(E_0)/E_0$.
Construction of the distribution $(E_p - E_0)/E_p$

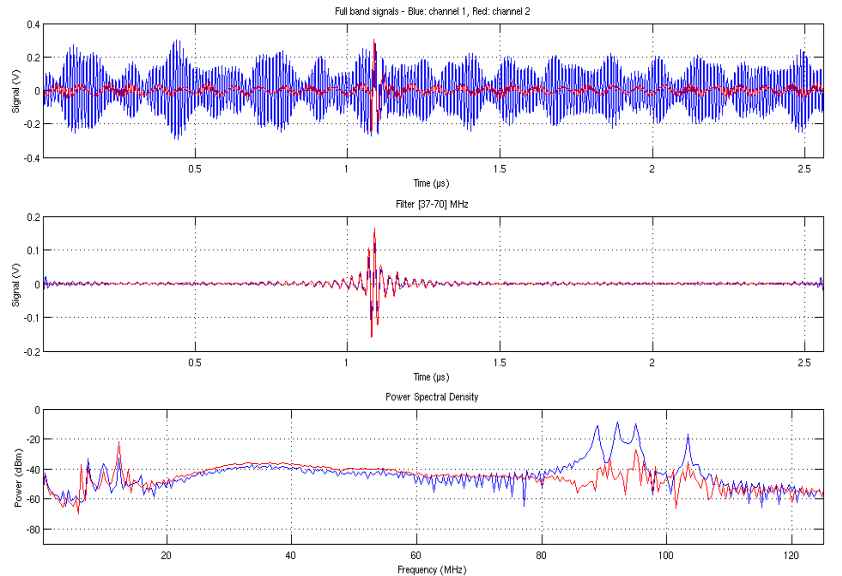
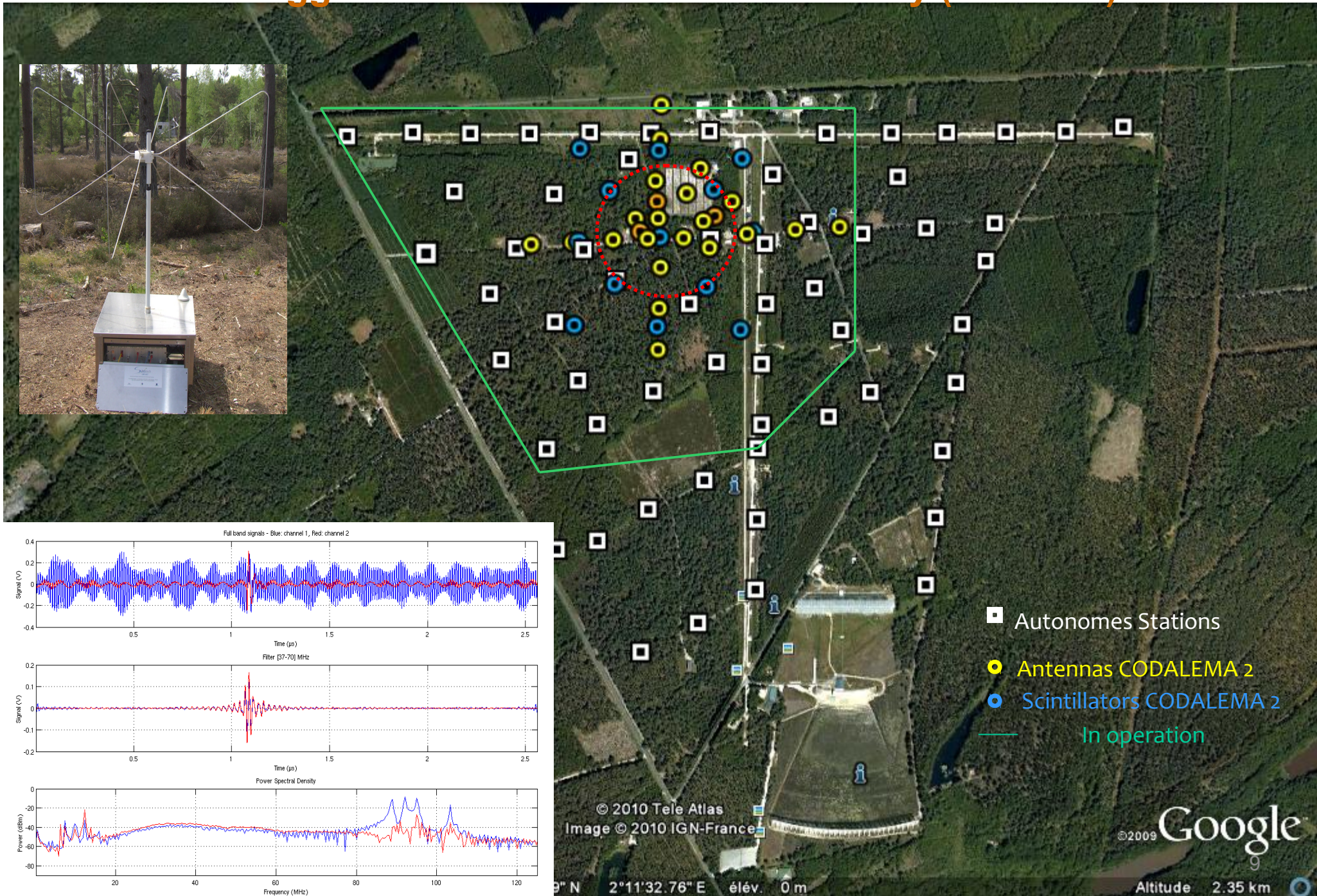
Without correction: $\sigma(E_0)/E_0 = 20\%$

=> an underestimated performance ?
=> More relevant RLDF will increase the correlation ?
=> Principle of the correction doesn't depend on form of the RLDF



CODALEMA 2011... @ Nançay

Self-triggered antennas + multi-scale array ($\Rightarrow 1 \text{ km}^2$)



- Autonomes Stations
- Antennas CODALEMA 2
- Scintillators CODALEMA 2
- In operation

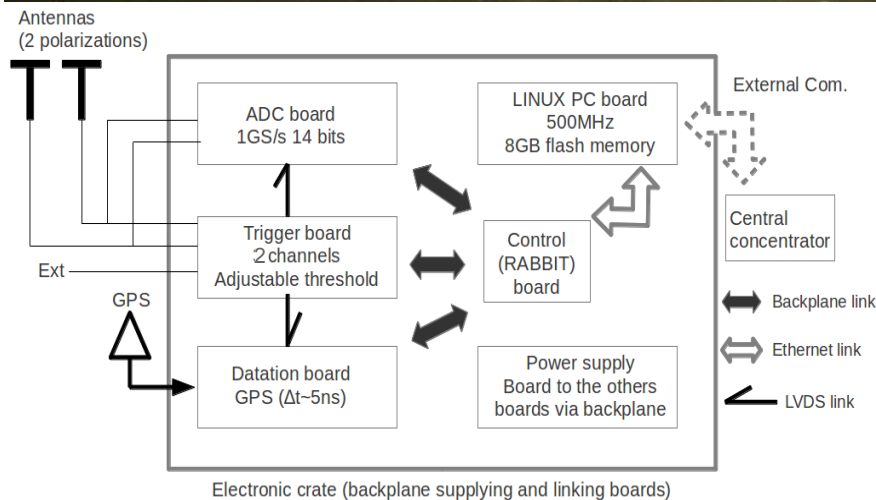
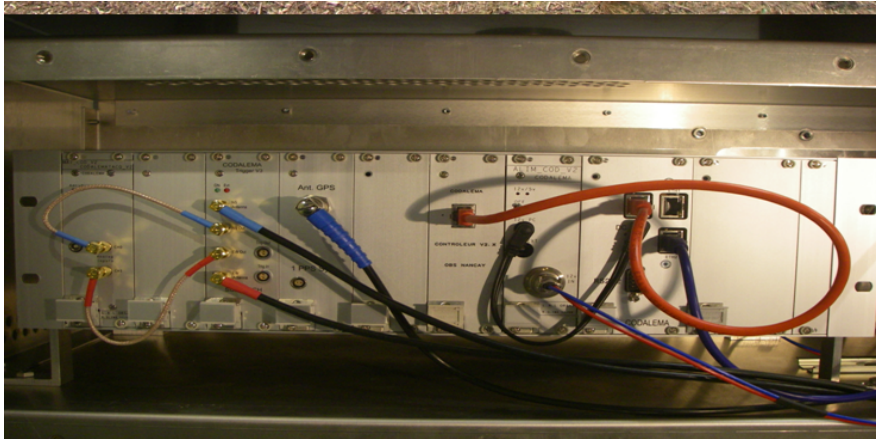
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Image © 2010 IGN-France

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9° N 2°11'32.76" E élév. 0 m

Altitude 2.35 km

The CODALEMA stand-alone station



- **1 ADC board** : 14 bits waveform, 1GS/s, $2.5 \mu s$, 1-250 MHz BW
- **1 dating board** : GPS timing resolution: $< 5 ns$
- **1 trigger board** : 1st level of trigger @ galactic threshold
- **1 PC board** : 2nd level of trigger in embedded PC + Data management + 8GB flash memory
- **Acq. rate: 25 evt/s**
- **Consumption: 20 W**

Implementation of the trigger

1st trigger level (hard)
based on analog filtering in the trigger board+ amplitude threshold
=>Waveform digitization in ADC

Send to CPU

Threshold update
+ Possible selections
Max of amplitude
Standard deviation
Duration
Recurrence

2nd trigger (soft)
based on waveform analysis through PC embedded
=> Data storage

Features of the detected events

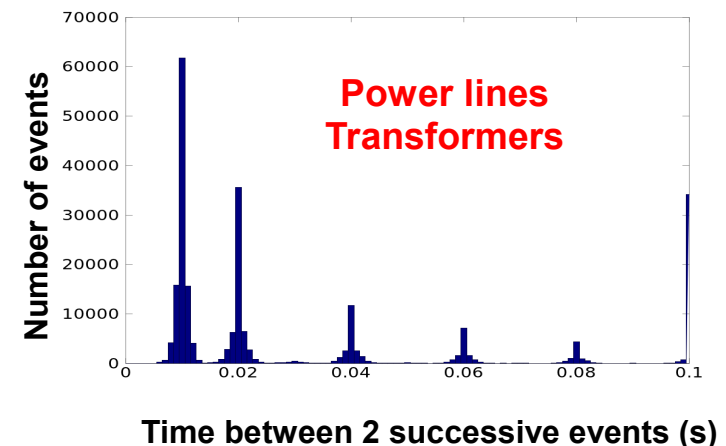
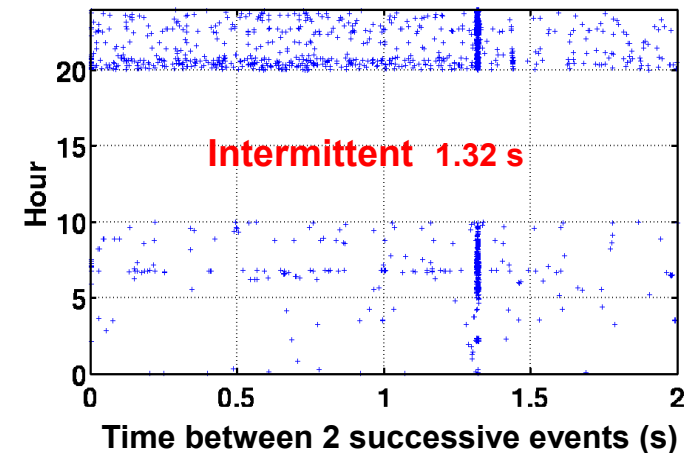
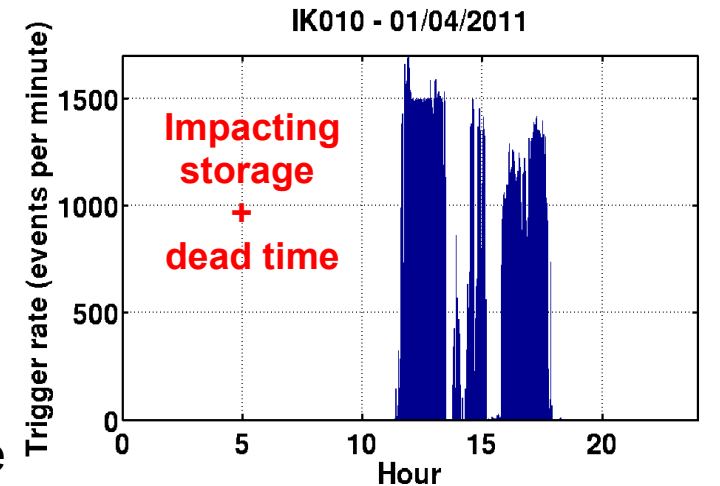
Dominated by anthropic emissions

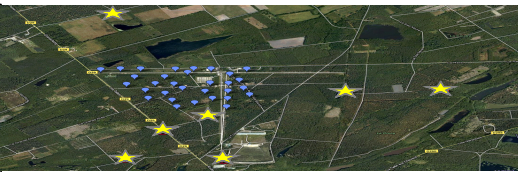
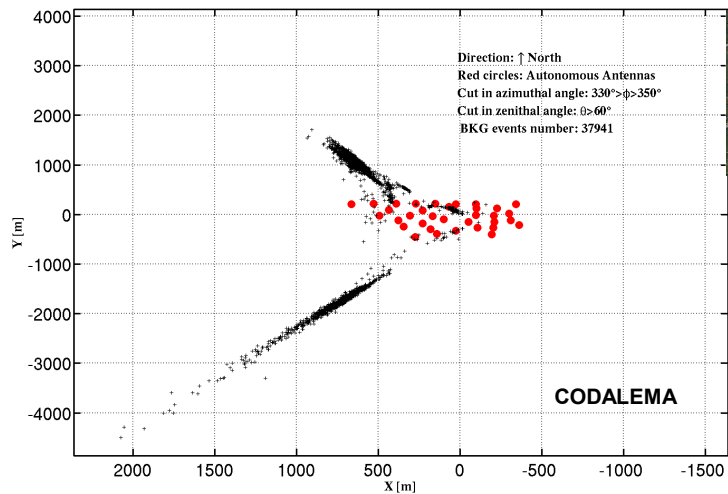
Intermittent emissions: eg. $T=1.32$ s, 1.4 s,... Filtering in the PC at level 2 (wave form analysis, recurrence transients...)
=> Reduce only storage, not the dead time

Daily recurring emissions (very annoying): @ 50 Hz & 100 Hz emissions affect most strongly the duty cycle
=> Power lines + transformers identified as major problem !

How to mitigate these effects ?

- Improving level 1 trigger card => identifying the recurrences before encoding ? => understand the mechanism of emission and its features
- Trying to solve the emission at the source level
=> modification of the power devices ?





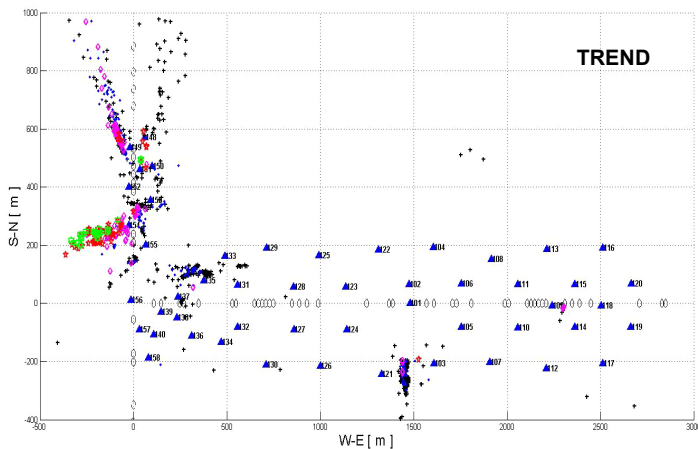
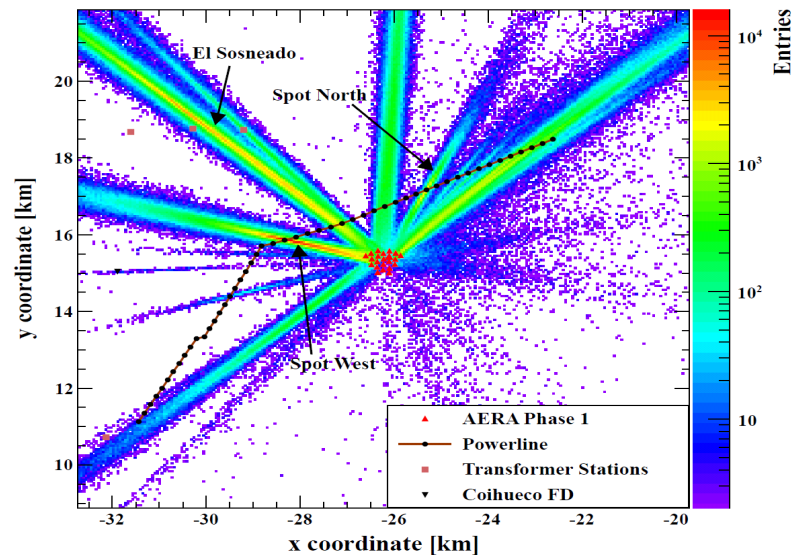
=> Observations of these RFI emissions by several experiences

=> Accurate source localizations expected using spherical reconstruction methods (Fixed source positions) (Number of events)



The problem will be analogous for shower events

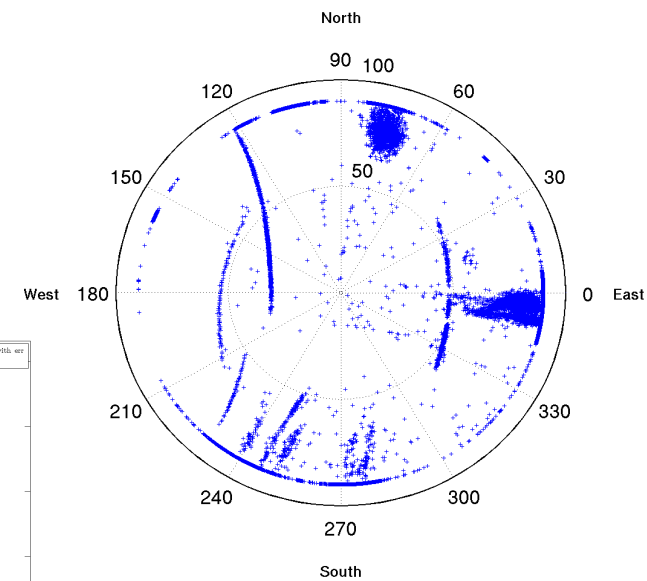
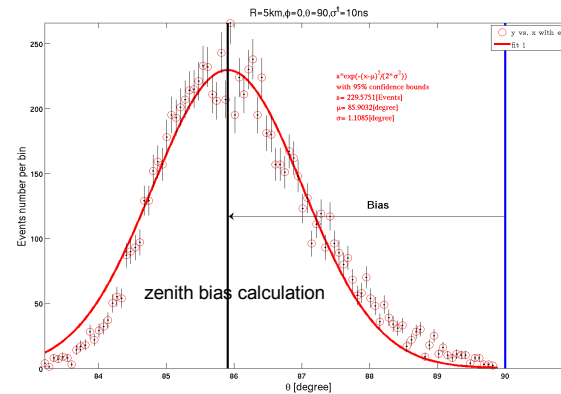
Except that in this case it is a single realization of the event parameters
=> Avoid a statistical estimate of the emitting center



Adjustments by planar wave fronts not relevant

=> Provides only arrival direction angles

- No localization of point-like sources
- Increasing bias with the zenith angles



Simulation of source reconstructions using spherical waves

=> Not only time resolutions are in question

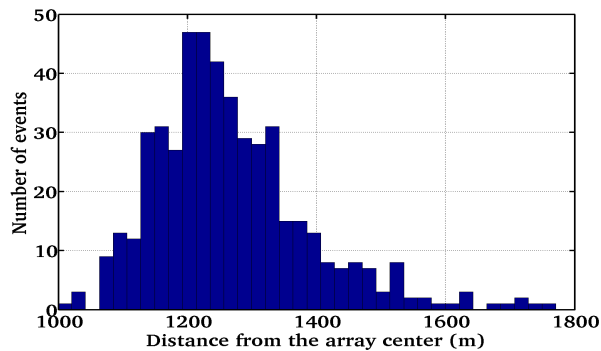
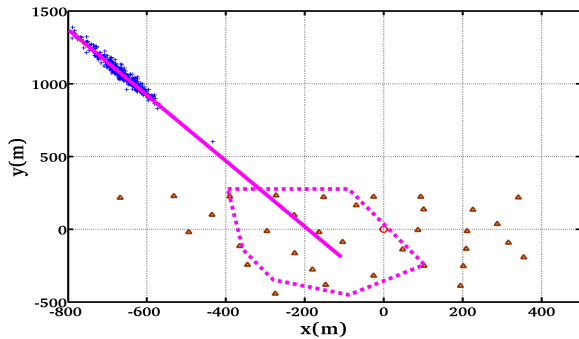
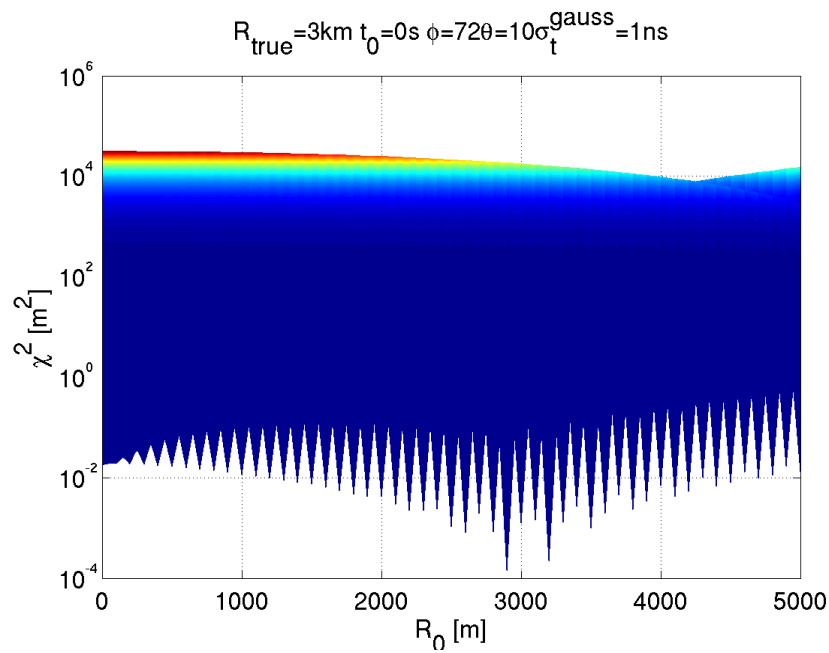
=> Also standard minimization methods are in question

+ results sensitive to

=> Algorithms (Simplex, Levenberg-Marquardt, Linear-search)

=> Initial values or the convergence

=> Detailed study of the minimization methods for the spherical reconstruction



=> Study of the objective function

(A. Rebai PhD thesis)

$$\chi^2 = \frac{1}{2} \sum_{i=1}^N \left[\|\vec{r}_s - \vec{r}_i\|_2^2 - (t_s^* - t_i^*)^2 \right]^2$$

Mathematical advanced analyzes

Study of Coercivity & Convexity (Hessian matrix)

χ^2 defined positive => at least one minimum exists

Local non-convexity

=> Existence of local minima

Not succeed to treat analytically the general case (3D array + source in sky) but some properties derived from particular cases

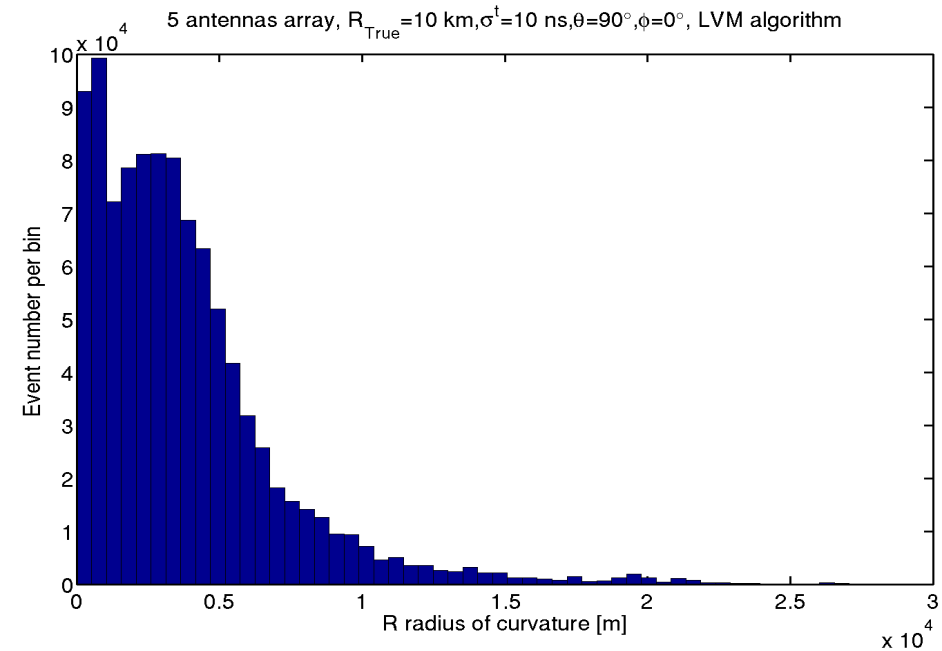
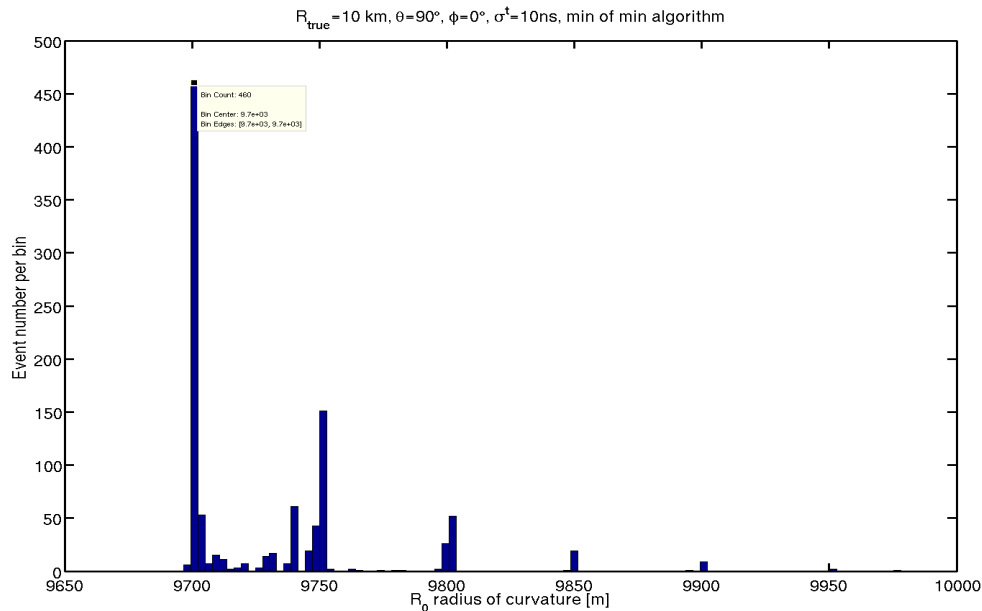
- Convex hull of the detector array
unique solution for the sources included
Degeneracies for external sources
- Critical points on a line (fired antenna barycenter-source location)

Conclusion => In this approach, spherical reconstruction seems an ill-posed problem !

Best method to bypass this problem not yet found...

=> Try to avoid trapping in some local minima => Search for the minimum of the objective function using a grid calculation

Simulated shower events



+ need bias estimation: in progress

- Try different approaches ?
- Use advanced statistical theories => additional information (signal amplitudes + functional of the RLDF, $\chi^2_{\text{Global}} = \chi^2_{\text{Time}} + \chi^2_{\text{Ampl}}$) ?

Work in progress...

Conclusion

Some earlier analyzes have been more deepened... leading to possible refinements of the interpretations

=> The electric field at shower core seems a relevant observable => provides an usable energy estimation after calibration by another detector

This observable « radio » of the energy could include a mix of several emission process (Charge excess ?, Coherence ?)

Triggering is challenging in autonomous mode
Recognition of emitting point sources is also challenging

Spherical reconstruction seems an ill-posed problem

=> Need to work further and maybe to interact with other disciplines (applied mathematics, earth sciences, GPS technics...)

Thank you for your attention