Measurement of a phase of a radio wave reflected from rock salt and ice irradiated by an electron beam for detection of ultra-high-energy neutrinos

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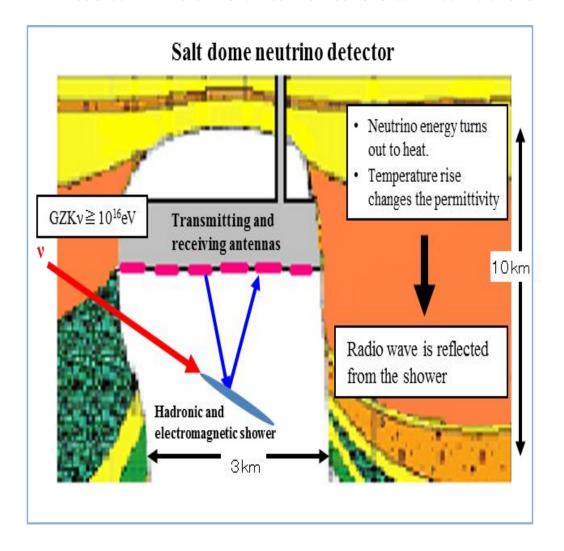
Ultra-high energy neutinos generated in GZK(Greisen-Zatsepin- $Log E(dN/dE) [km^{-2} yr^{-1} (2pi sr)^{-1}]$ Kuzmin) Process Yakutsk * Haverah Park Ultra-high-energy cosmic ray is generated in 104 △Fly's Eye Active Galactic Nuclei O AGASA 1000 AGASA Gamma-ray Burst etc. # HiRes GZK ν Protons (E>4*10¹⁹eV 10 Galactic 0.1 0.01 GZK threshol 0.001 $+\gamma(2.7K) \rightarrow \Delta^+ \rightarrow n$ 0.0001

- Detection of cosmic $v(10^{16} \sim 10^{21} \text{eV})$ using a rock salt formation or the Antarctic ice sheet
- ° ■The flux is extremely low ($\pm 1/\text{km}^2 \cdot \text{day}$), we need a gigantic mass of the detection medium of 50 Gt (± 10 events/year) a million times as large as Super Kamiokande.

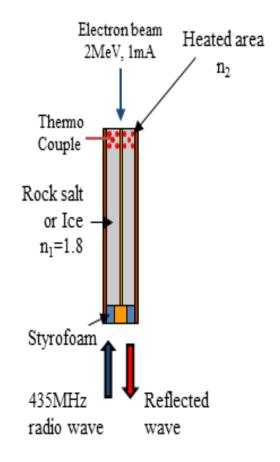
19

20

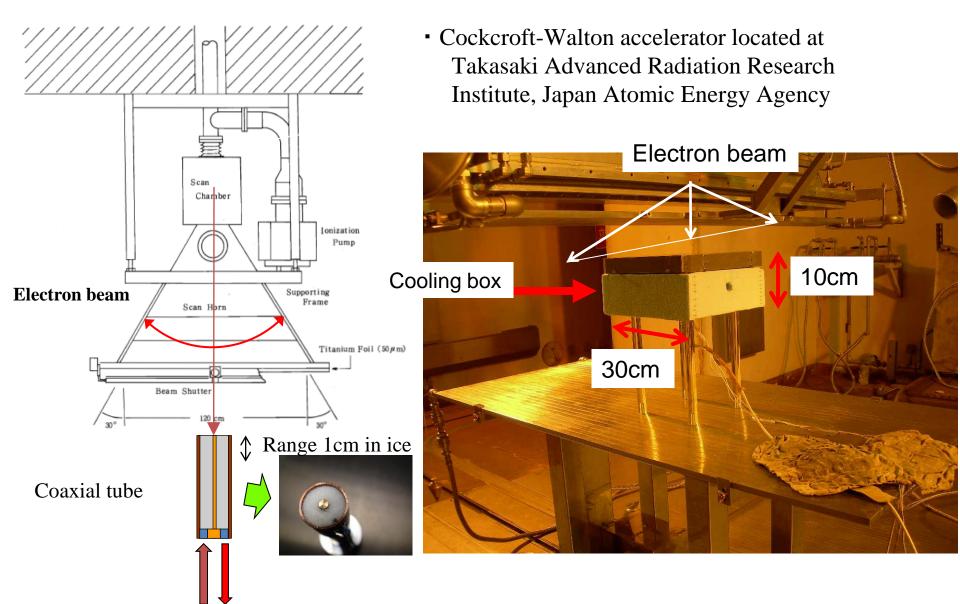
Radar method and a coaxial tube for ice and rock salt



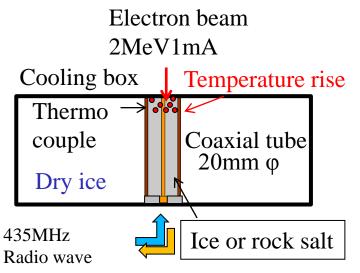
Coaxial tube with 2 cm diameter



Electron beam irradiation on ice or rock salt



Coaxial tube and radio-wave reflection

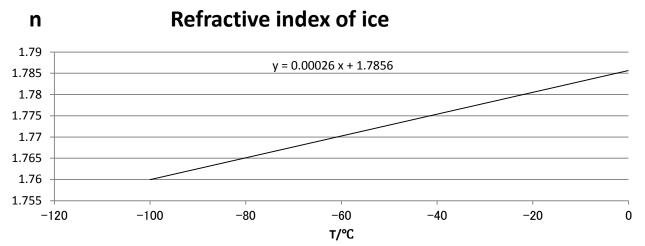


- •Coaxial tube was filled with ice or rock salt.
- •2MeV electron beam irradiate the open end of the coaxial tube.

$$(4 \text{ J/s} = 2 \times 10^{19} \text{ eV/s at 1mA})$$

- Temperature rise(ΔT)
 - \rightarrow Increase of refractive index(Δn)
 - → Radio wave reflection changes

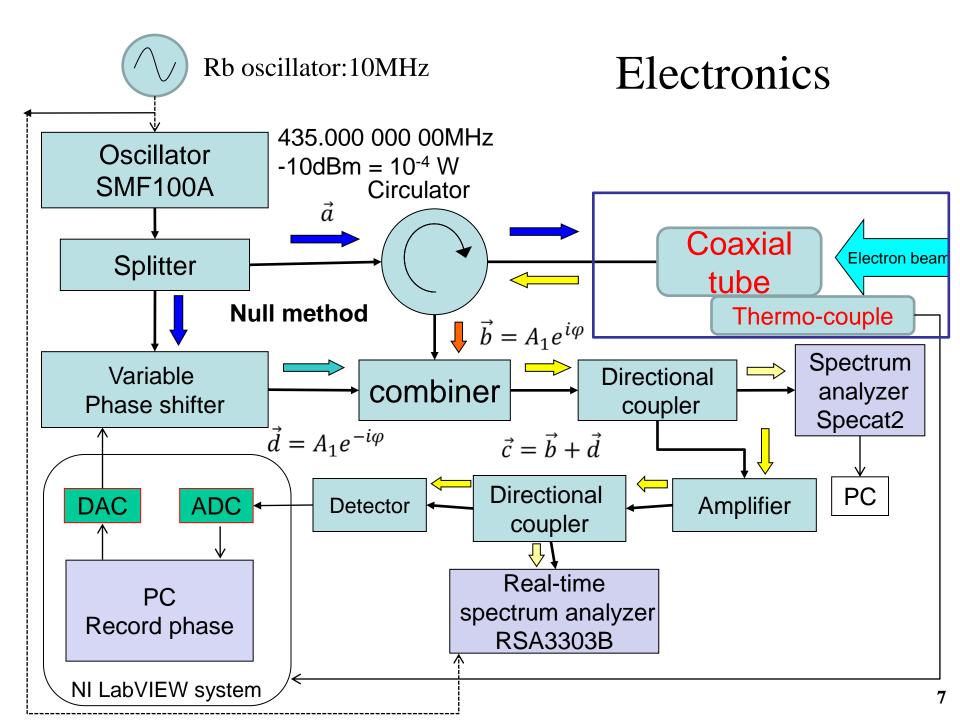
Power reflection rate with a change of refractive index



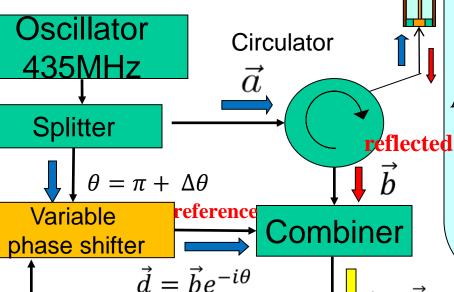
Matzler, C and Wegmuller, U., J. Appl. Phys. 80, 1623-1630(1987) Matsuoka, T, Fujita, S and Mae, S., J. Appl Phys. 80,5884-5890(1996)

- $\Delta n \propto \Delta T$ for ice and rock salt
- Power reflection rate (Γ) is proportional to ΔT^2 Fresnel equation

$$\Gamma = \frac{(n_2 - n_1)^2}{(n_2 + n_1)^2} \propto \Delta T^2$$
 n_1 : Before irradiation n_2 : Irradiation



Amplitude and Phase



Due to open end of coaxial tube

LabVIEW

system

Automatic tuning

Detector

$$|\vec{a}| / |\vec{b}| = 1$$

 $\int \vec{c} = \vec{b} + \vec{d}$

Splitter

RSA3303B

Amplitude measurement

Before irradiation: automatic tuning of θ is ON

$$\vec{c} = \vec{b}e^{-i\theta} + \vec{b}$$

reference reflected wave

$$\theta = \pi \qquad |\vec{c}|^2 = 0$$

At irradiation: automatic tuning OFF

$$\vec{c} = \vec{b}e^{-i\pi} + \vec{b}e^{-i\Delta\varphi}$$

reference reflected wave

$$\vec{c} = \vec{b}(e^{-i\Delta\varphi} - 1)$$

$$|\vec{c}|^2 = 2 |\vec{b}|^2 \{1 - \cos(\Delta \varphi)\}$$

Phase measurement

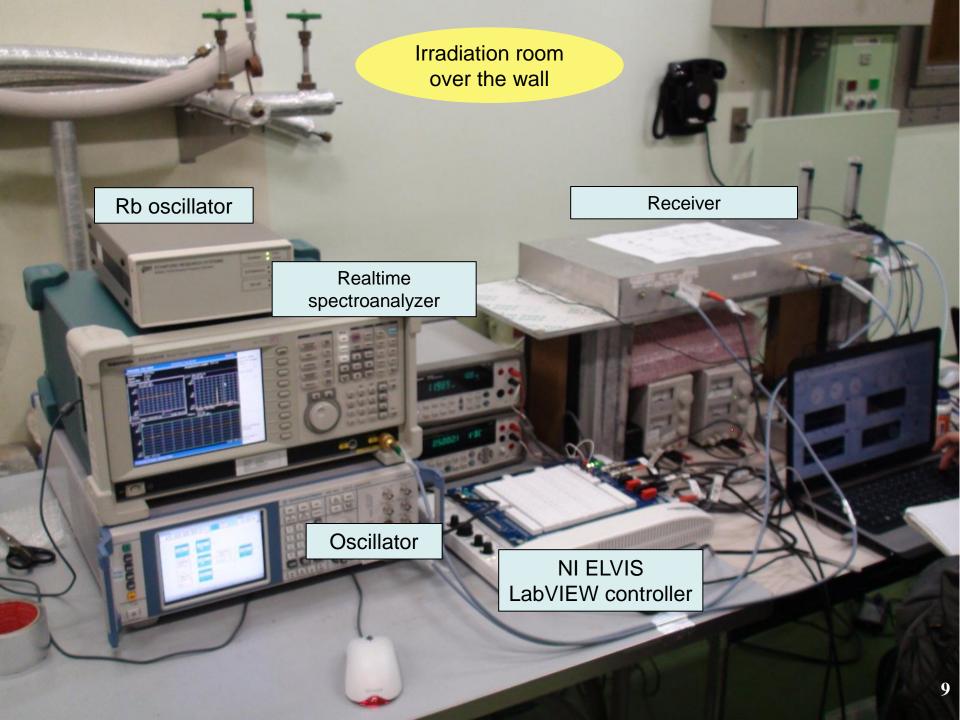
Before and at irradiation: automatic tuning ON

Tuned $\Delta\theta$ is recorded

$$\vec{c} = \vec{b}e^{-i(\pi + \Delta\theta)} + \vec{b}e^{-i\Delta\phi} = 0$$

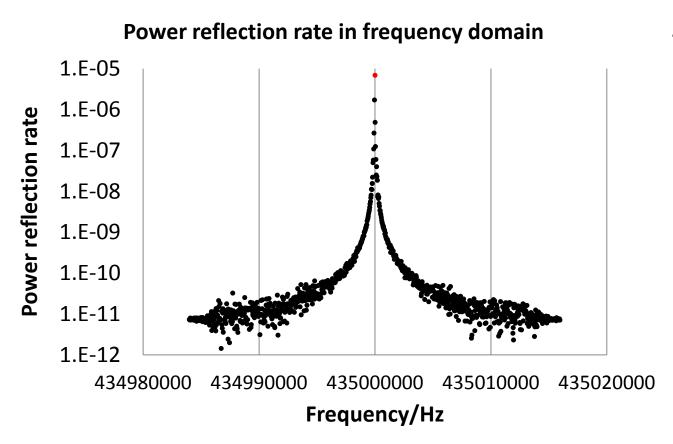
Reference Reflected wave

$$\Delta\theta = \Delta\phi$$
 then $\Delta\phi$ is measured



Frequency domain spectrum

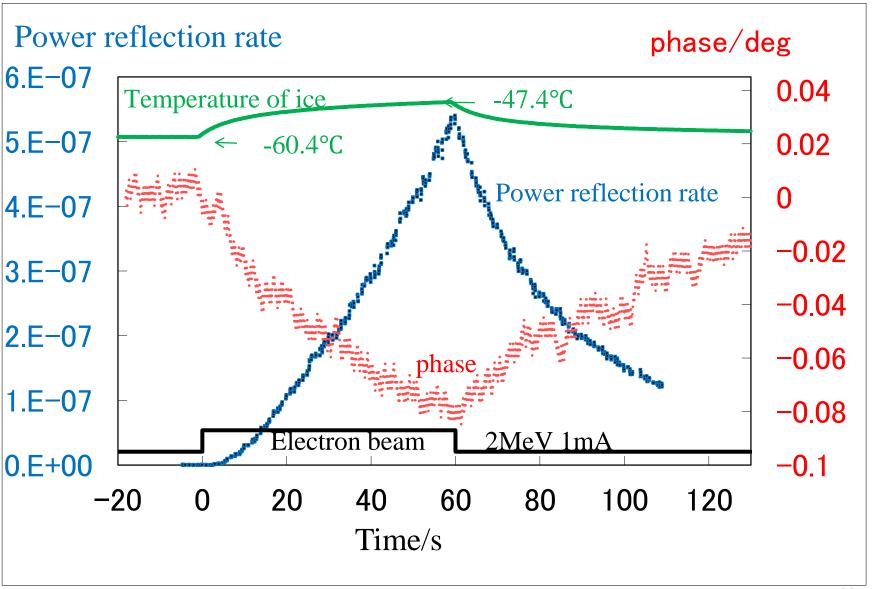
Rejection of noises from electric power source of 50 Hz.



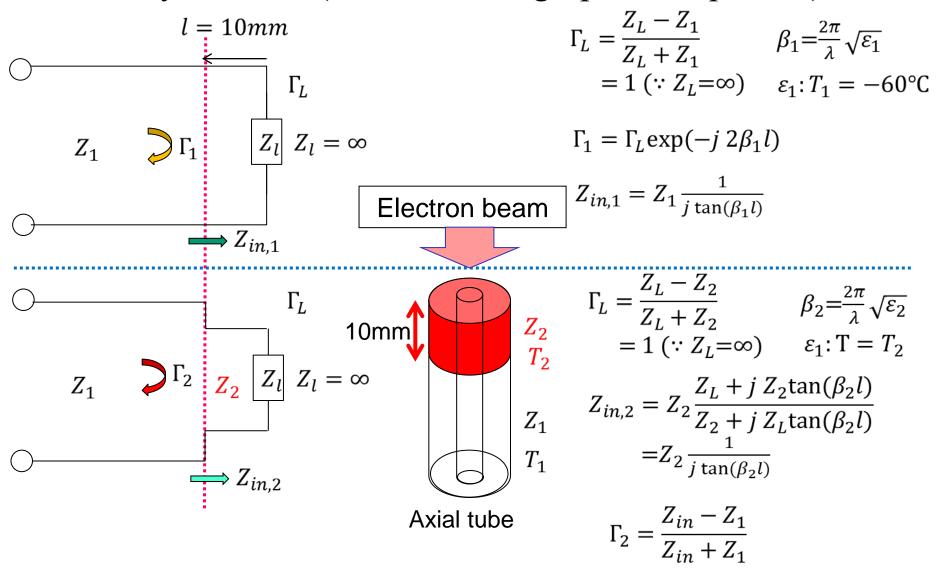
Data analysis of Real Time Spectrometer

- Time domain data of 1024points × 5 for 640 ms
- 1 datum for 128ms
- Conversion of time domain to frequency domain by FFT

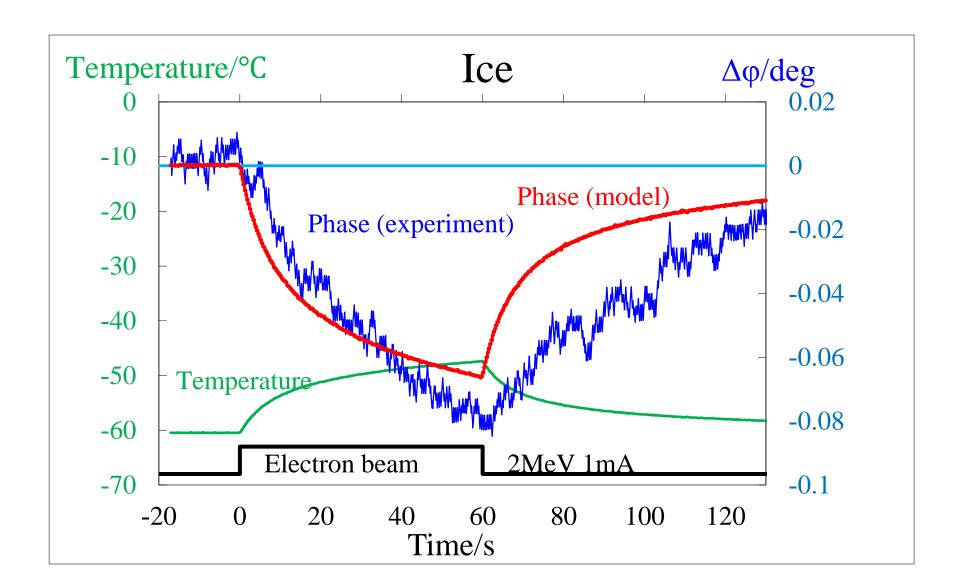
Reflection phase and rate in Ice



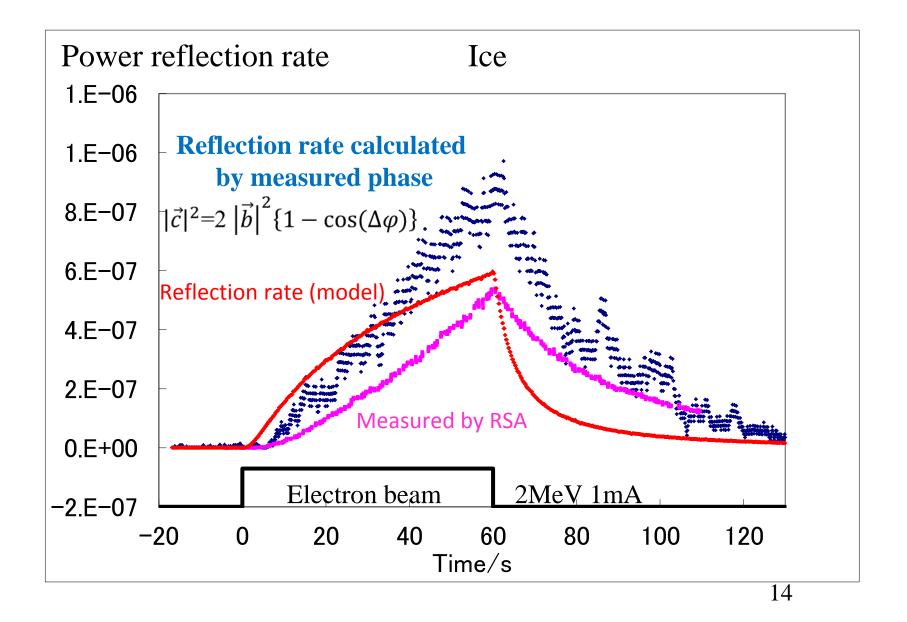
Calculation of phase difference of reflected wave by a model (based on Telegrapher's equation)



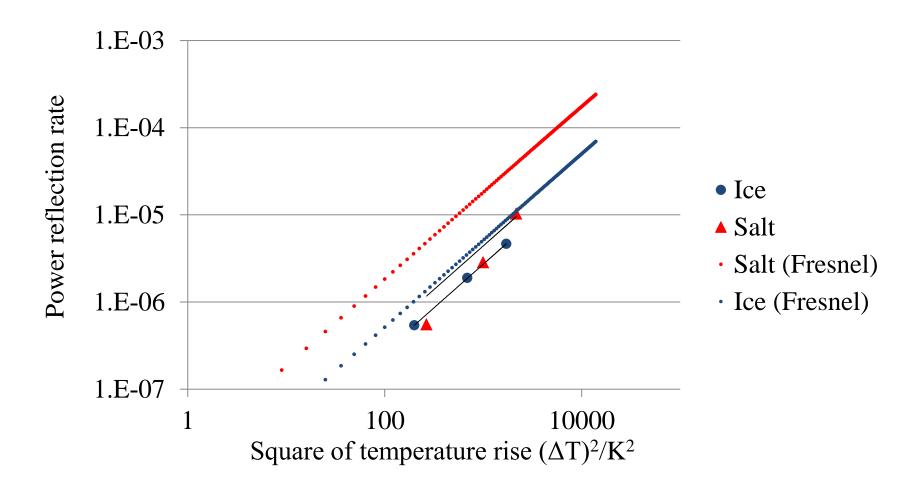
Comparison of phases between experiment and calculation



Reflection rate calculated by measured phase



Power reflection rate vs. square of temperature



Radiation detectors using heat effects.

	Bolometer	Cloud chamber	Bubble chamber	Acoustic detector	Radar chamber
Inventor	S.P. Langley	C.T.R. Wilson	D.A. Glaser	G.A. Askaryan	M. Chiba, et al.
Year	1878	1911	1952	1957	2007
Medium	Solid	Gas	Liquid	Solid, Liquid	Solid
Wave length	-	~500nm	~500nm	~1m	~10m
Body	Solid	Liquid particle,	Bubble,	Heated portion,	Heated portion,
Body size	-	\sim 0.5mm	\sim 0.1mm	$10\text{cm}\phi \times 5\text{m}$	$10\text{cm}\phi \times 5\text{m}$
Reflection or emission	-	Reflection	Reflection	Emission	Reflection
Operation	-	Decompression	Decompression	_	_
Process	Heating	Super cooling	Super heating	Heating	Heating
Amplification	Small heat capacity	Growth of liquid particle	Growth of bubble	_	Coherent reflection
Sensitivity	>1eV	\sim 100eV	~100eV	$> 10^{12} eV$	$>10^{12} eV$
Position reso.	-	\sim 0.5mm	~0.1mm	~30m	~30m
Detector size	$\sim 1 \mathrm{cm}$	\sim_{m}	\sim_{m}	\sim km	\sim km
Memory time	$\sim 1s$	\sim 10ms	\sim 1 μ s	_	~10s

Summary

- We found radio wave reflection from ice as well as rock salt with electron beam irradiation.
- •Constructing an automatic-negative-feedback circuit in which variable phase shifter was tuned to get the null output, we measured the phase of the reflected-radio wave.
- The reflected amplitude was reproduced by the measured phase and was explained by a model in which the temperature is taken into account.
- •Due to the long memory time, a stand-alone-GZK-neutrino detector could be constructed with a peak power of 1 GW (Equivalent Isotropic Radiation Power) radar supplied by a phased array antenna on a surface.
- •Rock salt formation, Antarctic ice sheet and the moon crust would be candidates for the detector media.
- A new radiation detector "Radar Chamber" is established utilizing for all dielectric media. It is not only for radiation detection but also for other purposes using materials with inhomogeneous refractive index in space and time.